



# CHORUS

This is the accepted manuscript made available via CHORUS. The article has been published as:

## Enhancing dark matter annihilation rates with dark bremsstrahlung

Nicole F. Bell, Yi Cai, James B. Dent, Rebecca K. Leane, and Thomas J. Weiler

Phys. Rev. D **96**, 023011 — Published 26 July 2017

DOI: [10.1103/PhysRevD.96.023011](https://doi.org/10.1103/PhysRevD.96.023011)

# Enhancing Dark Matter Annihilation Rates with Dark Bremsstrahlung

Nicole F. Bell,<sup>1,\*</sup> Yi Cai,<sup>1,†</sup> James B. Dent,<sup>2,‡</sup> Rebecca K. Leane,<sup>1,§</sup> and Thomas J. Weiler<sup>3,¶</sup>

<sup>1</sup>*ARC Centre of Excellence for Particle Physics at the Terascale  
School of Physics, The University of Melbourne, Victoria 3010, Australia*

<sup>2</sup>*Department of Physics, University of Louisiana at Lafayette, Lafayette, LA 70504-4210, USA*

<sup>3</sup>*Department of Physics and Astronomy, Vanderbilt University, Nashville, TN 37235, USA*

(Dated: July 10, 2017)

Many dark matter interaction types lead to annihilation processes which suffer from  $p$ -wave suppression or helicity suppression, rendering them sub-dominant to unsuppressed  $s$ -wave processes. We demonstrate that the natural inclusion of dark initial state radiation can open an unsuppressed  $s$ -wave annihilation channel, and thus provide the dominant dark matter annihilation process for particular interaction types. We illustrate this effect with the bremsstrahlung of a dark spin-0 or dark spin-1 particle from fermionic dark matter,  $\bar{\chi}\chi \rightarrow \bar{f}f\phi$  or  $\bar{f}fZ'$ . The dark initial state radiation process, despite having a 3-body final state, proceeds at the same order in the new physics scale  $\Lambda$  as the annihilation to the 2-body final state  $\bar{\chi}\chi \rightarrow \bar{f}f$ . This opens an unsuppressed  $s$ -wave at lower order in  $\Lambda$  than the well-studied lifting of helicity suppression via Standard Model final state radiation, or virtual internal bremsstrahlung. This dark bremsstrahlung process should influence LHC and indirect detection searches for dark matter.

## I. INTRODUCTION

The particle nature of dark matter (DM) remains unknown. In order to significantly probe its properties in indirect detection experiments, large or unsuppressed annihilation rates are desirable. The DM annihilation rate will generally be largest if it proceeds via an unsuppressed  $s$ -wave process. Unfortunately, there are a number of well motivated DM models in which the  $s$ -wave annihilation to Standard Model (SM) products,  $\bar{f}f$ , is absent or helicity suppressed. This renders indirect detection very unlikely, as the  $p$ -wave term is suppressed by a factor of the DM velocity squared (with  $v^2 \sim 10^{-6}$  in the present universe) while a helicity suppression factor of  $(m_f/m_\chi)^2$  can be significant for annihilation to light fermions. These suppressions are well-known features of neutralino annihilation in supersymmetric theories, but in fact are more general.

It is well known that such suppressions can be lifted via the bremsstrahlung of a SM particle. For example, an unsuppressed  $s$ -wave can be opened via the radiation of a photon [1–6] or electroweak gauge boson [7–14] during the DM annihilation processes. This has led to much recent work on the importance of SM radiative corrections in dark matter annihilation [1–33]. Despite the bremsstrahlung annihilation process having a 3-body final state, it can be the dominant annihilation channel in the universe today (if not at freeze out) because the suppression from additional coupling and phase space factors is small compared to the  $v^2 \sim 10^{-6}$  suppression of the  $p$ -wave contributions. Past work has primarily used final

state radiation (FSR) or virtual internal bremsstrahlung (VIB) to lift the suppression. If the DM is a SM gauge singlet, initial state radiation (ISR) of a SM particle is obviously not possible, however ISR of a  $W$  or  $Z$  boson from  $SU(2)$  charged DM is possible, and has been considered in [34–36].

An interesting possibility is that helicity or  $p$ -wave suppressions can instead be lifted by the ISR of a *dark sector* field. In this scenario, an initial state dark bremsstrahlung process can dominate over other suppressed channels. This will require that the dark sector contains more particles than just the DM candidate itself which, in fact, is very well motivated: the visible sector itself comprises more than one particle species, and likewise multiple dark sector fields are a common feature of many self-consistent, gauge-invariant, and renormalizable models. For example, mass generation in the dark sector can require the introduction of new fields, such as a dark Higgs, while DM stability may arise from a charge under a new dark sector gauge group, requiring the introduction of dark photons. More generally, models in which DM interactions are mediated by the exchange of only an axial-vector mediator are not gauge invariant. They require the addition of a dark Higgs to unitarize the longitudinal component of the gauge boson, and to give mass to both the gauge boson and DM [37–42]. Indeed, the simultaneous presence of both spin-1 and spin-0 mediators lead to new indirect detection phenomenology that does not arise in single mediator models [39–41]. Similarly, both scalar and pseudoscalar mediators can naturally appear together in complete theories [43–46].

In this paper, for the first time, we explore the possibility that helicity or  $p$ -wave suppressions of the DM annihilation process are lifted by dark bremsstrahlung from the initial state. We investigate the case where fermionic DM,  $\chi$ , radiates either a dark spin-1 field,  $Z'$ , or spin-0 field,  $\phi$ , to give the ISR processes  $\bar{\chi}\chi \rightarrow \bar{f}fZ'$  or  $\bar{\chi}\chi \rightarrow \bar{f}f\phi$ , respectively, as shown in Figure 1.

---

\* n.bell@unimelb.edu.au

† yi.cai@unimelb.edu.au

‡ jbdent@louisiana.edu

§ rebecca.lean@gmail.com

¶ tom.weiler@vanderbilt.edu

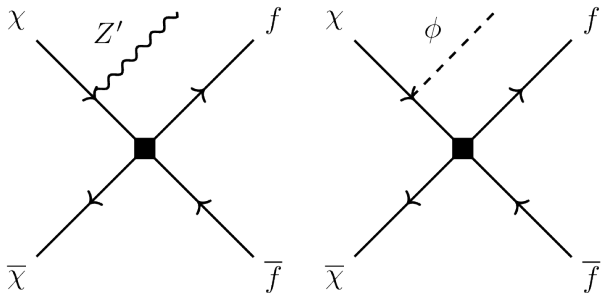


FIG. 1. Dark vector (left) and dark scalar (right) ISR. Note in both cases that there is also an additional diagram for emission from the other initial state DM.

Bremsstrahlung annihilation processes are very closely related to the mono- $X$  processes utilized in collider DM searches [47–79], as they are controlled by the same matrix element. For example, the radiation of photons from fermions in the FSR annihilation process  $\bar{\chi}\chi \rightarrow \bar{f}f\gamma$  is the analogue of the collider ISR mono-photon process  $\bar{f}f \rightarrow \bar{\chi}\chi\gamma$ . Likewise, the ISR of a dark spin-0 or spin-1 field from the initial state  $\chi$  in the  $\bar{\chi}\chi \rightarrow \bar{f}f\phi$  or  $\bar{\chi}\chi \rightarrow \bar{f}fZ'$  annihilation processes are then the analogue of the FSR mono- $Z'$  [80–84] or mono-dark Higgs [42] collider processes, respectively.

For the purpose of illustration, we shall assume the  $\bar{\chi}\chi \rightarrow \bar{f}f$  process is adequately described by an effective field theory (EFT operator) of the form  $(1/\Lambda^2)(\bar{\chi}\Gamma\chi)(\bar{f}\Gamma f)$ . We will see that the  $s$ -wave contribution to the ISR process scales as  $\langle\sigma v\rangle_{\text{ISR}} \propto \mathcal{O}(1/\Lambda^4)$ , i.e., the same order in  $\Lambda$  as the 2-body annihilation  $\bar{\chi}\chi \rightarrow \bar{f}f$ . In comparison, the well-studied lifting of helicity suppressions via FSR or VIB radiation can only produce unsuppressed  $s$ -wave cross sections at higher order in  $1/\Lambda$ , with cross sections scaling as  $\langle\sigma v\rangle_{\text{FSR, ISR}} \propto \mathcal{O}(1/\Lambda^8)$ <sup>1</sup>.

In Section II, we provide an overview of suppressions to fermionic DM annihilation cross sections, and discuss annihilation both directly to SM particles, and to dark mediators. In Section III we outline possible dark ISR annihilation processes, and investigate two interesting cases in more detail in Sections V and IV. We present our conclusions in Section VI.

## II. OVERVIEW OF FERMIONIC DARK MATTER ANNIHILATION

### A. Direct annihilation to SM particles

If DM is a Majorana fermion, the possible interactions which can mediate a  $\chi\chi \rightarrow \bar{f}f$  annihilation process are:

- $s$ -channel exchange of an axial-vector: helicity suppressed  $s$ -wave,
- $s$ -channel exchange of a scalar: no  $s$ -wave,
- $s$ -channel exchange of a pseudoscalar: unsuppressed  $s$ -wave, or
- $t$ -channel exchange of a sfermion-like scalar: helicity suppressed  $s$ -wave.

In the  $t$ -channel case, Fierz rearrangement to  $s$ -channel form gives  $A \otimes A$  and  $A \otimes V$  structures. The  $A \otimes A$  has a helicity suppressed  $s$ -wave, while the  $A \otimes V$  has no  $s$ -wave. For Majorana DM, we thus see that the  $s$ -channel exchange of a pseudoscalar is the only case of an unsuppressed  $s$ -wave. All other possibilities feature either helicity or  $v^2$  suppressions.

For Dirac DM, there are additional possibilities because vector couplings, which vanish for Majorana particles, are also allowed. Note, however, that while the exchange of a vector results in an unsuppressed  $s$ -wave annihilation cross section, these models are also well constrained because they lead to unsuppressed spin-independent scattering in direct detection experiments.

A summary of the cross section suppression factors, for both annihilation and scattering, for all possible Lorentz structures for  $\bar{\chi}\chi \rightarrow \bar{f}f$ , is given in Ref. [85].

### B. Direct annihilation to dark mediators

Table I details whether fermionic DM annihilation to two different mediators ( $\bar{\chi}\chi \rightarrow M_1M_2$ ) is  $s$ - or  $p$ -wave, depending on the Lorentz structures of the DM-mediator interactions. For annihilation to any two spin-1 mediators, the rate is  $s$ -wave. For any two spin-0 mediators, the rate is  $p$ -wave unless one scalar and one pseudoscalar are both present. For a mixed spin-0 and spin-1 final state, if the spin-1 is a vector, the processes are  $s$ -wave, while if the spin-1 is an axial-vector, the processes are  $p$ -wave<sup>2</sup>.

If one of these mediators is off-shell while the other is on-shell, it is equivalent to the dark ISR process discussed in the following section — where the on-shell mediator corresponds to the dark ISR, and the off-shell mediator has been integrated out to give the EFT vertex. As such, the annihilation type for dark ISR is related to the underlying Lorentz structures of the mediators. We now discuss the dark ISR processes in detail.

<sup>1</sup> This observation was also made in the case of ISR of a  $W/Z$  boson from  $SU(2)$  doublet DM [34–36].

<sup>2</sup> If there exist direct couplings between the dark mediators themselves, it is possible to have an  $s$ -wave process with a spin-0 plus spin-1 final state where the spin-1 is an axial-vector [39, 41]. Details of such couplings are model dependent, and we do not consider such processes in this work.

$\Gamma_{M_1} \otimes \Gamma_{M_2}$	$S \otimes S$	$S \otimes P$	$P \otimes P$	$V \otimes V$	$V \otimes A$	$A \otimes A$	$S \otimes V$	$S \otimes A$	$P \otimes V$	$P \otimes A$
$\bar{\chi}\chi \rightarrow M_1 M_2$	$v^2$	1	$v^2$	1	1	1	1	$v^2$	1	$v^2$

TABLE I. Suppression factors for fermionic DM annihilation to two different mediators  $M_1$  and  $M_2$ , which have varying Lorentz structures: vector (V), axial-vector (A), scalar (S) or pseudoscalar (P). The combination of mediators can be two spin-0 final states, two spin-1 final states, or a mixed spin-0 plus spin-1 final state. Note for Majorana DM, the  $V$  cases do not exist.

$\Gamma_\chi \otimes \Gamma_f$	$\bar{\chi}\chi \rightarrow \bar{f}f$	$\bar{\chi}\chi \rightarrow \bar{f}fZ'$		$\bar{\chi}\chi \rightarrow \bar{f}f\phi$	
		$\Gamma_{Z'} = V$	$\Gamma_{Z'} = A$	$\Gamma_\phi = S$	$\Gamma_\phi = P$
$V \otimes V$	1	1	1	1	1
$A \otimes V$	$v^2$	1	1	$v^2$	$v^2$
$V \otimes A$	1	1	1	1	1
$A \otimes A$	$(m_f/m_\chi)^2$	1	1	$v^2$	$v^2$
$S \otimes S$	$v^2$	1	$v^2$	$v^2$	1
$P \otimes S$	1	1	$v^2$	1*	$v^2$
$S \otimes P$	$v^2$	1	$v^2$	$v^2$	1*
$P \otimes P$	1	1	$v^2$	1	$v^2$

TABLE II. Suppression factors for fermionic DM annihilation processes, for varying Lorentz structures: vector (V), axial-vector (A), scalar (S) or pseudoscalar (P). In some cases, helicity suppression ( $\propto m_f^2/m_\chi^2$ ) or  $p$ -wave suppression ( $\propto v^2$ ) can be lifted to  $s$ -wave ( $\propto 1$ ) by including dark bremsstrahlung, with no relative suppression factors. Note that the presence of any vector or axial-vector ISR implies the existence of the competing  $s$ -wave process  $\bar{\chi}\chi \rightarrow Z'Z'$ . An asterisk \* indicates that the  $s$ -wave process  $\bar{\chi}\chi \rightarrow \phi_1\phi_2$  will inevitably be induced via scalar( $\phi_1$ )-pseudoscalar( $\phi_2$ ) mixing in a CP violating scenario. Note for Majorana DM, the  $V$  cases do not exist.

### III. DARK INITIAL STATE RADIATION

In this section, we consider the scenario where the ISR of a dark sector particle lifts helicity or  $p$ -wave suppression in fermionic DM annihilation processes.

Figure 1 demonstrates the dark sector ISR in DM annihilation. For the sake of illustration, we use an EFT to describe the interactions between DM and SM fermions. The qualitative effects we discuss are relevant for UV completions which map to the relevant cases. We assume one mediator is sufficiently heavy, such that the EFT description can safely be used without unitarity issues.

Table II details the annihilation type and relative suppression of all processes (whether they are  $s$ -wave,  $p$ -wave, or helicity suppressed). This reveals which Lorentz structures for particular dark ISR will lift suppression in DM annihilation.

We see, for example, that the radiation of a dark vector is a promising ISR scenario which lifts the suppression of DM annihilation for several Lorentz structures:  $S \otimes S$ ,  $S \otimes P$ ,  $A \otimes A$ , and  $A \otimes V$ . Radiating an axial-vector

lifts suppression in  $A \otimes A$  and  $A \otimes V$  annihilation processes. Radiating a scalar fails to lift any suppression of the annihilation cross section. Radiating a pseudoscalar, however, makes a process with a  $S \otimes S$  or  $S \otimes P$  structure  $s$ -wave. In the case of  $S \otimes P$  or  $P \otimes S$ , any scalar with such a structure will not have well-defined CP properties. Thus the mixing between the heavy scalar and the pseudoscalar is inevitable, and a  $2 \rightarrow 2$   $s$ -wave contribution,  $\bar{\chi}\chi \rightarrow \phi_1\phi_2$  will be induced, where  $\phi_1$  is a scalar and  $\phi_2$  is a pseudoscalar.

It is also important to note that once an additional dark sector field is included to allow dark ISR, there can also be  $s$ -wave annihilations of DM into the dark radiation. For spin-1 ISR, the direct annihilation to mediators  $\bar{\chi}\chi \rightarrow Z'Z'$  is  $s$ -wave for both vector and axial-vector couplings, and can dominate the total DM annihilation rate for some choices of the coupling strength or masses. In the case that the dark radiation is a spin-0 field, the  $t$ -channel annihilation process  $\bar{\chi}\chi \rightarrow \phi\phi$  is  $p$ -wave suppressed for both scalar and pseudoscalar couplings, and so can very naturally be sub-dominant to the

suppression-lifting ISR process. Also note that for sufficiently light dark radiation, Sommerfeld effects can be important.

To avoid the “dark radiation” contributing to the relic density, it must eventually decay to SM states. This can easily be arranged without introducing other consequences, e.g., via a gauge or Higgs portal to the SM, which can naturally appear for inclusion of a gauge boson or scalar, respectively. Such SM states can be a signal for indirect detection experiments (see, i.e., Refs. [39, 86]). We will assume that the couplings of the dark radiation to the SM are small, such that the  $2 \rightarrow 2$  exchange of dark radiation will be subdominant.

We now study in detail two particular cases of the lifting helicity or  $p$ -wave suppression:  $S \otimes S$  with dark pseudoscalar ISR and  $A \otimes A$  with dark vector ISR. We choose the former as it is the only scenario where introducing dark ISR to lift a  $p$ -wave cross section does not induce an additional competing  $2 \rightarrow 2$   $s$ -wave process. We choose the latter as an example of lifting helicity suppression. Other scenarios and Lorentz structures can dominate in particular regions of parameter space. Note also that in all the scenarios discussed, UV completions with the same Lorentz structures would map to the same results we present. Our results are not specific to EFTs, but rather to the underlying Lorentz structures.

#### IV. LIFTING P-WAVE SUPPRESSION IN $S \otimes S$ INTERACTIONS

In this section, we demonstrate how  $p$ -wave suppression can be lifted through dark pseudoscalar ISR, in the case of the Lorentz structure  $\Gamma_\chi \otimes \Gamma_f = S \otimes S$ . Such a structure is possible for both Majorana and Dirac DM, with the Majorana interaction terms differing by a factor of 1/2. We will also discuss any new competing annihilation processes.

##### A. $p$ -wave suppressed $\bar{\chi}\chi \rightarrow \bar{f}f$

For the Lorentz structure  $\Gamma_\chi \otimes \Gamma_f = S \otimes S$ , the DM interactions with SM fermions are described by the four-Fermi operator

$$\mathcal{L}_{\text{int}} \supset \frac{1}{\Lambda^2} (\bar{\chi}\chi)(\bar{f}f), \quad (1)$$

where  $\chi$  is a Dirac DM candidate,  $f$  are SM fermions and  $\Lambda$  is the cutoff scale for new physics, representing a heavy field which has been integrated out.

The operator in Eq. (1) yields a  $p$ -wave suppressed DM annihilation cross section for  $\bar{\chi}\chi \rightarrow \bar{f}f$  given by

$$\sigma v = \frac{v^2 m_\chi^2 \left(1 - m_f^2/m_\chi^2\right)^{3/2}}{8\pi\Lambda^4}. \quad (2)$$

where  $m_f$  is the mass of the SM fermion and  $m_\chi$  is the DM mass.  $v$  here denotes relative velocity. The  $v^2$  prefactor shows that this is clearly a  $p$ -wave suppressed process. We now explicitly show that, for such an operator, including dark pseudoscalar radiation lifts this suppression, and the dominant  $s$ -wave process can be  $\bar{\chi}\chi \rightarrow \bar{f}f\phi$ .

##### B. Dark pseudoscalar ISR, $\bar{\chi}\chi \rightarrow \bar{f}f\phi$

We consider a minimal setup in which the EFT operator of Eq. (1) is augmented by a coupling of the DM to a new pseudoscalar  $\phi$ ,

$$\mathcal{L}_{\text{int}} \supset \frac{1}{\Lambda^2} (\bar{\chi}\chi)(\bar{f}f) + i g_\phi \phi \bar{\chi}\gamma_5\chi, \quad (3)$$

where  $g_\phi$  is the coupling constant. While some complete model may have relations between the couplings and masses of the dark sector particles, for the sake of illustration we take all masses and couplings to be independent parameters. The annihilation cross section for the dark pseudoscalar ISR process  $\bar{\chi}\chi \rightarrow \bar{f}f\phi$  is given (in the  $m_f = 0$  limit) by

$$\langle\sigma v\rangle_{\chi\bar{\chi}\rightarrow\bar{f}f\phi} = \frac{g_\phi^2 m_\chi^2}{48\pi^3 \Lambda^4} \times \left\{ 1 + 24\rho_\phi^3 \sqrt{1 - \rho_\phi^2} (5\rho_\phi^2 - 2) \tan^{-1} \frac{\sqrt{1 - \rho_\phi^2}}{\rho_\phi} + 21\rho_\phi^2 - 105\rho_\phi^4 + 83\rho_\phi^6 + 12\rho_\phi^2 (1 - 9\rho_\phi^2 + 10\rho_\phi^4) \ln \rho_\phi \right\}, \quad (4)$$

where  $\rho_\phi = m_\phi/2m_\chi$ . Clearly this process is no longer velocity suppressed term, and so is  $s$ -wave. Furthermore, it scales  $\propto \mathcal{O}(1/\Lambda^4)$ , i.e., the same order in  $\Lambda$  as the annihilation to  $\bar{f}f$  which scales as  $\propto \mathcal{O}(v^2/\Lambda^4)$ . In contrast, FSR or VIB of a SM particle, e.g.  $\bar{\chi}\chi \rightarrow \bar{f}f\gamma$ , only allows unsuppressed  $s$ -wave annihilation at higher order in  $1/\Lambda$ , with a cross section scaling  $\propto \mathcal{O}(1/\Lambda^8)$ .

##### C. Competition with $\bar{\chi}\chi \rightarrow \phi\phi\phi$

Unlike the dark vector radiation case which will be discussed in the next section, there is no additional  $s$ -wave  $2 \rightarrow 2$  process induced for spin-0 fields because the

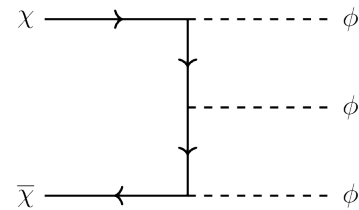


FIG. 2.  $s$ -wave process for DM annihilation to pseudoscalars,  $\bar{\chi}\chi \rightarrow \phi\phi\phi$ . Note that there are a total of six diagrams that contribute.

$\phi\phi$  final state is  $p$ -wave suppressed. In the limit  $m_\phi \ll m_\chi$ , the  $p$ -wave suppressed cross section for  $\bar{\chi}\chi \rightarrow \phi\phi$  is

$$\sigma v_{\bar{\chi}\chi \rightarrow \phi\phi} \simeq \frac{g_\phi^4 v^2}{384\pi m_\chi^2}. \quad (5)$$

Instead, the  $s$ -wave annihilation to 3 pseudoscalars  $\phi\phi\phi$  [87] shown in Fig. 2 will compete with the ISR  $\bar{f}f\phi$  channel. Note that both these  $2 \rightarrow 3$  process suffer the same 3-body phase space suppression. In the limit  $m_\phi \ll m_\chi$ , the  $s$ -wave cross section for  $\bar{\chi}\chi \rightarrow \phi\phi\phi$  is

$$\langle \sigma v \rangle_{\bar{\chi}\chi \rightarrow \phi\phi\phi} \simeq \frac{g_\phi^6 (7\pi^2 - 60)}{1536\pi^3 m_\chi^2}. \quad (6)$$

The full cross section is described in Appendix A.

Figure 3 displays the annihilation cross sections for the pseudoscalar ISR and  $\phi\phi\phi$  annihilation processes, illustrating potential regions of parameter space where either process is the dominant annihilation channel, depending on values of the couplings, DM and pseudoscalar masses. Other than providing kinematic thresholds for when the annihilations are allowed, the rates are effectively independent of the pseudoscalar mass. The pseudoscalar ISR process dominates over the three-body pseudoscalar process for much of the parameter space.

## V. LIFTING HELICITY SUPPRESSION IN $A \otimes A$ INTERACTIONS

In this section, we demonstrate how helicity suppression can be lifted through dark vector ISR, in the case of the Lorentz structure  $\Gamma_\chi \otimes \Gamma_f = A \otimes A$ . Such a structure is very natural for Majorana DM, but is also possible with Dirac DM. We will also discuss any new competing annihilation processes.

### A. Helicity suppressed $\bar{\chi}\chi \rightarrow \bar{f}f$

For the Lorentz structure  $\Gamma_\chi \otimes \Gamma_f = A \otimes A$ , the DM interactions with SM fermions are described by the four-Fermi operator

$$\mathcal{L}_{\text{int}} \supset \frac{1}{\Lambda^2} (\bar{\chi} \gamma^\mu \gamma^5 \chi) (\bar{f} \gamma^\mu \gamma^5 f), \quad (7)$$

where we shall again assume that the DM candidate  $\chi$  is a Dirac fermion.

The operator in Eq. (7) yields a helicity suppressed DM annihilation cross section for  $\bar{\chi}\chi \rightarrow \bar{f}f$ ,

$$\langle \sigma v \rangle_{\bar{\chi}\chi \rightarrow \bar{f}f} = \frac{m_f^2 \sqrt{1 - m_f^2/m_\chi^2}}{2\pi\Lambda^4}. \quad (8)$$

In the limit that  $m_f \rightarrow 0$ , this process is vanishing. We now explicitly show that including dark vector radiation of a  $Z'$  lifts this helicity suppression, such that the dominant  $s$ -wave process can be  $\bar{\chi}\chi \rightarrow \bar{f}fZ'$ .

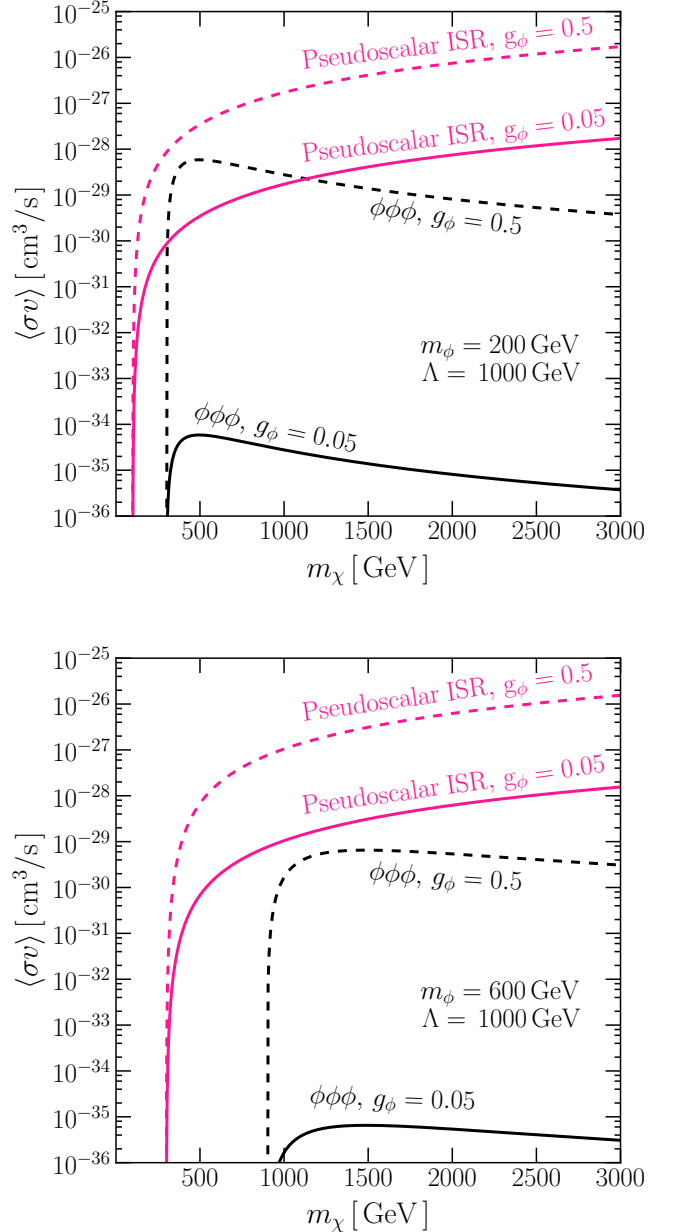


FIG. 3. Comparison of the  $s$ -wave cross sections for the pseudoscalar ISR (pink) and 3-body  $\phi\phi\phi$  (black) annihilation processes for  $m_\phi$ ,  $\Lambda$  and  $g_\phi$  as labeled, and  $m_f = 0$ . Note the largest value shown for  $m_\chi$  corresponds to  $m_\chi \sim \sqrt{4\pi}\Lambda$ , where the validity of the EFT within the perturbative regime is satisfied.

### B. Dark vector ISR, $\bar{\chi}\chi \rightarrow \bar{f}fZ'$

We again consider a minimal scenario where, in addition to the EFT interaction of Eq. (7), we include a coupling of the DM to a new spin-1 field,  $Z'$ ,

$$\mathcal{L}_{\text{int}} \supset \frac{1}{\Lambda^2} (\bar{\chi} \gamma_\mu \gamma_5 \chi) (\bar{f} \gamma^\mu \gamma_5 f) + g_{Z'} \bar{\chi} \gamma^\mu \chi Z'_\mu, \quad (9)$$



where  $g_{Z'}$  is the coupling constant. For the sake of illustration, we have chosen a  $Z'$  coupling of vector form to Dirac DM. Alternatively, an axial-vector  $Z'$  could be chosen (for either Dirac and Majorana DM) which would also open an unsuppressed  $s$ -wave, as shown in Table II.

The annihilation cross section for the dark vector ISR process  $\bar{\chi}\chi \rightarrow \bar{f}fZ'$  is given (in the  $m_f = 0$  limit) by

$$\langle\sigma v\rangle_{\bar{\chi}\chi\rightarrow\bar{f}fZ'} = \frac{g_{Z'}^2 m_\chi^2}{36\pi^3 \Lambda^4} \times \quad (10)$$

$$\left\{ 4 + 24\rho_{Z'}^3 (1 + 5\rho_{Z'}^2) \sqrt{1 - \rho_{Z'}^2} \tan^{-1} \frac{\sqrt{1 - \rho_{Z'}^2}}{\rho_{Z'}} \right.$$

$$\left. - 27\rho_{Z'}^2 - 60\rho_{Z'}^4 + 83\rho_{Z'}^6 + 12\rho_{Z'}^4 (10\rho_{Z'}^2 - 3) \ln \rho_{Z'} \right\},$$

where  $\rho_{Z'} = m_{Z'}/2m_\chi$ . As the cross section does not vanish in the limit  $m_f \rightarrow 0$ , this process no longer has the helicity suppressed ( $m_f^2/m_\chi^2$ ) dependence. It also has no velocity suppression, and so is  $s$ -wave. Again, we see the dark ISR cross section scales  $\propto \mathcal{O}(1/\Lambda^4)$ , i.e., the same order in  $\Lambda$  as the annihilation to  $\bar{f}f$  which scales  $\propto \mathcal{O}(m_f^2/(m_\chi^2 \Lambda^4))$ .

### C. Competition with $\bar{\chi}\chi \rightarrow Z'Z'$

The inclusion of the  $Z'$  vector induces an additional two-body annihilation process,  $\bar{\chi}\chi \rightarrow Z'Z'$ , as shown in Fig. 4. This process is also  $s$ -wave (irrespective of whether the  $Z'$  couplings are of vector or axial-vector form). Therefore, the dominant annihilation channel will be  $\bar{\chi}\chi \rightarrow Z'Z'$  or  $\bar{\chi}\chi \rightarrow \bar{f}fZ'$ , either of which may dominate depending on the region of parameter space. The cross section for the annihilation of Dirac DM to a pair of  $Z'$  is given by

$$\langle\sigma v\rangle_{\bar{\chi}\chi\rightarrow Z'Z'} = \frac{g_{Z'}^4 (1 - 4\rho_{Z'}^2)^{\frac{3}{2}}}{16\pi m_\chi^2 (1 - 2\rho_{Z'}^2)^2}, \quad (11)$$

where again  $\rho_{Z'} = m_{Z'}/2m_\chi$ .

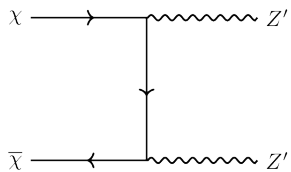


FIG. 4.  $s$ -wave process for DM annihilation to dark vectors. Note there is also a contribution from the  $u$ -channel diagram.

Figure 5 displays the annihilation cross sections for Dirac DM for both processes, illustrating regions of parameter space where either process may be the dominant annihilation channel. As the annihilation cross section to  $\bar{f}fZ'$  contains two powers of  $g_{Z'}$ , while that for  $Z'Z'$  contains four powers of this coupling, the dark ISR process

can easily dominate when  $g_{Z'}$  is small. As the coupling becomes larger,  $Z'Z'$  dominates for more of the parameter space. If the  $Z'$  is particularly heavy, the dark ISR process can also dominate over  $Z'Z'$  due to kinematic constraints.

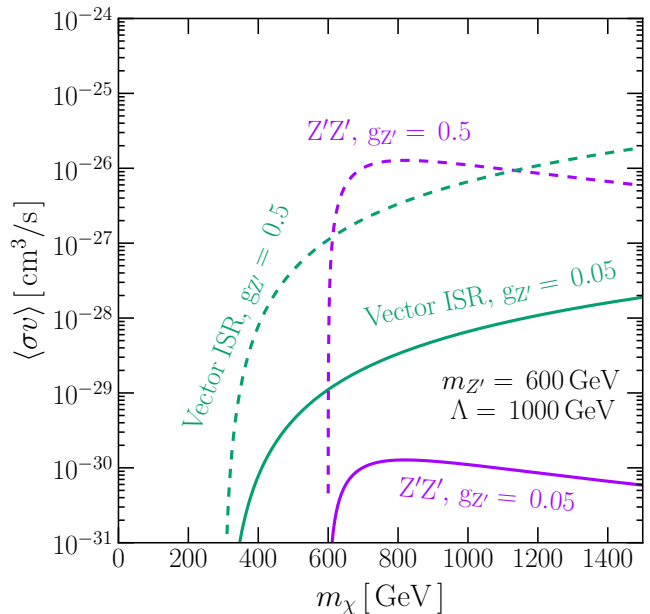
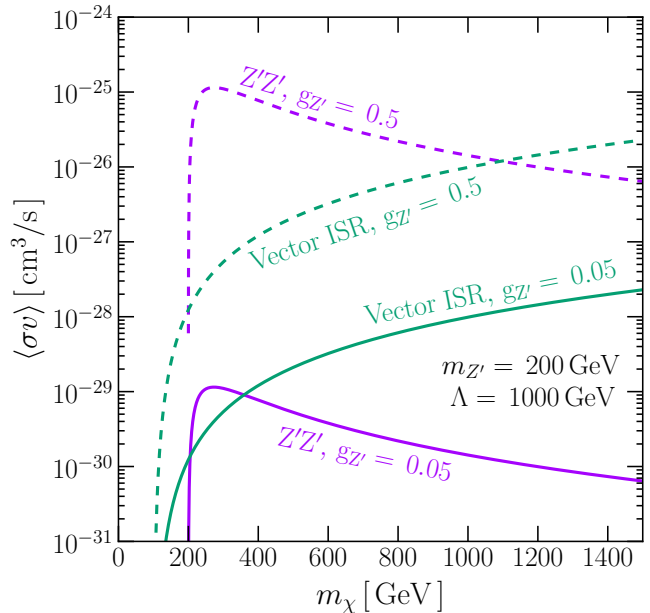


FIG. 5. Comparison of the  $s$ -wave cross sections for Dirac DM annihilating via the dark vector ISR process (cyan) and the 2-body  $Z'Z'$  process (purple), for  $m_{Z'}$ ,  $\Lambda$  and  $g_{Z'}$  as labeled, and  $m_f = 0$ . Note the largest value shown for  $m_\chi$  corresponds approximately to the largest value permitted within a gauge-invariant and perturbative framework, without other new physics appearing.

## VI. CONCLUSION

The observation of an unexplained flux of SM particles in the astrophysical sky can be interpreted as a DM signal. To probe the nature of DM via such an indirect detection signal, it is important to know which processes may provide the dominant contributions or strongest constraints. In this paper, for the first time, we have explored the possibility that dark ISR can open an unsuppressed  $s$ -wave annihilation channel. This can be the dominant DM annihilation mode in models where the lowest order processes are helicity or  $p$ -wave suppressed.

We found that dark ISR from the initial state  $\bar{\chi}\chi$  can lift such suppressions for several different types of dark radiation and DM interaction structures. For four-Fermi type interactions of DM with SM fermions, several Lorentz structures suffer suppressed  $2 \rightarrow 2$  annihilation processes —  $A \otimes A$ ,  $A \otimes V$ ,  $S \otimes S$ , and  $S \otimes P$ . The ISR of a dark vector opens an unsuppressed  $s$ -wave in all these cases. Radiating an axial-vector lifts the suppression of the  $A \otimes A$  and  $A \otimes V$  annihilation processes, while radiating a pseudoscalar opens an  $s$ -wave for the  $S \otimes S$  or  $S \otimes P$  interaction types.

An important feature of dark ISR is that the bremsstrahlung annihilation rate scales as  $\langle\sigma v\rangle_{\text{ISR}} \propto \mathcal{O}(1/\Lambda^4)$ , i.e., the same order in  $\Lambda$  as the 2-body annihilation  $\bar{\chi}\chi \rightarrow \bar{f}f$ . In comparison, the opening of an unsuppressed  $s$ -wave via FSR or VIB of a SM particle (e.g. as in the well studied  $\bar{\chi}\chi \rightarrow \bar{f}f\gamma$  process) occurs only at higher order in  $1/\Lambda$ , with a cross section scaling as  $\langle\sigma v\rangle_{\text{FSR, ISR}} \propto \mathcal{O}(1/\Lambda^8)$ .

When introducing a new field for dark ISR, additional competing annihilation processes are induced. For the ISR of a dark vector or axial-vector, a competing  $s$ -wave annihilation process is  $\bar{\chi}\chi \rightarrow Z'Z'$ . For the case of scalar or pseudoscalar ISR, there is no equivalent  $2 \rightarrow 2$   $s$ -wave process. However, the annihilation to 3 pseudoscalars,  $\bar{\chi}\chi \rightarrow \phi\phi\phi$ , is  $s$ -wave and can dominate in some regions of parameter space. As such, the interplay of several annihilation processes must always be considered.

The introduction of dark radiation should not be viewed as an additional or unnecessary complication to dark sector theories. It is highly likely that the dark sector, like the visible sector, has multiple field content. Dark vectors arise naturally when the DM stability is due to a charge under a new gauge group, while dark scalars are well motivated when considering mass generation in the dark sector. Indeed, self-consistent,

gauge invariant models frequently contain such features. We have shown that it is important to include dark radiative corrections in these scenarios, as they can be the dominant annihilation channel.

## ACKNOWLEDGMENTS

RKL thanks the Niels Bohr International Academy for their hospitality during the completion of this work. NFB, YC and RKL were supported in part by the Australian Research Council. TJW is supported in part by the Department of Energy (DoE) Grant No. DESC-0011981. Feynman diagrams are drawn using TIKZ-FEYNMAN [88].

## APPENDIX

### Appendix A: Three pseudoscalar cross section

The cross section for DM annihilation into three pseudoscalars,  $\bar{\chi}\chi \rightarrow \phi\phi\phi$ , is given by

$$\langle\sigma v\rangle_{\bar{\chi}\chi \rightarrow \phi\phi\phi} = \frac{g_\phi^6}{2^5 3! \pi^3} \int_{\rho_\phi}^{\frac{1-\rho_\phi}{2}} dx_2 \int_{x_1^{\min}}^{x_1^{\max}} dx_1 \quad (\text{A1})$$

$$\times \frac{\left(1 - 4(x_1 + x_2) + 4x_1x_2 + 4(x_1^2 + x_2^2) + 2\rho_\phi^2 - 3\rho_\phi^4\right)^2}{8m_\chi^2(x_1 - \rho_\phi^2)^2(x_2 - \rho_\phi^2)^2(1 - x_1 - x_2 - \rho_\phi^2)^2},$$

with  $x_{1,2} = E_{1,2}/2m_\chi$  where  $E_{1,2}$  is the energy of two of the three pseudoscalars in the center of mass frame of the DM pair. The integrand is clearly symmetric with respect to  $x_1$  and  $x_2$ , as expected. The integration limits  $x_1^{\min, \max}$  are functions of  $x_2$  and  $\rho_\phi$  expressed as

$$x_1^{\min, \max} = \frac{1 + 2x_2^2 + \rho_\phi^2 - x_2(3 + \rho_\phi^2) \mp A}{2(1 - 2x_2 + \rho_\phi^2)}, \quad (\text{A2})$$

where

$$A = \left[4x_2^4 - 4x_2(x_2^2 - \rho_\phi^2)(1 - \rho_\phi^2) + x_2^2(1 - 6\rho_\phi^2 - 3\rho_\phi^4) - \rho_\phi^2(1 - 2\rho_\phi^2 - 3\rho_\phi^4)\right]^{1/2}.$$

In the limit  $m_\phi \ll m_\chi$ , this produces the  $s$ -wave cross section shown in Eq. (6).

- 
- [1] Lars Bergstrom, “Radiative Processes in Dark Matter Photino Annihilation,” *Phys. Lett.* **B225**, 372–380 (1989).  
 [2] Ricardo Flores, Keith A. Olive, and Serge Rudaz, “Radiative Processes in Lsp Annihilation,” *Phys. Lett.*

**B232**, 377–382 (1989).

- [3] E. A. Baltz and L. Bergstrom, “Detection of leptonic dark matter,” *Phys. Rev.* **D67**, 043516 (2003), arXiv:hep-ph/0211325 [hep-ph].



- [4] Torsten Bringmann, Lars Bergstrom, and Joakim Edsjo, “New Gamma-Ray Contributions to Supersymmetric Dark Matter Annihilation,” *JHEP* **01**, 049 (2008), [arXiv:0710.3169 \[hep-ph\]](#).
- [5] Lars Bergstrom, Torsten Bringmann, and Joakim Edsjo, “New Positron Spectral Features from Supersymmetric Dark Matter - a Way to Explain the PAMELA Data?” *Phys. Rev.* **D78**, 103520 (2008), [arXiv:0808.3725 \[astro-ph\]](#).
- [6] V. Barger, Y. Gao, Wai Yee Keung, and D. Marfatia, “Generic dark matter signature for gamma-ray telescopes,” *Phys. Rev.* **D80**, 063537 (2009), [arXiv:0906.3009 \[hep-ph\]](#).
- [7] Nicole F. Bell, James B. Dent, Thomas D. Jacques, and Thomas J. Weiler, “W/Z Bremsstrahlung as the Dominant Annihilation Channel for Dark Matter,” *Phys. Rev.* **D83**, 013001 (2011), [arXiv:1009.2584 \[hep-ph\]](#).
- [8] Nicole F. Bell, James B. Dent, Ahmad J. Galea, Thomas D. Jacques, Lawrence M. Krauss, and Thomas J. Weiler, “W/Z Bremsstrahlung as the Dominant Annihilation Channel for Dark Matter, Revisited,” *Phys. Lett.* **B706**, 6–12 (2011), [arXiv:1104.3823 \[hep-ph\]](#).
- [9] Nicole F. Bell, James B. Dent, Thomas D. Jacques, and Thomas J. Weiler, “Dark Matter Annihilation Signatures from Electroweak Bremsstrahlung,” *Phys. Rev.* **D84**, 103517 (2011), [arXiv:1101.3357 \[hep-ph\]](#).
- [10] Paolo Ciafaloni, Marco Cirelli, Denis Comelli, Andrea De Simone, Antonio Riotto, and Alfredo Urbano, “On the Importance of Electroweak Corrections for Majorana Dark Matter Indirect Detection,” *JCAP* **1106**, 018 (2011), [arXiv:1104.2996 \[hep-ph\]](#).
- [11] Mathias Garny, Alejandro Ibarra, and Stefan Vogl, “Antiproton constraints on dark matter annihilations from internal electroweak bremsstrahlung,” *JCAP* **1107**, 028 (2011), [arXiv:1105.5367 \[hep-ph\]](#).
- [12] Nicole F. Bell, Amelia J. Brennan, and Thomas D. Jacques, “Neutrino signals from electroweak bremsstrahlung in solar WIMP annihilation,” *JCAP* **1210**, 045 (2012), [arXiv:1206.2977 \[hep-ph\]](#).
- [13] Andrea De Simone, Alexander Monin, Andrea Thamm, and Alfredo Urbano, “On the effective operators for Dark Matter annihilations,” *JCAP* **1302**, 039 (2013), [arXiv:1301.1486 \[hep-ph\]](#).
- [14] Paolo Ciafaloni, Denis Comelli, Andrea De Simone, Enrico Morgante, Antonio Riotto, and Alfredo Urbano, “The Role of Electroweak Corrections for the Dark Matter Relic Abundance,” *JCAP* **1310**, 031 (2013), [arXiv:1305.6391 \[hep-ph\]](#).
- [15] V. Berezhinsky, M. Kachelriess, and S. Ostapchenko, “Electroweak jet cascading in the decay of super-heavy particles,” *Phys. Rev. Lett.* **89**, 171802 (2002), [arXiv:hep-ph/0205218 \[hep-ph\]](#).
- [16] M. Kachelriess, P. D. Serpico, and M. Aa. Solberg, “On the role of electroweak bremsstrahlung for indirect dark matter signatures,” *Phys. Rev.* **D80**, 123533 (2009), [arXiv:0911.0001 \[hep-ph\]](#).
- [17] M. Kachelriess and P. D. Serpico, “Model-independent dark matter annihilation bound from the diffuse  $\gamma$  ray flux,” *Phys. Rev.* **D76**, 063516 (2007), [arXiv:0707.0209 \[hep-ph\]](#).
- [18] Nicole F. Bell, James B. Dent, Thomas D. Jacques, and Thomas J. Weiler, “Electroweak Bremsstrahlung in Dark Matter Annihilation,” *Phys. Rev.* **D78**, 083540 (2008), [arXiv:0805.3423 \[hep-ph\]](#).
- [19] James B. Dent, Robert J. Scherrer, and Thomas J. Weiler, “Toward a Minimum Branching Fraction for Dark Matter Annihilation into Electromagnetic Final States,” *Phys. Rev.* **D78**, 063509 (2008), [arXiv:0806.0370 \[astro-ph\]](#).
- [20] Paolo Ciafaloni and Alfredo Urbano, “TeV scale Dark Matter and electroweak radiative corrections,” *Phys. Rev.* **D82**, 043512 (2010), [arXiv:1001.3950 \[hep-ph\]](#).
- [21] Paolo Ciafaloni, Denis Comelli, Antonio Riotto, Filippo Sala, Alessandro Strumia, and Alfredo Urbano, “Weak Corrections are Relevant for Dark Matter Indirect Detection,” *JCAP* **1103**, 019 (2011), [arXiv:1009.0224 \[hep-ph\]](#).
- [22] Vernon Barger, Wai-Yee Keung, and Danny Marfatia, “Bremsstrahlung in dark matter annihilation,” *Phys. Lett.* **B707**, 385–388 (2012), [arXiv:1111.4523 \[hep-ph\]](#).
- [23] Keita Fukushima, Yu Gao, Jason Kumar, and Danny Marfatia, “Bremsstrahlung signatures of dark matter annihilation in the Sun,” *Phys. Rev.* **D86**, 076014 (2012), [arXiv:1208.1010 \[hep-ph\]](#).
- [24] Federica Giacchino, Laura Lopez-Honorez, and Michel H. G. Tytgat, “Scalar Dark Matter Models with Significant Internal Bremsstrahlung,” *JCAP* **1310**, 025 (2013), [arXiv:1307.6480 \[hep-ph\]](#).
- [25] Takashi Toma, “Internal Bremsstrahlung Signature of Real Scalar Dark Matter and Consistency with Thermal Relic Density,” *Phys. Rev. Lett.* **111**, 091301 (2013), [arXiv:1307.6181 \[hep-ph\]](#).
- [26] Mathias Garny, Alejandro Ibarra, Miguel Pato, and Stefan Vogl, “Internal bremsstrahlung signatures in light of direct dark matter searches,” *JCAP* **1312**, 046 (2013), [arXiv:1306.6342 \[hep-ph\]](#).
- [27] Torsten Bringmann and Francesca Calore, “Significant Enhancement of Neutralino Dark Matter Annihilation from Electroweak Bremsstrahlung,” *Phys. Rev. Lett.* **112**, 071301 (2014), [arXiv:1308.1089 \[hep-ph\]](#).
- [28] Torsten Bringmann, Ahmad J. Galea, and Parampreet Walia, “Leading QCD Corrections for Indirect Dark Matter Searches: a Fresh Look,” *Phys. Rev.* **D93**, 043529 (2016), [arXiv:1510.02473 \[hep-ph\]](#).
- [29] Gulab Bambhaniya, Jason Kumar, Danny Marfatia, Alekha C. Nayak, and Gaurav Tomar, “Vector dark matter annihilation with internal bremsstrahlung,” *Phys. Lett.* **B766**, 177–180 (2017), [arXiv:1609.05369 \[hep-ph\]](#).
- [30] Jason Kumar, Pearl Sandick, Fei Teng, and Takahiro Yamamoto, “Gamma-ray Signals from Dark Matter Annihilation Via Charged Mediators,” *Phys. Rev.* **D94**, 015022 (2016), [arXiv:1605.03224 \[hep-ph\]](#).
- [31] Jason Kumar, Jiajun Liao, and Danny Marfatia, “Dark matter annihilation with s-channel internal Higgsstrahlung,” *Phys. Lett.* **B759**, 277–281 (2016), [arXiv:1605.00611 \[hep-ph\]](#).
- [32] Feng Luo and Tevong You, “Enhancement of Majorana Dark Matter Annihilation Through Higgs Bremsstrahlung,” *JCAP* **1312**, 024 (2013), [arXiv:1310.5129 \[hep-ph\]](#).
- [33] Xue-lei Chen and Marc Kamionkowski, “Three body annihilation of neutralinos below two-body thresholds,” *JHEP* **07**, 001 (1998), [arXiv:hep-ph/9805383 \[hep-ph\]](#).
- [34] Paolo Ciafaloni, Marco Cirelli, Denis Comelli, Andrea De Simone, Antonio Riotto, and Alfredo Urbano, “Initial State Radiation in Majorana Dark Matter Annihilations,” *JCAP* **1110**, 034 (2011), [arXiv:1107.4453 \[hep-ph\]](#).

- [35] Mathias Garny, Alejandro Ibarra, and Stefan Vogl, “Dark matter annihilations into two light fermions and one gauge boson: General analysis and antiproton constraints,” *JCAP* **1204**, 033 (2012), [arXiv:1112.5155 \[hep-ph\]](#).
- [36] Paolo Ciafaloni, Denis Comelli, Andrea De Simone, Antonio Riotto, and Alfredo Urbano, “Electroweak Bremsstrahlung for Wino-Like Dark Matter Annihilations,” *JCAP* **1206**, 016 (2012), [arXiv:1202.0692 \[hep-ph\]](#).
- [37] James M. Cline, Grace Dupuis, Zuowei Liu, and Wei Xue, “The windows for kinetically mixed  $Z'$ -mediated dark matter and the galactic center gamma ray excess,” *JHEP* **08**, 131 (2014), [arXiv:1405.7691 \[hep-ph\]](#).
- [38] Felix Kahlhoefer, Kai Schmidt-Hoberg, Thomas Schwetz, and Stefan Vogl, “Implications of unitarity and gauge invariance for simplified dark matter models,” *JHEP* **02**, 016 (2016), [arXiv:1510.02110 \[hep-ph\]](#).
- [39] Nicole F. Bell, Yi Cai, and Rebecca K. Leane, “Dark Forces in the Sky: Signals from  $Z'$  and the Dark Higgs,” *JCAP* **1608**, 001 (2016), [arXiv:1605.09382 \[hep-ph\]](#).
- [40] Michael Duerr, Felix Kahlhoefer, Kai Schmidt-Hoberg, Thomas Schwetz, and Stefan Vogl, “How to save the WIMP: global analysis of a dark matter model with two s-channel mediators,” *JHEP* **09**, 042 (2016), [arXiv:1606.07609 \[hep-ph\]](#).
- [41] Nicole F. Bell, Yi Cai, and Rebecca K. Leane, “Impact of mass generation for spin-1 mediator simplified models,” *JCAP* **1701**, 039 (2017), [arXiv:1610.03063 \[hep-ph\]](#).
- [42] Michael Duerr, Alexander Grohsjean, Felix Kahlhoefer, Bjoern Penning, Kai Schmidt-Hoberg, and Christian Schwanenberger, “Hunting the dark Higgs,” (2017), [arXiv:1701.08780 \[hep-ph\]](#).
- [43] Dorival Goncalves, Pedro A. N. Machado, and Jose Miguel No, “Simplified Models for Dark Matter Face their Consistent Completions,” *Phys. Rev.* **D95**, 055027 (2017), [arXiv:1611.04593 \[hep-ph\]](#).
- [44] Nicole F. Bell, Giorgio Busoni, and Isaac W. Sanderson, “Self-consistent Dark Matter Simplified Models with an s-channel scalar mediator,” *JCAP* **1703**, 015 (2017), [arXiv:1612.03475 \[hep-ph\]](#).
- [45] Martin Bauer, Ulrich Haisch, and Felix Kahlhoefer, “Simplified dark matter models with two Higgs doublets: I. Pseudoscalar mediators,” (2017), [arXiv:1701.07427 \[hep-ph\]](#).
- [46] Seungwon Baek, P. Ko, and Jinmian Li, “Minimal renormalizable simplified dark matter model with a pseudoscalar mediator,” *Phys. Rev.* **D95**, 075011 (2017), [arXiv:1701.04131 \[hep-ph\]](#).
- [47] Morad Aaboud *et al.* (ATLAS), “Search for new phenomena in final states with an energetic jet and large missing transverse momentum in  $pp$  collisions at  $\sqrt{s} = 13\text{TeV}$  using the ATLAS detector,” *Phys. Rev.* **D94**, 032005 (2016), [arXiv:1604.07773 \[hep-ex\]](#).
- [48] Morad Aaboud *et al.* (ATLAS), “Search for new phenomena in events with a photon and missing transverse momentum in  $pp$  collisions at  $\sqrt{s} = 13\text{TeV}$  with the ATLAS detector,” *JHEP* **06**, 059 (2016), [arXiv:1604.01306 \[hep-ex\]](#).
- [49] Morad Aaboud *et al.* (ATLAS), “Search for dark matter produced in association with a hadronically decaying vector boson in  $pp$  collisions at  $\sqrt{s}=13\text{TeV}$  with the ATLAS detector,” (2016), [arXiv:1608.02372 \[hep-ex\]](#).
- [50] “Search for Dark Matter production associated with bottom quarks with  $13.3\text{fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 13\text{TeV}$  with the ATLAS detector at the LHC,” (2016), aTLAS-CONF-2016-086.
- [51] “Search for Dark Matter in association with a Higgs boson decaying to  $b$ -quarks in  $pp$  collisions at  $\sqrt{s} = 13\text{TeV}$  with the ATLAS detector,” (2016), aTLAS-CONF-2016-019.
- [52] CMS Collaboration (CMS), “Search for new physics in a boosted hadronic monotop final state using  $12.9\text{fb}^{-1}$  of  $\sqrt{s} = 13\text{TeV}$  data,” (2016), cMS-PAS-EXO-16-040.
- [53] CMS Collaboration (CMS), “Search for dark matter and graviton produced in association with a photon in  $pp$  collisions at  $\sqrt{s} = 13\text{TeV}$ ,” (2016), cMS-PAS-EXO-16-039.
- [54] CMS Collaboration (CMS), “Search for dark matter in  $Z + E_{\text{T}}^{\text{miss}}$  events using  $12.9\text{fb}^{-1}$  of 2016 data,” (2016), cMS-PAS-EXO-16-038.
- [55] CMS Collaboration (CMS), “Search for dark matter in final states with an energetic jet, or a hadronically decaying  $W$  or  $Z$  boson using  $12.9\text{fb}^{-1}$  of data at  $\sqrt{s} = 13\text{TeV}$ ,” (2016), cMS-PAS-EXO-16-037.
- [56] CMS Collaboration (CMS), “Search for dark matter in association with a Higgs boson decaying into a pair of bottom quarks at  $\sqrt{s} = 13\text{TeV}$  with the CMS detector,” (2016), cMS-PAS-EXO-16-012.
- [57] CMS Collaboration (CMS), “Search for Dark Matter Produced in Association with a Higgs Boson Decaying to Two Photons,” (2016), cMS-PAS-EXO-16-011.
- [58] CMS Collaboration (CMS), “Search for dark matter in association with a top quark pair at  $\sqrt{s}=13\text{TeV}$ ,” (2016), cMS-PAS-EXO-16-005.
- [59] CMS Collaboration (CMS), “Search for Dark Matter produced in association with bottom quarks,” (2016), cMS-PAS-B2G-15-007.
- [60] Linda M. Carpenter, Andrew Nelson, Chase Shimmin, Tim M. P. Tait, and Daniel Whiteson, “Collider searches for dark matter in events with a  $Z$  boson and missing energy,” *Phys. Rev.* **D87**, 074005 (2013), [arXiv:1212.3352 \[hep-ex\]](#).
- [61] Nicole F. Bell, James B. Dent, Ahmad J. Galea, Thomas D. Jacques, Lawrence M. Krauss, and Thomas J. Weiler, “Searching for Dark Matter at the LHC with a Mono- $Z$ ,” *Phys. Rev.* **D86**, 096011 (2012), [arXiv:1209.0231 \[hep-ph\]](#).
- [62] Linda Carpenter, Anthony DiFranzo, Michael Mulhearn, Chase Shimmin, Sean Tulin, and Daniel Whiteson, “Mono-Higgs-boson: A new collider probe of dark matter,” *Phys. Rev.* **D89**, 075017 (2014), [arXiv:1312.2592 \[hep-ph\]](#).
- [63] Alexey A. Petrov and William Shepherd, “Searching for dark matter at LHC with Mono-Higgs production,” *Phys. Lett.* **B730**, 178–183 (2014), [arXiv:1311.1511 \[hep-ph\]](#).
- [64] Asher Berlin, Tongyan Lin, and Lian-Tao Wang, “Mono-Higgs Detection of Dark Matter at the LHC,” *JHEP* **06**, 078 (2014), [arXiv:1402.7074 \[hep-ph\]](#).
- [65] Nicole F. Bell, Yi Cai, and Rebecca K. Leane, “Mono- $W$  Dark Matter Signals at the LHC: Simplified Model Analysis,” *JCAP* **1601**, 051 (2016), [arXiv:1512.00476 \[hep-ph\]](#).
- [66] Nicole F. Bell, Yi Cai, James B. Dent, Rebecca K. Leane, and Thomas J. Weiler, “Dark matter at the LHC: Effective field theories and gauge invariance,” *Phys. Rev.*

- D92**, 053008 (2015), [arXiv:1503.07874 \[hep-ph\]](#).
- [67] Ulrich Haisch, Felix Kahlhoefer, and Tim M. P. Tait, “On Mono-W Signatures in Spin-1 Simplified Models,” (2016), [arXiv:1603.01267 \[hep-ph\]](#).
- [68] Andreas Birkedal, Konstantin Matchev, and Maxim Perelstein, “Dark matter at colliders: A Model independent approach,” *Phys. Rev.* **D70**, 077701 (2004), [arXiv:hep-ph/0403004 \[hep-ph\]](#).
- [69] Yuri Gershtein, Frank Petriello, Seth Quackenbush, and Kathryn M. Zurek, “Discovering hidden sectors with mono-photon  $Z'$  searches,” *Phys. Rev.* **D78**, 095002 (2008), [arXiv:0809.2849 \[hep-ph\]](#).
- [70] Jessica Goodman, Masahiro Ibe, Arvind Rajaraman, William Shepherd, Tim M. P. Tait, and Hai-Bo Yu, “Constraints on Dark Matter from Colliders,” *Phys. Rev.* **D82**, 116010 (2010), [arXiv:1008.1783 \[hep-ph\]](#).
- [71] Jessica Goodman, Masahiro Ibe, Arvind Rajaraman, William Shepherd, Tim M. P. Tait, and Hai-Bo Yu, “Constraints on Light Majorana dark Matter from Colliders,” *Phys. Lett.* **B695**, 185–188 (2011), [arXiv:1005.1286 \[hep-ph\]](#).
- [72] Andreas Crivellin, Ulrich Haisch, and Anthony Hibbs, “LHC constraints on gauge boson couplings to dark matter,” *Phys. Rev.* **D91**, 074028 (2015), [arXiv:1501.00907 \[hep-ph\]](#).
- [73] Frank J. Petriello, Seth Quackenbush, and Kathryn M. Zurek, “The Invisible  $Z'$  at the CERN LHC,” *Phys. Rev.* **D77**, 115020 (2008), [arXiv:0803.4005 \[hep-ph\]](#).
- [74] Tongyan Lin, Edward W. Kolb, and Lian-Tao Wang, “Probing dark matter couplings to top and bottom quarks at the LHC,” *Phys. Rev.* **D88**, 063510 (2013), [arXiv:1303.6638 \[hep-ph\]](#).
- [75] Patrick J. Fox, Roni Harnik, Joachim Kopp, and Yuhsin Tsai, “Missing Energy Signatures of Dark Matter at the LHC,” *Phys. Rev.* **D85**, 056011 (2012), [arXiv:1109.4398 \[hep-ph\]](#).
- [76] Jose Miguel No, “Looking through the pseudoscalar portal into dark matter: Novel mono-Higgs and mono-Z signatures at the LHC,” *Phys. Rev.* **D93**, 031701 (2016), [arXiv:1509.01110 \[hep-ph\]](#).
- [77] Karim Ghorbani and Leila Khalkhali, “Mono-Higgs signature in fermionic dark matter model,” (2016), [arXiv:1608.04559 \[hep-ph\]](#).
- [78] A. J. Brennan, M. F. McDonald, J. Gramling, and T. D. Jacques, “Collide and Conquer: Constraints on Simplified Dark Matter Models using Mono-X Collider Searches,” *JHEP* **05**, 112 (2016), [arXiv:1603.01366 \[hep-ph\]](#).
- [79] Seng Pei Liew, Michele Papucci, Alessandro Vichi, and Kathryn M. Zurek, “Mono-X Versus Direct Searches: Simplified Models for Dark Matter at the LHC,” (2016), [arXiv:1612.00219 \[hep-ph\]](#).
- [80] Yang Bai, James Bourbeau, and Tongyan Lin, “Dark matter searches with a mono- $Z'$  jet,” *JHEP* **06**, 205 (2015), [arXiv:1504.01395 \[hep-ph\]](#).
- [81] Arpit Gupta, Reinard Primulando, and Prashant Saraswat, “A New Probe of Dark Sector Dynamics at the LHC,” *JHEP* **09**, 079 (2015), [arXiv:1504.01385 \[hep-ph\]](#).
- [82] Marcelo Autran, Kevin Bauer, Tongyan Lin, and Daniel Whiteson, “Searches for dark matter in events with a resonance and missing transverse energy,” *Phys. Rev.* **D92**, 035007 (2015), [arXiv:1504.01386 \[hep-ph\]](#).
- [83] Malte Buschmann, Joachim Kopp, Jia Liu, and Pedro A. N. Machado, “Lepton Jets from Radiating Dark Matter,” *JHEP* **07**, 045 (2015), [arXiv:1505.07459 \[hep-ph\]](#).
- [84] Mengchao Zhang, Minho Kim, Hye-Sung Lee, and Myeonghun Park, “Probing the chirality of dark matter at colliders with dark photon showering,” (2016), [arXiv:1612.02850 \[hep-ph\]](#).
- [85] Jason Kumar and Danny Marfatia, “Matrix element analyses of dark matter scattering and annihilation,” *Phys. Rev.* **D88**, 014035 (2013), [arXiv:1305.1611 \[hep-ph\]](#).
- [86] Asher Berlin, Stefania Gori, Tongyan Lin, and Lian-Tao Wang, “Pseudoscalar Portal Dark Matter,” *Phys. Rev.* **D92**, 015005 (2015), [arXiv:1502.06000 \[hep-ph\]](#).
- [87] Mohammad Abdullah, Anthony DiFranzo, Arvind Rajaraman, Tim M. P. Tait, Philip Tanedo, and Alexander M. Wijangco, “Hidden on-shell mediators for the Galactic Center  $\gamma$ -ray excess,” *Phys. Rev.* **D90**, 035004 (2014), [arXiv:1404.6528 \[hep-ph\]](#).
- [88] Joshua Ellis, “TikZ-Feynman: Feynman diagrams with TikZ,” *Comput. Phys. Commun.* **210**, 103–123 (2017), [arXiv:1601.05437 \[hep-ph\]](#).