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Quantum ripples over a semi-classical shock

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The evolution of an initially smooth spatial inhomogeneity in the density of a one-dimensional Fermi gas is well described by classical mechanics. The classical evolution leads to the formation of a shock wave: the density develops kinks in its coordinate dependence. We show that quantum corrections to the shock wave produce density ripples which run off the kinks. Despite their quantum origin, the amplitude and period of the ripples are expressed only in terms of classical objects derived from a smooth density profile.

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In one-dimensional many-body physics, numerous long-wavelength properties of a quantum system are faithfully represented by the dynamical properties of a continuous liquid. Such representation is at the heart of a powerful bosonization method [1-3]. Small-amplitude perturbations of the liquid's density may be considered in a harmonic approximation. In that approximation, waves of density are linear, and there is no difference between the quantum and the classical dynamics of the density perturbations. At larger amplitudes, waves become nonlinear; it is not clear then, if the classical and quantum dynamics remain indistinguishable even in the long-wavelength limit.

In a liquid made of fermions, higher density is associated with higher Fermi momentum and higher particles velocities. Described classically, a higher-density solitary segment moves faster than its lower-density periphery. That leads to a wave overturn and the formation of shocks in the density $\rho(x,t)$, *i.e.* points $x_{\gamma}(t)$ where the x-dependence of ρ looses its analyticity. Formation of shocks can be seen within a classical continuous medium description in terms of the Riemann-Hopf equation [4– 6]. It is insensitive to the interaction strength, which can be set to zero (the free fermions case). The classical description of propagation of shock waves in a gas of quantum particles is widely accepted across the fields of quantum physics from string theory [7] to cold atomic gases [8]. In the case of free fermions, however, one may also investigate the fully-quantum evolution of the many-body wave function, which is a Slater determinant of free-propagating single-particle states.

We consider the shock formation in the semiclassical limit of the quantum evolution of free fermions. Our main finding is the appearance of oscillatory structure in $\partial \rho / \partial x$ accompanying each of the shock points $x_{\gamma}(t)$. This structure, quantum in origin, has a characteristic scale of variation fully determined by the solution of the classical problem. (In that respect, the free-particle many-body wave function associated with the shock formation bears some resemblance to the WKB wave function of a single particle in an external potential.) In classical physics, the formation of shocks [6][9] is associated with the presence of velocity gradients. The inverse time to develop a shock can be estimated as $t_S \sim |dv/dx|$, where v(x) is the initial velocity distribution. If the velocity gradient comes from a "bump" $\Delta \rho$ in the density of fermions, then the corresponding density gradient, $d\rho/dx = (m/\hbar)dv/dx$ becomes large in the semiclassical limit, $\hbar \rightarrow 0$ at fixed dv/dx (hereinafter *m* is the mass of a fermion). Together with $|d\rho/dx|$, the total number of particles ΔN in the density bump scales as $1/\hbar$. We show that $1/\Delta N$ can be consistently used as a small parameter in the semiclassical expansion. The characteristic length of the oscillatory structure around a shock point scales as $(\Delta N)^{-2/3}$.

We are considering the time evolution of a solitary density maximum $\rho(x, t)$, which initially (at t = 0) has spatial extent Δx and amplitude $\Delta \rho$,

$$\rho(x,0) = n_0 + \Delta \rho f\left(\frac{x}{\Delta x}\right) \,. \tag{1}$$

Here the dimensionless positive function f(s) reaches a maximum, $f(0) \sim 1$, at s = 0, has spatial extent $|s| \sim 1$, and decays outside that region, $f(s \to \pm \infty) = 0$, revealing the background equilibrium density $n_0 = K_F/\pi$ of a Fermi sea $(|k| \leq K_F)$ of spinless particles. We aim to describe the motion of *semiclassical* perturbations. That leads us to assume that many particles go into the creation of the perturbation, $\Delta N = \Delta \rho \Delta x \gg 1$. Our further assumption is that the density perturbation is small compared to the equilibrium density, $\Delta \rho \ll n_0$. The two latter assumptions mean that $K_F \Delta x \gg \Delta N \gg 1$; therefore, the density perturbation consists of particles and holes residing in the narrow vicinities $\delta k \sim K_F \Delta \rho/n_0$ of the Fermi points $\pm K_F$.

There are several characteristic time scales for the evolution of the density perturbation. First, it will decompose into right- and left-moving modes, associated with particle-hole excitations around each one of the Fermi points. The characteristic time associated with this process is $t_{LR} = (m/\hbar)(\Delta x/n_0)$. Next, there is a characteristic time over which the wave packet representing a single fermion spreads over a spatial scale Δx of the initial many-body state. This time is given by $t_Q = (m/\hbar)(\Delta x)^2$. On this time scale the evolution ceases to be semiclassical. Finally, in the course of the semiclassical evolution of the density perturbation, it may develop a shock, in analogy to the one-dimensional classical hydrodynamics. This effect has been discussed, *e.g.*, in [10]. The corresponding characteristic time scale is $t_S = (m/\hbar)(\Delta x/\Delta \rho).$

Our assumptions $\Delta N \gg 1$ and $\Delta \rho \ll n_0$ lead to the time scales hierarchy, $t_{LR} \ll t_S \ll t_Q$. Namely, first the perturbation splits into left and right moving modes, then classical shocks have a chance to appear in each of the modes, and only at much later times the semiclassical nature of the evolution breaks down. We focus on the evolution occurring on time scale $\sim t_S$.

Since we are interested in the semiclassical regime, it is convenient to use the Wigner function,

$$W(x,k,t) = \int \left\langle \Psi | \psi^{\dagger} \left(x + \frac{y}{2}, t \right) \psi \left(x - \frac{y}{2}, t \right) | \Psi \right\rangle e^{-iky} dy,$$
(2)

where $|\Psi\rangle$ is the initial (t=0) state of the system.

In equilibrium, $|\Psi\rangle$ is a state where all single particle levels with momenta between $-\hbar K_F$ and $\hbar K_F$ are occupied, while all other states are empty. Here $K_F = \pi n_0$. This state leads to a Wigner function given by $W(x, k, t) = \theta (K_F - k) \theta (k + K_F)$. The Wigner function is interpreted as a distribution in phase space of the Fermions, which, in equilibrium, occupy the region $-K_F < k < K_F$ for all x, namely a band in phase space.

We can think about a semiclassical state in which a single mode (either right- or left-moving) of density perturbation is excited, as a state in which one of the Fermi wave numbers $(-K_F \text{ or } K_F)$ is allowed to vary with the space-time coordinates. From now on, and without loss of generality, we shall assume only a right moving mode, in which case the Fermi number K_F becomes time and space dependent, to be described by a function $k_F(x,t)$, while $-K_F$ stays fixed. This heuristic picture translates into the following *ansatz* for the Wigner function:

$$W(x,k,t) = \theta(k_F(x,t) - k)\theta(k + K_F).$$
(3)

We will see that Eq. (3) is a good approximation to the Wigner function everywhere except the vicinities of the shock points. Knowledge of the Wigner function allows one to compute the density, as these two are related as follows:

$$\rho(x,t) = \int W(x,k,t) \frac{dk}{2\pi}.$$
(4)

Given initial conditions for the Wigner function, one may compute it at later times making use of the evolution equation, $\partial_t W(x,k,t) + \frac{\hbar k}{m} \partial_x W(x,k,t) = 0$. The proof of this equation consists of the application of the Schrödinger equation to the fermions in Eq. (2) and integration by parts. The initial value problem for the evolution equation may be solved by the method of characteristics, which yields

$$W(x,k,t) = W\left(x - \frac{\hbar k}{m}t,k,0\right).$$
 (5)

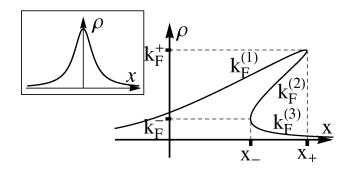


FIG. 1. Multi-valued function $k_F(x,t) = mv_F(x,t)/\hbar$ after the shock. Inset shows the function at initial time.

Applying Eq. (5) to ansatz (3) one sees that $k_F(x, t)$ must satisfy:

$$k_F(x,t) = k_F\left(x - \frac{\hbar}{m}k_F(x,t)t,0\right).$$
 (6)

This equation determines $k_F(x,t)$ implicitly, given the initial conditions $k_F(x,0)$.

The relation to the classical shock wave physics in one dimensional hydrodynamics may be realized by excluding \hbar from Eq. (6) with the transformation $v_F(x,t) = (\hbar/m)k_F(x,t)$ and noticing that $v_F(x,t)$ is in fact a solution of $\partial_t v_F(x,t) + v_F(x,t)\partial_x v_F(x,t) = 0$. This Riemann (or Riemann-Hopf) equation [6] is well known to give rise to shock waves.

In the following we make use of the inverse function, which is a solution of $k_F(x_F, t) = k$:

$$x_F(k,t) = x_F(k,0) + \frac{\hbar}{m}kt.$$
 (7)

If $x'_F(k,0)$ is negative, there is always a positive time at which $\partial x_F(k,t)/\partial k = 0$. This equation, together with Eq. (7) determines the time evolution of points where $|\partial k_F(x,t)/\partial x| = \infty$. The divergence is a manifestation of the shock phenomenon; it is also termed 'the gradient catastrophe'. After the first time the function $k_F(x,t)$ displays an infinite derivative, the solution to Eq. (6) becomes multi-valued as shown in Fig. 1. Without loss of generality, we assume that the multi-valued region is bounded by two points, $x_-(t)$ and $x_+(t)$, between which the function $k_F(x,t)$ has three branches, enumerated by $k_F^{(1)}(x,t) < k_F^{(2)}(x,t) < k_F^{(3)}(x,t)$.

Regardless of the fact that $k_F(x,t)$ is multi-valued, initial conditions such as (3) lead to a Wigner function which is 1 in a region of phase space between the curve $k_F(x,t)$ and the horizontal line $k = -K_F$, while the Wigner function is 0 outside this region. Here the curve of $k_F(x,t)$ is understood to contain all the branches of the function, just as the curve described in the right panel of Fig. 1.

The integral in Eq. (4) may easily be performed to obtain the density. The result is:

$$2\pi\rho^{cl}(x,t) = K_F + k_F^{(3)}(x,t) - k_F^{(2)}(x,t) + k_F^{(1)}(x,t), \ x \in I$$

$$2\pi\rho^{cl}(x,t) = K_F + k_F(x,t), \ x \notin I,$$
(8)

where $I = [x_{-}(t), x_{+}(t)]$ denotes the spatial segment of multi-valued behavior of k_F . The superscript cl in ρ^{cl} denotes that (8) gives only the result of classical theory, on which we want to improve in the vicinities of points $x = x_+$.

Next to $x_{-}(t)$ and $x_{+}(t)$ two branches of the function $k_F(x,t)$ meet. Explicitly, $k_F^{(1)}$ and $k_F^{(2)}$ (or $k_F^{(2)}$ and $k_F^{(3)}$) become equal to each other at $x_+(t)$ (or, respectively, at $x_{-}(t)), e.g.,$

$$k_F^{(1)}(x_+(t),t) = k_F^{(2)}(x_+(t),t) \equiv k_F^+(t).$$

The functions $k_F^{\pm}(t)$ are solutions of the equation $\partial x_F(k,t)/\partial k = 0$. The three branches $k_F^{(1,2,3)}$ display a square-root behavior, see Fig. 1, near respective meeting points $x_{\pm}(t)$, e.g.,

$$k_F^{(1,2)} - k_F^+(t) = \mp \alpha^+(t)\sqrt{x_+(t) - x} + \dots, \quad (9)$$

while $k_F^{(3)}(x,t)$ is regular at $x = x_+(t)$. Near the branch meeting points, $\rho(x,t)$ takes the approximate form:

$$\rho^{cl}(x,t) - \rho^{cl}(x_{\pm}(t),t) \simeq \pm \frac{\alpha^{\pm}(t)}{\pi} \sqrt{\mp (x - x_{\pm}(t))}.$$
(10)

These square-root kinks (which may be observed in Fig. 2 next to $x_{-}(t)$ and $x_{+}(t)$, albeit rounded by quantum corrections) point to the non-analytical behavior of $\rho(x,t)$

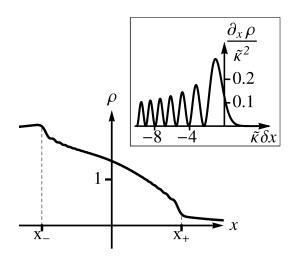


FIG. 2. A typical density profile. The classical trailing and leading shock points are $x_{-}(t)$ and $x_{+}(t)$, respectively; cf. Eq. (10). The quantum ripple effect is amplified in the derivative $\partial_x \rho(x,t)$. Inset: Normalized derivative, $-\partial_x \rho/\tilde{\kappa}^2$, as a function of $\delta x = x - x_{+}(t)$ in the region next to $x_{+}(t)$, see Eq. (21).

at $x = x_{\pm}(t)$. The corresponding coefficients are determined by initial conditions, $[\alpha^{\pm}(t)]^2 = 2(\partial_k^2 x_F)^{-1}$. According to Eq. (7), the derivative $\partial_k^2 x_F(k,t) =$ $d^2 x_F(k,0)/dk^2$ at fixed k is time-independent. To find $\alpha^{\pm}(t)$ one should evaluate the said derivative at k = $k_F^{\pm}(t)$.

Our goal is to evaluate quantum corrections to the classical density profile (10). To that end we start with introducing a family of many-body coherent states of free fermions [11] with smooth density profiles. We are interested in times $t \gg t_{LR}$, so it is sufficient to include only, say, right-movers in the consideration. The corresponding fermion density operator is $\hat{\rho}^R = \psi^{(R)\dagger}(x)\psi^{(R)}(x)$, and the field operators contain only Fourier harmonics with positive wave vector, $\psi^{(R)}(x) = \int_{k>0} \psi_k e^{ikx} \frac{dk}{2\pi}$. The coherent states, $|\Psi\rangle =$ $e^{i\int \Phi(x)\hat{\rho}^R(x)dx}|0\rangle$, are parametrized by the function $\Phi(x)$ with a transparent meaning, $2\pi\rho^{cl} = K_F + 2\pi\Phi'(x)$.

The standard bosonization methods [11] yield

$$\langle \Psi | \psi^{\dagger}(x_1)\psi(x_2) | \Psi \rangle = \frac{e^{i(\Phi(x_1) - \Phi(x_2))}}{i2\pi(x_1 - x_2)}, \qquad (11)$$

where we customarily dispensed with the additive Φ independent term which formally vanishes in the large- K_F limit, $K_F|x_1 - x_2| \to \infty$.

To evaluate the Wigner function, W(x, k, 0), we perform a gradient expansion $\Phi(x+y/2) - \Phi(x-y/2) \approx$ $(k_F''(x)y^3/24) + k_F(x)y$ in the exponent of Eq. (11) and then substitute the result in Eq. (2). The integration over y in Eq. (2) then yields [12]:

$$W(x,k,0) \approx \operatorname{Ai}_1\left(2^{2/3}\kappa(x,0)(k-k_F(x,0))\right),$$
 (12)

where

$$\operatorname{Ai}_{1}(x) = \int_{s}^{\infty} \operatorname{Ai}(s') ds' = \int e^{i\left(\frac{q^{3}}{3} + qx\right)} \frac{dq}{2\pi i q}, \quad (13)$$

and κ is the t = 0 value of

$$\kappa(x,t) = |\partial_x^2 k_F(x,t)/2|^{-1/3}$$
. (14)

The function Ai₁(s) changes monotonically from 1 at $s \rightarrow$ $-\infty$ to 0 at $s \to \infty$ on the scale $|s| \sim 1$. Therefore at large $\kappa(x,0)$ the Wigner function, Eq. (12), approaches its classical limit, $W(x, k, 0) \approx \theta(k_F(x) - k)$, as applied to right-movers, which confirms the ansatz (3).

Estimating $\kappa(x,0)$ with the help of Eq. (1) as $\kappa \sim$ $\Delta \rho^{-1/3} \Delta x^{2/3}$, we see that smearing of the Fermi stepfunction, as described by Eq. (12), is smaller by the parameter $(\Delta N)^{-2/3}$ than the shift ~ $\Delta N/\Delta x$ of the Fermi wave vector $k_F(x)$ from its equilibrium value. The same semiclassical parameter is needed to justify the gradient expansion which we used to derive Eq. (12). To assess the accuracy of that approximation, we may consider the correction, δW , to Eq. (12) resulting from the sub-leading term in the gradient expansion. An order of magnitude estimate yields $\delta W/W \sim \Delta N^{-2/3}$ at $|k - k_F(x)| \lesssim 1/\kappa(x, 0).$

We wish now to propagate the initial conditions (12) in time. It will be more convenient, however, first to switch to a representation involving only the inverse function, $x_F(k,t)$ rather than $k_F(x,t)$. This is achieved by performing a Taylor expansion, $k_F(x,0) - k = [x - x_F(k,0)]/x'_F(k,0) + \ldots$, and retaining only its first term in the argument of the Ai₁ function in Eq. (12). This is justified at $x - x_F(k,0) \ll \Delta x$, and at the same time sufficient to allow the Wigner function to vary between its limiting values of 0 and 1. We also use the identity $\partial_x^2 k_F(x,0)|_{x=F(k)} = -\frac{x''_F(k,0)}{x'_F(k,0)^3}$ in the definition of $\kappa(x,0)$, see Eq. (14), to finally transform Eq. (12) to:

$$W(x,k,0) \approx \operatorname{Ai}_1\left(2^{2/3}\tilde{\kappa}(k)(x-x_F(k,0))\right).$$
(15)

Here

$$\tilde{\kappa}(k) = \tilde{\kappa}(k,t) = \left|\partial_k^2 x_F(k,t)/2\right|^{-1/3} \tag{16}$$

is independent of time, according to Eq. (7).

We are now ready to compute the Wigner function W(x, k, t) and quantum corrections to density at later times. Combining Eqs. (5), (7), and (15) we easily find:

$$W(x,k,t) \approx \operatorname{Ai}_1\left(2^{2/3}\tilde{\kappa}(k)(x-x_F(k,t))\right).$$
 (17)

Given the validity of Eq. (17) for all times, it remains only to integrate it over k to obtain the density. Using the representation (13) of Ai₁, one arrives at:

$$\partial_x \rho(x,t) \approx \int e^{i \left[\frac{|x_F'(k)|}{24}q^3 + q(x - x_F(k,t))\right]} \frac{dq}{2\pi} \frac{dk}{2\pi}.$$
 (18)

Next we implement a gradient expansion of function $x_F(k,t)$ around a branch meeting point, $x_{\pm}(t)$. Using this expansion in Eq. (18) and changing variables there to $K = k + \frac{q}{2}$ and $Q = k - \frac{q}{2}$, we derive

$$\partial_x \delta \rho^{\pm}(x,t) \tag{19}$$

$$\approx \int e^{i \left[\frac{|x_F''(k_F^{\pm}(t))|}{6} (K^3 - Q^3) + (x - x^{\pm}(t))(K - Q)\right]} \frac{dQ}{2\pi} \frac{dK}{2\pi}.$$

Note that here we are computing only the singular contribution $\delta \rho^{\pm}(x,t)$ next to the branch meeting points,

$$\delta \rho^{\pm}(x,t) = \pm \left(\rho(x,t) - \rho^{cl}(x_{\pm}(t),t) \right), \qquad (20)$$

rather than full density $\rho(x,t)$. Integration over Q and P in Eq. (19) factorize, leading to the result:

$$\partial_x \rho^{\pm}(x,t) \approx -\tilde{\kappa}^{\pm}(t)^2 \left[\operatorname{Ai}(\tilde{\kappa}^{\pm}(t)\delta x_{\pm})\right]^2$$
. (21)

Here the distances, $\delta x_{\pm} = \mp (x - x_{\pm}(t))$, are measured from the classical shock points, and inverse length scales $\tilde{\kappa}^{\pm}(t)$ are related to the parameters of the classical solution $\tilde{\kappa}^{\pm}(t) = (\alpha^{\pm}(t))^{2/3}$, cf. Eq. (10). The density gradient (21) exhibits pronounced oscillations, "quantum ripples" in the vicinity of the classical shock points, see the inset in Fig. 2. At times $t \gtrsim t_S$, the distance between the shock points is $x_+(t) - x_-(t) \sim \Delta x$. The characteristic length scale for the ripples $1/\tilde{\kappa}^{\pm}(t) \sim \Delta x/(\Delta N)^{2/3}$ is parametrically smaller. Making use of the identity Ai" $(x) = xA_i(x)$, one may integrate Eq. (21) to find

$$\delta \rho^{\pm}(x,t)$$
(22)
$$\simeq \tilde{\kappa}^{\pm}(t) \left\{ \left[\operatorname{Ai}'(\tilde{\kappa}^{\pm}(t)\delta x_{\pm}) \right]^{2} - (\tilde{\kappa}^{\pm}(t)\delta x_{\pm}) \operatorname{Ai}^{2}(\tilde{\kappa}^{\pm}(t)\delta x_{\pm}) \right\}$$

Equation (22) together with the derivative (21) is the main result of this Letter. A typical plot of $\rho(x)$ is shown in Fig. 2, with the inset showing the graph of $\partial_x \delta \rho^{\pm}(x)$ around a branch meeting point.

Some elements of the solution and results which we developed for the one-dimensional free-fermion system may shed light on shock waves in higher dimensions and in one-dimensional quantum liquids of interacting particles.

The generalization onto a 3D and 2D free Fermi gases with a unidirectional density modulation and stronglyanisotropic initial momentum distribution is straightforward. Ripple effect should weaken in the case of a more isotropic momentum distribution. Experimental techniques to study the evolution of initial perturbations are available in such quasi 1D geometries [13, 14].

By continuity, the results obtained in this work are applicable to one-dimensional fermions with weak interparticle repulsion or attraction. Furthermore, a combination of the universal nonlinear Luttinger liquid theory [15] with the free-fermion theory developed here points to a similarity between the shock waves in generic onedimensional liquids of interacting fermions or bosons and in a free fermion system. The nonlinear Luttinger liquid description [15] is valid for small and smooth perturbations. Within that description, the interaction renormalizes the velocity, effective mass, and wave functions of the fermionic quasiparticles representing the excitations; the nonlinear dispersion relation of the quasiparticles is retained, allowing for shocks. Furthermore, the pair collisions of quasiparticles are ineffective in one dimension, thus strongly suppressing the dissipation compared to gases in higher-dimensions [3]. The formation of oscillatory patterns in non-dissipative shock wave physics is quite general, and thus we expect that interaction will not change the basic picture described here, namely that of oscillatory patterns following the formation of shocks.

The exact form of the oscillations in strongly interacting systems is beyond the scope of this letter (some results are available for long range interaction [10] or interacting bosonic systems [13, 16]). The dynamics of bosons with short-range repulsion in one dimension (the Lieb-Liniger model [17, 18]) is in fact directly related to the dynamics of fermions described by the nonlinear Luttinger liquid theory [3].

Numerics for free fermions was performed recently in Ref. [19]. The shape of the ripples in the entire interval $x_{-} < x_{+}$ is different from our predictions, probably be-

cause of a relatively small value of ΔN employed in the numerics [19]. Finally, there is quite close analogy between the structure of ripples we obtained with the ones predicted for one-dimensional spin liquids [20].

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