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Ajaz Mir, Sanat Tiwari, Abhijit Sen, Chris Crabtree, Gurudas Ganguli, and John Goree Phys. Rev. E **107**, 035202 — Published 9 March 2023 DOI: 10.1103/PhysRevE.107.035202

## <sup>1</sup> Synchronization of Dust Acoustic Waves in a forced Korteweg-de Vries-Burgers model

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(Dated: February 16, 2023)

The synchronization of dust acoustic waves to an external periodic source is studied in the framework of a driven Korteweg-de Vries-Burgers equation that takes into account the appropriate nonlinear and dispersive nature of low frequency waves in a dusty plasma medium. For a spatio-temporally varying source term the system is shown to demonstrate harmonic (1:1) and super-harmonic (1:2) synchronized states. The existence domains of these states are delineated in the form of Arnold tongue diagrams in the parametric space of the forcing amplitude and forcing frequency and their resemblance to some past experimental results is discussed.

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#### I. INTRODUCTION

The nonlinear phenomenon of frequency synchro-16 <sup>17</sup> nization is ubiquitous in many physical, chemical, and bi-18 ological systems and has been the subject of a large number of studies over the past several years [1–3]. The sim-19 plest mathematical model describing this phenomenon 20 consists of an ensemble of globally coupled nonlinear 21 point oscillators that adjust their intrinsic frequencies to 22 common collective frequency as the coupling strength 23 is increased [4–7]. Such a nonlinear phenomenon can also 24 be observed in a continuum medium (a fluid) where a self-25 excited oscillation or a wave can interact with a driving 26 force and adjust its oscillation or wave frequency [8–13]. 27 A plasma system with its wide variety of collective modes 28 and complex nonlinear dynamics provides a rich and chal-29 <sup>30</sup> lenging medium for the exploration of synchronization phenomena. A number of past experimental studies have 31 examined the driven response of a plasma to an external 32 frequency source [9–11, 14–23]. These studies include the 33 synchronization of waves and oscillations at ion and dust 34 dynamical scales as well as chaos and wave turbulence. 35 There have also been a few studies devoted to an inves-36 tigation of mutual synchronization between two plasma 37 38 devices [24–26].

More recently, synchronization phenomena have 39 been experimentally explored in dusty plasma devices 40 where it is easy to visualize the low-frequency wave ac-41 <sup>42</sup> tivity using fast video imaging. A dusty plasma is a fourcomponent plasma of electrons, ions, neutral gas atoms, 43 and micron-size particles of solid matter [27–29]. It can 44 be produced in a laboratory device like a glow discharge 45 plasma, by introducing micron sized solid particles [30– 46 47 33]. These small solid particles (dust) get negatively <sup>48</sup> charged by absorbing more electrons which have a higher <sup>49</sup> mobility than ions. Such a charged medium consisting of dust, ions and electrons, can sustain a variety of collec-50

<sup>51</sup> tive modes [29, 34–36]. The dust acoustic wave (DAW) <sup>52</sup> or dust density wave (DDW) first theoretically predicted <sup>53</sup> by Rao, Shukla and Yu [37] is one such well known low <sup>54</sup> frequency compressional mode that is analogous to the <sup>55</sup> ion acoustic wave [29, 38]. A DAW can be spontaneously <sup>56</sup> excited due to the onset of an ion-streaming instability. 57 The DAW has a very low frequency (typically 10-100 <sub>58</sub> Hz) [14, 30] due to the large mass of the dust parti-<sup>59</sup> cles and can consequently be visually observed; through <sup>60</sup> its images and video recording [31, 39–41]. The term <sup>61</sup> 'dust density wave' originated as a generalization of 'dust <sup>62</sup> acoustic wave', after observing wavefronts (visible in the <sup>63</sup> dust cloud) that appeared to be oblique with respect to <sup>64</sup> the ion drift direction [42]. Two key factors led to the <sup>65</sup> use of the term DDW, namely, the presence of ion drift <sup>66</sup> and an oblique orientation of the wavefront and its prop-<sup>67</sup> agation, with respect to the ion drift. Since then, many <sup>68</sup> research groups have used the term 'dust density wave' <sup>69</sup> and 'dust acoustic wave' synonymously [14, 31, 35, 43– <sup>70</sup> 45]. The present work focuses on the synchronization of 71 DAW using the forced Korteweg-de Vries-Burgers (fKdV-72 B) model.

Synchronization of dust acoustic waves has been 73 <sup>74</sup> studied in an anodic plasma [15], radio-frequency (RF) <sup>75</sup> and direct-current (DC) plasmas [14, 16, 46]. Pilch et <sup>76</sup> al. [15] reported the entrainment of DAWs through a <sup>77</sup> driving modulation to the anode. Ruhunusiri *et al.* [14] 78 reported observation of harmonic, super-harmonic, and 79 sub-harmonic synchrony of self-excited cnoidal DAWs. <sup>80</sup> This was achieved through the driven modulation of the <sup>81</sup> streaming ions in the dust cloud. Their experiments <sup>82</sup> showed parametric regions for the occurrence of such syn-<sup>83</sup> chrony in the form of Arnold tongue diagrams in the <sup>84</sup> state space of the driving frequency and driving ampli-<sup>85</sup> tude. They also observed features like the branching of <sup>86</sup> the tongues and the existence of an amplitude threshold <sup>87</sup> for synchronization to occur. Williams *et al.* [16] com-<sup>88</sup> pared DAW synchronization in RF and DC generated <sup>89</sup> plasmas. Their results suggested that in a RF plasma, <sup>90</sup> synchronization was restricted to a part of the dust cloud <sup>91</sup> volume unlike the complete dust cloud synchrony in a <sup>92</sup> DC discharge plasma. Deka et al. [46] observed the syn-

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<sup>93</sup> chronization of self-excited DDW, through the suppres-<sup>149</sup> a generalization of the fKdV model that was developed 94 95 96 97 98 99 100 <sup>101</sup> synchronization of a DDW driven by an ion flow. Unlike <sup>157</sup> framework for the study of synchronization in a realistic 102 103 104 105 perpendicular to the ion flow direction. 106

107 108 110 <sup>111</sup> els. One of the commonly employed mathematical model <sup>167</sup> collected at one spatial location. A parametric plot in <sup>112</sup> is the periodically forced Van der Pol (fVdP) oscilla- <sup>168</sup> the form of an Arnold tongue diagram shows multiple 113 tor [1, 3, 49],

<sup>114</sup> 
$$\frac{d^2x}{dt^2} - (c_1 - c_2 x^2) \frac{dx}{dt} + \omega_0^2 x = A_{dr} \cos(2\pi f_{dr} t)$$
 (1)

115 which describes the displacement x of a harmonic os-<sup>116</sup> cillator with a natural frequency  $\omega_0$ , with terms for a <sup>117</sup> nonlinear damping  $c_2 x^2 dx/dt$ , a source of energy for self-118 excitation  $c_1 dx/dt$ , and a periodic driving source at a fre-<sup>119</sup> quency  $f_{dr}$ . The fVdP oscillator can exhibit synchroniza-<sup>120</sup> tion not only at  $f_{dr}/f_0 \approx 1$ , which is called "harmonic" <sup>121</sup> synchronization, but at ratios that are rational numbers. <sup>122</sup> If  $f_{dr}/f_0 > 1$ , the synchronization is said to be "super-123 harmonic", whereas if  $f_{dr}/f_0 < 1$  it is "sub-harmonic". 124 Although the VdP oscillator model has been used in <sup>125</sup> the past as a reference for characterizing synchroniza-126 tion phenomena in plasmas and other media that support the propagation of waves [10, 11, 14, 18, 24, 46, 50]. 127 It should be pointed out that as a point oscillator model 128 its dynamics is restricted to nonlinear oscillations and it 129 cannot correctly represent nonlinear waves. This is also 130 <sup>131</sup> evident from the fact that the VdP model is an ordinary differential equation in time and therefore has no spatial 132 133 dynamics that characterizes a propagating wave. In ad-134 dition, for nonlinear dust acoustic or dust density waves dispersion plays an important role in defining their prop-136 agation characteristics and this is not built into the VdP <sup>137</sup> model. As a promising step in capturing spatial properties of a wave, one modelling approach to explain cluster  $_{190}$  Here n(x,t) is the dependent variable (the perturbed 138 139 140 141 142 143 describes the global synchronization of waves exhibiting  ${}^{196} \lambda_D$  and the dust plasma period  $\omega_{pd}^{-1}$ , respectively. 144 both nonlinearity and dispersion, in a plasma medium. <sup>197</sup> 145

146 147 to demonstrate synchronization of nonlinear dust acous- 199 ing term) has been shown to model the evolution of

sion mechanism, by modulating ion streaming using an 150 by Sen et al. [53] for driven nonlinear acoustic waves and external sinusoidal driver. Recently, Liu et al. [47] car- 151 subsequently extensively used to study nonlinear precurried out experiments in the Plasma Kristall-4 (PK-4) de- 152 sor solitons in dusty plasma experiments [54, 55]. For vice on board the International Space Station (ISS) un- 153 our study we include viscous dissipation in the model, der micro-gravity conditions and reported phase locking <sup>154</sup> an important feature of most laboratory studies of dusty for harmonic synchronization. The present work is mo- 155 plasmas [56, 57], which converts the fKdV to a fKdVtivated by Ruhunusiri et al. [14] experiment on global 156 B model. Such a model provides a proper theoretical the DDW in some experiments [42, 48], the wavefronts 158 dispersive plasma system that includes natural growth were not obliquely propagating, as the experiment was <sup>159</sup> and dissipation of waves. The driving term is chosen to designed to have a planar symmetry, provided by prox- 160 have an oscillatory form that has both a temporal and imity to a planar electrode, so the wavefronts were nearly 161 spatial periodicity. Our numerical solution of the model <sup>162</sup> equation, show clear signatures of harmonic (1:1) and Theoretical efforts towards interpretation and phys- 163 super-harmonic (1:2) synchronization. The characterisical understanding of these experimental results have so 164 tic features of the synchronization are delineated using far been limited to providing qualitative comparisons 165 power spectral density (PSD) plots, phase space plots with results obtained from very simple dynamical mod- 166 and Lissajous plots obtained from the time-series data <sup>169</sup> tongues, each corresponding to the existence region of a 170 harmonic or a higher-order super-harmonic synchronized <sup>171</sup> state. The harmonic tongue also show a branching be-172 haviour.

> The rest of the paper is organized as follows. Sec-173 <sup>174</sup> tion II briefly describes the fKdV-B model and the nu-<sup>175</sup> merical approach adopted to solve it. The section also 176 presents some numerical results for the undriven KdV 177 and KdV-B equations as background information on the 178 characteristic nonlinear features of the waves and to de-<sup>179</sup> scribe the diagnostic tools to be used for identifying syn-180 chronization phenomena. Section III presents our main 181 results on harmonic and super-harmonic synchronization 182 using the fKdV-B model. A brief summary and some <sup>183</sup> concluding discussion are provided in section IV.

#### THE FKDV-B EQUATION AND THE II. NUMERICAL APPROACH

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The fKdV-B equation, a one-dimensional driven 186 187 nonlinear partial differential equation, is of the form:

$$\frac{\partial n(x,t)}{\partial t} + \alpha n(x,t) \frac{\partial n(x,t)}{\partial x} + \beta \frac{\partial^3 n(x,t)}{\partial x^3} - \eta \frac{\partial^2 n(x,t)}{\partial x^2} = F_s(x,t).$$
(2)

or partial synchronization of propagating DDWs [51] un- 191 density in this case) and  $F_s(x,t)$  is an external spatioder microgravity conditions [42] used a chain of coupled 192 temporal forcing term.  $\alpha$ ,  $\beta$ , and  $\eta$  are positive quantities Van der Pol oscillators [52]. As a further advance, how- 193 representing the strength of nonlinearity, dispersion, and ever, there remains a need to develop a simple theoret- 194 viscous damping, respectively. The spatial coordinate x ical model, based on a wave equation, that successfully 195 and time t are normalized by the plasma Debye length

It should be mentioned that the KdV equation (*i.e.*, In this paper, we present such a model and use it 198 Eq. (2) in the absence of the viscous damping and driv-148 tic waves to an external driver. The fKdV-B model is 200 weakly nonlinear waves in dusty plasmas both in the

202 Hence it can correctly represent both nonlinear dust den- 258 of the driver. Equation (2) is solved for various values of 203 showed that the cnoidal solution of the KdV shows ex-204 cellent agreement with the DDW profiles observed in the 205 dusty plasma experiments [58, 59]. Theoretically the ex-206 perimental DDW evolution was modelled by the KdV 207 model in which the ion-streaming was taken into consid-208 <sup>209</sup> eration [38]. Earlier, a theoretical model based on the <sup>210</sup> fKdV equation [60] was used to explore the nonlinear <sup>211</sup> mixing of longitudinal dust lattice waves observed in the <sup>212</sup> dusty plasma experiment [61]. Nonlinear mixing means the natural mode and the external forcing mode retain 213 their identity after interaction and excited frequencies 214 are different combinations of addition, and subtraction 215 of the natural and forcing mode. The present theoretical 216 fKdV-B model is proposed to understand the global syn-218 chronization of the dust acoustic wave as was observed in the dusty plasma experiment [14]. Synchronization 219 <sup>220</sup> means the natural mode loses its identity and the sys-221 tem is controlled by the external driver. Here, we model <sup>222</sup> synchronization by incorporating the viscous damping 223 instead of nonlinear mixing as was done in Ref. [60]. The fKdV-B equation can be derived from the full fluid-224 Poisson set of equations, in the weakly nonlinear, dis-225 persive and dissipative regime by using a reductive per-226 turbation method [56, 62]. Such a derivation in the ab-227 sence of the viscosity term has been given in detail by 228 Sen et al. [53]. The KdV-B equation (i.e., Eq. (2) in 229 the absence of the driving term) is well known in the lit-230 <sup>231</sup> erature [56, 57, 62] and has been employed in the past to model oscillatory shocks in dusty plasmas [56, 63]. 232 The model has also been used to study temporal chaos 233 or spatial chaos by using a randomly time varying [64] 234 235 or randomly space varying [65] driving term. In earlier work by Sen *et al.* [53] the source term was taken to be a 236 constant, while in this work, we use a spatio-temporally 237 varying periodic source and carry out a numerical inves-238 tigation of Eq. (2) to study the synchronization of DAWs 239 based on the fKdV-B model. 240

The driving source is taken to be of the form of a 241 242 cnoidal-square travelling wave,

$$F_{s}(x,t) = A_{s}cn^{2}[2K(\kappa_{s})\{x/\lambda_{s} - f_{s}t\}; \kappa_{s}] \qquad (3)$$

<sup>244</sup> where cn is the Jacobi elliptic function,  $A_s$  is the driving <sup>245</sup> amplitude,  $\lambda_s$  is the spatial wave length and  $f_s$  is the <sup>246</sup> driving frequency.  $K(\kappa)$  is the complete elliptic integral <sup>247</sup> of first kind and the elliptic parameter  $\kappa$  is a measure of <sup>248</sup> the nonlinearity of the wave. The cnoidal-square travel-<sup>249</sup> ling wave is an exact solution of the KdV equation. It <sup>250</sup> can therefore mimic the driving of the system by a DAW <sup>251</sup> arising from an external (coupled) plasma source. For  $_{252}$  the numerical solution of Eq. (2) the initial waveform is <sup>253</sup> also taken to be of the form,

<sup>254</sup> 
$$n(x,t=0) = A_0 cn^2 [2K(\kappa_0) \{x/\lambda_0\}; \kappa_0],$$
 (4)

256 the driving source. The idea is to see whether the final 315 governed by the exact solution of the KdV and will be a

<sup>201</sup> presence [38] and in the absence [37] of ion-streaming. <sup>257</sup> driven modes of the system synchronize to the frequency sity and dust acoustic waves. Recently Liu et al. [38]  $_{259}$  f<sub>s</sub> and A<sub>s</sub> in order to find the regions of synchronization 260 in the parameter space of  $(A_s, f_s)$ .

Our numerical investigation of the fKdV-B equation 262 <sup>263</sup> is based on the pseudo-spectral method [66] and uses <sup>264</sup> periodic boundary conditions. The code is first bench <sup>265</sup> marked by reproducing earlier results [53, 67] obtained <sup>266</sup> for the fKdV equation. The various parameter values as-<sup>267</sup> sociated with the model are taken to be as follows: The <sup>268</sup> Jacobi elliptic parameters  $\kappa_0 = \kappa_s = 0.98$  for Eqs. (3)  $_{269}$  and (4). The wave vector of the initial perturbation *i.e.*,  $_{270} k_0 = 12k_m$  where  $k_m = (2\pi)/L_x$  being the minimum <sup>271</sup> wave vector associated with a system of length  $L_x = 6\pi$ . <sup>272</sup> The corresponding wavelength *i.e.*,  $\lambda_0 = (2\pi)/k_0$  and <sup>273</sup> amplitude  $A_0$  of the initial perturbation (*i.e.*, Eq. (4)) 274 are kept fixed throughout the analysis. We have taken  $k_s = 12k_m$  and  $k_s = 2 \times 12k_m$  for studying harmonic 276 (1:1), and super-harmonic (1:2) synchronization states. <sup>277</sup> The corresponding forcing wavelength is  $\lambda_s = (2\pi)/k_s$ . <sup>278</sup> Throughout the analysis, we have only varied the forc-<sup>279</sup> ing amplitude,  $A_s$  and forcing frequency,  $f_s$ . The co- $_{\rm 280}$  efficient  $\alpha$  in Eq. (2) is given by following expression  $_{281} \alpha = \left[\delta^2 + (3\delta + \sigma)\sigma + (\delta/2)(1 + \sigma^2)\right]/(\delta - 1)^2$  [55] and  $_{282} \beta = 0.5$ . We evaluate  $\alpha = 2.3$  with  $\sigma = T_{i0}/T_{e0} = 0.0036$ 283 where electron and ion temperatures are  $T_{e0} = 7$  eV and <sup>284</sup>  $T_{i0} = 0.025 \text{ eV}$ , respectively and  $\delta = n_{i0}/n_{e0} = 3.4$  where <sup>285</sup> electron and ion densities are  $n_{e0} = 2 \times 10^{14} \text{ m}^{-3}$  and <sup>286</sup>  $n_{i0} = 6.8 \times 10^{14} \text{ m}^{-3}$ , respectively. The nonlinearity pa-<sup>287</sup> rameters  $\alpha$  was measured from experimental parameters <sup>288</sup> reported by Flanagan *et al.* [58] for a wave experiment <sup>289</sup> using a setup similar to that of Ruhunusiri *et al.* [14]. <sup>290</sup> Since, there is no measurement of the viscosity param-<sup>291</sup> eter in Flanagan *et al.* [58] and no value is reported for <sup>292</sup> the experimental setup of Ruhunusiri *et al.* [14], we treat 293 the viscosity coefficient to be a free parameter, which <sup>294</sup> we adjust to obtain a good quantitative agreement with the signatures of dissipation in the experimental data of <sup>296</sup> Ruhunusiri et al. [14], namely the Arnold tongues. A <sup>297</sup> value of  $\eta = 0.0025$  best fits the experimental data. Us-<sup>298</sup> ing the experimental plasma parameters [58] and assum-<sup>299</sup> ing dust temperature  $T_d = 2$  eV, we calculate  $\Gamma = 92$ 300 and  $\kappa = 2.8$ . Referring to molecular dynamics simu-301 lations for dusty plasmas for the corresponding closest  $_{302}$   $\Gamma = 100$  and  $\kappa = 3$ , the value of normalized viscosity 303 is  $\eta^* = 0.04$  [68, 69]. This value of viscosity translates  $_{304}$  to  $\eta = 0.0027$  as per the KdV-B equation normaliza-305 tion, which is fairly close to our chosen value of viscosity 306 for the simulations of the fKdV-B model. Furthermore, 307 we take the same experimental values of the natural and  $_{308}$  driver frequencies as reported in the experiment [14] to 309 carry out numerical solutions of the fKdV-B model *i.e.*, <sup>310</sup> Eq. (2). Also, based on the chosen parameters  $\alpha$ ,  $\beta$ ,  $\kappa_0$ and  $k_0$ , the initial perturbation has amplitude  $A_0 = 46.32$  $_{312}$  and frequency  $f_0 = 22$  Hz, which is derived using the re-<sup>313</sup> lationship provided in Mir *et al.* [60]. The amplitude  $_{255}$  with the values of  $A_0$ ,  $f_0$  and  $\lambda_0$  different from those of  $_{314}$  of the initial perturbation chosen in this fashion will be



FIG. 1. PSD for the times-series of KdV (solid line) and KdV-B (dash-dotted line) equations with initial perturbation Eq. (4). Insets (I) and (II) show the phase space plots and time series, respectively for KdV (solid line) and KdV-B (dash-dotted line) models.

<sup>316</sup> stable solution of KdV for this particular amplitude.

We evolve the initial perturbation in Eq. (2) over 317 long times for these various different parameter values. 318 During the spatio-temporal evolution, we collect a time-319 series of the density field at a fixed spatial location and 320 use it to calculate the power spectral density. The PSD 321 provides a useful tool for distinguishing between synchro- <sup>358</sup> III. 322 nized and un-synchronized states. 323

As an illustrative example, we show in Fig. 1 the 359 324 325 326 <sup>327</sup> for  $\eta = A_s = 0$ ). The time-series data has been collected <sup>362</sup> by discussing harmonic (1:1) synchronization for which <sup>328</sup> up to  $t_{max} = 80 \ \omega_{pd}^{-1}$  with a time step  $dt = 10^{-5} \ \omega_{pd}^{-1}$ . <sup>363</sup> we choose the driving frequency to be slightly away from <sup>329</sup> The maximum sampling frequency  $f_S = 1/dt$  and the <sup>364</sup> the fundamental frequency of  $f_0 = 22$  Hz that is char-<sup>330</sup> Nyquist frequency is  $f_N = f_S/2$ . This leads to a fre-<sup>365</sup> acteristic fundamental frequency of the undriven system. 331 332 333 334 335 336 337 338 339 340 341 342 343 344 345 346 347 348 349 350 351  $_{352}$  of Eq. (2) with  $F_s = 0$  and the initial perturbation decays  $_{388}$  persive terms, but also a linear dissipative term, allowed 353 in time. The question is whether by driving the system 389 achieving synchronization of a wave. When we turned

<sup>354</sup> with a periodic source one can revive and sustain a non-<sup>355</sup> linear solution that is also synchronized with the driver. The answer is in the positive and we next present our 356 results on such a phenomenon. 357

### SYNCHRONIZATION IN FKDV-B MODEL

In this section we present the main results of our work, PSD, the time-series and the phase space plot of the so- 360 namely, the synchronization of the solutions of Eq. (2) to lution, obtained for a KdV (solid line) equation (Eq. (2) 361 an external driver of the form given by Eq. (3). We begin quency resolution of  $df = 1/t_{max}$  for the collected time-  $_{366}$  Two cases are considered, namely,  $f_s = 21$  and  $f_s = 23$ series. The time-series data corresponding to the first 308 Hz. The driving amplitude in both cases is taken to be few tens of periods is discarded to remove transient ef-  $_{369}$   $A_s = 0.40A_0$ . Figure 2 shows the attainment of harmonic fects while constructing the PSD. In Fig. 1 the nonlinear 370 (1:1) synchronization for both these cases with subplots character of the mode is evident from the presence of the  $_{371}$  (a, b) devoted to  $f_s = 21$  Hz and (c, d) to  $f_s = 23$  Hz, higher harmonics in the PSD and from the shape of wave 372 respectively. As can be seen from the time-series plots form in the time-series. The natural mode of KdV has a 373 in (a) and (c) the driven solutions are indeed locked to frequency  $f_0 = 22$  Hz. The single cycle phase space plot  $_{374}$  the driver. This is also clearly seen in the PSDs where (solid line) with its form resembling a separatrix curve 375 the fundamental frequencies of the driven solutions are indicates an undamped nonlinear periodic wave, in this 376 indeed at the frequency of the driver. Furthermore, the case the exact cnoidal-square wave. Also, for compari- 377 phase space plots in (b) and (d) show that these solutions son, we present in Fig. 1 the corresponding results for 378 constitute undamped nonlinear periodic waves that are the undriven KdV-B (dash-dotted) equation (Eq. (2) for 379 maintained by a balance between the nonlinear steep- $\eta = 0.0025$  and  $A_s = 0$ ) on top of the KdV (solid line)  $_{380}$  ening, dispersive broadening, viscous damping and amequation. The effect of viscous damping is seen in the fre- 381 plification due to the external pumping by the driving quency shift of the fundamental component in the PSD 382 term. The resultant phase space curve, that has the chartowards a lower value of  $f_0^{\eta} = 15$  Hz, the reduced am- 383 acteristic shape of a separatrix, represents a stationary plitude in the time-series and the spiralling of the phase 384 cnoidal wave solution. The presence of dissipation seems space plot (dash-dotted) towards the origin. It is clear 385 to be necessary for sustaining this synchronized driven that in the presence of finite viscosity the cnoidal-square 386 solution. We have found that in the partial differential wave can no longer be sustained as a nonlinear solution 387 equation Eq. (2), including not just nonlinear and dis-



FIG. 2. The harmonic (1:1) synchronization in the fKdV-B model with  $f_s < f_0$  and  $f_s > f_0$ . The time-series of the fKdV-B model (solid line) and the forcing (dash-dotted line) at driver frequency (a)  $f_s = 21$  Hz with threshold amplitude  $A_s = 0.40A_0$ and (b)  $f_s = 23$  Hz with threshold amplitude  $A_s = 0.40A_0$ . (c) PSD of times-series (a). (d) PSD of time-series (b). The inset (I) is the phase space plot and the inset (II) is the Lissajous figure which reflects the frequency locking at the driver frequency.



FIG. 3. The super-harmonic (1:2) synchronization in the fKdV-B model with  $f_s < 2f_0$  and  $f_s > 2f_0$ . The time-series of the fKdV-B model (solid line) and the forcing (dash-dotted line) at driver frequency (a)  $f_s = 43$  Hz with threshold amplitude  $A_s = 0.60A_0$  and (b)  $f_s = 45$  Hz with threshold amplitude  $A_s = 0.50A_0$ . (c) PSD of times-series (a). (d) PSD of time-series (b). The inset (I) is the phase space plot and the inset (II) is the Lissajous figure which reflects the frequency locking at half of the driver frequency.

<sup>390</sup> off dissipation, by setting the viscosity coefficient to zero <sup>393</sup> different from the case of a point oscillator, as described <sup>391</sup> in Eq. (2), we did not observe synchronization of the <sup>394</sup> by the Van der Pol oscillator Eq. (1), which requires a <sup>392</sup> wave, for the conditions that we studied here. This is <sup>395</sup> nonlinear dissipation term to obtain synchronization. In



FIG. 4. The Arnold tongue diagram for harmonic (1:1) and super-harmonic (1:2) synchronization states in the fKdV-B model. The amplitude is varied from  $A_s = 0.10A_0$  to  $A_s =$ 445  $0.70A_0$  for 1:1, and  $A_s = 0.20A_0$  to  $A_s = 0.70A_0$  for 1:2 synchronization.

<sup>396</sup> the absence of viscosity one only gets nonlinear mixing from the model as has been reported earlier in Mir et397 al. [60, 67]. The amount of viscosity also determines the 398 threshold condition for the driver amplitude. 399

To explore super-harmonic (1:2) synchronization we 400 401 402 403 quency  $2f_0 = 44$  Hz of the undriven system. The results 404 are shown in Fig. 3 where the subplots (a,c) are devoted 405 to  $f_s = 43$  Hz and (b,d) to  $f_s = 45$  Hz, respectively. 458 Ruhunusiri et al. [14]. In particular, comparison of our 406 As in the previous case of harmonic synchronization, we 459 theoretical Arnold tongue diagram with their experimen-<sup>407</sup> see clear evidence of super-harmonic (2:1) synchroniza-408 tion in the time-series plots, the PSDs and the phase 461 experimental Arnold tongue diagram we see the existence 409 411 state. One significant difference from the harmonic syn- 464 tant differences. With our model we have not been able 412 chronization case is that the minimum threshold ampli-465 to obtain sub-harmonic synchronization that have been 413 tude for the driver to achieve a 1:2 state is different for 466 observed in the experiment. Furthermore, our model uses 414 the cases  $f_s < 2f_0$  and  $f_s > 2f_0$ . They are  $A_s = 0.60A_0$  and external driver that closely resembles a nonlinear nat- $_{415}$  and  $A_s = 0.50A_0$ , respectively.

416 417 of the existence domain of these synchronized states in 470 used. However, it is not clear what form this driver takes  $_{418}$  the parameter space of the driver frequency  $f_s$  and driver  $_{471}$  inside the plasma system and whether it manifests it- $_{419}$  amplitude  $A_s$  in form of an Arnold tongue diagram. To  $_{472}$  self as a spatio-temporally varying perturbation. These  $_{420}$  obtain the Arnold tongue diagram,  $A_s$  is varied in steps of  $_{473}$  and other questions, such as the absence of sub-harmonic  $_{421}$  4.63 (which is  $0.10A_0$ ) from 0 to 32.42 (which is  $0.70A_0$ )  $_{474}$  synchronization in the equation, the neglect of dissipa- $_{422}$  while  $f_s$  is varied in steps of 0.5 Hz for harmonic syn- $_{475}$  tion arising from gas friction on the dust particles, etc., <sup>423</sup> chronization and 1.0 Hz for the super-harmonic case. <sup>476</sup> remain to be explored in the future in order to further <sup>424</sup> Fig. 4 shows the 1:1 and 1:2 entrained state tongues in <sup>477</sup> improve the model.

425 the fKdV-B model.

We observe several interesting features in the Arnold 426 <sup>427</sup> tongue diagram. To start with, there is always a thresh- $_{428}$  old amplitude  $A_s$  below which no synchronization occurs. <sup>429</sup> For the harmonic (1:1) synchronization it is  $A_s = 0.10A_0$  $_{430}$  for  $\eta = 0.0025$ . This is unlike the harmonic synchro-<sup>431</sup> nization phenomenon observed in a driven Van der Pol model where no such threshold is found [70]. Another im-432 portant feature is a distinctive branching of the Arnold tongue that is clearly seen for the (1:1) states at low 434 <sup>435</sup> forcing amplitudes marked with arrows. The branching 436 gives rise to a non-synchronized region between the fre- $_{437}$  quencies  $f_s = 22$  Hz to  $f_s = 18$  Hz at driver ampli-<sup>438</sup> tude  $A_s = 0.10A_0$ . This branching narrows down with  $_{439}$  the increase in  $A_s$ . Another branch is seen in between  $_{\rm 440}~f_s=18~{\rm Hz}$  and  $f_s=16.5~{\rm Hz}$  which also narrows down 441 with increase in  $A_s$ . A third feature is the asymmet-442 ric nature of the tongue structures about  $f_0$ . The fre-443 quency width over which synchronization can be obtained <sup>444</sup> is much broader for  $f_s < f_0$  compared to  $f_s > f_0$ .

#### SUMMARY AND CONCLUSIONS IV.

To summarize, we have studied the phenomenon 446 447 of synchronization of dust acoustic waves to an exter-448 nal periodic driver in a model system described by the <sup>449</sup> forced Korteweg-de Vries-Burgers equation. This equa-<sup>450</sup> tion provides a proper theoretical framework and a bet-451 ter physical model compared to the Van der Pol os-<sup>452</sup> cillator model for studying the dynamics of nonlinear <sup>453</sup> dust acoustic waves by properly accounting for nonlinagain consider two cases of  $f_s = 43$  Hz and  $f_s = 45$  Hz  $_{454}$  ear, dispersive and dissipative influences on the waves. which are slightly below and above the first harmonic fre- 455 Using the model, we have successfully demonstrated har-<sup>456</sup> monic (1:1) and super-harmonic (1:2) synchronization <sup>457</sup> states of DAWs for the experimental values reported by 460 tal one shows the following common features. As in the space plots. The Lissajous figures have a number eight- 462 of amplitude thresholds as well as clear evidence of the like trajectory which is indicative of a (1:2) synchronized 463 branching phenomena. However there are also impor-468 ural mode of the system whereas in the experiment a Finally, in Fig. 4 we present a consolidated picture 469 purely time varying external sinusoidal driver has been 479 480 dian Institute of Technology Jammu Seed Grant No. 486 SC0014566, NASA/JPL RSA No. 1672641, and National SG0012. AS is thankful to the Indian National Science <sup>487</sup> Science Foundation Grant No. PHY-1740379. 481 482 Academy (INSA) for the INSA Honorary Scientist posi-

<sup>483</sup> tion. CC and GG acknowledge NASA-JPL subcontract <sup>484</sup> No. 1573108 and NRL Base Funds. JG was supported Work done by ST and AM was supported by the In- 485 by United States Department of Energy Grant No. DE-

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