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Enhanced γ -Ray Emission from Neutron Unbound States Populated in β Decay

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Total absorption spectroscopy was used to investigate the β -decay intensity to states above the neutron separation energy followed by γ -ray emission in $^{87,88}\text{Br}$ and ^{94}Rb . Accurate results were obtained thanks to a careful control of systematic errors. An unexpectedly large γ intensity was observed in all three cases extending well beyond the excitation energy region where neutron penetration is hindered by low neutron energy. The γ branching as a function of excitation energy was compared to Hauser-Feshbach model calculations. For ^{87}Br and ^{88}Br the γ branching reaches 57% and 20% respectively, and could be explained as a nuclear structure effect. Some of the states populated in the daughter can only decay through the emission of a large orbital angular momentum neutron with a strongly reduced barrier penetrability. In the case of neutron-rich ^{94}Rb the observed 4.5% branching is much larger than the calculations performed with standard nuclear statistical model parameters, even after proper correction for fluctuation effects on individual transition widths. The difference can be reconciled introducing an enhancement of one order-of-magnitude in the photon strength to neutron strength ratio. An increase in the photon strength function of such magnitude for very neutron-rich nuclei, if it proved to be correct, leads to a similar increase in the (n, γ) cross section that would have an impact on r process abundance calculations.

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Neutron unbound states can be populated in the β decay of very neutron-rich nuclei, when the neutron separation energy S_n in the daughter nucleus is lower than the decay energy window Q_β . Given the relative strengths of strong and electromagnetic interactions these states decay preferentially by neutron emission. Beta delayed γ -ray emission from states above S_n was first observed in 1972 in the decay of ^{87}Br [1]. Since then it has been observed in a handful of cases: ^{137}I [2], ^{93}Rb [3], ^{85}As [4], ^{141}Cs [5], ^{95}Rb [6], ^{94}Rb [7], ^{77}Cu [8], and ^{75}Cu [9]. The paucity of information is related to the difficulty of detecting weak high-energy γ -ray cascades with the germanium detectors that are usually employed in β -decay studies. This problem has become known as the *Pandemonium* effect [10] and it also affects the accuracy of the data.

There is an analogy [11] between this decay process and neutron capture reactions which populate states in the compound nucleus that re-emit a neutron (elastic channel) or de-excite by γ rays (radiative capture). Indeed the

reaction cross section is parametrized in terms of neutron and γ widths, Γ_n and Γ_γ respectively, which also determines the fraction of β intensity above S_n that proceeds by neutron or γ emission. Radiative capture (n, γ) cross sections for very neutron-rich nuclei are a key ingredient in reaction network calculations used to obtain the yield of elements heavier than iron in the rapid (r) neutron capture process occurring in explosive-like stellar events. It has been shown [12–14] that the abundance distributions in different astrophysical scenarios are sensitive to (n, γ) cross sections. In the classical “hot” r process late captures during freeze-out modify the final element abundance. In the “cold” r process the competition between neutron captures and β decays determines the formation path. Cross section values for these exotic nuclei are taken from Hauser-Feshbach model calculations [15], which are based on a few quantities describing average nuclear properties: nuclear level densities (NLD), photon strength functions (PSF) and neutron transmission coefficients (NTC). Since these quantities are adjusted to

60 experiment close to β stability it is crucial to find means
61 to verify the predictions for very neutron-rich nuclei.

62 The Total Absorption Gamma-ray Spectroscopy
63 (TAGS) technique aims at detecting cascades rather than
64 individual γ rays using large 4π scintillation detectors.
65 The superiority of this method over high-resolution ger-
66 manium spectroscopy to locate missing β intensity has
67 been demonstrated before [16, 17]. However its appli-
68 cation in the present case is very challenging, since the
69 expected γ -branching is very small and located at rather
70 high excitation energies. As a matter of fact previous
71 attempts at LNPI [7] with a similar aim did not lead to
72 clear conclusions. In this Letter we propose and demon-
73 strate for the first time the use of the TAGS technique
74 to study γ -ray emission above S_n in β -delayed neutron
75 emitters and extract accurate information that can be
76 used to improve (n, γ) cross section estimates far from β
77 stability.

78 Neutron capture and transmission reactions have been
79 extensively used [18] to determine neutron and γ widths
80 (or related strength functions). An inspection of Ref. [18]
81 shows that in general Γ_n is orders-of-magnitude larger
82 than Γ_γ . In the decay of ^{87}Br , which is the best stud-
83 ied case [1, 19–21], a dozen states emitting single γ rays
84 were identified within 250 keV above S_n collecting about
85 0.5% of the decay intensity to be compared with a neu-
86 tron emission probability of 2.6%. The observation of
87 such relatively high γ -ray intensity was explained as be-
88 ing due to a nuclear structure effect: some of the levels
89 populated can only decay by emission of neutrons with
90 large orbital angular momentum l , which is strongly hin-
91 dered. In addition it has been pointed out [22] that a siz-
92 able γ -ray emission from neutron unbound states can be
93 a manifestation of Porter-Thomas (PT) statistical fluc-
94 tuations in the strength of individual transitions. The
95 role and relative importance of both mechanisms should
96 be investigated.

97 We present here the results of measurements for
98 three known neutron emitters, ^{87}Br [23], ^{88}Br [24] and
99 ^{94}Rb [25], using a newly developed TAGS spectrome-
100 ter. The results for ^{93}Rb , also measured, will be pre-
101 sented later [26]. The measurements were performed
102 at the IGISOL mass separator [27] of the University
103 of Jyväskylä. The isotopes were produced by proton-
104 induced fission of uranium and the mass-separated
105 beam was cleaned from isobaric contamination using the
106 JYFLTRAP Penning trap [28, 29]. The resulting beam
107 was implanted at the centre of the spectrometer onto a
108 movable tape which periodically removed the activity to
109 minimize daughter contamination. Behind the tape was
110 placed a 0.5 mm thick Si detector with a β -detection
111 efficiency of about 30%. The Valencia-Surrey Total Ab-
112 sorption Spectrometer *Rocinante* is a cylindrical 12-fold
113 segmented BaF_2 detector with a length and external di-
114 ameter of 25 cm, and a longitudinal hole of 5 cm diam-
115 eter. The separation between crystals is provided by a

116 thin optical reflector. The total efficiency for detecting
117 a single γ ray is larger than 80%. The spectrometer has
118 a reduced neutron sensitivity in comparison to NaI(Tl)
119 detectors, a key feature in the present application. It also
120 allows the measurement of multiplicities which helps in
121 the data analysis. In order to eliminate the detector in-
122 trinsic background and the ambient background we use
123 β -gated TAGS spectra in the present analysis. Neverthe-
124 less other sources of spectrum contamination need to be
125 characterized accurately.

126 In the first place the decay descendant contamination,
127 was computed using the Geant4 simulation toolkit [30].
128 In the case of the daughter decay we use an event gener-
129 ator based on the well known decay level scheme [23–25].
130 The calculated normalization factor was adjusted to pro-
131 vide the best fit to the recorded spectrum. The measure-
132 ment of ^{88}Br was accidentally contaminated by ^{94}Y , the
133 long-lived grand-daughter of ^{94}Rb , and was treated in
134 the same manner. The case of the contamination due
135 to the β -delayed neutron branch is more challenging.
136 The decay simulation must include the correct energy
137 sequence β -neutron- γ . Neutrons interact with detector
138 materials producing additional γ rays through inelastic
139 and capture processes. An event generator was imple-
140 mented which reproduces the known neutron energy dis-
141 tribution, taken from [31], and the known γ -ray intensity
142 in the final nucleus, taken from [23–25]. The event gen-
143 erator requires the β intensity distribution followed by
144 neutron emission $I_{\beta n}$ which was obtained from deconvo-
145 lution of the neutron spectrum. Another issue is whether
146 the interaction of neutrons with the detector can be sim-
147 ulated accurately. We have shown recently [32] that this
148 is indeed the case provided that Geant4 is updated with
149 the newest neutron data libraries and the original cap-
150 ture cascade generator is substituted by an improved one.
151 The normalization factor of the β -delayed neutron decay
152 contamination is fixed by the P_n value. Another impor-
153 tant source of spectrum distortion is the summing-pileup
154 of events. If more than one event arrives within the same
155 ADC event gate, a signal with the wrong energy is stored
156 in the spectrum. Apart from the electronic pulse pile-up
157 effect for a single detector module [33] one must consider
158 the summing of signals from different detector modules.
159 A new Monte Carlo (MC) procedure to calculate their
160 combined contribution has been developed. The proce-
161 dure is based on the random superposition of two stored
162 events within the ADC gate length. The normalization
163 of the resulting summing-pileup spectrum is fixed by the
164 event rate and the ADC gate length [33].

165 Several laboratory γ -ray sources were used to de-
166 termine the energy and resolution calibration of the
167 spectrometer. The highest calibration point was at
168 4.123 MeV. The measured singles spectra also served to
169 verify the accuracy of the spectrometer response simu-
170 lated with Geant4. The use of β -gated spectra in the
171 analysis required additional verifications of the simula-

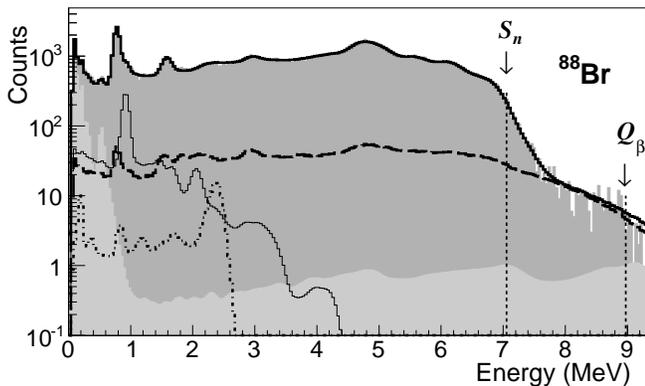


FIG. 1. Relevant histograms for ^{88}Br : parent decay (dark grey filled), daughter decay (dot-dashed line), summing-pileup (dashed line), β -delayed neutron decay (light grey filled), accidental contamination (thin continuous line), reconstructed spectrum (thick continuous line).

tion. Due to the existence of an electronic threshold in the Si detector (100 keV) the β -detection efficiency has a strong dependence with β -endpoint energy up to about 2 MeV. This affects the region of interest (see Fig. 1). To verify that the MC simulation reproduces this energy dependence we use the information from a separate experiment [34] measuring P_n values with the neutron counter BELEN and the same β detector. Several isotopes with different neutron emission windows $Q_\beta - S_n$ were measured, resulting in variations of the neutron-gated β efficiency as large as 25%. Geant4 simulations using the above mentioned β -delayed neutron decay generator are able to reproduce the isotope-dependent efficiency within better than 4%.

Figure 1 shows the β -gated TAGS spectrum measured during the implantation of ^{88}Br ions. Also shown is the contribution of the daughter ^{88}Kr decay, the neutron decay branch populating ^{87}Kr , the summing-pileup contribution and the accidental contamination of ^{94}Y . About 30% of the emitted neutrons produce a signal (light grey filled histogram). Most of the signals, concentrated below 1 MeV, are due to inelastic scattering. Only 1.5% of the neutrons undergo capture depositing energy up to 10 MeV. Notice the presence of net counts beyond the neutron separation energy, which can only be attributed to the decay feeding excited states above S_n which de-excite by γ -ray emission. In this region the major background contribution comes from summing-pileup which is well reproduced by the calculation as can be observed. Similar pictures were obtained for the decay of ^{87}Br and ^{94}Rb .

The analysis of the β -gated spectra follows the method developed by the Valencia group [35, 36]. The intensity distribution $I_{\beta\gamma}$ is obtained by deconvolution of the TAGS spectrum with the calculated spectrometer response to the decay. The response to electromagnetic

cascades is calculated from a set of branching ratios (BR) and the MC calculated response to individual γ rays. Branching ratios are taken from [23–25] for the low energy part of the decay level scheme. The excitation energy range above the last discrete level is treated as a continuum divided into 40 keV bins. Average BR for each bin are calculated from NLD and PSF as prescribed by the Hauser-Feshbach model. We use NLD from Ref. [37] as tabulated in the RIPL-3 library [38]. The PSF is obtained from Generalized Lorentzian (E1) or Lorentzian (M1, E2) functions using the parameters recommended in Ref. [38]. The electromagnetic response is then convoluted with the simulated response to the β continuum. The spin-parity of some of the discrete states at low excitation energy in the daughter nucleus is uncertain. They are however required to calculate the BR from the states in the continuum. The unknown spin-parities were varied and those values giving the best reproduction of the spectrum were adopted. There is also ambiguity in the spin-parity of the parent nucleus which determines the spin-parity of the levels populated in the continuum. Here we assume that allowed Gamow-Teller (GT) selection rules apply. Our choices, $3/2^-$ for ^{87}Br , 1^- for ^{88}Br and 3^- for ^{94}Rb , are also based on which values best reproduce the spectrum.

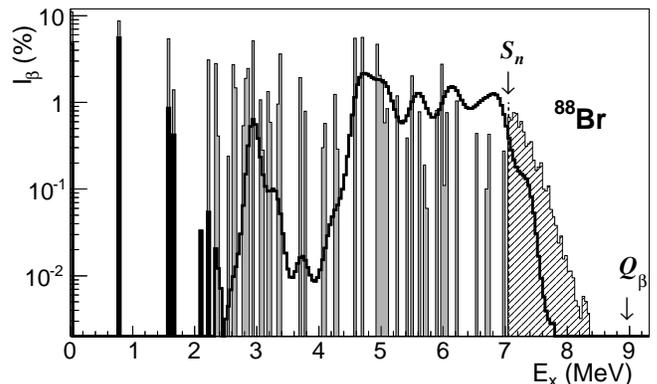


FIG. 2. Beta intensity distributions for ^{88}Br : TAGS result (continuous line), high-resolution γ spectroscopy (light grey filled histogram), from β -delayed neutron (hatched area).

As an example of the results of the analysis we show in Fig. 2 the $I_{\beta\gamma}$ intensity obtained for ^{88}Br . The spectrum reconstructed with this intensity distribution reproduces well the measured spectrum (see Fig. 1). The analysis for the other two isotopes shows similar quality in the reproduction of the spectra. We also include in Fig. 2 the intensity obtained from high-resolution measurements [24], showing a strong *Pandemonium* effect. The *Pandemonium* effect is even stronger in the case of ^{94}Rb and somewhat less for ^{87}Br . The complete $I_{\beta\gamma}$ and its impact on reactor decay heat [39] and antineutrino spectrum [40] summation calculations will be discussed elsewhere [41]. Here we concentrate on the portion of

that intensity located in the neutron unbound region. A sizable TAGS intensity is observed above S_n extending well beyond the first few hundred keV where the low neutron penetrability makes γ -ray emission competitive. For comparison Fig. 2 also shows $I_{\beta n}$ deduced from the neutron spectrum [31] as explained above. The $I_{\beta\gamma}$ above S_n adds up to $\sum I_{\beta\gamma} = 1.6(3)\%$, to be compared with the integrated $I_{\beta n}$ (or P_n) of 6.4(6)%. From the TAGS analysis for the other two isotopes we find a $\sum I_{\beta\gamma}$ of 3.5(5)% (^{87}Br) and 0.53(16)% (^{94}Rb) to be compared with P_n -values of 2.60(4)% and 10.18(24)% respectively. In the case of ^{87}Br we find 7 times more intensity than the high-resolution measurement [21]. The uncertainty quoted on $\sum I_{\beta\gamma}$ is dominated by systematic uncertainties. We did a careful evaluation of possible sources of systematic effects for each isotope. The uncertainty coming from assumptions in the BR varies from 1% to 5% (relative value) depending on the isotope. The impact of the use of different deconvolution algorithms [36] is in the range of 2% to 10%. The uncertainty in the energy dependence of the β efficiency contributes with 4%. The contribution of uncertainties in the width calibration ranges from 2% to 6%. A major source of uncertainty comes from the normalization of the background contribution, which at the energies of interest is dominated by the summing-pileup. We estimated that reproduction of spectra could accommodate at most a $\pm 15\%$ variation from the nominal value, which translates into uncertainties of 6% to 22%. The integral value $\sum I_{\beta\gamma}$ is affected also by the uncertainty in the integration range. The S_n value is known to better than 8 keV for all three isotopes and we estimate that the energy calibration in this region is correct to about one energy bin. This represents an additional uncertainty ranging from 11% to 15%.

Figure 3 shows the ratio $I_{\beta\gamma}/(I_{\beta\gamma} + I_{\beta n})$ in the range of energies analyzed with TAGS for all three cases. This ratio is identical to the average ratio $\langle \Gamma_\gamma / (\Gamma_\gamma + \Gamma_n) \rangle$ over all levels populated in the decay. The shaded area around the experimental value in Fig. 3 serves to indicate the sensitivity of the TAGS results to background normalization as indicated above. The average width ratio was calculated using the Hauser-Feshbach model. The results for the three spin-parity groups populated in GT decay are shown. The NLD and PSF values used in these calculations are the same as those used in the TAGS analysis. The new ingredient needed is the NTC, which is obtained from the Optical Model (OM) with the TALYS-1.4 software package [42]. OM parameters are taken from the so-called local parametrization of Ref. [43]. Neutron transmission is calculated for known final levels populated in the decay [23–25]. In order to compute the average width ratio we need to include the effect of statistical fluctuations in the individual widths [22]. We use the MC method to obtain the average of width ratios. The sampling procedure is analogous to that described in Ref. [35]. Level energies for each spin-parity are gener-

ated according to a Wigner distribution and their corresponding Γ_γ and Γ_n to individual final states are sampled from PT distributions. The total γ and neutron widths are obtained by summation over all possible final states and the ratio computed. The ratio is averaged for all levels lying within each energy bin. In order to suppress fluctuations in the calculated average, the sampling procedure is repeated between 5 and 1000 times depending on level density. Very large average enhancement factors were obtained, reaching two orders-of-magnitude when the neutron emission is dominated by the transition to a single final state.

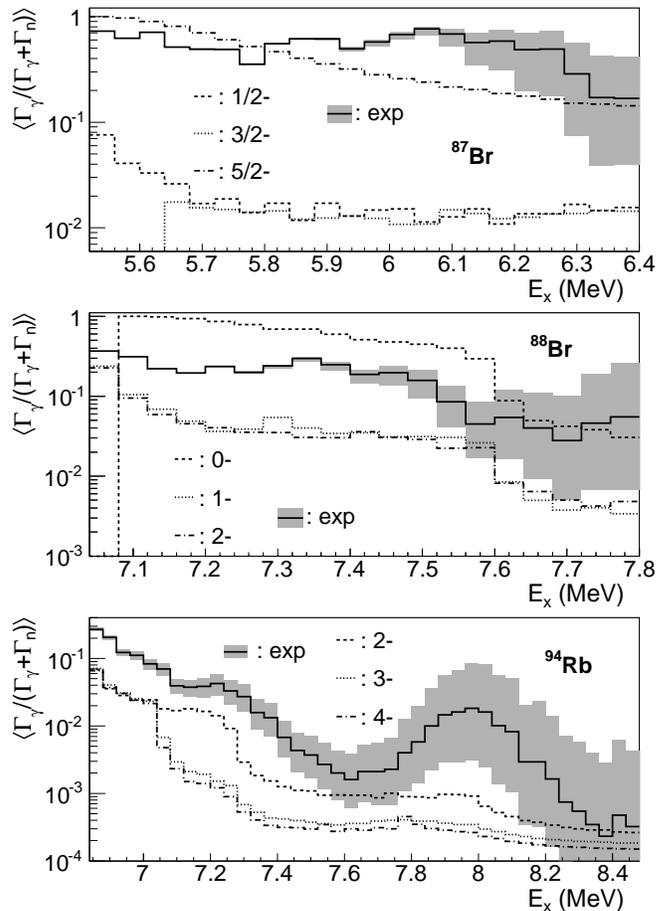


FIG. 3. Average gamma to total width ratio from experiment and calculated for the three spin-parity groups populated in allowed decays. The shaded area around the experimental value indicates the sensitivity to the background normalization (see text).

In the case of ^{87}Br $3/2^-$ decay one can see in Fig. 3 that the strong γ -ray emission above S_n can be explained as a consequence of the large hindrance of $l = 3$ neutron emission from $5/2^-$ states in ^{87}Kr to the 0^+ g.s. of ^{86}Kr , as pointed out in Ref. [1]. In the case of ^{88}Br 1^- decay a similar situation occurs for 0^- states in ^{88}Kr below the first excited state in ^{87}Kr at 532 keV, which require $l = 3$

321 to populate the $5/2^+$ g.s. in ^{87}Kr . For a more quanti-
 322 tative assessment one should know the distribution of β
 323 intensity between the three spin groups, which could be
 324 obtained from β -strength theoretical calculations. The
 325 case of ^{94}Rb 3^- decay is the most interesting. The fi-
 326 nal nucleus ^{93}Sr is five neutrons away from β stability.
 327 The γ intensity although strongly reduced, only 5% of
 328 the neutron intensity, is detectable up to 1.5 MeV be-
 329 yond S_n . The structure observed in the average width
 330 ratio, is associated with the opening of βn channels to
 331 different excited states. Note that the structure is re-
 332 produced by the calculation, which confirms the energy
 333 calibration at high excitation energies. In any case the
 334 calculated average gamma-to-total ratio is well below the
 335 experiment. In order to bring the calculation to the ex-
 336 perimental value one would need to enhance the PSF,
 337 or suppress the NTC, or any suitable combination of the
 338 two, by a very large factor. For instance we verified that
 339 a twenty-fold increase of the E1 PSF would reproduce
 340 the measurement assuming a β -intensity spin distribu-
 341 tion proportional to $2J + 1$. An enhancement of such
 342 magnitude for neutron-rich nuclei, leading to a similar
 343 enhancement of (n, γ) cross sections, will likely have an
 344 impact on r -process abundance calculations. Therefore
 345 it will be important to investigate the magnitude of pos-
 346 sible variations of the NTC.

347 In conclusion, we have confirmed the suitability of the
 348 TAGS technique to obtain accurate information on γ -ray
 349 emission from neutron unbound states and applied it to
 350 three known β -delayed neutron emitters. A surprisingly
 351 large γ -ray branching of 57% and 20% was observed for
 352 ^{87}Br and ^{88}Br respectively, which can be explained as a
 353 nuclear structure effect. In the case of ^{87}Br we observe 7
 354 times more intensity than previously detected with high
 355 resolution γ -ray spectroscopy, which confirms the need of
 356 the TAGS technique for such studies. In the case of the
 357 more neutron-rich ^{94}Rb the measured branching is only
 358 4.5% but still much larger than the results of Hauser-
 359 Feshbach statistical calculations, after proper correction
 360 for individual width fluctuations. The large difference
 361 between experiment and calculation can be reconciled by
 362 an enhancement of standard PSF of over one order-of-
 363 magnitude. To draw more general conclusions it will be
 364 necessary to extend this type of study to other neutron-
 365 rich β -delayed neutron emitters. Such measurements us-
 366 ing the TAGS technique are already underway and addi-
 367 tional ones are planned.

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