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QCD thermalization: *Ab initio* approaches and interdisciplinary connections

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QCD thermalization: ab initio approaches and interdisciplinary connections

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Abstract

Heavy-ion collisions at BNL's Relativistic Heavy-Ion Collider (RHIC) and CERN's Large Hadron Collider (LHC) provide strong evidence for the formation of a quark-gluon plasma, with temperatures extracted from relativistic viscous hydrodynamic simulations shown to be well above the transition temperature from hadron matter. Outstanding problems in QCD include how the strongly correlated quark-gluon matter forms in a heavy-ion collision, its properties off-equilibrium, and the thermalization process in the plasma. We review here the theoretical progress in this field in weak coupling QCD effective field theories and in strong coupling holographic approaches based on gauge-gravity duality. We outline the interdisciplinary connections of different stages of the thermalization process to non-equilibrium dynamics in other systems across energy scales ranging from inflationary cosmology, to strong field QED, to ultracold atomic gases, with emphasis on the universal dynamics of non-thermal and hydrodynamic attractors. We survey measurements in heavy-ion collisions that are sensitive to the early non-equilibrium stages of the collision and discuss the potential for future measurements. We summarize the current state-of-the art in thermalization studies and identify promising avenues for further progress.

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I. BIG PICTURE QUESTIONS AND OUTLINE OF THE REVIEW

Ultrarelativistic collisions of heavy nuclei at the Relativistic Heavy-Ion Collider (RHIC) and the Large Hadron Collider (LHC) produce several thousand particles in each event generating the hottest and densest matter on Earth (Adams *et al.*, 2005; Adcox *et al.*, 2005; Alver *et al.*, 2007; Arsene *et al.*, 2005; Foka and Janik, 2016a,b; Muller *et al.*, 2012; Roland *et al.*, 2014). At the highest LHC energies, temperatures of the order of five trillion Kelvin are attained (Adam *et al.*, 2016). Temperatures on this scale only previously existed at the earliest instants of our universe, a 10th of a microsecond after the Big Bang. Lattice gauge theory studies (Bazavov *et al.*, 2019b) show strongly interacting matter at these temperatures to be well over a crossover temperature from hadron matter to a regime where the degrees of freedom describing bulk thermodynamic quantities are the fundamental quark and gluon fields of Quantum Chromodynamics (QCD). The results of experimental and theoretical studies indicate that shortly after the heavy-ion collision, the produced quark-gluon fields form a strongly correlated state of matter, widely known as the quark-gluon plasma (QGP) (Shuryak, 1980).

The heavy-ion experiments at RHIC and LHC therefore provide us with a unique opportunity to study terrestrially the spacetime evolution of this non-Abelian QGP. A striking finding from the RHIC and LHC experiments is that the experimental data are consistent with a description of the QGP as a nearly perfect fluid with a very low value of shear viscosity to entropy density ratio of $\eta/s \leq 0.2$ (in natural units) (Romatschke and Romatschke, 2019). These values are very close to $\eta/s = 1/(4\pi)$, a universal property of a class of gauge theories with a large number of degrees of freedom at infinite coupling (Buchel and Liu, 2004; Iqbal and Liu, 2009; Kovtun *et al.*, 2005; Policastro *et al.*, 2001) and described in terms of a dual gravity picture (Maldacena, 1998).

While our understanding of the thermal properties of QGP matter and the flow of the nearly perfect fluid has developed significantly, progress in theoretical descriptions of the

early stages of heavy-ion collisions is relatively recent. In particular, there is a growing realization that the far-from-equilibrium dynamics that characterizes early time physics is extremely important in understanding collective phenomena in the heavy-ion experiments (Busza *et al.*, 2018; Mrówczyński *et al.*, 2017; Schlichting and Teaney, 2019). This review summarizes our perspective on the theoretical and phenomenological progress in this active research area and places these developments in a wider interdisciplinary context.

The QCD thermalization process represents an initial value problem in quantum field theory (QFT). It requires understanding the many-body correlations in the colliding hadrons, how such correlations influence multi-particle production as the collision occurs, and the subsequent effective loss of information of these many-body correlations during the thermalization process of the matter produced. While von Neumann entropy is conserved in the unitary quantum evolution of a nuclear collision in isolation, observables of interest may nevertheless approach (local) thermal equilibrium. The characteristic time scales for the corresponding effective loss of information and the extent to which the dynamics finally leads to an approach to (local) thermal equilibrium for key observables is the central topic of this review.

In particular, we will focus on the following key questions prompted by the dynamics of each stage of the spacetime evolution of quark-gluon matter in heavy-ion collisions¹:

- *What are the many-body correlations of strongly interacting matter in the colliding nuclei?*

The colliding nuclei produce the initial state for the subsequent thermalization process. In principle, there can be different thermalization scenarios for different initial conditions. Though many details of the quantum evolution are lost rather quickly, it is crucial to classify the range of initial conditions (such as underoccupied versus overoccupied) leading to a certain class of dynamical processes.

In QCD, a proton (or any other nucleus) must be viewed as a collection of short or long lived configurations of partons (quarks, antiquarks, and gluons), where each configuration carries the quantum numbers of the proton. When the proton or nucleus is boosted to high energies, short lived configurations typically containing large numbers

¹ For a complementary perspective on open questions in heavy-ion collisions, we refer the reader to Ref. (Busza *et al.*, 2018).

of partons live much longer due to time dilation. It is therefore more likely that a probe of the hadron at high energies will scatter off such many-body configurations of partons and their decay dominate the physics of multi-particle production in ultrarelativistic nuclear collisions. How precisely multi-particle production occurs requires a deep knowledge of the spatial and momentum distributions of partons in the boosted nuclei, the nature of their correlations, and how these correlations change with system size and with collision energy.

- *What is the physics of the first yoctosecond (10^{-24} seconds) of the collision?*

Parton configurations in a boosted nucleus have their momenta distributed between a few fast modes and more plentiful soft modes. In a heavy-ion collision, these fast modes in each of the nuclei interact relatively weakly with the other nucleus and populate the “fragmentation regions” corresponding to polar angles very close to the beam axes (Van Hove and Pokorski, 1975). The slower degrees of freedom interact more strongly with each other and produce strongly interacting gluon matter outside the fragmentation regions.

This spacetime picture of nuclear collisions was developed in a seminal paper by Bjorken to describe the subsequent hydrodynamic flow of the quark-gluon plasma (Bjorken, 1983), albeit he did not address how thermalization occurs in this scenario. An interesting question in this regard is whether the strong interactions of the soft modes with each other is due to strong coupling or whether it can be due to the large occupancy of these soft modes. The answer to this question may also influence the degree of transparency of the fast modes, in particular a “limiting fragmentation” scaling phenomenon seen in data.

A spacetime scenario in which both soft and hard modes in the nuclei interact very strongly and generate hydrodynamic flow was suggested by Landau. It is conceivable that there is a transition between these two spacetime pictures with energy (Casalderrey-Solana *et al.*, 2013; Gelis *et al.*, 2006b); if so, can they be distinguished by phenomena such as limiting fragmentation (Gonçalves *et al.*, 2019)?

- *Is there a unifying theoretical description of quark-gluon matter off-equilibrium?*

The quark-gluon matter formed in the first few yoctoseconds of the heavy-ion collision

is very far from equilibrium. A key question in its description is whether weak and strong coupling extrapolations to realistic values can lead to similar phenomenology.

A potentially rich line of inquiry is to isolate what features of the non-equilibrium evolution of strongly correlated/coupled quark-gluon matter are universal. One example is universal dynamics in the approach to local thermal equilibrium governed by viscous hydrodynamics. Another example is universality in time dependence across a class of non-equilibrium states for certain observables. In a weak coupling scenario, at high occupancies, these include far-from-equilibrium attractors associated with non-thermal fixed points (Berges *et al.*, 2014b,c, 2008a).

Far-off-equilibrium hydrodynamic attractors are observed to emerge in both strong and weak coupling (Heller and Spalinski, 2015; Romatschke, 2018). A related important set of questions concerns the use of effective theories like hydrodynamics for systems far away from equilibrium. Yet another line of inquiry is to determine how features of the dynamics evolve between the weakly coupled and strongly coupled regimes. An intriguing possibility to consider is whether the topological properties of strongly correlated systems may help provide unifying descriptions at both weak and strong coupling.

- *Can we cleanly isolate signatures of quark-gluon matter off-equilibrium?*

If matter in bulk locally equilibrates in heavy-ion collisions, the only information of the non-equilibrium evolution that survives is what is imprinted as initial conditions for its subsequent hydrodynamic evolution. The exceptions are electroweak and so-called “hard probes”; both of these are sensitive to the full history of the spacetime evolution of QCD matter.

A significant development in recent years is the vastly improved ability of the RHIC and LHC experiments to perform “event engineering” whereby final states can be studied by varying the “control parameters” corresponding to nuclear size, centrality of collision impact and final state multiplicities (triggering thereby on typical versus rare event configurations) across a wide range in energy and system size (Schukraft *et al.*, 2013). A challenging question is whether we can constrain the current state-of-the-art computational techniques to accurately reflect the systematics of this event

engineering, and further, to use these to isolate empirically the out-of-equilibrium dynamics?

- *Interdisciplinary connections*

The study of the out-of-equilibrium dynamics of strongly correlated systems is an important topic of significant contemporary interest in a number of sub-fields of physics. As we will discuss, the ideas and methods outlined in this review have significant overlap with these fields. Can one exploit these interdisciplinary connections to make progress?

We will address the outstanding questions listed above in two *ab initio* theoretical approaches to the problem of thermalization in QCD. One approach, the Color Glass Condensate Effective Field Theory (CGC EFT), is applicable at very high energies corresponding to a regime of very weak coupling $\alpha_S \ll 1$ and very high gluon occupancies f_g satisfying $\alpha_S f_g \sim 1$. This regime of weak coupling and high occupancies in QCD is characterized by a large emergent “saturation” scale much larger than the intrinsic non-perturbative scales corresponding to color confinement and chiral symmetry breaking.

The CGC EFT employs weak coupling many-body methods to separate (or factorize) these soft non-perturbative modes from the harder modes on the order of the saturation scale. Specifically, the requirement that physics be independent of the scale separation between soft and hard modes leads to renormalization group equations that describe how such non-perturbative information provided as an input at a given energy scale changes as it evolves. As one approaches asymptotic energies, the factorization of the hard and soft scales becomes increasingly robust and many of the properties of quark-gluon matter can be computed systematically. The quark-gluon matter in this limit is called the Glasma ([Gelis and Venugopalan, 2006c](#); [Lappi and McLerran, 2006](#)).

The other *ab initio* approach to thermalization is in the limit of strong 't Hooft coupling of $\alpha_S N_c \rightarrow \infty$, as the number of colors $N_c \rightarrow \infty$. In this limit, holographic approaches based on gauge-gravity duality ([Gubser *et al.*, 1998a](#); [Maldacena, 1998](#); [Witten, 1998a](#)) are robust and can be used to obtain exact results in non-Abelian gauge theories, the best understood example being $\mathcal{N} = 4$ superconformal Yang-Mills theory.

Neither of these theoretical approaches to the problem of thermalization are directly applicable to real world heavy-ion collisions at RHIC and LHC energies, where the relevant

couplings are likely neither particularly weak nor infinitely strong. Thus data-theory comparisons rely on phenomenological descriptions characterized by extrapolations of *ab initio* approaches well beyond their strict regimes of validity. By anchoring such phenomenological models in fundamental theory in well controlled limits, their success or failure in comparisons to data can then be traced to a particular set of assumptions in the extrapolations. We will clarify throughout the review whenever such phenomenological extrapolations are made.

We will begin in Section II by discussing the structure of matter within the colliding hadrons and heavy nuclei at high energies. After a brief introduction to QCD, and the associated parton picture of hadrons at high energies, we will focus our attention on what happens when the phase space density of partons in the wavefunctions of the colliding hadrons becomes large. Driving this physics is an emergent energy-dependent close packing “saturation” scale Q_S (Gribov *et al.*, 1983), which grows with energy and nuclear size, allowing for a systematic weak coupling description of the properties of saturated partons in high energy QCD. Specifically, we will discuss the CGC EFT, wherein the high energy hadron is described as a coherent state of static color sources and dynamical gluon fields. The saturation scale is manifest in the CGC EFT, allowing one to describe strongly correlated many-body parton correlations in the hadron wavefunctions (Gelis *et al.*, 2010; Kovchegov and Levin, 2012).

Non-perturbative soft modes of the high energy nuclei, their color charge distributions, and many-body correlations thereof, are represented by a density matrix at a given energy scale much smaller than those of the hard weakly coupled modes. While this non-perturbative density matrix has to be parametrized at the initial scale by physically plausible assumptions, a renormalization group (RG) framework (Iancu *et al.*, 2001; Jalilian-Marian *et al.*, 1998b) allows one to study systematically the energy evolution of parton many-body correlations as the hadron is boosted to higher energies.

In Section III, we will outline the problem of multi-particle production in quantum field theory in the presence of strong fields and discuss how this leads to a first principles description of the very early time evolution of the Glasma. Inclusive quantities such as multiplicities or energy densities, and their spacetime correlations, can be computed systematically in the Glasma in powers of the coupling $\alpha_S \ll 1$ at sufficiently high energy. At leading order in this power counting, the Glasma fields are highly occupied classical fields, with magnitude

$1/\alpha_S$.

At next-to-leading order (NLO), we discuss how quantum fluctuations, co-moving with the colliding nuclei, can be absorbed into the density matrices describing their non-perturbative many-body color charge distributions, as discussed in the previous section. In contrast, non-comoving quantum fluctuations produced after the collision in the Glasma are unstable and display quasi-exponential dynamical growth (Romatschke and Venugopalan, 2006a). We will describe how the physics of these unstable modes at very early proper times $\tau \lesssim \frac{1}{Q_S} \log^2(1/\alpha_S)$ is captured in a classical-statistical approximation of the quantum evolution with given quantum initial conditions.

Section IV describes the non-linear time evolution of far-from-equilibrium quark gluon matter for weak couplings relevant at very high energies. The range of validity of classical-statistical field theory descriptions for the evolution is discussed in the terms of the two-particle-irreducible (2PI) quantum effective action, which motivates fully 3+1-dimensional numerical simulations of the expanding Glasma fields.

The lattice field theory simulations demonstrate the emergence of a non-thermal attractor described by a self-similar gluon distribution, whose dependence on momentum, and an overall cooling rate, are characterized by universal numbers independent of the initial conditions. Because the numerical simulations correctly describe dynamics in the infrared, the attractor solution helps to identify the right effective kinetic theory amongst several competing options.

Kinetic theory increasingly captures the relevant dynamics of the thermalization process as the system expands and cools. In Section V, we discuss the leading order kinetic theory framework, progressively from elastic $2 \leftrightarrow 2$ scatterings, to effective collinear $1 \leftrightarrow 2$ processes, taking special note of interference and plasma instability effects. Phenomenological extrapolations to realistic couplings can also be explored in the language of hydrodynamic attractors, where the dependence on the coupling is replaced with the kinematic viscosity η/s . For values of the kinematic viscosity extracted from hydrodynamic simulations of heavy-ion collision, reasonable predictions are obtained for entropy production (Giacalone *et al.*, 2019), as well as hydrodynamic and chemical equilibration times (Kurkela and Mazeliauskas, 2019a; Kurkela *et al.*, 2019b).

In Section VI, we provide an overview of holography based on strong coupling approaches to thermalization in gauge theories. Our focus is on the conceptual features, universal mech-

anisms, and predictions from these studies. In particular, *ab initio* holographic computations predict the applicability of hydrodynamics over a time scale set by the local energy density, when the expanding matter in heavy-ion collisions settings is characterized by a large spatial anisotropy in its energy-momentum tensor (Chesler and Yaffe, 2010, 2011; Heller *et al.*, 2012b). This is at variance with the common presumption of local thermal equilibrium in applying hydrodynamics; in a paradigm shift, the transition to hydrodynamic flow is now referred to as hydrodynamization rather than thermalization (Casalderrey-Solana *et al.*, 2014b).

We will discuss, in particular, phenomenological attempts to apply these ideas to model heavy-ion collisions in the context of 1+1-dimensional boost invariant flow where hydrodynamization and hydrodynamic attractors were first discovered. We will also cover work on more realistic holographic descriptions of heavy-ion collisions that model confinement, the breaking of conformal invariance, the running of the coupling, and large N_c suppressed non-local correlations.

Section VII is devoted to a discussion of signatures of non-equilibrium dynamics in heavy-ion data. While electromagnetic and high transverse momentum strongly interacting final states are sensitive to early time dynamics, significant contributions to their rates accrue from all stages of the spacetime evolution of the system. Measurements of long range correlations amongst high momentum final states, offer promise in isolating the early time non-equilibrium dynamics of the Glasma from the late stage hydrodynamic flow. This can be achieved by “event engineering” the response of these final states to variations in energy and system size. We will also discuss how bulk observables, in combination with these final states, can constrain thermalization scenarios.

A striking example of the role of topology in heavy-ion collisions is the Chiral Magnetic Effect (CME) (Kharzeev *et al.*, 2008) corresponding to a vector current along the direction of an external magnetic field that is induced by topological transitions. The CME is primarily an early time effect; in this case as well, event engineering of multi-particle correlations offers the possibility of uncovering its role.

In Section VIII, we will address the question of the interdisciplinary connections of the thermalization process in heavy-ion collisions to that of other strongly correlated systems across energy scales. A striking similarity of strongly correlated flow in heavy-ion collisions to that of unitary Fermi gases was already noted shortly after the discovery of the QGP

perfect fluid. The Glasma likewise shares common features with other overoccupied systems across energy scales, from inflationary dynamics in the early universe, to a quantum portrait of black holes as highly occupied graviton states, to those of overoccupied ultracold Bose gases.

A concrete example of the influence of interdisciplinary ideas is that of the turbulent thermalization process underlying the non-thermal attractor in the Glasma, which is widely discussed in the context of reheating in the early universe following inflation ([Berges *et al.*, 2008a](#); [Micha and Tkachev, 2004](#)). The latter in turn is in the perturbative high-momentum regime a relativistic generalization of weak wave turbulence in fluids ([Zakharov *et al.*, 2012](#)). In the non-perturbative infrared regime, the Glasma attractor is nearly identical to that of overoccupied cold atomic gases, sharing the same scaling functions and exponents in a wide spectral range ([Berges *et al.*, 2015b](#)). This is suggestive of a classification of far-from-equilibrium systems into universality classes analogous to those for critical phenomena close to equilibrium ([Hohenberg and Halperin, 1977](#)). An exciting development with cross-disciplinary potential is the use of state-of-the-art cold atom experiments to provide deep insight into such universal dynamics ([Erne *et al.*, 2018b](#); [Glidden *et al.*, 2020](#); [Prüfer *et al.*, 2018](#)).

The search for effective theories far from equilibrium is also a major research direction in the theory of complex systems ranging from understanding entanglement, information loss and thermalization of closed quantum many-body systems, with insights to be gained from “tabletop” atomic and condensed matter systems ([Eisert *et al.*, 2015](#)). On the other end of the energy scale, are the connections to black holes and string theory with respect to general questions regarding the scrambling of information ([Lashkari *et al.*, 2013](#); [Maldacena *et al.*, 2016](#)) and the unitary dynamics underlying black hole formation and evaporation ([Almheiri *et al.*, 2019, 2020b](#); [Hawking, 1974, 1976](#); [Page, 1993](#); [Penington, 2019](#)).

Finally, the role of topology in heavy-ion collisions has interdisciplinary connections in the chiral magnetic effect which is now observed in condensed matter systems ([Li and Kharzeev, 2016](#)). Continual advances in laser technology also offer great promise in the precision study of anomalous currents off-equilibrium.

We end the review in Section IX with a brief summary and outlook towards future developments in our understanding of thermalization in QCD. As the outline suggests, thermalization in QCD is a rich field with many research directions and we have had to make

choices in our presentation due to space limitations. An important topic we do not address is the off-equilibrium dynamics of QCD matter in the vicinity of a critical point (Akamatsu *et al.*, 2019; Bluhm *et al.*, 2020; Bzdak *et al.*, 2020). Another is the related topic of hydrodynamic fluctuations (Akamatsu *et al.*, 2017; An *et al.*, 2019). Other noteworthy omissions in our presentation include the discussion of holographic deep inelastic scattering (Hatta *et al.*, 2008; Polchinski and Strassler, 2003; Shuryak and Zahed, 2018), holographic hard probes (Chesler *et al.*, 2009a,b, 2013; Gubser, 2006; Herzog *et al.*, 2006) and features of linear response theory (Herzog and Son, 2003; Kovtun and Starinets, 2005; Son and Starinets, 2002). Some aspects of holographic approaches that we omit or treat only partially are discussed in (Casalderrey-Solana *et al.*, 2014b; Chesler and van der Schee, 2015; DeWolfe *et al.*, 2014; Florkowski *et al.*, 2018a; Heller, 2016).

II. HADRON STRUCTURE AT HIGH ENERGIES

The initial value problem of the thermalization process in hadron-hadron collisions requires a deep understanding of the structure of QCD matter in the wavefunctions of the colliding hadrons. The spacetime picture since the early days of QCD is that the highly Lorentz contracted large x valence partons in the ultrarelativistic hadron wavefunctions go through unscathed in the collision, while their accompanying small x “fur coat of wee-parton vacuum fluctuations” (Bjorken, 1976) interacts strongly to form hot and dense matter (Bjorken, 1983). The wee parton phase space distributions evolve with energy and nuclear size; their properties determine key features of the bulk properties of the matter produced after the collision.

In this section, after a brief introduction to QCD and the parton model of hadrons at high energies, we will discuss significant developments in the description of hadron wavefunction in the CGC EFT. In particular, we will motivate how the semi-hard saturation scale Q_S arises in the nuclear wavefunctions, which justifies their description as highly occupied gluon shockwaves. As the largest scale in the problem, it not only sets the scale for many-body correlations in these shockwaves, and in the Glasma matter produced after the collision, but subsequently also determines the thermalization time and the initial temperature of the quark-gluon plasma.

A. Quantum Chromodynamics

Quantum Chromodynamics (QCD), the modern theory of the strong force in nature, is a nearly perfect theory, the only free parameters being the quark masses (Wilczek, 2000). The Lagrangian of the theory can be written compactly as

$$\mathcal{L}_{\text{QCD}} = -\frac{1}{4}F_{\mu\nu}^a F^{\mu\nu,a} + \sum_f \bar{\Psi}_i^f (i\gamma^\mu D_{\mu,ij} - m_f \delta_{ij}) \Psi_j^f. \quad (1)$$

Here $F_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a - gf^{abc}A_\mu^b A_\nu^c$ is the QCD field strength tensor for the color gauge fields A_μ^a that live in the adjoint representation of $SU(3)$, with $a = 1, \dots, 8$ and f^{abc} are the structure constants of the gauge group. The quark fields live in the fundamental representation of $SU(3)$ and are labeled by their color and flavor indices Ψ_i^f where the color index $i = 1, \dots, 3$ and f denotes the flavors of quarks with masses m_f . Finally, the Dirac matrix γ^μ is contracted with the covariant derivative $D_{\mu,ij} = \partial_\mu \delta_{ij} + ig t_{ij}^a A_\mu^a$, with t_{ij}^a being the generators of $SU(3)$ in the fundamental representation.

The theory is rich in symmetry. The structure of the Lagrangian is dictated by the invariance of the quark and gluon fields under local $SU(3)$ color gauge transformations. In addition, for massless quarks, the theory has a global chiral $SU(3)_L \times SU(3)_R$ symmetry, global baryon number $U(1)_V$ and axial charge $U(1)_A$ symmetries, and the quark and gluon fields are invariant under scale transformations. Not least, the Lagrangian is invariant under discrete parity, charge and time reversal symmetries.

All of these symmetries, except that of local $SU(3)$ color, are broken by vacuum/quantum effects that give rise to all the emergent phenomena in the theory, such as confinement, asymptotic freedom, quantum anomalies and the spontaneous breaking of chiral symmetry.

Because QCD is a confining theory, it is not analytically tractable in general and numerical methods are essential to uncover its properties. Euclidean lattice Monte Carlo methods can be applied to compute, with good accuracy, “static” properties of the theory such as the mass spectrum of hadrons, magnetic moments, and thermodynamic properties of QCD at finite temperature (Detmold *et al.*, 2019; Lin *et al.*, 2018).

These methods are however very limited in determining dynamical “real time” features of theory because of the contributions of a large number of paths to the QCD path integral in Minkowski spacetime. There are promising approaches to surmount this difficulty such as steepest descent Lefschetz thimble methods but they are currently only applicable to

problems in 1+1-dimensions (Alexandru *et al.*, 2017). Likewise, quantum computing offers an alternative paradigm to compute real time dynamics, but its applicability to QCD likely remains far in the future (Preskill, 2018).

One should note that the production of high transverse momentum and massive particles (jets, heavy quarkonia being two notable examples) can be computed with high precision in perturbative QCD (pQCD) (Collins *et al.*, 1989). This is because these processes correspond to very short transverse distances and asymptotic freedom tells us that the QCD coupling α_S is weak at these scales.

B. QCD at small x and high parton densities

A great success of pQCD is the QCD parton model (Bjorken and Paschos, 1969), wherein the complex dynamics of quark and gluon fields in hadrons can, at high energies and large momentum resolutions, be viewed as that of a weakly interacting gas of partons (single-particle quark, antiquark and gluon states). The cleanest way to access this sub-nucleon structure is in the deeply inelastic scattering (DIS) of electrons or other leptons off nucleons and nuclei, wherein a virtual photon emitted by the electron strikes a quark or antiquark inside the hadron.

For the thermalization process of interest in this review, the asymptotic high energy (or “Regge”) limit of DIS is most relevant. This limit corresponds to the Bjorken DIS variable $x_{\text{Bj}} \sim Q^2/s \rightarrow 0$ where Q^2 is the squared four-momentum transfer and s is the squared center-of-mass energy. In the parton model, $x_{\text{Bj}} \approx x$, where x is the light cone fraction of the momentum of the hadron carried by the struck parton². At small x , or equivalently at high energies, the number of partons in the hadron proliferate rapidly, as first observed in DIS experiments at the HERA collider in Germany (Abt *et al.*, 1993; Ahmed *et al.*, 1995; Derrick *et al.*, 1993, 1995; Lai *et al.*, 1995; Martin *et al.*, 1994). This growth is consistent with the predictions of the pQCD DGLAP (Altarelli and Parisi, 1977; Dokshitzer, 1977; Gribov and Lipatov, 1972; Lipatov, 1975) evolution equations.

The mathematical basis of the parton model in QCD follows from the observation that if one picks a lightcone³ gauge $A^+ = 0$, and quantizes the quark and gluon fields of QCD along

² In hadron-hadron collisions, it is more appropriate to speak in terms of momentum fractions, so we shall henceforth use x instead of x_{Bj} .

³ Lightcone coordinates are $k^\pm = (k^0 \pm k^3)/\sqrt{2}$ and lightcone fields are defined as $A^\pm = (A^0 \pm A^z)/\sqrt{2}$; we

a light front surface (say, $x^+ = 0$), the Hamiltonian of free quark and gluon fields shares the same vacuum⁴ as the fully interacting theory (Brodsky *et al.*, 1998). This allows one to construct the hadron wavefunction as a linear combination of a complete set of multi-parton eigenstates, each of which is an eigenstate of the free QCD Hamiltonian.

In this lightcone framework, the parton distribution functions measured in DIS experiments can be interpreted as one-body states of quarks and gluons that carry a lightcone momentum fraction $x = k^+/P^+$, where k^+ is the parton’s lightcone momentum and P^+ the lightcone momentum of the hadron. As first argued in (Gribov *et al.*, 1983; Mueller and Qiu, 1986), two-body “higher twist” gluon distributions, in a lightcone operator product expansion⁵ (OPE), grow as the square $(xG_A(x, Q^2))^2$ of the leading twist gluon distribution. For a fixed Q^2 , these two-body distributions become as large as the leading twist one-body distribution as $x \rightarrow 0$.

Importantly, the net effect of such many-body contributions⁶ is opposite to that of the leading term, softening the growth in the gluon distribution. When the gluon phase space density is maximal, of order $1/\alpha_S$, all n -body lightcone distributions contribute equally. This saturation of gluon distributions in a nucleus of radius R_A , corresponds to the generation of the saturation scale Q_S , where parametrically, for $Q^2 = Q_S^2$ the maximal occupancy is equated to the gluon phase space density as

$$\frac{1}{\alpha_S(Q_S)} = \frac{xG_A(x, Q_S^2)}{2(N_c^2 - 1)\pi R_A^2 Q_S^2}. \quad (2)$$

Fig. 1 illustrates the gluon saturation phenomenon and the interpretation of Q_S as the emergent “close packing” scale.

C. Effective Field Theory for high parton densities: the Color Glass Condensate

Since the usual formalism of pQCD relies on two-body and higher twist distributions being small, an alternative framework is necessary to understand the physics of gluon saturation and the emergence of the saturation scale in the nuclear wavefunction at high energies.

⁴ will work in the metric $g_{\pm, \mp} = 1; g_{i, j} = -1$, where i, j represent the two transverse coordinates.

⁵ In lightcone quantization, this argument requires a careful treatment of $k^+ = 0$ vacuum modes (Nakanishi and Yamawaki, 1977). For a perturbative treatment of lightcone wavefunctions, it may be sufficient to project out such modes (Collins, 2018; Fitzpatrick *et al.*, 2018).

⁶ In OPE language, these higher twist contributions are suppressed by powers of $1/Q^2$.

⁶ These include the screening of bremsstrahlung gluons by real and virtual gluons, and the recombination of softer gluons into harder gluons.

Fortuitously, the problem of high parton densities can be formulated as a classical effective field theory on the light front, which as noted, greatly simplifies the problem of heavy-ion collisions at high energies.

To understand this better, we will outline here an explicit construction performed for nuclei with large atomic number $A \gg 1$ (McLerran and Venugopalan, 1994a,b,c). An important ingredient in this construction (in the infinite momentum frame (IMF) $P^+ \rightarrow \infty$ of the nucleus) is a Born-Oppenheimer separation in time scales between the Lorentz contracted large x ($k^+ \sim P^+$) “valence” modes and the noted “wee fur” of small x ($k^+ \ll P^+$) gluons and “sea” quark-antiquark pairs. For partons of transverse momentum k_\perp , their lightcone lifetimes are given by

$$\begin{aligned} \tau_{\text{wee}} &= \frac{1}{k^-} = \frac{2k^+}{k_\perp^2} \equiv \frac{2xP^+}{k_\perp^2} \quad \text{and} \\ \tau_{\text{valence}} &\approx \frac{2P^+}{k_\perp^2} \longrightarrow \tau_{\text{wee}} \ll \tau_{\text{valence}}, \end{aligned} \quad (3)$$

suggesting that the valence parton modes are static over the times scales over which wee modes are probed. However one cannot integrate out the valence sources completely out of the theory because they are sources of color charge for wee partons and must couple to these in a gauge invariant manner.

Note further that since wee partons have large lightcone wavelengths ($\lambda_{\text{wee}} \sim 1/k^+ = 1/xP^+$), they can resolve a lot of color charge provided their transverse wavelength is not too large. The inequality

$$\lambda_{\text{wee}} \sim \frac{1}{k^+} \equiv \frac{1}{xP^+} \gg \lambda_{\text{valence}} \equiv \frac{R_A m_N}{P^+}, \quad (4)$$

where on the r.h.s the Lorentz contraction factor is P^+/m_N (with m_N the nucleon mass), suggests that wee partons with $x \ll A^{-1/3}$ resolve partons⁷ all along the longitudinal extent $2R_A \sim A^{1/3}$ in units of the inverse nucleon mass.

These charges will be random since they are confined to different nucleons and do not know about each other. A wee parton with momentum k_\perp resolves an area in the transverse plane $(\Delta x_\perp)^2 \sim 1/k_\perp^2$. The number of valence partons it interacts simultaneously with is

$$k \equiv k_{(\Delta x_\perp)^2} = \frac{N_{\text{valence}}}{\pi R_A^2} (\Delta x_\perp)^2, \quad (5)$$

⁷ Wee partons with wavelength $k_\perp \leq \Lambda_{\text{QCD}} \sim 1 \text{ fm}^{-1}$, see no color charge at all since color is confined (in nucleons!) on this scale. It is only wee partons with $k_\perp \gg \Lambda_{\text{QCD}}$ that see color charges from different nucleons along the longitudinal direction.

which is proportional to $A^{1/3}$ since $N_{\text{valence}} = 3 \cdot A$ in QCD. For a large nucleus with $k \gg 1$, one can show for $N_c \geq 2$ that the most likely color charge representation that the wee gluons couple to is a higher dimensional classical representation of order \sqrt{k} (Jeon and Venugopalan, 2004).

Thus wee partons couple to ρ , the classical color charge per unit transverse area of large x sources. On average, since the charge distributions are random, the wee partons will couple to zero charge; however, fluctuations locally can be large. These conditions can be represented as

$$\langle \rho^a(x_\perp) \rangle = 0; \langle \rho^a(x_\perp) \rho^b(y_\perp) \rangle = \mu_A^2 \delta^{ab} \delta^{(2)}(x_\perp - y_\perp), \quad (6)$$

where $a = 1, \dots, N_c^2 - 1$ and $\mu_A^2 = \frac{g^2 A}{2\pi R_A^2}$ is the color charge squared per unit area. For a large nucleus ($A \gg 1$), $\mu_A^2 \propto A^{1/3} \gg \Lambda_{\text{QCD}}^2$ is a large scale. Since it is the largest scale in the problem, $\alpha_S(\mu_A^2) \ll 1$. This result is remarkable because it provides a concrete example suggesting that QCD at small x is a weakly coupled EFT wherein systematic computations of its many-body properties are feasible.

We can now put together the kinematic and dynamical arguments above and write down the generating functional for the small x effective action as

$$\mathcal{Z}[j] = \int [d\rho] W_{\Lambda^+}[\rho] \left\{ \frac{\int^{\Lambda^+} [dA] \delta(A^+) e^{iS_{\Lambda^+}[A, \rho] - \int j \cdot A}}{\int^{\Lambda^+} [dA] \delta(A^+) e^{iS_{\Lambda^+}[A, \rho]}} \right\}. \quad (7)$$

Here Λ^+ denotes the longitudinal momentum scale that separates the static color sources from the dynamical gauge fields and the gauge invariant weight functional $W_{\Lambda^+}[\rho]$ describes the distribution of these sources at the scale Λ^+ , with its path integral over ρ normalized to unity.

The CGC effective action can be written in terms of the sources ρ and the fields A as

$$S_{\Lambda^+}[A, \rho] = \frac{1}{4} \int d^4x F_{\mu\nu}^a F^{\mu\nu, a} + \frac{i}{N_c} \int d^2x_\perp dx^- \delta(x^-) \text{Tr}(\rho U_{-\infty, \infty}[A^-]). \quad (8)$$

The first term here is the Yang-Mills action in the QCD Lagrangian given in Eq. (1). The dynamics of wee gluons in the CGC is specified by this term. The second term⁸ denotes the coupling of the wee gluon fields to the large x color charge densities ρ , which we have argued

⁸ This term can alternatively (Jalilian-Marian *et al.*, 2001) be written as $\text{Tr}(\rho \log(U_{-\infty, \infty}))$.

are static lightcone sources. Because the sources are eikonal sources along the lightcone, their gauge invariant coupling to the wee fields is described by the path ordered exponential along the lightcone time direction $U_{-\infty,\infty} = \mathcal{P} \exp \left(ig \int dx^+ A^{-,a} T^a \right)$. Physically, U corresponds to the color rotation of the color sources in the background of wee gluon fields.

The weight functional in the effective action (for the Gaussian random color charges in Eq. (6)), in what is now called the McLerran-Venugopalan (MV) model (Kovchegov, 1999; McLerran and Venugopalan, 1994a,b), can equivalently be written as⁹

$$W_{\Lambda^+}[\rho] = \exp \left(- \int d^2 x_{\perp} \frac{\rho^a(x_{\perp}) \rho^a(x_{\perp})}{2\mu_A^2} \right). \quad (9)$$

For each configuration of ρ 's in Eq. (7), the saddle point of the effective action is given by the Yang-Mills (YM) equations:

$$D_{\mu} F^{\mu\nu,a} = \delta^{\nu+} \delta(x^-) \rho^a(x_{\perp}), \quad (10)$$

whose solution is the non-Abelian analog of the Weizäcker–Williams (WW) fields in classical electrodynamics. The chromo-electromagnetic gluon field strengths are singular on the nuclear sheet of width $\Delta x^- \sim 2Rm_N/P^+$ and zero (pure gauge) outside.

The gauge field solutions in lightcone gauge are given by $A^- = 0$ and

$$A_{\text{cl}}^k = \frac{1}{ig} V(x^-, x_{\perp}) \nabla^k V^{\dagger}(x^-, x_{\perp}), \quad (11)$$

where $k = 1, 2$ are the transverse coordinates and $V = \mathcal{P} \exp \left(\int_{-\infty}^{x^-} dz^- \frac{1}{\nabla_{\perp}^2} \tilde{\rho}(z^-, x_{\perp}) \right)$. This solution of the equations of motion requires path ordering of the sources in x^- (Jalilian-Marian *et al.*, 1997; Kovchegov, 1999). Further, $\tilde{\rho}$ that appears in the solution is the color charge density in Lorenz gauge $\partial_{\mu} A'^{\mu} = 0$, where one has the solution $A'_{\text{cl}}{}^+ = \frac{1}{\nabla_{\perp}^2} \tilde{\rho}(x^-, x_{\perp})$, $A'_{\text{cl}}{}^- = A'_{\text{cl}\perp} = 0$. In fact, since the Jacobian of the transformation $[d\rho] \rightarrow [d\tilde{\rho}]$ is simple (Jalilian-Marian *et al.*, 1997), many-body distributions in lightcone gauge can be computed in terms of color charges in Lorenz gauge, a natural choice from the analogy to WW fields (Jackson, 1998).

As a simple example, the correlator of gauge fields in a large nucleus can be computed analytically in the MV model by averaging the solution in Eq. (11) with the weight functional W ,

$$\langle AA \rangle = \int [d\tilde{\rho}] A_{\text{cl}}.[\tilde{\rho}] A_{\text{cl}}.[\tilde{\rho}] W_{\Lambda^+}[\tilde{\rho}]. \quad (12)$$

⁹ Sub-leading terms are discussed in (Dumitru and Petreska, 2012; Jeon and Venugopalan, 2005).

This expression can be further Fourier decomposed to extract the number distribution of wee gluons $\frac{dN}{d^2k_\perp}$ and expressed¹⁰ in terms of Q_S . Specifically, for the occupation number $\phi = \frac{(2\pi)^3}{2(N_c^2-1)} \frac{dN}{\pi R^2 d^2k_\perp}$, one obtains $\phi \propto \frac{Q_S^2}{k_\perp^2}$ for $k_\perp \gg Q_S$, However for $k_\perp \ll Q_S$, the distribution is modified substantially from the WW distribution: $\phi \sim \frac{1}{\alpha_S} \log(Q_S/k_\perp)$. This softened infrared distribution in the CGC EFT provides a simple explanation of gluon saturation.

We are now in a position to understand the term Color Glass Condensate (CGC) (Gelis *et al.*, 2010; Iancu and Venugopalan, 2003) used to describe the ground state properties of a hadron/nucleus at very high energies. Color is obvious since the state is comprised primarily of a large number of gluons and sea quark-antiquark pairs. It is a glass because these small x gluons and sea quarks are generated by random sources with lifetimes much longer than the characteristic time scales of the scattering. This explains the structure of Eq. (7), where the path integral over the curly brackets is performed first for fixed color charge distributions and then averaged over an ensemble of such distributions. Finally, the state is a condensate because gluons have occupation numbers $\phi \sim 1/\alpha_S$, with momenta peaked at $k_\perp \sim Q_S$.

To take a specific example, consider the inclusive cross-section in the DIS scattering of a virtual photon on the nucleus, illustrated in Fig. 2. In the CGC EFT, it is expressed as the cross-section for a fixed distribution of sources convoluted with an ensemble of such sources:

$$\langle d\sigma \rangle = \int [\mathcal{D}\tilde{\rho}_A] W_{A^+}[\tilde{\rho}_A] d\hat{\sigma}[\tilde{\rho}_A]. \quad (13)$$

Thus on the time scale $t \sim 1/Q$ of the probe, it resolves a condensate of gluons with a well-defined number density of longitudinal modes down to $x \sim x_{Bj} \ll 1$. Due to time dilation (see Eq. (3)) the averaging over ρ_A with W takes place on much larger time scales. This two-stage averaging process clarifies how one reconciles gauge invariance with the presence of a colored condensate.

The CGC classical equations possess a ‘‘color memory’’ effect (Pate *et al.*, 2017) corresponding to the large gauge transformation V of a quark after interacting with the gluon shockwave, generating a transverse momentum kick $p_\perp \sim Q_S$ to the quark that can be measured in DIS experiments (Ball *et al.*, 2019). Remarkably, this is exactly analogous to the inertial displacement of detectors after the passage of a gravitational shockwave (Strominger and Zhiboedov, 2016). This gravitational memory is deeply related to asymptotic spacetime

¹⁰ In the MV model, this defines $Q_S^2 = c_A \mu_A^2$, where the coefficient c_A is determined numerically (Lappi, 2008).

symmetries and soft theorems in gravity and may also hold useful lessons for QCD¹¹.

D. Renormalization group evolution in the CGC EFT

We discussed thus far a classical EFT for large nuclei and Gaussian sources where the separation between fields (wee partons) and sources (valence sources) was picked randomly to be at the momentum scale Λ^+ . Physical observables such as the inclusive cross-section in Eq. (13) should not depend on Λ^+ . This invariance is the essence of the renormalization group and we will sketch below how it is realized in the EFT; a detailed discussion can be found in (Iancu and Venugopalan, 2003).

The important point to note is that real and virtual quantum fluctuations in the classical background field of the target, while apparently suppressed by $\mathcal{O}(\alpha_S)$ are actually $\alpha_S \log(\Lambda^+/\Lambda'^+)$ $\sim \mathcal{O}(1)$ from the phase space integration of these modes when $\Lambda'^+ = \Lambda^+ e^{-1/\alpha_S}$ (or equivalently, when $x_{\text{wee}} = x_{\text{val.}} e^{-1/\alpha_S}$). These large NLO contributions can be absorbed into the form of the LO cross-section in Eq. (13) at the scale Λ'^+ by redefining the weight functional $W_{\Lambda^+}[\tilde{\rho}] \rightarrow W_{\Lambda'^+}[\tilde{\rho}']$. Here $\tilde{\rho}' = \tilde{\rho} + \delta\tilde{\rho}$, is new color source density at Λ'^+ that incorporates the color charge density $\delta\tilde{\rho}$ induced by quantum fluctuations between Λ^+ and Λ'^+ .

One can thus write

$$\langle d\sigma_{\text{LO+NLO}} \rangle = \int [\mathcal{D}\tilde{\rho}_A] W_{\Lambda'^+}[\tilde{\rho}_A] d\hat{\sigma}_{\text{LO}}[\tilde{\rho}_A], \quad (14)$$

where

$$W_{\Lambda'^+}[\tilde{\rho}_A] = \left(1 + \log(\Lambda^+/\Lambda'^+) \mathcal{H}_{\text{LLx}}\right) W_{\Lambda^+}[\tilde{\rho}_A], \quad (15)$$

with the quantum fluctuations absorbed into we shall discuss further shortly.

Since the l.h.s of Eq. (14) should not depend on the arbitrary “factorization scale” Λ^+ , the derivative of both l.h.s and r.h.s with respect to it should be zero. From Eq. (15), one can therefore deduce the JIMWLK¹² RG equation (Iancu *et al.*, 2001; Jalilian-Marian *et al.*, 1998a,b)

$$\frac{\partial}{\partial Y} W_Y[\tilde{\rho}_A] = \mathcal{H}_{\text{LLx}} W_Y[\tilde{\rho}_A], \quad (16)$$

¹¹ An “infrared triangle” between asymptotic symmetries, memory and soft theorems in gravity (Strominger, 2017) also allows for an elegant interpretation of the infrared structure of QED (Bieri and Garfinkle, 2013; Kapec *et al.*, 2017). While color confinement implies such universal features do not apply to QCD in general, an emergent $Q_S \gg \Lambda_{\text{QCD}}$ suggests that they may be applicable in the Regge limit.

¹² JIMWLK is an acronym denoting the last names of the principal authors.

where the JIMWLK Hamiltonian (Weigert, 2002)

$$\mathcal{H}_{\text{LLx}} = \frac{1}{2} \int_{x_\perp, y_\perp} \frac{\delta}{\delta \tilde{\rho}^a(x_\perp)} \chi^{ab}(x_\perp, y_\perp) [\tilde{\rho}] \frac{\delta}{\delta \tilde{\rho}^b(y_\perp)}, \quad (17)$$

describes the evolution of the gauge invariant weight functional W with rapidity $Y = \log(\Lambda_0^+/\Lambda^+) \equiv \log(x_0/x)$, once the non-perturbative initial conditions for W are specified at an initial x_0 .

The Hamiltonian is computed in the CGC EFT, with $\chi^{ab}(x_\perp, y_\perp) [\tilde{\rho}] = \langle \delta \tilde{\rho}^a(x_\perp) \delta \tilde{\rho}^b(y_\perp) \rangle_{\tilde{\rho}}$ the two-point function of induced charge densities¹³ in the classical background field of the hadron. Note that with this computation of \mathcal{H}_{LLx} , the solution of Eq. (16) resums leading logarithms $\alpha_S \log(x_0/x)$ (LLx) to all orders in perturbative theory. Thus this powerful RG procedure extends the accuracy of computations of the cross-section from $\langle d\sigma_{\text{LO+NLO}} \rangle \rightarrow \langle d\sigma_{\text{LO+LLx}} \rangle$.

The JIMWLK RG equation can equivalently be expressed as a hierarchy of equations (the Balitsky-JIMWLK hierarchy independently derived in (Balitsky, 1996)) for the expectation value of an operator O :

$$\frac{\partial \langle O \rangle_Y}{dY} = \left\langle \frac{1}{2} \int_{x_\perp, y_\perp} \frac{\delta}{\delta \alpha^a(x_\perp)} \chi^{ab}(x_\perp, y_\perp) \frac{\delta}{\delta \alpha^b(y_\perp)} O[\alpha] \right\rangle_Y, \quad (18)$$

where $\alpha^a = \frac{1}{\nabla_\perp^2} \tilde{\rho}^a$. Remarkably, Eq. (18) has the form of a generalized Fokker-Planck equation in functional space, where Y is “time” and χ is the diffusion coefficient (Weigert, 2002).

There is no known analytical solution to the JIMWLK equation; as we shall discuss, it can be solved numerically. However good approximations exist in different limits. In a “weak field” (and leading twist) limit $g\alpha \ll 1$, one recovers for the number distribution (and the corresponding occupation number ϕ) extracted from Eq. (12), the celebrated LLx BFKL equation (Balitsky and Lipatov, 1978; Kuraev *et al.*, 1977) of pQCD. Another mean field “random phase” approximation (Iancu and McLerran, 2001; Weigert, 2002) allows one to evaluate the occupation number ϕ in the “strong field” limit of $g\alpha \sim 1$.

The longitudinal extent of the wee gluon cloud generated by the RG evolution has a width $x^- = \frac{1}{k^+} \sim \frac{1}{Q_S}$. This is much more diffuse relative to the width $e^{-1/\alpha_S} \frac{1}{Q_S}$ of valence modes. The RG evolution also predicts that the width of the wee gluon cloud shrinks with

¹³ Note that here and henceforth in this section, $\int_{x_\perp} = \int d^2x_\perp$ and $\int_{x_\perp, y_\perp} = \int d^2x_\perp d^2y_\perp$.

increasing boost (or rapidity) relative to an “observer” quark-antiquark pair, albeit at a slower rate than their larger x counterparts. Thus in the CGC EFT the scale for the overlap of the wave functions in the thermalization process is set by $\frac{1}{Q_s}$ rather than the the Lorentz contracted width of valence quarks, given by $\frac{1}{P^+}$.

E. DIS and the dipole model

In this sub-section, and the next, we will concretely relate the CGC EFT to the structure functions that are measured in DIS. These comparisons are essential for precision tests of the CGC EFT picture of high energy nuclear wavefunctions. They also play an important role in constraining the saturation scale and the shadowing of nuclear distributions that are key to determining the initial conditions for early time dynamics in heavy-ion collisions. These connections will become more evident in Section III C.

The inclusive cross-section can be expressed in full generality as $\langle d\sigma \rangle = L_{\mu\nu} W^{\mu\nu}$ where $L_{\mu\nu}$ is the well-known lepton tensor (Peskin and Schroeder, 1995) representing the squared amplitude for the emission of a virtual photon with four-momentum q^μ and $W^{\mu\nu}$ is the spin-averaged DIS hadron tensor which, for a nucleus in the IMF can be reexpressed as (McLerran and Venugopalan, 1999; Venugopalan, 1999)

$$W^{\mu\nu} = \frac{1}{2\pi} \frac{P^+}{m_N} \text{Im} \int d^2 X_\perp dX^- \int d^4 x e^{iq \cdot x} \times \langle \text{Tr}(\gamma^\mu S_A(X^- + \frac{x}{2}, X^- - \frac{x}{2}) \gamma^\nu S_A(X^- - \frac{x}{2}, X^- + \frac{x}{2})) \rangle, \quad (19)$$

where $S_A(x, y) = -i \langle \psi(x) \bar{\psi}(y) \rangle_A$ is the quark propagator in the gauge fields A^μ of the nucleus¹⁴.

In the CGC, the leading contribution is obtained by replacing the full QCD background field by the saturated classical background field: $A^\mu \rightarrow A_{\text{cl}}^\mu$, where A_{cl}^μ are the non-Abelian WW fields in Eq. (11). In $A^- = 0$ gauge¹⁵, the momentum space quark propagator in the classical background field is remarkably simple, given by (McLerran and Venugopalan, 1999) $S_{A_{\text{cl}}}(p, q) = S_0(p) \mathcal{T}_q(p, q) S_0(q)$, where the free Dirac propagator is $S_0 = \frac{i \not{p}}{p^2 + i\epsilon}$ and $\mathcal{T}_q(q, p) = \pm (2\pi) \delta(p^- - q^-) \gamma^- \int_{z_\perp} e^{-i(\mathbf{q}_\perp - \mathbf{p}_\perp) \cdot \mathbf{z}_\perp} V^{\pm 1}(\mathbf{z}_\perp)$ is the effective vertex corresponding

¹⁴ The second average in Eq. (19) corresponds to averaging over $\bar{\rho}$. We employ the relativistic normalization

$\langle P|P \rangle = \frac{P^+}{m_N} (2\pi)^3 \delta^3(0) \equiv \frac{P^+}{m_N} \int d^2 X_\perp dX^-$, where X^-, X_\perp are center-of-mass coordinates.

¹⁵ The solution of the YM equations is identical in this case to the solution in Lorenz gauge.

to the multiple scattering of the quark (or antiquark) off the shockwave background field represented by the eikonal path ordered phase V introduced previously after Eq. (11).

The DIS structure function is simply related to the inclusive cross-section. Plugging the dressed CGC propagator into Eq. (19), one can show, to this order of accuracy, that it can be expressed as (McLerran and Venugopalan, 1999),

$$F_2(x, Q^2) = \frac{Q^2}{4\pi^2\alpha_{\text{em}}} \int_0^1 dz \int_{r_\perp} | \Psi_{\gamma^* \rightarrow q\bar{q}} |^2 \sigma_{q\bar{q}A}(x, Q^2). \quad (20)$$

This expression can be simply interpreted to be the convolution of the probability of the virtual photon to split into a quark-antiquark pair (which can be computed in QED (Bjorken *et al.*, 1971)) with the ‘‘dipole’’ scattering cross-section of the quark-antiquark pair to scatter off the nucleus. For impact parameter $b_\perp = (x_\perp + y_\perp)/2$, it is given by

$$\sigma_{q\bar{q}A} = 2 \int d^2b_\perp \mathcal{N}_Y(b_\perp, r_\perp), \quad (21)$$

where the forward scattering amplitude $\mathcal{N}_Y(b_\perp, r_\perp) = 1 - \mathcal{S}_Y(b_\perp, r_\perp)$, with the S-matrix

$$\mathcal{S}_Y(r_\perp) = \frac{1}{N_c} \langle \text{Tr} (V(x_\perp) V^\dagger(y_\perp)) \rangle_Y. \quad (22)$$

One can compute the S-matrix explicitly in the MV model, which gives (Kovchegov, 1999; McLerran and Venugopalan, 1999; Venugopalan, 1999),

$$\mathcal{S}_Y(r_\perp) = \exp \left[-\alpha_S \frac{\pi^2}{2N_c} \frac{r_\perp^2 A x G_N(x, 1/r_\perp^2)}{\pi R_A^2} \right], \quad (23)$$

where G_N denotes the gluon distribution in the proton at the scale $\frac{1}{r_\perp^2}$.

One can expand out the exponential for very small values of r_\perp , and one observes that dipole cross-section is nearly transparent to color for small dipoles. As r_\perp grows, the S-matrix decreases; the saturation scale is defined as the value of r_\perp at which the S-matrix has a value that is significantly smaller than what one anticipates in pQCD. While there is some freedom in setting this scale, its growth with decreasing x is determined by the growth in the gluon distribution.

The MV result in Eq. (23) is the QCD Glauber model (Mueller, 1990) which gives the survival probability of a dipole after multiple independent scatterings off the nucleus. It can be refined by introducing an impact parameter distribution inside the proton (Bartels *et al.*, 2002), the so-called IP-Sat model, which can be further extended to model the S-matrix for nuclei (Kowalski *et al.*, 2008; Kowalski and Teaney, 2003).

The IP-Sat model provides very good agreement with a wide range of small x DIS data on e+p scattering at HERA (Rezaeian *et al.*, 2013). The latter constrains the parameters of this model, which in turn is an essential ingredient of the IP-Glasma model of the initial conditions for heavy-ion collisions. We will discuss the IP-Glasma model in Sec. III.C.3.

An advantage of the MV model formulation is that one can compute with relative ease (Blaizot *et al.*, 2004b; Dominguez *et al.*, 2011; Dusling *et al.*, 2018b; Fujii, 2002; Fukushima and Hidaka, 2017) not just the dipole Wilson line correlator but quadrupole and higher point correlators that appear in semi-inclusive final states in e+A and p+A collisions.

F. RG evolution and geometric scaling

The MV model is valid for a large nucleus at rapidities when the bremsstrahlung of soft gluons is not significant, namely, for $\alpha_S Y \leq 1$. The classical expressions we derived have no x dependence. For moderate x , one can introduce x dependence in framework along the lines of the IP-Sat model we discussed. However when $\alpha_S Y \gg 1$, the model is no longer applicable. In this regime, the RG evolution of the S-matrix in Eq. (22) is described by the Balitsky-JIMWLK hierarchy in Eq. (18) which, in addition to the coherent multiple scattering effects in the MV model, captures the the real and virtual quantum corrections we discussed previously.

Substituting the expectation value of the correlator of Wilson lines in Eq. (22) into the Balitsky-JIMWLK hierarchy in Eq. (18), leads, for $N_c, A \gg 1$, to the closed form In the limit Balitsky-Kovchegov (BK) (Balitsky, 1996; Kovchegov, 1999) equation for the RG evolution in rapidity of the dipole scattering amplitude:

$$\begin{aligned} \frac{\partial \mathcal{N}_Y(x_\perp, y_\perp)}{\partial Y} &= \bar{\alpha}_S \int_{z_\perp} \frac{(x_\perp - y_\perp)^2}{(x_\perp - z_\perp)^2 (z_\perp - y_\perp)^2} \\ &\times \left[\mathcal{N}_Y(x_\perp, z_\perp) + \mathcal{N}_Y(y_\perp, z_\perp) - \mathcal{N}_Y(x_\perp, y_\perp) \right. \\ &\quad \left. - \mathcal{N}_Y(x_\perp, z_\perp) \mathcal{N}_Y(z_\perp, y_\perp) \right]. \end{aligned} \quad (24)$$

The BK equation is the simplest RG equation that captures the physics of gluon saturation. For $\mathcal{N}_Y \ll 1$, the non-linear term in the last line above can be ignored and the equation reduces to the linear BFKL equation as anticipated previously. In this limit, the

amplitude has the solution,

$$\mathcal{N}_Y(r_\perp) \approx \exp\left(\omega\bar{\alpha}_s Y - \frac{\rho}{2} - \frac{\rho^2}{2\beta\bar{\alpha}_s Y}\right), \quad (25)$$

where $\omega = 4 \log 2 \approx 2.77$, $\beta = 28 \zeta(3) \approx 33.67$ and $\rho = \log(1/r_\perp^2 \Lambda_{\text{QCD}}^2)$. This solution gives the rapid ‘‘Markovian’’ growth of the dipole cross-section in rapidity due to the copious production of softer and softer gluons.

However when $\mathcal{N}_Y \sim 1$, the non-linear term arising from the fusion and screening of soft gluons completely saturates the growth of the dipole cross-section. If we impose a saturation condition $\mathcal{N}_Y = 1/2$, for $r_\perp = 2/Q_S$, on Eq. (25), the argument of the exponential vanishes for $\rho_s = \log(Q_S^2/\Lambda_{\text{QCD}}^2)$, with

$$Q_S^2 = \Lambda_{\text{QCD}}^2 e^{c\bar{\alpha}_s Y} \quad \text{where } c = 4.88. \quad (26)$$

Further, if we write $\rho = \rho_S + \delta\rho$, where $\delta\rho = \log(1/r_\perp^2 Q_S^2)$, one finds that (Iancu *et al.*, 2002)

$$\mathcal{N}_Y \approx (r_\perp^2 Q_S^2)^{\gamma_s}, \quad (27)$$

for $Q^2 < Q_S^4/\Lambda_{\text{QCD}}^2$, where $\gamma_s = 0.63$ is the BK anomalous dimension.

This ‘‘geometrical scaling’’ of the forward scattering amplitude means that Eq. (21) scales with $Q^2/Q_S^2(x)$ alone instead of x and Q^2 separately. Remarkably, this phenomenon was observed at HERA, providing a strong hint for the saturation picture (Stasto *et al.*, 2001). Moreover, the wider scaling window $Q^2 < Q_S^4/\Lambda_{\text{QCD}}^2$ stretching beyond Q_S provides a first principles explanation for a so-called ‘‘leading twist shadowing’’ of nuclear parton distributions relative to those in the proton (Frankfurt *et al.*, 2012). Such shadowed parton distributions are used to compute the rates of hard processes in heavy-ion collisions; understanding their microscopic origins is therefore important for quantifying hard probes of thermalization.

The BK equation, in a reaction-diffusion approximation, can be formally mapped into a well-known equation in statistical physics, the Fischer-Kolmogorov-Petrovsky-Piscounov (FKPP) equation (Munier and Peschanski, 2003). In this context, geometrical scaling appears as a late-time solution of a non-linear equation describing a traveling wavefront of constant velocity. In Fig. 3, we show numerical results for the unintegrated gluon distribution $\text{phi}(k_\perp^2) = \frac{\pi N_c k_\perp^2}{2\alpha_s} \int_0^{+\infty} d^2 r_\perp e^{ik_\perp \cdot r_\perp} [1 - \mathcal{N}_Y(r_\perp)]^2$, which displays this traveling wave front structure, with the evolution of the peaks of the wavefronts representing the evolution

of Q_S^2 with rapidity. The correspondence of high energy QCD to reaction-diffusion processes is very rich; specific applications to DIS have been discussed recently (Mueller and Munier, 2018a,b).

Q_S^2 in Eq. (26) (and the amplitude in Eq. (27)) grows very rapidly with rapidity, much faster than seen in the HERA data. However this is significantly modified by running coupling corrections, which are part of the next-to-leading-logs in x (NLLx) contributions to QCD evolution. The significant effect of these running coupling corrections is clearly seen in Fig. 3.

These give (Mueller and Triantafyllopoulos, 2002),

$$Q_{s,\text{running}}^2 = \Lambda_{\text{QCD}}^2 \exp\left(\sqrt{2b_0c(Y + Y_0)}\right), \quad (28)$$

where b_0 is the coefficient of the logarithm in the one loop QCD β -function¹⁶. The running coupling results are well approximated by a power law increase of the amplitude consistent with the HERA data. Further, qualitative features of geometric scaling persist, albeit the window for geometrical scaling is significantly smaller (Triantafyllopoulos, 2008).

For a large nucleus at the saturation boundary $Y_0 \propto \log^2(A^{1/3})$, one recovers the $A^{1/3}$ scaling of the saturation scale in the MV model from Eq. (28) for $Y_0 \gg Y$. A striking result, for $Y \gg Y_0$, is that the saturation scale for fixed impact parameter becomes independent of A . *In the asymptotic Regge limit, strongly correlated gluons in the nuclear wavefunctions lose memory of the initial conditions whereby they were generated.*

G. The state of the art in the CGC EFT

In previous sub-sections, we outlined a description of the wavefunction of a high energy nucleus in the CGC EFT, emphasizing a qualitative understanding of gluon saturation and key related analytical results. There have been significant developments since in the CGC EFT.

On the formal side, the Balitsky-JIMWLK framework for the LLx evolution of n -point Wilson line correlators, has been extended to NLLx (Balitsky and Chirilli, 2013b; Balitsky and Grabovsky, 2015; Caron-Huot, 2018; Kovner *et al.*, 2014a,b). For the two-point dipole correlator, which satisfies the LLx BK equation, the formalism has been extended

¹⁶ Sub-leading corrections in Y to Q_S have been computed to high order (Beuf, 2010).

to NLLx (Balitsky and Chirilli, 2008) and even (for $\mathcal{N} = 4$ supersymmetric Yang-Mills) to NNLLx in a recent *tour de force* computation (Caron-Huot and Herranen, 2018). The BFKL/BK kernel however receives large collinear contributions that need to be resummed in so-called small x resummation schemes for quantitative predictions (Ciafaloni *et al.*, 1999; Ducloué *et al.*, 2019; Iancu *et al.*, 2015; Salam, 1998).

While as we have discussed, there are good analytical approximations, a full analytical solution of the BK equation does not exist. Numerical simulations have however been known for some time for the LLx BK equation (Albacete *et al.*, 2004), the LLx+running coupling BK equation (Albacete *et al.*, 2005; Albacete and Kovchegov, 2007), and even more recently the full NLLx equation implementing collinear resummation (Ducloué *et al.*, 2020; Lappi and Mäntysaari, 2016). In particular, it is shown in (Ducloué *et al.*, 2020) that this NLLx framework provides very good agreement with the HERA data.

Numerical simulations have also been performed of higher point correlators in the Balitsky-JIMWLK hierarchy. As noted, Eq. (18) has the form of a functional Fokker-Planck equation. This can therefore be reexpressed as a Langevin equation in the space of Wilson lines (Blaizot *et al.*, 2003b; Weigert, 2002), allowing one to simulate the rapidity evolution of two-point Wilson line correlators (Rummukainen and Weigert, 2004) as well as four-point quadrupole and sextupole¹⁷ correlators (Dumitru *et al.*, 2011b; Lappi and Mäntysaari, 2013; Lappi and Ramnath, 2019). Fig. 4 shows a result for the dipole correlator from these simulations. Unfortunately, a similar Langevin representation is not known at present for the NLLx JIMWLK Hamiltonian.

Precision computations require not just higher order computations of the JIMWLK kernel but higher order computations of process dependent “impact factors” analogous to pQCD computations of coefficient functions that are convoluted, order-by-order, with the DGLAP splitting functions (Vermaseren *et al.*, 2005). For inclusive DIS, analytical expressions exist for the virtual photon impact factor $|\Psi_{\gamma^* \rightarrow q\bar{q}}|^2$ in Eq. (20) (Balitsky and Chirilli, 2013a). More recently, NLO impact factors have been computed for DIS exclusive diffractive light vector meson production (Boussarie *et al.*, 2017) and DIS inclusive photon+dijet production (Roy and Venugopalan, 2019, 2020). Numerical implementation of these results remains a formidable task and an essential component of precision studies of gluon saturation at a

¹⁷ These are probed in semi-inclusive DIS (Dominguez *et al.*, 2011) and in proton-nucleus collisions (Dusling *et al.*, 2018a,b; Kovner and Lublinsky, 2011).

future Electron-Ion Collider (EIC) (Accardi *et al.*, 2016; Aschenauer *et al.*, 2019).

An outstanding problem at small x is the impact parameter dependence of distributions. The BFKL kernel at large impact parameters contributes a Coulomb tail $\sim 1/b_{\perp}^2$; the conformal symmetry of the kernel and geometric scaling suggest a particular dependence of the saturation scale on the impact parameter (Gubser, 2011). The Coulomb tail is however not regulated by saturation and violates the Froissart bound on the asymptotic behavior of total cross-sections (Kovner and Wiedemann, 2003). This is only cured non-perturbatively by the generation of a mass gap in QCD. The Coulomb tail may be less of a problem in large nuclei with $\Lambda_{\text{QCD}}R_A \gg 1$ because the contribution of the Coulomb tail may be suppressed relative to protons, for which $\Lambda_{\text{QCD}}R_A \sim 1$.

III. NON-EQUILIBRIUM QCD MATTER AT HIGH OCCUPANCY

The CGC EFT provides us with powerful tools to address multi-particle production in heavy-ion collisions from first principles; the key organizing principle is the kinematic separation in the hadron wavefunction between static color sources at large x and small x gauge fields. In the following, we will sketch the elements of the formalism to follow the thermalization process through the overlap of two CGCs.

To apply this EFT framework to thermalization, one needs to understand first how to compute from first principles multi-particle production in the presence of strong fields¹⁸. The quark-gluon matter formed in this process is the Glasma (Gelis and Venugopalan, 2006c; Lappi and McLerran, 2006), a nonequilibrium state with high occupancy $f \sim \mathcal{O}(1/\alpha_S)$; this state decays and eventually thermalizes. The description of the temporal evolution of the Glasma can be classified systematically in weak coupling into LO, NLO, and so on.

Following our discussion of multi-particle production, we will describe the temporal evolution of the Glasma at LO. This corresponds to the solution of classical Yang-Mills equations with CGC initial conditions for the fields using both analytical approaches (valid for transverse momenta greater than the saturation scale) and a nonperturbative real time approach employing Hamilton’s equation on the lattice. The LO solutions are independent of rapidity, with the dynamics of the corresponding “flux tube” structures occurring entirely in

¹⁸ A well-known example of such a formalism is e^+e^- pair production in strong electromagnetic fields (Gelis and Tanji, 2016); another is that of Hawking radiation from the Black Hole horizon (Parikh and Wilczek, 2000).

the transverse plane of the collision. We will next discuss the IP-Glasma model of heavy-ion collisions, which combines the LO classical solutions with constraints on Q_S from DIS experiments on the proton and on nuclei.

However the LO description of the Glasma is limited because the classical fields are unstable to NLO quantum fluctuations that break boost invariance, growing exponentially in the square root of the proper time. As we will discuss, a careful treatment of such NLO modes shows that the dominant contributions can be resummed and absorbed into a classical-statistical description of the evolution. A key difference to the prior LO description is that the resummed classical-statistical evolution is now in 3+1-dimensions, involving both transverse and longitudinal degrees of freedom. This distinction is of fundamental importance in the subsequent description of the thermalization process in weak coupling.

In Section IV, we will discuss how this classical-statistical description fits into the general weak coupling classification of the evolution of quantum fields and shall outline the power counting that delineates the applicability of this approximation and its subsequent matching to kinetic theory. We will also describe there universal features of the Glasma that makes its study interesting in its own right.

A. Multi-particle production in strong fields

To compute multi-particle production systematically in the collision of the CGC gluon “shockwaves”, we will begin with the first principles Lehmann-Symanzik-Zimmerman (LSZ) formalism in QFT. For simplicity, we consider here a self-interacting ϕ^3 scalar theory; our discussion extends straightforwardly to the Yang-Mills case.

In the LSZ formalism, the amplitude for n -particles in the “out” state generated from the “in-vacuum” can be expressed as

$$\begin{aligned} \langle p_{1,\text{out}} \cdots p_{n,\text{out}} | 0_{\text{in}} \rangle &= \frac{1}{Z^{n/2}} \int \left[\prod_{i=1}^n d^4 x_i e^{ip_i \cdot x_i} \right. \\ &\left. \times (\partial_{x_i}^2 + m^2) \frac{\delta}{\delta J(x_i)} \right] \exp(i\mathcal{V}) . \end{aligned} \quad (29)$$

Here p_1, \dots, p_n denote the momenta of the produced particles and the “in-out” vacuum-amplitude $\langle 0_{\text{out}} | 0_{\text{in}} \rangle = \exp(i\mathcal{V})$, where \mathcal{V} is the sum of all connected vacuum-vacuum diagrams coupled to external sources. An illustration of multi-particle production for the

problem at hand is shown in Fig. 5.

In QFT computations, one usually sets $J = 0$ after the functional differentiation and $\langle 0_{\text{out}} | 0_{\text{in}} \rangle$ is a pure phase. When J is physical, $|\langle 0_{\text{out}} | 0_{\text{in}} \rangle|^2 = \exp(-2 \text{Im } \mathcal{V}) \neq 1$. In computing multi-particle production in this context, it is useful to employ¹⁹ the Schwinger-Keldysh (SK) QFT formalism (Keldysh, 1964; Schwinger, 1961). One introduces $+$ and $-$ vertices with opposite signs of the coupling in Feynman diagrams, and likewise for the sources J_{\pm} . The corresponding “ $+$ ” and “ $-$ ” fields live on the upper and lower segments of a closed time contour ranging forward in time from $t = -\infty$ on the upper contour and back to $-\infty$ on the lower contour, as shown in Fig. 5. Time ordered “ $++$ ” (anti-time ordered “ $--$ ”) Green’s functions “live” on the upper (lower) contour, and the mixed $+-$ “Wightman” functions connect the upper and lower contours.

Following LSZ, the probability to produce n -identical particles is

$$P_n = \frac{1}{n!} \prod_{i=1}^n \frac{d^3 p_i}{(2\pi)^3 2E_{p_i}} |\langle p_{1,\text{out}} \cdots p_{n,\text{out}} | 0_{\text{in}} \rangle|^2, \quad (30)$$

where $E_{p_i}^2 = p_i^2 + m^2$. Plugging the expression for the amplitude in Eq. (29) into the r.h.s, one can express the result as (Gelis and Venugopalan, 2006a)

$$P_n = \frac{1}{n!} \mathcal{D}^n \exp(i\mathcal{V}[J_+] - i\mathcal{V}[J_-]) |_{J_+=J_-=J}, \quad (31)$$

with

$$\mathcal{D} = \int_{x,y} Z G_{+-}^0(x,y) \frac{(\partial_{x_i}^2 + m^2)}{Z} \frac{(\partial_{y_i}^2 + m^2)}{Z} \frac{\delta}{\delta J_+(x)} \frac{\delta}{\delta J_-(y)}. \quad (32)$$

Here $\int_x = d^4x$, $G_{+-}^0(x,y) = \int \frac{d^3 p_i}{(2\pi)^3 2E_{p_i}} e^{ip \cdot (x-y)} \equiv \theta(p^0) \delta^{(3)}(x-y)$ and Z is the residue of the pole of the renormalized propagator.

The action of the operator \mathcal{D} can be understood as follows. The “ $+$ ” piece with $\frac{(\partial_{x_i}^2 + m^2)}{Z} \frac{\delta}{\delta J_+(x)}$ acts on a particular diagram in the connected sum of vacuum-vacuum connected diagrams $\mathcal{V}[J_+]$ by removing a source J_+ and then amputating the renormalized propagator to which it is attached. The same procedure is followed for the “ $-$ ” piece; the two amputated propagators are then sewn together by the renormalized “cut” propagator ZG_{+-}^0 .

¹⁹ For other discussions of the SK formalism in the context of the CGC and the Glasma, see (Jeon, 2014; Leonidov and Radovskaya, 2019; Wu and Kovchegov, 2018). For a recent discussion in the context of thermal field theory, see (Ghiglieri *et al.*, 2020).

Computing P_n in a theory with physical sources is hard because one also has to compute the disconnected vacuum-vacuum graphs for each n . However if we define a generating functional $F(z) = \sum_n z^n P_n$, Eq. (31) gives

$$F(z) = \exp(z\mathcal{D}) \exp(i\mathcal{V}[J_+] - i\mathcal{V}[J_-])|_{J_+=J_-=J}, \quad (33)$$

and successive differentiation of this equation with respect to z (and setting $z = 1$), generates the n -particle correlators $\langle n(n-1)(n-2)\cdots \rangle$. These moments do not require one compute the disconnected vacuum-vacuum graphs, since they also appear in the normalization of P_n and therefore cancel out²⁰ in the moments.

This is illustrated by expressing the r.h.s of Eq. (33) for $z = 1$ as

$$\exp(i\mathcal{V}_{\text{SK}}[J_+, J_-]) = \exp(\mathcal{D}) \exp(i\mathcal{V}[J_+] - i\mathcal{V}[J_-]), \quad (34)$$

where now $i\mathcal{V}_{\text{SK}}[J_+, J_-]$ represents the sum over all vacuum-to-vacuum connected graphs that live on the SK closed time contour. One can then express the inclusive multiplicity as (Gelis and Venugopalan, 2006a)

$$\langle N \rangle = \int_{x,y} Z G_{+-}^0(x,y) [\Gamma_+(x)\Gamma_-(y) + \Gamma_{+-}(x,y)]_{J_{\pm}=J}, \quad (35)$$

with the amputated one-point and two-point Green's functions in the Schwinger-Keldysh formalism defined respectively as

$$\Gamma_{\pm}(x) = \Delta_x^{\text{R}} \frac{\delta i\mathcal{V}_{\text{SK}}}{\delta J_{\pm}(x)}; \quad \Gamma_{+-}(x,y) = \Delta_x^{\text{R}} \Delta_y^{\text{R}} \frac{\delta^2 i\mathcal{V}_{\text{SK}}}{\delta^2 J_+(x) J_-(y)}, \quad (36)$$

with $\Delta_x^{\text{R}} = \frac{\partial_x^2 + m^2}{Z}$.

In summing over all the nodes of all the trees connecting $\Gamma_+(x)$ to the sources, the time (anti-time) ordered Feynman propagators in each tree on the upper (lower) SK contour are recursively converted to retarded propagators: $G_R = G_{++} - G_{+-} \equiv G_{-+} - G_{--}$. This is equivalent to solving the classical equations of motion with retarded boundary conditions when $J_{\pm} = J$! A further important result is that the renormalized cut propagator Γ_{+-} is obtained by solving the small fluctuation equations of motion in the classical background, also as an initial value problem with retarded boundary conditions.

²⁰ Such cancellations are seen in the Abramovsky-Gribov-Kancheli (AGK) rules (Abramovsky *et al.*, 1973) that implement the combinatorics of cut/uncut vacuum-to-vacuum graphs in Reggeon field theory (Gelis and Venugopalan, 2007).

As we discussed previously, the classical fields, and sources thereof, of the colliding CGC's are static shockwaves; as such, they do not spontaneously decay and are thus part of the nuclear wavefunction. After the collision, the colored sources become time dependent. Thus Γ_{\pm} in Eq. (35) corresponds to $\partial_x^2 \mathcal{A}_{\pm,\text{cl}}^{\mu}$, where $\mathcal{A}_{\pm,\text{cl}}^{\mu}$ is the time dependent $\mathcal{O}(1/g)$ Glasma field in the forward lightcone. The two-point function $\Gamma_{+-}(x, y)$ in Eq. (35) is $\mathcal{O}(1)$ and therefore NLO in the power counting for the inclusive multiplicity in the Glasma. The formalism can be extended to higher orders in α_S . Its generalization to higher multiplicity moments was developed in (Gelis and Venugopalan, 2006b).

B. The LO Glasma: classical gluon fields from shockwave collisions

Since at LO in our power counting only the product $\Gamma_+(x)\Gamma_-(y) \equiv \partial_x^2 \mathcal{A}_+^{\mu} \partial_y^2 \mathcal{A}_-^{\nu}$ in Eq. (35) contributes, one obtains for a fixed distribution of lightcone sources $\rho_{\pm,1,2} = \rho_{1,2}$ (where 1, 2 denote the two nuclei) (Gelis *et al.*, 2007)

$$\frac{d\langle N \rangle_{\text{LO}}}{dY d^2 p_{\perp}}[\rho_1, \rho_2] = \frac{1}{16\pi^3} \int_{x,y} \Delta_x^{\text{R}} \Delta_y^{\text{R}} \varepsilon_{\lambda}^{\mu} \varepsilon_{\lambda}^{\nu} \mathcal{A}_{\mu}(x) \mathcal{A}_{\nu}(y), \quad (37)$$

where repeated indices are summed over. Note too that $\mathcal{A}(x) \equiv \mathcal{A}_{\mu}[\rho_1, \rho_2](x)$ and $m = 0$ in $\Delta_{x,y}^{\text{R}}$. An integration by parts,

$$\int d^4 x e^{ip \cdot x} \partial_x^2 \mathcal{A}_{\mu}(x) = \int_{x^0 \rightarrow +\infty} d^3 x e^{ip \cdot x} (\partial_0 - iE_p) \mathcal{A}_{\mu}(x), \quad (38)$$

shows that Eq. (37) can be computed by solving the classical YM equations in Eq. (10) (with $J^{\mu} = \delta^{\mu+} \delta(x^-) \rho_1(x_{\perp}) + \delta^{\mu-} \delta(x^+) \rho_2(x_{\perp})$ and $\mathcal{A}_{\mu}(x)|_{x^0=-\infty} = 0$) to determine $\mathcal{A}_{\mu}(x)$.

In the discussion to follow, it will be convenient to introduce the (τ, η, x_{\perp}) coordinate system, where the proper time $\tau = \sqrt{(x^0)^2 - (x^3)^2}$ and the spacetime rapidity $\eta = \frac{1}{2} \log(\frac{x^0+x^3}{x^0-x^3})$, and $g_{\mu\nu} = \text{diag}(1, -\tau^2, -1, -1)$. A convenient gauge to solve the YM equations in the forward lightcone is the Fock-Schwinger gauge $\mathcal{A}^{\tau} \equiv x^+ \mathcal{A}^- + x^- \mathcal{A}^+ = 0$. In this gauge²¹, the solution to the YM equations are manifestly boost invariant: $\mathcal{A}^{\mu}(\tau, \eta, x_{\perp}) \equiv \mathcal{A}^{\mu}(\tau, x_{\perp})$ and one obtains (Gyulassy and McLerran, 1997; Kovner *et al.*, 1995a,b),

$$\mathcal{A}^i = A_{1,\text{cl}}^i + A_{2,\text{cl}}^i; \quad \mathcal{A}^{\eta} = \frac{ig}{2} [A_{1,\text{cl}}^i, A_{2,\text{cl}}^i], \quad (39)$$

²¹ A perturbative solution was also found in Lorenz gauge $\partial_{\mu} \mathcal{A}^{\mu} = 0$ (Kovchegov and Rischke, 1997).

with $\partial_\tau \mathcal{A}^i = 0$ and $\partial_\tau \mathcal{A}^\eta = 0$ at $\tau = 0^+$. This solution is obtained by matching the delta-functions on the lightcone wedges in Fig. 6 where .

Since the gauge fields are functionals of $\rho_{1,2}$, the full average inclusive multiplicity in the Glasma is obtained by averaging over many nuclear collisions, each with its distribution of color sources in the two nuclei²². This can be expressed as

$$\begin{aligned} \frac{d\langle\langle N \rangle\rangle_{\text{LO}}}{dY d^2 p_\perp} &= \int [D\rho_1][D\rho_2] W_{Y_{\text{beam}}-Y}^{\text{MV}}[\rho_1] W_{Y_{\text{beam}}+Y}^{\text{MV}}[\rho_2] \\ &\times \frac{d\langle N \rangle_{\text{LO}}}{dY d^2 p_\perp}[\rho_1, \rho_2], \end{aligned} \quad (40)$$

where $Y_{\text{beam}} = \log(\sqrt{s}/m_N)$ is the beam rapidity and $W_{Y_{\text{beam}}-Y}^{\text{MV}}$ ($W_{Y_{\text{beam}}+Y}^{\text{MV}}$) are the weight functionals in the MV model in Eq. (9) and at LO are independent of $Y_{\text{beam}} - Y$ ($Y_{\text{beam}} + Y$).

With the initial conditions in Eq. (39), the YM equations for $\tau = 0^+$ can be solved perturbatively to lowest non-trivial order in $\mathcal{O}(\frac{\rho_1}{\nabla_\perp^2} \frac{\rho_2}{\nabla_\perp^2})$; in this ‘‘dilute-dilute’’ approximation, one obtains for identical nuclei,

$$\frac{d\langle\langle N \rangle\rangle_{\text{LO}}}{dY d^2 p_\perp} = \pi R_A^2 \frac{g^6 \mu_A^4}{(2\pi)^4} \frac{2N_c(N_c^2 - 1)}{p_\perp^4} \mathcal{L}(p_\perp, \Lambda). \quad (41)$$

This result, which agrees with the pQCD bremsstrahlung formula first derived by Gunion and Bertsch (Gunion and Bertsch, 1982) is valid for $p_\perp \gg g^2 \mu_A$ and $\mathcal{L}(p_\perp, \Lambda)$ is a logarithmically divergent function, screened at $\Lambda \approx \Lambda_{\text{QCD}}$.

From our dipole model discussion (see Eq. (23) and related discussion), $Q_S^2 \propto G_A(x, p_\perp^2)$, where p_\perp is the momentum conjugate to the dipole size. This suggests that Eq. (41) (employing $Q_S^2 \propto \mu_A^2$, as noted in footnote 10) can be generalized to a ‘‘ k_\perp factorization’’ form $\frac{d\langle\langle N \rangle\rangle_{\text{LO}}}{dY d^2 p_\perp} \propto \alpha_S \int dk_\perp^2 \phi_A(x_1, k_\perp^2) \phi_B(x_2, (k_\perp - p_\perp)^2)$. Here $\frac{\phi_{A,B}(x, k_\perp^2)}{k_\perp^2}$ is the Fourier transform of the dipole scattering amplitude²³ in the each of the hadrons we discussed previously in Sec. II.F. This k_\perp factorization formula (Blaizot and Mueller, 1987; Gribov *et al.*, 1983) is widely used in phenomenological studies of hadron-hadron collisions.

The dilute-dilute analytical approximation for shockwave collisions can be generalized to compute the inclusive multiplicity to lowest order $\mathcal{O}(\frac{\rho_1}{\nabla_\perp^2})$ in one of the sources but to all orders $\mathcal{O}((\frac{\rho_2}{\nabla_\perp^2})^n)$ in the other. In this ‘‘dilute-dense’’ case as well, the inclusive gluon

²² Due to color confinement at distances scales $1/\Lambda_{\text{QCD}}$, one requires $\int_0^{1/\Lambda_{\text{QCD}}} d^2 x_\perp \rho_{1,2}^a = 0$ for each such configuration.

²³ This distribution is distinct from the WW-distribution and coincides with it only for large k_\perp (Blaizot *et al.*, 2004a; Kharzeev *et al.*, 2003).

multiplicity can be expressed as a k_{\perp} -factorized convolution of the unintegrated gluon distributions in the projectile and target. It is valid for $Q_{S,1}^2(x_1) \ll Q_{S,2}^2(x_2)$, corresponding to the forward (or backward) kinematic regions of the shockwave collision where the parton momentum fractions are $x_1 \gg x_2$. Alternately, it can be a good approximation in proton-nucleus collisions, where $Q_{S,A}^2 \sim A^{1/3} Q_{S,p}^2$ (Dumitru and McLerran, 2002; Kovchegov and Mueller, 1998).

C. Non-perturbative evolution of high occupancy fields

1. Real time evolution of boost invariant fields on the lattice

While analytical results for the inclusive multiplicity are available only in limited kinematic regions, the YM equations for shockwave collisions can be solved numerically to all orders $\mathcal{O}((\frac{\rho_{1,2}}{\sqrt{V_{\perp}}})^n)$ (Krasnitz and Venugopalan, 1998, 1999) to obtain the full non-perturbative result to Eq. (40) (Krasnitz *et al.*, 2001, 2003a; Krasnitz and Venugopalan, 2000, 2001; Lappi, 2003). Hamilton's equations are solved in Fock-Schwinger gauge $\mathcal{A}^{\tau} = 0$ with the initial conditions at $\tau = 0$ specified by Eq. (39). To preserve gauge invariance, lattice gauge theory techniques can be adapted to this problem. The boost invariance of the LO shockwave gauge fields provides a significant simplification whereby the 3+1-D Kogut-Susskind QCD lattice Hamiltonian (Kogut and Susskind, 1975) can be “dimensionally reduced” to the 2+1-D form (Krasnitz and Venugopalan, 1999)

$$aH = \sum_{\mathbf{x}} \left[\frac{g^2 a}{\tau} \text{tr} E^i E^i + \frac{2\tau}{g^2 a} (N_c - \text{Re tr} U_{1,2}) + \frac{\tau}{a} \text{tr} \pi^2 + \frac{a}{\tau} \sum_i \text{tr} (\Phi - \tilde{\Phi}_i)^2 \right]. \quad (42)$$

Here the trace refers to $SU(2)$ color and the sum is over all discretized cells with lattice spacing a in the transverse plane. For clarity, we have omitted the cell index j for all quantities in this expression. Further, the E^i with $i \in \{1, 2\}$ are the components of the transverse electric field living on each site; discretizing the initial conditions gives $E^i = 0$ at $\tau = 0$. The spatial plaquette of link variables U_j^i ,

$$U_{1,2}^j = U_j^1 U_{j+\hat{e}_1}^2 U_{j+\hat{e}_2}^{1\dagger} U_j^{2\dagger}, \quad (43)$$

(where $+\hat{e}_i$ indicates a shift from j by one lattice site in the $i = 1, 2$ transverse direction) represents the squared longitudinal magnetic fields in the Glasma. In Eq. (42), we have represented $A_\eta(\tau, x_\perp)$ as an adjoint scalar field Φ because, as a result of boost invariance, it transforms covariantly under η -dependent gauge transformations:

$$\tilde{\Phi}_i^j = U_j^i \Phi_{j+\hat{e}_i} U_j^{i\dagger}. \quad (44)$$

Finally, $\pi = E_\eta = \dot{\Phi}/\tau$ in Eq. (42) represents the longitudinal electric field.

The details of the numerical simulations of the real time evolution of gauge fields can be found in (Krasnitz and Venugopalan, 1999; Lappi, 2003). In the early work, only uniform sheets of nuclei were considered with constant (x independent) values of Q_S . These were subsequently relaxed to consider finite nuclei (Krasnitz *et al.*, 2003b,c); more realistic simulations with event-by-event simulations of RHIC and LHC collisions were developed later in the IP-Glasma model we shall discuss shortly (Schenke *et al.*, 2012b).

As anticipated, the numerical results reproduce the perturbative result in Eq. (41) at large $k_\perp \gg Q_S$. However, unlike that expression, there is no logarithmic factor $\mathcal{L}(k_\perp, \Lambda_{\text{QCD}})$. At momenta $k_\perp < Q_S$, the $1/k_\perp^4$ distribution is modified to a form that is well fit by a Bose-Einstein exponential distribution (Krasnitz *et al.*, 2003a). Even more remarkably, the non-linear dynamics generates a plasmon mass²⁴ that screens the momentum distribution in the infrared (Krasnitz and Venugopalan, 2001; Lappi and Peuron, 2018). The energy density is therefore well-defined at all proper times without infrared or ultraviolet divergences (Lappi, 2006).

2. Glasma flux tubes

An interesting consequence of the LO Glasma solution is that the Weizäcker-Williams plane polarized E and B fields in the colliding CGCs become purely longitudinal immediately after the collision at $\tau = 0^+$; $E_\eta, B_\eta \neq 0$ and $E_i, B_i = 0$. It was pointed out in (Kharzeev *et al.*, 2002) that this configuration satisfies the identity

$$Q_{\text{CS}} = \frac{\alpha_S}{2\pi} \int d^4x \text{Tr } E_\eta \cdot B_\eta, \quad (45)$$

²⁴ This plasmon mass is parametrically larger than the confining scale; its properties have been investigated recently in a number of approaches (Boguslavski *et al.*, 2019; Dumitru *et al.*, 2014).

where the topological charge $Q_{CS} = \frac{\alpha_S}{16\pi} \int d^3x K^0$ and K^μ is the Chern-Simons current. A neat interpretation (Chen *et al.*, 2015a; Lappi and McLerran, 2006) of this result is that the YM equations at $\tau = 0^+$ can be expressed as $\nabla \cdot E = \rho_{\text{el.}}$ and $\nabla \cdot B = \rho_{\text{mag.}}$, where $\rho_{\text{el.}}$, $\rho_{\text{mag.}}$ are respectively electric and magnetic charges densities²⁵ on the gluon shockwaves after the collision.

As sketched in Fig. 7, the induced electric and magnetic charges generate a “stringy” Glasma flux tube (Dumitru *et al.*, 2008a) of chromo-electromagnetic fields that is uniform in rapidity stretching between the fragmentation regions of the nuclei and are color screened (Krasnitz *et al.*, 2003c) on transverse distance scales $\geq 1/Q_S$.

One can straightforwardly compute the energy densities and pressures in the Glasma from the different components of the stress-energy tensor²⁶. We obtain $\mathcal{E} = 2\mathcal{P}_T + \mathcal{P}_L$ where,

$$\begin{aligned}\mathcal{P}_T &\equiv \frac{1}{2} (T^{xx} + T^{yy}) = \text{Tr} (F_{xy} + E_\eta^2) \\ \mathcal{P}_L &\equiv \tau^2 T^{\eta\eta} = \frac{1}{\tau^2} \text{Tr} (F_{\eta i}^2 + E_i^2) - \text{Tr} (F_{xy} + E_\eta^2) .\end{aligned}\quad (47)$$

At the earliest times after the collision $\tau = 0^+$, as noted, only the longitudinal E_η and $B_\eta = F_{xy}$ fields are non-zero. The above equation then immediately gives $\mathcal{P}_T = \mathcal{E}$ and $\mathcal{P}_L = -\mathcal{E}$. Thus at the earliest times, the pressure in the Glasma is purely transverse; after initial transverse dynamics, the longitudinal pressure $\mathcal{P}_L \rightarrow 0$ from below by $\tau \sim 1/Q_S$. Since the Glasma at LO is conformal, the energy density satisfies $\mathcal{E} = 2\mathcal{P}_T$ at this time.

Stringy models capture essential features of confining dynamics in QCD (Bali, 2001). In high energy collisions, they have a long history and capture the bulk features of the spectrum of multi-particle production (Andersson *et al.*, 1979; Artru, 1983); they underlie event generators such as PYTHIA (Andersson *et al.*, 1983). These models however screen color at distance scales $1/\Lambda_{\text{QCD}}$ and only carry electric flux and no magnetic flux; particle production is assumed to arise from the Schwinger mechanism (Andersson *et al.*, 1979). It is remarkable nevertheless to observe that similar stringy solutions emerge from the more fundamental framework of classical YM equations.

Motivated by this stringy picture, we expect the number of gluons per unit rapidity equals the number of flux tubes ($S_\perp/(1/Q_S^2)$) times the gluon occupancy in a flux tube

²⁵ These induced charge densities are proportional to the commutators $\delta^{ij}[A_{1,\text{cl.}}^i, A_{1,\text{cl.}}^j]$ and $\epsilon^{ij}[A_{1,\text{cl.}}^i, A_{1,\text{cl.}}^j]$ respectively.

²⁶ Note that

$$T^{\mu\nu} = -g^{\mu\alpha} g^{\nu\beta} g^{\gamma\delta} F_{\alpha\gamma} F_{\beta\delta} + \frac{1}{4} g^{\mu\nu} g^{\alpha\gamma} g^{\beta\delta} F_{\alpha\beta} F_{\gamma\delta} .\quad (46)$$

$(2(N_c^2 - 1)/\bar{\alpha}_S/(2\pi)^3)$ multiplied by a non-perturbative coefficient of $\mathcal{O}(1)$. Extracting the number density from the correlator of gauge fields at $\tau \sim 1/Q_S$ (Krasnitz and Venugopalan, 2001), one indeed finds that²⁷

$$\frac{dN_{\text{LO}}}{dY} = c_N \frac{2(N_c^2 - 1)}{(2\pi)^3} \frac{Q_S^2 S_\perp}{\bar{\alpha}_S}, \quad (48)$$

where S_\perp is the transverse area of the collision, $\bar{\alpha}_S = \alpha_S N_c/\pi$ and c_N is a gluon liberation coefficient (Mueller, 2000) estimated from the numerical simulations to be $c_N = 1.1$ with 10% accuracy (Lappi, 2008).

The YM simulations can also be extended to compute two particle correlations in the Glasma (Lappi *et al.*, 2010):

$$\frac{d^2 N_{\text{LO}}^{\text{conn.}}}{dY_1 d^2 p_\perp dY_2 d^2 k_\perp} = \frac{\kappa_2}{(N_c^2 - 1) Q_S^2 S_\perp} \frac{dN_{\text{LO}}}{dY_1 d^2 p_\perp} \frac{dN_{\text{LO}}}{dY_2 d^2 k_\perp}, \quad (49)$$

where κ_2 is a non-perturbative constant²⁸. Again, the numerical simulations bear out the Glasma flux tube interpretation: the likelihood that two particles are correlated is suppressed by the number of flux tubes, and non-factorizable color connected graphs by $\mathcal{O}(1/N_c^2)$. Perturbative arguments suggest that this picture can be extended to n -particle cumulants and that the n -particle multiplicity distribution that generates these cumulants is a negative binomial distribution (Gelis *et al.*, 2009). For n -particle multiplicities, this expectation is confirmed by non-perturbative numerical simulations (Schenke *et al.*, 2012a).

3. The IP-Glasma model

In the discussion thus far, color charge fluctuations on the scale $1/Q_S$ provide the only structure in the colliding gluon shockwaves. However nucleon distributions in nuclei are not uniformly smooth and can fluctuate from event to event. These fluctuations in nucleon positions are extremely important to understand key features of the data such as the azimuthal moments v_n of the flow distributions at low momenta (Alver and Roland, 2010; Alver *et al.*, 2010). Another important ingredient in the realistic modeling of heavy-ion collisions is the dependence of the saturation scale in the nuclei on x (or equivalently, \sqrt{s}), which describes the variations of particle multiplicities in energy and rapidity at RHIC and the LHC.

²⁷ Here and henceforth, for simplicity of notation, the path integral over gauge fields (moot at LO), and over sources, $\langle\langle\rangle\rangle$ is implicit.

²⁸ The results have a weak dependence on the ratio m/Q_S , where m is an infrared lattice regulator.

We will outline here the IP-Glasma model (Schenke *et al.*, 2012a,b, 2014a,b), and improvements thereof, which incorporates the fluctuations in the nucleon positions to construct event-by-event lumpy color charge distributions and corresponding gluon field configurations in the LO Glasma framework. As we will also discuss, the energy dependence of these configurations at a given Y or \sqrt{s} is determined by the saturation scales in the two nuclei.

An essential input is the dipole cross-section of the proton. Here we consider here the IP-Sat saturation model (Bartels *et al.*, 2002; Kowalski and Teaney, 2003) which, as discussed in Sec. II.E, is an impact parameter dependent generalization of the MV model. As noted, high precision combined data from the H1 and ZEUS collaborations (Aaron *et al.*, 2010; Abramowicz *et al.*, 2013) are used to constrain the parameters of the model and excellent fits are obtained (Rezaeian *et al.*, 2013).

The dipole cross-section for each nucleus at a given x is constructed by taking the product of the S-matrices corresponding to the dipole cross-sections of overlapping nucleons at a given spatial location \mathbf{x}_\perp . It can be expressed as (Kowalski *et al.*, 2008)

$$\frac{1}{2} \frac{d\sigma_{\text{dip}}^A}{d^2\mathbf{x}_\perp} = \left[1 - e^{-\frac{\pi^2}{2N_c} \mathbf{r}_\perp^2 \alpha_S(Q^2) x G(x, Q^2) \sum_{i=1}^A T_p(\mathbf{x}_\perp - \mathbf{x}_{T^i})} \right], \quad (50)$$

where T_p stands for the Gaussian thickness function for each of the A nucleons in each nucleus and $Q^2 = 4/\mathbf{r}_\perp^2 + Q_0^2$, with Q_0 fixed by the HERA inclusive data. The gluon distribution $xG(x, Q^2)$ is parametrized at the initial scale Q_0^2 and then evolved up to the scale Q^2 using LO DGLAP-evolution. We define the nuclear saturation scale $Q_S = 1/\sqrt{\mathbf{r}_{\perp,s}^2}$, at the $r_\perp = r_{\perp,s}$ for which the argument of the exponential in Eq. (50) equals one-half. To obtain the spatial dependence of Q_S , one self-consistently solves $x = 0.5 Q_S(\mathbf{x}_\perp, x)/\sqrt{s}$ for every \mathbf{x}_\perp .

The result of this procedure is a lumpy distribution of $Q_S^2(\mathbf{x}_\perp, x)$ denoting the sub-nucleon structure of the nucleus. Since the IP-Sat model is a simple generalization of the MV model, one can extract the variance of the color charge density $g^2\mu_A^2(\mathbf{x}_\perp)$ at each x from $Q_S^2(\mathbf{x}_\perp, x)$ (Lappi, 2008). One then samples random color charges $\rho^a(\mathbf{x}_\perp)$ on a transverse lattice,

$$\langle \rho_k^a(\mathbf{x}_\perp) \rho_l^b(\mathbf{y}_\perp) \rangle = \delta^{ab} \delta^{kl} \delta^2(\mathbf{x}_\perp - \mathbf{y}_\perp) \frac{g^2\mu_A^2(\mathbf{x}_\perp)}{N_y}, \quad (51)$$

where the indices $k, l = 1, 2, \dots, N_y$ label the N_y points of representing the width of the nucleus in x^- . The path ordered Wilson line in the dipole model S-matrix (see (22)) is

discretized as

$$V_{A(B)}(\mathbf{x}_\perp) = \prod_{k=1}^{N_y} \exp \left(-ig \frac{\rho_k^{A(B)}(\mathbf{x}_\perp)}{\nabla_T^2 - m^2} \right), \quad (52)$$

where m is a infrared cut-off and A, B distinguish the color charge distributions in the two colliding nuclei. The corresponding dipole distributions in each of the incoming nuclei for a particular configuration of color sources is shown in Fig. 8(a).

To each lattice site j , one then assigns two $SU(N_c)$ matrices $V_{(A),j}$ and $V_{(B),j}$, each of which defines a pure gauge configuration with the link variables $U_{(A,B),j}^i = V_{(A,B),j} V_{(A,B),j+\hat{e}_i}^\dagger$, where $+\hat{e}_i$ indicates a shift from j by one lattice site in the $i = 1, 2$ transverse direction. The link variables in the future lightcone U_j^i which are an input into Eqs. (43) and (44), are determined (Krasnitz and Venugopalan, 1999) from solutions of the lattice CYM equations at $\tau = 0$,

$$\text{tr} \left\{ t^a \left[(U_{(A)}^i + U_{(B)}^i) (1 + U^{i\dagger}) - (1 + U^i) (U_{(A)}^{i\dagger} + U_{(B)}^{i\dagger}) \right] \right\} = 0, \quad (53)$$

where t^a are the generators of $SU(N_c)$ in the fundamental representation. (The cell index j is omitted here.) The $N_c^2 - 1$ equations in Eq. (53) are highly non-linear and for $N_c = 3$ are solved iteratively. With these initial conditions, Hamilton's equations corresponding to Eq. (42), are solved to compute inclusive quantities in the LO Glasma. Fig. 8(b) shows the result for the energy density in the transverse plane at $\tau = 1/Q_S$

The IP-Glasma model gives a good description of bulk features of distributions at RHIC and the LHC (Schenke *et al.*, 2014a,b). In particular, when matched with the MUSIC relativistic viscous hydrodynamic code (Schenke *et al.*, 2011), the IP-Glasma+MUSIC model provides an excellent description of the multiplicity distributions, the inclusive centrality and p_\perp distributions, and not least, the v_n distributions in heavy-ion collisions putting strong constraints on the extracted transport coefficients of the quark-gluon plasma (Gale *et al.*, 2013; Ryu *et al.*, 2015).

There have been several developments since. Firstly, the model has been extended to include JIMWLK evolution of the sources $\rho(x_\perp) \rightarrow \rho(x_\perp, x^\mp)$ for nuclei with large P^\pm enabling one so study rapidity correlations of produced gluons (Dusling *et al.*, 2010; Schenke and Schlichting, 2016) and 3-D evolution of the LO Glasma fields (McDonald *et al.*, 2020; Müller, 2019; Schenke and Schlichting, 2016). Further, the extension of the IP-Glasma+MUSIC

model to hadron-hadron and hadron-nucleus collisions (Bzdak *et al.*, 2013) indicates that sub-nucleon shape fluctuations in the Glasma are essential in understanding final state contributions to two and multi-particle cumulants of azimuthal anisotropies for high multiplicity events in small systems (Schenke and Venugopalan, 2014), the so-called “ridge” correlations (Dusling *et al.*, 2016).

Data on incoherent diffraction from HERA are sensitive to such non-perturbative “shape” fluctuations (Mäntysaari, 2020; Mäntysaari and Schenke, 2016a,b); the framework developed here allows one to constrain the latter with HERA data and in future likely more precisely with the EIC. Numerical simulations suggest that long range two particle correlations in the Glasma (Lappi *et al.*, 2016) when combined with hydrodynamic flow can explain the systematics of high multiplicity azimuthal moments in small systems (Schenke *et al.*, 2016, 2020b).

D. The Glasma at NLO

Thus far, we focused on the leading order dynamics of classical fields $\mathcal{A} \equiv \mathcal{O}(1/g)$ in the Glasma. As we shall discuss now, quantum fluctuations that are parametrically $\mathcal{O}(1)$ and contribute to Γ_{+-} in Eq. (35) play a big role both before ($p^\eta = 0$ modes) and after ($p^\eta \neq 0$ modes) the collision²⁹. We discussed the former previously in the context of the small x evolution of the hadron wavefunctions. We will discuss here the role of these modes after the collision. The $p^\eta \neq 0$ modes only appear after the collision; as we shall discuss subsequently, they play a fundamental role in the thermalization of the Glasma.

1. Dynamics of $p^\eta = 0$ modes: QCD factorization and energy evolution

At NLO ($\mathcal{O}(1)$ relative to the leading $\mathcal{O}(1/\alpha_S)$ contribution) for the inclusive multiplicity in Eq. (35), one of the two terms is the amputated small fluctuations propagator Γ_{+-} and the other is a one loop correction to Γ_{\pm} (or equivalently, the classical field). The $p^\eta = 0$ modes lie close to the beam rapidities $\pm Y_{\text{beam}}$; before the collision, they can be visualized as the fur of wee gluon modes accompanying the valence partons moving along the light cone.

After the collision, the valence partons are stripped of the small x wee gluon modes which

²⁹ p^η is the Fourier conjugate of the spacetime rapidity η .

then populate $p^\eta \neq 0$. The surviving $p^\eta = 0$ modes are valence modes and the quasi-static cloud of large x partons than accompany them into the fragmentation region of the nuclear collision. Thus $p^\eta = 0$ modes after the collision are likely not very interesting from the perspective of thermalization at central rapidities.

Before the collision, all one has are the $p^\eta = 0$ modes. These modes are further separated into sources and fields with the latter dynamically absorbed into the former via the “small x ” evolution of the weight functionals $W_{Y_{\text{beam}\pm Y}}[\rho_{1,2}]$ of each of the comoving nuclei. This however requires a factorization of the quantum fluctuations of each of the two nuclei from each other.

The resulting factorized form of Eq. (40) can be proven to leading logarithmic accuracy in x (Gelis *et al.*, 2008a,b). An important ingredient in the proof is the structure of the cut propagator $G_{+-}(u, v) \propto \int \frac{d^2 k_\perp dk^+}{k^+} e^{ik^+(u^- - v^-) + i\frac{k_\perp^2}{2k^+}(u^+ - v^+)}$. If the spacetime points u and v reside on one of the nuclei, say moving along x^+ , then $u^- \approx v^-$ and one of the phases vanishes. The other phase oscillates rapidly when $k^+ \rightarrow 0$ giving a convergent contribution. However for $k^+ \rightarrow \infty$, it converges to unity, and one obtains a logarithmic divergence dk^+/k^+ which is the source of the large logs resummed in the small x evolution of the nucleus.

In the case where quantum fluctuations in the two nuclei could “talk” to each other before the collision, the spacetime points u and v reside respectively on the lightcones of the incoming nuclei corresponding to $u^\pm - v^\pm \neq 0$. The phases therefore oscillate rapidly when either $k^\pm \rightarrow \infty$ and there are no logarithmic divergences from such contributions. The only possible region where such fluctuations may contribute is when the nuclei overlap. The area of this region is $x^+ x^- = \frac{1}{P^+ P^-} \sim \frac{1}{s}$; such contributions are therefore suppressed by the squared c.m. energy.

Thus the factorized form in Eq. (40) at LLx is satisfied to high accuracy, and one can replace $W_{Y_{\text{beam}\pm Y}}^{\text{MV}}[\rho_{1,2}] \rightarrow W_{Y_{\text{beam}\pm Y}}[\rho_{1,2}]$, where the latter satisfies the JIMWLK equation in Eq. (16). This allows one to go beyond the boost invariant MV expression and to treat the dynamical evolution (in Y) of the weight functionals in the two nuclei. While our arguments are suggestive that the factorization theorem can be extended to NLLx, a formal proof is lacking.

As Y_{beam} increases with increasing energy, the W 's in Eq. (40) describe the energy evolution of the inclusive multiplicity³⁰. Running coupling corrections, that are part of the NLLx

³⁰ This LLx result is implicitly assumed in the 3+1-D IP-Glasma simulations (Schenke and Schlichting,

contributions, improve the accuracy of the computations significantly. In future, one may anticipate using the NLLx JIMWLK Hamiltonian as a systematic improvement to describing energy evolution and rapidity correlations in heavy-ion collisions.

The details of the factorization of the W 's, and their energy evolution, are crucial for phenomenology because they dictate concretely the dependence of final state observables (such as the energy density and correlators thereof) on the saturation scales in the wavefunctions of the colliding nuclei.

2. Dynamics of $p^n \neq 0$ modes: plasma instabilities and the classical-statistical approximation

The $p^n \neq 0$ modes are generated right after the collision when the sources become time dependent and produce gluon modes away from the rapidities of the beams. At NLO, their contribution to the gluon spectrum, for a fixed distribution of color sources, can be written as (Gelis *et al.*, 2007)

$$\begin{aligned} \frac{dN_{\text{NLO}}}{dY d^2p_\perp} &= \frac{1}{16\pi^3} \int d^4x d^4y e^{ip \cdot (x-y)} \partial_x^2 \partial_y^2 \sum_\lambda \epsilon_\mu^\lambda \epsilon_\nu^\lambda \\ &\times [\mathcal{A}^\mu(x) \delta \mathcal{A}^\nu(y) + \delta \mathcal{A}^\mu(x) \mathcal{A}^\nu(y) + G_{+-}(x, y)] , \end{aligned} \quad (54)$$

where ϵ_μ^λ is a gluon polarization vector of helicity λ . The first two terms in this expression represent the NLO contribution to $\Gamma_+(x)\Gamma_-(y)$ in Eq. (35), with $\delta \mathcal{A}$ the one-loop correction to the classical field $\mathcal{A} \equiv \mathcal{A}[\rho_1, \rho_2]$, and the last term represents Γ_{+-} , which first appears at NLO.

Let's first consider the cut propagator term G_{+-} in this expression. Its contribution to the NLO multiplicity can be written as

$$\sum_{\lambda, \lambda'} \int \frac{d^3k}{(2\pi)^3 2E_k} \left| \int_{x^0 \rightarrow \infty} d^3x e^{ip \cdot x} (\partial_x^0 - iE_q) \epsilon_\mu^\lambda a_{\lambda'k}^\mu \right|^2 , \quad (55)$$

where $a_{\lambda'ak}^\mu(x)$ is a small fluctuation field of $\mathcal{O}(1)$ about \mathcal{A}^μ with the plane wave initial condition $e_{\lambda'}^\mu T^a e^{ik \cdot x}$, where T^a are the $SU(3)$ generators in the adjoint representation³¹. Note that the structure above is analogous to Eq. (38) except that the classical field is replaced by the small fluctuation field. The latter obeys the small fluctuation equations of

2016).

³¹ For compactness, we will suppress color indices henceforth.

motion, and its solution can be expressed as

$$a^\mu(x) = \int_{\tau=0^+} d^3u [a(y) \cdot \mathbf{T}_y] \mathcal{A}^\mu(x), \quad (56)$$

where \mathbf{T}_y is a linear operator that corresponds to a shift of the initial data on the classical fields and their derivatives (Dusling *et al.*, 2011b; Gelis *et al.*, 2008a),

$$a(y) \cdot \mathbf{T}_y = a^\mu(y) \frac{\delta}{\delta \mathcal{A}^\mu(y)} + (\partial^\nu a^\mu(y)) \frac{\delta}{\delta (\partial^\nu \mathcal{A}^\mu)}, \quad (57)$$

on the initial spacelike surface at $\tau = 0^+$.

The key insight provided by Eq. (56) is that to compute the small fluctuation field at a spacetime point x in the forward lightcone, it is sufficient to know the small fluctuation field at $\tau = 0^+$, rather than solve the small fluctuation equations on a time-dependent background. We will return to this point shortly.

Plugging Eq. (56) into Eq. (55), and thence into Eq. (54), one obtains

$$\begin{aligned} \frac{dN_{\text{NLO}}}{dY d^2p_\perp} &= \left[\int_{\Sigma_y} [\delta \mathcal{A}(y) \cdot \mathbf{T}_y] + \int_{\Sigma_y, \Sigma_z} [\Gamma_2(y, z) \cdot \mathbf{T}_y \mathbf{T}_z] \right]_{\tau=0^+} \\ &\times \frac{dN_{\text{LO}}}{dY d^2p_\perp}, \end{aligned} \quad (58)$$

where $\Sigma_y = \int d^3y$ denotes the initial spacelike surface $\tau = 0^+$ and

$$\Gamma_2(y, z) = \sum_\lambda \int \frac{d^3k}{(2\pi)^3 2E_k} a_{+k\lambda}(y) a_{-k\lambda}(z), \quad (59)$$

is the small fluctuation propagator evaluated on this surface³².

This NLO result is however not suppressed parametrically by $\mathcal{O}(\alpha_S)$ relative to the LO result because the LO Glasma is very unstable to small fluctuations:

$$\mathbf{T}_y \mathcal{A}(x) \sim \frac{\delta \mathcal{A}(x)}{\delta \mathcal{A}(y)} \sim g e^{\sqrt{\gamma_{\text{inst.}} \tau}}, \quad (60)$$

where $\gamma_{\text{inst.}}$, parametrically of order Q_S , denotes the growth rate of the instability. This exponential growth of small fluctuations in Eq. (56) with $\sqrt{\tau}$ is clearly demonstrated in Fig. 9 from 3+1-D numerical simulations of the YM equations for an η -dependent fluctuation $a(\eta)$ on top of the boost invariant Glasma background (Romatschke and Venugopalan, 2006a,b).

³² Discussions of the computation of this propagator at $\tau = 0^+$ can be found in (Dusling *et al.*, 2011b; Epelbaum and Gelis, 2013; Fukushima *et al.*, 2007).

The very small values of g in the plot³³ are chosen to ensure that the classical-statistical approximation is satisfied in the numerical simulations. This point is discussed further in Section IV.

The existence of such instabilities was previously predicted (Mrówczyński, 1993) and studied with the context of a finite temperature hard thermal loop effective field theory (Atems *et al.*, 2013; Rebhan *et al.*, 2005). They are understood to be analogous to Weibel instabilities familiar in plasma physics (Arnold *et al.*, 2003a); for a recent review, we refer the reader to (Mrówczyński *et al.*, 2017).

As a result of the instability, the exponentially growing small fluctuations can become of the order of the LO classical field for $\tau \sim \frac{1}{\gamma_{\text{inst}}} \log^2 \frac{1}{\alpha_S}$. In a so-called classical-statistical approximation (Aarts and Berges, 2002), these leading instabilities can be resummed to all orders, modifying Eq. (58) to

$$\frac{dN_{\text{resum}}}{dY d^2p_{\perp}} = \int [Da] F[a] \frac{dN_{\text{LO}}}{dY d^2p_{\perp}} [\mathcal{A} + a], \quad (61)$$

where $F[a] \sim \exp\left(-\int_{\Sigma_y \Sigma_z} a(y) \Gamma_2^{-1}(y, z) a(z)\right)$.

To conclude our discussion of the classical-statistical approximation, as a final step, we need to perform the average of the color sources to obtain the inclusive multiplicity distribution at early times in the Glasma:

$$\begin{aligned} \frac{\langle\langle dN \rangle\rangle}{dY d^2p_{\perp}} &= \int [D\rho_1][D\rho_2] W_{Y_{\text{beam}-Y}}[\rho_1] W_{Y_{\text{beam}+Y}}[\rho_2] \\ &\times \int [Da] F[a] \frac{dN_{\text{LO}}}{dY d^2p_{\perp}} [\mathcal{A} + a]. \end{aligned} \quad (62)$$

This result of course applies to other inclusive quantities such as components of stress-energy tensor given in Eq. (46).

In the classical-statistical approximation, the one loop correction to the classical field ($\delta\mathcal{A}$) is suppressed at early times relative to the G_{+-} term we consider here. In general, the classical-statistical approximation does not account for the full quantum evolution of the Glasma fields. In the next section, we will discuss the dynamical power counting of quantum fields within the framework of the two-particle irreducible (2PI) effective action that specifies

³³ At RHIC (LHC) energies, $g^2\mu \propto Q_S \sim 1 - 2$ GeV on the x-axis of Fig. 9. With these values, $\tau \gg 10$ fm, the typical life time of such a collision. However for $g \sim 10^{-5}$, from QCD running, $g^2\mu$ is larger than the Planck scale. The takeaway message from Fig. 9 is the functional form of the fit and not the absolute values.

the range of validity of the classical-statistical approximation, the nature of the corrections beyond, as well as numerical results from the implementation of this approximation and the consequences thereof.

IV. FAR-FROM-EQUILIBRIUM GLUON AND QUARK PRODUCTION: FROM PLASMA INSTABILITIES TO NON-THERMAL ATTRACTORS

We have seen in the previous section that the overoccupied Glasma is unstable with respect to small quantum fluctuations which break longitudinal boost invariance. As noted there, the growth of fluctuations is caused by primary (Weibel-like ([Mrówczyński *et al.*, 2017](#))) instabilities ([Fukushima and Gelis, 2012](#); [Romatschke and Venugopalan, 2006a,b](#)). However, there are also secondary instabilities that arise due to the nonlinear interactions of unstable modes ([Berges and Schlichting, 2013](#)). The fluctuations that are initially small grow with time and an over-occupied plasma emerges on a time scale $Q_S \tau \sim \log^2(\alpha_S^{-1})$.

At this stage, the details about the initial spectrum of fluctuations is effectively lost as a consequence of the strongly nonlinear evolution. The apparent loss of information at such an early stage gives rise to decoherence towards a more isotropic equation of state in this prethermalization regime ([Arnold *et al.*, 2005](#); [Berges *et al.*, 2004](#); [Dusling *et al.*, 2011a](#)). Subsequently, a universal scaling behavior emerges far from equilibrium with increasing anisotropy ([Berges *et al.*, 2015b](#)), which is described in terms of non-thermal attractor solutions ([Berges *et al.*, 2014b,c](#)), representing the first stage of the “bottom-up” thermalization scenario ([Baier *et al.*, 2001](#); [Bodeker, 2005](#)).

In the following, we will describe how this nonlinear behavior emerges starting from the underlying quantum field theory, formulated as an initial value problem in time. Essential aspects of the far-from-equilibrium quantum evolution can be approximated by a controlled weak-coupling expansion around the full (non-perturbative) classical-statistical theory, first pointed out in the context of scalar field theories ([Aarts and Berges, 2002](#); [Khlebnikov and Tkachev, 1996](#); [Son, 1996](#)) and then extended to include fermions ([Aarts and Smit, 1999](#); [Berges *et al.*, 2011](#); [Borsanyi and Hindmarsh, 2009](#); [Kasper *et al.*, 2014](#); [Saffin and Tranberg, 2011](#)).

In strong field QCD, this corresponds to an expansion in $\alpha_S \equiv g^2/(4\pi)$, where the leading order contribution includes the full classical-statistical theory of gluons as described in

Sec. III. The next-to-leading order contributions take into account the back-reaction of the quarks onto the gluons, and encodes important quantum effects such as anomalies. The non-equilibrium time evolution of gluons with dynamical quarks has been studied numerically on the lattice in Refs. (Gelfand *et al.*, 2016; Gelis *et al.*, 2006a; Tanji and Berges, 2018).

Such an expansion around the full classical-statistical field theory breaks down on the time scale $Q_S \tau \sim \alpha_S^{-3/2}$ (Baier *et al.*, 2001; Berges *et al.*, 2014b), where typical gluon occupancies become of order unity. To continue further and capture the late-time evolution towards local thermal equilibrium, one employs a resummed perturbative description of quantum field theory in an on-shell approximation. This also underlies the effective kinetic theory we will discuss in Sec. V.

The range of validity of both approximation schemes, the expansion around the classical-statistical theory at early times, as well as effective kinetic theory employed at late times with their common overlap at intermediate times (Jeon, 2005; Mueller and Son, 2004), can be efficiently discussed using the two-particle irreducible (2PI) quantum effective action (Baym, 1962; Cornwall *et al.*, 1974) on the closed time path (Berges, 2004a; Calzetta and Hu, 1988).

A. Non-equilibrium time evolution equations from the quantum effective action

Quantum evolution equations can be formulated in terms of expectation values of field operators, such as the macroscopic field $\mathcal{A}(x)$ and the connected two-point correlation function or propagator $G(x, y)$ on the closed time contour \mathcal{C} we introduced in Sec. III. In practice, the space-time evolution of the one-point, two-point or higher-point correlation functions cannot be computed for the full quantum theory without approximations. However one can formally write down exact evolution equations, which provide an efficient starting point justifying the applicability of systematic expansion schemes.

Writing for simplicity only the gauge field part, the evolution equations for connected one and two-point correlation functions follow from the stationarity of the 2PI effective action (Baym, 1962; Cornwall *et al.*, 1974)

$$\begin{aligned} \Gamma[\mathcal{A}, G] = & S[\mathcal{A}] + \frac{i}{2} \text{tr}(\ln G^{-1}) + \frac{i}{2} \text{tr}(G_0^{-1}(\mathcal{A}) G) \\ & + \Gamma_2[\mathcal{A}, G] + \text{const} , \end{aligned} \quad (63)$$

where $iG_{0;ab}^{-1,\mu\nu}(x, y; \mathcal{A}) \equiv \delta^2 S[\mathcal{A}] / \delta \mathcal{A}_\mu^a(x) \delta \mathcal{A}_\nu^b(y)$ is the inverse propagator with Lorentz in-

dices μ, ν and color indices $a, b = 1, \dots, N_c^2 - 1$ for $SU(N_c)$ gauge theories with classical action $S[\mathcal{A}]$. Here $\Gamma_2[\mathcal{A}, G]$ contains all two-particle irreducible contributions, which lead to the self-energy $\Pi_{ab}^{\mu\nu}(x, y) \equiv 2i\delta\Gamma_2[\mathcal{A}, G]/\delta G_{\mu\nu}^{ab}(x, y)$. Higher n -point correlation functions can be obtained from $\Gamma[\mathcal{A}, G]$ by functional differentiation with respect to the fields, once the solutions for \mathcal{A} and G are known.

1. Macroscopic field, spectral and statistical functions

The full quantum evolution equation for the macroscopic field is obtained from the stationarity of $\Gamma[\mathcal{A}, G]$ with respect to variations in $\mathcal{A}(x)$, and is given by

$$\frac{\delta S[\mathcal{A}]}{\delta \mathcal{A}_\mu^a(x)} = -J_a^\mu(x) - \frac{i}{2} \text{tr} \left[\frac{\delta G_0^{-1}(\mathcal{A})}{\delta \mathcal{A}_\mu^a(x)} G \right] - \frac{\delta \Gamma_2[\mathcal{A}, G]}{\delta \mathcal{A}_\mu^a(x)}. \quad (64)$$

For the discussion of the evolution equations for two-point functions, it is convenient to introduce spectral and statistical components by

$$G_{\mu\nu}^{ab}(x, y) \equiv F_{\mu\nu}^{ab}(x, y) - \frac{i}{2} \rho_{\mu\nu}^{ab}(x, y) \text{sgn}_c(x^0 - y^0) \quad (65)$$

where the spectral function $\rho(x, y)$ is associated with the expectation value of the commutator of two fields and the statistical function $F(x, y)$ by the anti-commutator for bosons³⁴ (Berges, 2004a). A similar decomposition can be done for the self-energy, $\Pi(x, y) \equiv -i\Pi^{(0)}(x)\delta(x-y) + \Pi^{(F)}(x, y) - i\Pi^{(\rho)}(x, y)\text{sgn}_c(x^0 - y^0)/2$, where $\Pi^{(0)}$ describes a local contribution to the self-energy. With this notation, the equations for spectral and statistical two-point correlation functions, which follow from the stationarity of $\Gamma[\mathcal{A}, G]$ with respect to variations in G , can be written as (Berges, 2004a)

$$\begin{aligned} [iG_{0,ac}^{-1,\mu\gamma}(x; \mathcal{A}) + \Pi_{ac}^{(0),\mu\gamma}(x)] \rho_{\gamma\nu}^{cb}(x, y) &= - \int_{y^0}^{x^0} dz \Pi_{ac}^{(\rho),\mu\gamma}(x, z) \rho_{\gamma\nu}^{cb}(z, y), \\ [iG_{0,ac}^{-1,\mu\gamma}(x; \mathcal{A}) + \Pi_{ac}^{(0),\mu\gamma}(x)] F_{\gamma\nu}^{cb}(x, y) &= - \int_{t_0}^{x^0} dz \Pi_{ac}^{(\rho),\mu\gamma}(x, z) F_{\gamma\nu}^{cb}(z, y) + \int_{t_0}^{y^0} dz \Pi_{ac}^{(F),\mu\gamma}(x, z) \rho_{\gamma\nu}^{cb}(z, y), \end{aligned} \quad (66)$$

³⁴ In terms of the Keldysh components of the propagator employed in Sec. III, this reads:

$$\begin{aligned} G_{++}(x, y) &= F(x, y) - i\rho(x, y)\text{sgn}(x^0 - y^0)/2, \\ G_{--}(x, y) &= F(x, y) + i\rho(x, y)\text{sgn}(x^0 - y^0)/2, \\ G_{+-}(x, y) &= F(x, y) + i\rho(x, y)/2, \\ G_{-+}(x, y) &= F(x, y) - i\rho(x, y)/2. \end{aligned}$$

Here we denote $\int_a^b dz \equiv \int_a^b dz^0 \int d^3z \sqrt{-g(z)}$ with given initial time t_0 and g is the determinant of the metric. The inverse propagator enters Eq. (66) as

$$\begin{aligned} iG_{0,ab}^{-1,\mu\nu}(x; \mathcal{A}) &= (-g)^{-\frac{1}{2}} D_\gamma^{ac}(\mathcal{A}) (-g)^{\frac{1}{2}} g^{\gamma\alpha} g^{\mu\nu} D_\alpha^{cb}(\mathcal{A}) \\ &- (-g)^{-\frac{1}{2}} D_\gamma^{ac}(\mathcal{A}) (-g)^{\frac{1}{2}} g^{\gamma\nu} g^{\mu\alpha} D_\alpha^{cb}(\mathcal{A}) - g f^{abc} \mathcal{F}_c^{\mu\nu}(\mathcal{A}) \end{aligned}$$

with the covariant derivative $D_\mu^{ab}(\mathcal{A}) = \delta^{ab} \partial_\mu - g f^{abc} \mathcal{A}_\mu^c$ and $\mathcal{F}_{\mu\nu}^a(\mathcal{A}) = \partial_\mu \mathcal{A}_\nu^a - \partial_\nu \mathcal{A}_\mu^a + g f^{abc} \mathcal{A}_\mu^b \mathcal{A}_\nu^c$ is the field strength tensor.

The non-zero spectral and statistical parts of the self-energy $\Pi^{(\rho/F)}(A, F, \rho)$ on the r.h.s and the space-time local part $\Pi^{(0)}(F)$ on the l.h.s of these coupled set of equations make the evolution equations nonlinear in the fluctuations. In general, they contain contributions from the interaction vertices of QCD, where in addition to the standard three- and four-vertices there is a three-gluon vertex associated with the presence of a non-vanishing field expectation value. The explicit expressions for the derivatives on the r.h.s of Eq. (64) and the self-energy contributions entering Eq. (66) are given to three loop order (g^6) in Ref. (Berges, 2004b), and the corresponding expressions in co-moving (τ, η) coordinates can be found in Ref. (Hatta and Nishiyama, 2012). The inclusion of quark degrees of freedom follows along the same lines and can also be found in Ref. (Berges, 2004b).

The non-equilibrium initial conditions for the coupled evolution equations Eq. (64) and Eq. (66) can be formulated in (τ, η) coordinates (and Fock-Schwinger gauge $\mathcal{A}_\tau = 0$) for the Glasma initial conditions we discussed in the previous section. The gauge field expectation values in Eq. (39) correspond to the Glasma background fields, while the spectral and statistical two-point functions describe the fluctuations. The former satisfy at all times the equal-time commutation relations

$$\begin{aligned} \rho_{\mu\nu}^{ab}(x, y) \Big|_{x^0=y^0} &= 0, \\ \partial_{x^0} \rho_{\mu\nu}^{ab}(x, y) \Big|_{x^0=y^0} &= -\delta^{ab} \frac{g_{\mu\nu}}{\sqrt{-g(x)}} \delta(\vec{x} - \vec{y}), \\ \partial_{x^0} \partial_{y^0} \rho_{\mu\nu}^{ab}(x, y) \Big|_{x^0=y^0} &= 0. \end{aligned} \tag{67}$$

2. Resummed evolution equations to leading order

In order to isolate the leading contributions one has to take into account the strong external currents $J \sim \mathcal{O}(1/g)$ in the Glasma, which induce non-perturbatively large background

fields $\mathcal{A} \sim \mathcal{O}(1/g)$. In contrast, the statistical fluctuations F originate from the vacuum and are therefore initially $\mathcal{O}(1)$. The spectral function ρ encodes the equal-time commutation relations and is therefore parametrically $\mathcal{O}(1)$ at any time.

Considering only the leading contributions in a weak coupling expansion, the evolution equation Eq. (64) reduces to the classical Yang-Mills equation for the classical Glasma field \mathcal{A} , and the equations for the spectral and statistical two-point correlation functions read

$$\begin{aligned} iG_{0,ac}^{-1,\mu\gamma}(x; \mathcal{A}) \rho_{\gamma\nu}^{cb}(x, y) &= 0, \\ iG_{0,ac}^{-1,\mu\gamma}(x; \mathcal{A}) F_{\gamma\nu}^{cb}(x, y) &= 0. \end{aligned} \tag{68}$$

Here sub-leading contributions are suppressed by at least a factor of g^2 relative to the leading contribution.

At this order the evolution of the Glasma background fields decouples from that of the fluctuations. The evolution of vacuum fluctuations of the initial state is taken into account by Eq. (68) to linear order in the fluctuations. This was an important assumption in the derivation in Sec. III.D.2 and was exploited in Ref. (Dusling *et al.*, 2011b; Epelbaum and Gelis, 2013) to obtain the spectrum of initial fluctuations right after the collision. These approximations are therefore only valid for short enough evolution times such that the fluctuations have parametrically small values.

In general, it is difficult to find suitable approximation schemes for the 2PI effective action in gauge theories beyond the linear regime (Arrizabalaga and Smit, 2002). However it provides a formal justification of a resummed coupling expansion of the quantum field theory around the full classical-statistical solution; as we will soon discuss, this scheme can be implemented numerically on a lattice to describe dynamics far from equilibrium.

Furthermore, as we shall also discuss below, the different dynamical stages of the Glasma undergoing a non-equilibrium instability at early times can be conveniently understood analytically from power counting in the 2PI effective action beyond the linear regime (Berges and Schlichting, 2013). Not least, the 2PI effective action approach allows for efficient on-shell approximations employing a gradient expansion; these lead to effective kinetic equations describing non-equilibrium evolution at later times (Blaizot and Iancu, 2002). We will discuss these equations and their numerical solutions in Sec. V.

B. Nonlinear evolution of plasma instabilities

In Sec. III.D.2, we demonstrated that the highly anisotropic state of the Glasma is unstable with respect to small quantum fluctuations. In the language of Eq. (68), these correspond to the quasi-exponential growth of the statistical function, (Berges *et al.*, 2014e; Fukushima, 2007; Fukushima and Gelis, 2012; Romatschke and Venugopalan, 2006a,b)

$$F_{\mu\nu}^{ab}(\tau, \tau, x_T, y_T, \nu) \sim \exp \left[\Gamma(\nu) \sqrt{g^2 \mu \tau} \right], \quad (69)$$

where we recall $g^2 \mu \propto Q_S$ and $\Gamma(\nu)$ is a function of order unity for characteristic modes ν that are Fourier coefficients with respect to the relative rapidity³⁵

$$F_{\mu\nu}^{ab}(x, y) = \int \frac{d\nu}{2\pi} F_{\mu\nu}^{ab}(x, y, \nu) e^{i\nu(\eta_x - \eta_y)}. \quad (70)$$

1. Dynamical power counting

The behavior of the quantum evolution beyond the linear regime is captured by a dynamical power counting scheme (Berges *et al.*, 2012, 2009, 2008b; Berges and Serreau, 2003). Self-energy corrections are classified according to powers of the coupling constant g , of the background field \mathcal{A} , and of the statistical fluctuations F . Thus a generic self-energy contribution is of order $g^n F^m \mathcal{A}^l \rho^k$ and contains the suppression factor from powers of the coupling constant (n) as well as the enhancement due to a parametrically large background field (l) and large fluctuations (m). The “weight” of the spectral function (k) remains parametrically of order one at all times as encoded in the equal-time commutation relations, see Eq. (67).

For the strong macroscopic fields $\mathcal{A} \sim 1/g$ in the Glasma, sizable self-energy corrections occur once fluctuations grow to be as large as $F \sim 1/g^{(n-l)/m}$ for characteristic modes. This yields a hierarchy of time scales, where diagrammatic contributions with smaller values of $r = (n - l)/m$ become important at earlier times (since $g \ll 1$) compared to contributions with larger values of r .

The quasi-exponential growth stops when fluctuations become of $\mathcal{O}(1/g^2)$, where they saturate. At $\mathcal{O}(1/g^2)$ the fluctuations lead to sizable contributions from every given loop-order and the perturbative power-counting scheme breaks down. The corresponding time scale may be estimated from the one-loop correction

³⁵ Here ν is equivalent to the momentum p^η in the (τ, η) coordinate system.

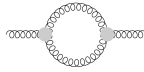


which has $r = 2$ ($n = 2, l = 0, m = 1$). Using the quasi-exponential growth behavior Eq. (69) the factor of $\sim g^2$ from the vertex is compensated by the propagator line $F \sim \mathcal{O}(1/g^2)$ at time

$$\tau_{\text{occ}} \stackrel{g \ll 1}{\sim} \frac{1}{Q_S} \log^2(g^{-2}) , \quad (71)$$

which denotes the characteristic time for the end of the instability regime.

The earliest time for nonlinear amplification to set in can be inferred from the diagram with the lowest value of r . For our problem, this is realized by the one-loop contribution with $r = 1$ ($n = 2, l = 0, m = 2$),



which already becomes sizable when $F \sim \mathcal{O}(1/g)$, where the two propagator lines compensate for the two powers of the coupling. Using again the quasi-exponential growth behavior Eq. (69) of the primary unstable modes, the time at which this $\mathcal{O}(1/g)$ correction becomes important relative to the $\mathcal{O}(1/g^2)$ in Eq. (71) is $\sim \tau_{\text{occ}}/4$ in the weak-coupling limit. This is followed by a series of higher-loop corrections, all leading to a fast broadening of the primary unstable range in rapidity wave number ν (Berges and Schlichting, 2013).

2. Classical-statistical field theory limit

The evolution of the Glasma to later times than τ_{occ} is non-perturbative. While in scalar quantum field theories there are different ways to address it, an example being large- N resummation techniques (Aarts *et al.*, 2002; Berges, 2002), for gauge theories the most frequently employed approach is the classical-statistical approximation. The latter can be understood starting from the full quantum 2PI effective action by a set of well-defined approximations.

One first notes that a given propagator line of a diagram may be associated to either the statistical (F) or the spectral (ρ) correlation function. The set of diagrams included in the classical-statistical approximation can be identified as those corrections that contain the most powers of the statistical function relative to powers of the spectral function for each

type of diagram (Aarts and Berges, 2002). This corresponds to resumming the leading effects of the instability to all orders in the coupling constant (Dusling *et al.*, 2011b; Epelbaum and Gelis, 2013).

Therefore, in contrast to expansions at fixed loop-orders, the classical-statistical approach provides a controlled approximation scheme which is particularly well suited for problems involving large statistical fluctuations. Specifically, for the large $F \sim \mathcal{O}(1/g^2)$ values encountered at the end of the plasma instability regime, neglecting powers of $\rho \sim \mathcal{O}(1)$ compared to those of F represents a systematic weak-coupling approximation of a system that is strongly correlated because of the large fluctuations.

While leading order in this expansion corresponds to the full non-equilibrium classical-statistical field theory for the gauge fields, genuine quantum corrections for the dynamics arise. As we will soon discuss, the dynamical evolution of quarks and anti-quarks represent a class of such genuine quantum corrections (Tanji and Berges, 2018).

We can conclude from this discussion that for the far-from-equilibrium overoccupied Glasma there is a well-controlled mapping of the weak-coupling quantum dynamics for correlation functions onto a classical-statistical field theory. The latter can be simulated numerically on a lattice. In principle, starting with large field amplitudes, the mapping involves two steps: I) The field is separated into a large coherent part and a small fluctuation part in which one linearizes the field evolution equations. The set of linearized equations is given by Eq. (68). II) Though small initially, the fluctuations grow because of plasma instabilities. Once they become sizable, the time evolution of the linearized equations is stopped and the results are used as input for a subsequent classical-statistical simulation which is fully non-linear.

A virtue of the two-step procedure of mapping the original quantum theory to the classical description is that it has a well-defined continuum limit, enabling one to recover the full physical results for certain quantities in the weak-coupling limit (Aarts and Smit, 1998). In scalar field theories, this is well tested by comparisons to fully quantum calculations using 2PI effective action techniques (Aarts and Berges, 2002) and likewise, when scalar fields are coupled to fermions (Berges *et al.*, 2014d). The mapping was first applied in cosmology in the context of post-inflationary scalar preheating dynamics (Khlebnikov and Tkachev, 1996; Son, 1996).

The two-step procedure is in practice replaced by a simplified description whereby one

starts with the fully non-linear classical-statistical description already from the initial time in the strong-field regime. This can be well controlled, for a given regularization with lattice spacing a in the weak coupling limit, by ensuring that vacuum fluctuations from modes with momenta near the cutoff $\sim 1/a$ do not dominate the dynamics. Several studies investigated the range of validity of this simplified “one-step” mapping of the original quantum theory onto the classical-statistical description—see for instance Ref. (Epelbaum *et al.*, 2014); the limitations of the classical-statistical approximation have been studied in detail in Ref. (Berges *et al.*, 2014a) for scalar field theories.

Fig. 10 provides snapshots of the time evolution of the gluon distribution for an analytically computed initial spectrum of fluctuations given in Ref. (Epelbaum and Gelis, 2013) employing the fully non-linear classical-statistical description already from the initial time in the strong-field regime. The non-equilibrium evolution is computed numerically using the Wilson formulation of lattice gauge theory in real time (Berges *et al.*, 2014e). In addition to gauge invariant quantities, (Coulomb type) gauge fixed distribution functions can be extracted for comparison to effective descriptions such as kinetic theory. The definition of the distribution function shown in Fig. 10 employs the two-point correlation function of the gauge field following Ref. (Berges *et al.*, 2014b). While the gluon distribution as a function of transverse momentum p_T and rapidity wave number ν is dominated by the boost-invariant ($\nu = 0$) background at early times $Q_S\tau \sim 1$, an over-occupied plasma emerges on a time scale $Q_S\tau \sim \log^2(\alpha_S^{-1})$.

A corresponding evolution is found irrespective of the details of the fluctuations in the initial conditions. Fig. 11 shows the example of the gauge-invariant longitudinal pressure-pressure correlation function for different rapidity wave numbers ν , averaged over transverse coordinates, as a function of time (Berges and Schlichting, 2013). The evolution starts from initial conditions with simplified initial fluctuations taken as an additive contribution to the strong background gauge fields. While primary unstable modes at non-zero rapidity wave number exhibit quasi-exponential amplification first, secondary instabilities with enhanced growth rates set in with a delay for higher momentum modes due to the nonlinear processes described above. Subsequently the instability propagates towards higher momenta until saturation occurs and the system exhibits a much slower dynamics (Berges and Schlichting, 2013; Romatschke and Venugopalan, 2006b). This behavior is similar to that observed in non-expanding gauge theories (Berges *et al.*, 2009, 2008b) and cosmological models for scalar

field evolution (Berges and Serreau, 2003).

C. Non-thermal attractor

The plasma instabilities lead to a far-from-equilibrium state at time $Q_S \tau_{\text{occ}} \sim \log^2(\alpha_S^{-1})$, which exhibits an over-occupied gluon distribution whose characteristic properties may be parametrized as

$$f(p_T, p_z, \tau_{\text{occ}}) = \frac{n_0}{2g^2} \Theta\left(Q - \sqrt{p_T^2 + (\xi_0 p_z)^2}\right). \quad (72)$$

Here n_0 denotes the magnitude of the initial over-occupancy of the plasma, averaged over spin and color degrees of freedom up to the momentum Q . The momentum scale Q is of comparable magnitude, albeit non-trivially related, to the saturation scale Q_S . The degree of anisotropy of the gluon distribution in momentum space is described by the parameter ξ_0 .

While Eq. (72) does not capture all details of the state at τ_{occ} , a precise matching to the Glasma appears inessential because of the existence of an attractor solution for the subsequent dynamics. In fact, variation of the parameters of Eq. (72) can be used to visualize attractor properties.

Fig. 12 illustrates the evolution of the plasma in the occupancy–anisotropy plane, originally introduced in Refs. (Kurkela and Moore, 2011a,b). The horizontal axis shows the characteristic “hard scale” occupancy $n_{\text{Hard}}(\tau) = f(p_{\perp} \simeq Q, p_z = 0, \tau)$, while the vertical axis shows the momentum-space anisotropy, which can be characterized in terms of the ratio of typical longitudinal momenta (Λ_L) to the typical transverse momenta (Λ_T). These typical longitudinal and transverse momentum scales are gauge invariant quantities expressed as ratios of the product of covariant derivatives of the field strength tensor normalized by the energy density (Berges *et al.*, 2014c). In a weak coupling limit, these are proportional to $\langle p_{\perp} \rangle$ and $\langle p_z \rangle$ for a single particle distribution $f(p_{\perp}, p_z, \tau)$.

The blue lines in Fig. 12 show a projection of lattice simulation results onto the anisotropy–occupancy plane. The different initial conditions are indicated by blue dots. After some time all curves exhibit a similar evolution along the diagonal, clearly illustrating the presence of a non-thermal attractor independent of the initial conditions. The attractor has a number of interesting properties associated to non-thermal fixed points that we shall discuss in Sec. IV.C.1–Sec. IV.C.3.

1. *Far-from-equilibrium universal scaling*

Apart from the insensitivity to details of the initial conditions, the Glasma's evolution exhibits a universal scaling behavior such that the dynamics in the vicinity of the attractor becomes self-similar. In the weak coupling limit, the gluon distribution can be expressed in terms of a time independent scaling function f_S (Berges *et al.*, 2014b),

$$f(\tau, p_T, p_z) = \frac{(Q\tau)^\alpha}{\alpha_S} f_S\left((Q\tau)^\beta p_T, (Q\tau)^\gamma p_z\right). \quad (73)$$

This scaling behavior is characteristic of the phenomenon of wave turbulence and has been observed in a variety of systems far from equilibrium (Berges *et al.*, 2015a; Micha and Tkachev, 2004). As shown in Fig. 13, the moments of the longitudinal momentum distribution at different times in the evolution (top) collapse into universal curves for each moment m of the single particle distribution. One observes a corresponding behavior for moments of the transverse momentum distribution. This self-similar behavior of the distribution allows one to extract numerically the values of the scaling exponents in Eq. (73) to be $\alpha \simeq -2/3$, $\beta \simeq 0$ and $\gamma \simeq 1/3$ (Berges *et al.*, 2014b).

These values are consistent with those obtained also analytically from small-angle elastic scattering as the dominant process and confirm the onset of the “bottom-up” thermalization scenario (Baier *et al.*, 2001). The competition between longitudinal momentum broadening via small-angle scattering and the red-shift due to the longitudinal expansion leads to a decrease of the typical longitudinal momenta as $p_z/Q \sim (Q\tau)^{-1/3}$, while the typical transverse momenta remain approximately constant, $p_T/Q \sim \text{const}$. At the same time, the gluon occupancy decreases as $f(\tau, p_T \sim Q) \sim \alpha_S^{-1}(Q\tau)^{-2/3}$ and becomes of order unity on a time scale $Q\tau_{\text{quant}} \sim \alpha_S^{-3/2}$ when quantum effects can no longer be neglected. Beyond τ_{quant} , the classical-statistical framework becomes inapplicable and one may resort to an effective kinetic description as will be discussed in Sec. V.

2. *Identifying the weak-coupling thermalization scenario*

In Fig. 12, we showed the predictions of various thermalization scenarios for the momentum anisotropy with decreasing occupancy. These thermalization scenarios are based on estimates in effective kinetic theory and differ primarily in how infrared momentum modes

are treated. Clearly, these differences lead to very different paths in the thermalization process. As the system evolves with decreasing occupancy from the initial $f \sim \alpha_S^{-1}$, classical-statistical field theory simulations accurately capture the physics of the infrared regime. This may be used to distinguish whether a particular thermalization scenario is indeed realized, especially since lattice simulations and effective kinetic theory have an overlapping regime of validity when $1 < f < \alpha_S^{-1}$.

The gray lines in Fig. 12 indicate the different thermalization scenarios put forward in Refs. (BMSS) (Baier *et al.*, 2001), (BD) (Bodeker, 2005), (KM) (Kurkela and Moore, 2011a) and (BGLMV) (Blaizot *et al.*, 2012). Besides the BMSS scenario, which is consistent with the lattice simulation results and is discussed in detail in Sec. V, the BD scenario considers the possibility that plasma instabilities lead to an overpopulation $f \sim 1/\alpha_S$ of modes with $|\mathbf{p}| \lesssim m_D$. The coherent interaction of hard excitations with the soft sector then causes an additional momentum broadening such that the longitudinal momenta of hard excitations fall at a slower rate. A possible variant of the impact of plasma instabilities for the subsequent quantum evolution underlies also the KM scenario. In the BGLMV scenario, elastic scattering is argued to be highly efficient in reducing the anisotropy of the system. This would generate an attractor with a fixed anisotropy such that Λ_L/Λ_T remains constant in time.

The selection of the appropriate effective kinetic theory using lattice simulation data represents the state of the art, and is the basis for the thermalization discussion of Sec. V. The justification of the kinetic description solely based on perturbation theory in its range of validity raises important open questions on how to incorporate the effects of infrared modes.

3. Non-thermal attractors in scalar field theories

Non-thermal attractors in overoccupied weakly coupled field theories have been studied earlier in the context of cosmological (p)reheating and thermalization after inflation in the early universe (Berges *et al.*, 2008a; Micha and Tkachev, 2003, 2004). A large class of inflationary models employs scalar field theories, where an initially coherent inflaton field decays due to non-equilibrium instabilities. These may originate from tachyonic/spinodal dynamics or parametric resonance (Berges and Serreau, 2003; Kofman *et al.*, 1994; Traschen and Brandenberger, 1990). The instabilities lead to overoccupied excitations, whose transient

dynamics can exhibit self-similar evolution.

The dynamics is in general spatially isotropic on large scales, in contrast to the longitudinal expansion relevant for heavy-ion collisions. To compare the two, if we impose the isotropic case of no expansion with overoccupied initial conditions for gauge fields, the gluon distribution function in the self-similar regime obeys $f(t, p) = t^{-4/7} f_S(t^{-1/7} p)$ in three spatial dimensions. This is characteristic of an energy cascade towards higher momentum scale due to weak wave turbulence (Kurkela and Moore, 2011b, 2012; Schlichting, 2012).

In the fixed box case for a relativistic real scalar field theory in the self-similar regime, the distribution function obeys $f^\phi(t, p) = t^{-(d+1)/(2l-1)} f_S^\phi(t^{-1/(2l-1)} p)$ for l -vertex scattering processes (Micha and Tkachev, 2004). For quartic ($l = 4$) self-interactions, the exponents are identical to the gauge theory with the same geometry. However in the presence of spontaneous symmetry breaking, the non-zero field expectation value leads to effective three-vertex scattering processes off the macroscopic field. These analytical estimates have been numerically verified using 2PI effective action techniques in Refs. (Berges and Wallisch, 2017; Shen and Berges, 2020) for a N -component scalar field theory with quartic self-interactions. In classical-statistical simulations, which construct the ensemble averages from individual runs with a non-zero field value, the observed scaling exponents are consistent with the estimates in the presence of an effective three-vertex (Micha and Tkachev, 2004).

In Ref. (Berges *et al.*, 2015b) longitudinally expanding N -component scalar field theories are analyzed starting from over-occupied initial conditions. In the vicinity of the non-thermal attractor, very similar scaling behavior as for the non-Abelian gauge theory is observed. The universal scaling exponents and shape of the scaling function agree well with those obtained for the early stage of the bottom-up thermalization process for gauge theories for not too late times.

As an example, Fig. 14 shows results for the $N = 4$ component scalar theory for intermediate transverse momentum $p_T \sim Q/2$, where the normalized scaling distribution as a function of the rescaled longitudinal momentum is given. All data curves at different times in the scaling regime collapse onto a single curve using the scaling exponents $\alpha = -2/3$, and $\gamma = 1/3$. This scaling curve is seen to be indistinguishable from the corresponding scaling curve for non-Abelian gauge theory, which shares the same scaling exponents. The results provide a striking manifestation of universality far from equilibrium.

D. Far-from-equilibrium separation of scales and ultrasoft scale dynamics

The weakly coupled QCD plasma exhibits a hierarchy of scales in thermal equilibrium at high temperature T , with the separation of hard momenta $\sim T$ dominating the system's energy density, soft (electric screening/Debye) momenta $\sim gT$, and ultrasoft (magnetic) momenta $\sim g^2T$ for $g^2 = 4\pi\alpha_S \ll 1$. A similar separation of scales exists far from equilibrium in the vicinity of the non-thermal attractor, where for comparison we will consider the spatially isotropic case without longitudinal expansion.

Starting from over-occupied initial conditions, in this fixed-box case the gluon distribution function in the self-similar regime obeys $f(t,p) = t^{-4/7} f_s(t^{-1/7}p)$ in three spatial dimensions (Kurkela and Moore, 2011b, 2012; Schlichting, 2012). Accordingly, the time-dependent hard momentum scale dominating the energy density is given by $\Lambda(t) \sim t^{1/7}$. The Debye scale $m_D(t) \sim g\sqrt{\int d^3p f(t,p)/p} \sim t^{-1/7}$ decreases with time (Berges *et al.*, 2014c; Boguslavski *et al.*, 2018; Kurkela and Moore, 2012; Lappi and Peuron, 2017; Mace *et al.*, 2016).

At even lower scales, the dynamics becomes non-perturbative for momenta $K(t)$ where the occupancy reaches $\sim 1/\alpha_S$, and the perturbative notion of a gluon distribution function becomes problematic in this ultrasoft regime. As suggested in Ref. (Kurkela and Moore, 2012), the evolution of the ultrasoft scale may be estimated approximately as $K(t) \sim t^{-2/7}$ using the power law form of the occupation number distribution extracted in the perturbative regime. While initially all characteristic momentum scales are of the same order Q_S , this suggests that during the self-similar evolution a dynamical separation of these scales $K(t) \ll m_D(t) \ll \Lambda(t)$ occurs as time proceeds.

1. Non-equilibrium evolution of the spatial Wilson loop

A proper description of the non-perturbative low momentum regime can be based on gauge-invariant quantities. This should take into account that the infrared excitations of non-Abelian gauge theories are extended objects, which can be computed from Wilson loops (Berges *et al.*, 2019, 2017a, 2008b; Dumitru *et al.*, 2014; Mace *et al.*, 2016). At the magnetic scale, spatial Wilson loops capture the long-distance behavior of gauge fields \mathcal{A} ,

defined as

$$W = \frac{1}{N_c} \text{Tr} \mathcal{P} e^{-ig \int_{\mathcal{C}} \mathcal{A}_i(\mathbf{z}, t) dz_i}, \quad (74)$$

where the index i labels spatial components (Montvay and Munster, 1997). Here \mathcal{P} denotes path ordering along a closed line \mathcal{C} , and the trace is in the fundamental representation of $SU(N_c)$.

The behavior of the spatial Wilson loop for large areas $A \gg 1/Q_s^2$ enclosed by the line \mathcal{C} reflects the long-distance or infrared properties of the strongly correlated system. Similarly to the large-distance behavior of the spatial Wilson loop in a high-temperature equilibrium plasma, the spatial Wilson loop exhibits an area law in the overoccupied regime of the non-equilibrium plasma, i.e. $-\log\langle W \rangle \sim A$ (Berges *et al.*, 2008b; Dumitru *et al.*, 2014; Mace *et al.*, 2016).

However, here the area-law behavior occurs in the self-similar regime of the non-equilibrium evolution. This is demonstrated in Fig. 15, which shows the logarithm of the Wilson loop as a function of the time-rescaled area $\sim t^{-\zeta}A$ with universal scaling exponent ζ (Berges *et al.*, 2019, 2017a). Results for both $SU(2)$ and $SU(3)$ gauge groups are displayed. After taking into account the Casimir color factors, normalizing the data points with $C_F = (N_c^2 - 1)/(2N_c)$ discloses a very similar behavior for $N_c = 2$ and $N_c = 3$ (Berges *et al.*, 2017a). The scaling exponent $\zeta = 0.54 \pm 0.04$ (stat.) ± 0.05 (sys.) agrees for both gauge groups to very good accuracy (Berges *et al.*, 2019). This value of the scaling exponent for the ultra-soft scale $\sqrt{\sigma}$ obtained from lattice simulations and the perturbatively motivated result for the scaling of $K(t)$ (Kurkela and Moore, 2012) are rather close, corroborating $\sqrt{\sigma} \sim K$.

The positive value for ζ signals evolution towards larger length scales, with a growing characteristic area $A(t) \sim t^\zeta$. For large A/t^ζ one observes from Fig. 15 the generalized area-law behavior (Berges *et al.*, 2019, 2017a)

$$-\log\langle W \rangle \sim A/t^\zeta. \quad (75)$$

This implies a time-dependent string tension scale $\sigma(t) = -\partial \log\langle W \rangle / \partial A \sim t^{-\zeta}$.

In Ref. (Mace *et al.*, 2016), this behavior is related to the rate of topological transitions, the so-called sphaleron transition rate:

$$\Gamma_{\text{sphaleron}} = C \sigma^2, \quad (76)$$

where C is a number of order unity. The picture that emerges is that the rate of topological transitions is large at early times, $\Gamma_{\text{sphaleron}} \sim Q_S^4$, but subsequently decreases with time at a rate dictated by the universal scaling exponent ζ . One expects this rate to converge from above to the thermal rate for sphaleron transitions in a high-temperature plasma (Moore and Tassler, 2011). We will return to the implications of these results for the evolution of anomalous currents in Sec. IV.E.

Fig. 16 summarizes the behavior of the different characteristic scales in the self-similar regime far from equilibrium. Apart from the perturbative behavior of the hard scale, classical-statistical lattice simulations results are given for the Debye and the non-perturbative string tension scale (Mace *et al.*, 2016). The result clearly demonstrate the dynamical separation of scales as a function of time.

2. Effective condensate dynamics

The traced Wilson loop Eq. (74) may be directly related to correlation functions of a gauge-invariant scalar field (Ford *et al.*, 1998; Gasenzer *et al.*, 2014; Mitreuter *et al.*, 1998). In thermal equilibrium, this scalar field serves as an order parameter for the confinement-deconfinement phase transition of the underlying gauge theory (Braun *et al.*, 2010; Fister and Pawłowski, 2013). In the self-similar scaling regime of the non-thermal attractor, the dynamical evolution of the scalar order-parameter field modes towards the infrared bears many similarities (Berges *et al.*, 2019) with the dynamics of Bose condensation in non-relativistic field theories far from equilibrium (Berges and Sexty, 2012; Chantesana *et al.*, 2019; Piñeiro Orioli *et al.*, 2015). Even quantitatively, the values for the infrared scaling exponents in the different theories agree well within errors (Berges *et al.*, 2019).

The non-equilibrium infrared dynamics for scalars starting from over-occupation has been studied in great detail (Berges *et al.*, 2015b, 2008a; Berges and Sexty, 2011, 2012; Boguslavski and Piñeiro Orioli, 2019; Chantesana *et al.*, 2019; Deng *et al.*, 2018; Moore, 2016; Nowak *et al.*, 2012, 2011; Piñeiro Orioli and Berges, 2019; Piñeiro Orioli *et al.*, 2015; Scheppach *et al.*, 2010; Shen and Berges, 2020; Walz *et al.*, 2018). The emergence of self-similar scaling behavior is closely related to the existence of non-thermal fixed points (Berges and Hoffmeister, 2009; Berges and Mesterhazy, 2012; Berges *et al.*, 2008a; Corell *et al.*, 2019). For scalar N -component theories, the behavior can be approximately described by a large- N effective

kinetic theory to next-to-leading order, which describes the perturbative higher momentum regime as well as the non-perturbative infrared dynamics (Walz *et al.*, 2018).

Both relativistic and non-relativistic scalar theories can show the same infrared scaling and condensation properties (Piñeiro Orioli *et al.*, 2015). This is true even for the anisotropic dynamics of relativistic scalars with longitudinal expansion along the z -direction; the latter geometry is relevant in the context of heavy-ion collisions, and both scalars and gauge theories show also a very similar behavior for higher momenta in this case (Berges *et al.*, 2015a). Because of the strong enhancement in the overoccupied infrared regime, the low momentum modes exhibit essentially isotropic properties despite longitudinal expansion.

A remarkable development in this regard is that table-top experiments with ultracold quantum gases have discovered universal transport processes towards the infrared starting from initial overoccupation of bosonic excitations of trapped atoms (Erne *et al.*, 2018b; Prüfer *et al.*, 2018), similar to the case discussed here. This is discussed further in Sec. VIII.

E. Early-time fermion production and quantum anomalies

In the high energy limit, strong gauge fields dominate the earliest stages of the plasma's space-time evolution. However, the Bose enhancement from over-occupied gluons can lead to a rapid production of quarks with important phenomenological consequences for heavy-ion collisions, such as direct photon production from the electrically charged quarks (Chatterjee *et al.*, 2010) or the breaking of classical symmetries due to anomalies, a prominent example being the chiral magnetic effect (Kharzeev *et al.*, 2016; Koch *et al.*, 2017). At early times these processes occur far from equilibrium and require suitable techniques for their computation. We will discuss here these techniques and their consequences for the production and evolution of fermions off-equilibrium.

1. Real-time simulations for fermions and gauge fields beyond the classical-statistical approximation

Since identical fermions cannot occupy the same state, their quantum nature is in general highly relevant and a consistent quantum treatment of their dynamics is of crucial importance. In the QCD Lagrangian in Eq. (1), quarks appear as bilinear fields. Their real-time

quantum dynamics may therefore be computed by numerically solving the operator Dirac equation coupled to the gluon fields.

This can be achieved in an approximation where the gauge fields are treated using classical-statistical field theory and by employing a mode function analysis of the operator Dirac equation for quarks with available lattice simulation techniques (Aarts and Smit, 1999; Berges *et al.*, 2011; Borsanyi and Hindmarsh, 2009; Kasper *et al.*, 2014; Saffin and Tranberg, 2011). For strong gauge fields $\sim 1/g$, this approximate description amounts to a systematic expansion of the quantum dynamics in $\alpha_S \equiv g^2/(4\pi)$, where the leading order includes the full (non-perturbative) classical-statistical theory of gluons, and the next-to-leading order takes into account back-action of the quarks onto the gluons, which is controlled by $\sim \alpha_S N_f$ for N_f quark flavors.

This can be also directly understood from the path integral formulation of the quantum theory as described in detail in Ref. (Kasper *et al.*, 2014) for Abelian and non-Abelian gauge theories with fermions on a lattice. Performing the Gaussian integration for the quark fields in QCD analytically yields a path integral for the gauge fields A^\pm on the forward (+) and backward (−) part of the closed time contour (see Sec. III) with an effective action

$$S_{\text{eff}}[A^+, A^-] = \text{Tr} \log \Delta^{-1}[A^+, A^-] + iS_{\text{YM}}[A^+, A^-]. \quad (77)$$

The term $\text{Tr} \log \Delta^{-1}[A^+, A^-]$ arises from the Gaussian integral over the quarks, where $i\Delta^{-1}[A^+, A^-]$ denotes the inverse fermion propagator in the presence of the gauge fields. Here $S_{\text{YM}}[A^+, A^-]$ is the Yang-Mills action of the pure gauge theory evaluated on the upper and lower branch of the closed time contour.

The power counting for strong gauge fields is most efficiently done by a rotation of the \pm basis for the gauge fields, splitting the gauge fields into a “classical” part \bar{A} and a “quantum” one \tilde{A} , according to

$$A^+ = \frac{1}{g}\bar{A} + \frac{g}{2}\tilde{A} \quad , \quad A^- = \frac{1}{g}\bar{A} - \frac{g}{2}\tilde{A}. \quad (78)$$

Expressed thus in terms of \bar{A} and \tilde{A} , the interaction terms of S_{YM} can be similarly decomposed into classical and quantum parts.

This is illustrated in Fig. 17, which indicates the classical three-vertex $\sim \bar{A}^2\tilde{A}$ and four-vertex $\sim \bar{A}^3\tilde{A}$ parts of S_{YM} , which are linear in the quantum field \tilde{A} . Fig. 18 gives the corresponding quantum three-vertex $\sim g^4\tilde{A}^3$ and four-vertex $\sim g^4\bar{A}\tilde{A}^3$ parts of S_{YM} , which

are cubic in the quantum field \tilde{A} and suppressed by two powers of α_S compared to their classical counterparts.

A similar analysis can be done for the $\text{Tr} \log \Delta^{-1}[\bar{A}, \tilde{A}]$ contribution coming from the quark fluctuations. Expanding this contribution in powers of the quantum field \tilde{A} yields (Kasper *et al.*, 2014)

$$\text{Tr} \log \Delta^{-1}[\bar{A}, \tilde{A}] \sim g^2 \text{Tr} \left(j_q[\bar{A}] \tilde{A} \right) + g^4 \mathcal{O}(\tilde{A}^3). \quad (79)$$

The linear term in \tilde{A} is proportional to the quark vector-current in the presence of the classical gauge field, $j_q[\bar{A}]$ (Aarts and Smit, 1999; Berges *et al.*, 2011; Borsanyi and Hindmarsh, 2009; Kasper *et al.*, 2014; Saffin and Tranberg, 2011).

Correspondingly, in this formulation the limit $g = 0$ represents the classical-statistical field theory limit of pure Yang-Mills theory. In fact, the rescalings with the gauge coupling employed in Eq. (78) reflect the fact that for classical-statistical field theory the coupling can always be scaled out by suitable field re-definitions, while this not possible in the presence of quantum corrections. Since fermions are genuinely quantum, one cannot scale out the coupling from their contributions, as seen in Eq. (79) which starts at order α_S .

According to the above analysis, genuine quantum corrections to the dynamics in pure Yang-Mills theory enter only at order α_S^2 . Both the classical-statistical field contribution for the Yang-Mills part, and the lowest contribution from quark fluctuations to S_{eff} , are linear in \tilde{A} . Neglecting higher-order corrections coming from terms with higher powers of \tilde{A} , the stationarity condition $\delta S_{\text{eff}}[\bar{A}, \tilde{A}]/\delta \tilde{A} = 0$ yields the classical Yang-Mills evolution equation for \bar{A} with the quark current as a source term. This can be efficiently implemented numerically with sampling techniques using the Wilson plaquette formulation on a lattice (Aarts and Smit, 1999; Berges *et al.*, 2011; Borsanyi and Hindmarsh, 2009; Kasper *et al.*, 2014; Saffin and Tranberg, 2011).

Numerical solutions of the non-equilibrium time evolution of gluons with dynamical quarks have been obtained in Ref. (Gelis *et al.*, 2006a) from 2+1 dimensional boost invariant simulations, in Ref. (Gelfand *et al.*, 2016) in 3+1 space-time dimensions for a non-expanding system, and in Ref. (Tanji and Berges, 2018) for the realistic case with longitudinal expansion. The calculations provide important first principles results on early quark production and the approach towards chemical equilibrium. The results for the gluon sector are in line with earlier simulations without quarks as expected at weak couplings, including self-similar scaling characteristic of the first stage of the bottom-up thermalization scenario (Baier *et al.*,

2001; Berges *et al.*, 2014b). Several properties of the quark number distributions are carried over from the gluon distributions, such as longitudinal momentum broadening (Tanji and Berges, 2018; Tanji and Venugopalan, 2017).

We also note recent work on the real-time propagation of heavy quarks in the Glasma that are important for a first-principles understanding of quarkonium production in heavy-ion collisions (Lehmann and Rothkopf, 2020).

Classical-statistical lattice simulations cannot correctly describe the late-time thermalization dynamics, when typical gluon occupancies become of order unity. The evolution may then be continued with effective kinetic descriptions as reported in Sec. V.E.3.

2. Real-time off-equilibrium dynamics of quantum anomalies

The pair production of quarks and antiquarks lead to macroscopic manifestations of quantum anomalies, corresponding to the breaking of classical symmetries by quantum effects. These may be observable in heavy-ion collisions in the form of a Chiral Magnetic Effect (CME) whereby topological transitions in the very strong electromagnetic B fields at early times generate a vector current in the direction of the B field (Fukushima *et al.*, 2008; Kharzeev *et al.*, 2008). The prospects for the discovery of this and related phenomena are reviewed in Refs. (Kharzeev *et al.*, 2016; Koch *et al.*, 2017).

The key idea is that transitions between different topological sectors of the non-Abelian gauge theory can induce a net axial charge asymmetry j_a^0 of light quarks, which can fluctuate on an event-by-event basis. In off-central heavy-ion collisions, where strong electromagnetic \vec{B} -fields are present, this axial charge asymmetry can be converted into an electric current $\vec{j} \sim j_a^0 \vec{B}$ that is potentially observable. Since the large “magnetar strength” B fields die off very quickly after the collision (Skokov *et al.*, 2009), the CME is most pronounced at the earliest times after the collision.

The non-equilibrium dynamics of topological transitions in a highly occupied, albeit non-expanding, Glasma were studied in Ref. (Mace *et al.*, 2016) by performing classical-statistical simulations and employing a cooling technique to isolate infrared dominated topological transitions. Since gluon saturation generates a large scale $Q_S \gg \Lambda_{\text{QCD}}$, so-called sphaleron transitions generate real-time transitions between configurations characterized by integer valued topological charge that may be separated by an energy barrier.

Interestingly, the boost invariant Glasma configurations discussed in Sec. III.C.2 do not correspond to integer valued configurations of topological charge (Kharzeev *et al.*, 2002); sphaleron transitions therefore go hand in hand with the explosive growth of plasma instabilities that break boost invariance, a phenomenon dubbed as “exploding sphalerons” (Shuryak and Zahed, 2003). As noted in Eq. (76), the sphaleron transition rate is controlled by the spatial string tension in the Glasma.

While off-equilibrium topological transitions are an essential ingredient, the CME in heavy-ion collisions is mediated by the transport of quarks in this topological background and in the presence of external B fields. To address this problem of anomaly transport in such backgrounds, real-time lattice simulations have been performed with dynamical fermions for 3+1 dimensional Abelian and non-Abelian gauge theories in Refs. (Mace *et al.*, 2017; Müller *et al.*, 2016) for given background gauge fields. Transient anomalous charge production in strong-field QCD has also been studied in Refs. (Tanji, 2018; Tanji *et al.*, 2016).

Anomalies have been investigated for Abelian theories off-equilibrium for the fully dynamical situation, including the back-reaction of the fermions onto the gauge fields, in one (Kharzeev and Kikuchi, 2020; Zache *et al.*, 2019), two (Ott *et al.*, 2019) and three (Mace *et al.*, 2019; Mueller *et al.*, 2016) spatial dimensions. In Refs. (Kharzeev and Kikuchi, 2020; Zache *et al.*, 2019) dynamical topological transitions in the massive Schwinger Model with a θ -term, as a prototype model for CP-violation, are studied. A dynamical order parameter for quantum phase transitions between different topological sectors is established, which can be accessed through fermion two-point correlators. Using exact diagonalization techniques, it is shown that the topological transitions persist beyond the weak-coupling regime (Zache *et al.*, 2019).

Quantum fluctuations lead to an anomalous violation of parity symmetry in quantum electrodynamics for an even number of spatial dimensions, which is studied in Ref. (Ott *et al.*, 2019) using the lattice simulation techniques described above. While the leading parity-odd electric current vanishes in vacuum, a non-cancellation of the anomaly for strong electric fields off-equilibrium is observed with distinct macroscopic signatures.

The non-linear dynamics of the CME in QED has been computed in Ref. (Mueller *et al.*, 2016) using real-time lattice simulations. For field strengths exceeding the Schwinger limit for pair production, one encounters a highly absorptive medium with anomaly-induced dynamical refractive properties. An intriguing tracking behavior is found, where the system

spends longest times near collinear field configurations with maximum anomalous current.

An interesting phenomenon observed in such simulations of off-equilibrium QED plasmas is that of chiral instabilities proceeding through the primary and secondary instabilities we discussed previously culminating in a self-similar turbulent magnetic helicity transfer to macroscopic length scales (Mace *et al.*, 2019), see also Ref. (Buividovich and Ulybyshev, 2016).

V. EQUILIBRATION IN QCD KINETIC THEORY

A. The quasi-particle description of QCD plasmas

In order to solve the quantum equations of motion, Eq. (66), for the late time evolution towards local thermal equilibrium, an effective description with a well defined range of validity at some (long) time and distance scales is needed. A well known example is kinetic theory, which describes the state of the system in terms of phase space distributions of particles. Such effective kinetic description of the plasma may be obtained from n -particle irreducible quantum effective action techniques following along the lines of Refs. (Berges, 2004b; Blaizot and Iancu, 2002; Carrington and Kovalchuk, 2009).

The derivation of kinetic theory from the underlying quantum field theory involves a series of approximations. First, for the notion of particles with a well defined position and momentum between collisions to be valid, the de Broglie wavelength of the (quasi-)particles must be small compared to the mean free path between collisions. Likewise, quantum interference effects between successive scattering events should not spoil a description in terms of independent scatterings. For the weakly coupled QCD plasma at high temperature these questions have been addressed in a series of works culminating into the kinetic theory formulation by Arnold, Moore and Yaffe (Arnold *et al.*, 2003b).

The phase space distribution functions employed in kinetic descriptions are derived from two-point correlation functions of the underlying quantum field theory. In local thermal equilibrium, the system is locally homogeneous and time independent. Therefore all two-point functions can only depend on the relative coordinate $s^\mu = x^\mu - y^\mu$. For slow variations in space and time of the central coordinates $X^\mu = (x^\mu + y^\mu)/2$, one considers the evolution in X as given by a gradient expansion of Eq. (66) for the spectral function ρ and the statistical

function F . To the lowest order in gradients, the evolution equation for ρ is not dynamical, and a quasi-particle description emerges from an on-shell spectral function ρ in the weak coupling limit (Berges, 2004b).

Here we consider the temperature T of the QCD plasma to be the single dominant energy scale in the problem. Already at leading order in the coupling, the self-energy receives contributions from an infinite number of perturbative loop diagrams with hard $\mathcal{O}(T)$ internal momentum—Hard Thermal Loops (HTL) (Braaten and Pisarski, 1990). This results in quasi-particles acquiring a screening mass $m \sim gT$.

The equation of motion for the statistical function is solved by generalizing the Kubo-Martin-Schwinger (KMS) relation to introduce a non-equilibrium distribution function $f(X, p)$

$$F(X, p) = -i \left(\frac{1}{2} \pm f(X, p) \right) \rho(X, p), \quad (80)$$

where “+” is for bosons and “−” for fermions, and the quasi-particle momentum p^μ is the Fourier conjugate to the relative coordinate s^μ . In general, there can be separate distributions for different color, spin and polarisation components of the two-point correlation function.

From the equation of motion for the statistical function one obtains the kinetic Boltzmann equation for the distribution function, which is written as³⁶

$$p^\mu \partial_\mu f(X, p) = -C[f]. \quad (81)$$

The leading order collision term $C[f]$ is obtained by a systematic power counting in the coupling constant; this computation is non-trivial and various diagrammatic approaches have been employed to derive the relevant collision processes. For a systematic derivation of kinetic theory from the underlying field theory see (Calzetta *et al.*, 2000; Jeon, 1995; Jeon and Yaffe, 1996) for the scalar case and (Aarts and Martinez Resco, 2005; Gagnon and Jeon, 2007) for Abelian field theories.

For non-Abelian gauge theories at high temperatures, the leading order collision kernel appears at g^4 order. However in addition to elastic scattering processes, there are collinear splitting processes, which contribute at the same order. The importance of the latter was recognized only later (Arnold *et al.*, 2001; Aurenche *et al.*, 1998). The corresponding vertex

³⁶ Keeping interactions with strong background gauge fields leads to more general equations (Mrówczyński *et al.*, 2017).

corrections for the underlying quantum field theory can be formulated using higher n PI effective actions (Berges, 2004b; Carrington and Kovalchuk, 2009).

Once relevant physics processes are accounted for at the given order, Eq. (81) describes the non-equilibrium evolution of QCD plasmas with the coupling constant g as the only free parameter at high temperature (with the possible exception of heavy quark masses). In particular, one can use linearized kinetic theory to compute various transport properties of the plasma around thermal equilibrium: shear and bulk viscosities, conductivity, diffusion and higher order transport coefficients (Arnold *et al.*, 2006, 2003c; York and Moore, 2009). For a recent comprehensive review on perturbative thermal QCD techniques in kinetic theory, and beyond, see Ref. (Ghiglieri *et al.*, 2020). As we will discuss in detail below, the QCD kinetic theory also provides a phenomenologically successful picture of QCD thermalization in heavy-ion collisions (Abraao York *et al.*, 2014; Keegan *et al.*, 2016a; Kurkela and Lu, 2014; Kurkela and Zhu, 2015). For a complementary review, see (Schlichting and Teaney, 2019).

1. Chiral kinetic theory

In the following sections, we will discuss in detail the equilibration of QCD in the framework of spin and color averaged kinetic theory. Spin and color dependent kinetic descriptions require extensions of phase space distributions (Berezin and Marinov, 1977; Mueller and Venugopalan, 2019). Such theories must include a relativistic covariant description of Berry curvature and of the dynamics of the chiral anomaly for spinning and colored particles in external background fields (Chen *et al.*, 2015b; Son and Yamamoto, 2013; Stephanov and Yin, 2012).

Recent work in this direction include a Wigner function approach (Gao and Liang, 2019; Hattori *et al.*, 2019; Sheng *et al.*, 2020; Weickgenannt *et al.*, 2019; Yang *et al.*, 2020), chiral effective field theory (Carignano *et al.*, 2019) and a worldline formalism (Mueller and Venugopalan, 2017). An important question to resolve in this context is the relation of the dynamics of Berry's phase to that of the chiral anomaly (Fujikawa, 2006; Mueller and Venugopalan, 2018; Yee and Yi, 2020). A common goal of these approaches is a consistent framework to describe anomalous transport in QCD which can be matched to an anomalous relativistic hydrodynamic description at late times (Inghirami *et al.*, 2020). These studies

have strong interdisciplinary connections to chiral transport across energy scales ranging from Weyl and Dirac semi-metals to astrophysical phenomena (Landsteiner, 2016).

B. Leading order kinetic theory

We will briefly recap here the main ingredients of QCD effective kinetic theory at leading order in the coupling constant (Arnold *et al.*, 2003b). We will consider the time evolution of the color and spin/polarization averaged distribution function f_s with effective $2 \leftrightarrow 2$ scatterings and $1 \leftrightarrow 2$ collinear radiation terms. For a transversely homogeneous and boost invariant system (applicable at early times in central heavy-ion collisions), the phase space distribution $f^s(\tau, \mathbf{p}) \equiv f_{\mathbf{p}}^s$ is a function only of Bjorken time $\tau = \sqrt{t^2 - z^2}$ and momentum. The resulting Boltzmann equation is

$$\left(\partial_\tau - \frac{p^z}{\tau} \frac{\partial}{\partial p^z} \right) f_{\mathbf{p}}^s = -C_{2 \leftrightarrow 2}^s[f](\mathbf{p}) - C_{1 \leftrightarrow 2}^s[f](\mathbf{p}) \quad (82)$$

with the massless³⁷ dispersion relation, $p^0 = |\mathbf{p}| = p$. Consequently, this kinetic theory describes a conformal system with temperature T as the only dimensionful scale. The index s refers to different particle species in the theory, for example, quarks and gluons in $SU(N_c)$ gauge theory with N_f fermion flavors. The second term on the left hand side is due to the longitudinal gradients in a boost invariant expansion (Baym, 1984). The expansion redshifts the distribution in the p^z direction making it more anisotropic along the longitudinal direction. Different stages of the thermalization process are defined by the competition between the expansion which drives the system away from equilibrium and the collision terms that isotropize and equilibrate the system.

1. Elastic two-body scattering

The $2 \leftrightarrow 2$ collision term for particle species $s = a$ is

$$\begin{aligned} C_{2 \leftrightarrow 2}^a[f](\mathbf{p}) &= \frac{1}{4p\nu_a} \sum_{bcd} \int \frac{d^3\mathbf{k}d^3\mathbf{p}'d^3\mathbf{k}'}{(2\pi)^9 2k 2p' 2k'} \\ &\times \{ f_{\mathbf{p}}^a f_{\mathbf{k}}^b (1 \pm f_{\mathbf{p}'}^c) (1 \pm f_{\mathbf{k}'}^d) - f_{\mathbf{p}'}^c f_{\mathbf{k}'}^d (1 \pm f_{\mathbf{p}}^a) (1 \pm f_{\mathbf{k}}^b) \} \\ &\times |\mathcal{M}_{cd}^{ab}|^2 (2\pi)^4 \delta^{(4)}(p^\mu + k^\mu - p'^\mu - k'^\mu), \end{aligned} \quad (83)$$

³⁷ At leading order, we can neglect the thermal mass correction $m_s \sim gT$ to the dispersion relation $p = \sqrt{|\mathbf{p}|^2 + m_s^2}$ for hard momenta $|\mathbf{p}| \sim T$ on external legs.

where \sum_{bcd} is the sum over all particle and antiparticle species. The second line represents the phase space loss and gain terms. $|\mathcal{M}_{cd}^{ab}|^2$ is the $2 \leftrightarrow 2$ scattering amplitude squared and summed over spin and color degrees of freedom of the external legs, with $\nu_g = 2(N_c^2 - 1)$ for gluons and $\nu_q = 2N_c$ for quarks.

The scattering matrix element $|\mathcal{M}_{cd}^{ab}|^2$ in Eq. (83) should be calculated using in-medium corrected propagators and vertices from the HTL effective Lagrangian (Ghiglieri *et al.*, 2020). At leading order in the coupling constant, and for hard $p \sim T$ external legs, the scattering matrix element coincides with the tree level vacuum matrix element; for instance, in the case of two gluon scattering,

$$|\mathcal{M}_{gg}^{gg}|^2 = 8\nu_g N_c^2 g^4 \left(3 - \frac{st}{u^2} - \frac{su}{t^2} - \frac{tu}{s^2} \right), \quad (84)$$

where s, t and u are the Mandelstam variables. In-medium corrections become relevant when $-t, -u \sim (gT)^2$ is small, but s is large, as is the case for the small angle scattering of hard particles. When the exchanged gluon (or quark) is soft, so that $q = |\mathbf{p}' - \mathbf{p}| \ll T$ in the t -channel (and similarly in the u -channel), the vacuum collision matrix suffers from a soft Coulomb divergence $|\mathcal{M}|^2 \propto 1/(q^2)^2$. Therefore the problematic scattering matrix elements in this region need to be reevaluated using the non-equilibrium propagators for internal lines, which regulate the divergence (Arnold *et al.*, 2003b).

For isotropic distributions and hard $p \sim T$ external legs, the soft self-energy (which cuts off the long range Coulomb interactions) is proportional to the in-medium effective masses of hard gluons and quarks (Arnold *et al.*, 2003b). For gluons, it is given by (assuming $f_{\mathbf{p}}^q = f_{\mathbf{p}}^{\bar{q}}$)

$$m_g^2 = 2g^2 \int \frac{d^3\mathbf{p}}{(2\pi)^3 p} [N_c f_{\mathbf{p}}^g + N_f f_{\mathbf{p}}^q], \quad (85)$$

However for anisotropic distributions, the HTL resummed gluon propagator³⁸ develops poles at imaginary frequency indicating the presence of a soft gauge instability (Mrówczyński, 1997; Mrówczyński *et al.*, 2017). Formally, this restricts the applicability of kinetic theory to parametrically small anisotropies (Arnold *et al.*, 2003b).

The rich physics of plasma instabilities has been studied extensively (Mrówczyński *et al.*, 2017). While such instabilities are of fundamental importance at early times, remarkably, classical-statistical simulations of the non-equilibrium field dynamics of the Glasma (discussed in Sec. IV.C) show that such instabilities do not play a significant role *at late times*

³⁸ Note that there are no unstable fermionic modes in anisotropic plasmas (Mrówczyński, 2002; Schenke and Strickland, 2006).

in expanding 3+1 dimensional non-Abelian plasmas. Motivated by these findings, phenomenological approaches in kinetic theory simulations for anisotropic distributions use an isotropic screening prescription (Abraao York *et al.*, 2014; Kurkela and Zhu, 2015).

2. Fokker-Planck limit of elastic scatterings

For isotropic distributions, the elastic collision kernel for soft momentum exchange can be rewritten as a drag and diffusion process in momentum space (Blaizot *et al.*, 2014; Ghiglieri *et al.*, 2016, 2018a; Ghiglieri and Teaney, 2015; Hong and Teaney, 2010; Moore and Teaney, 2005; Schlichting and Teaney, 2019). First, one needs to separate the full collision kernel into a diffusion term for soft momentum transfers $q < \mu$ and large-angle scatterings $q > \mu$, where the cutoff scale μ satisfies $gT \ll \mu \ll T$:

$$C_{2\leftrightarrow 2}^g[f](\mathbf{p}) = C_{\text{diff}}^g[f](\mu)|_{q < \mu} + C_{2\leftrightarrow 2}^g[f](\mathbf{p})|_{q > \mu}. \quad (86)$$

The physics of the diffusion term is that of hard particles being kicked around by the fluctuating soft gauge fields generated by other particles. For an isotropic non-equilibrium plasma, the expectation value of such gauge field fluctuations can be related to equilibrium fluctuations with the help of an effective temperature T_* (taking $f_{\mathbf{p}}^g = f_{\mathbf{p}}^{\bar{g}}$)

$$T_* \equiv \frac{g^2}{m_g^2} \int \frac{d^3\mathbf{p}}{(2\pi)^3} [N_c f_{\mathbf{p}}^g (1 + f_{\mathbf{p}}^g) + N_f f_{\mathbf{p}}^q (1 - f_{\mathbf{p}}^q)]. \quad (87)$$

Note that although $T_* = T$ in equilibrium, T_* is distinct from the effective temperature defined by the energy density and used in Sec. V.E and Sec. VI. Evaluating the collision kernel in the limit of soft momentum transfer and isotropic distributions, results in a Fokker-Planck type collision term

$$C_{\text{diff}}^g[f](\mu) = \eta_D(p) \hat{p}^i \frac{\partial}{\partial p^i} (f_{\mathbf{p}}^g (1 + f_{\mathbf{p}}^g)) + \frac{1}{2} q^{ij} \frac{\partial^2 f_{\mathbf{p}}^g}{\partial p^i \partial p^j}, \quad (88)$$

where η_D is the drag coefficient, $q^{ij} = \hat{q}_L \hat{p}^i \hat{p}^j + \frac{1}{2} \hat{q} (\delta^{ij} - \hat{p}^i \hat{p}^j)$ is the diffusion tensor and $\hat{p}^i = p^i/p$ is the unit vector.

The transport coefficients \hat{q} and \hat{q}_L can be evaluated using the resummed HTL propagators, while η_D is constrained by the Einstein relation and the requirement that Eq. (88) vanish in equilibrium (Arnold, 2000a,b; Ghiglieri *et al.*, 2016; Moore and Teaney, 2005). The leading

order result for \hat{q} is

$$\hat{q}(\mu) = \frac{g^2 N_c T_* m_g^2}{2\pi} \log \frac{\mu^2}{2m_g^2}. \quad (89)$$

The UV divergence in the diffusion term is canceled by the corresponding IR divergence in the large-angle scattering term in Eq. (86).

We can now specify the isotropic screening prescription for regulating the elastic collision kernel for anisotropic distributions: for a soft gluon exchange in the t -channel (similarly for the u -channel), the divergent term is replaced by IR regulated term $t \rightarrow t(q^2 + \xi_g^2 m_g^2)/q^2$, where $\xi_g = e^{5/6}/2$ is a numerical constant fixed such that the new matrix element reproduces the full HTL result for the drag and momentum diffusion properties of soft gluon scattering (Abraao York *et al.*, 2014).

Similarly, one can regulate divergent soft fermion exchanges to reproduce gluon to quark conversion $gg \rightarrow q\bar{q}$ at leading order for isotropic distributions (Ghiglieri *et al.*, 2016; Kurkela and Mazeliauskas, 2019b). Formally, this regulated collision kernel is accurate for small couplings and for near-isotropic systems. However in practice, numerical simulations for phenomenological applications are often performed for stronger couplings $g \approx 1$ and anisotropic systems.

3. Effective collinear one-to-two processes

In addition to the momentum diffusion of hard particles, soft gluon exchange can also take the particle slightly off shell and make it kinematically possible for it to split into two nearly collinear hard particles. Naively, such a $2 \rightarrow 3$ process has an additional vertex relative to elastic $2 \leftrightarrow 2$ scattering making it subleading in the coupling constant. However both the soft gluon exchange and the perturbed off-shell hard particle have $\sim 1/(g^2 T^2)$ enhancements from the propagators. These compensate for the additional vertex insertion and the nearly-collinear emission phase space (Arnold *et al.*, 2001). For the same reason, multiple soft scatterings $N + 1 \rightarrow N + 2$ also have to be summed over.

Physically, this means that the nearly on-shell hard particle lives long enough before splitting to receive multiple kicks from the plasma, which destructively interfere, leading to the suppression of emissions from very energetic particles. This phenomenon is known as the Landau-Pomeranchuk-Migdal (LPM) effect (Landau and Pomeranchuk, 1953a,b; Migdal,

1955, 1956). Collectively these processes are described as an effective $1 \leftrightarrow 2$ matrix element. In Eq. (82) it is denoted by $C^{1 \leftrightarrow 2}[f](\mathbf{p})$ and has the explicit form,

$$\begin{aligned}
C_{1 \leftrightarrow 2}^a[f](\mathbf{p}) &= \frac{(2\pi)^3}{2\nu_a p^2} \sum_{bc} \int_0^\infty dp' dk' \left[\right. \\
&\quad \gamma_{bc}^a(p; p', k') \delta(p - p' - k') \\
&\quad \times \left\{ f_{p\hat{\mathbf{p}}}^a [1 \pm f_{p'\hat{\mathbf{p}}}^b] [1 \pm f_{k'\hat{\mathbf{p}}}^c] - f_{p'\hat{\mathbf{p}}}^b f_{k'\hat{\mathbf{p}}}^c [1 \pm f_{p\hat{\mathbf{p}}}^a] \right\} \\
&\quad - 2\gamma_{ac}^b(p'; p, k') \delta(p' - p - k') \\
&\quad \left. \times \left\{ f_{p'\hat{\mathbf{p}}}^b [1 \pm f_{p\hat{\mathbf{p}}}^a] [1 \pm f_{k'\hat{\mathbf{p}}}^c] - f_{p\hat{\mathbf{p}}}^a f_{k'\hat{\mathbf{p}}}^c [1 \pm f_{p'\hat{\mathbf{p}}}^b] \right\} \right], \tag{90}
\end{aligned}$$

where the unit vector $\hat{\mathbf{p}} = \mathbf{p}/|\mathbf{p}|$ defines the splitting direction and $\gamma_{bc}^a(p; p', k')$ is the effective collinear splitting rate.

As required by detailed balance, Eq. (90) describes both particle splitting $p \leftrightarrow p' + k'$ and fusion $p + k' \leftrightarrow p'$ processes. Factoring out the kinematic splitting function $P_{g \rightarrow g} \left(z = \frac{p'}{p} \right) = N_c \frac{1+z^4+(1-z)^4}{z(1-z)}$ for the gluonic process $g \rightarrow gg$, this rate is given by

$$\gamma_{gg}^g(p; p', k') = P_{g \rightarrow g}(z) \frac{\nu_g g^2}{4\pi} \int \frac{d^2 h}{(2\pi)^2} \frac{2\mathbf{h} \cdot \text{Re} \mathbf{F}_g(\mathbf{h}; p, p', k')}{4(2\pi)^3 p p'^2 k'^2}, \tag{91}$$

where the integral has mass dimension two and is proportional to the virtuality acquired by the hard particle due to interactions with the soft gauge field. The complex two-dimensional function $\mathbf{F}_g(\mathbf{h}; p, p', k')$ (with mass dimension one) solves the integral equation (Arnold *et al.*, 2001, 2002, 2003b),

$$\begin{aligned}
2\mathbf{h} &= i\delta E(\mathbf{h}) \mathbf{F}_g(\mathbf{h}) + g^2 \frac{N_c T_*}{2} \int \frac{d^2 \mathbf{q}_\perp}{(2\pi)^2} \mathcal{A}(\mathbf{q}_\perp) \\
&\quad \{ 3\mathbf{F}_g(\mathbf{h}) - \mathbf{F}_g(\mathbf{h} - k' \mathbf{q}_\perp) - \mathbf{F}_g(\mathbf{h} - p' \mathbf{q}_\perp) - \mathbf{F}_g(\mathbf{h} + p \mathbf{q}_\perp) \}, \tag{92}
\end{aligned}$$

where the energy difference between the incoming and the outgoing states is

$$\delta E(\mathbf{h}; p, p', k') \equiv \frac{m_g^2}{2k'} + \frac{m_g^2}{2p'} - \frac{m_g^2}{2p} + \frac{\mathbf{h}^2}{2pk'p'}, \tag{93}$$

and $\mathbf{h} = (\mathbf{p}' \times \mathbf{k}') \times \hat{\mathbf{p}}$ quantifies the transverse momentum in the near collinear splitting.

The second term on the r.h.s. of Eq. (92) can be interpreted as a linearized collision integral with loss and gain terms describing the probability of a particle to scatter in and out of transverse momentum \mathbf{h}/p . The scattering rate $\mathcal{A}(\mathbf{q}_\perp)$ is proportional to the mean

square fluctuation of soft gauge fields; for isotropic distributions it is given by (Aurenche *et al.*, 2002b),

$$\mathcal{A}(\mathbf{q}_\perp) = \frac{1}{\mathbf{q}_\perp^2} - \frac{1}{\mathbf{q}_\perp^2 + 2m_g^2}. \quad (94)$$

Even with this isotropic approximation, Eq. (92) is highly non-trivial. Various numerical schemes have been proposed for solving it (Aurenche *et al.*, 2002a; Ghiglieri and Moore, 2014; Ghisoiu and Laine, 2014).

4. Bethe-Heitler and LPM limits of collinear radiation

We will now discuss two limiting cases of the soft gluon radiation $z = \frac{p'}{p} \ll 1$. In the first case, the so-called Bethe-Heitler (BH) limit, the interference between successive scattering events can be neglected. This corresponds to the first (decoherence) term in Eq. (92) being much larger than the scattering integral, $pzg^2T_*/m_g^2 \ll 1$. In this case, the equation can be solved iteratively. One obtains (Ghiglieri *et al.*, 2018a)

$$\gamma_{gg}^g(p; p', k') \Big|_{\text{BH}}^{z \ll 1} = P_{g \rightarrow g}(z) \frac{\nu_g \alpha_S}{(2\pi)^4} \frac{\hat{q}(\mu)p}{m_g^2} \Big|_{\mu=em_g}, \quad (95)$$

where $\hat{q}(\mu)$ is given in Eq. (89). In the opposite limit $zpT_*/m_g^2 \gg 1$ (but still $z \ll 1$), the successive scatterings by the medium interfere destructively reducing the emission rate to

$$\gamma_{gg}^g(p; p', k') \Big|_{\text{LPM}}^{z \ll 1} = P_{g \rightarrow g}(z) \frac{\nu_g \alpha_S}{(2\pi)^4} \left(\frac{\hat{q}(\mu)p}{z} \right)^{1/2}, \quad (96)$$

where, at next-to-leading-logarithmic order, μ solves $\mu^2 = 2\sqrt{2}e^{2-\gamma_E+\pi/4}\sqrt{\hat{q}(\mu)pz}$ (Arnold and Dogan, 2008).

Due to soft gauge field instabilities, collinear radiation in anisotropic plasmas contains unstable modes (Hauksson *et al.*, 2018, 2020). In phenomenological applications these unstable modes are neglected and the isotropic approximation in Eq. (94) is employed instead.

C. Bottom-up thermalization

1. Initial conditions

Baier *et al.* (BMSS) (Baier *et al.*, 2001) spelled out a bottom-up scenario for thermalization beginning from the overoccupied Glasma discussed previously in Sections II, III and

IV. In this framework, the momentum modes $p \sim Q_S$ can be interpreted as quasi-particles with a well-defined anisotropic distribution after time³⁹ $\tau Q_S \geq \log^2 \alpha_S^{-1}$. The initial gluon distribution in kinetic simulations of this scenario is parametrized at $Q_S \tau_0 = 1$ as (Kurkela and Zhu, 2015)

$$f_{\mathbf{p}}^g = \frac{2A}{g^2 N_c} \frac{\langle p_T \rangle_c}{\sqrt{p_{\perp}^2 + p_z^2 \xi_0^2}} e^{-\frac{2}{3} \frac{p_{\perp}^2 + \xi_0^2 p_z^2}{\langle p_T \rangle_c^2}}. \quad (97)$$

The normalization A is chosen to reproduce the co-moving energy density $\tau \mathcal{E} = \langle p_T \rangle_c dN_g / d^2 \mathbf{x}_{\perp} dY$. In this expression the gluon number density at a fixed initial rapidity is determined at LO by numerical simulations of the Glasma and the result can be read off Eq. (48). To obtain the first principles input for the initial gluon production as a function of rapidity, one further needs to solve the JIMWLK equations described previously in Sec. II.D. Likewise, one can determine the average transverse momentum $\langle p_T \rangle_c \approx 1.8 Q_S$ (Lappi, 2011). Finally, the anisotropy parameter ξ_0 is varied to quantify our ignorance of the longitudinal momentum distribution.

For the evolution of the overoccupied and highly anisotropic initial state, specified at its initial time by Eq. (97), the typical gluon occupancy and the deviation from isotropy can be monitored by computing the following ratios,

$$\frac{\langle pf \rangle}{\langle p \rangle} = \frac{\int \frac{d^3 p}{(2\pi)^3} p f_{\mathbf{p}}^g f_{\mathbf{p}}^g}{\int \frac{d^3 p}{(2\pi)^3} p f_{\mathbf{p}}^g}, \quad \frac{\mathcal{P}_T}{\mathcal{P}_L} = \frac{\frac{1}{2} \int \frac{d^3 p}{(2\pi)^3} p_{\perp}^2 f_{\mathbf{p}}^g}{\int \frac{d^3 p}{(2\pi)^3} p_z^2 f_{\mathbf{p}}^g}. \quad (98)$$

2. Stage one: collisional broadening

The solution of the collisionless Boltzmann equation in the boost invariant expansion is a simple rescaling of initial longitudinal momentum which does not change the typical occupancy, but increases the anisotropy quadratically in time. However in the presence of elastic collisions, gluons scatter into the longitudinal momentum direction thus broadening the distribution. The longitudinal momentum diffusion for anisotropic distributions can be estimated from the Fokker-Planck equation Eq. (88)

$$\left(\partial_{\tau} - \frac{p^z}{\tau} \frac{\partial}{\partial p^z} \right) f_{\mathbf{p}}^g = \frac{\hat{q}}{4} \frac{\partial^2 f_{\mathbf{p}}^g}{\partial p_z^2}, \quad (99)$$

³⁹ Plasma instabilities operational over shorter time-scales are well described by classical-statistical simulations – see Sec. IV.B.

where we kept the dominant term on the right hand side. Note that for a highly occupied anisotropic system $\hat{q} \sim \int_{\mathbf{p}} (f_{\mathbf{p}}^g)^2$, this equation admits the scaling solution Eq. (73); as discussed in Sec. IV.C.2, this solution is singled out in the classical-statistical simulations.

The physical picture is that the longitudinal momentum diffuses as $\langle p_z^2 \rangle \sim \hat{q}\tau$, where $\hat{q} \sim \alpha_S^2 n_g^2 / (Q_S^2 \sqrt{\langle p_z^2 \rangle})$ and the hard gluon number density per rapidity is constant $\alpha_S n_g \tau Q_S^{-2} \sim 1$. From this, it follows that the longitudinal momentum decreases as

$$\langle p_z^2 \rangle \sim Q_S^2 (Q_S \tau)^{-2/3}. \quad (100)$$

This clearly shows that the increase in anisotropy is milder than in the free streaming case. One obtains, $\mathcal{P}_T/\mathcal{P}_L \propto (\tau/\tau_0)^{2/3}$ and $\langle pf \rangle/\langle p \rangle \propto (\tau/\tau_0)^{-2/3}$ in agreement with the scaling behavior of the non-thermal attractor of Sec. IV.C.1. The typical occupancy becomes of $\mathcal{O}(1)$ at the time

$$\tau Q_S \geq \alpha_S^{-3/2}. \quad (101)$$

This is the first stage of bottom-up thermalization. As discussed previously, this corresponds to a “quantum breaking” time where the classical-statistical approximation breaks down definitively. After this time, hard gluons with $p_T \sim Q_S$ are no longer overoccupied, although they still carry most of the energy and particle number.

3. Stage two: collinear cascade

Once the typical hard parton occupancy becomes $\mathcal{O}(1)$, the diffusion coefficient scales as $\hat{q} \sim \alpha_S^2 n_g$, where we still have $\alpha_S n_g \tau Q_S^{-2} \sim 1$. At this time, the longitudinal momentum diffusion rate and the expansion rate are comparable, with the result that the longitudinal momentum reaches the constant value

$$\langle p_z^2 \rangle \sim \alpha_S Q_S^2. \quad (102)$$

This ensures that the momentum anisotropy remains constant in the second bottom-up stage.

In this stage, in addition to elastic scatterings, medium induced collinear radiation becomes important, that rapidly increases the population of soft gluons.

The soft gluon multiplicity can be estimated using the Bethe-Heitler formula Eq. (95); integrating over soft momentum $m_D < p < \sqrt{\langle p_z^2 \rangle}$ and neglecting logarithmic factors, one

obtains $n_g^{\text{soft}} \sim \tau \frac{\alpha_S^3}{m_g^2} (n_g^{\text{hard}})^2$. The screening mass is now dominated by soft isotropic gluons, $m_g^2 \sim \alpha_S n_g^{\text{soft}} / \sqrt{\langle p_z^2 \rangle}$. Using the expression above for the longitudinal momentum, we can show that the soft and hard gluon multiplicities are of the same order at times

$$Q_S \tau \geq \alpha_S^{-5/2}. \quad (103)$$

At this time, the soft gluons have thermalized amongst themselves forming a bath with an effective temperature. This marks the end of the second stage of bottom-up thermalization.

4. Stage three: mini-jet quenching

Even though the soft gluons have thermalized, the hard gluons still dominate the energy density. They are however highly diluted $\langle fp \rangle / \langle p \rangle \ll 1$; the non-equilibrium modes are now underoccupied as opposed being overoccupied in the first bottom-up stage. Although soft gluon emission is very efficient in populating the infrared, the successive $z \sim 1/2$ branching of modes is more efficient for energy transfer. Such branching suffers from the LPM suppression. The hard gluons are finally absorbed by the thermal bath in a “mini-jet” quenching that is formally identical to the jet quenching formalism that is typically applied to describe much harder modes.

The system finally thermalizes when the energy in soft and hard components becomes comparable. This happens at the time

$$\tau_{\text{thermal}} = C_1 Q_S^{-1} \alpha_S^{-13/5}, \quad (104)$$

with the thermalization temperature $T = C_2 \alpha_S^{2/5} Q_S$. Here C_1 and C_2 are $\mathcal{O}(1)$ constants (Baier *et al.*, 2011, 2002). This time scale is parametrically $\alpha_S^{-1/10}$ longer than when stage two ends and therefore only cleanly distinguishable at asymptotically small values of the coupling.

The bottom-up thermalization scenario provides an intuitive picture of equilibration at weak coupling. It is remarkable, given the complexity of the thermalization process in QCD, that this scenario allows one to relate the final thermalization time and temperature to the scale for gluon saturation in the nuclear wavefunctions.

Asymptotic freedom tells us that the coupling constant must run with Q_S , which is the relevant hard scale in the problem. Therefore an interesting consequence of Eq. (104) is

that $\tau_{\text{thermal}} \sim \log^{13/5}(Q_S)/Q_S \rightarrow 0$ as $Q_S \rightarrow \infty$. Thus contrary to naive expectations, the bottom-up thermalization scenario predicts that thermalization in the asymptotic Regge limit of QCD will occur nearly instantaneously relative to the size of the system.

5. Numerical realization of bottom-up thermalization

The thermalization time scales in the discussion above were only parametric estimates. We will discuss now the results of a numerical implementation of the bottom-up kinetic evolution from the overoccupied initial phase space distribution in Eq. (97) to the Bose-Einstein distribution (Kurkela and Zhu, 2015).

For 't Hooft coupling $\lambda = N_c g^2 = 1$ and initial anisotropy $\xi_0 = 10$, we show in Fig. 19(a-c) the evolution of the gluon distribution (integrated over spherical angle) with different momentum weights. The three panels correspond respectively to the distribution of the gluon energy density $d\mathcal{E}/dp$, the number density dn/dp and the screening mass dm_g^2/dp as a function of gluon momentum. To factor out the dilution due to expansion, all of these quantities are normalized by the total gluon number density n . The lines correspond to different times $\tau Q_S \approx 1, 10, 10^3$.

We see that at early times $\tau Q_S \approx 1 - 10$ the hard $p > Q_S$ modes dominate both the energy and particle number, and even have significant contributions to the screening mass. At very late times $\tau Q_S \approx 10^3$, the particle number and the screening mass are completely dominated by the soft sector, but there is still a noticeable contribution to the energy density from the modes with $p > Q_S$.

It is interesting to compare the momentum distributions in Fig. 19 with the anisotropy and occupancy evolution in Fig. 20 (which is a kinetic theory extension of the lattice computation in Fig. 12). We mark the times $\tau Q_S \approx 1, 10, 10^3$ with a diamond, a circle and a triangle respectively on the $\lambda = 1, \xi_0 = 10$ simulation trajectory (blue solid line). We observe that typical occupancies drop rather quickly below unity and we see a slight increase of anisotropy as it happens. However the slope of the anisotropy increase is different from the naive expectation in the first stage of bottom-up thermalization and is dependent on the choice of initial conditions.

The anisotropy plateaus of the second stage is reached, already at $\tau Q_S \approx 10$, somewhat quicker than what the parametric estimates would suggest. Finally, because the soft sector

is more isotropic than that of hard gluons, we observe that as the gluon number shifts towards lower momentum (see Fig. 19(b)), the anisotropy starts falling sharply in Fig. 20. This marks the onset of the third stage of bottom-up thermalization. Although dilute hard modes still contribute significantly to the energy density, the balance shifts towards more densely populated soft modes whose occupancy steadily increases as the system isotropizes.

The bottom-up process ends finally when the system isotropizes. In practice, the third stage of bottom-up equilibration is significantly longer than the second stage, in contrast to the $\alpha_s^{-1/10}$ difference in the parametric time scales.

For an initial distribution with different initial anisotropy values $\xi_0 = 4, 10$ (dashed and solid lines) in Fig. 20, the evolution follows a qualitatively similar path. Although we expect all initial conditions to converge at thermal equilibrium, it is remarkable that different initializations merge already at rather large values of the anisotropies $\mathcal{P}_T/\mathcal{P}_L \approx 10$, when the system is still far away from local thermal equilibrium. This precocious collapse to a universal curve, independent of the initial conditions, is termed a “hydrodynamic attractor”. This phenomenon is discussed further in Sec. V.E and Sec. VI.D.

Finally, Fig. 20 also shows kinetic equilibration with increasing coupling constant (and decreasing shear viscosity η/s). For $\lambda \geq 5$, corresponding to small values of $\eta/s \lesssim 2$, (and for which the initial occupancy is already below unity), the system starts to isotropize almost immediately and the distinct stages of the bottom-up scenario are no longer clearly discernible.

D. Self-similar evolution in the high-occupancy regime

1. Self-similar scaling

When characteristic field occupancies are sufficiently large for the classical-statistical approximation to be valid, but small enough for the perturbative kinetic expansion to apply, there is an overlapping regime where both approximations to the dynamics of the system are valid (Aarts and Smit, 1998; Jeon, 2005; Mueller and Son, 2004).

As discussed in Sec. IV.C and Sec. V.C, the non-equilibrium dynamics of the overoccupied plasma undergoes a remarkable simplification in complexity by exhibiting self-similar evolution. In kinetic theory language, the self-similar behavior refers to the situation when

the particle distributions at different times can be related by rescaling the momentum arguments and the overall normalization – see Eq. (73), where α, β, γ denote the universal scaling exponents. The relations between the exponents are constrained by conservation laws and the Boltzmann equation Eq. (81), for which Eq. (73) provides a solution.

Longitudinally expanding systems are anisotropic and subject to soft gauge instabilities. Therefore from a perturbative viewpoint it is very surprising that plasma instabilities do not seem to affect the late time evolution of the classical-statistical real time simulations, as shown in Fig. 12. The self-similar evolution near the non-thermal attractor is consistent with the bottom-up thermalization scenario and numerical QCD kinetic theory simulations (Kurkela and Zhu, 2015), which explicitly neglects plasma instabilities. How to consistently solve the effective kinetic theory in anisotropic plasmas is an important open question (Mrówczyński *et al.*, 2017).

Finally, as mentioned in Sec. IV.C.2, in the case of the non-expanding isotropic systems, the self-similar direct energy cascade plays an important role in equilibration of overoccupied bosons. The same scaling exponents and the scaling function are also reproduced in kinetic theory simulations (Abraao York *et al.*, 2014; Kurkela and Lu, 2014). Fermions are never overoccupied and chemical equilibration takes place over longer timescales than the direct energy cascade (Kurkela and Mazeliauskas, 2019b).

2. Pre-scaling phenomenon

In Ref. (Mazeliauskas and Berges, 2019) it was found that the far-from-equilibrium QGP already exhibits a self-similar behavior before the scaling exponents attain their constant values $\alpha = -2/3$, $\beta = 0$ and $\gamma = 1/3$. The pre-scaling phenomenon is realized through the time dependent rescaling of the distribution function and its arguments (c.f. work by (Micha and Tkachev, 2004)),

$$f_{\mathbf{p}}^g \stackrel{\text{pre-scaling}}{=} \frac{(Q\tau)^{\alpha(\tau)}}{\alpha_S} f_S \left((Q\tau)^{\beta(\tau)} p_{\perp}, (Q\tau)^{\gamma(\tau)} p_z \right), \quad (105)$$

where $\alpha(\tau)$, $\beta(\tau)$ and $\gamma(\tau)$ are generic time dependent functions.

Figure 21 shows the evolution of time dependent scaling exponents in QCD kinetic theory at very small couplings and overoccupied initial conditions (Mazeliauskas and Berges, 2019). The value of the exponents is calculated from the time dependence of various moments of

the distribution,

$$n_{m,n}(\tau) = \int \frac{d^3p}{(2\pi)^3} p_T^m |p_z|^n f_{\mathbf{p}}^g. \quad (106)$$

The different lines of the same color in Fig. 21 correspond to integrals with different powers of the momentum. It is important to note that the rescaling in Eq. (105) is implicitly assumed to be valid in a certain physically relevant momentum range. Therefore a finite set of moments of Eq. (106) contains all the physically relevant information in the distribution. As shown in Fig. 21, different extractions rapidly collapse onto each other and a unique set of scaling exponents emerge that govern the time evolution of all probed moments.

The time dependent scaling exponents provide a more differential picture of how self-similar behavior and information loss emerges near the non-thermal attractor. Here the scaling exponents act as effective degrees of freedom, whose slowly varying evolution constitutes a hydrodynamic description of the system around the non-thermal attractor. In particular, the time dependent exponents could be well suited to study the evolution away from the attractor in equilibrating systems even if the non-thermal attractor is never fully reached, for instance, at larger values of the coupling. For related studies in scalar field theory, see also Ref. (Schmied *et al.*, 2019).

E. Extrapolation to stronger couplings

We have discussed thus far a non-equilibrium QCD evolution scenario which is strictly valid only for $g \ll 1$. However the coupling constant is not parametrically small even at the Z boson mass scale, where $\alpha_S(M_Z^2) \approx 0.1179 \pm 0.0010$ ($g = \sqrt{4\pi\alpha_S} \approx 1.2$) (Tanabashi *et al.*, 2018). In the case of finite temperature perturbation theory, the expansion parameter is $\sim \alpha_S T/m_D \sim g$ – the convergence is therefore very slow (Blaizot *et al.*, 2003a). In this section, we will therefore discuss phenomenological extrapolations of the QCD kinetic theory to “realistic” couplings.

The first calculation at next-to-leading order for QGP transport properties was performed for heavy quark diffusion and the corrections were found to be large (Caron-Huot and Moore, 2008). On the other hand, the NLO contributions to the photon emission nearly cancel and the overall contribution is only $\sim 20\%$ (Ghiglieri *et al.*, 2013). Recently, computations of the shear viscosity, quark diffusion and second order transport coefficients have been extended to include higher order contributions (dubbed “almost NLO” in (Ghiglieri *et al.*, 2018a,b))

thanks to the breakthrough idea of evaluating HTL correlations on the lightcone (Caron-Huot, 2009). From Fig. 22, we see that NLO results for the specific shear viscosity η/s can be a factor five smaller than the leading order result at the accessible QGP temperatures $T \lesssim 1$ GeV. It is conceivable that a better reorganization of the perturbative expansion will result in an improved convergence at NLO (Ghiglieri *et al.*, 2018a).

Nevertheless, for phenomenological applications in heavy-ion collisions, the strong coupling constant value $\alpha_S \approx 0.3$ ($g \approx 2$) is commonly used in leading order calculations. Examples of these include thermal photon emission (Paquet *et al.*, 2016), heavy quark transport (Yao *et al.*, 2020) or parton energy loss (Burke *et al.*, 2014). At this point, it is fair to admit that the leading order kinetic theory applications to equilibration processes in the QGP do not provide a controlled expansion at realistic energies and therefore have large theoretical uncertainties.

On the other hand, QCD kinetic theory does contain the necessary physical processes, such as elastic and inelastic scatterings, to describe QCD thermalization at weak coupling. Therefore in the absence of real time non-perturbative QCD computations, extrapolating the weak coupling results to larger couplings provides a useful baseline, which can be systematically improved upon.

As we will discuss below, the dependence on the coupling constant is better replaced by the value of shear viscosity η/s —a physical property of the QGP. The relaxation to equilibrium is naturally controlled by the strength of dissipative processes. Therefore rescaling weakly coupled kinetic theory dynamics to small values of η/s (favored by hydrodynamic modeling of QGP) can be fairly compared to heavy-ion phenomenology and other microscopic models. This includes the genuinely strongly coupled systems that will be discussed in Sec. VI. This may indicate that lessons learned from QGP equilibration in leading order kinetic theory are more robust than the LO expansion itself.

There have been a number of phenomenological applications of kinetic theory to the study of thermalization in QCD. Early notable examples include Refs. (Biro *et al.*, 1993; Geiger and Muller, 1992; Hwa and Kajantie, 1986). Numerical implementations of classical kinetic theory including elastic $gg \leftrightarrow gg$ and inelastic $gg \leftrightarrow ggg$ gluon scatterings were pioneered in Ref. (El *et al.*, 2008; Xu and Greiner, 2005). Below we will focus on the results from the numerical implementations of the quantum kinetic theory including all leading order processes, which was discussed in Sec. V.B.

1. *Hydrodynamic attractors in QCD kinetic theory*

The universal macroscopic effective theory close to local thermal equilibrium is given by fluid dynamics consisting of the conservation laws and constitutive equations (Landau and Lifshits, 1959)

$$\partial_\mu T^{\mu\nu} = 0, \quad T^{\mu\nu} = T_{\text{hydro}}^{\mu\nu}(\mathcal{E}, u^\mu, \dots). \quad (107)$$

The only surviving information is contained in the macroscopic fluid variables, the energy density \mathcal{E} and fluid velocity u^μ ; all other information about the initial conditions is lost.

The surprising phenomenological success of viscous hydrodynamics in describing many soft hadronic observables in heavy-ion collisions leads one to consider the possibility whether a fluid dynamic description is applicable to systems with significant deviations from local thermal equilibrium. This topic was first investigated in the strongly coupled holographic models, and subsequently in the relaxation time approximation (RTA) kinetic theory and hydrodynamic models; see reviews (Florkowski *et al.*, 2018a; Romatschke and Romatschke, 2019) and Sec. VI.

In the QCD kinetic theory simulations of boost invariant expansion of homogeneous plasmas (Heller *et al.*, 2018; Keegan *et al.*, 2016b; Kurkela *et al.*, 2019a), it was observed that the energy-momentum tensor quickly becomes a sole function of time measured in units of the characteristic kinetic relaxation time⁴⁰ $\tau_R \sim \eta/(sT)$, i.e.,

$$\tilde{w} \equiv \frac{\tau T}{4\pi\eta/s}. \quad (108)$$

In such case the evolution of the energy-momentum tensor can be characterized by $\mathcal{P}_L/\mathcal{E}$ as a function of \tilde{w} (Heller *et al.*, 2012b, 2018). Because \tilde{w}^{-1} is proportional to the Knudsen number—the natural expansion parameter for deviations from equilibrium—one would expect that for large \tilde{w} the kinetic theory will agree with viscous hydrodynamic result $\mathcal{P}_L/\mathcal{E} = \frac{1}{3} - \frac{16}{9} \frac{\eta/s}{\tau T}$. Surprisingly, the simplest viscous constitutive relation is satisfied already for $\tilde{w} \approx 1$, when viscous correction is comparable to the equilibrium pressure. Such effective hydrodynamic description of systems substantially away from equilibrium is called nowadays the hydrodynamic attractor. This notion is in fact much richer and its further aspects are discussed in Sec. VI.D.

⁴⁰ The effective temperature can be defined as a function of the energy density that would play the role of the temperature in equilibrium. In conformal models it is given by the fourth root of the energy density $T = (\mathcal{E}/(\nu_{\text{eff}}\pi^2/30))^{1/4}$. For an ideal gas of quarks and gluons, $\nu_{\text{eff}} = 47.5$ and $\nu_{\text{eff}} = 16$ for gluons only.

Figure 23 shows the pressure anisotropy $\mathcal{P}_L/\mathcal{P}_T$ as a function of rescaled time in an expanding homogeneous system for different values of the coupling constant. The system is prepared in an equilibrium state at initial time and then is allowed to undergo a boost invariant expansion which drives the system away from equilibrium. However as the expansion slows down, it relaxes back to isotropy, satisfying $\mathcal{P}_L/\mathcal{P}_T = 1$.

Note that the kinetic simulations for different couplings (which corresponds to very different kinetic relaxation times) collapse onto each other even when the pressure anisotropy $\mathcal{P}_L/\mathcal{P}_T$ is significant. Overall, the kinetic evolution is very close to that of an infinitely strongly coupled system. Although neither a weakly coupled kinetic theory, nor an infinitely strongly coupled supersymmetric Yang-Mills theory is an exact description of QCD in heavy-ion collisions, Fig. 23 gives some indication that in the rescaled time units \tilde{w} the final stages of QCD equilibration could follow a very similar hydrodynamic attractor curve.

To map the hydrodynamic attractor evolution in dimensionless time \tilde{w} to that in physical units, one needs to fix the interaction strength by setting the shear viscosity over entropy ratio η/s and the dimensionful temperature scale. Extensive hydrodynamic model comparisons to data constrain the shear viscosity to rather small values of $4\pi\eta/s \sim 2$ close to $T_c \approx 155$ MeV, although its value at higher temperatures is not well determined (Bernhard *et al.*, 2019; Devetak *et al.*, 2019). The characteristic temperature scale in the hydrodynamic stage is well constrained by the transverse entropy density per rapidity $(s\tau)_{\text{hydro}} \sim (T^3\tau)_{\text{hydro}}$, which is directly proportional to the produced particle multiplicity, and hence can be inferred from the experimental measurements (Hanus *et al.*, 2019). Inverting Eq. (108), we can relate the dimensionless time \tilde{w} in a longitudinally expanding conformal plasma to Bjorken time τ via the relation

$$\tau = \kappa^{1/2} \tilde{w}^{3/2} (4\pi\eta/s)^{3/2} (s\tau)_{\text{hydro}}^{-1/2} . \quad (109)$$

The proportionality coefficient $\kappa = (s\tau)_{\text{hydro}}/(\tau T^3)$ becomes a numerical constant in thermal equilibrium, where $\kappa = \nu_{\text{eff}} 4\pi^2/90$. Because the kinetic simulations converge towards conventional viscous hydrodynamics predictions for $\tilde{w} \gtrsim 1$, it was estimated in Ref. (Kurkela *et al.*, 2019a,b) that the hydrodynamic description becomes applicable for times $\tau \gtrsim 1$ fm/c for $\eta/s \approx 0.16$ and typical entropy densities found in central PbPb collisions (Kurkela *et al.*, 2019a,b). This is consistent with the early hydrodynamization picture employed in the modeling of heavy-ion collisions.

2. Entropy production and initial energy density

At even earlier times $\tilde{w} \lesssim 1$, kinetic simulations with very different initial conditions might have not yet collapsed onto a single curve (Almaalol *et al.*, 2020; Kurkela and Zhu, 2015). Nevertheless one may employ the hydrodynamic attractor curve, which is regular for $\tilde{w} \rightarrow 0$, for a macroscopic fluid dynamic description far from equilibrium (Romatschke, 2018) (see also Sec. VI.C). In kinetic theory at early times, such an attractor curve has vanishingly small longitudinal pressure $P_L \approx 0$ and constant energy density per rapidity $\mathcal{E}\tau = \text{const}$. Such initial conditions are typical for kinetic evolution in the bottom-up picture discussed in Sec. V.C. Figure 24 shows the energy density \mathcal{E} normalized by the equilibrium evolution $(\mathcal{E}\tau^{4/3})_{\text{hydro}}/\tau^{4/3}$ for different hydrodynamic attractors obtained from QCD and Yang Mills (YM) kinetic theory (Kurkela and Mazeliauskas, 2019a,b; Kurkela *et al.*, 2019a,b), AdS/CFT (Heller *et al.*, 2012b; Heller and Spalinski, 2015; Romatschke, 2018) and Boltzmann RTA (Behtash *et al.*, 2019b; Blaizot and Yan, 2018; Heller *et al.*, 2018; Strickland, 2018; Strickland *et al.*, 2018). All attractors approach the universal viscous hydrodynamic description at late times $\tilde{w} > 1$, while at early times they follow $\mathcal{E} \sim \tau^{-1}$, corresponding to “free-streaming” behavior⁴¹, which can be expressed as

$$\frac{\mathcal{E}\tau^{4/3}(\tilde{w} \ll 1)}{(\mathcal{E}\tau^{4/3})_{\text{hydro}}} = C_\infty^{-1}\tilde{w}^{4/9}. \quad (110)$$

Here the dimensionless constant C_∞ quantifies the amount of work done.

A directly observable consequence of the equilibration process is the particle multiplicity which is a measure of the entropy produced in heavy-ion collisions (Muller and Schafer, 2011). For a given hydrodynamic attractor, the final entropy for boost invariant expansion is proportional to the initial energy and is given by a simple formula (Giacalone *et al.*, 2019)

$$(s\tau)_{\text{hydro}} = \frac{4}{3}C_\infty^{3/4} \left(4\pi\frac{\eta}{s}\right)^{1/3} \kappa^{1/3} (\mathcal{E}\tau)_0^{2/3}, \quad (111)$$

Ref. (Giacalone *et al.*, 2019) shows that combining the entropy production from hydrodynamic attractors with initial initial state energy deposition in the CGC framework gives a good description of the centrality dependence of measured particle multiplicities. In particular, one can extend the original Bjorken estimate (Bjorken, 1983) of the initial energy density in heavy-ion collisions to much earlier times. For central Pb-Pb collisions at

⁴¹ The presence of scattering terms in Eq. (81) are crucial for the early time anisotropy evolution, but not for the energy density. According to the equations of motion $\partial_\tau(\tau\mathcal{E}) = -P_L$, and $\tau\mathcal{E} \approx \text{const}$ as long as $P_L/\mathcal{E} \ll 1$.

$\sqrt{s_{\text{NN}}} = 2.76$ TeV one finds $\mathcal{E}(\tau_0) = 270$ GeV/fm³ at $\tau_0 = 0.1$ fm/ c — nearly a thousand times larger energy density than at the QCD crossover temperature.

3. Chemical equilibration of QGP

The early quark production from strong gauge fields was discussed in Sec. IV.E. However once the gluon fields are no longer overoccupied, chemical equilibration has to be described using QCD effective kinetic theory. A study of light quark flavor (*up*, *down* and *strange*) chemical equilibration in isotropic and longitudinally expanding systems were recently presented in (Kurkela and Mazeliauskas, 2019a,b). At leading order, there are two fermion production channels: gluon fusion $gg \rightarrow q\bar{q}$ and splitting $g \rightarrow q\bar{q}$. It was found that quark production processes are slower than gluon self-interactions. Therefore, for example, the gluon self-similar energy cascade seen in non-expanding isotropic systems is over well before an appreciable number of fermions is produced. Similarly, gluons maintain an approximate kinetic equilibrium among themselves, while fermions attain a Fermi-Dirac distribution at much later times.

The longitudinal expansion drives both gluons and fermions from the kinetic equilibrium, ensuring that equilibrium distributions can only be approached at late times when the expansion rate slows down. However the expansion does not seem to affect fermion production; therefore chemical equilibrium is achieved *before* thermal equilibrium. For massless quarks, the quark-gluon plasma satisfies the conformal equation of state $\mathcal{P} = \frac{1}{3}\mathcal{E}$ and the chemical composition of the plasma has little effect on the total evolution of the energy-momentum tensor. Therefore hydrodynamization, chemical and thermal equilibrium are achieved sequentially (Kurkela and Mazeliauskas, 2019a,b) satisfying

$$\tau_{\text{hydro}} < \tau_{\text{chem}} < \tau_{\text{therm}} . \quad (112)$$

Figure 25 shows the total energy density (red solid line), gluon energy density (green dotted line) and quark energy density (blue dashed line) as a function of time. Gluons, which dominate initially, are quickly overtaken by quarks and the approximate chemical equilibrium energy ratios are reached by $\tau = 1.5$ fm/ c . This supports assumption of chemical equilibrium in the lattice equation of state used in hydrodynamic simulations of the quark-gluon plasma.

Finally, an important piece of evidence for the formation of a chemically equilibrated QGP in heavy-ion collisions is the enhanced production of hadrons carrying strange quarks (Andronic *et al.*, 2018). It is believed that in small collision systems such as proton-proton collisions, strange quarks are not produced thermally in sufficient numbers and therefore strange hadron production is suppressed. Although in the kinetic description above the three light flavors are all taken to be massless, the chemical equilibration rate can be used to estimate the necessary life time (and system size) for the creation of the chemically equilibrated QGP. The results in Ref. (Kurkela and Mazeliauskas, 2019a) show that the plasma may reach chemical equilibrium for particle multiplicities down to $dN_{\text{ch}}/d\eta \sim 10^2$. Strange hadron production in such high multiplicity proton-proton collisions will be tested in future runs of the LHC (Citron *et al.*, 2019).

4. Equilibration of spatially inhomogeneous systems

Up to this point, we discussed the equilibration of longitudinally expanding but otherwise homogeneous systems. Realistic heavy-ion collisions create initial conditions which are not homogeneous in the transverse plane. Such geometric deformations are strongly believed to be the source of the multi-particle correlations observed experimentally (Heinz and Snellings, 2013). In the weak coupling picture discussed in Sec. III, the spatial fluctuations are the result of the uneven color charge distributions in the colliding nuclei. On the largest scale ~ 10 fm it is determined by the overlap of the average nuclear profiles. On nucleon scales ~ 1 fm one can resolve event-by-event fluctuations of individual colliding nucleons. On yet smaller scales ~ 0.1 fm one has stochastic fluctuations of color charges in the internal structure of a nucleon.

Equilibration in kinetic theory, of small transverse perturbations around the homogeneous far-from-equilibrium background, has been investigated in several works (Keegan *et al.*, 2016a; Kurkela *et al.*, 2019a,b). Relevant information on the complicated kinetic evolution of the particle distribution $f_{\mathbf{p}}^s$ can be captured by the linearized energy-momentum tensor

response functions $G_{\alpha\beta}^{\mu\nu}$

$$\begin{aligned} \delta T_{\mathbf{x}}^{\mu\nu}(\tau_{\text{hydro}}, \mathbf{x}) &= \int d^2\mathbf{x}' G_{\alpha\beta}^{\mu\nu}(\mathbf{x}, \mathbf{x}', \tau_{\text{hydro}}, \tau_{\text{EKT}}) \\ &\times \delta T_{\mathbf{x}}^{\alpha\beta}(\tau_{\text{EKT}}, \mathbf{x}') \frac{\overline{T}_{\mathbf{x}}^{\tau\tau}(\tau_{\text{hydro}})}{\overline{T}_{\mathbf{x}}^{\tau\tau}(\tau_{\text{EKT}})}. \end{aligned} \quad (113)$$

Here the Green functions $G_{\alpha\beta}^{\mu\nu}(\mathbf{x}, \mathbf{x}', \tau_{\text{EKT}}, \tau_{\text{hydro}})$ describe the evolution and equilibration of energy-momentum tensor perturbations from an early time τ_{EKT} to a later time τ_{hydro} .

Remarkably, the linearized response functions are to a good approximation universal functions of the dimensionless time \tilde{w} , similar to the hydrodynamic attractor describing the background equilibration. This provides a practical tool—the linearized pre-equilibrium propagator KØMPØST—for a pre-equilibrium kinetic description of heavy-ion collisions based on QCD kinetic theory—(Kurkela *et al.*, 2019a,b). For the first time, the combination of the initial state IP-Glasma model discussed in Sec. III.C.3, kinetic equilibration and viscous hydrodynamics evolution make it possible to describe all the early stages of heavy-ion collisions in a theoretically complete setup. Experimental signatures of such setups are currently being investigated (Gale *et al.*, 2020; Schenke *et al.*, 2020a).

Similarly to the evolution of the background, the equilibration of linearized perturbations in QCD kinetic theory shares universal features with other microscopic descriptions (Broniowski *et al.*, 2009; Liu *et al.*, 2015; Romatschke, 2015; van der Schee *et al.*, 2013). Thanks to this universal behavior, “universal pre-flow” is guaranteed to grow linearly with time for small gradients $\nabla\mathcal{E}/\mathcal{E} \ll 1$ (Keegan *et al.*, 2016a; Kurkela *et al.*, 2019a; Vredevoogd and Pratt, 2009)

$$\vec{v} \approx -\frac{1}{2} \frac{\vec{\nabla}\mathcal{E}}{\mathcal{E} + \mathcal{P}_T} \tau, \quad (114)$$

where, for long wavelength perturbations, $\vec{\nabla}\mathcal{E}/(\mathcal{E} + \mathcal{P}_T) = \text{const}$ in conformal theories (Keegan *et al.*, 2016a). These response functions have been compared directly in Yang-Mills and RTA kinetic theories (Kamata *et al.*, 2020).

QCD kinetic theory simulations beyond the linearized regime have not been accomplished to date, albeit there exist phenomenological studies of parton transport simulations based on perturbative QCD matrix elements (Greif *et al.*, 2017). To what extent the macroscopic description in terms of hydrodynamics can be applicable in inhomogeneous systems with non-linear transverse expansion is still an open question (see Sec. VI.E.2 for a discussion in holography). However, encouragingly, the results of several works (Kurkela *et al.*, 2020,

2018b, 2019c,d) have demonstrated that for transversely expanding systems, the hydrodynamic attractor remains a good description of local equilibration until the evolution time becomes comparable to the transverse system size.

VI. AB INITIO HOLOGRAPHIC DESCRIPTION OF STRONG COUPLING PHENOMENA

A. Holography and heavy-ion collisions

The preceding sections were concerned with the description of heavy-ion collisions in a weak coupling QCD framework. This section will present what currently constitutes the only approach capable of describing real time phenomena in genuinely strongly coupled (1+3)-dimensional quantum field theories in a fully *ab initio* manner – holography (Gubser *et al.*, 1998a; Maldacena, 1998; Witten, 1998a).

The available description in this case does not make visible use of the gauge field degrees of freedom. Instead, it is based on the notion of a correspondence to higher dimensional geometries, which arise as solutions of Einstein’s equations with negative cosmological constant and appropriate matter fields.

The guiding principle for our presentation will be universality. We will be interested in phenomena shared across strongly-coupled quantum field theories and seek in them theoretical lessons and phenomenological implications for thermalization in QCD.

A prime example of such a quantity is the aforementioned $\eta/s = 1/(4\pi)$ in *all* holographic QFTs as long as they are described by two-derivative gravity theories. One purpose of this article is to review other kinds of universalities that exist in the genuine non-equilibrium regime.

B. Controlled strong coupling regime

The best-known holographic gauge theory is the $\mathcal{N} = 4$ super Yang-Mills theory. At the Lagrangian level, it can be viewed as the gluon sector of $SU(N_c)$ QCD coupled in a maximally supersymmetric way to 4 Weyl fermions and 6 real scalars, both in the adjoint representation of the gauge group (Ammon and Erdmenger, 2015). This theory, as opposed

to QCD, is conformally invariant; the coupling constant does not run with the energy and becomes an external parameter that defines the theory.

In the planar $N_c \rightarrow \infty$ limit for asymptotically large values of the 't Hooft coupling constant

$$\lambda \equiv 4\pi\alpha_S N_c \rightarrow \infty, \quad (115)$$

the degrees of freedom in the $\mathcal{N} = 4$ super Yang-Mills theory reorganize themselves in such a way that correlation functions of certain operators including the energy-momentum tensor in a whole class of interesting states can be computed using a 5-dimensional Einstein gravity action with a negative cosmological constant

$$S_{\text{grav}} = \frac{1}{16\pi G_N} \int d^5x \sqrt{\det g} \left\{ R - 2 \left(-\frac{6}{L^2} \right) \right\} \quad (116)$$

and supplemented with matter fields. Here R is the Ricci scalar and L is the length scale set by the cosmological constant. For the $\mathcal{N} = 4$ super Yang-Mills theory at $\lambda \rightarrow \infty$ one has

$$\frac{L^3}{G_N} = \frac{2 N_c^2}{\pi} \quad (117)$$

and a particular matter sector. They both follow from relevant string theory considerations (Maldacena, 1998).

One should view the Einstein gravity description to be applicable only when $\lambda \rightarrow \infty$. Since the QFT coupling constant does not appear in any form in Eq. (116), it indicates that the coupling constant dependence drops from all the QFT quantities one can describe in this way for $\lambda \rightarrow \infty$. When the coupling constant is large, but not infinite, the relevant description becomes Einstein gravity supplemented with higher-curvature terms like the fourth power of the curvature. The form of these terms follows again from string theory considerations and in controllable situations they should be necessarily treated as small corrections. Due to equations of motion becoming generically higher order in derivatives, the uncontrollable extrapolation of the kind one does in, e.g., kinetic theory can be done here only in a very limited number of cases (Woodard, 2015). We will discuss these topics in Sec. VI.F.2.

The “vanilla” setting in holography is 5-dimensional gravity with negative cosmological constant encapsulated by Eq. (116), which provides a consistent dual holographic description of an infinite class of strongly coupled conformal field theories (CFTs) with a large number of microscopic constituents (Bhattacharyya *et al.*, 2008). Specifically, it describes a class of

states in strongly coupled CFTs in which the only local operator acquiring an expectation value is the energy-momentum tensor $T^{\mu\nu}$. The most comprehensive holographic results on heavy-ion collisions concern this particular case.

A generic 5-dimensional metric can be always brought to the form

$$ds^2 = \frac{L^2}{u^2} (-du^2 + g_{\mu\nu}(u, x) dx^\mu dx^\nu). \quad (118)$$

Here u is an additional direction emerging on the gravity side interpreted as a scale in a dual QFT. Einstein's equations put conditions on acceptable forms of $g_{\mu\nu}(u, x)$. The most symmetric solution for gravity with negative cosmological constant has $g_{\mu\nu}(u, x) = \eta_{\mu\nu}$, which is the 4-dimensional Minkowski metric. This is the empty AdS_5 (anti-de Sitter) solution, which represents in gravitational language the time development of the vacuum in holographic CFTs. The surface $u = 0$ acts as a boundary of AdS_5 and, more generally, $g_{\mu\nu}(u = 0, x)$ has the interpretation of a metric in which the corresponding QFT lives.

The expectation value of the energy-momentum tensor arises by looking at the subleading behaviour of $g_{\mu\nu}(u, x)$ close to the boundary ([Balasubramanian and Kraus, 1999](#); [de Haro et al., 2001](#)). This is particularly simple for CFTs living in Minkowski space:

$$g_{\mu\nu}(u, x) = \eta_{\mu\nu} + \frac{4\pi G_N}{L^3} \langle T_{\mu\nu} \rangle(x) u^4 + \dots \quad (119)$$

The ellipsis denotes higher order terms in the small- u expansion that turn out to contain only even powers of u with the coefficients being polynomials in $\langle T^{\mu\nu} \rangle$ and its derivatives. One cannot a priori exclude terms like $\exp(-1/u)$ which were considered for example in Ref. ([Heller et al., 2013b](#)), but a general understanding of such terms is lacking. In the following, we will refer to the interior of AdS spacetimes as “bulk physics” and the QFT physics as “boundary physics”.

We are interested here in discussing time dependent states in Minkowski spacetime that model the dynamics of heavy-ion collisions. Given Eq. (119), such states can be probed through their expectation value of the energy-momentum tensor by solving the equations of motion of Eq. (116) as an initial value problem. This is achieved using numerical relativity techniques ([Chesler and Yaffe, 2014](#); [Heller et al., 2012a](#); [Liu and Sonner, 2018](#)) and requires one to specify initial conditions and the solutions are subject to boundary conditions at $u = 0$.

There are two natural ways (with pros and cons) of studying the non-equilibrium physics of quantum field theories using holography, see Fig. 26. The first approach circumvents

the problem of finding initial conditions, a key reason for its use in early works on the subject (Chesler and Yaffe, 2009, 2010). Moreover, this approach allows one to compare equilibration across theories by starting with the same kind of initial state (for example, the vacuum or a thermal state) and perturbing it in a defined manner. In particular, it underlies a significant body of research on understanding features of linear response theory in different microscopic models (Grozdanov *et al.*, 2016; Kovtun and Starinets, 2005; Kurkela and Wiedemann, 2019; Romatschke, 2016). As an example, Ref. (Keegan *et al.*, 2016b) discussed in Sec. V.E – see Fig. 23 – compared the approach to hydrodynamics across models (including holography) using fully non-linear kicks. The drawbacks to perturbing simple states are that firstly, the approach to hydrodynamics is so rapid that it is hard to disentangle exciting the system from its subsequent relaxation; secondly, the class of states one obtains in this way is rather limited.

The second method, in which one solves gravity equations for different initial conditions, allows one to access a larger range of transient behavior. In particular, since we do not know which initial conditions are closest to the physics realized in experiment, one may wish to scan as many of these initial conditions as possible to obtain a comprehensive picture. The downside is that in most cases this way of phrasing the problem is very specific to the geometric language of describing strongly coupled QFTs similarly to the one-particle distribution function being very specific to the weak coupling language. It does not allow for controllable comparisons with other frameworks akin to Ref. (Keegan *et al.*, 2016b). This can be somewhat ameliorated in holographic collisions in which the initial conditions for gravity originate from superimposing two exact solutions corresponding to individual projectiles approaching each other.

Thermalization at strong coupling is a process in which one starts with an excited geometry in the bulk and after some time it becomes locally very close to a black hole geometry. This encapsulates the notion of thermalization of expectation values of local operators. Non-local observables discussed in Sec. VI.F.3 can still show traces of non-equilibrium behaviour after local thermalization occurs. This should not come as a surprise since the thermalization of non-local observables is necessarily constrained by causality.

The discussion thus far was quite generic but the explicit formulas were provided for strongly coupled CFTs. While QCD is not a CFT, holography does not pose any conceptual problems in studying strongly coupled gauge theories with a non-trivial RG provided the

theory remains strongly coupled at all scales. This can be realized by introducing relevant deformations to holographic CFTs, modifying their Lagrangian by $\int d^4x J O(x)$ with the (scaling) dimension $\Delta < 4$ of $O(x)$. This triggers a non-trivial bulk metric dependence on u , providing the gravitational counterpart of an RG flow.

In holography, the bulk object corresponding to O is a scalar field ϕ appearing in the matter sector that supplements the universal sector in Eq. (116). This scalar field is non-zero because the J of the QFT translates into its asymptotic boundary conditions; the latter generates a non-trivial profile for ϕ when solving the bulk equations of motion. Of course, the action for the bulk matter fields equips ϕ with a potential and the form of the potential determines the physics of the RG flow in the corresponding QFT (including the information about Δ). We will review representative results in Sec. VI.F.1.

To close, holography provides an *ab initio* window to study strongly-coupled QFTs, which include conformal and non-conformal gauge theories. The conceptual problem of fully non-perturbative real time evolution of a whole class of QFTs reduces in this setting to a technical challenge of solving a set of coupled partial differential equations in higher number of dimensions, which is well within reach of the existing numerical relativity methods.

The holographic approach is very general and can be equally well applied to the problem of time evolution of the nuclear medium in heavy-ion collisions, as well as problems originating in branches of physics (Ammon and Erdmenger, 2015; Hartnoll *et al.*, 2016). Finally, we stress again that holography as a tool for QFT comes with its own limitations illustrated by the fact that one needs to work in regimes where the gravity description is classical or semi-classical.

C. Early times in Bjorken flow at strong coupling

Bjorken flow (Bjorken, 1983) without transverse expansion in a CFT setting is arguably the best studied example of a nonlinear non-equilibrium phenomenon in holography⁴². Because of the conservation of the energy-momentum tensor, all the non-trivial information about the dynamics can be extracted from $\langle T^{\tau\tau} \rangle \equiv \mathcal{E}(\tau)$. This parametrization is useful to describe the early time physics relevant for modeling initial stages of ultra-relativistic

⁴² Recently devised hyperbolic quenches (Mitra *et al.*, 2019) adopt effectively (1+1)-dimensional boost invariant geometry of heavy-ion collisions in the context of condensed matter physics.

heavy-ion collisions.

Towards this end, Ref. (Beuf *et al.*, 2009) noticed that combining Eq. (119) (expanded to sufficiently high order in u) with a general Taylor series ansatz for $\mathcal{E}(\tau)$ around $\tau = 0$ does not lead to singular bulk metric coefficients in the limit $\tau \rightarrow 0$ as long as the early time expansion contains only positive even powers of proper time:

$$\mathcal{E}(\tau \approx 0) = \mathcal{E}_0 + \mathcal{E}_2\tau^2 + \mathcal{E}_4\tau^4 + \dots . \quad (120)$$

The coefficients in the above equation are not entirely arbitrary but they are related one-to-one to the near-boundary expansion of the bulk metric that satisfies the constraints on the initial time slice, as encapsulated by Eqs. (118) and Eq. (119). The early time series (120) turns out to have a non-zero, but finite radius of convergence, which allows one to reliably study the initial dynamics of the system. However, as shown in Ref. (Beuf *et al.*, 2009), and later corroborated in Ref. (Heller *et al.*, 2012a) using the full numerical solution of bulk Einstein's equations, the radius of convergence of (120) is much too small to see the transition to hydrodynamics. This point is illustrated in Fig. 27 using the effective temperature⁴⁰. Furthermore, simple analytic continuations of the series (120) based on the Padé approximants method provide unreliable extrapolations.

One lesson therefore is that the only method to obtain $\langle T^{\mu\nu} \rangle$ in strongly coupled QFTs beyond the early time limit examples is to use numerical relativity. Before we proceed there, a few more comments related to Eq. (120) are in order. Firstly, the analysis of Ref. (Beuf *et al.*, 2009) uses regularity of the initial metric on a particular constant time slice of the bulk geometry, namely the one dictated by the coordinates chosen in Eq. (118). It is therefore logically possible⁴³ that there are initial metrics defined on other bulk constant time slices that give rise to energies densities of the form different than dictated by Eq. (120). Second, note that in Eq. (120) any number of the lowest order terms can vanish and the energy density at early time can behave like, e.g., $\mathcal{E}|_{\tau \approx 0} \sim \tau^2$ (Grumiller and Romatschke, 2008).

Another point is that there are various reasons why one may not want to start the evolution at $\tau = 0$. The most obvious one is related to creating either non-equilibrium initial states from the vacuum or thermal states, as discussed in Fig. 26. In these cases, the sources will need some non-zero time to act (Chesler and Yaffe, 2010). The other reason

⁴³ Ref. (Jankowski *et al.*, 2014) chose initial surfaces in the bulk as in Fig. 26 with results consistent with Eq. (120).

is more conceptual and is related to the observation that while one should not expect the infinitely strongly coupled approach to be a phenomenologically viable description at $\tau = 0$, it may become one from some $\tau > 0$ onward. Note that from the gravity point of view, it is far from clear that all the initial conditions set in the bulk for $\tau > 0$ are extendable to $\tau = 0$ and, as a result, one can view them as, a priori, containing richer behavior.

Because of this issue, it is unclear whether all well behaved initial conditions for numerical relativity simulations actually describe genuine states in underlying QFTs. As opposed to Refs. (Heller *et al.*, 2012a,b; Jankowski *et al.*, 2014), the papers initializing their codes at later times with turned off sources include Refs. (Kurkela *et al.*, 2020; Romatschke, 2018; Wu and Romatschke, 2011). In particular, Ref. (Romatschke, 2018) finds initial conditions at some early but non-zero τ such that $\mathcal{E} \sim \frac{1}{\tau}$ initially, which is clearly very different from Eq. (120).

As already discussed in Sec. V.E, the transition to hydrodynamics can be observed in the cleanest way upon introducing the scale-invariant time variable \tilde{w} defined in Eq. (108) and using $\mathcal{P}_T/\mathcal{P}_L$, $\mathcal{P}_L/\mathcal{E}$ or any reasonable function of this ratio such as

$$\mathcal{A} = \frac{\mathcal{P}_T - \mathcal{P}_L}{\mathcal{E}/3} = \frac{3\frac{\mathcal{P}_T}{\mathcal{P}_L} - 3}{2\frac{\mathcal{P}_T}{\mathcal{P}_L} + 1} \quad (121)$$

introduced in Refs. (Florkowski *et al.*, 2018a; Heller *et al.*, 2012b; Jankowski *et al.*, 2014) as a function of $w \equiv \tau T$. Note that in strongly coupled limit of holography $4\pi\eta/s = 1$ and we will simply denote then \tilde{w} as w .

It is well understood by now that at late time $\mathcal{A}(w)$ acquires the form of a trans-series (Aniceto *et al.*, 2019b; Aniceto and Spaliński, 2016; Heller *et al.*, 2013a; Heller and Spalinski, 2015) known from the studies of asymptotic expansions in mathematical and quantum physics, see Refs. (Aniceto *et al.*, 2019a; Dorigoni, 2019) for reviews. The hydrodynamic part is a series in inverse powers of w and has a vanishing radius of convergence⁴⁴. Its first few terms read

$$\begin{aligned} \mathcal{A}(w) = & \frac{2}{\pi}w^{-1} + \frac{2 - 2\log 2}{3\pi^2}w^{-2} \\ & + \frac{15 - 2\pi^2 - 45\log 2 + 24\log^2 2}{54\pi^3}w^{-3} + \dots, \end{aligned} \quad (122)$$

⁴⁴ The same applies to Gubser (Denicol and Noronha, 2019b) and cosmological (Buchel *et al.*, 2016) flows, but is not the case for Bjorken flow with fine-tuned transport coefficients (Denicol and Noronha, 2019a). Furthermore, Ref. (Heller *et al.*, 2020b) used the results of (Grozdanov *et al.*, 2019a,b; Withers, 2018) to show that divergence of the hydrodynamic gradient expansion is a generic feature of the linear response theory.

see Refs. (Booth *et al.*, 2009; Florkowski *et al.*, 2018a; Heller and Janik, 2007; Heller *et al.*, 2012b, 2009; Janik, 2007; Jankowski *et al.*, 2014; Kinoshita *et al.*, 2009a,b; Nakamura and Sin, 2006). This relation should be understood as expressing the energy-momentum tensor in terms of hydrodynamic constitutive relations to the third lowest order. The first term carries information about the first derivative of flow velocity and the shear viscosity, the second term is a contribution from second derivatives of velocity and associated transport coefficients. The third term is the last one that is known analytically. The current state of the art is set by Ref. (Casalderrey-Solana *et al.*, 2018) which, improving on the earlier efforts of Ref. (Heller *et al.*, 2013a), computed numerically the lowest 380 terms in the expansion Eq. (122). On top of the power law late time (w) expansion come exponentially suppressed terms that represent transient phenomena visible also in the linear response theory (Heller *et al.*, 2013a; Heller and Spalinski, 2015; Heller and Svensson, 2018; Janik and Peschanski, 2006b).

Fig. 27 illustrates time evolution of the effective temperature $T(\tau)$. Hydrodynamics is applicable at a time after which the pressure anisotropy deviates from Eq. (122) by very little. As discussed in detail in Ref. (Heller *et al.*, 2012a), the precise moment of applicability of hydrodynamics depends on the desired accuracy of the match to Eq. (122) and the order of the truncation. Of course, the latter aspect should be understood in the sense of an asymptotic series.

The main message from the studies in Refs. (Chesler and Yaffe, 2010; Heller *et al.*, 2012a,b; Jankowski *et al.*, 2014; Kurkela *et al.*, 2020; Romatschke, 2018; Wu and Romatschke, 2011) and related works is that low order hydrodynamic constitutive relations (see Eq. (122)) become applicable at strong coupling after $\tau = \mathcal{O}(1/T)$. This is the regime where the pressure anisotropy in the system is sizable, as illustrated in Fig. 28. Since the system is still far away from local thermal equilibrium, the word hydrodynamization was coined in (Casalderrey-Solana *et al.*, 2014b) to distinguish the applicability of viscous hydrodynamics constitutive relations from local thermalization. In particular, the latter phenomenon occurs at strong coupling for times which can be even 10 times larger than the hydrodynamization time.

The modern perspective on hydrodynamics, viewing in particular the gradient expansion as a part of a trans-series, is reviewed in detail in Ref. (Florkowski *et al.*, 2018a). In the following, we will discuss an alternative way of thinking about the applicability of hydrodynamics using the concept of hydrodynamic attractors. These objects already made their

appearance in Sec. V.E and bear a structural similarity to non-thermal attractors (fixed points) discussed in Sec. IV.C.

D. Hydrodynamic attractors in holography

Hydrodynamic attractors proposed in Ref. (Heller and Spalinski, 2015), and developed in many works including Refs. (Almaalol *et al.*, 2020; Almaalol and Strickland, 2018; Aniceto and Spaliński, 2016; Basar and Dunne, 2015; Behtash *et al.*, 2019a, 2018, 2019c; Blaizot and Yan, 2018, 2020; Brewer *et al.*, 2019; Casalderrey-Solana *et al.*, 2018; Chattopadhyay and Heinz, 2020; Dash and Roy, 2020; Denicol and Noronha, 2018, 2019a,b; Florkowski *et al.*, 2018b; Heller *et al.*, 2020a; Jaiswal *et al.*, 2019; Kurkela *et al.*, 2020; Romatschke, 2017, 2018; Shokri and Taghinavaz, 2020; Spaliński, 2018; Spaliński, 2018; Strickland, 2018; Strickland *et al.*, 2018; Strickland and Tantary, 2019) can be viewed as a way of approaching the problem of information loss about the underlying state from the point of view of observations restricted to the energy-momentum tensor $\langle T^{\mu\nu} \rangle$.

Reexamining Fig. 28 through these lenses, we see that a set of different states considered there follows to a good approximation a single profile of $\mathcal{A}(w)$ from a certain value of w onward. This is the notion of attraction between different initial conditions as seen by an *effective* phase space covered by \mathcal{A} at a fixed value of w . While this observation does not call for invoking a truncated gradient expansion, the emerging universality seen in Fig. 28 agrees very well with hydrodynamic gradient expansion truncated at low order. These observations lie behind the name hydrodynamic attractor and parallel the discussion in Sec. V.E.1.

Let us step back and review this phenomenon from a broader perspective advocated recently in Ref. (Heller *et al.*, 2020a). To proceed, we will utilize the aforementioned notion of phase space introduced in this context in Ref. (Behtash *et al.*, 2018). Specifically, one should think of \mathcal{A} as a particularly clean (scale invariant) way of representing information about $\langle T^{\mu\nu} \rangle$ and w as a useful way of parametrizing time evolution, adjusted to the fact that transient phenomena in conformal theories occur over time scales set by the energy density.

Of course, knowing \mathcal{A} at a given value of w does not allow one to predict its value later, since the true microscopic variable is the bulk metric. A bigger chunk of information is provided by considering \mathcal{A} and some of its derivatives with respect to w (or \mathcal{E} and its derivatives with respect to τ). Such sets of variables form the notion of an effective phase

space. In fact, there is a limit to how big such phase space needs to be – numerical solutions of Einstein’s equations displayed in Fig. 28 require typically specifying a few functions on several dozen grid points.

One can then assign a metric to an effective phase space, i.e. the distance between points representing here classes of solutions, and track how such a distance changes as time evolves. The loss of information is expected to make a set of solutions reduce its volume in the effective phase space. For example, in Fig. 28 one introduces the notion of proximity between two solutions $|\mathcal{A}_1(w) - \mathcal{A}_2(w)|$. With respect to this notion, various solutions from the chosen set eventually approximately collapse to a point in \mathcal{A} at a fixed value of w . It should be clear that the hydrodynamic attractor at a given value of w is not a notion relevant for all states. It needs to be regarded as a statement about properties of some class of states initialized prior to that.

Furthermore, assigning a distance measure to phase spaces allows one to define the notion of slow evolution. It was introduced to this topic in Ref. (Heller and Spalinski, 2015) under the name slow roll approximation, which originates from the field of inflationary cosmology (Liddle *et al.*, 1994). For example, the distance notion discussed above leads to the magnitude of velocity of a given state being $|\mathcal{A}'(w)|$ and slowly evolving solutions (note that Ref. (Heller *et al.*, 2020a) was defining rather regions of slow evolution) are those which lead to the flattest form of $\mathcal{A}(w)$. In Fig. 28, such a solution found in Ref. (Romatschke, 2018) by fine tuning initial conditions is denoted by orange. What is quite remarkable is that this solution at early times has \mathcal{A} very close to $\frac{3}{2}$. This corresponds to free streaming $\mathcal{P}_L = 0$, which evades the study of initial conditions behind Eq. (120) reported in Ref. (Beuf *et al.*, 2009).

It is also important to stress that the notion of slowly evolving solutions is a priori independent from the notion of convergence (attraction). However in full phase space, or at least a representative projection of it, one can make a thermodynamic-like argument, as in Ref. (Heller *et al.*, 2020a), in favor of typical states residing in the slow roll region. One can think of the slow evolution as a generalization of the notion of the gradient expansion that does not involve an expansion with individual terms badly behaving at very early times, namely, as inverse powers of w in Eq. (122).

Finally, the approach to the hydrodynamic attractor at strong coupling and mechanisms that govern it were questions raised in Ref. (Kurkela *et al.*, 2020) by looking at results of

simulations with different initialization time. This is depicted in Fig. 29. The idea behind it, building on earlier results in Refs. (Blaizot and Yan, 2018, 2020), is that information loss can be driven by at least two distinct mechanisms. The first one are exponentially suppressed corrections to Eq. (122), which stem from the linear response theory physics. The characteristic feature of them is that their decay rates do not depend on w . The second mechanisms driving the information loss is expansion, which for the comoving velocity $u^\mu \partial_\mu \equiv \partial_\tau$ gives $\nabla_\mu u^\mu = \frac{1}{\tau}$. What one therefore expects is that information loss predominantly driven by the expansion is going to be faster at earlier times (smaller w) and slower at later times. Indeed, such feature was seen in Ref. (Kurkela *et al.*, 2020) for hydrodynamic models and for the kinetic theory for early initialization times. However, in holography this does not seem to be the case and the approach to the hydrodynamic attractor takes roughly a fixed amount of time regardless of the chosen initialization time, see Fig. 29, which is consistent with it being governed by transients.

E. Holographic collisions

In CFTs, Bjorken flow in the absence of transverse expansion has a high degree of symmetry that allows for comprehensive studies of hydrodynamization and associated phenomena. In particular, the numerical approach pursued in Refs. (Chesler and Yaffe, 2010, 2014; Jankowski *et al.*, 2014) fully determines the evolution of $\langle T^{\mu\nu} \rangle$ as a function of proper time τ upon specifying one positive number (initial energy density \mathcal{E}) and a single function of the AdS direction u , see Eq. (119). As a result, it was possible to comprehensively scan over initial states in search of universal behavior.

If one relaxes these symmetry assumptions and allows for dynamics in the transverse plane (van der Schee, 2013; van der Schee *et al.*, 2013), the space of initial conditions becomes too big to allow for a comprehensive analysis. Therefore one would like to have another guiding principle to arrive at interesting configurations for modeling non-equilibrium evolution of $\langle T^{\mu\nu} \rangle$ in holographic heavy-ion collisions. This key idea is to study holographic collisions of localized lumps of matter (Casalderrey-Solana *et al.*, 2013, 2014a; Chesler, 2015, 2016; Chesler *et al.*, 2015; Chesler and Yaffe, 2011, 2015; Grumiller and Romatschke, 2008).

The localized objects (shockwaves) in question move at the speed of light and are char-

acterized by the following non-zero components of $\langle T^{\mu\nu} \rangle$,

$$\langle T^{00} \rangle = \langle T^{33} \rangle = \pm \langle T^{03} \rangle = \mu_{\pm}(\mathbf{x}_{\perp}) h(x^0 \mp x^3), \quad (123)$$

where x^0 is the lab-frame time, x^3 is the direction along which the object is moving (specified by \mp in the argument of h), $\mu_{\pm}(\mathbf{x}_{\perp}) \geq 0$ is an arbitrary function specifying the transverse profile and $h(x^0 \mp x^3) \geq 0$ is another arbitrary function specifying the longitudinal profile (Chesler, 2015). While a single projectile defined by Eq. (123) is exact, the superposition of two projectiles approaching each other and overlapping in the transverse plane leads to a non-trivial collisional process.

Such collisions should not be regarded as literal models of the early stages of heavy-ion collisions, since the projectiles do not originate from QCD. (See however (Gubser *et al.*, 2008b; Lin and Shuryak, 2009; van der Schee and Schenke, 2015).) Instead, one should treat holographic shockwaves collisions as illustrating possible far-from-equilibrium phenomena accessible in a fully *ab initio* way at strong coupling that goes well beyond the Bjorken flow geometry discussed previously.

1. Planar shocks

The simplest settings to consider are collisions of planar shockwaves – objects defined by Eq. (123) with μ_{\pm} constant. Following Ref. (Casalderrey-Solana *et al.*, 2013), one can consider a Gaussian longitudinal profile for h of the form,

$$h(x^0 \mp x^3) = \frac{N_c^2}{2\pi^2} \varrho^4 e^{-\frac{(x^0 \mp x^3)^2}{2d^2}}, \quad (124)$$

and recognize that, in heavy-ion collisions, the dimensionless product of the amplitude ϱ (not to be confused with the charge density discussed in previous sections) and the width d decreases as $\gamma^{-1/2}$ as the total center-of-mass energy of the collision ($\sqrt{s} = 2\gamma M_{\text{ion}}$) increases.

Within this analogy, high energy collisions correspond to collisions of very thin shockwaves⁴⁵. The collisions of projectiles defined by Eq. (123) do not lead to longitudinal boost invariance since the initial state of the two projectiles is not boost invariant even when they are infinitely thin. The extent to which this is the case was explored in (Casalderrey-Solana

⁴⁵ The problem of colliding planar projectiles in Eq. (123) with $h(x^0 \mp x^3) \sim \delta(x^0 \mp x^3)$ was posed originally in (Janik and Peschanski, 2006a) and addressed in an early time expansion akin to Eq. (120) in (Grumiller and Romatschke, 2008).

et al., 2013) and, quite remarkably, the results fit well (Gubser and van der Schee, 2015) with complex deformations of the purely boost invariant flow introduced in (Gubser, 2013).

As it turns out, the features of the collision change as a function of γ . Firstly, the collision of “low- γ ” (thick) shockwaves proceed such that the two blobs of matter first merge and their subsequent evolution is approximated well by viscous hydrodynamics. This is referred to (Casalderrey-Solana *et al.*, 2013) as to the Landau scenario (Belenkij and Landau, 1956; Landau, 1953). As can be seen in Fig. 30, the “high- γ ” regime of thin shocks leads to a rich set of transient physics before hydrodynamics becomes applicable. Another important phenomenon discussed in (Casalderrey-Solana *et al.*, 2014a; Müller *et al.*, 2020; Waeber *et al.*, 2019) is the notion of longitudinal coherence. This notion applies to the “centre-of-mass” frame of high energy collisions and states that the longitudinal structure of projectiles does not leave an imprint on the transient form of the energy-momentum tensor in the post-collision region provided it is sufficiently localized. Finally, despite the differences between thin and thick shocks’ collisions at transient times after the remnants dissolve, which take much longer time than shown in Fig. 30, the structure of the late time hydrodynamic flow is very similar in the two cases (Chesler *et al.*, 2015).

2. Transverse dynamics in holography

Studies of hydrodynamization in the presence of transverse expansion in (Chesler, 2015, 2016; Chesler and Yaffe, 2015) still define the state-of-the-art in numerical applied holography. Fig. 31 illustrates the profile of the energy density in such collisions. The main lesson from these works is the early applicability of viscous hydrodynamics not just for very large longitudinal gradients of the energy-momentum tensor (as for Bjorken flow and for planar shocks), but also in the presence of large transverse gradients generating transverse expansion.

From the perspective of these strong coupling results, the applicability of hydrodynamics in pA and even pp collisions (Chesler, 2016) is as natural as the applicability of hydrodynamics in Bjorken flow and can be explained in terms of fast decaying contributions to the trans-series for $\langle T^{\mu\nu} \rangle$. Further, these works corroborate studies in (van der Schee, 2013) by providing successful tests of the early time radial expansion model proposed in (Vredevoogd and Pratt, 2009). Towards this end, Ref. (Chesler and Yaffe, 2015) found very small elliptic

flow despite off-central collision and confirmed that near mid-rapidity the energy flux grew linearly with proper time, as predicted in Ref. (Vredevoogd and Pratt, 2009).

As discussed earlier in Sec. V.E.4, such “universal flow” at small wavenumbers is also reproduced by weak coupling kinetic theory. It would be interesting to compare if the full transverse response functions of the energy-momentum tensor in strong coupling agrees with those discussed in Sec. V.E.4 in the context of kinetic theory.

F. Other aspects of thermalization at strong coupling

1. Non-conformal strongly-coupled QFTs

All the strong coupling results reviewed thus far concerned well defined QFTs without a scale. As reviewed in Sec. VI.B, in holography there are no conceptual obstacles to breaking conformal symmetry. However considering QFTs with non-trivial renormalization group flows does make gravitational calculations more involved due to the presence of field(s) in addition to gravity that one needs to solve for and due to the more involved near-boundary analysis that generalizes Eq. (119). All-in-all, the number of results on this front relevant for thermalization in QCD is significantly lower than in the conformal case, but still allows one to draw lessons.

Broadly speaking, there are two approaches to this problem. The first is top-down and studies renormalization group flows originating from turning on a relevant deformation in a known holographic CFT. The prime example is the so-called $\mathcal{N} = 2^*$ gauge theory arising as a deformation of $\mathcal{N} = 4$ super Yang-Mills theory by adding masses to half of its fields (Buchel *et al.*, 2007). The advantage of this approach is that one makes sure that one is studying well defined features of a strongly coupled QFT. The drawback is that such well understood examples are scarce and might have rigid features that do not exist in QCD.

The other class are so-called bottom-up models that couple AdS gravity to a bulk scalar field or fields whose Lagrangian is chosen by insisting on it reproducing some desired feature of QCD. One such approach was introduced in (Gursoy and Kiritsis, 2008; Gursoy *et al.*, 2008) using the QCD β -function as a guideline; another model (Gubser *et al.*, 2008a) uses as a benchmark reproducing the QCD equation of state at vanishing baryon density.

Furthermore, one can also introduce confinement by making the geometry end smoothly

in the bulk (Witten, 1998b). One can think of it as the manifestation of a mass gap, with no excitations below the lowest bound state energy.

The breaking of conformal symmetry introduces an additional scale in the problem of thermalization and changes hydrodynamization times, although in none of the setups explored to date by an order of magnitude or more with respect to the strong coupling CFT prediction of $\sim 1/T$ (Buchel *et al.*, 2015; Janik *et al.*, 2015). This also indicates that \tilde{w} defined in Eq. (108) plays a less prominent role in non-conformal QFTs than it does in strongly coupled CFTs.

Furthermore, the hydrodynamic gradient expansion acquires new transport terms, most notably, the bulk viscosity ζ . Hydrodynamization and (on a much later time scale) isotropization still do occur, but there are now two more emergent time scales related to i) the applicability of the equation of state and ii) the expectation value of the operator breaking conformal symmetry reaching its thermal value. The relation between these scales depends on the details of the model (Attems *et al.*, 2018, 2017a,b).

Finally, confinement represented holographically as the appearance of an infrared wall leads to the new physical effect in which excitations of the bulk geometry and matter fields bounce back and forth as in a cavity (Bantilan *et al.*, 2020; Craps *et al.*, 2015). Such an effect was not present in the studies reviewed earlier and is not yet explored in the context of expanding plasmas.

2. Away from the strong coupling regime

Another important direction studied in the context of thermalization in strongly coupled gauge theories concerns corrections from finite values of the coupling constant. In the context of the $\mathcal{N} = 4$ super Yang-Mills, the leading correction in the inverse power of the 't Hooft coupling constant behaves as $\lambda^{-3/2}$; on the gravity side, it arises at least in part due to a particular expression quartic in the curvature (Gubser *et al.*, 1998b). Such a higher curvature gravity action when treated exactly is ill-behaved due to the Ostrogradsky instability (Woodard, 2015). It is however not meant to be considered as such since it is just an effective field theory truncated at a fixed order in the derivative expansion.

Treating these higher curvature terms as small contributions to the Einstein's equations with negative cosmological constant allows one to derive the leading order corrections to

various holographic predictions at $\lambda \rightarrow \infty$. For example, they increase the shear viscosity of the $\mathcal{N} = 4$ super Yang-Mills from $\eta/s = 1/(4\pi)$ at $\lambda \rightarrow \infty$ (Policastro *et al.*, 2001) to $\eta/s = 1/(4\pi) \times (1 + 15 \zeta(3) \lambda^{-3/2})$ for large, but finite λ (Buchel, 2008a,b).

The quartic term discussed above is the first higher order term appearing for the $\mathcal{N} = 4$ super Yang-Mills, but one should remember that the Einstein-Hilbert action with negative cosmological constant describes infinitely many strongly coupled CFTs. For some of these (Buchel *et al.*, 2009), the leading correction to Eq. (116) is quadratic in curvature and can be written as the so-called Gauss-Bonnet term

$$\delta S_{\text{grav}}^{GB} = \frac{\lambda_{GB}}{2} L^2 (R^2 - 4R_{ab}R^{ab} + R_{abcd}R^{abcd}) . \quad (125)$$

This contribution has $|\lambda_{GB}| \ll 1$ in top-down settings and the sign of λ_{GB} can be either positive or negative.

As a result, there are bona fide holographic CFTs for which the ratio of shear viscosity to entropy density is slightly lower than $1/(4\pi)$ (Buchel *et al.*, 2009; Kats and Petrov, 2009). This important result showed that the celebrated value of $1/(4\pi)$ *is not* the lower bound in nature as originally conjectured in Ref. (Kovtun *et al.*, 2005), although the existence of another lower bound cannot be excluded.

Furthermore, the combined gravity action of (116) and (125) leads to, at least superficially, second order equations of motion. While it is known that microscopically this does not correspond to a well behaved QFT outside the regime $|\lambda_{GB}| \ll 1$ (Camanho *et al.*, 2016), in the spirit of bottom-up models discussed in Sec. VI.F.1 one can treat it, at least in some cases, as a model of QFT at a finite value of the “coupling constant”.

In the context of planar shockwaves collisions discussed in Sec. VI.E, perturbative calculations in λ_{GB} predict less stopping and more energy deposited close to the light-cone (Folkestad, Asmund and Grozdanov, Sašo and Rajagopal, Krishna and van der Schee, Wilke, 2019; Grozdanov, Sašo and van der Schee, Wilke, 2017). There appears to be also a correlation between the shear viscosity and hydrodynamization times, as encapsulated by (108).

Furthermore, linear response calculations performed exactly in λ_{GB} reveal that the singularity structure of real time correlators in equilibrium can change drastically as the coupling is varied (Grozdanov *et al.*, 2016). In particular, the results seem to mimic features expected from a kinetic theory, such as the appearance of branch cuts (Kurkela and Wiedemann,

2019; Romatschke, 2016), rather than single pole singularities known in strongly coupled QFTs (Kovtun and Starinets, 2005).

The situation at a nonlinear level is more complicated. While the equations of motion are second order, the coefficients in front of the highest derivative terms are complicated and can vanish in regions of spacetime. This signals a breakdown of the initial value problem. Overcoming this obstacle is currently an active topic of research in the relativity community (Cayuso *et al.*, 2017; Ripley and Pretorius, 2019, 2020a,b).

Finally, we wish to bring the reader's attention to a more phenomenological set of hybrid approaches (Ecker *et al.*, 2018; Iancu and Mukhopadhyay, 2015; Kurkela *et al.*, 2018a; Mukhopadhyay *et al.*, 2016) in which gravity is used to model the IR of a QFT and a weak coupling framework is put to work to represent the UV. Both frameworks are coupled to each other and predictions rely on a subtle interplay between the two combined models. Such a setting bears structural similarity to (Gursoy and Kiritsis, 2008; Gursoy *et al.*, 2008) discussed in the previous section. However it uses the gravitational description only where it can be trusted, which is the regime where the coupling constant is large.

3. *Non-local correlators*

All the quantities discussed by us at strong coupling until now concerned one-point functions of gauge invariant operators. Because of the underlying large- N_c hierarchy, the problem of finding connected two- and higher-point functions correlation functions decouples from the problem of finding the one-point functions discussed so far. Such correlation functions can be thought of as correlation functions of the bulk free (for two-point functions) or weakly interacting (for higher-point functions) quantum fields⁴⁶ living on top of gravitational backgrounds when the insertion points of the bulk correlators are taken to the boundary (Banks *et al.*, 1998). In the following, we will focus on two-point functions.

Since we are talking about time dependent setups and, hence, Lorentzian correlators, the distinction between Wightman, retarded, etc correlators is appropriate (Herzog and Son, 2003; Skenderis and van Rees, 2008, 2009; Son and Starinets, 2002). Towards this end, the retarded correlator depends only on the gravitational background and captures the

⁴⁶ They should not be confused with the underlying strongly coupled QFT for which both the classical bulk background, as well as free bulk quantum fields are an effective description.

response of the strongly coupled QFT to sources. However the Wightman correlator depends both on the constructed gravitational background and the state of the bulk quantum field. Therefore its calculation is challenging in time dependent processes and, unless one creates a non-equilibrium state using sources exciting the vacuum or a thermal state (Chesler and Teaney, 2011, 2012; Keranen and Kleinert, 2015, 2016), one has to deal with an additional freedom of initial conditions to scan.

It should perhaps not come as a surprise that to date there were no studies of such correlators in an expanding plasma. Noteworthy works in this area are Refs. (Chesler and Teaney, 2011, 2012; Keranen and Kleinert, 2015, 2016), which studied equilibration of scalar operator two-point functions under a spatially uniform quench.

Many references use a proxy for correlators being a bulk geodesic spanned between the insertion points appropriate for operators of large scaling dimension in the Euclidean signature. However, in Lorentzian signature, this is an uncontrollable approximation (Headrick *et al.*, 2014; Keranen and Kleinert, 2015; Louko *et al.*, 2000). On the other hand, the comparison between Wightman functions calculated according to the correct microscopic prescription, and the geodesic proxy, led to qualitatively similar results (Keranen and Kleinert, 2015, 2016).

If one takes this as an indication of the geodesic proxy as capturing the relevant physics, then one lesson following from such studies is that the symmetrized correlator with small spacelike separation between its insertion points thermalizes earlier than the one with larger separation (Balasubramanian *et al.*, 2011a,b). This is also natural from the point of view of causality.

Furthermore, Ref. (Keranen and Kleinert, 2016) observed a relation between the equilibration time scale of the spatially Fourier transformed Wightman function and the equilibration time scale of $1/T$ governing hydrodynamization at strong coupling and discussed in Sec. VI.C. This study was done for a scalar operator, which does not exhibit a hydrodynamic tail.

It is natural to conjecture that the energy-momentum tensor or a $U(1)$ current Wightman function would take longer to equilibrate due to the presence of hydrodynamic modes, but such studies have not been yet performed. Finally, as noted in Ref. (Keranen and Kleinert, 2016), we should stress that the aforementioned momentum space features of equilibration do not translate easily to the real-space properties. This is so because sharp features in the

correlator do not necessarily reside at small distances.

VII. SIGNATURES OF NON-EQUILIBRIUM QCD

The experimental heavy-ion collision programs at BNL and CERN, combined with advances in theory and empirically motivated models have, over the last couple of decades, greatly advanced our understanding of deconfined QCD matter. The successful multi-observable data-to-model comparisons provide ample evidence that a new phase of matter is created with the thermodynamic properties predicted by lattice QCD (Andronic *et al.*, 2018; Bazavov *et al.*, 2019a; Bellwied *et al.*, 2020; Bernhard *et al.*, 2019; Ding *et al.*, 2016; Gardim *et al.*, 2019; Pang *et al.*, 2018). While thermodynamic features of QCD can possibly also be extracted from neutron star physics—a spectacular recent example being the gravitational radiation pattern of neutron star mergers (Weih *et al.*, 2020)—heavy-ion collisions are likely the only place in the universe where the *non-equilibrium* many-body properties of QCD can be explored.

We will not discuss here signatures of high parton density matter in the hadron wavefunctions that have been discussed elsewhere (Blaizot, 2017). Uncovering definitive evidence for and systematic study of gluon saturation is a major goal of the Electron-Ion Collider (EIC) (Accardi *et al.*, 2016; Aschenauer *et al.*, 2019). We note that diffractive and exclusive signatures of gluon saturation at the EIC are especially promising (Mäntysaari *et al.*, 2020; Mäntysaari and Venugopalan, 2018).

Our focus here will be on quark-gluon matter formed after the collision. In the high parton density framework of the CGC EFT, the Glasma matter at the earliest times is most sensitive to the physics of gluon saturation. Indeed, if the contributions of the initial state can be isolated from that of the final state, heavy-ion collisions could present definitive evidence for gluon saturation.

However, as we will discuss, a clean separation of initial and final state effects in the complex spacetime evolution of the heavy-ion collision is challenging (Adolfsson *et al.*, 2020). Nevertheless, data from both light and heavy-ion collisions at RHIC and the LHC can help constrain key features of gluon saturation, an example being the energy and nuclear dependence of the saturation scale Q_S .

A. Electromagnetic and hard probes

Since the Glasma matter is likely to be far off-equilibrium at the earliest instants of the heavy-ion collision, its features can be extracted most directly in probes that are the least sensitive to the later stages of the collision. The primary candidates here are electromagnetic probes of the medium such as photons and dileptons which, once emitted, do not interact with the medium.

The problem here is that photons and di-leptons are produced continuously through out the spacetime evolution of the quark-gluon matter and from the subsequent hadronic phase as well. Current models of heavy-ion collisions, which include photon yields from the pre-hydro kinetic theory phase tend to under predict the produced photon yields (Churchill *et al.*, 2020; Gale *et al.*, 2020); for an alternative mechanism, see (Oliva *et al.*, 2017).

Photons emitted from the highly occupied Glasma have been suggested as an additional source of radiation (Berges *et al.*, 2017b). While phenomenological model comparisons show a significant Glasma contribution (Garcia-Montero, 2019), the theoretical modeling of photon rates at present carries sizable uncertainty.

Besides photons and di-leptons, inclusive yields of high momentum strongly interacting final states are also sensitive to gluon saturation and to early time dynamics in the heavy-ion collision. These include hadrons at high transverse momenta, jets and heavy quarkonia. Gluon saturation influences the production rates for these processes and rescattering in the Glasma influences their dynamics. These effects are most pronounced for $p_{\perp} \sim Q_S$. We discussed heavy quark pair production in the Glasma in Section IV. The diffusion coefficient of these heavy quarks has been computed recently in this framework, and scales as Q_S^3 (Boguslavski *et al.*, 2020). Heavy quark diffusion in Glasma-like environments and their subsequent evolution have also been explored recently in several works (Carrington *et al.*, 2020; Liu *et al.*, 2019; Mrówczyński, 2018). A non-trivial problem is distinguishing this early-time evolution of heavy quarks from their late time evolution (Akamatsu *et al.*, 2018; Brambilla *et al.*, 2019; Rapp and van Hees, 2010). Similar considerations also hold for the propagation of jets⁴⁷ in the Glasma (Asakawa *et al.*, 2011; Carrington *et al.*, 2017; Dumitru *et al.*, 2008b; Ipp *et al.*, 2020).

⁴⁷ The final stage of “bottom up” thermalization corresponds to the “jet quenching” of partons of momentum $\sim Q_S$ that are quenched to the thermal medium; this framework also explains key features of the quenching of very high momentum jets in the QGP (Blaizot *et al.*, 2013, 2016).

Higher point correlations of hard probes, add significant sensitivity to the dynamics of quark-gluon matter off-equilibrium. An example is the potential of two-particle Hanbury-Brown–Twiss (HBT) photon interferometry to study early time dynamics (Garcia-Montero *et al.*, 2019). Such measurements are sensitive to the large longitudinal-transverse anisotropies that are not reflected in photon yields. However experimental measurements of soft photon correlations are very challenging experimentally and high statistics will be needed to disentangle the signal.

B. Long-range rapidity correlations

Long-range rapidity correlations are an important tool in unentangling initial and final state effects in hadron-nucleus and nucleus-nucleus collisions. This is because causality dictates that the latest time that a correlation can be induced between two particles A and B that freeze-out is given by

$$\tau = \tau_{\text{freeze-out}} \exp\left(-\frac{|y_A - y_B|}{2}\right). \quad (126)$$

Thus two particles that are long-range in rapidity $|y_A - y_B| \gg 1$ would be correlated at very early times in the collision (Dumitru *et al.*, 2008a). A particular example is the so-called “ridge” effect, reviewed in (Dusling *et al.*, 2016) that correlates two particles not just in rapidity but also in relative azimuthal angle (Dumitru *et al.*, 2011a). A recent summary of the physics of initial state correlations can be found in Ref. (Altinoluk and Armesto, 2020).

However if hydrodynamic flow also sets in early, this ridge could be a final state effect (Shuryak and Zahed, 2013) due to the underlying boost-invariance of the hydrodynamic fluid. A way forward to disentangling initial state physics of CGCs and the Glasma at early times from late time dynamics is to look at the evolution of two-particle correlations with their rapidity separation (Bzdak and Dusling, 2016) Another is to study the long range correlations of particles with large transverse momenta that do not follow hydrodynamically (Dusling and Venugopalan, 2013; Martinez *et al.*, 2019).

C. Bulk observables

We discussed previously limiting fragmentation of hadron distributions and its potential to distinguish initial and final state effects in hadron-hadron collisions (Gonçalves *et al.*,

2019). We will now discuss other bulk observables in high energy nucleus-nucleus, hadron-nucleus and hadron-hadron collisions that can help constrain the properties of saturated gluons and their early-time evolution. In the smaller systems, even if the system hydrodynamizes quickly, the large shape fluctuations of partons will provide insight into multi-parton correlations in the initial state (Mäntysaari, 2020); understanding these from first principles is a challenging problem (Dumitru *et al.*, 2020) that may also require the EIC to resolve.

A number of works have explored applications of holographic ideas to the study of bulk observables in heavy-ion collisions. A universal prediction of holography is that of hydrodynamization being distinct from local thermalization. A specific phenomenological investigation implementing this idea used holographic boost invariant dynamics with transverse expansion as a successful model of preflow (van der Schee *et al.*, 2013). Another development is (van der Schee and Schenke, 2015), which treated the planar shockwave collisions discussed in Sec. VI.E.1 as an explicit model of initial state physics. While this study recovered qualitative features of soft particle spectra, the rapidity distribution of produced particles are too narrow relative to experimental data. It would be very interesting to explore more complicated holographic models of heavy-ion collisions and constrain them with experimental data.

In a thermalizing system, the loss of information of the initial conditions manifest itself as production of entropy. Therefore if the system locally thermalizes, and its flow is nearly isentropic, the measured number of particles probes the entropy produced during the non-equilibrium evolution of quark-gluon matter. The CGC framework accounts for the increase of particle multiplicity with increasing collision energy with the growth of the saturation scale Q_S (Albacete and Marquet, 2014). Recent calculations of entropy production in the equilibration processes using hydrodynamic attractors provides a quantitative relation between the energy deposition in the CGC picture and the final particle numbers (Giacalone *et al.*, 2019).

On the other hand, the energy of the observed particles depends on the work done during the whole expansion and therefore has different dependencies on the dynamics of the pre-equilibrium stage. Comparing these two robust experimental measurements (energies and multiplicities) already casts doubts on complete equilibration of QGP in peripheral nucleus-nucleus collisions (Giacalone *et al.*, 2019; Kurkela *et al.*, 2019c).

Many of the experimental signatures of QGP (strangeness enhancement, jet suppression,

flow harmonics,...) show a smooth dependence on system size from central to peripheral nucleus-nucleus collisions, proton-nucleus and proton-proton collisions. As the system size shrinks, so does its lifetime, corresponding to an increase in the relative importance of non-equilibrium QCD processes increases.

Equilibration studies in large systems already put a lower bound below which the system will not reach hydrodynamization or chemical equilibrium (Kurkela and Mazeliauskas, 2019a; Kurkela *et al.*, 2019b). Therefore explaining observed signals of collectivity (or absence thereof) in small collisions systems requires a proper treatment of non-equilibrium QCD dynamics. Some recent examples of work in this direction include studies of flow harmonics (Kurkela *et al.*, 2019d; Schenke *et al.*, 2020b), parton energy loss (Andres *et al.*, 2020) and heavy-quark evolution (Mrówczyński, 2018). Furthermore, as we discussed in sec. VI.E.2, hydrodynamization without equilibration, of small systems is very natural in holography.

Also noteworthy is recent phenomenological work (Huang *et al.*, 2018) quantifying the role of non-equilibrium dynamics in the Chiral Magnetic Effect we discussed in Section IV. A topic that demands further investigation is the origin of the very large vorticities measured in off-central heavy-ion collisions, as extracted from measurements of the polarization of Λ -baryons (Becattini and Lisa, 2020). The vorticities are introduced on macroscopic scales on the order of the system size; how these propagate efficiently down to the microscopic scales of the Λ remains to be understood.

D. Future prospects

A recent recommendation from the European Strategy for Particle Physics report emphasized that the main physics goal of future experiments with heavy-ion and proton beams at the LHC will be *a detailed, experimentally tested dynamical understanding of how out-of-equilibrium evolution occurs and equilibrium properties arise in a non-Abelian quantum field theory* (Citron *et al.*, 2019; Ellis *et al.*, 2019). The scheduled runs 3 and 4 of the LHC will mark a decade of high-statistics data across system sizes at the highest achievable collision energies.

In the United States, the continued operation of RHIC will provide further insight into several of the signatures we have discussed. In particular, with the anticipated commission-

ing of the sPHENIX detector (Roland, 2019), hard probes of QCD off-equilibrium will be studied in a dynamical range that is complementary to that of the LHC.

Looking further to the future, the Electron-Ion Collider (EIC) project has received Critical Mission Zero (CD0) approval from the US Department of Energy. The EIC will explore with high precision the landscape of hadron structure at high energies (Accardi *et al.*, 2016; Aschenauer *et al.*, 2019).

One may therefore anticipate that this decade and the next will bring many opportunities to exploit the signatures we have articulated here, and likely several novel ones, of the properties of QCD off-equilibrium.

VIII. INTERDISCIPLINARY CONNECTIONS

Understanding the thermalisation process in QCD associated to heavy-ion collisions addresses some of the most fundamental questions in quantum dynamics, with exciting interdisciplinary connections to very different many-body systems. The transient “fireball” expanding in vacuum explores far-from-equilibrium conditions at early times, followed by a series of characteristic stages which are finally expected to lead to a fluid-like behavior governing the approach to local thermal equilibrium. Very similar questions of equilibration and the emergence of collective behavior from the underlying unitary quantum dynamics are relevant for diverse applications ranging from high-energy and condensed matter physics to practical quantum technology. For reviews in the context of condensed matter physics, see (Borgonovi *et al.*, 2016; D’Alessio *et al.*, 2016; Gogolin and Eisert, 2016).

Several non-equilibrium phenomena were first proposed in the context of QCD matter in extreme conditions, and then explored and experimentally probed in alternative quantum many-body systems. For instance, the phenomenon of prethermalization (Berges *et al.*, 2004) with the rapid establishment of an effective equation of state during the early stages of heavy-ion collisions (Arnold *et al.*, 2005; Dusling *et al.*, 2012) has been explored for early-universe inflaton dynamics (Podolsky *et al.*, 2006), or condensed matter systems (Langen *et al.*, 2016; Moeckel and Kehrein, 2008; Mori *et al.*, 2018), and experimentally discovered in ultracold quantum gases on an atom chip (Smith *et al.*, 2013).

In turn, aspects of entanglement represent one of the major overarching schemes in contemporary physics of quantum-many body systems, and gravity in and out of equilibrium,

while investigations about its relevance to the thermalization process in QCD are comparably recent. There are many excellent topical reviews on entanglement and we refer the reader to Refs. (Calabrese and Cardy, 2009; Casini and Huerta, 2009; Eisert *et al.*, 2010; Rangamani and Takayanagi, 2017), while we discuss some aspects of entanglement in our context in more detail below.

To capture the thermalisation dynamics in QCD related to heavy-ion collisions, detailed comparisons take into account that the coupling of non-Abelian gauge theories is not a constant but changes with characteristic energy or momentum scale in a particular way. While strong at low scales, the coupling becomes weak at sufficiently high energies because of the phenomenon of asymptotic freedom (Gross and Wilczek, 1973; Politzer, 1973). Even in the high-energy limit, where the gauge coupling is weak, one is facing a strongly correlated system because a plasma of gluons with high occupancy $f(Q_S) \sim 1/\alpha_S(Q_S)$ is expected to form, see Sec. III. Such a transient over-occupation leading to strong correlations even for weakly coupled systems can be found in a variety of physical applications far from equilibrium. Examples include the pre-heating scenario for the very early stages of our universe after a period of strongly accelerated expansion called inflation (Kofman, 2008), or the relaxation dynamics in table-top setups with ultracold quantum gases following a sudden change in external control parameters such as magnetic fields (Prüfer *et al.*, 2018).

The very high level of control in experiments with synthetic quantum systems, such as ultracold quantum gases, enables dedicated quantum simulations. These systems provide very flexible testbeds, which can realize a wide range of Hamiltonians with variable interactions and degrees of freedom based on atomic, molecular and optical physics engineering (Bloch *et al.*, 2008). Since these setups can be well isolated from the environment, they offer the possibility of studying fundamental aspects such as the thermalisation process from the underlying unitary quantum evolution.

While digital quantum simulations based on a Trotterized time evolution on a universal quantum computer are challenging to scale up, present large scale analog quantum simulators using ultracold quantum gases already explore the many-body limit described by quantum field theory (Bernien *et al.*, 2017; Bloch *et al.*, 2008; Eckel *et al.*, 2018; Erne *et al.*, 2018a; Feng *et al.*, 2019; Gring *et al.*, 2012; Haller *et al.*, 2010; Hu *et al.*, 2019; Hung *et al.*, 2013; Keesling *et al.*, 2019; Langen *et al.*, 2015; Murthy *et al.*, 2019; Navon *et al.*, 2015, 2016; Parsons *et al.*, 2016; Prüfer *et al.*, 2019; Prüfer *et al.*, 2018; Schweigler *et al.*, 2017; Zache *et al.*,

2020). In principle, with quantum simulators also non-universal aspects of the dynamics of gauge theories can be studied. This has been first achieved for Abelian gauge theory with digital quantum simulations, such as using trapped ions (Martinez *et al.*, 2016) or with superconducting qubits (Klco *et al.*, 2018).

An interesting possibility to consider is applying a hybrid quantum-classical framework to real time problems. This has been discussed in a “single particle” digital strategy for scattering problems whereby higher loop quantum contributions can be simulated digitally and the background gauge field treated in principle on a quantum simulator (Mueller *et al.*, 2019, 2020). It is also important to note that scalable analog systems for the quantum simulations of gauge theories using ultracold atoms have been reported (Kokail *et al.*, 2019; Mil *et al.*, 2020). We anticipate significant progress in all of these approaches to quantum computation of real time problems in the decade ahead.

A. Strong interactions: Unitary Fermi gas

A paradigmatic example for the interdisciplinary cross-fertilization among the different physical applications is the work on collective motion of a unitary Fermi gas. Near unitarity, the s -wave scattering length, which characterises the two-body interaction strength, becomes very large and the effective scale invariance of the interaction at unitarity can lead to universal behavior (Chin *et al.*, 2010), which can also be accessed out of equilibrium (Eigen *et al.*, 2018). Many similarities for dynamical properties, such as a low ratio of shear viscosity to entropy density, have been discussed in this context in comparison to QCD. See the discussion in Sec. V.

We noted that heavy-ion experiments indicate that the hot quark-gluon plasma may be described as the most perfect fluid realized in nature (Aad *et al.*, 2012; Aamodt *et al.*, 2010; Adams *et al.*, 2005; Adcox *et al.*, 2005; Chatrchyan *et al.*, 2011). The only serious experimental competitors are ultracold quantum gases at temperatures that differ by twenty orders of magnitude! Strong interactions play also a central role in holographic approaches, which is addressed in Sec. VI, and there exist concrete proposals on how to realize holographically systems resembling unitary Fermi gases starting with Refs. (Balasubramanian and McGreevy, 2008; Son, 2008). A comprehensive review of common aspects of QCD, unitary Fermi gases and holography is provided by Ref. (Adams *et al.*, 2012).

B. Highly occupied systems I: Preheating in the early universe

The dilution of matter and radiation during the inflationary period of the early universe leads to an extreme condition, which may be well characterised by a pure state with vacuum-like energy density carried by a time dependent coherent (inflaton) field with large amplitude (Kofman, 2008). A wide class of post-inflationary models with weak couplings exhibit the subsequent decay of the inflaton field amplitude via non-equilibrium instabilities (Kofman *et al.*, 1994; Traschen and Brandenberger, 1990). The detailed mechanisms for the origin of an instability and the scattering processes are different than in QCD with strong color fields.

However the rapid growth of fluctuations from the inflaton decay leads to a non-linear time evolution that follows along similar lines as outlined in Sec. IV for QCD. For instance, for scalar fields with weak quartic interaction $\lambda \ll 1$, a corresponding overoccupation $\sim 1/\lambda$ up to a characteristic momentum scale is achieved after the instability. Likewise, at this stage, the prethermalization (Arnold *et al.*, 2005; Berges *et al.*, 2004) of characteristic properties, such as an effective equation of state, is observed in these scalar models (Podolsky *et al.*, 2006).

Moreover, a self-similar attractor solution is approached subsequently as discussed in Sec. IV.C.3. As compared to the longitudinally expanding QCD plasma, a major difference stems from the isotropic expansion of the universe. Some aspects of isotropic expansion can be lifted for the inflaton field dynamics by introducing suitably rescaled (conformal) time and field amplitudes, such that the dynamics is essentially that of Minkowski spacetime without expansion (Micha and Tkachev, 2003). In fact if compared to QCD dynamics without expansion, then characteristic dynamical properties such as the values of scaling exponents in the attractor regime agree with what is found for self-interacting scalar field dynamics with quartic interactions in the absence of spontaneous symmetry breaking (Berges and Wallisch, 2017).

This concerns both the gauge theory's direct energy cascade towards the perturbative high-momentum regime (Kurkela and Moore, 2011b, 2012; Schlichting, 2012), as well as the inverse particle cascade towards low momenta in the non-perturbative regime associated with non-thermal fixed points (Berges *et al.*, 2019). In turn, scalar fields with longitudinal expansion seem to exhibit several universal features shared with QCD dynamics in the

transient scaling regime (Berges *et al.*, 2015a). In particular, the inverse cascade essentially follows the behavior of the corresponding non-expanding system because of the strong Bose enhancement of rates at low momenta (Berges *et al.*, 2015a); see also Sec. IV.D.2.

C. Highly occupied systems II: Bose gases far from equilibrium

Though the inflaton dynamics is described by a relativistic field theory, the self-similar scaling behavior at sufficiently low momenta below the screening mass scale is predicted to exhibit universal properties of a non-relativistic system (Piñeiro Orioli *et al.*, 2015). The non-equilibrium infrared dynamics for scalars starting from overoccupation has been theoretically studied in great detail (Berges *et al.*, 2015b, 2008a; Berges and Sexty, 2011, 2012; Boguslavski and Piñeiro Orioli, 2019; Chantesana *et al.*, 2019; Deng *et al.*, 2018; Moore, 2016; Nowak *et al.*, 2012, 2011; Piñeiro Orioli and Berges, 2019; Scheppach *et al.*, 2010; Shen and Berges, 2020; Walz *et al.*, 2018). However important aspects of this far-from-equilibrium dynamics can be probed experimentally using Bose gases in an optical trap. For the example of an interacting, non-relativistic Bose gas of density n in three spatial dimensions, this concerns the dilute regime, $\sqrt{na^3} \ll 1$, with a characteristic inverse coherence length given by the momentum scale $Q = \sqrt{16\pi an}$. Here Q plays a similar role as the saturation scale for gluons in the gauge theory case, and the diluteness $\sqrt{na^3}$ provides the dimensionless coupling parameter. An overoccupied Bose gas then features large occupancies $\sim 1/\sqrt{na^3}$ for modes with momenta of order Q (Piñeiro Orioli *et al.*, 2015).

Universal scaling far from equilibrium associated with non-thermal fixed points has been experimentally discovered using different cold atom systems (Erne *et al.*, 2018a; Prüfer *et al.*, 2018). For instance, in Refs. (Prüfer *et al.*, 2019; Prüfer *et al.*, 2018) the non-equilibrium dynamics of magnetic hyperfine excitations of a spin one Bose gas is studied in an elongated trap, following a sudden change in the applied magnetic field as an external control parameter. Fig. 32 exemplifies the scaling dynamics of the measured transversal spin for three different initial conditions. After an initial non-equilibrium instability regime, all data in the self-similar scaling regime are seen to collapse to a single curve after rescaling with time using universal scaling exponents. While this example concerns infrared scaling, bi-directional scaling including a self-similar evolution towards higher momenta with subsequent thermalization has been experimentally analyzed in Ref. (Glidden *et al.*, 2020).

D. Highly occupied systems III: Classicalization and unitarization of gravitational amplitudes

An intriguing idea is that of black holes as long lived states of highly occupied gravitons ($f \gg 1$) that satisfy the condition $\alpha_{\text{gr}} f = 1$ (Dvali and Gomez, 2013). Here $\alpha_{\text{gr}} = L_P^2/R_S^2$, where L_P is the Planck length and R_S denotes the Schwarzschild radius. A dynamical picture of the formation of such a black hole state is in $2 \rightarrow N$ scattering of gravitons at trans-Planckian energies. In the Regge limit, as first discussed in Ref. (Lipatov, 1991), and subsequently in Ref. (Amati *et al.*, 1987), the scattering is dominated by the formation of $N - 2$ soft quanta. The argument of Dvali and collaborators is that the copious production of soft gravitons leads to perturbative unitarization of the scattering cross-section precisely when $\alpha_{\text{gr}} f = 1$.

This “classicalization of amplitudes” was shown explicitly (Dvali *et al.*, 2015) using the tree level Kawai-Lewellen-Tye (KLT) relations (Kawai *et al.*, 1986) that express N -point tree level gravity amplitudes in terms of sums of products of Yang-Mills N -point tree amplitudes. These results are in remarkable agreement with computations in Lipatov’s EFT approach (Addazi *et al.*, 2017).

The ideas of the classicalization and unitarization of $2 \rightarrow N$ gravitational amplitudes are remarkably similar to the discussion of the CGC EFT in Sec. II and Sec. III. The BFKL results on $2 \rightarrow N$ gluon scattering are likewise reproduced in the semi-classical CGC EFT. A path forward is to employ so-called “double copy” methods that exploit a color-kinematics duality between gravity and QCD amplitudes (Bern *et al.*, 2019). Such a correspondence was prefigured in the high energy limit in Ref. (Lipatov, 1991) and discussed further more recently (Liu, 2019; Sabio Vera *et al.*, 2012).

Of particular interest in our context is of a “classical double copy” between classical Yang-Mills equations and classical gravity (Goldberger and Ridgway, 2017; Monteiro *et al.*, 2014). This points to a concrete correspondence between collisions of the classical gluon shock waves producing the Glasma and that of gravitational shock waves that produce black holes (Dvali and Venugopalan, 2020). It would also be interesting to understand if this correspondence shares universal features at the unitarity limit with that of the holographic gravitational shock waves discussed in Sec. VI.

E. Anomalous currents in non-equilibrium QED: Condensed matter systems and strong laser fields

Strong color fields as well as strong electromagnetic fields are an essential ingredient for the understanding of the early stages of the plasma's space-time evolution in off-central heavy-ion collisions. Strong gauge fields lead to a wealth of intriguing phenomena related to quantum anomalies, such as the chiral magnetic effect (Kharzeev *et al.*, 2016; Koch *et al.*, 2017) described in Sec. IV. As we discussed there, there are strong connections between the transport properties of anomalous currents in hot QCD and in strongly correlated condensed matter systems, in particular Dirac/Weyl semimetals with applied fields (Li and Kharzeev, 2016).

Here we wish to note that the similar questions can also be addressed in future strong laser field experiments that will be able to explore QED dynamics in extreme conditions (Di Piazza *et al.*, 2012). For instance, for QED field strengths exceeding the Schwinger limit for pair production, a highly absorptive medium with quantum anomaly-induced dynamical refractive properties related to the chiral magnetic effect is predicted (Mueller *et al.*, 2016).

F. Thermalization and entanglement

While the time evolution of isolated quantum systems is unitary, relevant observables in non-equilibrium quantum field theory can approach thermal equilibrium values at sufficiently late times, without the need for any coarse-graining or reference to a reduced density operator. Thermalization in quantum field theory has been demonstrated for scalar quantum field theories in various spatial dimensions (Arrizabalaga *et al.*, 2005; Berges, 2002; Berges and Cox, 2001; Juchem *et al.*, 2004) and with fermions (Berges *et al.*, 2003; Shen *et al.*, 2020), see Ref. (Berges, 2004a) for an introductory review⁴⁸. In gauge theories at strong coupling, thermalization from unitary dynamics was observed using holographic approaches, as we discussed in Sec. VI.

It has been analyzed in detail how, in particular, locally defined quantities of isolated quantum many-body systems can exhibit thermal features (Deutsch, 1991; Rigol *et al.*, 2008; Srednicki, 1994). In such time-dependent processes, entanglement entropy of spatial

⁴⁸ For thermalization studies in classical-statistical field theories for given regularization, see Ref. (Aarts *et al.*, 2001).

subregions – the von Neumann entropy of spatially reduced density matrices – was seen to reach the value predicted by thermal states after exhibiting a period of growth, see, e.g., Refs. (Abajo-Arrastia *et al.*, 2010; Alba and Calabrese, 2017; Calabrese and Cardy, 2005; Cotler *et al.*, 2016; Kaufman *et al.*, 2016; Liu and Suh, 2014). Understanding why and how this happens has been an active sub-field of research in lattice systems, quantum field theory and holography.

Ref. (Berges *et al.*, 2018a,b) applied similar considerations to a model of e^+e^- collisions and pursued the idea to view entanglement as a source of an apparent thermal behavior seen in multiparticle production in such events as discussed in Refs. (Andronic *et al.*, 2009; Becattini, 1996). Recently an entanglement entropy measure devised for proton-proton collisions at the LHC has been argued to be consistent with the data; the latter is at variance with expectations from Monte-Carlo simulations (Tu *et al.*, 2020). In the same vein, Ref. (Ecker *et al.*, 2016) explored the behavior of the entanglement entropy in a holographic model of heavy-ion collisions discussed in Sec. VI.E and found it can serve as an order parameter distinguishing between the Landau (full stopping) and Bjorken (transparency) scenarios.

The notion of entanglement plays the key role in tensor networks methods which represent quantum-many body wave functions and density matrices of physical interest yet low enough entanglement allowing for their efficient manipulation on classical computers, see Ref. (Orus, 2014) for a review. Such methods are robust in describing ground states and low-lying excited states in 1+1 dimensions (Hastings, 2006; Vidal, 2008) and considerable progress has been made in the past few years on using them for condensed-matter physics applications in 2+1 dimensions (Corboz, 2016a,b; Corboz *et al.*, 2018; Rader and Läuchli, 2018; Vanderstraeten *et al.*, 2016).

In the context of the present review, we want to highlight a number of recent developments in applying tensor networks to QCD and heavy-ion collision motivated problems in (1+1)-dimensional settings ranging from the applications to gauge theories reviewed in Ref. (Bañuls and Cichy, 2020) to non-equilibrium processes in interacting QFTs on a lattice (Banuls *et al.*, 2020; Buyens *et al.*, 2017, 2016; Pichler *et al.*, 2016). In the latter cases, the aforementioned growth of entanglement with time is a bottle neck for simulations being able to reach late times.

Finally, entanglement entropy in holography arises as a Bekenstein-Hawking entropy of a

special class of surfaces (Dong *et al.*, 2016; Hubeny *et al.*, 2007; Lewkowycz and Maldacena, 2013; Ryu and Takayanagi, 2006). This discovery has led to new insight into quantum gravity by bringing quantum information tools to the mix. An impressive results in this direction is the quantitative understanding of the time evolution of the entropy of Hawking radiation from an evaporating black hole (Almheiri *et al.*, 2019, 2020a,b; Penington, 2019; Penington *et al.*, 2019). These works point to a new mechanism towards resolving Hawking’s information paradox (Hawking, 1976; Page, 1993). From the point of view of the present review, they can be thought of as including finite- N_c effects in holographic studies of a class of thermalization processes at very late times.

IX. SUMMARY AND OUTLOOK

In 1974, T.D. Lee suggested that *it would be interesting to explore new phenomena by distributing a high amount of energy or high nuclear density over relatively large volume* (Baym *et al.*, 1975). Forty six years later we are beginning to come to grips with the richness of many-body QCD dynamics, made possible by experimental programs in nucleus-nucleus collisions in the decades since, culminating in the discovery of the quark-gluon plasma at RHIC and the LHC. As demonstrated at these colliders, the non-Abelian QGP is a nearly perfect fluid showing little resistance to pressure gradients.

This conclusion is a consequence of the remarkable and apparently unreasonable success of relativistic viscous hydrodynamics in the description of the heavy-ion data from RHIC and LHC. However the quantitative phenomenological success of hydrodynamical models also owes a great deal to our improved understanding of the initial conditions for hydrodynamic evolution, in particular in the modeling of event-by-event fluctuations in the nuclear geometry, as well as a deepening understanding of how the quark-gluon matter is released in the heavy-ion collisions and thermalizes to form the QGP.

With regard to the latter, comparisons of the hydrodynamical models to data require that thermalization occurs very rapidly on time scales on the order of three yoctoseconds – approximately a tenth of the lifetime of the nuclear collision. These very short lifetimes and the nearly perfect fluidity of the subsequent flow of the QGP suggest that the non-equilibrium matter formed is very strongly correlated. The quest to understand *ab initio* the structure of strongly correlated QCD matter in nuclear wavefunctions at high energies,

and how this matter is released, decoheres, and thermalizes, has motivated a large body of work over the last couple of decades, right from the inception of the RHIC program to the present.

Strongly correlated QCD matter can arise either in weak coupling when the occupancies of the constituents are very large, or in strong coupling. Further, since the coupling runs towards strong coupling as the system evolves, both weak and strongly couplings may be realized in the fluid. In this review, we have summarized the theoretical ideas and techniques in both strong and weak coupling frameworks that address the thermalization process in heavy-ion collisions.

We emphasized the emergence of attractors in both the weak coupling EFT and in the holographic approaches that may be universal across a wide range of energy scales. We also noted concomitantly the very concrete interdisciplinary connections of strongly correlated QCD (and QCD-like) matter off-equilibrium to dynamical features of phenomena ranging from pre-heating in inflationary cosmology, pair-production in laser induced strong QED fields, and to non-equilibrium dynamics in ultracold atomic gases.

In particular, we discussed an intriguing universality in the non-thermal attractor discovered in simulations of overoccupied expanding Glasma to that discovered in identically prepared simulations of the self-interacting scalar fields that model the ultracold systems. Remarkably, cold atom experiments have discovered such a non-thermal attractor, albeit with a different geometry than that of a heavy-ion collision. This opens up the exciting prospect of extending the program underway of the “tabletop engineering” of ultracold atom systems as analog quantum simulators of the ground state properties of gauge theories to uncover far-from-equilibrium properties of non-Abelian gauge theories.

We also discussed the signatures for QCD matter off-equilibrium and the challenges of disentangling these from contributions at later stages of the heavy-ion collision. On-going and near-term experiments at both RHIC and the LHC will greatly enhance these prospects both through novel measurements and larger data sets than present ones. The EIC will provide information complementary to those of the heavy-ion experiments to further tease out and make more precise our understanding of the initial state. Further progress will also depend on theoretical developments in the weak and strong coupling frameworks and the convergence between the two when extrapolated to the realistic couplings of the heavy-ion experiments.

Computations of the properties of saturated gluons in the CGC EFT are now at next-to-leading-order and next-to-leading log accuracy for a few processes. We expect this trend to continue, which will allow for very precise extractions of the saturation scale in DIS and proton-nucleus collisions. A more conceptual challenging problem is to understand the large fluctuations in the large x initial conditions that may generate very anisotropic shape distributions of small x partons. As we noted briefly, such studies may benefit from the universality between the non-linear equations that describe high energy QCD evolution and those that describe reaction diffusion processes in statistical mechanics.

In the description of the Glasma, a straightforward but technically challenging problem is to extend several of the computations in fixed box geometries to the more realistic longitudinally expanding case. A more difficult challenge is to implement fully quantum contributions beyond the classical-statistical approximation. While there is considerable insight gained from on-going studies of scalar field theories in this regard, further progress will require further conceptual breakthroughs. A noteworthy feature of the overoccupied Glasma is the emergence of infrared structures that may have non-trivial topological features ([Spitz *et al.*, 2020](#)). This may be universal to other many-body systems leading to novel potential synergies in addition to those discussed in this review.

Recent numerical simulations using QCD effective kinetic theory have painted a detailed picture of the different equilibration stages in longitudinally expanding, albeit homogeneous, QCD matter. However the kinetic description of inhomogeneous systems with rapid radial expansion needs further development. This is especially important for studies of collisions of light nuclei or in proton-nucleus collisions, where tantalizing signals of collective behavior have been seen. It will be interesting within this framework to understand whether a unified many-body description emerges which smoothly interpolates from a few parton scatterings in the smallest collision systems to the emergent fluid-like behavior in the largest systems.

On the more formal side, computations of various transport properties of the QGP beyond leading order have higher order corrections that are large for all but extremely small values of the coupling constant. Finite temperature resummation techniques may help improve the convergence of the perturbative expansion. A potential path forward is to combine a non-perturbative description of the infrared sector with kinetic theory in the UV.

A key part of our review was devoted to developments in holographic approaches to off-equilibrium dynamics in QCD like theories. An important discovery is that the hydrody-

dynamic gradient expansion is an asymptotic series, which allows one to view the applicability of hydrodynamics through the emergent universal behavior of a hydrodynamic attractor.

An open problem is the existence of hydrodynamic attractors for flows with transverse expansion and/or broken conformal symmetry. It would be very interesting to make a clear-cut statement to what extent these phenomena appear in a tractable manner outside idealizations of the geometry of ultrarelativistic heavy-ion collisions or highly-symmetric cosmologies. Another important future direction is to address collisions in holographic models that incorporate confinement following recent promising work in this direction. Not least, it would be interesting to reconsider expanding plasma setups and, more broadly, thermalization at strong coupling in the context of Gauss-Bonnet gravity discussed in Section VI.F.2. First steps in this direction relied on treating the Gauss-Bonnet term as a small correction. Going beyond this regime, which challenging from many perspectives, can reveal genuinely new effects in holographic setups like non-thermal fixed points discussed in Section IV.C. Finally, an important open question in holography is to understand if long-range “ridge-like” correlations can naturally arise at strong coupling and whether they can survive till late time.

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FIGURES

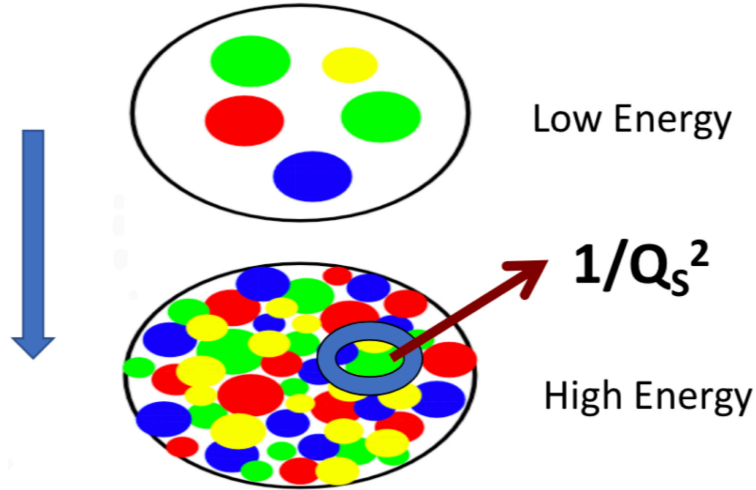


FIG. 1 Transverse hadron profile resolved in scattering with fixed squared momentum transfer Q^2 and increasing center-of-mass energy \sqrt{s} . The requirement that proliferating soft gluons have maximal occupancy $1/\alpha_S$ generates the close packing saturation scale Q_S . Figure adapted from (Iancu and Venugopalan, 2003).

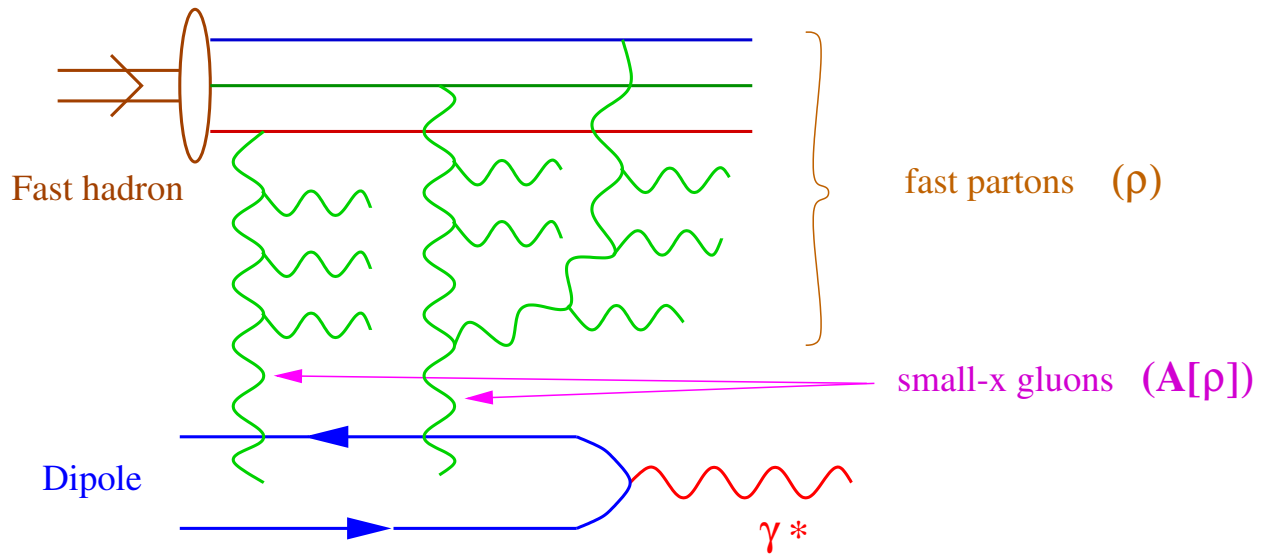


FIG. 2 DIS in the dipole picture. The virtual photon emitted by the electron splits into a $q\bar{q}$ dipole which scatters off dynamical small x gauge fields coupled to the static large x lightcone sources. Figure from (Iancu and Venugopalan, 2003).

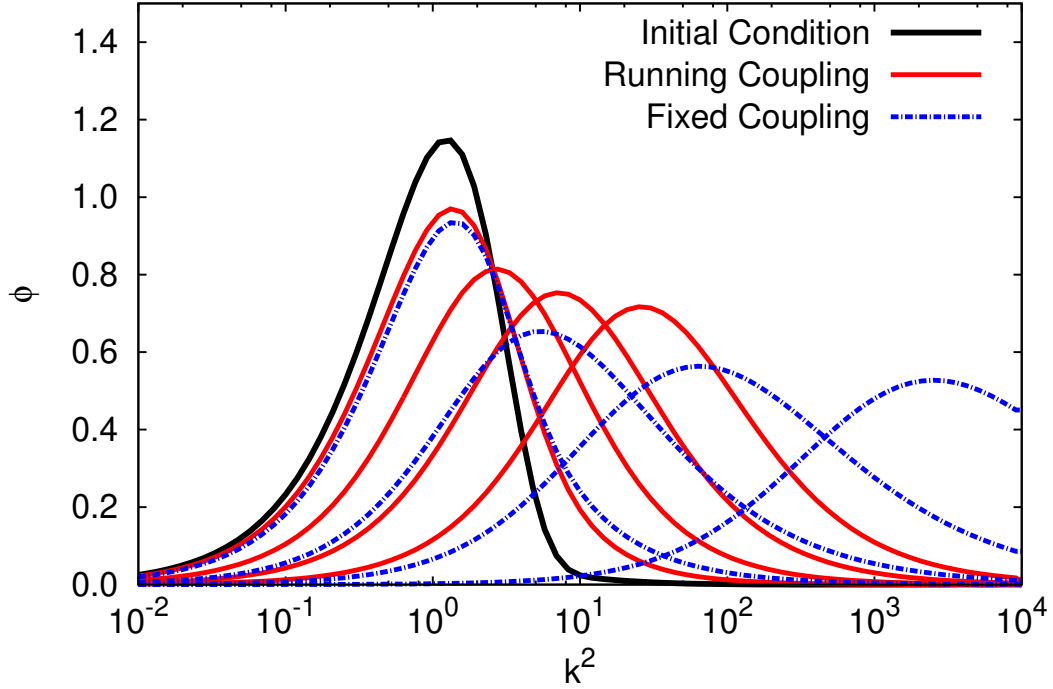


FIG. 3 The unintegrated gluon distribution (in units of the transverse area) versus squared transverse momentum (normalized to its value at the peak of the initial condition curve) from the solution of the Balitsky-Kovchegov equation. The different curves represent increasing rapidities (left to right) for fixed and running coupling. Figure from (Dusling *et al.*, 2010).

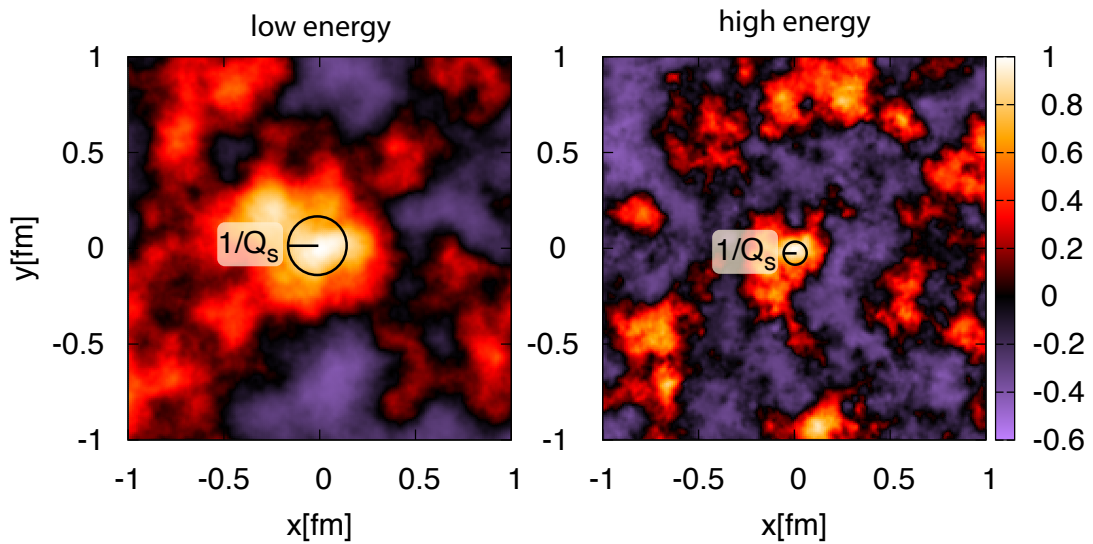


FIG. 4 Solution of JIMWLK equation for the correlator of Wilson lines $V(x_\perp)V^\dagger(y_\perp)$ probed by the DIS dipole (Dumitru *et al.*, 2011b). As the nucleus is boosted from low energy (or rapidity) to high energy, the regions with large values of these correlator shrink spatially, corresponding to larger values of Q_S .

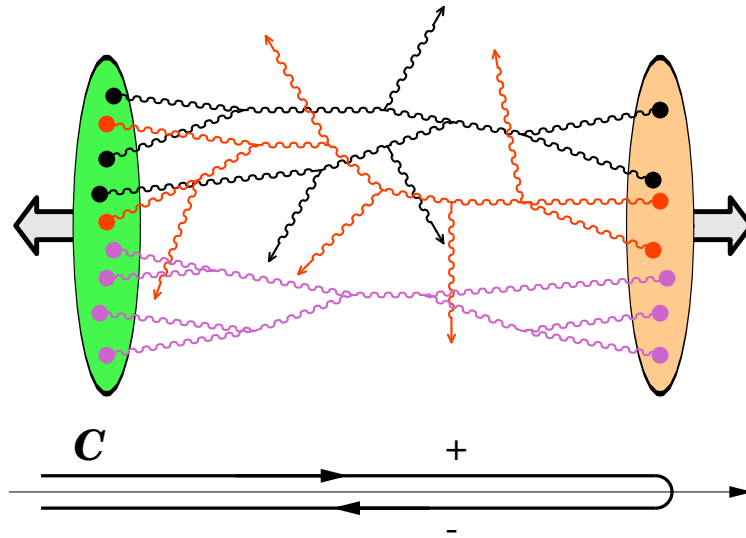


FIG. 5 a) Multi-particle production from cut “vacuum-vacuum” graphs connecting time dependent sources of the two nuclei after the collision. Figure from (Gelis *et al.*, 2010). b) The Schwinger-Keldysh closed time contour on which the sources and fields are defined.

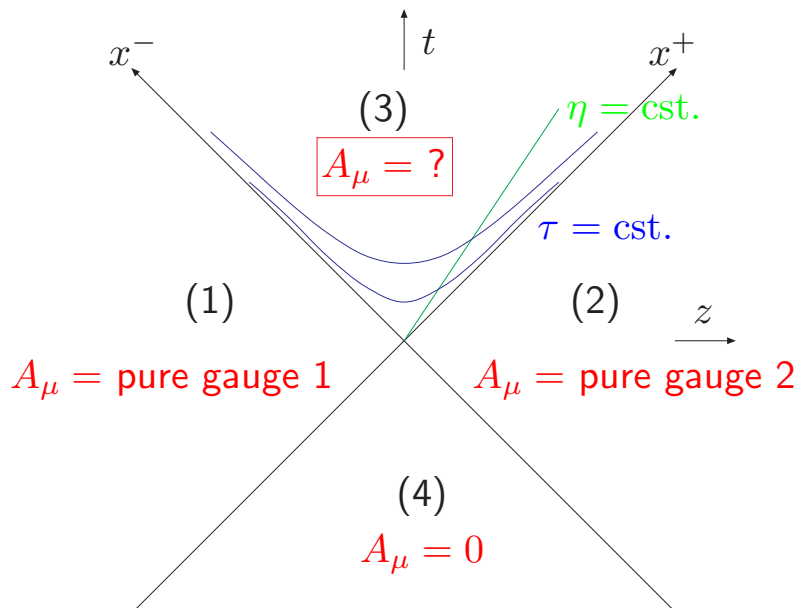


FIG. 6 Spacetime diagram of gauge field configurations. Before the collision, the gauge fields are pure gauge solutions with zero field strength. (In the text, the pure gauge solution of the right moving nucleus is denoted by $A_{1,\text{cl.}}^i$ and that of the left moving nucleus by $A_{2,\text{cl.}}^i$.) After the collision, the gauge field solution ($\mathcal{A}^{i,\eta}$ in text) correspond to finite field strengths in the Glasma. Figure from (Lappi and McLerran, 2006).

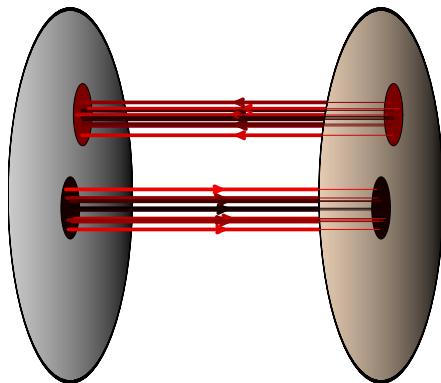


FIG. 7 Glasma flux tubes: Boost invariant LO Glasma configurations of transverse size $1/Q_S$ at $\tau = 0^+$ with parallel E_η and B_η , corresponding to finite Chern-Simons charge. Such configurations decay rapidly and are unstable to quantum fluctuations. Figure from (Dumitru *et al.*, 2008a).

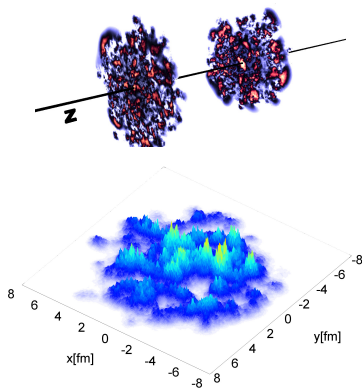


FIG. 8 a) Collisions of nuclei with sub-nucleon color charge fluctuations determined by the IP-Sat model. b) The LO energy density in the Glasma at $\tau = 1/Q_S$. Figures from (Schenke *et al.*, 2012a).

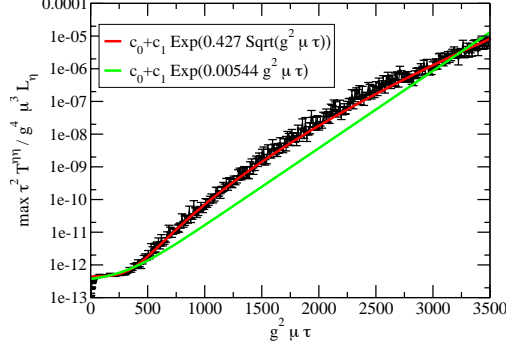


FIG. 9 Growth of the maximally unstable Fourier mode of the longitudinal pressure $P_L = \tau^2 T^{\eta m}$. Note that since $g^2 \mu \propto Q_S$, the results are in units of Q_S^3/g^2 , with $g \sim 10^{-5}$ and $L_{\eta} = 1.6$. Figure from (Romatschke and Venugopalan, 2006a).

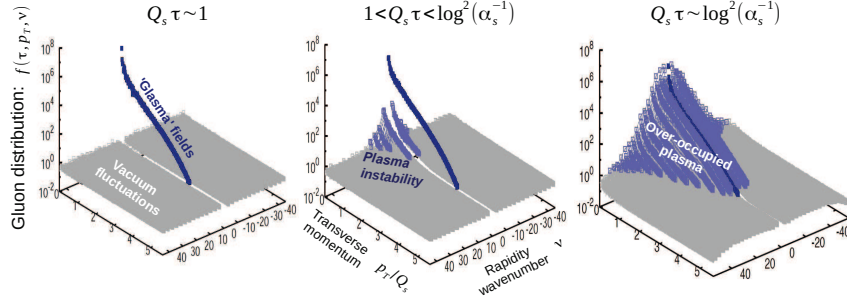


FIG. 10 Time evolution of the gluon distribution at early times $0 \lesssim Q_S \tau \lesssim \log^2(\alpha_S^{-1})$ from next-to-leading order CGC initial conditions (Epelbaum and Gelis, 2013) at very weak coupling ($\alpha_S \sim 10^{-6}$). Fig. taken from Ref. (Berges *et al.*, 2014e).

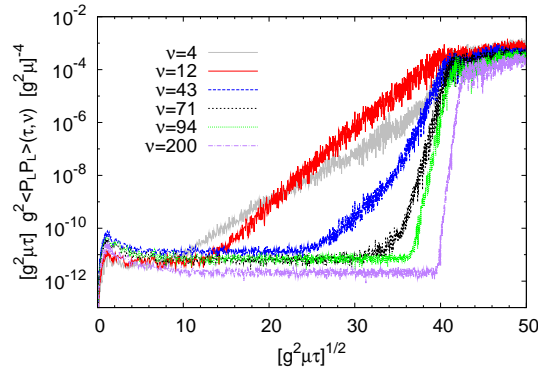


FIG. 11 Time evolution of the longitudinal pressure-pressure correlator for different rapidity wave numbers ν with parameters as described in Ref. (Berges and Schlichting, 2013). Once the initial fluctuations have grown larger, one observes the emergence of secondary instabilities at larger ν with enhanced growth rates.

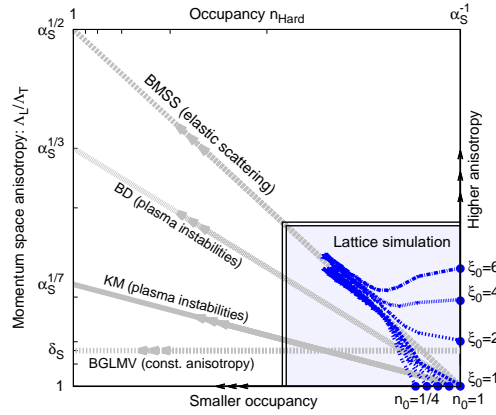


FIG. 12 Evolution in the occupancy–anisotropy plane, taken from Ref. (Berges *et al.*, 2014b). Indicated are the thermalization scenarios proposed in (BMSS) (Baier *et al.*, 2001), (BD) (Bodeker, 2005), (KM) (Kurkela and Moore, 2011a) and (BGLMV) (Blaizot *et al.*, 2012). The blue lines show the results of classical-statistical simulations for different initial conditions (Berges *et al.*, 2014b,e).

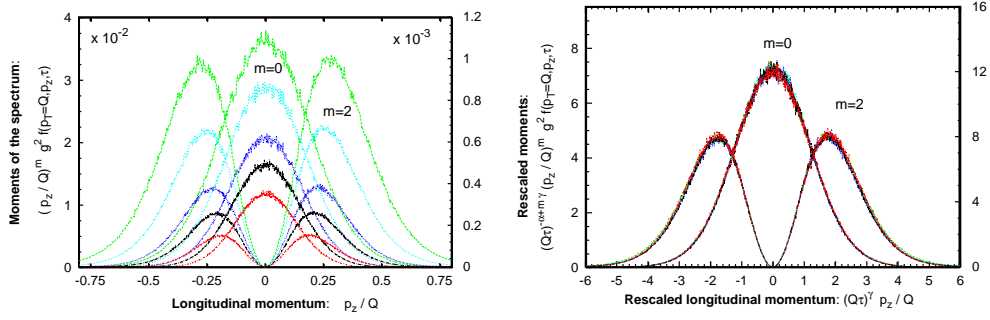


FIG. 13 (color online) (**top**) Moments of the single particle distribution function as a function of longitudinal momenta. The longitudinal spectra are evaluated at transverse momentum $p \simeq Q$. The different curves correspond to different times $Q\tau = 750, 1000, 1500, 2000, 3000$ (top to bottom) of the evolution. (**bottom**) The rescaled moments of the distribution function are found to collapse onto a single curve when plotted as a function of the rescaled longitudinal momentum variable. Fig. taken from Ref. (Berges *et al.*, 2014b).

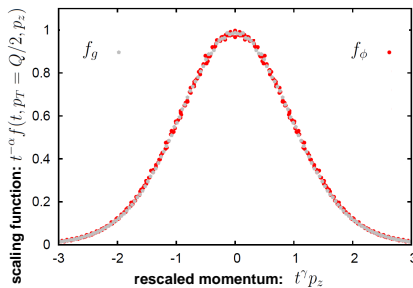


FIG. 14 The normalized distribution for the scalar theory (f_ϕ) as a function of the rescaled longitudinal momentum at different times in the self-similar regime compared to the gauge theory (f_g) (Berges *et al.*, 2015b).

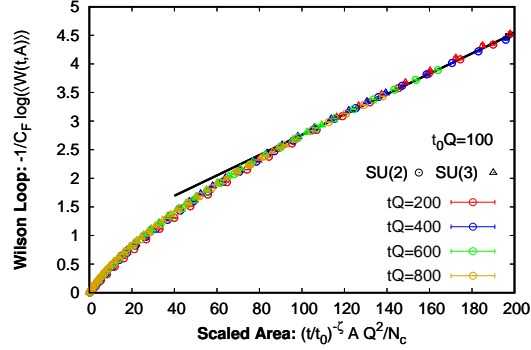


FIG. 15 Self-similar behavior of the spatial Wilson loop as a function of the time-rescaled area $\sim t^{-\zeta}A$ with universal scaling exponent ζ for gauge groups $SU(N_c)$ with $N_c = 2$ (circles) and $N_c = 3$ (triangles) (Berges *et al.*, 2017a).

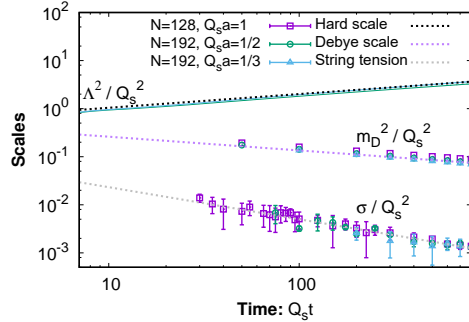


FIG. 16 Time evolution of the hard scale (Λ^2), the electric screening scale (m_D^2), and the spatial string tension (σ). Symbols for m_D^2 and σ denote lattice results for different lattice spacings employed, while the solid green line represents a continuum extrapolation for the hard scale. The dotted lines represent perturbative estimates for the hard and Debye scales, and a fit to the lattice results for the non-perturbative string tension scale (Mace *et al.*, 2016).

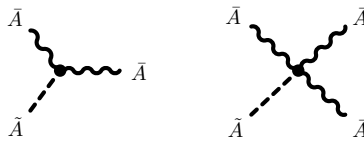


FIG. 17 Illustration of rescaled “classical” three- and four-vertices, which are independent of the coupling. Fig. taken from Ref. (Kasper *et al.*, 2014).

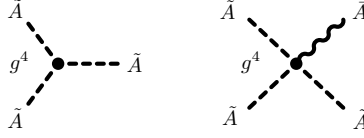


FIG. 18 Illustration of rescaled “quantum” three- and four-vertices, which are $\sim g^4$. Fig. taken from Ref. (Kasper *et al.*, 2014).

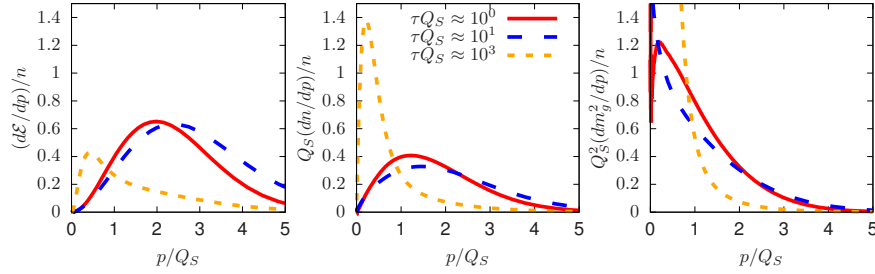


FIG. 19 Momentum differentiated plots of the (a) energy density (b) number density and (c) screening-mass at different stages of bottom-up thermalization. All curves are normalized by the instantaneous number density n and lines correspond to different times $\tau Q_S = 10^0, 10^1, 10^3$.

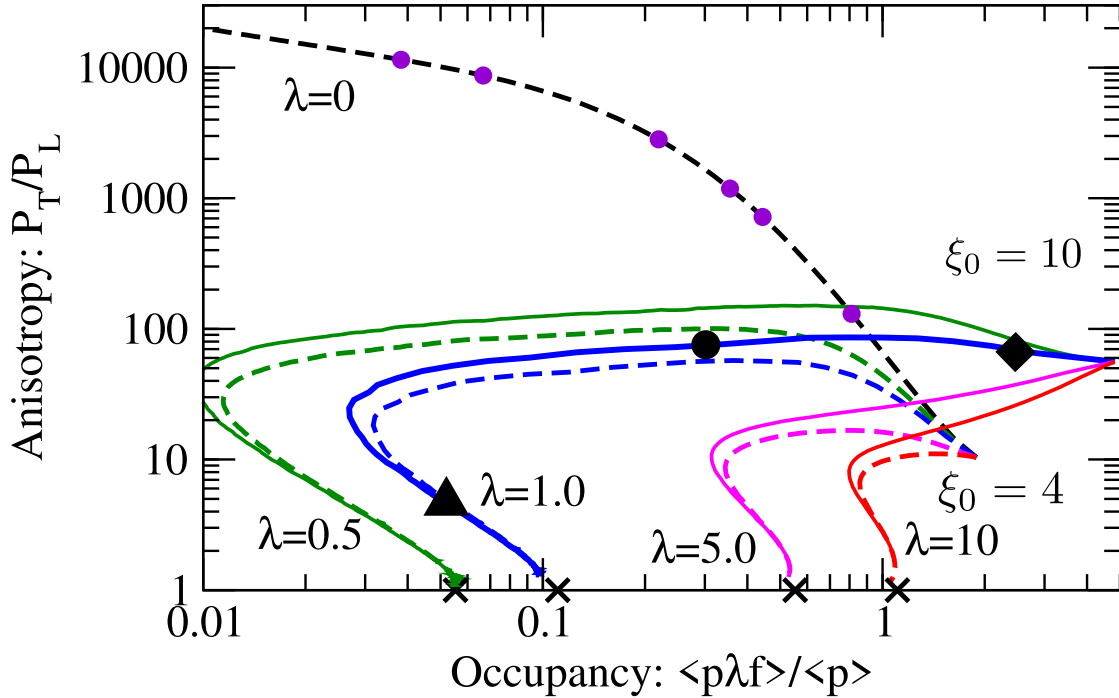


FIG. 20 The gluon kinetic theory equilibration in anisotropy-occupancy plane for initial anisotropy $\xi_0 = 10$ and different values of the coupling constant. Times corresponding to $\tau Q_S = 10^0, 10^1, 10^3$ are indicated by black symbols. Simulations with smaller initial anisotropy $\xi_0 = 4$ is shown by dashed curves. Figure adapted from (Kurkela and Zhu, 2015).

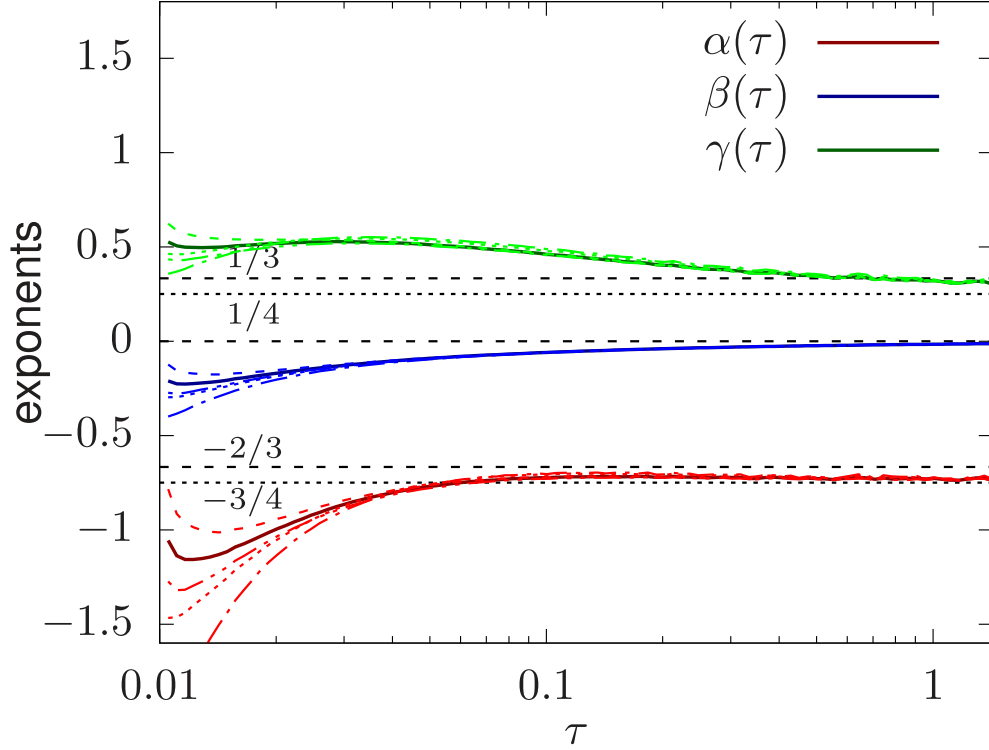


FIG. 21 The time evolution of instantaneous scaling exponents extracted from different sets of integral moments of the distribution. Horizontal lines indicate possible asymptotic values. Figure is taken from (Mazeliauskas and Berges, 2019).

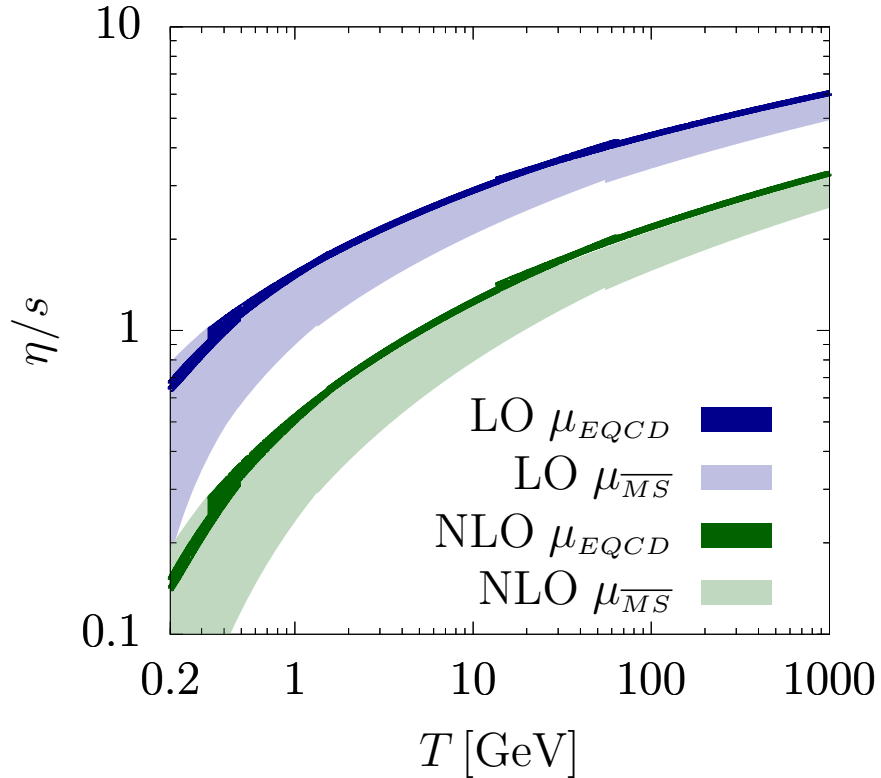


FIG. 22 The shear viscosity over entropy ratio as a function of temperature at leading (LO) and (nearly) next to leading order (NLO) thermal QCD. The bands correspond to the scale variation of running coupling prescriptions. Figure taken from (Ghiglieri *et al.*, 2018a).

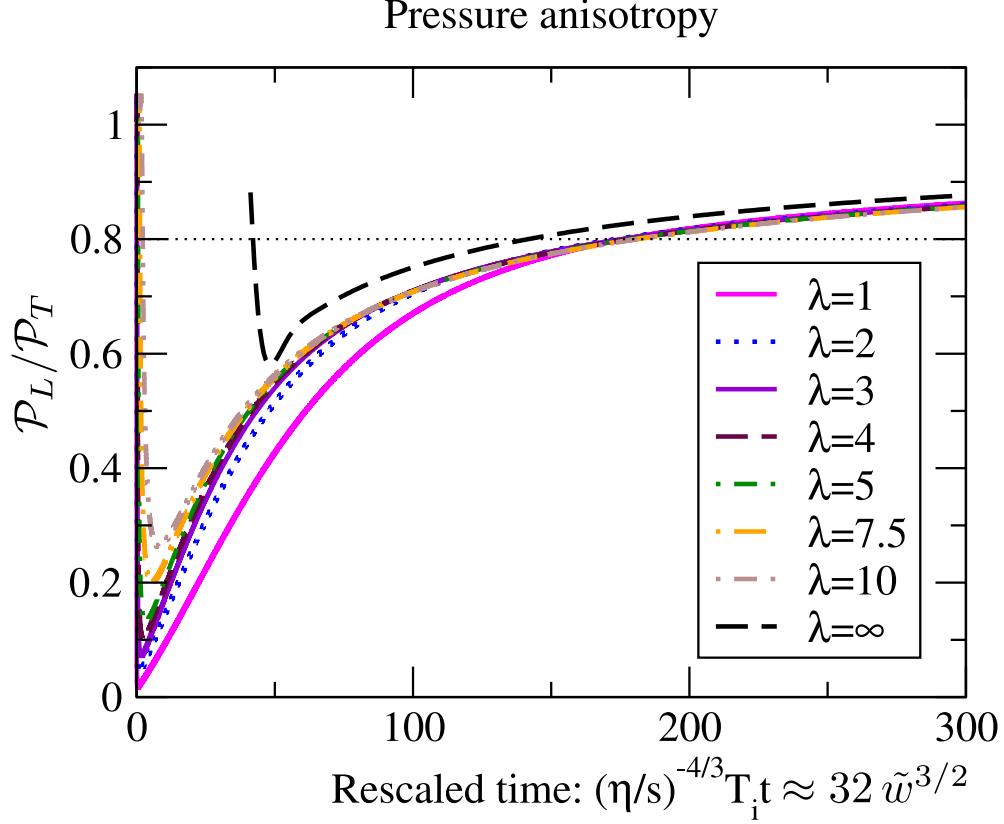


FIG. 23 The pressure anisotropy evolution in expanding geometry. Gluon kinetic theory simulations $\lambda = 1, \dots, 10$ are compared to a holographic model of supersymmetric Yang-Mills ($\lambda = \infty$). Note that here T_i is initial temperature and so at late times $(\eta/s)^{4/3} T_i t \approx 32 \tilde{w}^{3/2}$. Figure adapted from (Keegan *et al.*, 2016b).

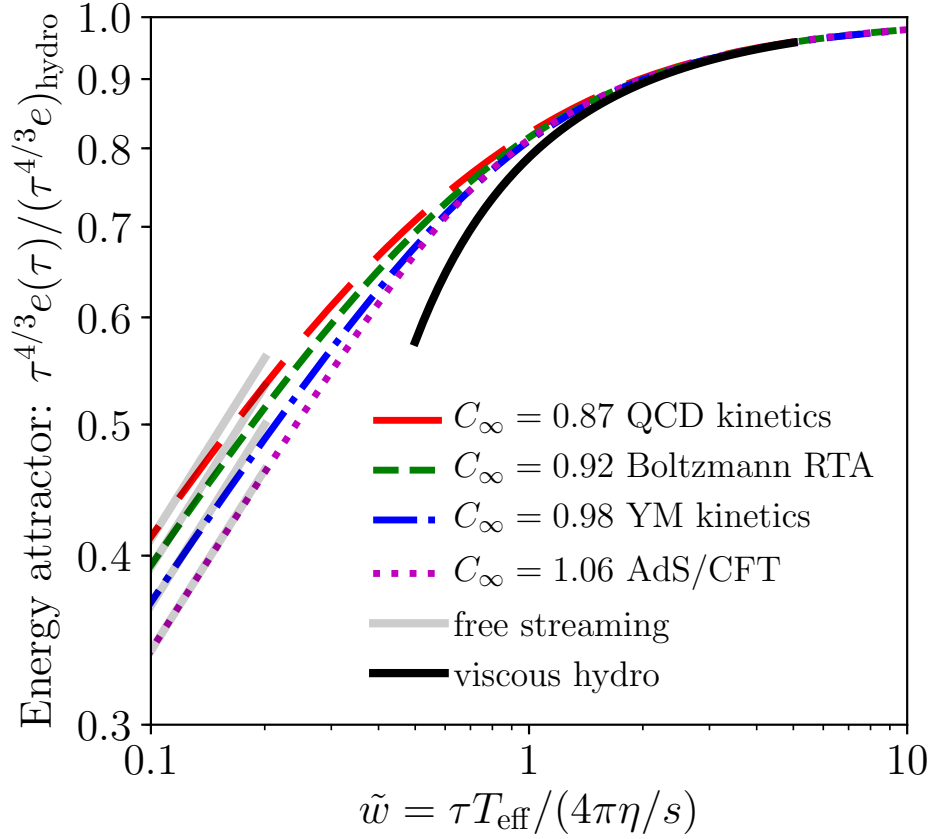


FIG. 24 Hydrodynamic attractors for pre-equilibrium evolution of energy density for different microscopic theories. Figure taken from (Giacalone *et al.*, 2019).

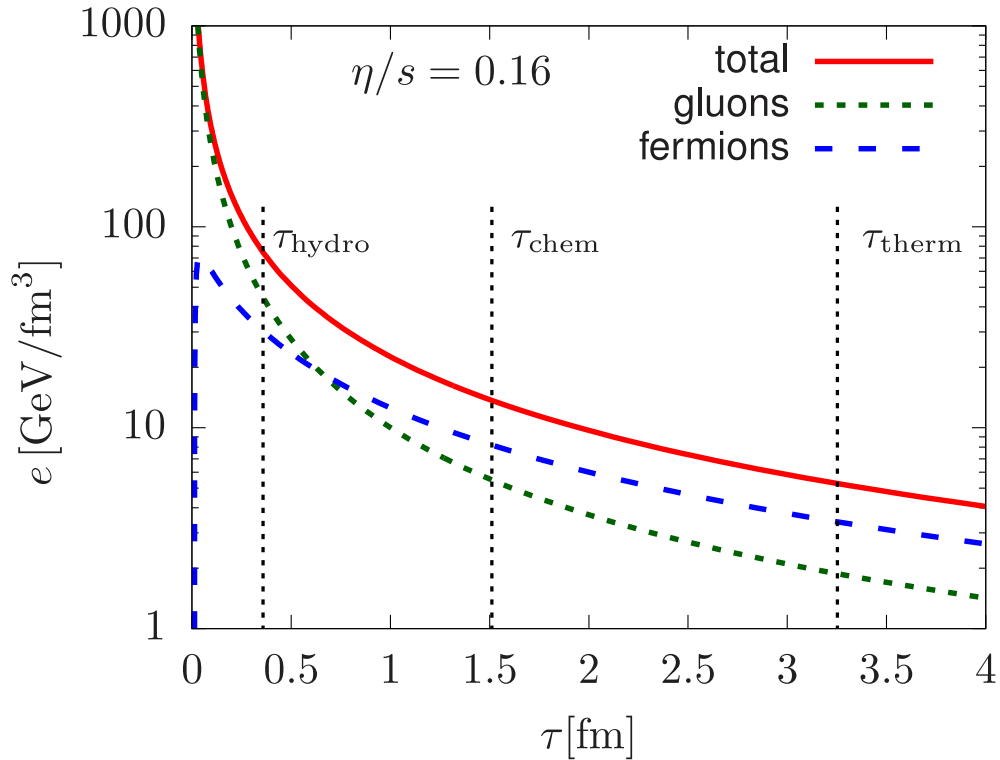


FIG. 25 The energy density evolution in a chemically equilibrating quark-gluon plasma. The vertical lines indicate the times of approximate hydrodynamic, chemical and thermal equilibriums. Figure taken from (Kurkela and Mazeliauskas, 2019a).

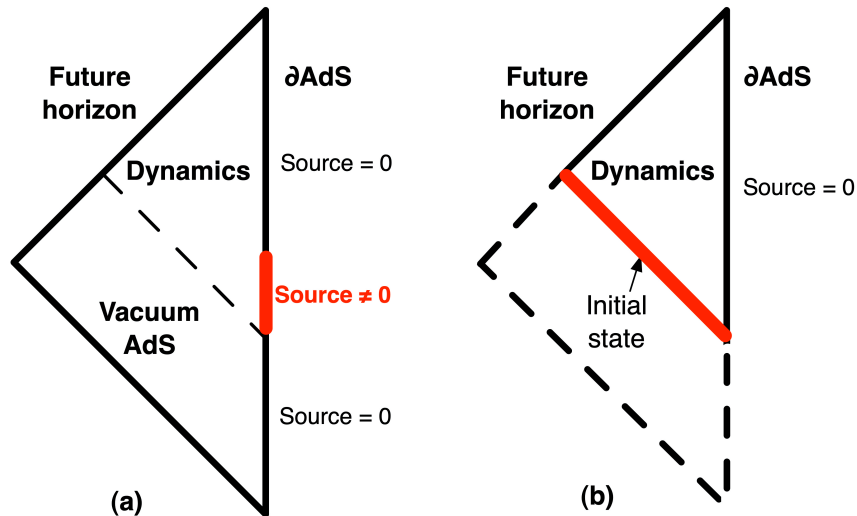


FIG. 26 Penrose diagrams dual to far-from-equilibrium states in strongly coupled QFTs. Left plot: The system starts in the vacuum with known bulk geometry and is perturbed by a non-trivial source, which appears as an asymptotic boundary condition in gravity. After the source is turned off, the QFT is in a non-equilibrium state modeled by a time dependent geometry. Right plot: The sources are always off, but one instead specifies non-trivial initial conditions for the bulk metric. Figure adapted from Ref. (Heller *et al.*, 2012c).

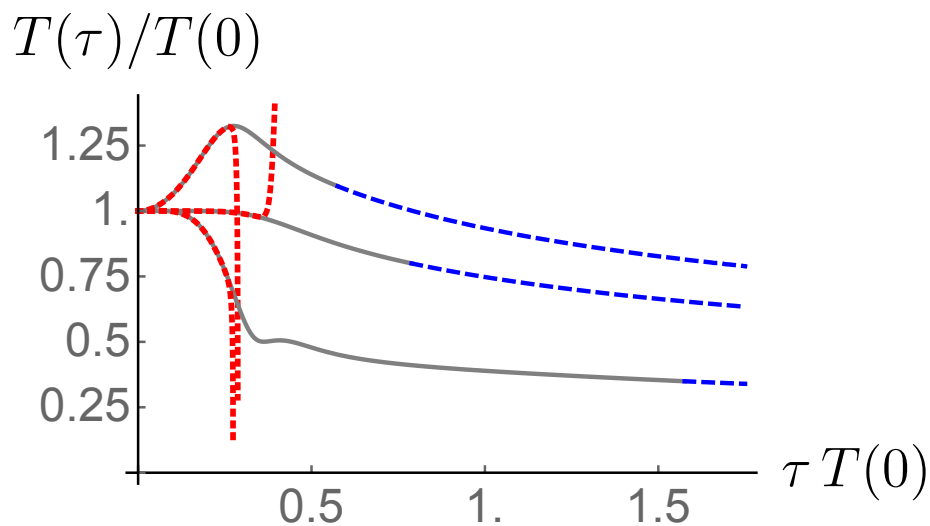


FIG. 27 Evolution of the effective temperature as a function of time for 3 different states with nonzero initial energy density. The gray curves denote the far-from-equilibrium regime. The blue dashed parts extending indefinitely to the right mark the applicability of viscous hydrodynamic relations truncated at the third order in derivatives (122). The red dotted curves denote the series in Eq. (120) extracted using Ref. (Beuf *et al.*, 2009). The figure is adapted from Ref. (Heller *et al.*, 2012a).

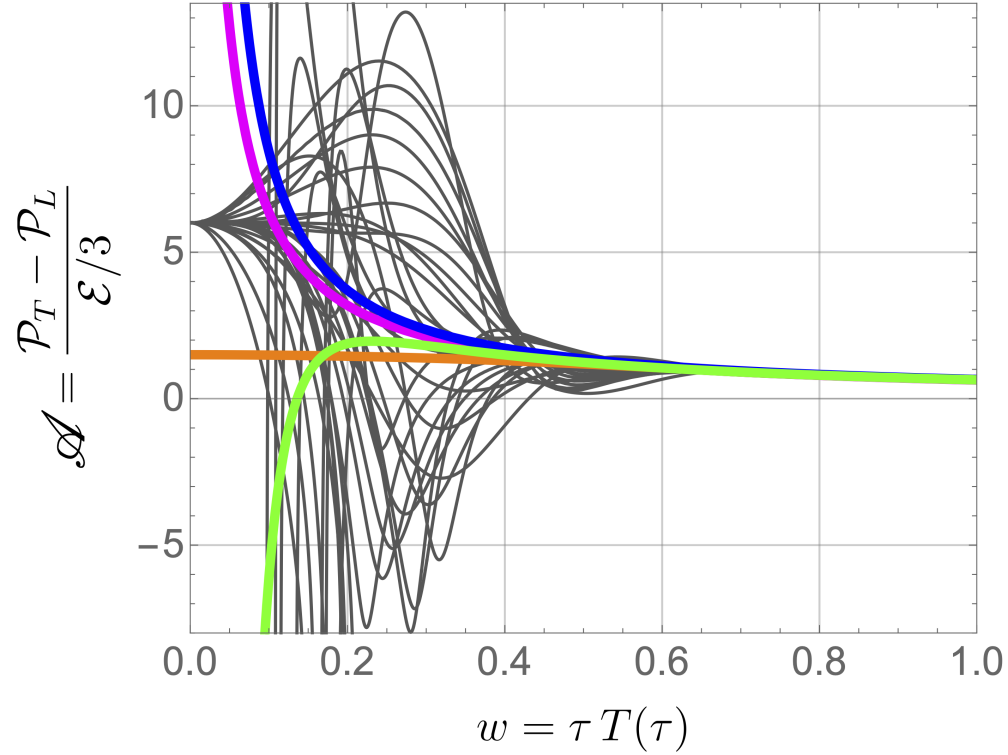


FIG. 28 $\langle T^{\mu\nu} \rangle$ in a holographic CFT as a function of the dimensionless clock variable w for 29 different initial states (grey curves). Magenta, blue and green curves denote predictions of hydrodynamic constitutive relations truncated respectively at first, second and third order (122). The orange curve is the hydrodynamic attractor (Romatschke, 2018). The figure is adapted from Refs. (Heller *et al.*, 2012a,b; Romatschke, 2018).

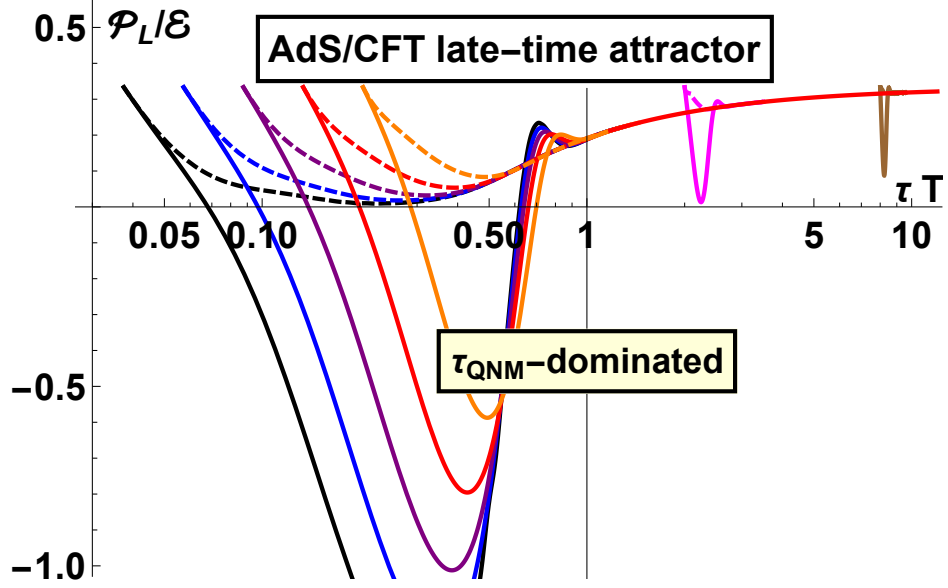


FIG. 29 Hydrodynamization of states whose gravity dual has initially support close to the boundary (dashed curves) or deep in the bulk (solid curves) initialized at different times (different colors), see text for details. The figure is adapted from Ref. (Kurkela *et al.*, 2020).

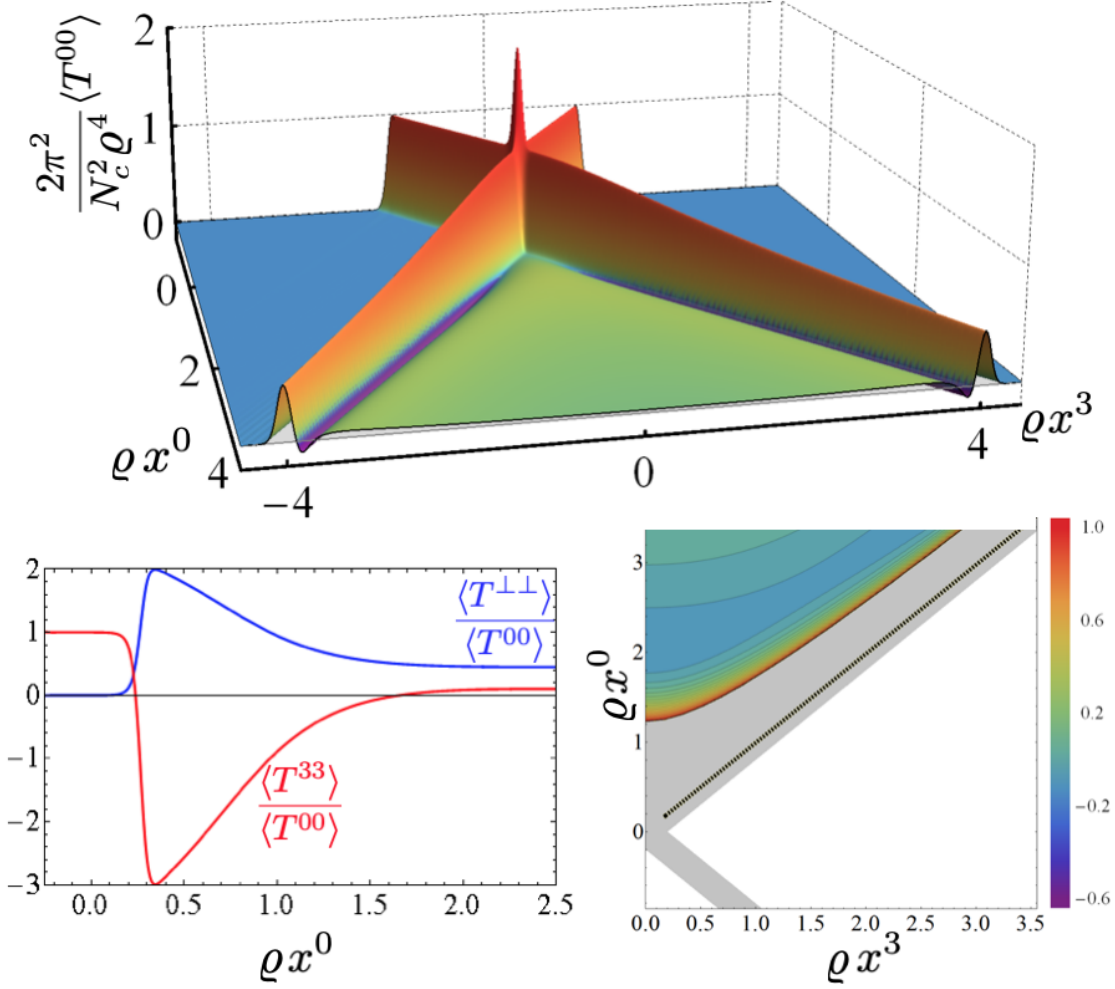


FIG. 30 $\langle T^{\mu\nu} \rangle$ resulting from a collision of thin planar shocks with $\varrho d = 0.08$ (124). TOP: Lab-frame energy density as a function of time x^0 and longitudinal position x^3 . Between the remnants and the central rapidity region, there are small regions of negative energy density. BOTTOM-LEFT: At mid-rapidity, the transverse and longitudinal pressure after the collision are consistent with $\langle T^{00} \rangle \sim \tau^2$ in Eq. (120). BOTTOM-RIGHT: The color encoding denotes deviations from constitutive relations and points to the applicability of hydrodynamics. The post-collision $\langle T^{\mu\nu} \rangle$ does not have a rest frame in the gray region (Arnold *et al.*, 2014). Plots adapted from Ref. (Casalderrey-Solana *et al.*, 2013).

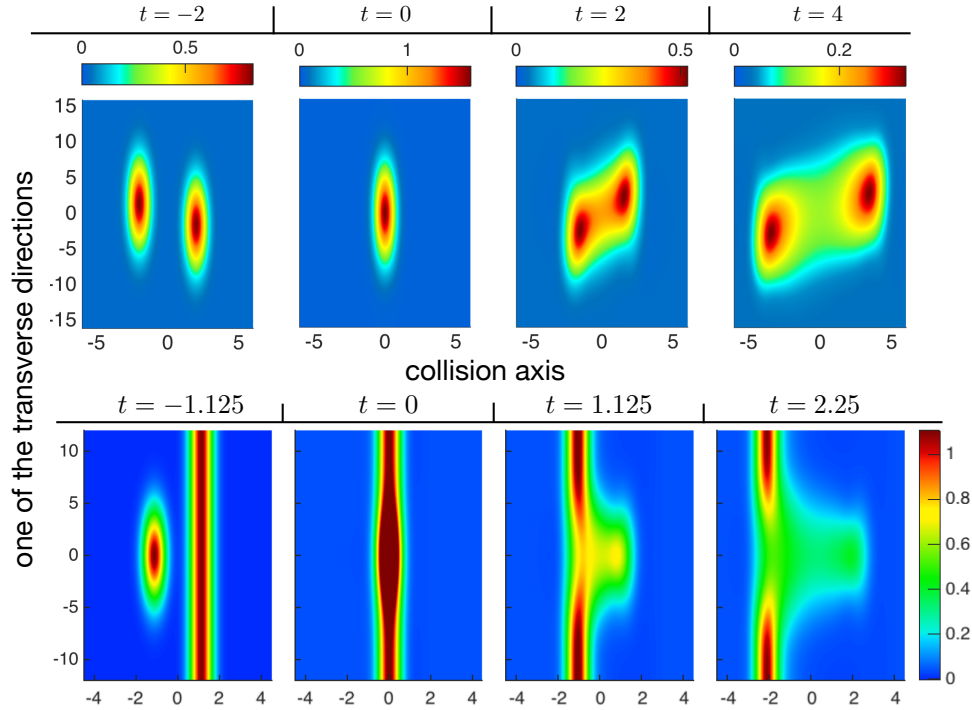


FIG. 31 Energy density in holographic heavy-ion collisions with transverse dynamics. TOP: Off-central collision with modest elliptic flow. BOTTOM: Proton-nucleus collision, as modeled by a shockwave with a small Gaussian extent in the transverse plane (left projectile) and a planar shock (right projectile). The smaller projectile punches out a hole in the larger projectile and excites matter at mid-rapidity, leading to substantial radial flow. Plots adapted from (Chesler and Yaffe, 2015) and (Chesler, 2015).

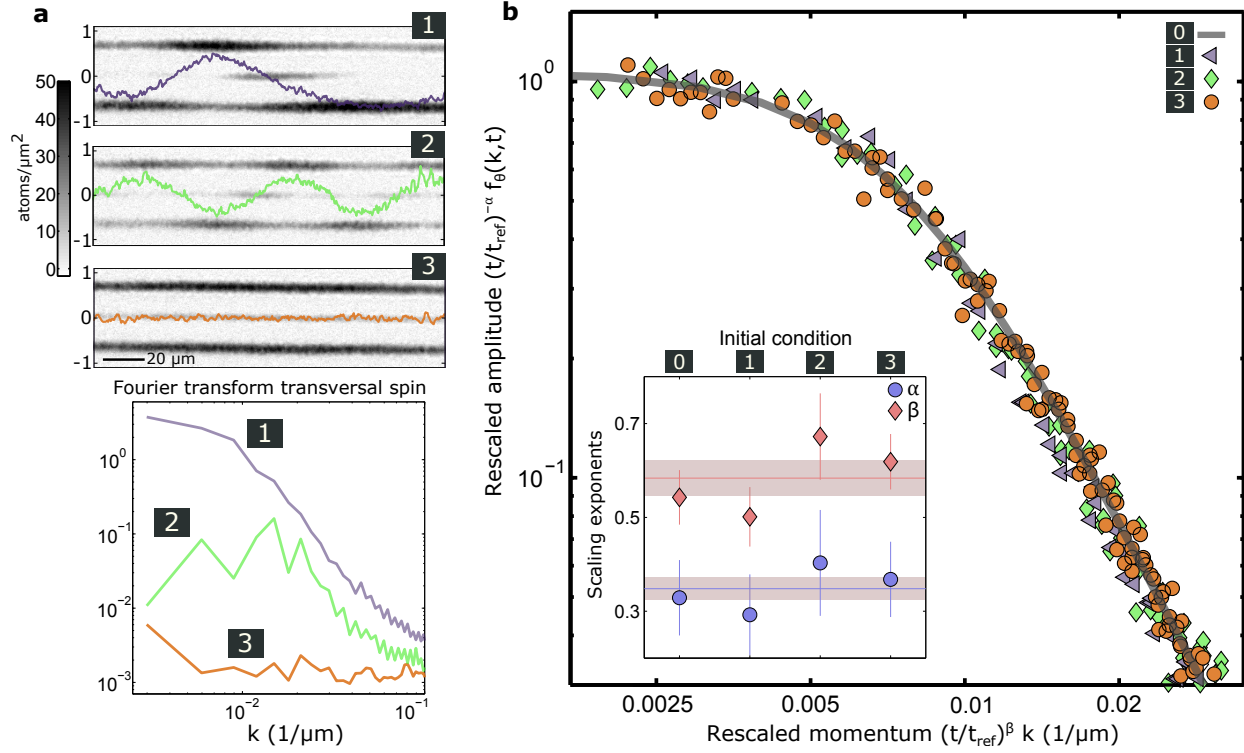


FIG. 32 (a) Absorption images of different magnetic hyperfine states of a spin-one Bose gas with the extracted transversal spin (solid lines) for three different far-from-equilibrium initial conditions. (b) All initial conditions lead to the same universal scaling behavior, such that all data points collapse onto a single curve after rescaling with time using the universal exponents α and β . Plots taken from Ref. (Prüfer *et al.*, 2018).