

## CHCRUS

This is the accepted manuscript made available via CHORUS. The article has been published as:

## Highly Resolved Measurements of a Developing Strong Collisional Plasma Shock

Hans G. Rinderknecht, H.-S. Park, J. S. Ross, P. A. Amendt, D. P. Higginson, S. C. Wilks, D.
Haberberger, J. Katz, D. H. Froula, N. M. Hoffman, G. Kagan, B. D. Keenan, and E. L. Vold Phys. Rev. Lett. **120**, 095001 — Published 2 March 2018 DOI: 10.1103/PhysRevLett.120.095001

## 1

2

3

4

5

6

7

8

q

10

11

12

13

14

15

## Highly-resolved measurements of a developing strong collisional plasma shock

Hans G. Rinderknecht,<sup>1,\*</sup> H.-S. Park,<sup>1</sup> J.S. Ross,<sup>1</sup> P.A. Amendt,<sup>1</sup> D.P. Higginson,<sup>1</sup> S.C. Wilks,<sup>1</sup> D.

Haberberger,<sup>2</sup> J. Katz,<sup>2</sup> D.H.Froula,<sup>2</sup> N.M. Hoffman,<sup>3</sup> G. Kagan,<sup>3</sup> B.D. Keenan,<sup>3</sup> and E.L. Vold<sup>3</sup>

<sup>1</sup>Lawrence Livermore National Laboratory, Livermore, California 94550

<sup>2</sup>Laboratory for Laser Energetics, Rochester, New York 14623

<sup>3</sup>Los Alamos National Laboratory, Los Alamos, NM 87545

(Dated: February 2, 2018)

The structure of a strong collisional shock front forming in a plasma is directly probed for the first time in laser-driven gas-jet experiments. Thomson scattering of a 526.5 nm probe beam was used to diagnose temperature and ion velocity distribution in a strong shock ( $M \sim 11$ ) propagating through a low-density ( $\rho \sim 0.01 \text{ mg/cc}$ ) plasma composed of hydrogen. A forward-streaming population of ions traveling in excess of the shock velocity was observed to heat and slow down on an unmoving, pre-shocked population of cold protons, until ultimately the populations merge and begin to thermalize. Instabilities are observed during the merging, indicating a uniquely plasma-phase process in shock front formation.

Shocks are ubiquitous phenomena in high-energy-16 density (HED) plasmas, and are important both in as-17 trophysics and laser-plasma experiments such as inertial 18 confinement fusion (ICF). At distances large relative to 19 the shock front width  $\Delta x$ , the shocked plasma state may 20 be calculated from the unshocked (upstream) density and 21 pressure and the shocked (downstream) fluid velocity.[1] 22 However, conditions near the shock front are often im-23 portant: for example, the radially-converging shock in 24 an ICF implosion inevitably violates this condition when 25 reaching a radius  $R \leq \Delta x$ . Moreover, hydrodynamic 26 treatments[2] are insufficient to calculate the structure 27 of a strong shock front with Mach number  $M \gtrsim 1.5$ , de-28 fined as the ratio of the shock velocity to the upstream 29 sound speed  $(u_{sh}/c_s)$ .[3] 30

In strong collisional plasma shocks, kinetic ion dis-31 tributions at the discontinuity extend to tens of times 32 the ion thermal mean-free-path in the shocked plasma 33  $(\lambda_{ii})$ .[4, 5] Strong collisional[6–8], collisionless,[9–16] and 34 magnetized shocks<sup>[17]</sup> have been studied in laboratory<sup>55</sup> 35 plasmas, but despite substantial theoretical effort, few <sup>56</sup> 36 measurements of collisional shock-front structure – pro- 57 37 files of temperature, density, and velocity distribution 58 38 within the front – have been performed. [18] Profiles of 59 39 electric field were recently measured in strong plasma 60 40 shocks; [19] however, the extremely small length- and 61 41 time-scales of collisions in most experimental plasmas 62 42 make measurements of shock structure particularly diffi- 63 43 cult. 44

This Letter presents the first measurements of strong, <sup>65</sup> 45 collisional plasma shock-front structure in the formation <sup>66</sup> 46 stage. In experiments using the OMEGA laser, [20] strong <sup>67</sup> 47 shocks  $(M \sim 11)$  were driven into a volume of hydrogen <sup>68</sup> 48 gas, injected by a gas-jet system prior to the laser fir- 69 49 ing. The volume was interrogated using a 526.5 nm (2 $\omega$ ) 70 50 probe beam impulse.[21] The Thomson-scattered light 71 51 from this probe was imaged and used to infer spatially-72 52 resolved temperature, density, and flow velocity within 73 53 the shock front. [22] These experiments demonstrate for 74 54

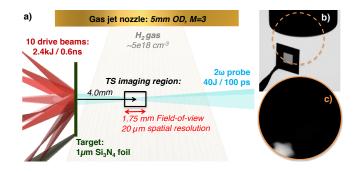


FIG. 1. (color online) (a) Experimental design to probe plasma shock front structure. Laser beams drive a Si<sub>3</sub>N<sub>4</sub> foil, launching a strong shock into a H<sub>2</sub> gas jet. Thomsonscattered light from a  $2\omega$  beam aligned with the foil axis is imaged in a 1.75 mm region, 4.0 mm from the foil. (b) View of the target foil aligned near the gas-jet nozzle. (c) Pinhole camera image showing x-rays ( $h\nu > 1.5$  keV) from the foil.

the first time ion velocity separation within a plasma shock.

In general, strong collisional plasma shock formation can be understood as follows. Electron conduction creates a preheat layer with increased electron temperature ahead of the density jump. The thickness of this layer is predicted to be ~  $\lambda_{ii}\sqrt{m_i/m_e}$ , set by the difference in electron and ion thermal velocities.[23] The increased temperature reduces stopping power in the preheat region, allowing the most energetic shocked ions with the largest mean-free-paths  $[\lambda(\epsilon) = (\epsilon/T_i)^2 \lambda_{ii}]$  to stream forward. The balance of ion stopping power, electron-ion thermalization, ion-ion drag and collisional heating establishes the shock-front structure.

The experimental layout is shown in Figure 1. A gasjet nozzle (5 mm diameter) injects a Mach-3 cone of hydrogen gas (atomic density  $\sim 5 \times 10^{18}$  cm<sup>-3</sup>) into the target chamber. Ten laser beams containing 2.4 kJ in a 1 mm diameter spot drive a 1 µm-thick silicon nitride ablator foil positioned near the gas jet, with a max-

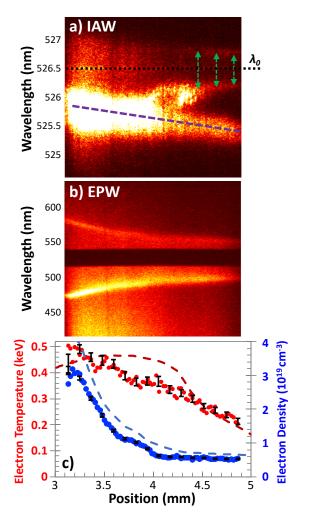


FIG. 2. (color online) Thomson scattering images of the shocked plasma: (a) IAW and (b) EPW features recorded  $^{\rm 46}$ 4.1 ns after the laser drive. Spatial dimension is horizontal,  $^{\rm 47}$ with the shock propagating to the right. The IAW image 48 captures the shock transition, showing a forward-streaming 49 population (purple) slowing on and heating a cold, still pop- 50 ulation (green). (c) Fits to the EPW spectra provide electron  $_{51}$ temperature (red) and density (blue), showing the preheat 52 region followed by the density jump. Error bars indicate typical fitting uncertainty. Trends from PIC simulations (dashed)  $^{53}$ capture density jump position and temperature magnitude. 55

imum intensity of 500 TW/cm<sup>2</sup>. The driven foil ex-<sup>57</sup> 1 plodes, launching a strong shock into the gas. After 58 2 a delay of 4.1 ns, a  $2\omega$  laser impulse containing 40 J 3 in 100 ps is injected along the foil axis to probe the 4 state of the plasma by Thomson scattering. The di-5 mensionless scattering parameter  $\alpha \equiv 1/k\lambda_{Debye} > 1$ , 6 so the spectrum is dominated by collective scattering. [24] 7 The scattered spectrum was recorded using both narrow-8 and wide-band spectrometers, for the ion acoustic wave 9 (IAW) and electron plasma wave (EPW) features, re-10 spectively. The IAW feature encodes information about 11 the flow velocity, density, and temperature of ion popu-12 lations, whereas the EPW feature encodes the electron 13

density and temperature. [25] The spectra were imaged 14 along one spatial dimension, recording a 1.75 mm fieldof-view along the probe axis with 20  $\mu$ m resolution.[26] 16 The scattering k-vector was oriented  $\sim 60^{\circ}$  from the direction of flow. In this geometry, plasma flowing away from the ablator produces a blue-shift in the scattered 19 light. X-ray pinhole cameras recorded self-emission from the irradiated targets, confirming the target survives exposure to the gas jet (Fig. 1c); x-rays from the target preheat the hydrogen gas ahead of the shock. 23

15

17

18

20

21

22

24

25

26

27

28

29

30

31

32

33

34

35

36

37

38

39

40

41

42

43

44

Figures 2a-b show Thomson scattering images of the shocked plasma. These images record the formation of the shock front, and notably resolve the ion velocityspace evolution within the shock front. To the best of our knowledge, these data represent the first record of velocity structure within a strong collisional plasma shock. A qualitative discussion reveals many interesting details of strong plasma shock formation: ions streaming in excess of the shock velocity (dashed purple line) first interact with the cold proton gas (two peaks symmetric around the initial wavelength, indicating negligible flow velocity). This interaction heats the pre-shocked ions (green arrows, broadening of unshocked feature) and electrons (broadening of electron feature). The streaming and unstreaming populations merge at a flow velocity of  $\sim 750 \mu m/ns$ , with rapid changes in the IAW feature.

This qualitative picture is supported by quantitative analysis, performed by forward-fitting a scatteredspectrum model to the data. [24] Figure 2c shows fits to the EPW data. Fitting uncertainty was calculated using a reduced chi-squared method. This data shows the characteristic shock features: [5] an electron preheat layer is observed in which the temperature increases from 220 to 350 eV, followed by an increase in density from 0.5to  $3.1 \times 10^{19}$  cm<sup>-3</sup> and temperature to 480 eV. Moreover, the IAW spectra in the electron preheat region verify the primary prediction of the kinetic theory of strong shocks: an energetic, forward-streaming ion population extends throughout the electron preheat layer.

Using these measurements, plasma parameters of interest are calculated in both the preheat and shock region (Table I). The thermal mean-free-path for hydrogen ions in the shocked plasma is estimated to be  $170 \pm 20 \mu m$ : over  $10^3$  times the estimated Debye length, confirming the assumption of quasineutrality for this plasma.

TABLE I. Plasma parameters in shock and preheat regions. Parameter Shock Preheat

Sound speed, $c_s$ ( $\mu$ m/ns)	285	90
Debye length, $\lambda_{De}$ (µm)	0.019	0.016
Plasma frequency, $\omega_{pe}  (\mathrm{ns}^{-1})$	$3{\times}10^5$	$1.3{ imes}10^5$
Ion skin depth, $c/\omega_{pi}$ (µm)	47	99
Hydrogen thermal mean-free-path, $\lambda_H$ (µm)	170	10
Flow velocity, v ( $\mu m/ns$ )	750	0

Kinetic theory [4, 27] and simulations [5] suggest that 1 the width of the electron preheat region should be  $\Delta x \approx$ 2  $\lambda_{ii}\sqrt{m_i/m_e} \approx 8$  mm. This prediction exceeds the in-3 strument field of view, and indeed the distance from 4 the foil to the imaged area. While the images do not 5 record the entire electron preheat region, linear extrap-6 olation implies the region extends  $\sim 1$  mm beyond the 7 imaged area, indicating a shock width of 3 mm: ap-8 proximately one-third of the predicted scaling. The rea-9 son for this discrepancy is likely that the shock has not 10 vet fully formed. Vidal et al.<sup>[5]</sup> show that a shock re-11 quires spatial separation of approximately  $2\Delta x$  from the 12 pusher to reach steady state, before which the shock is in 13 a transient formation stage, asymptotically approaching 14 its final width from below. Given the velocity difference 15 between the front and the shocked fluid  $[(u_{sh} - v_1)] =$ 16  $u_{sh}(\gamma - 1)/(\gamma + 1) \approx u_{sh}/4$ , the time required to reach 17 a steady state is estimated as  $t_S \approx 8\Delta x/u_{sh}$ . Taking 18  $u_{sh} = x_{meas}/t_{meas} \sim 1000 \ \mu {\rm m/ns},$  the steady-state time 19 is  $t_S \approx 24$  ns, six times longer than the sample time. 20 This estimate corroborates our observation of a narrow 21 shock compared to the analytical prediction. 22

A Thomson scattering model with multiple ion popu-23 lations was forward-fit to interpret the IAW data. Fig-24 ure 3 shows representative fits in the preheat and shock 25 regions, and the results of these fits. The model includes 26 the effects of finite optics. [28] [29] In the preheat (4.5– 27 5 mm) and density-jump (3–4 mm) regions, the hydrogen 28 distribution is well represented as the sum of two flow-29 ing Maxwellians: a hot population streaming forward at 30  $\sim 1000 \ \mu m/ms$ , and a cold population not flowing with 31 respect to the lab frame. Additionally, a small popula-32 tion of forward-streaming Si and N ions is required to 33 match the narrow, blue-shifted peak observed through-34 out the imaged region. Accurate fits are obtained assum-35 ing these ions share the same composition as the fully-36 ionized target foil.[30] Given the relatively high density 37 of the Si<sub>3</sub>N<sub>4</sub> pusher, some pusher ions are predicted to 38 stream into the low-density hydrogen plasma while the 39 shock forms. [31] The velocity of the  $Si_3N_4$  ions is similar 40 to the free-streaming velocity (Fig. 3e). 41

Within the preheat region, the calculated sound-42 speed of the background protons  $(T_i = 50 \text{ eV} \Rightarrow c_s =$ 43 90  $\mu$ m/ns), confirms a strong shock ( $M \approx 11$ ). The hot 44 streaming population constitutes approximately 20% of 45 the protons. These are substantially hotter ( $\sim 3 \text{ keV}$ )<sup>57</sup> 46 and faster ( $\sim 1200 \ \mu m/ns$ ) than the shocked plasma, and <sup>58</sup> 47 are also hotter than the electrons. The high fraction <sup>59</sup> 48 and temperature of the streaming population is consis- 60 49 tent with a forming shock front: in steady-state,  $T_e > T_i^{61}$ 50 and  $f_{hot} \sim 10\%$  in the preheat region, whereas in a  $^{62}$ 51 forming shock the streaming ions are hotter and more 63 52 numerous.[5] 53

At the beginning of the density jump, the data show 65 acceleration and heating of the cold protons as the hot 66 population slows. The hot population density exceeds 67

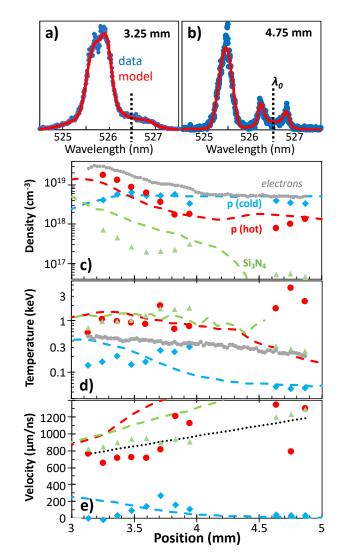


FIG. 3. (color online) (a,b) Example lineouts and fits to IAW spectra in the shock and preheat region. (c–e) Density, temperature, and velocity from fits to the IAW spectra. Data was accurately matched using three ion populations: cold background hydrogen (blue diamonds), hot streaming hydrogen (red circles), and streaming ablator ions (green triangles). Electron density and temperature (grey) and free-streaming velocity (black dotted) are included for reference. Results of PIC simulations (dashed) capture trends in proton density and temperature.

the cold population near 3.5 mm, and continues to grow toward the foil, reaching an asymptotic temperature ( $\sim$ 1 keV) and velocity ( $\sim$ 700 $\mu$ m/ns). In contrast, the cold population density drops as the ion shock forms. Behind the density jump, the remaining cold population ( $\sim$ 20%) has increased in temperature from the pre-shock value by 4×.

Immediately ahead of the density jump (4.0–4.5 mm), the scattered light spectrum varies rapidly, and accurate fits of the three-population model could not be found. In particular, near  $\lambda = 526$  nm a flashing pattern is

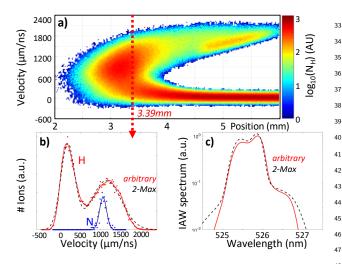


FIG. 4. (color online) Results from PIC simulation: (a)<sup>40</sup> Hydrogen velocity-position phase-space at the experimental<sup>49</sup> sample time. (b) Hydrogen (red) and Nitrogen distributions<sup>50</sup> (blue), 3.39 mm from the foil, with arbitrary (red solid) and <sup>51</sup> 2-Maxwellian (black dashed) fits to the H and Maxwellian fit <sup>52</sup> (blue solid) to N. (c) Simulated Thomson-scattering spectrum <sup>53</sup> at 3.39 mm using arbitrary and thermal fits.<sup>54</sup>

55

56

observed: four peaks separated by  $40 \pm 10 \mu m$ , with in- 57 1 creasing brightness toward the density jump. Vidal et 58 2 al. [5] report that electrostatic instability growth is possi-3 ble within a Mach 5 shock for times  $0.15 < t/t_S < 0.43$ . 4 The present experiment falls within this range, suggest- 61 5 ing the flashing is a signature of instability growth. The 62 6 fastest-growing wavelength of the ion-ion two-stream in-7 stability in this plasma is on the order of a few microns, <sub>64</sub> 8 smaller than the instrument resolution and thus not di-  $_{65}$ 9 rectly observable.[32] However the growth rate of such 66 10 modes is rapid ( $\sim \omega_{pi} > 300 \text{ ns}^{-1}$ ). The intensity of the  $_{67}$ 11 feature grows exponentially as ~ exp $(x[20 \ \mu m^{-1}])$ . Com-12 paring to the shock velocity, this implies a growth rate 69 13 on the order of  $10^4$  ns<sup>-1</sup>. 14 70

For comparison to the data, a simulation was per-71 15 formed using the particle-in-cell (PIC) code LSP.[33] 72 16 This simulation is similar to Ref. [34], with fully-73 17 implicit kinetic ions undergoing binary collisions, and 74 18 fluid electrons. [35] A Si<sub>3</sub>N<sub>4</sub> plasma pushes a 50 eV, 75 19  $5 \times 10^{18}$  cm<sup>-3</sup> hydrogen plasma. The initial conditions 76 20 for the  $Si_3N_4$  were taken from a planar 1D simulation  $_{77}$ 21 using the radiation-hydrodynamic code HYADES[36] at 78 22 1.0 ns. The ion distributions produced 4.1 ns after the  $_{79}$ 23 initial laser drive were fit using a bi-Maxwellian model for 80 24 the hydrogen and a thermal model for the silicon and ni- 81 25 trogen, as shown in Figure 4a–b; the results are shown as <sup>82</sup> 26 dashed lines in Figs. 2 and 3. The model captures many <sup>83</sup> 27 features observed in the data, including the electron den- 84 28 sity and temperature (Fig. 2c), the relative trends of den- 85 29 sity in the hot- and cold-proton populations (Fig. 3c), 86 30 and the heating and velocity of the cold population (d, 87 31 e). However the model fails to capture the density of the \*\* 32

 $\rm Si_3N_4$  ions, and the velocities of the forward-streaming ions (which are consistent with the free-streaming velocity). It is notable that, while the simulated electron density jump leads the data by 0.1–0.2 mm, the increase in density of the hot protons lags behind the data by 0.2 mm. This discrepancy indicates that the code is not accurately capturing the dynamics of shock stagnation, which is more rapid in the experiment.

The simulations confirm bi-Maxwellian fits are a reasonable approximation. However, the Thomson scattering form factor is in principle sensitive to non-thermal distributions. To assess whether this affects the data, a non-thermal model was developed, composed of evenlyspaced Maxwellians with amplitudes fit to the simulated hydrogen distribution. Fig. 4b-c shows a best fit of this arbitrary model (using 20 peaks) and its effect on the Thomson-scattered spectrum. This model performed better than the 2-Maxwellian fit primarily within the density jump (3.2–3.6 mm in the simulation, 3.4–3.8 mm in the data). In this region, the 2-Maxwellian fit overestimates the tail of the distribution: with the higherresolution model, the edges of the scattered light spectrum fall off more rapidly. Notably, the 2-Maxwellian model is highly accurate ahead of the density jump (3.7– 4.2 mm in the simulation, 3.9–4.4 mm in the data), despite the fact that accurate fits to the data could not be found. This discrepancy, in combination with the more rapid increase in the hot proton fraction in the data, suggests that instability in the experiment stagnates the ion flows more efficiently than the simulation predicts.

In conclusion, the structure of a strong  $(M \approx 11)$ collisional plasma shock front has been measured for the first time using optical Thomson scattering. Threepopulation fits to the data demonstrate the kinetic structure of strong-shock formation: a hot population of ions streaming through the cold background in the preheat region, heating and drag of the cold background, and rapid increase of the hot population within the density jump. The relatively short preheat region (~  $\lambda_H \sqrt{m_i/m_e/3}$ ) observed in data and simulation confirms that strong shocks approach their steady-state width from below. Kinetic simulations reproduce the density, electron temperature, and proton population trends observed in the data, but do not capture the flashing observed prior to the density jump, and under-predict the rate of shock stagnation, suggesting two-stream instability plays a role in the shock formation. This data provides an unprecedented level of detail in examining ion collisional processes in highenergy-density plasmas. In future studies, the ion meanfree-path will be controlled to obtain a scaling of shock width and formation time with Mach number. Simultaneous measurement of electric field structures [19] will further improve understanding of these phenomena. The relatively long time- ( $\sim 20$  ns) and length-scales ( $\sim cm$ ) needed for full shock formation may require laser energy on the scale of the National Ignition Facility, [37]

where an Optical Thomson Scattering (OTS) diagnostic <sup>59</sup>
 is now available. [38, 39] This research program offers a <sup>60</sup>
 new challenge to high-fidelity physics codes, for improved <sup>61</sup>
 accuracy in modeling plasmas of interest to ICF and lab <sup>62</sup> oratory astrophysics.
 The authors would like to acknowledge valuable con-

versations with Luis Chacon, Andrei Simakov and Will

<sup>8</sup> Taitano. This work performed under the auspices of the <sup>67</sup>

<sup>8</sup> Taitano. This work performed under the auspices of the <sup>67</sup>

<sup>9</sup> U.S. Department of Energy by LLNL under Contract <sup>68</sup>
 <sup>10</sup> DE-AC52-07NA27344, and supported by the LLNL Lab-<sup>69</sup>

<sup>10</sup> DE-AC52-07NA27344, and supported by the LLNL Lab-<sup>69</sup> <sup>11</sup> oratory Directed Research and Development Program<sup>70</sup>

<sup>11</sup> oratory Directed Research and Development Program <sup>71</sup> (LDRD) (17-ERD-060, 18-ERD-047). Laser system time

 $_{13}$  provided by the LBS program at the LLE Omega Facility.  $_{73}^{72}$ 

<sup>14</sup> \* rinderknecht1@llnl.gov

- [1] W. J. M. Rankine, Philosophical Transactions of the <sup>79</sup> Royal Society of London 160, 277 (1870).
- [2] M. Y. Jaffrin and R. F. Probstein, Physics of Fluids 7, 81
   1658 (1964).
- [3] B. D. Keenan, A. N. Simakov, L. Chacon, and W. T. 83
   Taitano, Physical Review E 96, 053203 (2017).
- <sup>21</sup> [4] M. S. Greywall, Physics of Fluids **18**, 1439 (1975).
- [5] F. Vidal, J. P. Matte, M. Casanova, and O. Larroche, <sup>86</sup>
   Physics of Fluids B: Plasma Physics (1989-1993) 5, 3182 <sup>87</sup>
   (1993).
- [6] R. A. Bosch, R. L. Berger, B. H. Failor, N. D. Delamater, <sup>89</sup>
   G. Charatis, and R. L. Kauffman, Physics of Fluids B: <sup>90</sup>
   Plasma Physics 4, 979 (1992). <sup>91</sup>
- [7] E. C. Merritt, A. L. Moser, S. C. Hsu, J. Loverich, and <sup>92</sup>
   M. Gilmore, Physical Review Letters **111**, 085003 (2013). <sup>93</sup>
- [8] A. L. Moser and S. C. Hsu, Physics of Plasmas 22, 055707 94
   (2015). 95
- [9] A. R. Bell, P. Choi, A. E. Dangor, O. Willi, D. A. Bassett, 96
   and C. J. Hooker, Physical Review A 38, 1363 (1988). 97
- [10] R. L. Berger, J. R. Albritton, C. J. Randall, E. A. 96
   Williams, W. L. Kruer, A. B. Langdon, and C. J. Hanna, 99
   Physics of Fluids B: Plasma Physics 3, 3 (1991). 100
- I11] L. Romagnani, S. V. Bulanov, M. Borghesi, P. Aude-101
   bert, J. C. Gauthier, K. Löwenbrück, A. J. Mackinnon,102
   P. Patel, G. Pretzler, T. Toncian, and O. Willi, Physical103
   Review Letters 101, 025004 (2008). 104
- [12] Y. Kuramitsu, Y. Sakawa, T. Morita, C. D. Gregory,<sup>105</sup>
   J. N. Waugh, S. Dono, H. Aoki, H. Tanji, M. Koenig,<sup>106</sup>
   N. Woolsey, and H. Takabe, Physical Review Letters<sup>107</sup>
   106, 175002 (2011).
- [13] H.-S. Park, D. D. Ryutov, J. S. Ross, N. L. Kugland, S. H.<sup>109</sup>
  Glenzer, C. Plechaty, S. M. Pollaine, B. A. Remington,<sup>110</sup>
  A. Spitkovsky, L. Gargate, G. Gregori, A. Bell, C. Mur-<sup>111</sup>
  phy, Y. Sakawa, Y. Kuramitsu, T. Morita, H. Takabe,<sup>112</sup>
  D. H. Froula, G. Fiksel, F. Miniati, M. Koenig, A. Rava-<sup>113</sup>
- sio, A. Pelka, E. Liang, N. Woolsey, C. C. Kuranz, R. P.<sup>114</sup>
   Drake, and M. J. M. J. Grosskopf, High Energy Density<sup>115</sup>
- <sup>52</sup> Physics **8**, 38 (2012). 116
- [14] H. Ahmed, M. E. Dieckmann, L. Romagnani, D. Do-117
   ria, G. Sarri, M. Cherchez, E. Ianni, I. Kourakis, A. L.118
   Giesecke, M. Notley, R. Prasad, K. Quinn, O. Willi, and119
   M. Borghesi, Physical Review Letters 110, 205001 (2013).120
- <sup>57</sup> [15] J. S. Ross, D. P. Higginson, D. Ryutov, F. Fiuza,<sub>121</sub> <sup>58</sup> R. Hatarik, C. M. Huntington, D. H. Kalantar, A. Link,

B. B. Pollock, B. A. Remington, H. G. Rinderknecht,
G. F. Swadling, D. P. Turnbull, S. Weber, S. Wilks, D. H.
Froula, M. J. Rosenberg, T. Morita, Y. Sakawa, H. Takabe, R. P. Drake, C. Kuranz, G. Gregori, J. Meinecke,
M. C. Levy, M. Koenig, A. Spitkovsky, R. D. Petrasso,
C. K. Li, H. Sio, B. Lahmann, A. B. Zylstra, and H. S.
Park, Physical Review Letters 118, 185003 (2017).

- [16] T. Morita, Y. Sakawa, Y. Kuramitsu, S. Dono, H. Aoki, H. Tanji, J. N. Waugh, C. D. Gregory, M. Koenig, N. C. Woolsey, and H. Takabe, Physics of Plasmas 24, 072701 (2017).
- [17] S. V. Lebedev, L. Suttle, G. F. Swadling, M. Bennett, S. N. Bland, G. C. Burdiak, D. Burgess, J. P. Chittenden, A. Ciardi, A. Clemens, P. De Grouchy, G. N. Hall, J. D. Hare, N. Kalmoni, N. Niasse, S. Patankar, L. Sheng, R. A. Smith, F. Suzuki-Vidal, J. Yuan, A. Frank, E. G. Blackman, and R. P. Drake, Physics of Plasmas **21**, 056305 (2014).
- [18] R. E. Center, Physics of Fluids 10, 1777 (1967).

74

75

76

77

78

85

- [19] R. Hua, H. Sio, S. C. Wilks, F. N. Beg, C. McGuffey, M. Bailly-Grandvaux, G. W. Collins, and Y. Ping, Applied Physics Letters **111**, 034102 (2017).
- [20] T. R. Boehly, D. L. Brown, R. S. Craxton, R. L. Keck, J. P. Knauer, J. H. Kelly, T. J. Kessler, S. A. Kumpan, S. J. Loucks, S. A. Letzring, F. J. Marshall, R. L. Mc-Crory, S. F. B. Morse, W. Seka, J. M. Soures, and C. P. Verdon, Optics Communications 133, 495 (1997).
- [21] A. J. Mackinnon, S. Shiromizu, G. Antonini, J. Auerbach, K. Haney, D. H. Froula, J. Moody, G. Gregori, C. Constantin, C. Sorce, L. Divol, R. L. Griffith, S. Glenzer, J. Satariano, P. K. Whitman, S. N. Locke, E. L. Miller, R. Huff, K. Thorp, W. Armstrong, R. Bahr, W. Seka, G. Pien, J. Mathers, S. Morse, S. Loucks, and S. Stagnitto, Review of Scientific Instruments **75**, 3906 (2004).
- [22] J. Katz, J. S. Ross, C. Sorce, and D. H. Froula, Journal of Instrumentation 8, C12009 (2013).
- [23] Y. B. Zel'dovich and Y. P. Raizer, *Physics of Shock Waves and High-Temperature Hydrodynamic Phenomena*, Dover Books on Physics (Dover Publications, 2002).
- [24] D. H. Froula, S. H. Glenzer, N. C. Luhmann, and J. Sheffield, eds., *Plasma Scattering of Electromagnetic Radiation*, 2nd ed. (Academic Press, Boston, 2011).
- [25] J. S. Ross, H. S. Park, P. Amendt, L. Divol, N. L. Kugland, W. Rozmus, and S. H. Glenzer, Review of Scientific Instruments 83, 10E323 (2012).
- [26] J. S. Ross, D. H. Froula, A. J. Mackinnon, C. Sorce, N. Meezan, S. H. Glenzer, W. Armstrong, R. Bahr, R. Huff, and K. Thorp, Review of Scientific Instruments 77, 10E520 (2006).
- [27] M. Casanova, O. Larroche, and J.-P. Matte, Physical Review Letters 67, 2143 (1991).
- [28] R. K. Follett, J. A. Delettrez, D. H. Edgell, R. J. Henchen, J. Katz, J. F. Myatt, and D. H. Froula, Review of Scientific Instruments 87, 11E401 (2016).
- [29] Velocities are assumed to be along the target axis; the  $\vec{k} \cdot \vec{V}$  projection is removed from the reported value.
- [30] Simulations predict the N/Si ratio is constant (4/3) until the ions slow significantly.
- [31] C. Bellei and P. A. Amendt, Physics of Plasmas 24, 040703 (2017).
- [32] J. D. Huba, "NRL Plasma Formulary," (online), Naval Research Laboratory, Washington, D.C. (2006).

- [33] D. R. Welch, D. V. Rose, B. V. Oliver, and R. E. Clark, 12 1 Nuclear Instruments and Methods in Physics Research 13 2
- Section A: Accelerators, Spectrometers, Detectors and 14 3 Associated Equipment 464, 134 (2001). 4 15
- [34] C. Bellei, H. Rinderknecht, A. Zylstra, M. Rosenberg, 16 5 H. Sio, C. K. Li, R. Petrasso, S. C. Wilks, and P. A. 17 6
- Amendt, Physics of Plasmas 21, 056310 (2014). 7 [35] Flux limiter = 0.2. Simulations with kinetic electrons  $_{19}$ 8
- were highly consistent with the fluid case. 9 20 J. T. Larsen and S. M. Lane, Journal of Quantitative 21
- [36]10 22
- Spectroscopy and Radiative Transfer 51, 179 (1994). 11

[37] E. I. Moses, Nuclear Fusion 49, 104022 (2009).

18

- [38] Ross, J S, Datte, P, Divol, L, Galbraith, J, Froula, D H, Glenzer, S H, Hatch, B, Katz, J, Kilkenny, J, Landen, O, Manuel, A M, Molander, W, Montgomery, D S, Moody, J D, Swadling, G, and Weaver, J, Review of Scientific Instruments 87, 11E510 (2016).
- [39]J. Galbraith, P. Datte, S. Ross, G. Swadling, S. Manuel, B. Molander, B. Hatch, D. Manha, M. Vitalich, and B. Petre, in SPIE Optical Engineering + Applications, edited by J. A. Koch and G. P. Grim (SPIE, 2016) p. 99660E.