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# Improved constraints on an axion-mediated force

S.A. Hoedl,\* F. Fleischer, E.G. Adelberger, and B.R. Heckel

*Center for Experimental Nuclear Physics and Astrophysics,*

*Box 354290, University of Washington,*

*Seattle, Washington 98195-4290, USA*

## Abstract

Low mass pseudo-scalars, such as the axion, can mediate macroscopic parity and time-reversal symmetry violating forces. We searched for such a force between polarized electrons and unpolarized atoms using a novel, magnetically unshielded torsion pendulum. We improved the laboratory bounds on this force by more than 10 orders of magnitude for pseudo-scalars heavier than 1 meV, and for the first time, have constrained this force over a broad range of astrophysically interesting masses (10  $\mu\text{eV}$  to 10 meV).

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7 Low mass pseudo-scalars are predicted by string theories [1] and many extensions to the  
8 standard model. Such particles are typically associated with the spontaneous breaking of  
9 a new symmetry at a very high energy scale. The most well developed pseudo-scalar, the  
10 axion [2], is a consequence of Peccei and Quinn’s solution to the “Strong-CP” problem:  
11 QCD explicitly violates the product of charge and parity (CP) reversal symmetry, so that  
12 the CP violating parameter,  $\theta_{QCD}$ , is expected to be of order unity; however, bounds on the  
13 electric dipole moment of the neutron [3] and the Hg atom [4] constrain  $\theta_{QCD} \leq 3 \times 10^{-10}$ .  
14 To solve this fine-tuning problem, Peccei and Quinn proposed a new symmetry [5] that  
15 spontaneously broke in the very early Universe, dynamically minimized  $\theta_{QCD}$  and generated  
16 the axion. Today, these axions would compose at least some of the cold dark matter [6].

17 Cosmology, astrophysics and laboratory experiments constrain the axion’s mass and/or  
18 couplings. The present dark matter density imposes a model-dependent lower bound of  
19 about  $10 \mu\text{eV}$  [7]. Estimations of stellar and supernovae energy loss rates impose an upper  
20 bound of about  $10 \text{ meV}$  [8]. Heavier axions that couple so strongly that they do not escape  
21 from stars or supernovae are precluded by laboratory-based particle physics experiments [9]  
22 and by limits on hot dark matter [10]. The constraints on the mass of pseudo-scalars other  
23 than axions must be evaluated for each model individually. Nevertheless, the axion mass  
24 bounds define an “axion-window” where most axion and pseudo-scalar search efforts are  
25 focused. In the remainder of this paper, we refer to both axions and similar pseudo-scalars  
26 as axion-like particle (ALPs).

27 Most ALP searches look for the conversion of a galactic [11], solar [12] or laboratory  
28 [13] origin ALP into a photon in the presence of a static magnetic field. However, any ALP  
29 that couples with both scalar and pseudo-scalar vertices to fundamental fermions would also  
30 mediate a parity and time-reversal symmetry violating (PTV) force [14] between a polarized  
31 electron and an unpolarized atom, described by the potential:

$$V(\hat{\sigma}, \mathbf{r}) = \frac{\hbar^2(\hat{\sigma} \cdot \hat{r})}{8\pi m_e} \left( \frac{g_s^a g_p^e}{\hbar c} \right) \left( \frac{1}{r\lambda_{\text{ALP}}} + \frac{1}{r^2} \right) e^{-r/\lambda_{\text{ALP}}} \quad (1)$$

32 where  $\mathbf{r}$  is the electron-atom separation vector,  $\lambda_{\text{ALP}} = m_{\text{ALP}}/\hbar c$  is the ALP Compton wave-  
33 length,  $\hat{\sigma}$  and  $m_e$  are the spin unit-vector and mass of the polarized electron respectively.  
34 The ALP pseudo-scalar coupling constant to a polarized electron is  $g_p^e$ , and the ALP scalar  
35 coupling constant to an unpolarized atom is  $g_s^a = Z(g_s^e + g_s^p) + Ng_s^n$ , where  $Z$  and  $N$  are  
36 the proton and neutron numbers respectively, and  $g_s^e$ ,  $g_s^p$  and  $g_s^n$  are the scalar couplings to

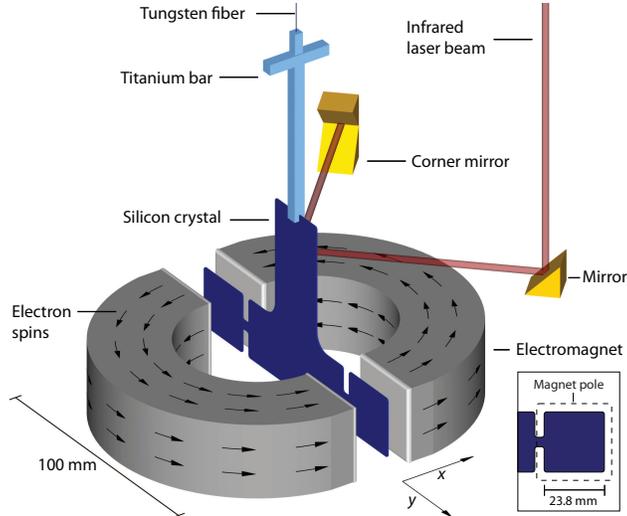


FIG. 1. (color online). A scale drawing of the torsion pendulum apparatus. The gap between the magnet halves, either 2.62 mm or 5.44 mm, is exaggerated for clarity. The inset shows the location of one cut-out in the silicon crystal with respect to the pole-face edges.

37 electrons, protons and neutrons respectively. For simplicity, we assume that  $g_s^e = 0$  and that  
 38  $g_s^p = g_s^n = g_s^N$ , where  $g_s^N$  is the scalar coupling constant to a nucleon, so that  $g_s^a = (Z+N)g_s^N$ ;  
 39 the DFSZ [15] axion model, motivated by grand-unified theories, makes the same assump-  
 40 tions. Although  $g_s^N g_p^e$  is expected to be very small (for the axion,  $g_s^N \propto \theta_{QCD}$ ), PTV force  
 41 searches have three advantages over more conventional ALP experiments: they do not rely on  
 42 cosmological or astrophysical sources of ALPs; they are sensitive to ALPs that do not couple  
 43 to photons, and most importantly, they can simultaneously probe the entire axion-window,  
 44 which corresponds to a force with a range between 0.02 mm and 20.0 mm.

45 Previous PTV searches [16, 17] minimized background magnetic forces with magnetic  
 46 shields that necessarily limited their sensitivity to a short-range PTV force. We overcame  
 47 this limitation with an apparatus, shown in Fig. 1, that allowed us to observe an unshielded,  
 48 highly non-magnetic torsion pendulum suspended by a tungsten fiber between two halves of  
 49 a stationary split toroidal electromagnet. The pendulum consisted of semiconductor-grade  
 50 silicon single-crystal attached to an ultra-pure titanium bar. An autocollimator monitored  
 51 the pendulum's twist oscillation. The magnet-off pendulum period and Q were typically  
 52 280 s and 2500 respectively. The magnet-on Q was position dependent and ranged from 100  
 53 at  $|x| = 1.2$  mm, to 1000 at  $|x| = 0.0$  mm. A double layer mu-metal cylinder surrounded  
 54 the magnet and reduced laboratory magnetic fields by a factor of  $10^4$ .

55 A PTV force between the polarized electrons in the electromagnet and the unpolarized  
 56 silicon atoms in the pendulum generates a magnetic-field-dependent (MFD) torque on the  
 57 pendulum, given by  $(g_s^N g_p^e / \hbar c) G(x, \lambda_{\text{ALP}}) B$ , where  $G(x, \lambda_{\text{ALP}})$  is a geometrical factor calcu-  
 58 lated by integrating the PTV force over the polarized electrons in the magnet and the silicon  
 59 atoms in the pendulum,  $x$  is the distance between the pendulum and the symmetry plane  
 60 between the magnet halves, and  $B$  is the magnetic field in the gap between the magnet  
 61 halves. The magnetic field generates two spurious MFD torques. First, frozen ferromag-  
 62 netic impurities (FFIs) on the surface or in the bulk of the silicon create a MFD torque, given  
 63 by  $\mu_{\text{FFI}} B$ , where  $\mu_{\text{FFI}}$  is the fixed total magnetic moment perpendicular to the magnetic field  
 64 and the fiber axis. Second, because silicon is diamagnetic, the pendulum seeks a rotation  
 65 angle that minimizes the integral of  $B^2$  over the pendulum's volume. The magnetic field,  
 66 therefore, acted as a torsion spring, the “magnet spring,” that was typically much stronger  
 67 than the fiber spring itself. The magnet spring could mimic a MFD torque if its equilibrium  
 68 angle were correlated with the magnetic field polarity.

69 Our pendulum design suppressed these spurious MFD torques. Metallic impurities in  
 70 semi-conductor grade silicon are typically less than 1 ppb by atom [18]; they are also diffuse  
 71 and unlikely to be FFIs because of the single-crystal nature of the material. We removed  
 72 FFIs on the silicon surface by cleaning the pendulum according to the RCA1 and RCA2  
 73 protocols [19] that typically leave less than  $10^{10}$  metal atoms per  $\text{cm}^2$ . The strength of  
 74 the magnet spring was reduced by cut-outs in the silicon that minimized material in the  
 75 magnetic field gradients at the pole-face edges (see inset of Fig. 1). A  $300\text{\AA}$  thick coating  
 76 of paramagnetic terbium canceled the silicon's diamagnetism to within 5%. For a magnet  
 77 gap of 5.44 mm, additional fine-tuning of the magnet spring was necessary and achieved by  
 78 cooling the pendulum to  $5^\circ\text{C}$  (terbium's susceptibility is a strong function of temperature  
 79 [20]).

80 The MFD torque was determined by observing the motion of the pendulum while repeat-  
 81 edly cycling between opposite magnet current states (Fig. 2a). For each magnet state, we  
 82 monitored the pendulum's oscillation for an integral number of periods and then determined  
 83 its equilibrium angle, oscillation amplitude, phase and period by a non-linear least squares  
 84 fit. The data were divided into 12-cycle runs. The magnet spring equilibrium angles,  $\theta_m^\pm$ , and  
 85 spring constants,  $\kappa_m^\pm$ , could differ in the clockwise (+) and counter-clockwise (−) magnet  
 86 states. The observed equilibrium angles of the pendulum,  $\theta^\pm$ , must satisfy the equations:

$$\pm \bar{B} \left( (g_s^N g_p^e / \hbar c) G(x, \lambda_{\text{ALP}}) + \mu_{\text{FFI}} \right) + \kappa_m^\pm (\theta^\pm - \theta_m^\pm) + \kappa_f (\theta^\pm - \theta_f) = 0, \quad (2)$$

87 where  $\theta_f$  and  $\kappa_f$  are the equilibrium angle and spring constant of the fiber spring re-  
 88 spectively, and  $\bar{B}$  is the average absolute value of the magnetic field (measured with a Hall  
 89 probe). We inferred the total spring constant,  $\kappa^\pm = \kappa_f + \kappa_m^\pm$ , from the pendulum period in  
 90 each magnet state. Before and after each data run, we measured  $\theta_f$  and  $\kappa_f$ . From Eq. 2,  
 91 one can show that any MFD torque will generate a normalized torque asymmetry, defined  
 92 as

$$\tilde{N} = \frac{1}{\bar{B}} \left[ \bar{\kappa} \Delta\theta - \kappa_f \frac{\Delta\kappa(\bar{\theta} - \theta_f)}{\bar{\kappa} - \kappa_f} \right], \quad (3)$$

93 where  $\Delta\theta = (\theta^+ - \theta^-)/2$  was determined using a filter that corrected for linear drift,  
 94  $\bar{\theta} = (\theta^+ + \theta^-)/2$ ,  $\Delta\kappa = (\kappa^+ - \kappa^-)/2$  and  $\bar{\kappa} = (\kappa^+ + \kappa^-)/2$ . The second term corrected for  
 95 the spurious MFD torque created by a misalignment between the fiber and magnet-spring  
 96 equilibrium angles. (The correction was typically smaller than the scatter among the cycles  
 97 in each run). A PTV force, a FFI, and an asymmetry in the magnet spring equilibrium angle  
 98 could all contribute to  $\tilde{N}$  according to  $\tilde{N} = (g_s^N g_p^e / \hbar c) G(x, \lambda_{\text{ALP}}) + \mu_{\text{FFI}} + \Delta\theta_m (\bar{\kappa} - \kappa_f) / \bar{B}$ ,  
 99 where  $\Delta\theta_m = (\theta_m^+ - \theta_m^-)/2$ . Nevertheless, the PTV force has a unique signature: its finite  
 100 range requires that  $G(x, \lambda_{\text{ALP}})$  must increase approximately as  $\cosh(x/\lambda_{\text{ALP}})$  as the pendulum  
 101 approaches either magnet half. Because spurious MFDs should not exhibit this behavior (we  
 102 test for this below), we could identify a true PTV force in the presence of spurious forces by  
 103 measuring  $\tilde{N}$  at different values of  $x$ .

104 We found that the pendulum's equilibrium angle depended on its horizontal position so  
 105 that the autocollimator could only observe the pendulum when it was positioned on two  
 106 lines in the horizontal plane (see inset, Fig. 2b). We measured  $\tilde{N}$  at eight positions on these  
 107 lines. Figure 2b plots a typical set of  $\tilde{N}$ s. Their position dependence was not consistent  
 108 with a PTV force, but could be modeled by a function of the form  $\tilde{N} = ax + by + c$ , where  $x$   
 109 and  $y$ , defined in Fig. 1, were measured from the axial symmetry axis of the magnet halves,  
 110 and  $a \ll b$ . At a fixed pendulum position,  $\tilde{N}$  also depended on the history of the magnet. In  
 111 particular, after demagnetizing the magnet by applying a linearly decreasing (9000 s decay  
 112 time) harmonic current (36 s period),  $\tilde{N}$  could change by up to ten times the 1- $\sigma$  uncertainty

113 expected given the scatter among the 12 cycles in each run. The demagnetization-induced  
 114 changes in  $\tilde{N}$  at different pendulum positions were correlated.

115 The demagnetization behavior and the linear  $y$  dependence can be explained by a degauss-  
 116 dependent asymmetry in  $\Delta\theta_m$ . The magnet spring can be considered as two effective linear  
 117 springs connected between the pendulum and the inner edge of each magnet pole-face. An  
 118 asymmetry in the magnetic fields at the pole-face edges would create a difference in the  
 119 equilibrium position of these linear springs in the two magnet states that would in turn  
 120 create a  $\Delta\theta_m$ . Moving the pendulum along  $y$  changed the distance between the fiber axis  
 121 and each linear spring. We expect  $\Delta\theta_m$  to depend linearly on  $y$  for pendulum displacements  
 122 that are small compared to the distance between the fiber axis and the pole-face edges. (The  
 123 pendulum was translated by  $\pm 1$  mm along  $y$  while the fiber is 29 mm from the pole-face  
 124 edge).

125 We constrained  $g_s^N g_p^e / \hbar c$  in the presence of the demagnetization behavior and linear  $y$   
 126 dependence by segmenting the data-taking procedure into position scans, each of which  
 127 consisted of a demagnetization, followed by measurements of  $\tilde{N}$  at the different positions,  
 128 and a subsequent fit of the linear model to this data. We repeated the position scan 20  
 129 times with fields of 0.387 T and 0.193 T and 2.62 mm magnet gap, and 23 times with a  
 130 field of 0.109 T and a 5.44 mm magnet gap. The average of the residuals from each fit at a  
 131 fixed  $x$  formed our ALP observable. Figure 2c shows the ALP observable as a function of  $x$   
 132 for the 2.62 mm gap data. The large-gap data extended our sensitivity to larger values of  
 133  $\lambda_{\text{ALP}}$  and also provided an important systematic check of the small gap data. The  $1\text{-}\sigma$  error  
 134 of the ALP observable (calculated from the scatter among the residuals in each data set)  
 135 was dominated by the non-reproducible effects of the demagnetizing procedure (the thermal  
 136 noise is a factor of 100 smaller). For a given  $\lambda_{\text{ALP}}$ , we fit the predicted  $x$ -dependence of the  
 137 PTV force to the ALP observables. (The predicted force conservatively accounted for the  
 138 uncertainties in the pendulum's dimensions and relative position). The fits yielded larger  
 139 than expected  $\chi^2$ . To provide a 95% confidence limit, we inflated the ALP observable error  
 140 bars for each fit so that  $\chi^2/\nu = 1$ , where  $\nu$  is the number of degrees of freedom, and then  
 141 found values of  $g_s^N g_p^e$  so that  $\Delta\chi^2 = 3.95$ . The inflation factors were typically 2.50 and 2.08  
 142 for the 2.62 mm and 5.44 mm gap data respectively. We have no evidence for a PTV force.  
 143 Figure 3 shows our exclusion bounds.

144 The  $\mu_{\text{FFI}}$  contribution is not expected to mimic a PTV force because  $\bar{B}$  differs by less

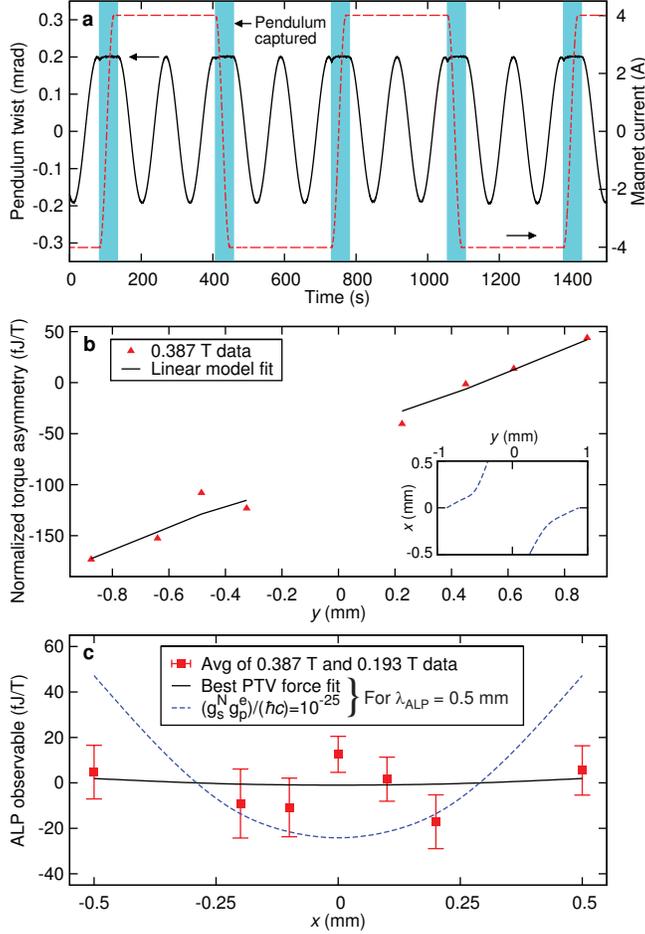


FIG. 2. (color online). (a) A typical interval of slightly more than two complete cycles of 0.193 T data. The shaded regions illustrate when an electrostatic feedback loop stabilized the pendulum. (b) Typical average  $\tilde{N}$  data. The inset shows the positions where the autocollimator could observe the pendulum. (c) The ALP observables, best fit signal and a hypothetical PTV signal for the 2.62 mm gap data set.

145 than a few parts in  $10^3$  over the pendulum's positions. However, the  $\Delta\theta_m$  contribution  
 146 to  $\tilde{N}$  could depend on the pendulum's  $x$ -position in a manner that mimicked a true PTV  
 147 force signal and generated a systematic error. Because the pendulum's equilibrium angle  
 148 depended on its location within the magnet halves, a  $x$ -dependent  $\Delta\theta_m$  would occur if the  
 149 tilt of the apparatus about the  $x$  or  $y$ -axis, the magnet temperature, or the absolute value of  
 150 the magnetic field were correlated with the magnet state. (Because of the magnetic forces,  
 151 the relative position of the magnet halves could depend on the absolute value of the magnetic  
 152 field.) A laboratory magnetic field or field gradient that leaked through the magnetic shields

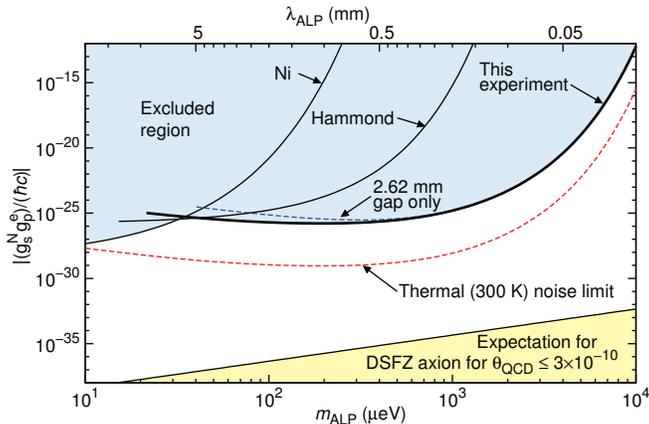


FIG. 3. (color online). The experimental 95% confidence upper limit on  $|g_s^N g_p^e / \hbar c|$ . The force mediated by the DFSZ [15] axion would appear below the bottom-most line. The thermal noise limit represents an ideal torsion pendulum with a magnet-on Q of 3000. See Ref. [21] for bounds on the PTV force between polarized and unpolarized nucleons.

153 would also create a  $\Delta\theta_m$  that could depend on the pendulum position.

154 For each of these systematic errors, we modified the experiment to exaggerate the effect,  
 155 measured  $\tilde{N}$  at eight pendulum positions and solved for a PTV force using the same method  
 156 employed to analyze the ALP observable. The systematic error corrections listed in Table 1  
 157 were calculated by multiplying  $g_s^N g_p^e / \hbar c$ , extracted from the exaggerated data, by the ratio  
 158 of the normal to the exaggerated effects. The total systematic error was less than its  $1\text{-}\sigma$   
 159 uncertainty, which itself was a factor of thirty smaller than the bound on  $g_s^N g_p^e / \hbar c$  for all  $\lambda_{\text{ALP}}$   
 160 so that corrections to the statistical confidence bounds plotted in Fig. 3 were not required.

161 We have substantially improved the bounds on a PTV force between polarized electrons  
 162 and unpolarized nucleons over most of the axion-window, tightening existing constraints  
 163 on ALPs heavier than 1 meV by more than a factor of  $10^{10}$ . Our experimental sensitivity  
 164 was limited by demagnetization scatter and by deviations from a simple model of the lin-  
 165 ear position dependence of the normalized torque asymmetry. We hypothesize that slight  
 166 asymmetries in the magnetic field at the pole-face edges generated both effects. Further  
 167 improvement could be achieved by constructing a magnet from a laminated ferromagnetic  
 168 material that generates a more homogeneous and reproducible magnetic field profile and by  
 169 using a pendulum constructed of a denser material such as germanium. Such an experi-  
 170 ment would more closely approach the thermal limit and could ultimately yield constraints

TABLE I. Systematic error summary for the 2.62 mm gap data. The systematic errors for the 5.44 mm gap data are less than those listed here. The uncertainties for other values of  $\lambda_{\text{ALP}}$  scale with the statistical sensitivity plotted in Fig. 3.

Systematic error	Size of effect		Correction to $g_s^N g_p^e / \hbar c$ for $\lambda_{\text{ALP}} = 0.5$ mm
$y$ -axis tilt	$+2.20 \pm$	3.30 nrad	$(+4.60 \pm 6.90) \times 10^{-28}$
$x$ -axis tilt	$-0.10 \pm$	1.60 nrad	$(-0.23 \pm 3.68) \times 10^{-28}$
$ B $	$+7.0 \pm$	0.8 $\mu\text{T}$	$(0.00 \pm 1.96) \times 10^{-28}$
Magnet Tem.	$-0.32 \pm$	0.27 mK	$(-9.63 \pm 9.49) \times 10^{-29}$
Lab. $B_y$	$+27 \pm$	1 $\mu\text{T}$	$(+1.46 \pm 1.29) \times 10^{-29}$
Lab. $\nabla B_x$	$+3.7 \pm$	0.8 $\mu\text{T}/\text{mm}$	$(-2.55 \pm 9.68) \times 10^{-30}$
Lab. $B_x$	$+23 \pm$	2 $\mu\text{T}$	$(+0.74 \pm 7.77) \times 10^{-29}$
Lab. $\nabla B_y$	$+0.6 \pm$	0.6 $\mu\text{T}/\text{mm}$	$(-1.05 \pm 1.82) \times 10^{-30}$
Total			$(+3.51 \pm 8.12) \times 10^{-28}$

171 a factor of 1000 better than those presented here.

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177 \* sethh@uw.edu

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