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Temperature-dependent ^{29}Si incorporation during deposition of highly enriched ^{28}Si films

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21	In this study, we examine the mechanisms leading to ²⁹ Si incorporation into highly
22	enriched ²⁸ Si films deposited by hyperthermal ion beams at elevated temperatures in the
23	dilute presence of natural abundance silane (SiH ₄) gas. Enriched ²⁸ Si is a critical material
24	in the development of quantum information devices because ²⁸ Si is free of nuclear spins
25	that cause decoherence in a quantum system. We deposit epitaxial thin films of ²⁸ Si
26	enriched in situ beyond 99.99998 % 28 Si onto Si(100) using an ion beam deposition
27	system and seek to develop the ability to systematically vary the enrichment and measure
28	the impact on quantum coherence. We use secondary ion mass spectrometry to measure
29	the residual ²⁹ Si isotope fraction in enriched samples deposited from ≈ 250 °C up to 800

°C. The ²⁹Si isotope fraction is found to increase from $< 1 \times 10^{-6}$ at the lower 30 temperatures, up to > 4 x 10^{-6} at around 800 °C. From these data, we estimate the 31 temperature dependence of the incorporation fraction, s, of SiH₄, which increases sharply 32 from about 2.9 x 10^{-4} at 500 °C to 2.3 x 10^{-2} at 800 °C. We determine an activation 33 energy of 1.00(8) eV associated with the abrupt increase in incorporation and conclude 34 that below 500 °C, a temperature independent mechanism such as activation from ion 35 36 collisions with adsorbed SiH₄ molecules is the primary incorporation mechanism. Direct 37 incorporation from the adsorbed state is found to be minimal.

38

I. INTRODUCTION

In solid state quantum information (QI), enriched ²⁸Si is a critical material for the 39 40 further development of silicon based quantum computing architectures, e.g., quantum dots, quantum wells, and few dopant atoms in Si. By eliminating ²⁹Si nuclei, which have 41 a non-zero nuclear spin and are present with roughly a 5 % natural abundance, pure 28 Si 42 43 becomes an ideal spin-free environment in which to place the electron and nuclear spins of qubits. Without the randomly fluctuating nuclear spins present, donor spins in ²⁸Si 44 45 interact with their environment far less than in natural silicon leading to a greatly enhanced coherence time (T_2^*). Consequently, ²⁸Si has been dubbed a "semiconductor 46 vacuum"¹. Theoretical modeling and bulk electron spin resonance (ESR) experiments 47 predicted the enhancement in T_2^* to be proportional to the reduction in ²⁹Si 48 concentration^{2,3}, which further spurred interest in exploiting ²⁸Si experimentally. 49 50 Numerous research groups have shown through bulk ESR and nuclear magnetic resonance (NMR) experiments of ³¹P spins in ²⁸Si that nuclear and electron spin 51

coherence (echoed, T_2) times can easily exceed seconds^{1,4–6}. Si and Si/Ge based quantum 52 computing also can benefit from utilization of ²⁸Si, e.g., quantum dots formed in Si and in 53 quantum wells within ²⁸Si/SiGe heterostructures^{7,8}. A few of these groups have begun to 54 show both long T_2 times and coherent manipulation in ²⁸Si for both bulk donor spins⁹ as 55 well as single electron spins in quantum wells⁸ and quantum dots^{4,10}. Additionally, qubit 56 57 manipulation schemes, which have been proposed for arrays of quantum dot qubits, and which involve tuning the qubit ESR frequency via a Stark shift¹¹, have been demonstrated 58 in single quantum dots in ²⁸Si. This Stark shifting mechanism relies on qubit spins that 59 have very narrow inhomogeneous ESR linewidths of a few kHz^{1,12}, which have only been 60 shown in a material with homogeneous mass such as highly enriched ²⁸Si. 61

Despite these advantages, only a limited amount of the most highly enriched ²⁸Si 62 (99.995 % ²⁸Si) is available within the solid state quantum computing community for use 63 in ensemble spin QI experiments. Historically, ²⁸Si has primarily been produced at great 64 65 cost and effort through international collaborations such as the International Avogadro Coordination (IAC)¹³, which produced bulk crystals. Other sources of ²⁸Si include 66 67 enriched epilayers grown on natural abundance Si substrates, which are more abundant within the community than bulk ²⁸Si but are typically less highly enriched (≈ 99.9 % 68 ²⁸Si). QI experiments on single donor and dot spins largely utilized ²⁸Si epilayer samples 69 and demonstrated the benefit of enriched ²⁸Si to QI. A helpful review of some of the 70 different sources of ²⁸Si that have been used in the QI field has been presented by Itoh 71 and Watanabe¹⁴. The lack of simply produced, readily available, and consistently highly 72 enriched ²⁸Si has led us to develop a mass selected ion beam deposition system capable 73 74 of taking natural abundance silane (SiH₄) gas and enriching it *in situ* to an extremely high

15 level of ²⁸Si just prior to depositing it epitaxially on a target Si(100) substrate. Using this 176 system, thin films of highly enriched materials including ²⁸Si have previously been grown 177 amorphously on Si substrates at room temperature^{15,16} as well as epitaxially at higher 178 temperatures (this work). Several other groups have also previously demonstrated 179 enriched ²⁸Si thin film deposition using an ion beam system, though not generally with a 180 focus on high quality material (highly enriched, chemically pure) for QI^{17–19}.

In addition to the general scarcity of ²⁸Si for semiconductor quantum computing 81 82 research, a specific need exists for enriched silicon with targeted levels of enrichment to facilitate mapping the dependence of T_2^* on ²⁹Si concentration in the few-spin regime. 83 Recent ESR measurements^{4,10} of T_2^* for single ³¹P spins in ²⁸Si have disagreed with the 84 theoretical predictions for the same systems². Our ion beam deposition system provides 85 us with a unique opportunity to produce ²⁸Si material at targeted levels of enrichment, 86 which can enable the mapping of T_2^* as a function of enrichment in the single spin 87 regime. Understanding the sources of residual ²⁹Si in our films, and how to control them, 88 89 is a necessary step towards targeting specific enrichments.

In order for enriched ²⁸Si to be useful as a host material for gubit spins, it must not 90 91 only be isotopically pure, but chemically pure and highly crystalline so as to avoid 92 introducing other sources of decoherence such as impurity nuclear spins and dangling bonds. Achieving single crystal epitaxial deposition of ²⁸Si films on Si(100) substrates 93 94 requires elevated substrate temperatures during deposition. We previously showed that for amorphous room temperature deposition, the ²⁹Si in the film could be accounted for 95 by ²⁹SiH₄ incorporation into ²⁸Si films due to a physical adsorption process which was 96 coincident with the ion beam deposition¹⁶. Depositing with an elevated substrate 97

98	temperature is expected to lead to a change in the activation of the ²⁹ Si contamination
99	through a chemical activation process described by a reactive sticking coefficient that is
100	similar in nature to chemical vapor deposition (CVD). Therefore, the goal of this study is
101	to determine how the ²⁹ Si and ³⁰ Si incorporation is affected by increasing the substrate
102	deposition temperature. Measurements of residual ²⁹ Si and ³⁰ Si isotope fractions
103	(amount-of-substance fraction in mol/mol, or isotopic concentration) are made using
104	secondary ion mass spectrometry (SIMS) which is extremely sensitive to isotope ratios.
105	This means that our monoisotopic ion beam deposition system offers a unique method of
106	measuring the activated sticking of natural abundance SiH ₄ on Si substrates because any
107	isotopic contaminants (29 Si and 30 Si) incorporated via sticking from SiH ₄ are easily
108	distinguished from the background of pure ²⁸ Si being deposited from the ion beam.
109	In this article, we deposit epitaxial thin films of ²⁸ Si and measure the residual ²⁹ Si
110	and ³⁰ Si isotopes in samples deposited in different SiH ₄ partial pressures while varying
111	the substrate deposition temperature of each sample. From these measurements, we
112	extract the temperature dependence of the incorporation fraction, s , of ²⁹ SiH ₄ due
113	primarily to reactive sticking and determine the associated activation energy, E_c . These
114	results give us an understanding of, and thus control over, the ²⁹ Si concentration in ²⁸ Si
115	films.

116 II. EXPERIMENTAL METHODS

117 A. lon beam system and selectivity

118 Enriched ²⁸Si thin films are deposited using a hyperthermal energy ion beam 119 deposition system. This system injects commercial, natural abundance SiH_4 source gas 120 into a high vacuum Penning-type ion source generating a Si^+ ion beam and then mass

121	filters the ions in a magnetic field before depositing them onto heated Si(100) substrates.
122	The mass separation principle and ion beam system used here have been described in
123	more detail elsewhere ^{16,20} . Previously, we analyzed SiH ₄ incorporation in samples
124	deposited at room temperature, but now we examine the temperature dependence of $s(T)$
125	at temperatures required to facilitate epitaxial growth. The SiH ₄ used here is unenriched
126	and is thus assumed to have a Si isotopic abundance roughly the same as natural Si (92.2
127	$\%$ $^{28}Si,$ 4.7 $\%$ $^{29}Si,$ and 3.1 $\%$ $^{30}Si).$ This SiH_4 is also 99.999 % pure according to the gas
128	vendor. This SiH_4 is used to form a plasma within the ion source that both cracks and
129	ionizes the SiH ₄ molecules. From the plasma, singly charged Si ⁺ ions (charge of 1 e , the
130	elementary charge) are extracted into the beamline at high voltage and enter a system of
131	electrostatic optics, which steer them into the magnetic field of a 90° sector mass
132	analyzer. ²⁸ Si has a mass \approx 28 u (unified atomic mass units), and by tuning to the
133	appropriate magnetic field, ions with a mass-to-charge ratio $\approx 28 \text{ u/}e \ (\pm 0.16 \text{ u/}e \text{ at } 28 \text{ u})$
134	pass through an aperture while those with different mass-to-charge ratios are rejected.
135	Beyond the aperture, ions are focused onto the target substrate in an ultra-high vacuum
136	(UHV) deposition chamber. In parallel, gas diffusion of SiH_4 from the ion source to the
137	sample location in the deposition chamber provides the source of the ²⁹ Si and ³⁰ Si
138	contamination concerning this report.
139	A mass spectrum of the constituents of the ion beam is generated by sweeping the

magnetic field of the mass analyzer while monitoring the intensity of the ion beam
current at the target. Each of the resulting series of peaks in the current vs. magnetic field
signal correspond to different integer mass-to-charge ratios of atomic or molecular ions.
In the mass spectrum of SiH₄, a series of mass peaks are formed starting at 28 u, which is

²⁸Si. The adjacent peak at 29 u corresponds to a combination of ²⁹Si and ²⁸SiH. We 144 calculate the intrinsic geometric mass selectivity (i.e. mass resolution) of the ion beam 145 system from the spectrum to estimate the amount of ²⁹Si potentially contaminating the 146 ²⁸Si beam. These calculations use Gaussian fits to the mass peaks to determine the 147 148 overlap of the 29 u peak on the 28 u peak. They give a peak separation of about 11 σ 149 (standard deviations) and a resolving power $m/\Delta m \approx 80$ for m = 28 u, which yields a lower bound on the fraction of ions at 28 u consisting of ²⁹Si of roughly 10⁻²⁵, as 150 previously discussed elsewhere¹⁶. However, this argument neglects gas scattering effects, 151 152 which would likely be a dominant contributing factor to the contamination compared to 153 this extremely small geometric component. Gas scattering causes an ion at mass 29 u to 154 lose sufficient energy to be incorporated into the 28 u trajectory and pass through the 155 selecting aperture. This scattering tail effect, or abundance selectivity, for a single magnet system can be estimated from literature to contribute $\approx 1 \times 10^{-6}$ of the higher mass peak to 156 the lower mass peak (for a mass of 28 u) $^{21-23}$, but it is difficult to measure experimentally 157 in our system. That scattering fraction combined with the ²⁹Si natural abundance gives an 158 estimate for an upper bound on the 29 Si concentration in the 28 Si beam of roughly 10^{-7} . 159 160 This concentration may be significant for the samples measured in this study with the lowest ²⁹Si isotope fractions approaching 10⁻⁷. However, we do not see any evidence that 161 we are reaching this scattering enrichment limit, e.g. attenuation of the ³⁰Si isotope 162 fractions compared to ²⁹Si, discussed further below. For the purposes of this paper, we 163 164 assume the scattering tail contribution is negligible in these experiments and the ion beam is pure 28 Si. In this study, we consider the difference between the expected (100 % 165

enriched) and measured enrichment by identifying only the natural abundance SiH_4 gas diffusing from the ion source into the deposition chamber as the source of ²⁹Si and ³⁰Si.

168

B. Temperature and pressure determinations

169 In order to extract the temperature dependence of the incorporation fraction, s, of 29 SiH₄ on 28 Si(100), a good estimate of the SiH₄ partial pressure at the sample location 170 171 during deposition is needed. The base pressure of the deposition chamber was measured to be approximately 1.3×10^{-8} Pa (1×10^{-10} Torr) for these experiments. During operation 172 of the ion beam, SiH₄ is leaked into the ion source at a pressure of about 4×10^{-4} Pa (3 x 173 10^{-6} Torr), and some SiH₄ gas diffuses into the deposition chamber, which typically sees 174 175 a factor of 50 to 100 increase from the base pressure. We estimate the partial pressure of 176 SiH₄ at the sample from measurements of the individual gas components using a residual 177 gas analyzer (RGA) in the chamber. Typically, while operating the ion beam, the fraction of total pressure increase due to SiH₄ and other Si hydrides is estimated to be about 28 %, 178 179 while the rest is mostly H_2 . This is because a lot of SiH₄ is cracked into SiH_x, where SiH_x 180 is a combination of Si hydrides (1 < x < 4), and thus results in a large amount of byproduct H₂. The SiH_x partial pressure is then estimated to vary from 1.4×10^{-7} Pa to 9.6 181 x 10^{-7} Pa (1.1 x 10^{-9} Torr to 7.2 x 10^{-9} Torr) across the high temperature samples in this 182 183 study. The previous room temperature samples were generally deposited in higher partial pressures up to 4.4×10^{-5} Pa (3.3 x 10^{-7} Torr) for two samples. From the partial pressures, 184 we get an important quantity in this analysis, the SiH_x molecular gas flux, F_a , which is 185 186 calculated using the estimated partial pressure during deposition and the Hertz-Knudsen 187 equation,

188
$$F_g = p(2\pi m k_B T_g)^{-1/2},$$
 (1)

189 where *p* is the pressure, *m* is the mass of the gas (SiH₄), k_B is Boltzmann's constant, and 190 T_g is the gas temperature (assumed to be 21°C). The pressures for the high temperature 191 samples correspond to F_g gas flux values between 4 x 10¹¹ cm⁻²s⁻¹ and 3 x 10¹² cm⁻²s⁻¹.

192 The sample temperature during deposition was also carefully measured to ensure 193 an accurate mapping of enrichment versus temperature and determination of *s*.

194 Temperature was measured in this study using an infrared pyrometer that viewed the 195 sample through a window from outside the vacuum chamber. The temperature readings 196 were calibrated for our system by monitoring eutectic samples in the chamber near their 197 melting temperature and adjusting the pyrometer emissivity to match the known melting point temperatures. The two temperature standards used here were a Au-Si eutectic²⁴ and 198 199 an Al-Si eutectic, which were each held at their melting temperatures of 363 °C and 577 200 °C respectively while calibrating the pyrometer. Multiple calibrations are needed because 201 the emissivity of Si is not constant with temperature. The emissivity as measured through 202 the chamber window changes from about 0.25 at 363 °C up to 0.42 at 577 °C and is 203 expected to reach a high value of 0.68 within the range of temperatures used in this study. 204 This range of values is similar to emissivity values for Si surfaces reported in the literature (≈ 0.1 at 100 °C and 0.68 at > 800 °C)²⁵. Including uncertainties in the 205 206 calibration, the pyrometer temperature readings of the substrate are estimated to have a 5 207 % relative uncertainty due to fluctuations in the current used for sample heating as well 208 as temperature gradients across the sample.

209 C. Substrate preparation and deposition

²⁸Si samples were deposited epitaxially on a variety of natural abundance Si(100)
substrates including p-type, n-type, and undoped (intrinsic) wafers that were cleaved into

212 chips measuring 4 mm by 10 mm. Substrates were cleaned ex situ using standard Si 213 cleaning procedures for metals and organics used in complementary metal-oxide-214 semiconductor (CMOS) technology consisting of a piranha etch, hydrofluoric acid (HF) strip, and "standard clean-2" $(SC2)^{26}$. The chips are capped with a thin protective oxide 215 216 during the final SC2 cleaning step. After cleaning, the chips were immediately mounted 217 onto sample holders and loaded into the vacuum chamber via a load lock. Substrates were 218 then prepared for deposition in situ by first degassing them overnight at 600 °C and then 219 flash annealing them to 1200 °C for \approx 10 s several times. This flash removes the oxide and produces a clean (2x1) reconstructed Si(100) surface on which to deposit ²⁸Si 220 221 epitaxially. Typically, flashed substrates were inspected using a UHV scanning tunneling 222 microscope (STM) to ensure a clean surface. The substrate temperature was then elevated 223 to the growth temperature prior to exposure to the ion beam for deposition. To map out the ²⁹Si temperature dependence, samples were deposited at substrate temperatures 224 225 ranging from 249 °C up to 812 °C in increments of roughly 100 °C. Also included in this 226 study for qualitative comparison are data from previous amorphous samples deposited at 227 room temperature (≈ 21 °C) on substrates that were only prepared *ex situ* with HF. For the higher temperature samples, ²⁸Si ions were deposited onto the substrates 228

in the hyperthermal energy regime with an average ion energy at the target of ≈ 40 eV. This energy is selected to stay as high as possible to minimize space charge effects while keeping the net sputter yield ≈ 0 . Typical ion beam currents of around 500 nA were achieved over an area on the chip of about 6 mm², which corresponds to an average ion flux, F_i , that varied from 3 x 10¹³ cm⁻²s⁻¹ to 3 x 10¹⁴ cm⁻²s⁻¹. F_i was calculated from the film thickness measured *post facto* in combination with the deposition time for each

sample. These fluxes correspond to deposition rates between 0.3 nm/min and 3.9 nm/min.
The thicknesses of the deposited films were taken from the calibrations of the SIMS
depth profiles and ranged from about 50 nm to over 300 nm depending on the sample and
the measurement location on the deposition spot.

239

D. SIMS measurements

Measurements of the enrichment of the ²⁸Si films grown at different temperatures 240 were made *ex situ* using SIMS. The samples were sputter eroded using an O_2^+ primary 241 beam at an impact energy of 8 keV and a current of 1 nA while monitoring counts of ²⁸Si, 242 ²⁹Si, and ³⁰Si to determine their relative abundances. The beam was focused to a probe 243 244 size of a few micrometers in diameter, and it was raster-scanned over a 50 µm x 50 µm 245 area. The analyzer's magnetic field was cycled to allow the positive secondary ions for 246 each isotope to be detected by a secondary electron multiplier. The mass resolving power 247 for the measurement conditions was $m/\Delta m \approx 6000$ measured at 10 % of the peak maximum. This resolving power is necessary to cleanly separate the ²⁹Si signal from the 248 ²⁸SiH signal that arises due to the SIMS process. Under these conditions, we estimate that 249 less than 10⁻⁵ of the ²⁸SiH signal contributes to the ²⁹Si measurement. Uncertainties of the 250 251 isotope ratios were determined from the standard deviation of the mean of the 252 measurements. A profilometer was used to calibrate the depth scale in the measurements, and that allowed for determination of ²⁸Si film thicknesses and growth rates. 253 254

255 III. ANALYTICAL APPROACH

In this study, we evaluate a multi-mechanism gas sticking deposition model to correlate the SIMS measurements of enrichment to the deposition conditions (e.g. SiH_x

partial pressure and deposition rate) for samples deposited between ≈ 21 °C and 850 °C. 258 259 We consider two distinct sources of Si atoms that contribute to the films. The dominant (high flux) source is the ion beam, assumed to be pure ²⁸Si as discussed above, and the 260 261 second is the diffusive partial pressure of SiH_x from the ion source, which contains all 262 three Si isotopes in their natural abundance. The SIMS measurements provide the resultant isotopic concentrations for ²⁹Si or ³⁰Si as a fraction of the total Si deposited. The 263 measured isotope fractions of ²⁹Si and ³⁰Si are modelled by the mixed ion beam 264 deposition and gas sticking, to provide a combined deposition model, c_z (with z denoted 265 as 29 for ²⁹Si and 30 for ³⁰Si), given by: 266

$$267 c_z = \frac{F_g a_z s}{F_g s + F_i}, (2)$$

where F_g is the SiH_x gas flux, F_i is the ²⁸Si ion flux, a_z is the natural abundance of ²⁹Si or ³⁰Si in the SiH₄, and *s* is an effective incorporation fraction. We simplify c_z by defining the SiH₄ flux ratio $d = F_g/F_i$ that correlates the isotope concentrations to deposition conditions:

$$c_z = \frac{a_z s d}{1 + s d},\tag{3}$$

where c_z increases approximately linearly with *d* in the dilute regime ($d \ll 1$) where most experiments were performed. Additionally, when $F_g \gg F_i$ (never true in our experiments), then the natural abundance source dominates and $c_z \rightarrow a_z$, the natural abundance ratio. To add statistical weight and simultaneously consider the ²⁹Si and ³⁰Si data, we generalize the isotope specific model of Eq. (3) by dividing it by each isotope's natural abundance so that ²⁹Si and ³⁰Si data can be fit together within the same model. This has the effect of changing the units from an isotope specific incorporation to total 280 (all isotopes) adsorbed SiH₄, which gives a total gas sticking deposition model, $c_{tot.}$, 281 where

282
$$c_{tot.} = \frac{c_z}{a_z} = \frac{sd}{1+sd}$$
. (4)

Eq. (4) allows us to determine the incorporation fraction *s* for each sample deposited at different temperatures, to get the trend of *s* vs. *T*. Additionally, since *s* represents the fraction of diffusing SiH₄ gas that become permanently incorporated in the film, at the single molecule level, *s* is the probability that a specific molecule becomes incorporated on the timescale of an arriving ion.

288 To describe the anticipated phenomenological behavior of s vs. T, we define a 289 temperature dependent incorporation model, s(T), that considers two classical gas incorporation mechanisms: a sticking term (physisorption), s_p , and a higher temperature 290 reactive mechanism, s_c , (e.g., hydrogen cracking or chemisorption). Both s_c and s_p are 291 expected to be thermally activated, where s_p decreases with increasing temperature as 292 more molecules escape (desorb), and s_c increases with increasing temperature as more 293 molecules react and bond to the surface. Since c_{tot} , is normalized by the total flux, s(T)294 295 is the probability per molecule that SiH₄ is incorporated. We define these individual 296 components to be:

297
$$s_p = 1 - A_p \exp(-E_p/k_B T)$$
 and $s_c = A_c \exp(-E_c/k_B T)$, (5)

where E_p is the activation energy for "physisorption," E_c is the activation energy for "chemisorption," k_B is the Boltzmann constant, and *T* is the substrate temperature during deposition. The prefactors A_p and A_c are free parameters that account for the average site occupancy and the time integral over many activation attempts that occur at molecular

302 vibrational frequencies and other atomistic factors. From this, the total incorporation

303 fraction at a given temperature is the sum of the two sticking components,

304
$$s(T) = s_p + s_c + s_0 = 1 - A_p \exp(-E_p/k_B T) + A_c \exp(-E_c/k_B T) + s_0.$$
 (6)

305 An s_0 term accounts for temperature independent incorporation, like activation from a

306 collision with an ion in the depositing flux. The simple sum assumes that the two

307 mechanisms are independent, e.g., physisorption is not a requirement for chemisorption,

308 and that $E_c \gg E_p$ such that $s(T) \le 1$ for all *T*.

309 IV. RESULTS AND DISCUSSION

310 A. ²⁸Si enrichment and pressure correlation

311 For all deposition temperatures (except room temperature and 249 °C), epitaxial growth of the ²⁸Si films was achieved. Interface widths of a few nanometers were found 312 313 for samples grown between roughly 350 °C and 420 °C, while the surface roughness 314 increased significantly for higher temperature samples. The sample with the best enrichment and lowest ²⁹Si isotope fractions in this study was deposited at 502 °C. A 315 316 SIMS depth profile for this sample is shown in Fig. 1. Between 40 nm and 280 nm into the film, the averaged isotope fractions are; 28 Si: 99.9999819(35) %, 29 Si: 1.27(29) x 10⁻⁷, 317 and 30 Si: 5.5(19) x 10⁻⁸. The average values for 29 Si and 30 Si are represented by dashed 318 319 lines in Fig. 1 and fall below the data because of many zero counts on the SIMS detector 320 for those measurement cycles. At a depth of around 300 nm in this sample, the sputter 321 beam erodes into the substrate and the isotope fractions return to the natural values.

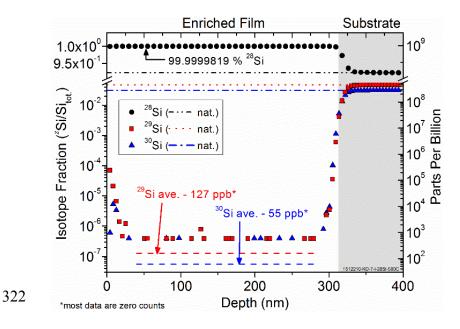
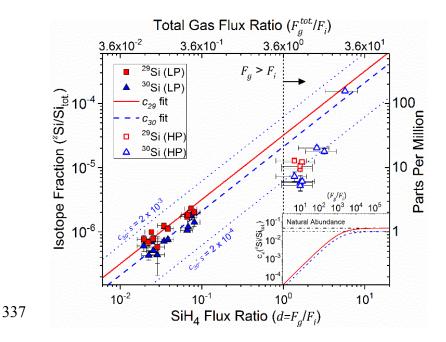


FIG. 1. (Color online) Semi-log SIMS depth profile of the 502 °C sample showing the isotope fractions of ²⁸Si (circles), ²⁹Si (squares), and ³⁰Si (triangles). The sharp increase in ²⁹Si and ³⁰Si isotope fractions to the natural abundance levels (dotted lines) at 300 nm corresponds to reaching the substrate (gray shade). The average isotope fractions (from 40 nm to 280 nm) are shown as dashed lines. These averages lie below the visible data because many of the data were zero counts.

Temperature is not the only significant experimental parameter affecting enrichment; the enrichment is also seen to depend linearly on the SiH_x partial pressure (when $F_g \ll F_i$). We plot the raw SIMS data for the room temperature samples as a function of the SiH₄ flux ratio, $d = F_g/F_i$, in Fig. 2, with ²⁹Si (squares) and ³⁰Si (triangles) isotope fractions plotted together. The top axis of Fig. 2 shows the total gas flux ratio, $F_g^{tot.}/F_i$, using the total measured pressure increase during deposition without subtracting out the estimated H₂ fraction in the gas. The hydrogen subtraction only shifts



338 FIG. 2. (Color online) Correlation plot of isotope fraction from SIMS vs. SiH₄ flux ratio $d = F_a/F_i$ shown on a log-log scale for samples deposited at room temperature (≈ 21 339 °C); ²⁹Si (squares) and ³⁰Si (triangles). The top axis is the flux ratio for the total gas flux 340 during deposition. The gas sticking deposition model, c_z is fit to the ²⁹Si and ³⁰Si data 341 (solid and dashed line respectively) and gives a value of $s = 6.8(3) \times 10^{-4}$. c_z is linearly 342 343 proportional to s over the range of the data, but asymptotes (see inset) to the natural abundance values when $F_q \gg F_i$ (dash-dotted lines). To demonstrate the sensitivity to 344 changes in the free parameter s, c_{30} for two other values of s (blue dotted lines) are also 345 shown. "LP" and "HP" denote data from two different experimental configurations, only 346 the "LP" data is used for quantitative analysis. Horizontal error bars are dominated by the 347 348 uncertainty in the pressure measurements, and vertical error bars represent the standard 349 deviation of the SIMS data.

351 the axis laterally and both the 29 Si and 30 Si isotope fractions have a strong linear

352 correlation with the SiH₄ flux fraction (deposition conditions), showing that a higher SiH₄

- 353 flux fraction produces a larger isotope fraction in the sample. This high correlation is
- 354 strong evidence that the diffusive SiH_x is the primary source of the minor isotopes.

355	To determine the probability of SiH_x being incorporated during growth, we fit the
356	data in Fig. 2 using Eq. (3) to get c_{29} and c_{30} with s as the only free parameter. These fits
357	are shown in Fig. 2 as solid and dashed lines respectively. They are approximately linear
358	over the range of the data with a slope proportional to <i>s</i> . For $d > 10^4$, c_z asymptotes to
359	the natural abundance values (see Fig. 2 inset) of 4.7 % for 29 Si and 3.1 % for 30 Si. The
360	best fit to the data yields a room temperature incorporation fraction of $s = 6.8(3) \times 10^{-4}$.
361	The points at $d > 1$ were deposited in a different vacuum chamber with poorer pressure
362	measurement but are included to show continuation of the qualitative trend to much
363	higher pressures. Only the "LP" data is used for quantitative analysis at room temperature
364	and this configuration was used for all the higher temperature work. $c_{30}(d)$ is plotted for
365	two other values of s which span an order of magnitude around the fit value so that the
366	reader can see the sensitivity of the fit to <i>s</i> . Note that when viewed as a log-log plot as in
367	Fig. 2, c_z does not change slope as s is varied, it only changes vertical offset.
368	Affirmation that the diffusive SiH ₄ partial pressure is the dominant source of the
369	minor isotopes is found in the measured isotope ratios ²⁹ Si/ ³⁰ Si for each sample, shown in
370	Fig. 3. If the minor isotope contribution originated from the ion beam, ³⁰ Si would be
371	attenuated compared to ²⁹ Si and increase the ²⁹ Si/ ³⁰ Si ratio above the natural value, e.g.
372	datum 31. Instead, the measured isotope ratios for most samples are very close to the
373	natural value of about 1.5, indicating that the source of ²⁹ Si and ³⁰ Si has a natural
374	abundance of Si isotopes, e.g., the SiH ₄ source gas.

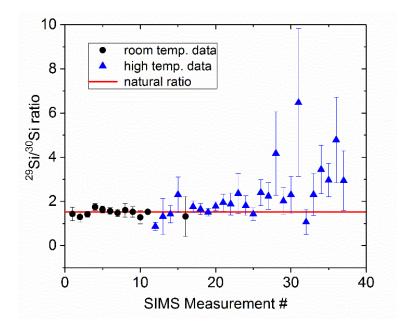


FIG. 3. (Color online) 29 Si/ 30 Si isotope ratios for samples deposited at room temperature (circles) and elevated temperatures (triangles). The ratios of these samples agree with the natural abundance ratio of 1.5 (line) indicating that the source of 29 Si and 30 Si is naturally abundant, probably the SiH₄ gas. Measurement numbers 28, 31, and 36, which lie above a 29 Si/ 30 Si ratio of four, suffer from discrete counting noise in the SIMS measurements due to a total 30 Si count < 10, which makes the ratio highly volatile. Error bars are derived from the standard deviation of the SIMS data.

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385 **B.** Temperature dependence of ²⁹Si and s

Isotope fractions were measured by SIMS on samples grown at many different temperatures. For each SIMS measurement at each temperature, an average isotope fraction is found in the depth region where the minor isotope counts reach a minimum. The SIMS measurements from ²⁸Si samples grown at the low end of the epitaxial temperature range (249 °C) had a residual ²⁹Si isotope fraction of 7.9(12) x 10⁻⁷. The sample grown at the highest temperature, 812 °C, has a ²⁹Si isotope fraction of 4.32(46) x

 10^{-6} . This factor of five increase in isotope fraction is the focus of our analysis and 392 discussion. We note that for the samples deposited at 705 °C, 708 °C, and 812 °C, 393 394 significant morphological roughness may have resulted in some substrate mixing during 395 the SIMS measurement. Therefore, these data are an upper bound since the SIMS may 396 not have reached the minimum isotope fraction before breaking into the substrate. In this 397 report, we excluded data that was clearly influenced by this effect (did not reach a stable 398 isotope fraction minima), however, it is possible that this measurement artifact still 399 weakly contributed to the measured values of the isotope ratios reported here for the 400 highest temperature samples.

In this work, we primarily report the dependence of the ²⁹Si and ³⁰Si incorporation 401 from the SiH₄ partial pressure on the different substrate temperatures. The raw 29 Si 402 isotope fraction increases rapidly in the range from 502 °C (1.27 x 10^{-7}) to 812 °C (4.32 x 403 10⁻⁶), however, maintaining an identical SiH₄ partial pressure and ion beam flux was not 404 possible for each sample. Using the c_z function though, we can adjust the isotope fraction 405 406 values at each temperature for the variations in deposition parameters. To do this, the raw SIMS measurements are adjusted to a common SiH₄ flux ratio $d = F_g/F_i$. We choose the 407 d value corresponding to an area of the sample deposited at 502 °C, i.e., the lowest 408 measured isotope fractions of ²⁹Si and ³⁰Si. 409

The procedure for the adjustment is as follows: multiple values of the isotope fractions are determined from SIMS plots at each temperature, e.g., Fig 1. Each SIMS value is adjusted by solving Eq. (3) using the corresponding *d* value to determine *s* for each datum independently, denoted as s_T^n , where *n* is the *n*th datum measured for a sample deposited at temperature *T*. Then, using the specific s_T^n value put back into Eq

415 (3), we can adjust the isotope ratio to the *d* value of the 502 °C sample ($d_{502} = 0.0073$) 416 and denote the adjusted value as $c_z(s_T^n, d_{502})$.

In Fig. 4, the isotope fractions adjusted for pressure and ion beam flux (deposition rate) are plotted as a function of temperature to isolate the enrichment's temperature dependence. The adjusted isotope fractions appear to initially trend downwards from \approx 7.9 x 10^{-7 29}Si at 249 °C to a minimum at the 502 °C average of about 1.3 x 10^{-7 29}Si. The room temperature data appears to deviate from the low temperature trend, which is likely

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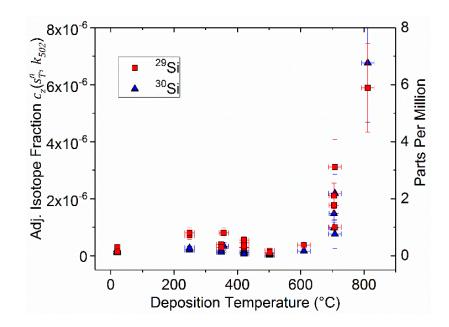


FIG. 4. (Color online) Adjusted isotope fraction, $c_z(s_T^n, d_{502})$, versus temperature for ²⁹Si 424 (squares) and ³⁰Si (triangles). The raw isotope fractions are adjusted to the deposition 425 426 conditions (F_a and F_i) of the 502 °C sample to account for differences in pressure and 427 deposition rate amongst different samples. This shows the expected increase in isotope 428 fraction at a given temperature for a sample deposited under the same conditions as the 429 502 °C sample. Horizontal error bars are due to uncertainty in the pyrometer calibration 430 and temperature fluctuation during deposition, and vertical error bars represent the 431 standard deviation of the SIMS data.

432 due to unaccounted for systematic variations, e.g., because those samples were 433 amorphous and were grown in a different experimental configuration. Surface orientation and crystallinity are known to affect the adsorption of SiH_4 on Si surfaces²⁷, which may 434 435 lead to a lower effective sticking coefficient compared to that of the crystalline samples. Above 502 °C, the adjusted ²⁹Si isotope fraction sharply increases up to 5.9×10^{-6} ²⁹Si at 436 437 812 °C. This increase is emblematic of the thermal activation of a chemical process, 438 perhaps similar to a CVD reaction, and appears to dominate over the incorporation 439 mechanism at lower temperatures.

We can analyze these mechanisms better by evaluating the values of *s* at each temperature and comparing them to the temperature dependent incorporation model. The measured isotope fractions are first converted to isotope independent SiH₄ fractions using their natural abundance values according to the generalization to obtain Eq. (4) for the total gas sticking deposition model, $c_{tot.}$. The generalized ²⁹Si and ³⁰Si data for each deposition temperature are then plotted together in Fig. 5 against their SiH₄ flux ratios and fit together using $c_{tot.}$ to get a single s_T value for each temperature.

447 The values of *s* determined from data (slopes of the lines in Fig. 5) are shown in 448 Fig. 6. These incorporation fractions are a total net sticking probability; i.e., the 449 probability that a SiH_x molecule that struck the surface was permanently incorporated into the film. The dependence of s on temperature closely follows the trend of the 29 Si 450 and ³⁰Si isotope fractions in Fig. 4. In the lower temperature regime, s trends downwards 451 slightly from a value of 1.6(2) x 10^{-3} at 249 °C to a minimum of 2.9(4) x 10^{-4} at 502 °C. 452 453 In this temperature range, the decrease of s suggests reduced incorporation due to 454 increasing thermally activated escape from physisorption (s_p) (The room temperature

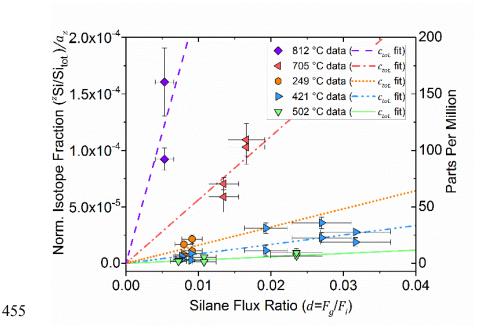
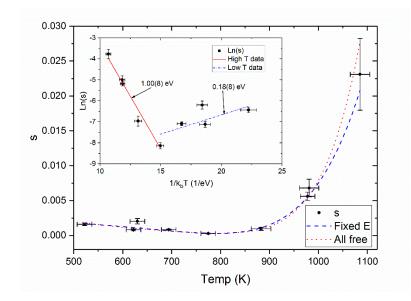


FIG. 5. (Color online) Correlation plots of the generalized isotope fractions c_{tot} , vs. SiH₄ 456 457 flux ratio shown on a linear scale for samples deposited at several elevated temperatures. The raw SIMS isotope fractions for ²⁹Si and ³⁰Si are each generalized using their natural 458 abundance, a_z , to get the estimated total adsorbed SiH₄. Then c_{tot} is fit to the data for 459 460 each temperature (solid, dashed, dotted lines) to determine s_T . The fits originate at the 461 point (0, 0) because zero SiH₄ flux results in an adsorbed SiH₄ fraction of zero. 462 Horizontal error bars are dominated by the uncertainty in the pressure measurements, and 463 vertical error bars are derived from the standard deviation of the SIMS data.

465 datum is excluded since the amorphous nature of the sample will not have the same

- 466 adsorption kinetics). Then as T is increased above 600 °C, s rapidly increases to 2.3(5) x
- 467 10^{-2} at 812 °C. This increase suggests thermal activation of the reactive sticking
- 468 coefficient, i.e., chemisorption, s_c , which increases with increasing temperature. These
- 469 values of *s* are consistent with previously reported values of the reactive sticking
- 470 coefficient of silanes on Si surfaces, although there is a large variation in the literature,
- 471 e.g., Si CVD studies have shown the reactive sticking coefficient, s_r , to range from 5 x



473

474 FIG. 6. (Color online) Incorporation fraction, s (circles) vs. deposition temperature. s is determined from the fits of c_{tot} to the data for each temperature in Fig 5. The fit to Eq. 475 476 (6) is shown in the red dotted line with all parameters free. The inset presents the same data in Arrhenius form, $\ln(s)$ vs. $1/k_B T$, with linear fits to the activation energies above 477 and below 502 °C ($1/k_B T \approx 15 \text{ eV}^{-1}$). Using the energies determined in the inset as fixed 478 479 inputs to Eq. (6), the model is refit and shown as the blue dashed line. Horizontal error 480 bars are due to uncertainty in the pyrometer calibration and temperature fluctuation 481 during deposition, and vertical error bars represent the standard deviation of the fit values 482 of Fig. 5.

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484 10^{-4} to 5 x 10^{-3} for polycrystalline Si deposition at 600 °C to 800 °C²⁸, and from 1 x 10^{-3}

485 to 3 x 10⁻⁵ for Si(111) surfaces below 500 °C^{29,30}.

Next, we fit the data in Fig. 6 using the temperature dependent incorporation model of Eq. (6). This model considers physisorption (s_p) with thermally activated escape (desorption), thermally activated incorporation (s_c) , e.g., due to reaction, as well as temperature independent process like ion activation (s_0) . The fit to Eq. (6) while leaving all five parameters free is shown as a red dotted line in Fig. 6. Visually, this

491 appears to be a good fit, but several aspects of the fit parameters suggest problems, e.g., 492 the numerical value for $E_p \approx 2 \ge 10^{-4}$ eV, $s_0 \ll 0$, and several parameters have a high-493 degree of cross-correlation, suggesting too many degrees of freedom.

494 In order to break the correlations and isolate E_p and E_c , we replot the data of Fig. 6 in an Arrhenius form, $\ln(s)$ vs. $1/k_BT$ in the inset of Fig. 6. One can see two regimes 495 of data, corresponding to above and below the 502 °C datum at $1/k_BT \approx 15 \text{ eV}^{-1}$. We 496 497 approximate each regime as being dominated by a distinct physical mechanism so that each segment individually can be fit to a line $\ln(s) = \ln(A) - E(1/k_BT)$ where the 498 499 slope is the effective activation energy in that regime. The fit line for the higher 500 temperature data is shown as a solid red line with a slope of $1.00 \text{ eV} \pm 0.08 \text{ eV}$ while the 501 fit to the lower temperature data is shown in a dash-dotted blue line with a slope of 0.18 $eV \pm 0.08 eV$ and have R² values of 0.98 and 0.61, respectively. The higher temperature 502 503 value of 1.00 eV is consistent with reported activation energies for SiH₄ CVD between 504 600 °C and 800 °C. The literature values are found to vary between about 0.4 eV to 2.2 505 eV depending heavily on experimental conditions such as surface orientation, gas pressure, and hydride species^{27–29,31–34}. 506

507 Using the energy values determined from the slopes in the Arrhenius fits, we can 508 input these as known values back into Eq. (6) supposing that they correspond to the 509 energies in that model and perform the fit with fewer degrees of freedom. A plot of this 510 fit is shown as a blue dashed line in Fig. 6. The best fit values are $s_0 = -0.9977(3)$, 511 $A_p = 0.035(6)$, $A_c = 1023(141)$ with some correlation remaining between s_0 and A_p , 512 but the fit yields reasonable relative errors and a $R^2 = 0.91$.

513 Examining the values of the best fit parameters yield some physical insights. 514 Firstly, if we consider the zero-temperature limit (albeit a substantial extrapolation), then $s(T \rightarrow 0) = s_0 + 1 = 0.0023(6)$, as opposed to our initial assumption that $s(T \rightarrow 0) =$ 515 1 and $s_0 \ge 0$, i.e. all the SiH₄ would be incorporated in the zero-temperature limit. The 516 implication of needing $s_0 < 0$ to fit the data is this initial assumption is incorrect, and 517 518 that adsorbing SiH_4 is not guaranteed to be incorporated. Rather, under these conditions, 519 only $\approx 1/400$ adsorbed molecules would become incorporated. Ultimately, another 520 mechanism (besides adsorption) must be acting as a gateway to incorporation, accounting for the $s_0 + 1$ terms discussed above. We then look to possible temperature *independent* 521 522 activation processes that can lead to incorporation, the most obvious being a collision 523 from the ion beam. Focusing on the low temperature regime where we neglect (s_c) , we can write a differential equation (below the activation of s_c) where 524

525
$$\frac{dn_s}{dt} = (1 - n_s)F_g - n_s v \exp\left(-\frac{E_p}{k_B T}\right) - p n_s F_i , \qquad (7)$$

with n_s being the probability a surface site is occupied, v is the molecular vibration 526 527 frequency and p corresponds to the probability that an ion collision results in the 528 incorporation of a SiH₄ molecule. The first term is the source term from the gas flux, 529 which becomes diminished if a substantial portion of the surface is covered by SiH₄; the second term is thermally activated desorption (similar to s_p but without the implicit 530 531 integral over the gas flux rate) and the last term represents collisional activation from the 532 ion beam. We set this equation to zero (corresponding to the steady state condition during 533 growth) and then look at the zero temperature limit for the conditions corresponding to d_{502} at each atomic site ($F_g = 0.0013 \text{ s}^{-1}$, $F_i = 0.179 \text{ s}^{-1}$) and with p = 1. We find a 534

steady state occupation probability $n_s = \frac{F_g}{pF_i + F_g} = 0.0072$, and assuming all other sites 535 would become occupied by ²⁸Si from the ion beam, then $c_{tot.} = n_s$. Considering the 536 isotopic abundance $a_{29} = 0.047$, then the probability that a surface site is occupied by a 537 ²⁹SiH₄ is $n_{s-29} = 3.4 \times 10^{-4}$, and the corresponding ²⁹Si isotope ratio would also be 538 3.4 x 10⁻⁴. By comparison, using the alternate T = 0 limit from the fit of Eq. (6) above 539 for $s(T \to 0) = 0.0023$, and using $d_{502} = \frac{F_g}{F_i} = 0.0073$ we can extrapolate that the zero 540 temperature concentration of silicon from the gas $c_{tot.}$ (fit) $\approx d_{502} \cdot s(T \rightarrow 0) =$ 541 1.7×10^{-5} and the ²⁹Si isotope fraction would be 0.8×10^{-6} . These two zero temperature 542 limits can be reconciled if we calculate $c_{tot.} = n_s \cdot p$ and drop the assumption of p = 1, 543 where we use $c_{tot.} = 1.7 \times 10^{-5}$ from the Eq. (6) fit and $n_s = 0.0072$ following Eq. (7) 544 above. From this, we find the implied ion collision activation probability p = 0.0023, 545 546 which is numerically equal to the $s(T \rightarrow 0)$. Considering the value p = 0.0023, an 547 activation probability of 1 in 400 is plausible (although probably low) for a 40 eV ion at 548 normal incidence, but considering the uncertainties present and the extreme extrapolation 549 to well below the limits of the data, this value should not be taken too seriously. 550 We do feel that the data provide good evidence for 1) a thermally activated 551 incorporation mechanism with an activation energy of 1.00(8) eV, probably 552 corresponding to breaking an H bond; 2) a temperature-independent activation 553 mechanism proportional to the surface population, probably ion collisional activation; 554 and 3) a diminishing surface population due to thermal desorption that reduces the 555 absolute incorporation from (2). 556

557 V. SUMMARY AND CONCLUSIONS

We analyzed the measured enrichment from SIMS (i.e. residual ²⁹Si and ³⁰Si 558 559 isotope fractions) of samples deposited at temperatures ranging from 249 °C to 812 °C to 560 understand how enrichment changes as a function of temperature due to SiH₄ 561 incorporation. From this analysis, we determined the temperature dependence of the incorporation fraction, s, and modeled it using two sticking terms. The lowest ²⁹Si 562 isotopic concentration was found for deposition at 502 °C at 1.27(29) x 10^{-7} . A reactive 563 564 incorporation mechanism due to CVD-like chemisorption is observed and increases 565 minor isotope concentration between 502 °C up to 812 °C. While we achieve epitaxial growth for all samples deposited above 249 °C, the temperature range above 502 °C 566 567 coincides with increased film roughness and crystalline defect formation, although 568 several mechanisms are believed to contribute to this behavior. In this work, the optimal deposition temperature for minimizing the ²⁹Si isotopic concentration is found to be \approx 569 570 500 °C, however, we consistently use 450 °C to produce high quality, smooth epitaxial 571 films with reduced risk of chemical contamination and expect to suffer little or no 572 decrease in enrichment.

We find an activation energy for this reactive incorporation of $E_c = 1.00(8)$ eV. Below the activation of the reactive process, the data suggest that incorporation directly from adsorption is rare and that a temperature independent mechanism like an ion collision is likely. Understanding the role of SiH₄ gas sticking for a range of deposition temperatures is the first step in enabling production of ²⁸Si samples with targeted levels of enrichment (²⁹Si isotope fractions) facilitating a study of T₂ coherence times as a function of ²⁹Si concentration.

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