

CHCRUS

This is the accepted manuscript made available via CHORUS. The article has been published as:

Rare math xmlns="http://www.w3.org/1998/Math/MathML" display="inline">mrow>mmultiscripts>mrow>mi mathvariant="normal">K/mi>/mrow>mprescripts>/mpres cripts>none>/none>mrow>mn>40/mn>/mrow>/mmultisc ripts>/mrow>/math> Decay with Implications for Fundamental Physics and Geochronology M. Stukel et al. (KDK Collaboration) Phys. Rev. Lett. **131**, 052503 — Published 31 July 2023 DOI: 10.1103/PhysRevLett.131.052503

Rare ⁴⁰K decay with implications for fundamental physics and geochronology

M. Stukel,¹ L. Hariasz,¹ P.C.F. Di Stefano,^{1, *} B.C. Rasco,² K.P. Rykaczewski,² N.T. Brewer,^{2,3}

D.W. Stracener,² Y. Liu,² Z. Gai,⁴ C. Rouleau,⁴ J. Carter,⁵ J. Kostensalo,⁶ J. Suhonen,⁷ H. Davis,^{8,9}

E.D. Lukosi,^{8,9} K.C. Goetz,¹⁰ R.K. Grzywacz,^{2,3,11} M. Mancuso,¹² F. Petricca,¹² A. Fijałkowska,¹³

M. Wolińska-Cichocka,^{2,3,14} J. Ninkovic,¹⁵ P. Lechner,¹⁵ R.B. Ickert,¹⁶ L.E. Morgan,¹⁷ P.R. Renne,^{5,18} and I. Yavin

(KDK Collaboration)

¹Department of Physics, Engineering Physics & Astronomy,

Queen's University, Kingston, Ontario K7L 3N6 Canada

²Physics Division, Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831, USA

³ Joint Institute for Nuclear Physics and Application, Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831, USA

⁴Center for Nanophase Materials Sciences, Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831, USA

⁵Berkeley Geochronology Center, Berkeley, California 94709, USA

⁶Natural Resources Institute Finland, Joensuu FI-80100, Finland

⁷Department of Physics, University of Jyväskylä, Jyväskylä FI-40014, Finland

⁸Department of Nuclear Engineering, University of Tennessee, Knoxville, Tennessee 37996, USA

⁹ Joint Institute for Advanced Materials, University of Tennessee, Knoxville, Tennessee 37996, USA

¹⁰Nuclear and Extreme Environments Measurement Group,

Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831, USA

¹¹Department of Physics and Astronomy, University of Tennessee, Knoxville, Tennessee 37996, USA

¹² Max-Planck-Institut für Physik, Munich D-80805, Germany

¹³Faculty of Physics, University of Warsaw, Warsaw PL-02-093

¹⁴Heavy Ion Laboratory, University of Warsaw, Warsaw PL-02-093

¹⁵MPG Semiconductor Laboratory, Munich D-80805, Germany

¹⁶Department of Earth, Atmospheric, and Planetary Sciences,

Purdue University, West Lafayette, Illinois 47907, USA

¹⁷U.S. Geological Survey, Geology, Geophysics, and Geochemistry Science Center, Denver, Colorado 80225, USA

¹⁸Department of Earth and Planetary Science, University of California, Berkeley 94720, USA

(Dated: May 10, 2023)

Potassium-40 is a widespread, naturally-occurring isotope whose radioactivity impacts subatomic rare-event searches, nuclear structure theory, and estimated geological ages. A predicted electron-capture decay directly to the ground state of argon-40 has never been observed. The KDK (potassium decay) collaboration reports the first, successful observation of this rare decay mode. A blinded analysis reveals a non-zero ratio of intensities of ground-state electron-captures ($I_{\rm EC^0}$) over excited-state ones ($I_{\rm EC^*}$) of $I_{\rm EC^0}/I_{\rm EC^*} = 0.0095 \stackrel{\rm stat}{\pm} 0.0022 \stackrel{\rm sys}{\pm} 0.0010$ (68%CL), with the null hypothesis rejected at 4σ . In terms of branching ratio, this signal yields $I_{\rm EC^0} = 0.098\% \stackrel{\rm stat}{\pm} 0.023\% \stackrel{\rm sys}{\pm} 0.010\%$, roughly half of the commonly used prediction, with consequences for various fields [Hariasz *et al.*, arXiv:2211.10343].

Potassium-40 (40 K) is a long-lived radioactive isotope figuring prominently in a variety of fields from geology to searches for exotic subatomic particles and processes. ⁴⁰K is one of the three naturally occurring K isotopes. It is fairly common, representing about a ten-thousandth of natural potassium, and contributing roughly half of the radioactivity in the human body. It plays an important role in geochronology, through K/Ar and ${}^{40}\text{Ar}/{}^{39}\text{Ar}$ dating. In addition, its commonness makes ⁴⁰K a challenging background in many particle-physics experiments looking for rare processes, such as dark matter interactions or neutrinoless double-beta decay $(0\nu\beta\beta)$. Lastly, its decay scheme involving all three types of beta decay is rare, and impactful to nuclear-structure analyses [1]. Surprisingly, among those decays, the electron-capture (EC) transitions of ⁴⁰K to ⁴⁰Ar are incompletely known.

From the standpoint of particle physics exploring the fundamental nature of our universe, a host of experiments

are looking for putative dark-matter particles that could make up the bulk of matter [2]. Low-energy X-rays and Auger electrons emitted after ⁴⁰K electron capture fall in the signal region expected in many dark-matter models: as such, electron capture is a significant background in many experiments of this type. In the case of decays to an excited stated of 40 Ar (EC^{*}), the high-energy γ -ray emitted during de-excitation provides a means of identifying and rejecting the problematic X-ray; however, in the case of electron-capture decay to ground state (EC^0) , there is no such means to tag this background. Due to the chemical similarity between Na and K, trace amounts of this ⁴⁰K are always present even in ultra-radiopure NaI scintillating crystals used for dark matter detection [3–6]. The DAMA/LIBRA experiment employs such a detector and has claimed an observation of a signal consistent with particle dark matter for over 20 years [7]. However, it has been noted that the lack of an experimental verification

of the ground state electron capture of 40 K can pose a challenge to any interpretation of the DAMA/LIBRA results in terms of a dark matter model [8].

With a total half-life of over a billion years, ⁴⁰K is one of the longest-lived naturally-occurring radioactive isotopes on Earth, motivating its longstanding use as a dating tool in geology and archaeology. Moreover, measuring samples with various techniques provides insight into the thermal history of the Earth, and variation in K/Ar and ⁴⁰Ar/³⁹Ar ages can significantly affect our understanding of terrestrial and solar-system evolution. Advances in the analytical precision of the K/Ar [9] and ${}^{40}\text{Ar}/{}^{39}\text{Ar}$ [10] techniques have caused the field to address systematic uncertainties, including those in the total and electroncapture decay rates. Though there have long been calls for the EC^0 contribution to be determined [11, 12], many commonly-used reviews in the field ignore it [13, 14]. However, it has recently been pointed out that this omission could lead to an overestimation of a sample age by tens of millions of years [15].

The overall ⁴⁰K decay scheme, including the measurement of this work, is displayed in Fig. 1a. The principal disintegration mode is by β^- to 40 Ca. It dominates electron capture, which is mainly to an excited state of ⁴⁰Ar, with a rarer, and previously unobserved, decay to the ground state which is the focus of this work. There is also a much weaker decay by positron emission to ⁴⁰Ar. With our work, the ground-state electron capture of ⁴⁰K is the only observed third-forbidden unique electron-capture decay [16]. For such a transition the theories are challenging to validate [17, 18] leading to a wide spread of reported intensity values, $I_{\rm EC^0} = (0.0-0.8)\%$ for the assumed or predicted branching ratio [8, 13, 15, 19– 22]. In a regime which is difficult to access theoretically, the KDK (potassium decay) measurement informs the extent of suppression of $0\nu\beta\beta$ processes probing physics beyond the standard model [1].

The focus of this work is the first measurement of the electron-capture decay to ground state, carried out by the KDK collaboration [25]. This experiment is challenging due to the high forbiddenness of the direct groundstate electron capture and the dominating, much more frequent excited-state electron capture, necessitating exceptionally strong background rejection. Experiments wishing to make this measurement must be able to distinguish between excited and ground-state electron-capture events with exceptionally high precision. In KDK, this is achieved by combining a very sensitive detector to trigger on X-rays from both forms of electron capture occurring in a ⁴⁰K source, with a high efficiency tagger to identify the γ -rays from EC^{*}, and thus distinguish both types of decays (Fig. 1b). This enables us to determine the ratio of their intensities, $\rho = I_{\rm EC^0}/I_{\rm EC^*}$.

The discrimination between excited and ground-state events was achieved through a unique detector configuration ([26] and Fig. 2 of [27]). When an electron capture occurs in the source, a high-resolution, low-energy threshold X-ray detector is available to observe the associated ~ 3 keV characteristic X-ray. Following such a trigger in the inner detector, a coincidence window is opened with a large, outer, γ -ray tagger which completely surrounds the source and inner detector. If the decay is to the excited state, the tagger should identify the accompanying γ -ray. In practice, various efficiencies and backgrounds affect this scheme, and events are referred to either as coincident or anti-coincident. As described further on, our analysis untangles how many are actually EC⁰ decays as opposed to EC^{*} decays.

The inner detector is a 100 mm² active surface area silicon drift detector [28] (SDD). The detector is contained inside the γ -ray tagger, the Modular Total Absorption Spectrometer (MTAS) from Oak Ridge National Laboratory [29]. MTAS is a metric tonne array of NaI(Tl) scintillators surrounding the inner detector. The tagging efficiency for 1.46 MeV γ rays is ~ 98%. Details of the setup, including the energy calibration, can be found in our technical publication [26].

The ⁴⁰K source was made from enriched KCl (16.1(6)% ⁴⁰K abundance in K) thermally deposited over 1 cm diameter onto a graphite substrate. The ~ 9×10^{17} atoms of ⁴⁰K in the source have an activity of ~ 16 Bq, equivalent to that found in two medium-sized bananas [30], and the source is 5.1(9) μ m thin to allow the X-rays to escape from it. The source rests directly in front of the SDD and is centered inside MTAS.

Data presented in what follows represent our full 34 live-day data set opened in January 2022. To avoid biases during the analysis, the anti-coincident SDD spectrum was blinded over the electron capture signal region and the silicon escape peak region while cuts and analysis methods were established. To help understand backgrounds, data were analyzed at 3 coincidence windows (1, 2 and 4 μ s), with the middle one chosen in advance for the main result. Low-level data analysis is described elsewhere [26, 27].

To determine the ratio of intensities between the ground and excited-state electron captures, our analysis compares the number of coincident and anti-coincident Ar X-rays. In order to achieve this we need to understand the respective SDD spectra in Fig. 2. In the energy region of interest, they each contain a continuum and several X-ray lines, all dominated by the source and its related interactions.

Thanks to the exceptional resolution of the SDD [26], the elemental X-ray lines, each modelled with a Gaussian, are easily identifiable. The shape for each element is the same in both the coincident and anti-coincident spectra, though the intensities differ. The Ar X-ray lines at ~ 3 keV from electron capture are the focus of this work; the coincident spectrum unambiguously provides the position and shape of the sought-after anti-coincident line. The relative coincident and anti-coincident inten-

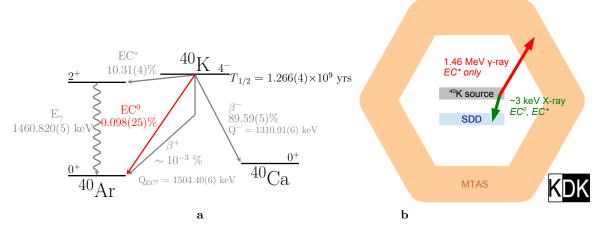


FIG. 1. (a) Decay scheme of ⁴⁰K. The branching ratios and half-life were calculated from our determination of $I_{\rm EC^0}/I_{\rm EC^*}$ and from literature values for T^- (partial half-life of the β^- decay) and T^* (partial half-life of EC^{*}) [23]. The γ transition energy is taken from [22], $Q_{\rm EC^0}$ and Q^- are taken from [24]. (b) Schematic of how the KDK experiment distinguishes EC^{*} from EC⁰ (not to scale).

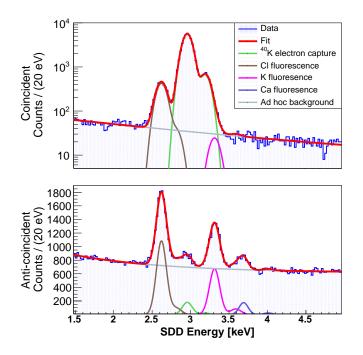


FIG. 2. SDD coincidence and anti-coincidence spectra. Results of simultaneous fit to coincident (top) and anticoincident (bottom) SDD spectra at a 2 μ s coincidence window. Signal counts are shown in green. Various fluorescence peaks and an exponential background model are included. The total minimization has an associated goodness-of-fit of p = 0.4.

sities of this contribution depend on the branching ratios, known X-ray emission probabilities of EC^0 and EC^* branches [31], and additionally on experimental effects which are i) the MTAS tagging efficiency, ii) the possibility of spurious coincidences with the MTAS background, and iii) the possibility of γ interactions with the SDD. At energies of a few keV, the continuous spectral elements are modelled with exponential and flat components. The prime contributor to these events are the 40 K β^- , possibly in coincidence with MTAS background. Due to shielding from MTAS, the natural background rate in the SDD is extremely suppressed. The β^+ branch of 40 K is also a negligible background. Particulars of the background Cl, K and Ca X-ray fluorescence lines are discussed in Sec. IIA of [27].

Using the spectral elements discussed, a joint likelihood function was developed to describe the coincident and anti-coincident binned spectra. It was then minimized to provide an estimator of the ratio of intensities (ρ) , a confidence interval, and goodness-of-fit. An example of a full fit with details of the components is shown in Fig. 2, with details of the anti-coincident signal region in Fig. 3. The overall result we obtain is:

$$\rho = I_{\rm EC^0} / I_{\rm EC^*} = 0.0095 \stackrel{\rm stat}{\pm} 0.0022 \stackrel{\rm sys}{\pm} 0.0010.$$

All uncertainties in this work correspond to a 68% level confidence interval. This result is consistent over the various coincidence windows, and corresponds to about 500 observed EC⁰ events.

As part of the estimation of systematic errors, the fit was repeated for various binnings and histogram ranges. Other contributions to the systematics include uncertainties on physical parameters, the most important of which is the uncertainty on the tagging efficiency, and uncertainties on the shape of the background continuum. Details are provided in Sec. IID of [27].

We have also performed a likelihood ratio test comparing the null hypothesis ($\rho = 0$, no ground state decay) to the alternative hypothesis (ρ is free). Results are shown in Fig. 3 for the anti-coincident signal region. If the null hypothesis was true, the probability of obtaining a canon-

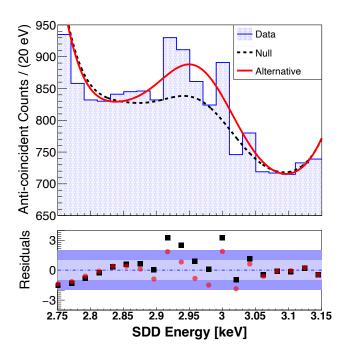


FIG. 3. The anti-coincident X-ray signal region corresponding to Fig. 2, showing a fit assuming a null hypothesis of $\rho = 0$ (dashed black) and an alternate fit with ρ free (solid red), returning $\rho = I_{\rm EC^0}/I_{\rm EC^*} = 0.0095 \pm 0.0022 \pm 0.0010$. A likelihood ratio test between the two hypotheses returns $p = 2 \times 10^{-5}$ which is equivalent to a 4σ significance. For each bin count o and associated fit value e the residual is defined as $(o - e)/\sqrt{e}$.

ical test statistic as large as the one obtained from the data is $p = 2 \times 10^{-5}$. The significance is equivalent to the probability that a Gaussian fluctuation will be at least 4σ above its mean. Our results in terms of central value of ρ and of significance are stable to changes in fitting procedures and models, including those that affect the goodness-of-fit. This first attempt to measure the elusive ⁴⁰K electron-capture decay to ground state yields compelling evidence for its existence.

Following the formulation discussed elsewhere (Sec. IIIA of [27]), we re-evaluate the decay scheme of 40 K (Fig. 1a) using the novel ρ parameter and the most recent data evaluations of the partial decay constants of the β^- and EC^{*} branches ($\lambda^- = 0.4904 \pm 0.0019$ Ga⁻¹ and $\lambda^* = 0.05646 \pm 0.00016 \text{ Ga}^{-1}$, Sec. 5.2 of [23]). Our measurement yields a ground-state decay electron-capture branching ratio of $I_{\rm EC^0} = 0.098\% \stackrel{\rm stat}{\pm} 0.023\% \stackrel{\rm sys}{\pm} 0.010\%$, which is about 50% smaller and 5 times more precise than the value obtained using the generally accepted comparative half-life $(\log ft)$ prediction [21]. This ground state branching ratio is robust within uncertainties for various commonly used sets of decay constants [13, 23, 32]. Depending on whether experimental or theoretical inputs are added, the branching ratio for β^+ can vary by a factor 2, but remains small

and does not affect the rest of the decay scheme within uncertainties.

Our measured EC^0 branching ratio is approximately a factor of two smaller than theoretical predictions [8, 15, 15]19-22], with the exception of our calculation (Fig. 6 and Sec. IIIB of [27]), 0.058(22)%. As our work represents the first measurement of a third-forbidden unique electroncapture, the variance may indicate that fine-tuning is needed in theoretical modelling. This fine-tuning impacts the traditional modelling of nuclear matrix elements of the neutrinoless double- β decay since the present theoretical analyses of the three 40 K decay branches, β^- , EC^{*}, and EC^0 , imply that the contributions of forbidden transitions to these matrix elements are heavily suppressed. This increases the computed neutrinoless double-beta decay half-life of 48 Ca by an estimated factor of $7{}^{+3}_{-2}$, as discussed in Sec. IIIB of [27], adding to the challenge of detecting this rare decay mode.

Moreover, unlike the other theoretical approaches, the calculation of this work does not depend on predicting the $I_{\rm EC^0}/I_{\beta^+}$ ratio and is therefore not reliant on the few, difficult, measurements of the I_{β^+}/I_{β^-} ratio [19, 33–37]. Our work, combined with theoretical input $I_{\rm EC^0}/I_{\beta^+} = 215.0(31)$ [17], implies $I_{\beta^+}/I_{\beta^-} = (4.9 \pm 1.2) \times 10^{-6}$, less than half the commonly-used value [19].

From the standpoint of rare-event searches, our novel measurement enables quantification of the irreducible ⁴⁰K background in signal regions that can not be tagged by the 1.46 MeV γ -ray. In particular, experiments using NaI are going to great lengths to deal with the lowenergy emissions from electron capture; measures include extreme purification of crystals [5, 6], veto systems [4, 5], and cryogenic particle identification [3]. In addition, the EC^0 decay of ${}^{40}K$ may affect the longstanding, but controversial, claim for dark-matter discovery by the DAMA experiment [8]. This argument relies not only on $I_{\rm EC^0}$, but also on detailed assumptions regarding shapes and intensities of other backgrounds in DAMA [17], and the 1.46 MeV tagging efficiency of that experiment. Our measurement is roughly half the value assumed in [8], tending to relax this type of constraint on the darkmatter interpretation of DAMA.

In geochronology, precision in age-determination has reached $\leq 1\%$ [38, 39]. Figure 4a displays the effect of including the KDK measurement of EC⁰ with commonlyused decay constants [13], which reduces K/Ar ages by about 1%. Shown also is the effect of fully updating the decay scheme by combining this work with recently reevaluated lifetimes for other ⁴⁰K transitions [23], which points to underestimation of ages by up to 20 million years. The inclusion of EC⁰ has a less dramatic *direct* effect on ⁴⁰Ar/³⁹Ar ages, inducing a variation of about 0.1%, and updating the full decay scheme leads to no change within error. However, there is an *indirect* effect on ⁴⁰Ar/³⁹Ar ages calibrated with K/Ar-dated standards, such as the widely-used Fish Canyon sanidine reference material [40], as seen in Figs. 4b,c. For instance, Fig. 4c re-evaluates the age of the Acapulco meteorite [41], one of the oldest known objects in the solar system. The ages derived from the more recent decay constants [13, 23] are statistically consistent with ages obtained using the Pb/Pb [42] and Sm/Nd [43] techniques. The variability in the other partial decay branches is now the limiting factor in the ⁴⁰K decay scheme, which remains a significant source of uncertainty in ⁴⁰Ar/³⁹Ar dating.

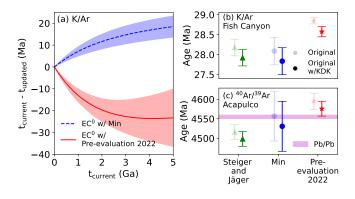


FIG. 4. Effect of updates to ⁴⁰K decay scheme on geochronology. (a) Variation in K/Ar age when adding the EC⁰ branch to commonly-used reference decay constants [13] (Min, blue, dashed, errors from EC⁰ only) and to the β^- and EC^{*} branches of a recent literature evaluation (Sec. 5.2 of [23], Kossert pre-evaluation, red, solid, all errors). In all cases $t_{\rm current}$ represents time calculated using Min decay constants neglecting EC⁰. (b) Systematic change in K/Ar age of the Fish Canyon sanidine reference material when EC⁰ is added to the previous decay constants and to those in [44]. (c) Updated ⁴⁰Ar/³⁹Ar age of the Acapulco meteorite using the decay constants and reference ages shown in (b). Pb/Pb age is updated from [42]. All uncertainties correspond to a 68% confidence level.

We have presented unambiguous experimental evidence for the previously questioned electron-capture decav to ground state of ⁴⁰K. This represents the first measured third-forbidden unique electron-capture decay, and will thus be of crucial importance for testing theoretical predictions regarding the high-multipole matrix elements relevant for neutrinoless double- β decay. Our work provides a better understanding of low-energy backgrounds in rare-event searches in non-accelerator particle physics. From the standpoint of geochronology, this first measurement of EC^0 as an additional production channel changes the determined age at the order of the current analytical precision for both K/Ar and ${}^{40}\text{Ar}/{}^{39}\text{Ar}$ methods. Our novel measurement revealing an extremely rare decay of a ubiquitous isotope has vast implications in a variety of fields.

We are grateful to Xavier Mougeot of LNHB for drawing our attention to his latest evaluation of the decay scheme of ⁴⁰K. Engineering support has been contributed by Miles Constable and Fabrice Rétière of TRIUMF, as

well as by Koby Dering through the NSERC/Queen's MRS. Funding in Canada has been provided by NSERC through SAPIN and SAP RTI grants, as well as by the Faculty of Arts and Science of Queen's University, and by the McDonald Institute. Work was performed at Oak Ridge National Laboratory, managed by UT-Battelle, LLC, for the U.S. Department of Energy under Contract DE-AC05-00OR22725. Thermal deposition was conducted at the Center for Nanophase Materials Sciences, which is a DOE Office of Science User Facility. This manuscript has been authored by UT-Battelle, LLC under Contract No. DE-AC05-00OR22725 with the U.S. Department of Energy. The United States Government retains and the publisher, by accepting the article for publication, acknowledges that the United States Government retains a non-exclusive, paid-up, irrevocable, world-wide license to publish or reproduce the published form of this manuscript, or allow others to do so, for United States Government purposes. The Department of Energy will provide public access to these results of federally sponsored research in accordance with the DOE Public Access Plan (http://energy.gov/downloads/doepublic-access-plan). US support has also been supplied by the Joint Institute for Nuclear Physics and Applications, and by NSF grant EAR-2102788. This material is based upon work supported by the U.S. Department of Homeland Security under grant no. 2014-DN-077-ARI088-01. Disclaimer: The views and conclusions contained in this document are those of the authors and should not be interpreted as necessarily representing the official policies, either expressed or implied, of the U.S. Department of Homeland Security. Any use of trade, firm, or product names is for descriptive purposes only and does not imply endorsement by the U.S. Government. J.C., L.E.M., and P.R.R. acknowledge support from NSF grant 2102788.

* distefan@queensu.ca

- H. Ejiri, J. Suhonen, and K. Zuber, Neutrino-nuclear responses for astro-neutrinos, single beta decays and double beta decays, Physics Reports **797**, 1 (2019).
- [2] M. Schumann, Direct detection of WIMP dark matter: concepts and status, J. Phys. G Nucl. Phys. 46, 103003 (2019).
- [3] G. Angloher *et al.* (COSINUS collaboration), Simulation-based design study for the passive shielding of the COSINUS dark matter experiment, Eur. Phys. J. C 82, 248 (2022).
- [4] P. Adhikari *et al.* (COSINE-100 collaboration), Background model for the NaI (Tl) crystals in COSINE-100, Eur. Phys. J. C 78, 490 (2018).
- [5] M. Antonello *et al.* (SABRE collaboration), Monte Carlo simulation of the SABRE PoP background, Astropart. Phys. **106**, 1 (2019).
- [6] J. Amaré et al. (ANAIS-112 collaboration), Annual

modulation results from three-year exposure of ANAIS-112, Phys. Rev. D $\mathbf{103},$ 102005 (2021).

- [7] R. Bernabei *et al.* (DAMA/LIBRA collaboration), First model independent results from DAMA/LIBRA– Phase2, Universe 4, 116 (2018).
- [8] J. Pradler, B. Singh, and I. Yavin, On an unverified nuclear decay and its role in the DAMA experiment, Phys. Lett. B 720, 399 (2013).
- [9] G. J. Wasserburg and R. J. Hayden, A⁴⁰-K⁴⁰ dating, Geochim. Cosmochim. Acta 7, 51 (1955).
- [10] C. Merrihue and G. Turner, Potassium-argon dating by activation with fast neutrons, J. Geophys. Res. 71, 2852 (1966).
- [11] Th. F. Nägler and I. M. Villa, In pursuit of the ⁴⁰K branching ratios: K-Ca and ³⁹Ar-⁴⁰Ar dating of gem silicates, Chem. Geol. **169**, 5 (2000).
- [12] F. Begemann, K. R. Ludwig, G. W. Lugmair, K. Min, L. E. Nyquist, P. J. Patchett, P. R. Renne, C.-Y. Shih, I. M. Villa, and R. J. Walker, Call for an improved set of decay constants for geochronological use, Geochim. Cosmochim. Acta 65, 111 (2001).
- [13] K. Min, R. Mundil, P. R. Renne, and K. R. Ludwig, A test for systematic errors in ⁴⁰Ar/³⁹Ar geochronology through comparison with U/Pb analysis of a 1.1-Ga rhyolite, Geochim. Cosmochim. Acta 64, 73 (2000).
- [14] P. R. Renne, R. Mundil, G. Balco, K. Min, and K. R. Ludwig, Joint determination of ⁴⁰K decay constants and ⁴⁰Ar*/⁴⁰K for the Fish Canyon sanidine standard, and improved accuracy for ⁴⁰Ar/³⁹Ar geochronology, Geochim. Cosmochim. Acta 74, 5349 (2010).
- [15] J. Carter, R. B. Ickert, D. F. Mark, M. M. Tremblay, A. J. Cresswell, and D. C. W. Sanderson, Production of ⁴⁰Ar by an overlooked mode of ⁴⁰K decay with implications for K-Ar geochronology, Geochronology 2, 355 (2020).
- [16] B. Singh, J. L. Rodriguez, S. S. M. Wong, and J. K. Tuli, Review of log ft values in β decay, Nuclear Data Sheets 84, 487 (1998).
- [17] X. Mougeot, Improved calculations of electron capture transitions for decay data and radionuclide metrology, Appl. Radiat. Isot. 134, 225 (2018).
- [18] X. Mougeot, Towards high-precision calculation of electron capture decays, Appl. Radiat. Isot. 154, 108884 (2019).
- [19] D. W. Engelkemeir, K. F. Flynn, and L. E. Glendenin, Positron Emission in the Decay of K⁴⁰, Phys. Rev. **126**, 1818 (1962).
- [20] N. B. Gove and M. J. Martin, log-f tables for beta decay, Atomic Data and Nuclear Data Tables 10, 205 (1971).
- [21] LOGFT (accessed August 28th, 2020), with original code from Gove and Martin 1971. available at: https://www. nndc.bnl.gov/logft/.
- [22] J. Chen, Nuclear data sheets for A= 40, Nuclear Data Sheets 140, 1 (2017).
- [23] K. Kossert, Y. Amelin, D. Arnold, R. Merle, X. Mougeot, M. Schmiedel, and D. Zapata-García, Activity standardization of two enriched ⁴⁰K solutions for the determination of decay scheme parameters and the half-life, Appl. Radiat. Isot. 188, 110362 (2022).
- [24] M. Wang, W. J. Huang, F. G. Kondev, G. Audi, and S. Naimi, The AME 2020 atomic mass evaluation (II). Ta-

bles, graphs and references, Chinese Physics C **45**, 030003 (2021).

- [25] P. C. F. Di Stefano *et al.* (KDK collaboration), The KDK (potassium decay) experiment, in *J. Phys. Conf. Ser.*, Vol. 1342 (IOP Publishing, 2020) p. 012062.
- [26] M. Stukel *et al.* (KDK collaboration), A novel experimental system for the KDK measurement of the ⁴⁰K decay scheme relevant for rare event searches, Nucl. Instrum. Methods Phys. Res. A **1012**, 165593 (2021).
- [27] L. Hariasz *et al.*, First observation of the ground state electron-capture of 40 K, Submitted to Phys. Rev. C (2022).
- [28] P. Lechner *et al.*, Silicon drift detectors for high resolution room temperature X-ray spectroscopy, Nucl. Instrum. Methods Phys. Res. A **377**, 346 (1996).
- [29] M. Karny, K. P. Rykaczewski, A. Fijałkowska, B. C. Rasco, M. Wolińska-Cichocka, R. K. Grzywacz, K. C. Goetz, D. Miller, and E. F. Zganjar, Modular total absorption spectrometer, Nucl. Instrum. Methods Phys. Res. A 836, 83 (2016).
- [30] B. Hoeling, D. Reed, and P. B. Siegel, Going bananas in the radiation laboratory, Am. J. Phys. 67, 440 (1999).
- [31] X. Mougeot, BetaShape: A new code for improved analytical calculations of beta spectra, in *EPJ Web of Conferences*, Vol. 146 (EDP Sciences, 2017) p. 12015.
- [32] M. M. Bé et al., Table of radionuclides (Vol. 5–A= 22 to 244) (2010).
- [33] P. R. Bell and J. M. Cassidy, Gamma-rays of K⁴⁰, Phys. Rev. 77, 409 (1950).
- [34] P. R. Bell and J. M. Cassidy, Measurement of the gammaray energy of K⁴⁰, Phys. Rev. **79**, 173 (1950).
- [35] S. A. Colgate, The positron activity of K⁴⁰, Phys. Rev. 81, 1063 (1951).
- [36] D. R. Tilley and L. Madansky, Search for positron emission in K⁴⁰, Phys. Rev. **116**, 413 (1959).
- [37] H. Leutz, G. Schulz, and H. Wenninger, The decay of potassium-40, Zeitschrift für Physik 187, 151 (1965).
- [38] E. M. Niespolo, D. Rutte, A. L. Deino, and P. R. Renne, Intercalibration and age of the Alder Creek sanidine ⁴⁰Ar/³⁹Ar standard, Quat. Geochronol. **39**, 205 (2017).
- [39] I. McDougall and P. Wellman, Calibration of GA1550 biotite standard for K/Ar and ⁴⁰Ar/³⁹Ar dating, Chem. Geol. 280, 19 (2011).
- [40] K. F. Kuiper, A. Deino, F. J. Hilgen, W. Krijgsman, P. R. Renne, and J. R. Wijbrans, Synchronizing rock clocks of earth history, Science **320**, 500 (2008).
- [41] P. R. Renne, ⁴⁰Ar/³⁹Ar age of plagioclase from Acapulco meteorite and the problem of systematic errors in cosmochronology, Earth and Planetary Science Letters 175, 13 (2000).
- [42] C. Göpel and G. Manhès, The thermal history of the Acapulco meteorite and its parent body deduced from U/Pb systematics in mineral separates and bulk rock fragments, Comptes Rendus Geoscience 342, 53 (2010).
- [43] A. Prinzhofer, D. A. Papanastassiou, and G. J. Wasserburg, Samarium-neodymium evolution of meteorites, Geochim. Cosmochim. Acta. 56, 797 (1992).
- [44] R. H. Steiger and E. Jäger, Subcommission on geochronology: Convention on the use of decay constants in geo-and cosmochronology, EPSL 36, 359 (1977).