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Phys. Rev. Lett. **109**, 162503 — Published 18 October 2012

DOI: [10.1103/PhysRevLett.109.162503](https://doi.org/10.1103/PhysRevLett.109.162503)

Observation of large scissors resonance strength in actinides

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The orbital M1-scissors resonance (SR) has been measured for the first time in the quasi-continuum of actinides. Particle- γ coincidences are recorded with deuteron and ^3He induced reactions on ^{232}Th . The residual nuclei $^{231,232,233}\text{Th}$ and $^{232,233}\text{Pa}$ show an unexpectedly strong integrated strength of $B_{M1} = 11 - 15 \mu_N^2$ in the $E_\gamma = 1.0 - 3.5$ MeV region. The increased γ -decay probability in actinides due to the SR is important for cross-section calculations for future fuel cycles of fast nuclear reactors and may also have impact on stellar nucleosynthesis.

PACS numbers: 23.20.-g, 24.30, 27.90.+b

In nuclear systems the width and density of excited states have long been known to increase rapidly with excitation energy creating a region of high level density which is known as the quasi-continuum below the separation energies. The statistical character of the γ -ray decay from the quasi-continuum can only be measured and examined in terms of average electromagnetic properties, referred to as the radiative strength function (RSF).

The importance of the RSF cannot be overstated as it is one of the critical input parameters for calculating neutron-induced reaction cross sections starting from the low keV range. The RSF is relevant to the design of future and existing nuclear power reactors where sophisticated simulations heavily depend on evaluated nuclear data of the many nuclear reactions involved [1]. The RSF also plays a central role in elemental formation during stellar nucleosynthesis [2, 3]. Calculations have shown that relatively small changes to the overall shape of the RSF such as pygmy resonances can have an order of magnitude effect on the rate of elemental formation in the r-process [4]. Often measured RSF are completely lacking or are insufficient to carry out calculations and extrapolations have to be made which can lead to significant uncertainties in the calculated results. Experimental constraints of the RSF in the actinide region will improve the predictive power of reaction modelling.

The γ decay of excited atomic nuclei is to a large extent governed by collective transitions. The softest collective M1 mode, the scissors resonance (SR), appears when the deformed proton and neutron clouds oscillate against each other like the blades of a scissors. Such an isovector collective motion was first predicted by Lo Iudice and Palumbo [5].

The particular situation with the SR built on the ground state has been extensively studied in (γ, γ') and (e, e') reactions. Recently, a review of these experiments and various models has been given [6]. The microscopic description of the SR is based on collective single-particle transitions between orbitals of the same angular momentum ℓ but different $j = \ell \pm 1/2$. For deformed rare-earth nuclei one finds exper-

imentally integrated strengths of $B_{M1} = 3 - 4 \mu_N^2$. However, the SR is not only built on the ground state, but on all states in the nucleus according to the Brink hypothesis [7]. Measurements of the γ -decay between levels in the quasi-continuum show significant higher SR strength. Here, the two-step cascade method and the Oslo method give integrated strengths of $6 - 7 \mu_N^2$ [8, 9].

The spins and parities of some SR states in ^{232}Th and ^{238}U have been determined in (γ, γ') and (e, e') reactions [10]. In addition, (γ, γ') reactions on $^{235,236}\text{U}$ have been reported [11, 12]. Again the measured strengths are only $B_{M1} \sim 3 \mu_N^2$. Since these experiments rest on the identification of single states in an energy region housing $10^4 - 10^5$ levels per MeV, one could expect that not all the strength has been resolved as γ -lines.

In this Letter we report for the first time on the observation of the full SR strength in the quasi-continuum of actinides nuclei. The data show a clear splitting of the SR strength in ^{233}Th . Furthermore, the RSF is found to be independent of the excitation energy in the 3 - 5 MeV region, thus supporting the Brink hypothesis [7].

The Oslo nuclear physics group has developed a method to determine simultaneously the nuclear level density and the RSF from particle- γ coincidences [13, 14]. These quantities provide information on the average properties of excited nuclei in the quasi-continuum and are essential in nuclear reaction theories as they are the only quantities needed for a complete description of the γ decay at higher excitation energies.

The experiments were conducted at the Oslo Cyclotron Laboratory (OCL) with a 12-MeV deuteron and a 24-MeV ^3He beam bombarding a self-supporting target of ^{232}Th with thickness of 0.968 mg/cm². Particle- γ coincidences were measured with the SiRi particle telescope and the CACTUS γ -detector systems [15, 16]. The SiRi detectors were placed in backward direction, covering eight angles from $\theta = 126$ to 140° relative to the beam axis. The front and end detectors had a thickness of 130 μm and 1550 μm , respectively. The CACTUS array consists of 28 collimated $5'' \times 5''$ NaI(Tl) detectors with a total efficiency of 15.2% at $E_\gamma = 1.33$ MeV.

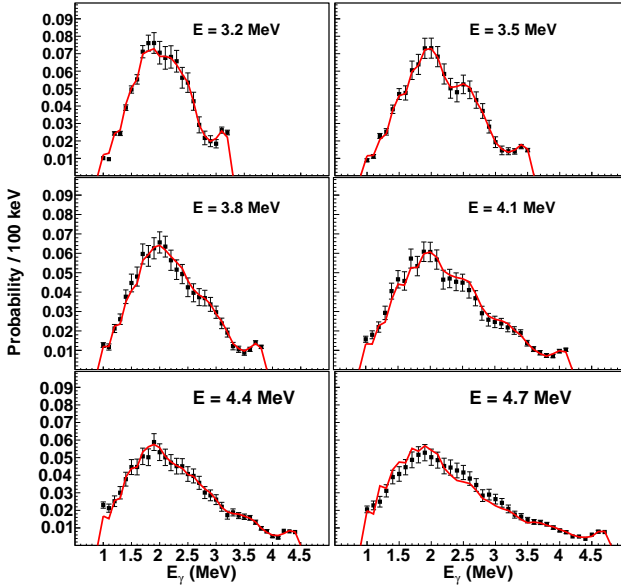


FIG. 1: (Color online) First-generation γ spectra for ^{233}Th . The experimental spectra P (squares with error bars) are compared with the ones obtained from multiplying the extracted \mathcal{F} and ρ functions (red line). The initial excitation energy bins E are 100 keV wide.

The particle- γ coincidences with time information are sorted event by event. From the known charged-particle type and the kinematics of the reaction, the energies deposited in the telescopes can be translated to initial excitation energy E in the residual nucleus. For each energy-bin E , the γ -spectra are then unfolded using the NaI-response function. The resulting matrix describes the γ -ray energy distribution at each bin E and is then the starting point for the Oslo method.

An iterative subtraction technique has been developed to separate out the first-generation (primary) γ transitions from the total γ cascade [17]. It is assumed that the γ -decay pattern is the same whether the levels were initiated directly by the nuclear reaction or by decay from higher-lying states. The technique gives the first-generation matrix $P(E, E_\gamma)$ where E is the initial excitation energy from which the γ cascade starts. In the quasi-continuum, we assume that P is proportional to the level density at the final excitation energy $\rho(E - E_\gamma)$ in accordance with Fermi's golden rule [18]. Furthermore, the decay is governed by the γ -transmission coefficient $\mathcal{F}(E \rightarrow E - E_\gamma)$, which according to the Brink hypothesis [7], is independent of excitation energy:

$$P(E, E_\gamma) \propto \mathcal{F}(E_\gamma)\rho(E - E_\gamma). \quad (1)$$

This allows a simultaneous least χ^2 -fit of the two one-dimensional functions \mathcal{F} and ρ to the two-dimensional matrix P . To test the ansatz (1), we have compared the experimental first-generation spectra and the ones obtained by multiplying the extracted \mathcal{F} and ρ functions, see Fig. 1. The agreement is very good and holds also for the other 11 spectra included in

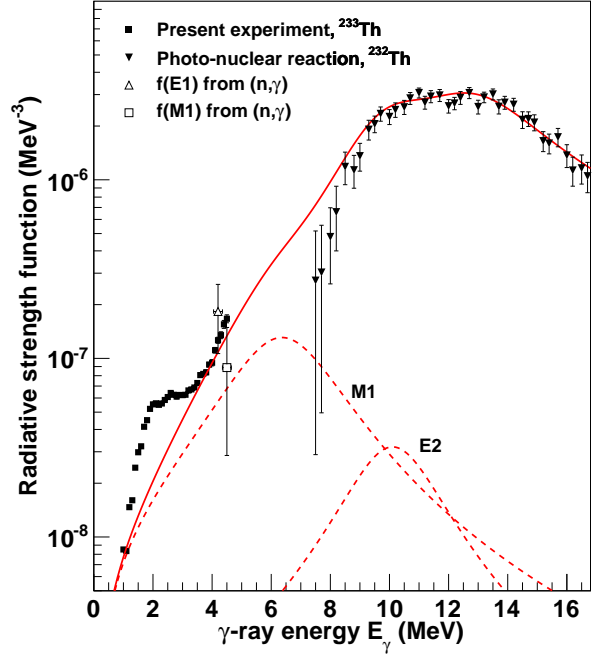


FIG. 2: (Color online) Experimental RSF for ^{233}Th (squares). For comparison, the (γ, abs) reaction on ^{232}Th [24] are shown (triangles). The solid red line is an estimate of the tail of the ^{232}Th GEDR (including the GMDR) used for the "background" of the SR. Also two data points (open symbols) from the (n, γ) reactions are shown [23]. The resonance parameters (ω , Γ and σ) of the EGLO model [23] are (10.9 MeV, 5.66 MeV, 211.4 mb) and (13.87 MeV, 6.68 MeV, 314.3 mb) for the two GEDRs using $T = 0$ MeV. The GMDR parameters are (6.66 MeV, 4.0 MeV, 9.83 mb). The GEQR contribution is negligible.

the global fit (not shown).

The two functions have to be normalized, which means that the unknown parameters A , B and α in the transformations

$$\tilde{\rho}(E - E_\gamma) = A \exp[\alpha(E - E_\gamma)]\rho(E - E_\gamma), \quad (2)$$

$$\tilde{\mathcal{F}}(E_\gamma) = B \exp(\alpha E_\gamma)\mathcal{F}(E_\gamma) \quad (3)$$

must be determined from other experiments. The parameters A and α can be determined by normalizing ρ to known levels at low excitation energy [19] and to the level density extracted from neutron resonance spacings D . We assume a spin distribution [20]

$$g(E, I) \simeq \frac{2I+1}{2\sigma^2} \exp[-(I+1/2)^2/2\sigma^2], \quad (4)$$

where E is excitation energy, I is spin and σ is the spin cut-off parameter. For the actinides studied here, we have typically $\sigma(S_n) = 6 - 7\hbar$, which gives significantly more high-spin states than populated in the light ion reactions used [21]. Thus, the total experimental level density has to be multiplied

TABLE I: Parameters used for normalization of level density and radiative strength function (see text).

Nucleus	S_n (MeV)	a (MeV ⁻¹)	E_1 (MeV)	$\sigma(S_n)$	D (eV)	$\rho(S_n)$ (10 ⁶ MeV ⁻¹)	$x\rho(S_n)$ (10 ⁶ MeV ⁻¹)	$\langle\Gamma_\gamma(S_n)\rangle$ (meV)
²³¹ Th	5.118	23.91	-0.408	6.20	9.6(15)	8.1(16)	4.7(8)	26(2)
²³² Th	6.438	24.00	0.673	6.28	1.17(35) ^a	12.7(38) ^a	6.9(21) ^a	33(10) ^a
²³³ Th	4.786	24.09	-0.389	6.13	16.5(4)	4.4(6)	2.21(29)	24(2)
²³² Pa	5.553	24.00	-1.155	6.52	0.53(16) ^a	42(12) ^a	20(6) ^a	33(10) ^a
²³³ Pa	6.529	24.09	-0.181	6.54	0.42(8)	44(10)	24(5)	33(10) ^a

^a) Estimated values from systematics

with a reduction factor to serve as normalization to the experimental ρ by $x \sim \sum_{I=I_{\min}}^{I_{\max}} g(S_n, I)$, where I_{\min} and I_{\max} define the reaction spin window.

The last parameter B can be determined by reproducing the total γ -radiation width $\langle\Gamma_\gamma\rangle$ from neutron resonance data. In the present work we have followed the normalization procedure of [22] and references therein. In the cases where neutron resonance data are missing, we use values based on the systematics of neighboring nuclei. The parameters applied for the normalizations are listed in Table I. The level density parameter a and back-shift parameter E_1 are used to estimate the total level density ρ from the level density spacing D from neutron resonance capture.

Provided that dipole radiation is dominant in the quasi-continuum, the RSF can be calculated from the normalized transmission coefficient by [23]

$$f(E_\gamma) = \frac{1}{2\pi} \frac{\tilde{\mathcal{T}}(E_\gamma)}{E_\gamma^3}. \quad (5)$$

Figure 2 shows the RSF for ²³³Th together with the GEDR data [24] on ²³²Th.

The observed excess in the RSF is interpreted as the SR for several reasons. It is positioned around $E_\gamma = 2.2$ MeV, which fits the systematics from nuclei studied in the rare-earth region. Also, previous measurements for the SR built on the ground state [10–12] reveal centroids around 2.2 MeV of excitation energy, and several states in these studies are proven to be populated by M1 transitions. To our knowledge, the SR is the only known candidate for a soft collective mode at these energies.

In order to extract the SR component of the total RSF, we use the extrapolated tail of the giant electric dipole (GEDR) and giant magnetic dipole (GMDR) resonances, see solid line of Fig. 2. This tail is tailored to fit the low and high γ -energy part of the experimental RSF data of the various actinides studied. Of course, this is not an ideal situation, and only photo-nuclear reaction data for ²³²Th exist in literature. On the other hand, the background contributes only $\sim 30\%$ to the total RSF. The uncertainty in f_{GEDR} will introduce some errors in the total SR strength, but is negligible for the centroid and the width of the SR energy distribution.

The present ²³²Th(d,p)²³³Th experiment gave a rich data set. In fact, two statistically independent data sets of the

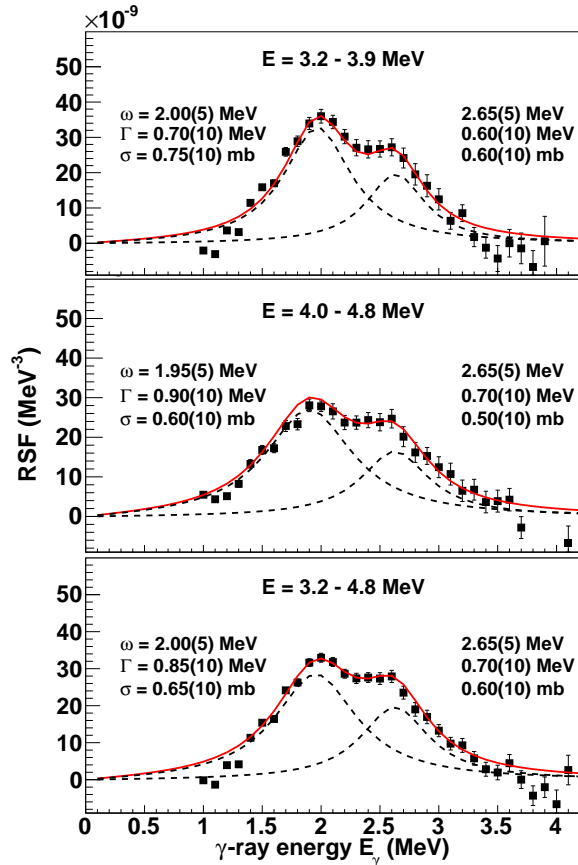


FIG. 3: (Color online) The observed SR for various excitation-energy regions of ²³³Th. The strengths are obtained by subtracting the underlying tail of the GEDR and GMDR (see Fig. 2). The data of the two upper panels are based on statistically independent data sets. The RSF for the lower panel is extracted for simultaneous fitting of the two data sets together, giving approximately the average of the two upper RSFs. The resonance centroid ω , width Γ and strength σ are given for the lower and higher resonance components.

$P(E, E_\gamma)$ matrix could be analyzed and compared. Figure 3 shows the SR energy distributions and resonance parameters of ²³³Th from initial excitation energies $E = 3.2 - 3.9$ MeV and $E = 4.0 - 4.8$ MeV. In the lower panel, the result from a fit

TABLE II: Scissors mode parameters (see text).

Nuclide	δ	ω_{M1} MeV	B_{M1} μ_N^2	$\omega_{M1}S_{-1}$ μ_N^2
^{231}Th	0.183	2.49(20)	11.2(30)	17.4
^{232}Th	0.192	2.23(20)	13.8(40)	15.8
^{233}Th	0.200	2.24(10)	15.3(20)	16.0
^{232}Pa	0.192	2.14(20)	14.7(40)	15.1
^{233}Pa	0.192	2.29(20)	12.7(30)	16.3

to the whole excitation region, $E = 3.2 - 4.8$ MeV, is shown for comparison. The centroids and the strength of the two SRs are almost identical. The width increases somewhat in the higher excitation-energy region, but a corresponding reduction in the σ parameter maintains the strength. We conclude that the similarity of the two distributions supports the validity of the Brink hypothesis in this energy and mass region.

It is evident from the two data sets that the SR is split into two Lorentzians. From the resonance parameters of Fig. 3, the integrated B_{M1} strengths of the components can be calculated by

$$B_{M1} = \frac{9\hbar c}{32\pi^2} \left(\frac{\sigma\Gamma}{\omega_{M1}} \right), \quad (6)$$

giving for the whole excitation region a strength of the first resonance of $B_{M1} = 9.7(15)\mu_N^2$ and the second of $5.6(7)\mu_N^2$. The corresponding energy splitting is $\Delta\omega_{M1} = 0.65(3)$ MeV.

There are various models for the SR properties [6], and we choose here the sum-rule approach [25]. By multiplying the resonance centroid ω_{M1} with the inverse energy-weighted sum rule S_{-1} we obtain the SR strength:

$$B_{M1} = \omega_{M1}S_{-1} = \omega_{M1} \frac{3}{16\pi} \Theta_{IV}(g_p - g_n)^2 \mu_N^2. \quad (7)$$

We use bare gyromagnetic factors for the protons ($g_p = 1$) and neutrons ($g_n = 0$). Since our measurements are in the quasi-continuum, the isovector moment of inertia Θ_{IV} is taken as the rigid-body moment of inertia $\Theta_{\text{rigid}} = \frac{2}{5}m_N r_0^2 A^{5/3}(1 + 0.31\delta)$ with $r_0 = 1.15$ fm and δ is the nuclear quadrupole deformation taken from [26]. Figure 4 displays the extracted SR energy distributions for the five nuclei measured in this work. The centroids ω_{M1} and strengths B_{M1} are summarized in Table II. The agreement with the predicted sum-rule strength is gratifying.

Although not completely evident for all five nuclei, we assume that the SRs have two components as shown as dashed Lorentzians in Fig. 4. Typically, the experimental splitting is $\Delta\omega_{M1} \sim 0.7$ MeV, and the ratio of the strengths between the lower and upper resonance components is $B_2/B_1 \sim 2$.

The splitting could be due to γ deformation, which splits the SR into three components [25] where the two first have

$$\omega_1 = (\cos\gamma + \eta \sin\gamma) \omega_{M1}, \quad B_1 = \frac{1}{2} (\cos\gamma + \eta \sin\gamma) B_{M1}, \quad (8)$$

$$\omega_2 = (\cos\gamma - \eta \sin\gamma) \omega_{M1}, \quad B_2 = \frac{1}{2} (\cos\gamma - \eta \sin\gamma) B_{M1}, \quad (9)$$

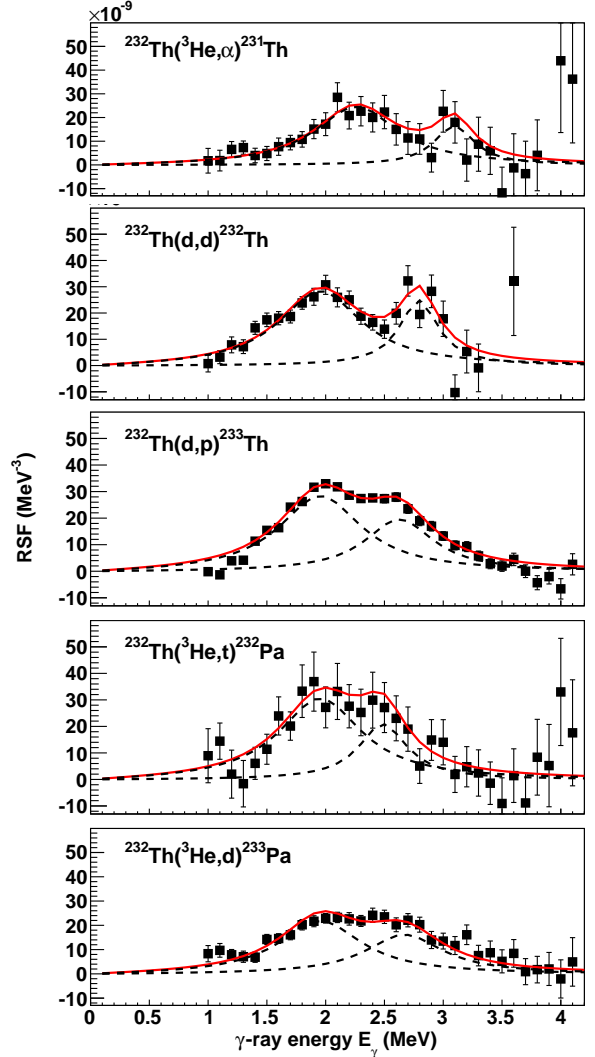


FIG. 4: (Color online) The SR for the five nuclei studied. Approximate GEDR tails are subtracted from the total RSFs as demonstrated in Fig. 2.

with $\eta = \sqrt{1/3}$. In order to describe the observed splitting of ~ 0.7 MeV, a deformation of $\gamma \sim 15^\circ$ is required. With this choice, we obtain theoretically $B_2/B_1 \sim 0.7$, which is not in accordance with the observed ratio. The third SR component is fragmented around $\omega_3 = 2\eta \sin\gamma \omega_{M1} \sim 0.7$ MeV and carries a strength of $B_3 \sim 0.5\mu_N^2$, only. This prediction is very difficult (if at all possible) to verify experimentally.

In conclusion, a large integrated SR strength is found in several actinides with centroids around $\omega_{M1} \sim 2.2$ MeV. The strength, which is well described by the inverse-energy weighted sum rule, is about three times larger than the GEDR contribution.

It is reasonable to believe that the SR will appear for all deformed nuclei in this mass region. The presence of the SR will effectively enhance the γ -decay probability for ex-

citations above the neutron binding energy. As a result, the increased calculated (n, γ) cross sections for actinides with insufficient or lacking experimental data could have a significant impact on fuel-cycle calculations of fast nuclear reactors. In addition, it has the potential of improving the nuclear-physics aspect of the nucleosynthesis in the actinide region.

The energy splitting of the SR could indicate a deformation of $\gamma \sim 15^\circ$. However, theory predicts that the higher SR component has the strongest strength in contradiction with the observations. Therefore, the splitting may be due to other mechanisms.

We would like to thank E.A. Olsen, A. Semchenkov and J. Wikne at the Oslo Cyclotron Laboratory for providing the stable and high-quality deuterium and ^3He beams during the experiment and the Lawrence Livermore National Laboratory for providing the ^{232}Th target. This work was supported by the Research Council of Norway (NFR), the French national research programme GEDEPEON, the US Department of Energy under Contract No. DE-AC52-07NA27344, and the National Research Foundation of South Africa.

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