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SU(3) semiclassical representation of quantum dynamics of interacting spins

Shainen M Davidson and Anatoli Polkovnikov

Department of Physics, Boston University, Boston, MA 02215, USA

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We present a formalism for simulating quantum dynamics of lattice spin-one systems by first introducing local hidden variables and then doing semiclassical (truncated Wigner) approximation in the extended phase space. In this way we exactly take into account the local on-site Hamiltonian and approximately treat spin-spin interactions. In particular, we represent each spin with eight classical SU(3) variables. Three of them represent the usual spin components and five others are hidden variables representing local spin-spin correlations. We compare our formalism with exact quantum dynamics of fully connected spin systems and find very good agreement. As an application we discuss quench dynamics of a Bose-Hubbard model near the superfluid-insulator transition for a 3D lattice system consisting of 1000 sites.

Recent experiments in such areas as ultra-cold gases, coupled atom-photon systems, ultrafast pump-probe spectroscopy in solids, and others has stimulated active theoretical research in quantum dynamics of interacting systems. While there has been significant progress in various directions, our understanding is still quite limited. Partly this is due to a lack of reliable and controllable numerical methods. Perhaps the most powerful methods in equilibrium based on guantum Monte-Carlo are not readily adopted to non-equilibrium systems. The other available methods include: simulations based on exact diagonalization; dynamical renormalization group and related matrix product states methods [1]; dynamical mean field theory based methods [2, 3] mostly developed for fermions and only recently applied to bosons [4]; quantum kinetic equations and Keldysh diagrammatic technique [5]; and phase space methods. The latter has recently become a major tool for studying dynamics of various systems, from interacting atomic clocks to the early Universe [6–12]. These methods are based on mapping the density matrix and operators into classical functions, which depend on phase space variables like coordinates and momenta, complex wave amplitudes, or classical spin degrees of freedom.

Phase space methods are very efficient for systems near the classical or non-interacting limit, where the quantum evolution is well described by classical trajectories. The main idea of the present work is that we can significantly improve the accuracy of these methods by extending classical phase space by introducing new (hidden) phase space variables. In this way we can take into account local quantum fluctuations exactly and treat other degrees of freedom approximately. We illustrate this idea by focusing on an example of spin-one coupled systems where, in addition to three variables representing the x, y, z components of the spin, we introduce five additional hidden variables representing local spin-spin correlations. We believe these ideas can be further extended and applied to study a large class of interacting systems both in and out of equilibrium.

Before proceeding with our ideas we briefly review the

phase space representation of quantum dynamics of interacting bosonic systems. For concreteness we focus on the Wigner-Weyl quantization (see Supplementary Material for a brief overview [13]). Any operator of a quantum system that we would normally represent through a function of boson operators \hat{a} and \hat{a}^{\dagger} can be mapped to a function over the classical phase space of (complex) canonical variables α and α^* . This function is called the Weyl symbol of the operator and it is uniquely defined. The Weyl symbol of the density matrix is known as the Wigner function [10, 16]. It plays the role of the (quasi)-probability distribution. In general, time evolution of the Wigner function is given by a Fokker-Planck equation with high derivatives, which is hard to handle. However, near the classical limit or for non-interacting systems one can use the so-called truncated Wigner approximation (TWA) [6, 7, 10], where the Wigner function, like a classical probability distribution, is conserved on classical trajectories. Then the expectation value of any operator can be straightforwardly computed at any moment in time:

$$\langle \hat{\Omega}(t) \rangle \approx \int d\vec{\alpha}_0 d\vec{\alpha}_0^* W(\vec{\alpha}_0, \vec{\alpha}_0^*) \Omega_W(\vec{\alpha}_{cl}(t), \vec{\alpha}_{cl}^*(t)), \quad (1)$$

where $W(\vec{\alpha}_0, \vec{\alpha}_0^*)$ is the Wigner function representing the initial state of the system, $\Omega_W(\vec{\alpha}, \vec{\alpha}^*)$ is the Weyl symbol of the operator $\hat{\Omega}$, and the classical paths are found by solving Hamilton's equations,

$$i\dot{\alpha}_{cl} = \frac{\partial H_W}{\partial \alpha_{cl}^*}.$$
 (2)

TWA becomes more accurate as the number of particles increases. It is always exact when the Hamiltonian is non-interacting, i.e. quadratic in boson operators.

One way to implement TWA with spin operators is to use the Schwinger boson representation for spin, where the spin operators are built from 2 creation/annihilation pairs of boson operators contracted with the Pauli matrices:

$$\hat{S}_{\alpha} = \hat{a}_{i}^{\dagger} \frac{1}{2} \sigma_{\alpha}^{ij} \hat{a}_{j}.$$
(3)

The boson pairs "inherit" the commutation relations of the matrices they are contracted with, in this case the spin algebra of SU(2). Representing the spin Hamiltonian in this way, we can use the machinery of coherent state TWA to approximate the dynamics (see [10] Section 3.3). Since TWA dynamics is exact for a Hamiltonian that is quadratic in boson operators, it is also exact for a system linear in spin operators.

We can similarly represent any SU(N) group of operators with commutation relations $[\hat{X}_{\alpha}, \hat{X}_{\beta}] = i f_{\alpha\beta\gamma} \hat{X}_{\gamma}$ using N boson operator pairs and the N dimensional matrices T_{α}^{ij} with the corresponding algebra:

$$\hat{X}_{\alpha} = \hat{a}_i^{\dagger} T_{\alpha}^{ij} \hat{a}_j.$$

$$\tag{4}$$

Note that any Hamiltonian that acts on a Hilbert space of dimension N (or indeed any Hermitian operator of dimension N) can be represented as a linear superposition of the generators of the SU(N) group and an identity matrix. Again, the TWA dynamics of any Hamiltonian linear in SU(N) variables will be exact. This statement can be alternatively understood based on the linearity of the Heisenberg equations of motion as was recently explored in Ref. [17].

Taking the Wigner-Weyl transform of (4), we can define real number valued variables

$$X_{\alpha} = \alpha_i^* T_{\alpha}^{ij} \alpha_j, \tag{5}$$

and the classical equations of motion for the canonical variables (2) will lead to dynamics for these variables as (see Supplementary Material for details [13])

$$\dot{X}_{\alpha}^{cl} = f_{\alpha\beta\gamma} \frac{\partial H_W}{\partial X_{\beta}^{cl}} X_{\gamma}^{cl}.$$
(6)

In the case of the SU(2) group representing e.g. spin in a magnetic field the structure constants are given by $f_{\alpha\beta\gamma} = \epsilon_{\alpha\beta\gamma}$, where ϵ is the fully antisymmetric tensor. Then it is easy to see that the equation above represents the standard Bloch equations $\vec{X} = \vec{X} \times \vec{B}$, where $H = -\vec{B}\vec{X}$ and \vec{B} is a (possibly time-dependent) magnetic field.

If we have a single spin 1/2 degree of freedom then its Hamiltonian can always be represented as a linear superposition of Pauli matrices (generators of the SU(2)group) and thus semiclassical dynamics are exact. Consider now the slightly more complicated situation of an isolated spin-one degree of freedom. A generic Hamiltonian for spin-one can include interactions, i.e. terms non-linear in spin operators. Just to be specific consider the interactions of the type \hat{S}_z^2 such that

$$\hat{H}_I = -\vec{B}\vec{S} + (U/2)\hat{S}_z^2.$$
(7)

Note that both \vec{B} and U can explicitly depend on time. If the interaction coupling U is zero we are back to the situation discussed before, where the Hamiltonian is a linear superposition of SU(2) generators and semiclassical dynamics are exact. However, for finite interactions the semiclassical approximation breaks down at long times. However, since spin-one operators exist in a 3 dimensional Hilbert space, we can make the semiclassical dynamics exact by enlarging the group to SU(3) by using the 8 3-dimensional matrix generators of SU(3) as the matrices T_{α} in (4). Those can be chosen to be Gell-Mann matrices, but it is more convenient to use linear combination such that the first three generators are the spin-one spin matrices [18], e.g.

$$T_1 = \begin{pmatrix} 0 & \frac{1}{\sqrt{2}} & 0\\ \frac{1}{\sqrt{2}} & 0 & \frac{1}{\sqrt{2}}\\ 0 & \frac{1}{\sqrt{2}} & 0 \end{pmatrix}$$
(8)

and thus $\hat{X}_1 = \left(\hat{a}_1^{\dagger}\hat{a}_2 + \hat{a}_3^{\dagger}\hat{a}_2 + h.c.\right)/\sqrt{2}$ and so on (see Supplementary Material for details [13]).

Because any Hermitian 3×3 matrix can be expressed through a linear combination of SU(3) matrices (and the identity) the interaction term also becomes linear in the SU(3) representation. For example, for our choice of SU(3) generators we have $\hat{S}_z^2 = (2 - \sqrt{3}\hat{X}_8)/3$.

Thus the whole spin-1 Hamiltonian becomes linear in the generators of SU(3) and the equations of motion (6) become exact. These equations can be interpreted as an exact classical representation of quantum dynamics of a spin-one system in the eight-dimensional phase space spanned by the classical phase space variables $X_1, \ldots, X_8[19]$. X_1, X_2, X_3 represent the three spin components, while the remaining five variables represent nonlinear spin terms. These are effectively hidden variables in our approach. The advantage of our approach becomes apparent when we start coupling spin-one systems together and start doing approximations.

Note that the eight equations (6) are not completely independent: they satisfy constraints set by conservation of the Casimir operators:

$$C_1 = \sum_{\alpha} X_{\alpha}^2, \ C_2 = \sum_{\alpha\beta\gamma} d_{\alpha\beta\gamma} X_{\alpha} X_{\beta} X_{\gamma}, \tag{9}$$

where $d_{\alpha\beta\gamma}$ are the symmetric structure constants of the SU(3) group. Interestingly $X_1^2 + X_2^2 + X_3^2$, representing the sum of squares of the classical spin components, is not conserved under SU(3) dynamics. There is no paradox here: recall all operators including $\hat{S}_{x,y,z}^2$ are represented through linear combination of operators \hat{X}_{α} so e.g. X_1^2 does not have the simple physical interpretation of S_x^2 even though X_1 does represent S_x .

To illustrate the difference between the "naive" SU(2)TWA and the new SU(3) TWA we consider the Hamiltonian (7) with $B_x = B_y = 0$ and $B_z = 1$. The SU(2) and SU(3) Weyl symbols corresponding to this Hamiltonian

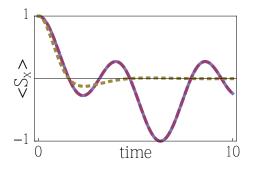


FIG. 1. Comparison of the dynamics of $\langle \hat{S}_x \rangle$ for a spin-one particle initially pointed in the x-direction subjected to the Hamiltonian $\frac{1}{2}\hat{S}_z^2 - \hat{S}_z$. The dynamics are calculated with exact diagonalization (solid, blue), SU(3) TWA (dashed, red), and SU(2) TWA (dotted, yellow). (Color online.)

are

$$(H_I)_W^{SU(2)} = (U/2)X_3^2 - X_3 - 1/2, (H_I)_W^{SU(3)} = (U/6)(2 - \sqrt{3}X_8) - X_3.$$
(10)

Note that in the SU(2) case we chose the spin-one representation of the spin operators given by the first three operators of the SU(3) representation. The additional constant term -1/2 in the SU(2) Hamiltonian comes from $(\hat{X}_k^2)_W = X_k^2 - \operatorname{tr}(T_k^2)/4$. For concreteness we choose U = 1 and start with the spin pointing along the xdirection and observe the expectation value of \hat{S}_x as a function of time. In Fig. 1 we show comparison of the resulting exact dynamics with SU(2) and SU(3) TWA approximations. As expected the SU(3) TWA is exact while the SU(2) semiclassical dynamics are only accurate at short times. The difference comes from the fact that any interaction terms in the SU(2) case are represented by non-linearity while in the SU(3) case they are represented by additional (hidden) variables, which in turn have their own quite complex dynamics.

Next let us consider a more complicated setup, where we deal with a system of interacting spin-one degrees of freedom such that the Hamiltonian becomes

$$\hat{H} = \sum_{n} \hat{H}_{I}^{(n)} + \hat{H}_{C} \tag{11}$$

where $H_I^{(n)}$ is the local spin-one Hamiltonian (7) describing *n*-th spin and

$$\hat{H}_{C} = -J \sum_{n \neq m} (\hat{S}_{x}^{n} \hat{S}_{x}^{m} + \hat{S}_{y}^{n} \hat{S}_{y}^{m}).$$
(12)

We have chosen a fully connected Hamiltonian to allow for comparison of TWA and exact dynamics for larger system sizes.

The Weyl symbol of the coupling term is the same for the SU(2) and the SU(3) representations because it does not involve local nonlinear spin-operators,

$$(H_C)_W = -J \sum_{n \neq m} (X_1^n X_1^m + X_2^n X_2^m).$$
(13)

For multi-spin systems we do not use the exact Wigner function, which is defined and integrated over the coherent state variables. Instead, we use a multivariate Gaussian distribution $f^n(X_1^n, \ldots, X_{N^2-1}^n)$ for each site n and integrate over the $N^2 - 1$ SU(N) variables:

$$\langle \hat{\Omega}(t) \rangle \approx \int \prod_{n} dX^{n} f^{n}(\vec{X}^{n}) \Omega_{W}(\vec{X}^{cl\,n}(t)), \qquad (14)$$

where the mean and covariance matrix for each f is fixed by the quantum expectation values of the initial state of the system, $\langle X^n_{\alpha} \rangle_{f^n} = \langle \hat{X}^n_{\alpha} \rangle$ and $\langle X^n_{\alpha} X^n_{\beta} \rangle_{f^n} =$ $\langle (\hat{X}^n_{\alpha} \hat{X}^n_{\beta} + \hat{X}^n_{\beta} \hat{X}^n_{\alpha})/2 \rangle$ (see Supplementary Material for details [13]). We use this best Gaussian approximation for two reasons. First, the exact Wigner function will in general have negative values, so the integration depends on the cancellation of positive and negative contributions, which numerically requires more sample points to converge. Secondly, we numerically found that the best Gaussian TWA results are consistently more accurate. Formally this Gaussian scheme is justified if we increase the spin size (proportional to the conserved value of the Casimir operator). For an initial correlated (not product) state one should use the multivariate Gaussian which correctly reproduces both local and non-local correlation functions like $\langle X^m_{\alpha} X^n_{\beta} \rangle$. For observables, instead of Weyl ordering one can use direct quantum classical substitution $\hat{X}_m^n \to X_m^n$ because any onsite observable is linear in \hat{X} and for linear operators this substitution is exact [10].

In Fig. 2 we show the dynamics of the spin fluctuations $\langle \hat{S}_{z}^{2} \rangle$ per site obtained by exact diagonalization, and SU(2) and SU(3) TWA. The system is initially prepared with all spins pointing in the x-direction. We compare the dynamics for a fully connected system for different values of the coupling J and for different system sizes. As the coupling is lowered and the on-site term in the Hamiltonian becomes more dominant, the SU(3) TWA becomes a better approximation, while the SU(2) becomes worse. When the on-site term is 5 times as dominant as the coupling term, the SU(3) TWA is indistinguishable from exact quantum dynamics. As the system size increases, and hence each site is connected to more sites, the SU(3) TWA dynamics approach exact quantum dynamics. Similarly to the SU(2) case, SU(3) TWA fails to describe quantum revivals, which occur later and later in time as the system size increases.

As a more practical example, we model the Bose-

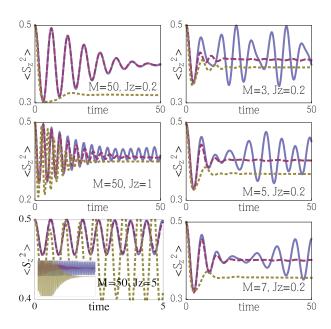


FIG. 2. Comparison of the dynamics of $\langle \hat{S}_z^2 \rangle$ for a fully connected system with M spins due to Hamiltonian (11) calculated with exact diagonalization (solid, blue), SU(3) TWA (dashed, red), and SU(2) TWA (dotted, yellow). Initially all spins are pointed in the x-direction. On the left are the results for various coupling strengths Jz, where z = M - 1 is the connectedness (the inset in the bottom plot shows the same plot for times 0 to 50). On the right are the results for different system sizes. (Color online.)

Hubbard model using the effective Hamiltonian [20]

$$H_{eff} = \frac{U}{2} \sum_{i} (\hat{S}_{z}^{i})^{2} - J\bar{n} \sum_{\langle ij \rangle} (\hat{S}_{x}^{i} \hat{S}_{x}^{j} + \hat{S}_{y}^{i} \hat{S}_{y}^{j}) - \mu \sum_{i} \hat{S}_{z}^{i},$$
(15)

where \bar{n} is the mean particle density. This truncation of the Hilbert space to three dimensions per site is acceptable in the vicinity of the Mott insulating state [21]. We use SU(3) TWA to determine the dynamics of the order parameter $\rho_s = \sum_{i \neq j} \langle \hat{S}_i^+ \hat{S}_j^- \rangle / M^2$ (representing superfluid density) for a 3D system with $M = 10^3$ sites in a cubic lattice with periodic boundary conditions. The results are plotted in Fig. 3.

First we quench from the Mott insulator phase, i.e. a Fock state on each site (the ground state for J = 0). In terms of the effective spin Hamiltonian, this corresponds to a product state of $|\hat{S}_z = 0\rangle$. The dynamics arise from an instantaneous quench to a finite coupling, either $J\bar{n}z/U = 0.2$ or $J\bar{n}z/U = 1$. In each case, the system moves away from a pure Mott insulator state; for a smaller coupling, there is some oscillation which is absent for a larger coupling. The superfluid density remains small, as a sudden quench leads to a high temperature state which does not exhibit long range order [22].

We also show a quench from the superfluid phase (the ground state for U = 0), which in terms of the effec-

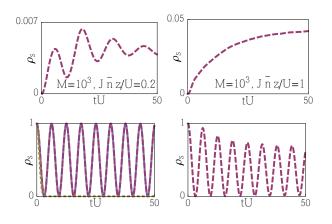


FIG. 3. The dynamics of the order parameter ρ_s for the effective Bose-Hubbard model (15), calculated with SU(3) TWA. The top row shows the results starting in the Mott insulator phase. The bottom row begins in the superfluid phase: in the left-hand plot, the system is quenched to J = 0, while in the right hand-hand plot the system is quenched to $J\bar{n}z/U = 0.1$. (Color online.)

tive spin Hamiltonian corresponds to a product state of $|\hat{S}_x = +1\rangle$. When the system is quenched to J = 0, each site precesses independently. Thus we can calculate the dynamics using exact diagonalization, SU(3) TWA, and SU(2) TWA. Since the on-site Hamiltonian can be linearized in terms of SU(3) variables, the SU(3) TWA reproduces the exact quantum dynamics, including quantum recurrences, while the SU(2) TWA decays. When we instantaneously quench to $J\bar{n}z/U = 0.1$, the SU(3) TWA still reproduces the oscillations of quantum recurrences, damped by the coupling to the larger system.

In summary, we have introduced a semiclassical formalism for simulating the quantum dynamics of strongly interacting coupled-spin systems. We have shown that by increasing the phase space and introducing new (hidden) degrees of freedom one can partially account for local quantum fluctuations and significantly improve the accuracy of the semiclassical description of the dynamics. We have argued and shown numerically that the accuracy of this method increases as we increase connectivity of the system. We have demonstrated numerically that this method accurately reproduced results of quench dynamics of coupled spin-one systems in a broad range of parameters including the strong coupling regime. As another illustration we analyzed quench dynamics across the superfluid-insulator transition in a three-dimensional Bose-Hubbard model.

While here we only presented results for SU(3) variables, this formalism can be straightforwardly applied to any SU(N) group (albeit with a larger phase space), where N is the dimension of a local Hamiltonian. Thus we can use classical dynamics to exactly do local quantum dynamics which are linear in any SU(N) representation. This should allow one to take into account quan-

tum fluctuations within larger clusters and then use the TWA approximation to treat inter-cluster coupling. We are planning to analyze this possibility in a future work. An important and open question is finding the optimal way of introducing hidden variables keeping their number larger than in the naive classical limit, yet much smaller than the Hilbert space size.

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