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Thermofield dynamics and Gravity

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Abstract

Thermofield dynamics is presented in terms of a path-integral using coherent states, equivalently, using a coadjoint orbit action. A field theoretic formulation in terms of fields on a manifold $\mathcal{M} \times \tilde{\mathcal{M}}$ where the two components have opposite orientation is also presented. We propose formulating gravitational dynamics for noncommutative geometry using thermofield dynamics, doubling the Hilbert space modeling the noncommutative space. We consider 2+1 dimensions in some detail and since \mathcal{M} and $\tilde{\mathcal{M}}$ have opposite orientation, the commutative limit leads to the Einstein-Hilbert action as the difference of two Chern-Simons actions.

1 Introduction

It is well established by now that thermofield dynamics (TFD) gives a natural framework to analyze time-dependent processes at finite temperature [1, 2]. But the formalism goes beyond this limited context. It is ultimately a method for describing mixed states as pure states in an enlarged Hilbert space. Therefore, it is the natural theoretical framework for physical contexts where entropy plays an important role. And nowhere is entropy more central or more mysterious than in gravity and so it is not surprising that the description of states in a black hole background involves thermofield dynamics [3, 4, 5].

While most discussions of TFD focus on algebraic aspects of the formalism, or on detailed Feynman diagram techniques, in this paper we will write a general path-integral for TFD in terms of functional integration over unitary matrices. For simplicity, we will consider finite-dimensional Hilbert spaces, with a large dimension taken at the end. An alternate representation would be in terms of an auxiliary fermion field. These representations of TFD are interesting in their own right, but part of our motivation is to apply this to gravity on noncommutative spaces [6, 7].

The existence of entropy for empty gravitational backgrounds such as de Sitter space suggests the idea of assigning a set of states to space itself. In turn, this leads to the notion of noncommutative or fuzzy spaces, since the basic premise of the latter is that some smooth manifolds can be obtained as an approximation to a Hilbert space of states as some parameter is taken to be very large. (If we start with a finite-dimensional Hilbert space, this parameter is usually the dimension itself.) The possibility that gravity might emerge from how this limit is taken was pointed out some time ago [8]. What emerges naturally is Chern-Simons gravity. Here we will argue that one can get the Einstein-Hilbert action if the whole problem is set within TFD and we assign gravitational fields of opposite chirality to the physical system and the tilde system. In other words, our basic suggestion is that one must double the Hilbert space modeling the noncommutative geometry and construct dynamics using thermofield dynamics. (The extension of TFD for fields and the related diagrammatic perturbation theory on a noncommutative space have been considered before [9]. This is different from what we propose here.)

There are, of course, many alternate approaches to gravity using noncommutative spaces going back to the original suggestion by Connes, which leads to the spectral action of Connes and Chamseddine [10], Other approaches include the use of a θ -deformed differential geometry, with and without Drinfeld twists [11, 12] and the emergence of spacetime and gravity from matrix models [13]. (The matrix model approach has similarities to ours, but is still significantly different.)

There has also been more general interest in noncommutative spaces. In fact, field theories on noncommutative spaces have been an important topic of research for a long time now [14]. Such spaces can arise as brane solutions in certain contexts in string theory and in the matrix version of M-theory [15]. Gauge theories on such spaces can describe

fluctuations of the brane solutions and this has also contributed to interest about field theories on fuzzy and noncommutative spaces.

Among noncommutative spaces, there is a subclass which can be described by finitedimensional matrices; these are the fuzzy spaces and, by now, there are many examples of such spaces [6, 7]. When the dimension of the matrices becomes large, these spaces tend to their smooth versions in terms of both the geometry and the algebra of functions on such spaces. For much of what we do in this paper, we will use finite-dimensional Hilbert spaces, so the discussion is within this class of spaces. A more general starting point is possible, but we may note that the finiteness of entropy for the de Sitter universe suggests that the use of a finite-dimensional Hilbert space to describe all degrees of freedom is worthy of consideration as a basic premise for physical theories. We also point out another facet of this, namely, entropy considerations are also very suggestive of treating the event horizon as a fuzzy sphere [16]. The idea that entropy may play a stronger role has also been suggested recently, namely, that gravity itself can be related to entropy and the first law of thermodynamics [17], or from the first law for CFTs [18], or that it may be an entropic force [19]. There is also a recent approach based on a matrix model description of the $\mathcal{N} = 4$ supersymmetric Yang-Mills theory which starts with a finite number of states, see [20].

The time-evolution of a single quantum system starting from a given pure state can be expressed by a path-integral using coherent states or in terms of unitary matrices. The relevant action is in the form of a coadjoint orbit action [21]. In the next section, we rewrite TFD for a single quantum system in a similar way. This requires a doubled set of coherent states and the action relevant for this path-integral involves the trace over a matrix P which has the eigenvalue +1 for the system under consideration and -1 for the tilde system. We can rephrase this in terms of an action defined on a closed contour on a cylinder $\mathbb{R} \times S^1$ with a single winding around the S^1 cycle. Multiple holonomies around this cycle can be related to the Rényi entropy.

We then re-express this in a more familiar form as the functional integral of a field theory, the fields being two fermions. The fermions are defined on two copies of a suitable Kähler manifold \mathcal{M} with opposite orientation. The fields are also coupled to a background field which is a multiple of the Kähler form on \mathcal{M} . A variant of this formulation is to take the fields to be spinors and the action to be the massless Dirac action,. A limit $c \to \infty$ has to be taken at the end, where c plays the role of the speed of light in the action. One can also generalize this to a multipartite system by considering multiple copies of the fields.

In section 3, we discuss how these ideas may apply to gravity, mostly in the setting of the three-dimensional (or 2+1 dimensional) case. A doubling of the Hilbert space defining the fuzzy geometry is introduced. The basic (Euclidean) symmetry under consideration is $SU(2)_L \times SU(2)_R$. While the time-component of a gauge field for this symmetry appears in the thermofield action, the spatial components arise as auxiliary fields which give a simple way to encode the symmetry in the large N limit. These have to be eliminated, i.e., a specific choice for these gauge fields has to be made, at the end. One way to do this would be to choose them as extrema of the action in the large N limit. Gravitational field equations arise as this choice.

One of the basic suggestions in this paper is that, for the doubled Hilbert space defining the geometry, the gauge fields of, say, $SU(2)_L$ couple to one component while the gauge fields of $SU(2)_R$ couple to the other, the tilde part. The action in the large N limit becomes a Chern-Simons action for the $SU(2)_L$ fields and a similar one for the $SU(2)_R$ fields, with a crucial negative sign, for the tilde part. The full action is thus the difference of two Chern-Simons actions, which is equivalent to Einstein gravity in three dimensions.

We close with a brief discussion of the Minkowski signature and comments on comparison of the results with the literature. A short appendix elaborates on the Hall effect connection for one version of the TFD path-integral.

2 Generalizing thermofield dynamics

2.1 Action and functional integral for thermofield dynamics

We start by considering the thermal average of an observable \mathcal{O} at temperature β^{-1} defined by

$$\langle \mathcal{O} \rangle = \operatorname{Tr}(\rho \,\mathcal{O}) = \frac{1}{Z} \operatorname{Tr}\left(e^{-\beta H} \mathcal{O}\right), \qquad Z = \operatorname{Tr}\left(e^{-\beta H}\right)$$
(1)

The density matrix ρ corresponds to a mixed state. The basic idea of thermofield dynamics is to represent the average $\langle \mathcal{O} \rangle$ as the expectation value of the operator \mathcal{O} for a pure state. This will require a doubling of the Hilbert space of states. It is easy to see that we cannot represent ρ as a pure state without doubling, since $\rho^2 = \rho$ for a pure state and we have $\rho^2 \neq \rho$ for the thermal density matrix and no unitary transformation can change this property. If \mathcal{H} denotes the Hilbert space of states (which can and will be taken to be finite dimensional for most of the discussion), then the Hilbert space for thermofield dynamics is $\mathcal{H} \otimes \tilde{\mathcal{H}}$, where $\tilde{\mathcal{H}}$ is a copy of \mathcal{H} itself. A general state in $\mathcal{H} \otimes \tilde{\mathcal{H}}$ is of the form $|m, \tilde{n}\rangle$. The thermal vacuum is then defined as [1]

$$|\Omega\rangle = \frac{1}{\sqrt{Z}} \sum_{n} e^{-\frac{1}{2}\beta E_n} |n, \tilde{n}\rangle$$
⁽²⁾

We have used a basis of eigenstates of the Hamiltonian. In $|\Omega\rangle$ the corresponding states from each Hilbert space contribute to this sum. With this choice, it follows easily that

$$\langle \Omega | \mathcal{O} | \Omega \rangle = \frac{1}{Z} \sum_{m,n} e^{-\frac{1}{2}\beta(E_n + E_m)} \langle m | \mathcal{O} | n \rangle \langle \tilde{m} | \tilde{n} \rangle$$

$$= \frac{1}{Z} \sum_n e^{-\beta E_n} \langle n | \mathcal{O} | n \rangle = \operatorname{Tr}(\rho \mathcal{O})$$

$$(3)$$

where we have used the fact that \mathcal{O} only acts on \mathcal{H} , and $\langle \tilde{m} | \tilde{n} \rangle = \delta_{mn}$. The thermal average is thus expressed as the expectation value over the pure state $|\Omega\rangle$. Thermodynamic

entropy for ρ will arise as an entanglement entropy as we restrict the description to just one component, namely \mathcal{H} , of this doubled Hilbert space.

The tilde Hilbert space is usually chosen as a copy of the dual space of \mathcal{H} , namely \mathcal{H}^* . The motivation for this choice is that time-evolution in $\tilde{\mathcal{H}}$ is then given by -H, with the full Hamiltonian being

$$\check{H} = H - \check{H} = H \otimes \mathbf{1} - \mathbf{1} \otimes H \tag{4}$$

The state $|\Omega\rangle$ obeys $\dot{H}|\Omega\rangle = 0$ and is independent of time. Thus for a free bosonic field with single particle energies ω_k ,

$$\check{H} = \sum_{k} \omega_k \left(a_k^{\dagger} a_k - \tilde{a}_k^{\dagger} \tilde{a}_k \right) \tag{5}$$

It is possible to introduce a Bogolyubov transformation from $a, a^{\dagger}, \tilde{a}, \tilde{a}^{\dagger}$ to A, A^{\dagger} such that $|\Omega\rangle$ is defined by $A_k|\Omega\rangle = 0$ [2]. We will not go into this at this point.

A general density matrix ρ obeys the Liouville equation

$$i\frac{\partial\rho}{\partial t} = H\,\rho - \rho\,H\tag{6}$$

It is possible to construct an action whose variational equation of motion is (6). This is given by

$$S = \int dt \operatorname{Tr} \left[\rho_0 \left(U^{\dagger} \, i \frac{\partial U}{\partial t} - U^{\dagger} \, H \, U \right) \right] \tag{7}$$

where ρ_0 is the initial density matrix and U is a unitary matrix on \mathcal{H} . The dynamical variable in (7) is U and ρ is defined as $\rho = U \rho_0 U^{\dagger}$. Notice that (7) is in the form of a coadjoint orbit action. It may be viewed as a general action for a general quantum system and is very useful in extracting effective actions for collective phenomena [22]. It is particularly suited for semiclassical and large N expansions. (A few papers which focus on this aspect, by no means an exhaustive list, are [22, 23].)

Our first goal is to construct a similar action for the calculation of averages and correlators in the thermofield state $|\Omega\rangle$. Although we initiated the discussion using $\rho = Z^{-1} e^{-\beta H}$, the idea of thermofield states can be used with any density matrix, so we will consider this more general situation. We start by introducing a different notation which will help to simplify the calculations. Let $\{|n\rangle\}$ denote a basis of states for the *N*-dimensional Hilbert space, initially chosen to be finite-dimensional as mentioned in the introduction. It is convenient to represent these in terms of coherent states which we denote by $|z\rangle$ (and a similar set $|w\rangle$ for the tilde part). The states $|n\rangle$ can then be described using the coherent state wave functions $\langle z|n\rangle = \phi_n(z)$, and $\langle w|n\rangle = \chi_n(w)$. (We use wave functions rather than the states $|z\rangle$, $|w\rangle$ since they will be more appropriate for the path integral version.)

Explicitly, there are many ways to construct the coherent states, for example, in terms of the rank r SU(2) representations (with N = r + 1), or rank 1 representation of U(N), or using other subgroups of U(N) if the dimensions are compatible. The coherent states

may be viewed as sections of an appropriate line bundle over a suitable Kähler space \mathcal{M} (of real dimension 2 d) (which would be $\mathbb{CP}^1 = SU(2)/U(1)$ for the SU(2) coherent states and \mathbb{CP}^{N-1} for the rank 1 U(N) coherent states). Put another way, we consider \mathcal{M} as a phase space and choose the symplectic structure to be a multiple of the fundamental Kähler form on \mathcal{M} and carry out quantization. In the Bargmann polarization, which is most convenient for us, the wave functions are holomorphic and correspond to the coherent states. For a compact space of finite volume, the dimension of the Hilbert space will be finite and will carry a representation of the isometry group of the space, appropriately U(N) (for \mathbb{CP}^{N-1}) or SU(2) (for \mathbb{CP}^1) as mentioned above. For the case of \mathbb{CP}^{N-1} , the wave functions for the basis of states is of the form $z_k, k = 1, 2, \dots, N$, where z_k are the homogeneous coordinates; the local coordinates in one coordinate patch would be $z_k/z_N, k = 1, 2, \dots, (N-1)$. For the case of \mathbb{CP}^1 with local coordinates z, \bar{z} (for one coordinate patch), we get

$$\phi_n(z,\bar{z}) = \left[\frac{(r+1)!}{n!\,(r-n)!}\right]^{\frac{1}{2}} \frac{z^n}{(1+\bar{z}z)^{r/2}}, \qquad n=0,1,\cdots,r$$
(8)

which obey

$$\left(\partial_{\bar{z}} + \frac{r}{2} \frac{z}{(1+\bar{z}z)}\right) \phi_n(z) = 0 \tag{9}$$

This is the holomorphicity condition appropriate to the space \mathbb{CP}^1 . We shall shortly return to the explicit formulae for $\phi_n(z)$ and $\chi_n(w)$, but for now, we observe that, for the coherent states, we can choose an orthonormal basis with

$$\int_{\mathcal{M}} d\mu(\bar{z}, z) \,\phi_n^* \,\phi_m = \delta_{nm}, \qquad \int_{\mathcal{M}} d\mu(\bar{w}, w) \,\chi_n^* \,\chi_m = \delta_{nm} \tag{10}$$

(The integration measure is the phase volume; thus, for example, $d\mu = dz d\bar{z}/(\pi (1 + \bar{z}z)^2)$, for \mathbb{CP}^1 .) The thermofield state $|\Omega\rangle$ can then be represented as

$$|\Omega\rangle = \chi_n^* \left(\sqrt{\rho}\right)_{nm} \phi_m \tag{11}$$

Here we take $\sqrt{\rho}$ to act on the ϕ 's, i.e., on \mathcal{H} . Thus $\sqrt{\rho} \phi$ is an element of \mathcal{H} . The action of an operator \mathcal{O} on $|\Omega\rangle$ is given by

$$\mathcal{O}\left|\Omega\right\rangle = \chi^{\dagger}\sqrt{\rho} \ \mathcal{O}\phi \tag{12}$$

with the expectation value

$$\langle \Omega | \mathcal{O} | \Omega \rangle = \int (\phi^{\dagger} \sqrt{\rho} \chi) (\chi^{\dagger} \sqrt{\rho} \mathcal{O} \phi) = \int \phi_{a}^{*} (\sqrt{\rho})_{ab} \chi_{b} \chi_{c}^{*} (\sqrt{\rho})_{cd} (\mathcal{O} \phi)_{d}$$

$$= \operatorname{Tr}(\rho \mathcal{O})$$

$$(13)$$

Our notation also makes explicit some of the well known properties of TFD, particularly its relation to C^* -algebras and the Tomita-Takesaki theory or the HHW version of the same [24]. (For a concise review of this aspect of TFD, see [2].) A key result is the existence of an antilinear operation J, referred to as the modular conjugation, and a so-called modular operator Δ . These obey the properties

$$\Delta^{\dagger} = \Delta, \qquad \Delta > 0, \qquad \Delta |\Omega\rangle = |\Omega\rangle$$
$$J^{\dagger} = J, \qquad J^{2} = 1, \qquad J |\Omega\rangle = |\Omega\rangle$$
(14)

In our notation (12) for $|\Omega\rangle$, these are given by $\Delta = e^{-\beta \check{H}}$ (for the thermal state $\rho = Z^{-1}e^{-\beta H}$), and

$$J\phi = \chi^*, \qquad J\chi = \phi^*, \qquad J\lambda\phi = \lambda^*\chi^*$$
 (15)

Thus

$$J |\Omega\rangle = J \left(\chi^{\dagger} \sqrt{\rho} \ \mathcal{O}\phi\right) = \phi_a(\sqrt{\rho})^*_{ab} \chi^*_b = \chi^{\dagger} \sqrt{\rho^{\dagger}} \phi = \chi^{\dagger} \sqrt{\rho} \phi$$
$$= |\Omega\rangle$$
(16)

where we used the fact that $\rho^{\dagger} = \rho$. All the properties (14) can be easily verified.

We now turn to the time-evolution of the states. For this purpose, we will rewrite Ω in a slightly different notation as

$$\Omega(\bar{z},\bar{u}) = \sum_{n\,m} \psi_n(\bar{u}) \left(\sqrt{\rho}\right)_{nm} \phi_m(\bar{z}) \tag{17}$$

Since we use the wave function representation for the states from now on, we write Ω rather than $|\Omega\rangle$. Also, we write ψ in place of χ^* . The time-evolved wave functions can be represented as a path integral,

$$\phi_n(\bar{z},t) = \int [\mathcal{D}z] \, e^{i \, S(z,\bar{z},t \, | z',\bar{z}')} \, \phi_n(\bar{z}',0) \tag{18}$$

where $S(z, \bar{z}, t | z', \bar{z}')$ is the action for the coherent states integrated from z', \bar{z}' to z, \bar{z} over time t. For the full thermofield state Ω we find

$$\Omega(\bar{z},\bar{u},t) = \int [\mathcal{D}z\,\mathcal{D}u]\,e^{i\,S(z,\bar{z},t\,|z',\bar{z}')}\,e^{i\,\tilde{S}(u,\bar{u},t\,|u',\bar{u}')}\,\Omega(\bar{z}',\bar{u}',0) \tag{19}$$

Thus a vacuum-to-vacuum amplitude has the form

$$F = \int [\mathcal{D}z \,\mathcal{D}u] \,\Omega^*(z, u) \,e^{i\,S(z,\bar{z},t\,|z',\bar{z}')} \,e^{i\,\tilde{S}(u,\bar{u},t\,|u',\bar{u}')} \,\Omega(\bar{z}',\bar{u}') \tag{20}$$

In the operator notation

$$\int \phi_k^*(z) \ e^{i S(z,\bar{z},t \mid z',\bar{z}')} \ \phi_l(\bar{z}') = \langle k \mid e^{-iH_z t} \mid l \rangle$$
(21)

where H_z is the Hamiltonian for the (z, \bar{z}) -system. In other words, in the coherent state basis,

$$e^{iS(z,\bar{z},t\,|z',\bar{z}')} = \langle z|\,e^{-iH_zt}\,|z'\rangle \tag{22}$$

The amplitude F can therefore be written as

$$F = \sum (\sqrt{\rho})_{kl}^* \langle k | e^{-iH_z t} | a \rangle \langle l | e^{-iH_u t} | b \rangle (\sqrt{\rho})_{ab}$$
(23)

We choose $H_u = -H^T$, so that

$$F = \operatorname{Tr}\left(\sqrt{\rho^{\dagger}} e^{-iHt} \sqrt{\rho} e^{iHt}\right)$$
(24)

If we introduce operators A, B in the (z, \overline{z}) -sector (or on \mathcal{H}), we can write, for $t > t_1 > t_2$,

$$\langle A(t_1) B(t_2) \rangle = \operatorname{Tr}\left(\sqrt{\rho} \ U(t,t_1) \ A \ U(t_1,t_2) \ B \ U(t_2,0) \ \sqrt{\rho} \ U^{\dagger}(t,0)\right)$$
(25)

In terms of a time-contour, this may be represented as shown in Fig. 1. This is not the more familiar Schwinger-Bakshi-Mahanthappa-Keldysh (SBMK) closed time contour [25]. That has insertion of ρ at t = 0 and identity at t = T. For equilibrium cases, it is equivalent to the F given here. If ρ is not the equilibrium choice, then there are several amplitudes (with different physical meanings) one can consider. We can define a more general state specified by a matrix K as

$$\Omega_K = \sum_{n\,m} \psi_n(\bar{u}) K_{nm} \,\phi_m(\bar{z}) \tag{26}$$

The SBMK contour is then obtained for the amplitude

$$F_{1\rho} = \int [\mathcal{D}z \,\mathcal{D}u] \,\,\Omega_1^*(z,u) \,\, e^{i\,S(z,\bar{z},t\,|z',\bar{z}')} \,\, e^{i\,\tilde{S}(u,\bar{u},t\,|u',\bar{u}')} \,\,\Omega_\rho(\bar{z}',\bar{u}') \tag{27}$$

This is the physically relevant amplitude (with insertions of operators such as A, B as required) for time-evolution of a closed statistical system starting with a given ρ . However, if we view the physical system under consideration as a subsystem within a larger closed system which is in a pure state which is an eigenstate of the total Hamiltonian, then the amplitude (25) would be the relevant one.

We now turn to the action which will give the expected behavior for the (z, \bar{z}) - and (u, \bar{u}) -sectors. It is given by

$$S = \int dt \left[\left(i \, \bar{z}_k \dot{z}_k - \bar{z}_k \, H_{kl} z_l \right) + \left(i \, \bar{u}_k \dot{u}_k + \bar{u}_k \, H_{kl}^T u_l \right) \right] \tag{28}$$

with the constraints

$$\bar{z}_k \, z_k = 1, \qquad \bar{u}_k \, u_k = 1 \tag{29}$$



Figure 1: Time contours. The left side shows the contour for equation (25) and the right side for equation (42).

This action is easily quantized in terms of geometric quantization. For the (z, \bar{z}) -sector, the canonical one-form and two-form are given by

$$A = \frac{i}{2}(\bar{z}_k \, dz_k - d\bar{z}_k \, z_k), \qquad \omega = i \, d\bar{z}_k \wedge dz_k \tag{30}$$

The polarization condition on the wave functions can be chosen as

$$\nabla_z \Psi = \left(\frac{\partial}{\partial z_k} + \frac{\bar{z}_k}{2}\right) \Psi = 0 \tag{31}$$

which leads to wave functions of the form

$$\Psi = e^{-z_k \bar{z}_k/2} f(\bar{z}) \tag{32}$$

with z_k acting as $\partial/\partial \bar{z}_k$ on the the f's. The constraint shows that the f can have one power of \bar{z} , which implies that $f(\bar{z}) \sim \bar{z}_k$. There are exactly N states, giving the rank 1 representation of U(N). The Hamiltonian operator is

$$H = \bar{z}_k H_{kl} \frac{\partial}{\partial \bar{z}_l} \tag{33}$$

We see that matrix elements of this Hamiltonian reproduce H_{kl} . For the (u, \bar{u}) -sector, we again have $\Psi = \exp(-u \cdot \bar{u}/2) f(\bar{u})$ with u_k acting as $\partial/\partial \bar{u}_k$. The Hamiltonian is

$$H = -\bar{u}_k H_{kl}^T \frac{\partial}{\partial \bar{u}_l}, \qquad \langle k | H | l \rangle = -H_{kl}^T$$
(34)

The operation $H \to -H^T$ represents charge conjugation in the Lie algebra of U(N). It is useful to define

$$z_k = \xi_{k1}, \quad \bar{u}_k = w_k = \xi_{k2}, \qquad P = \begin{bmatrix} 1 & 0 \\ 0 & -1 \end{bmatrix}$$
 (35)

With a partial integration on the term involving the time-derivative of u_k in (28), we can bring the action to the form

$$S = \int dt \sum_{\alpha,\beta=1,2} P_{\alpha\beta} \left(i\bar{\xi}_{k\beta} \dot{\xi}_{k\alpha} - \bar{\xi}_{k\beta} H_{kl} \xi_{l\alpha} \right)$$
$$= \int dt \operatorname{Tr} \left[P \left(i \xi^{\dagger} \dot{\xi} - \xi^{\dagger} H \xi \right) \right]$$
(36)

The variables z_k , u_k with the constraint (29) define (two copies of) \mathbb{CP}^{N-1} , so it is useful to use a notation in terms of group elements. Writing $\xi_{k\alpha} = U_{k0}^{(\alpha)} = \langle k | U^{(\alpha)} | 0 \rangle$, for two unitary matrices $U^{(\alpha)}$, we find

$$S = \int dt \left[\left(i U^{(1)\dagger} \dot{U}^{(1)} - U^{(1)\dagger} H U^{(1)} \right)_{00} - \left(i U^{(2)\dagger} \dot{U}^{(2)} - U^{(2)\dagger} H U^{(2)} \right)_{00} \right]$$
(37)

The N states forming the fundamental representation of SU(N) can be viewed as being generated from a highest weight state, usually the vacuum, by the action of various operators. Group theoretically, they can be constructed from the elements $\langle k|U|0\rangle$ for a chosen state $|0\rangle$. We have explicitly included this in our notation. The action (37) is close to the form (7), but the key difference is that it is the difference of two actions of the form (7). In the same notation, the state Ω can be written as

$$\Omega = \bar{z}_k \sqrt{\rho}_{kl} w_l = \bar{\xi}_{k1} \sqrt{\rho}_{kl} \xi_{l2} = \langle 0 | U^{(1)\dagger} \sqrt{\rho} U^{(2)} | 0 \rangle$$
(38)

Although we have used coherent states to arrive at (37) and (38), these results are no longer dependent on any specific representation for the states. In terms of (38), we can represent (20) as

$$F = \int [\mathcal{D}U] \ e^{iS} \ \langle 0|U^{(2)\dagger}(t)) \ \sqrt{\rho^{\dagger}} \ U^{(1)}(t)|0\rangle \ \langle 0|U^{(1)\dagger}(0) \ \sqrt{\rho} \ U^{(2)}(0)|0\rangle \tag{39}$$

with S as given in (37) and $[\mathcal{D}U]$ denotes functional integration for the two copies of \mathbb{CP}^{N-1} .

In calculating F as in (20), or correlators as in (25), effectively we have a closed timecontour with insertions of Ω at two points corresponding to t = 0 and t = T with $T \to \infty$ eventually. This suggests a neat way to rewrite these results. We will use a complex time variable $\tau = t + i\theta$, with the identitication $\theta \sim \theta + 2\pi$, so that τ describes a cylinder $\mathbb{R} \times S^1$. Further we introduce a one-form

$$\mathcal{A} = -i \left[H \, dt + \frac{i}{2\pi} \log \rho \, d\theta \right] \tag{40}$$

The contour of integration C is then taken as starting at $(t, \theta) = (0, 0)$ going to $(0, \pi)$ to (T, π) to $(T, 2\pi) \sim (T, 0)$ to (0, 0), as shown in the second figure in Fig. 1. This goes around S^1 once. We then consider the $\mathcal{P}e^{\oint_C \mathcal{A}}$, where \mathcal{P} denotes path-ordering along C. For the segments along the *t*-direction, we can use (22) to obtain the action as before. For the segments (0, 0) to $(0, \pi)$ and (T, π) to $(T, 2\pi)$ along the imaginary time-direction, we get factors of $\sqrt{\rho}$. For this, we use the same result as (22) with $(i/2) \log \rho$ in place of the Hamiltonian H. Thus

$$e^{iS(w,z',\pi)} = \langle w | e^{\frac{1}{2}\log\rho} | z' \rangle \tag{41}$$

Notice that the factor $\int U^{\dagger} \dot{U} d\tau$ is insensitive to whether we integrate along real or imaginary directions. We also take H to be independent of θ so that $H(\text{at } \theta = \pi) = H(\text{at } \theta = 0)$. With this result, we can write the path integral for thermofield dynamics as

$$F_J = \int [\mathcal{D}U] \exp\left[\oint_C \left(-U^{\dagger} \dot{U} + U^{\dagger} \mathcal{A} U\right)_{00} + \oint AJ + BJ'\right]$$
(42)

where we have also introduced sources which facilitate the insertion of operators A, B, etc. Setting J = J' = 0 gives the vacuum to vacuum transition amplitude. In operator notation, this is given by

$$F(C) = \operatorname{Tr} \mathcal{P}e^{\oint_C \mathcal{A}} \tag{43}$$

Thus F(C) is the Wilson loop integral for the connection (40) taken along the contour Cwhich winds once around the S^1 direction of the cylinder. Consider now a small deformation of the contour $C \to C'$ at a point τ_0 along the path. Evidently

$$F(C') - F(C) = \operatorname{Tr} \mathcal{P} \exp\left(\int_{\tau_0}^{(0,2\pi)}\right) \mathcal{F}\sigma \ \mathcal{P} \exp\left(\int_{(0,0)}^{\tau_0}\right)$$
(44)

where $\sigma = \delta t \wedge \delta \theta$ is the infinitesimal area between C' and C at τ_0 . The curvature \mathcal{F} is given by

$$\mathcal{F} = \frac{1}{2\pi} \frac{\partial}{\partial t} \log \rho + i \frac{\partial H}{\partial \theta} + \frac{i}{2\pi} \left[\log \rho, H \right]$$
(45)

In our case, $\partial H/\partial \theta = 0$ and $\partial \rho/\partial t = 0$. (We have insertions of the same ρ at t = 0 and t = T.) Thus, the curvature \mathcal{F} vanishes if $[\rho, H] = 0$. For the equilibrium case where this is obtained, we can do deformation of the contour C reducing F(C) to a simple holonomy of the connection (40) around the S^1 -direction of the cylinder $\mathbb{R} \times S^1$. (This is without insertions of operators; for correlators there will be a nontrivial segment along the *t*-direction.)

We can go further and consider multiple windings around the S^1 component of the cylinder. We can consider windings inserted at any value of t, defining $W(C, n, t) = F(C + C_n(t))$ where $C + C_n(t)$ corresponds to the contour C with extra n windings around S^1 at the point t. Thus the Rényi entropy is defined in terms of the holonomy as

$$S_R(t) = -\left(\frac{W(C, n, t) - 1}{n - 1}\right)$$
(46)

Since thermofield dynamics uses pure states, entropy must be defined in terms of the basic observables of the formalism. For the vacuum amplitude, i.e., in the absence of insertions of operators on the real time line, the holonomy is the only observable we have and hence equation (46) gives the natural definition.

2.2 A field theoretic representation

We have represented thermofield dynamics in (39) as a functional integral over group elements. There are other equivalent ways of representing it, one of which is as the functional integral of a field theory. This is will prove useful as well. The basic idea is that the quantum system can be viewed as the one-particle sector of a field theory. The time-evolution matrix element (21) can be written as

$$\langle k|e^{-iHt}|l\rangle = \langle 0|a_k e^{-iHt} a_l^{\dagger}|0\rangle = \langle 0|T a_k(t) a_l^{\dagger}(0)|0\rangle$$

$$= \mathcal{N} \int [da \, da^*] e^{iS} a_k(t) a_l^{\dagger}(0)$$

$$(47)$$

$$S = \int dt \left[a_k^*(i\partial_0) a_k - a_k^* H_{kl} a_l \right]$$
(48)

and \mathcal{N} is the standard normalization factor,

$$\mathcal{N}^{-1} = \int [da \ da^*] \ e^{iS} \tag{49}$$

Further, we can introduce a (z, \bar{z}) -dependent field (on \mathcal{M}) given by $\psi(z, \bar{z}, t) = \sum_k a_k z_k$, $\psi^{\dagger}(z, \bar{z}, t) = \sum_k a_k^{\dagger} \bar{z}_k$. (We use ψ^* when we integrate over the c-number versions of these fields in the functional integral.) The diagonal coherent state representation of operators also allows us to introduce $H(z, \bar{z})$ such that

$$H_{kl} = \int_{\mathcal{M}} d\mu(z, \bar{z}) \ \bar{z}_k \ H(z, \bar{z}) \ z_l \tag{50}$$

The action (48) can thus be written as

$$S = \int dt \, d\mu(z,\bar{z}) \, \left[\psi^*(i\partial_0)\psi - \psi^\dagger \, H(z,\bar{z}) \, \psi \right] \tag{51}$$

The part of the action (36) with the negative eigenvalue for P can be represented in a similar way, with a field ϕ, ϕ^{\dagger} .

The fields ψ , ψ^* (and ϕ , ϕ^*) are restricted since ψ is holomorphic and ψ^* is antiholomorphic. Thus the functional integration over these fields have also to be suitably restricted. A convenient way to extend the functional integration over all fields and still restrict the dynamics to holomorphic ψ 's is to use the Landau level trick. We take ψ to describe a charged particle on \mathcal{M} with a background magnetic field which is constant in a suitable basis. This can be done for by taking the field to be proportional to the Kähler two-form ω on \mathcal{M} and consider the action

$$S = \int dt \, d\mu(z,\bar{z}) \, \left[\psi^* \left(i \, \partial_0 - A_0(z,\bar{z}) + \frac{D^2 + E_0}{2m} \right) \, \psi \right] \tag{52}$$

where $D_i = \nabla_i - iA_i$ is the gauge and Levi-Civita covariant derivative on ψ and $d(A_i dx^i) = n \omega$ for some parameter n. E_0 is the lowest eigenvalue of $-D^2$. Further, $A_0(z, \bar{z}) = H(z, \bar{z})$ is the Hamiltonian of the original theory. m is a parameter which we will take to be zero eventually. The Hamiltonian for ψ is $H' + A_0$, with

$$H' = -\frac{D^2 + E_0}{2m}$$
(53)

The eigenstates of H' are the Landau levels, the lowest of which obeys a holomorphicity condition, and has zero eigenvalue since we subtracted E_0 (which is the lowest eigenvalue of $-D^2$). The higher states are not holomorphic, but will decouple as $m \to 0$. The dynamics will thus be restricted to the lowest state which corresponds to holomorphic ψ 's. Finally, we introduce

$$\Omega(\psi^*, \phi^*) = \int_{\mathcal{M}} d\mu(z, \bar{z}) d\mu(w, \bar{w}) \ \psi^*(z) \ \phi^*(w) \ z_k \sqrt{\rho}_{kl} \ w_l$$
(54)

Collecting these results together, we conclude that thermofield dynamics is given by

$$F = \mathcal{N} \int [d\psi d\psi^* \, d\phi d\phi^*] \, e^{iS} \, \Omega^*(t) \, \Omega(0)$$

$$S = \int dt \, d\mu(z,\bar{z}) \left[\psi^* \left(i \, \partial_0 - A_0(z,\bar{z}) + \frac{D^2 + E_0}{2m} \right) \psi - \phi^* \left(i \, \partial_0 - A_0(z,\bar{z}) + \frac{D^2 + E_0}{2m} \right) \phi \right]$$

$$= \int dt \int_{\mathcal{M}} d\mu(z,\bar{z}) \, \psi^* \left(i \, \partial_0 - A_0(z,\bar{z}) + \frac{D^2 + E_0}{2m} \right) \psi + \int dt \int_{\tilde{\mathcal{M}}} d\mu(z,\bar{z}) \, \phi^* \left(i \, \partial_0 - A_0(z,\bar{z}) + \frac{D^2 + E_0}{2m} \right) \phi$$
(55)
(55)

As usual, \mathcal{N}^{-1} is given by the integral of e^{iS} over all fields. In the second line of (56) we take $\tilde{\mathcal{M}}$ to have the orientation opposite to that of \mathcal{M} . The fields may be taken to be bosonic, but it will turn out to be more convenient to take them as fermionic fields.

A simple extension will give a generalization of this result to a multipartite system with identical components. Going back to (48), we take the states to be of the form $|k\rangle = |\alpha I\rangle \in \mathcal{H}_1 \otimes \mathcal{H}_2$ and define a set of fermion fields $\psi_I = \sum_{\alpha} a_{\alpha I} z_{\alpha}$. The action may now be written as

$$S = \int dt \int_{\mathcal{M}} d\mu(z,\bar{z}) \ \psi_{I}^{*} \left(i \,\partial_{0}\delta_{IJ} - (A_{0}(z,\bar{z})_{IJ} + \frac{D^{2} + E_{0}}{2m}\delta_{IJ} \right) \psi_{J} + \int dt \int_{\tilde{\mathcal{M}}} d\mu(z,\bar{z}) \ \phi_{I}^{*} \left(i \,\partial_{0}\delta_{IJ} - (A_{0}(z,\bar{z}))_{IJ} + \frac{D^{2} + E_{0}}{2m}\delta_{IJ} \right) \phi_{J}$$
(57)

We may interpret the labels I, J as corresponding to some internal symmetry or degrees of freedom.

There is one more improvement we can do on this formula. If $\mathcal{M} \times \mathbb{R}$ admits spinors, we can replace the action by the Dirac type action

$$S = \int dt \int_{\mathcal{M}} d\mu(z,\bar{z}) \ \bar{\Psi}_I(i\gamma^{\mu}D_{\mu})_{IJ}\Psi_J + \int dt \int_{\tilde{\mathcal{M}}} d\mu(z,\bar{z}) \ \bar{\Phi}_I(i\gamma^{\mu}D_{\mu})_{IJ}\Phi_J \tag{58}$$

where Ψ and Φ are spinors, γ^{μ} are the standard Dirac matrices and $\bar{\Psi} = \Psi^{\dagger}\gamma^{0}$, $\bar{\Phi} = \Phi^{\dagger}\gamma^{0}$. The Hamiltonian for Ψ/Φ now has the form $H' + A_{0}$ with $H' = -i\gamma^{0}\gamma^{i}D_{i}$. The eigenstates of H' are again Landau levels; there are zero modes for H' which satisfy a holomorphicity condition. The other levels are separated by a gap of order of the magnetic field $\sim c n \omega$, where c plays the role of the speed of light for the action (58), relating $\partial/\partial t$ and $\partial/\partial x$. Taking $c \to \infty$ all the nonzero eigenstates decouple¹. (We assume, as usual, that the negative energy levels are all filled.) The degeneracy of states is controlled by $n\omega$ and remains finite in this limit. There is also another limit we can take, namely, $n \to \infty$; in this case, not only the nonzero levels decouple, the degeneracy of the zero modes also tends to infinity. This is equivalent to an expansion in inverse powers of the background field.

2.3 A quick summary

A quick recapitulation of the results in this section will be useful. We have expressed thermofield dynamics for a single quantum system as a functional integral over a unitary group, the action for which is given in (37). This action involves the trace over a matrix Pwhich has the eigenvalue +1 for the system under consideration and -1 for the tilde system. Equivalently, one can use an action defined on a closed contour on a cylinder $\mathbb{R} \times S^1$ with one winding around the S^1 cycle. Multiple holonomies around this cycle can be related to the entropy.

¹This is not the usual nonrelativistic limit since we have set the mass to zero in (58).

One can also express thermofield dynamics as field theoretic functional integral with two fields defined on two copies of a suitable Kähler manifold with opposite orientation. The fields are subject to a background field which is a multiple of the Kähler form on the manifold \mathcal{M} . The fields can be taken to be spinors and the action to be the massless Dirac action, with a limit $c \to \infty$ at the end, where c plays the role of the speed of light in the action. Additional degrees of freedom can be incorporated by considering multiple copies of the fields, with, generally, a nonabelian symmetry acting on them.

3 Gravity from noncommutative spaces

A particularly interesting situation to which the foregoing analysis can be applied is a formulation of gravity using noncommutative or fuzzy spaces. The relevance of fuzzy spaces starts with the question: What if we quantized gravity? In that case, we would have a Hilbert space of states, and the continuous manifold description would be obtained as an approximation for large number of degrees of freedom. This is clearly the realm of noncommutative geometry or fuzzy spaces.

A scenario for implementing this idea would be as follows. We can model noncommutative spaces in terms of the lowest Landau level of a quantum Hall system [26, 27]. For example, we can think of the complex projective space \mathbb{CP}^k as SU(k+1)/U(k). It is thus possible to consider uniform background fields on \mathbb{CP}^k which are valued in the algebra of U(k) and which are proportional to the curvatures of \mathbb{CP}^k . Single particle wave functions in this background will fall into SU(k+1) multiplets. The lowest such set of states can be represented by holomorphic wave functions. Put another way, these wave functions are sections of a holomorphic U(k) bundle on \mathbb{CP}^k . They form an N-dimensional Hilbert space \mathcal{H}_1 which may be viewed as a model for the fuzzy version \mathbb{CP}_F^k of the differential manifold $\mathcal{M} = \mathbb{CP}^k$. Functions on the fuzzy space are $N \times N$ matrices viewed as linear transformations on \mathcal{H}_1 . In the large N limit, we recover the usual commutative algebra under pointwise multiplication of functions on \mathbb{CP}^k . This limit can be analyzed by using classical functions (on \mathbb{CP}^k) to represent operators and using *-products to represent operator products.

The isometry group for \mathbb{CP}^k is SU(k + 1). Hence, background fields valued in the algebra of U(k) amount to connections and curvatures for (at least part of) the isometry group. Thus these background fields, since they correspond to gauging the isometries, can be viewed as describing gravitational degrees of freedom. An action for these fields, which may be derived from (57) or (58) (or from (A6) given in the appendix), would thus be a gravitational action.

This point of view regarding gravity was suggested many years ago [8]. The action used in that case was (7) or (A4), not the thermofield case with both positive and negative eigenvalues for P. Simplifying it using *-products of \mathbb{CP}^k led to the Chern-Simons term for the U(k)-valued gauge fields as the leading term in the action. However, there were several points which were not clear at that stage. A priori, since we start with \mathcal{H}_1 , the choice of modeling this as \mathbb{CP}_F^k is arbitrary. The leading CS term which emerges in the large N limit, being topological, is not sensitive to the metrical details of \mathbb{CP}^k , but the subleading ones are. Further, for gravity on a 2k-dimensional spatial manifold (+ time), we need gauge fields valued in the algebra of the corresponding Poincaré group, (or de Sitter group, including a cosmological constant), with the gauge fields corresponding to the translations being the frame fields and those for the Lorentz transformations giving the spin connection. Thus, with Euclidean signature, we need SO(2k+2) rather than U(k). Finally, it was not clear how one could get Einstein gravity rather than CS gravity. With the thermofield approach, one can improve on some of these problems.

In this paper, we make two basic suggestions. The first is that in discussing gravity using the large N limit of a Hilbert space \mathcal{H}_1 , we should use thermofield dynamics. This is motivated by the well-founded expectation that entropy should play an important role in gravity and that thermofield dynamics, which can incorporate entropy within a formalism of pure states, is therefore a natural framework. We thus double the Hilbert space to $\mathcal{H}_1 \otimes \mathcal{H}_1^*$. The large N approximation of these spaces by a manifold \mathcal{M} will introduce gauge fields corresponding to the frame fields and spin connection of \mathcal{M} . Our second suggestion is that these gauge fields for the physical system and the tilde system should be considered as parity conjugates of each other. This is related to the fact that the orientation of the manifold used to define the states, say in (57) or (58), is reversed for the tilde system. Equivalently, this emerges from the matrix P in (A6). With this choice, the action, as we will see below, leads to Einstein gravity rather than Chern-Simons gravity.

Some of the key ingredients are then the following.

- 1. We need a Hilbert space \mathcal{H} which carries a representation of a group G. The latter will eventually become the isometry group of the continuous spacetime $\mathcal{M} \times \mathbb{R}$ which will emerge as we take the limit of some parameter $\theta \to 0$. (θ could be $\sim N^{-\alpha}$, for some $\alpha > 0$.) G = U(k) for the discussion given above, but could be more general, and, in fact, will be SO(4) for the case of three-dimensional Euclidean gravity.
- 2. The Hilbert space will have three components, $\mathcal{H}_1 \otimes \mathcal{H}_2 \otimes \mathcal{H}_3$ with states of the form $|\alpha, a, I\rangle$, where \mathcal{H}_3 refers to any matter system of interest. For gravity, as a first approximation, we will not need to consider excitations of the matter system, which means that we can restrict the matter fields to the ground state. In this case, the states in \mathcal{H} will be taken to be of the form $|\alpha, a, 0\rangle$ corresponding to a representation $R_1 \otimes R_2$ of G with the transformation

$$|\alpha, a, 0\rangle' = g^{(1)}_{\alpha\beta} g^{(2)}_{ab} |\beta, b, 0\rangle$$
(59)

We will consider R_2 to be a fixed representation, and take the dimension of R_1 to be very large to approximate \mathcal{H}_1 by a continuous manifold. In the case of an infinite dimensional representation R_1 , we take a limit of the highest weight vector to approximate to a continuous manifold. R_1 must be a highest weight representation to define symbols and *-products. These representations have to be unitary as required by quantum mechanics.

It is possible to reduce the product $R_1 \otimes R_2$ in terms of irreducible representations but this will not be important for us. (A priori, one could consider different groups G_1 and G_2 acting on \mathcal{H}_1 and \mathcal{H}_2 respectively. We have no good argument to exclude this, except that the minimal way is to use the same group.)

3. To define the functions representing operators and *-products, we will need a set of wave functions. These will be obtained by quantizing the Kähler space G/H for some suitable $H \in G$. This will also define R_1 (and \mathcal{H}_1). Gauge fields will emerge as part of the the procedure for the large N limit.

Given this structure, it is possible to simplify the trace of an operator, say, $\operatorname{Tr}(iH)$. It gives a Chern-Simons form of the appropriate dimension. For a spatial manifold M of dimension 2k, the action (A6) will simplify as the integral of a Chern-Simons (2k+1)-form with gauge group G, more precisely as the difference of Chern-Simons actions for the two chiralities. The gauge fields on the continuous spacetime are introduced to express the unitary transformations on \mathcal{H} in terms of G-transformations on the continuum $\mathcal{M} \times \mathbb{R}$. In a sense, these gauge fields define the small θ (or large N) limit we are taking. The natural question then is whether we can choose them in some optimal fashion. As the optimization requirement, we extremize the limit of the action (A6) with respect to A_{μ} . These optimization conditions are to be considered as the equations of motion for gravity.

Equivalently, we can use the field theoretic action (58) with the spinor fermionic fields of the form $\Psi_{i,0}(z, \bar{z}, t)$, $\Phi_{i,0}(z, \bar{z}, t)$. The gauge fields are matrices of the form

$$A_{ab} = A_i dx^i \,\delta_{ab} + (A_\mu)_{ab} dx^\mu \tag{60}$$

where $d(A_i dx^i) = n\omega$ and $(A_\mu)_{ab}$ are valued in the Lie algebra of G; they are matrices in the representation R_2 .

We now show how these ideas can be applied to gravity in three dimensions. In three Euclidean dimensions, it is well known that Einstein gravity can be formulated as a Chern-Simons theory [28]. This is true for Minkowski signature as well, but we will first discuss the Euclidean case. For this, we will use the generators of SO(4) in the spinor representation, with 4×4 hermitian γ -matrices. The translation generators P_a and the rotation generators S_{ab} are defined as

$$P_a = \frac{\gamma_3 \gamma_a}{2il}, \qquad S_{ab} = \frac{1}{4i} \left(\gamma_a \gamma_b - \gamma_b \gamma_a \right), \qquad a, b = 0, 1, 2.$$
(61)

Here l is a quantity with the dimensions of length which is related to the cosmological constant. The action is then given by

$$S = -\frac{l}{32\pi G} \int \operatorname{Tr}\left[\gamma_5 \left(A \, dA + \frac{2}{3}A^3\right)\right] \tag{62}$$

The gauge field A is built of the frame field $e^a = e^a_\mu dx^\mu$ and the spin connection $\omega^{ab} = \omega^{ab}_\mu dx^\mu$ as

$$A = -i P_a e^a - \frac{i}{2} \omega^{ab} S_{ab} \tag{63}$$

The trace has the property that it gives a pairing between the P_a and S_{bc} only,

$$\operatorname{Tr}(\gamma_5 P_a S_{bc}) = \frac{1}{l} \epsilon_{abc}$$
(64)

This pairing, rather than the usual Cartan metric of SO(4), is crucial in being able to write Einstein gravity as a Chern-Simons theory. In fact, simplifying (62) using (63, 64) gives

$$S = \frac{1}{16\pi G} \int d^3x \,\det e \,\left(R - \frac{3}{2l^2}\right) \tag{65}$$

Since the SO(4) generators split into two chiral SO(3)'s corresponding to the subspaces with $\gamma_5 = 1$ and $\gamma_5 = -1$, i.e., $SO(4) \sim SO(3)_L \times SO(3)_R$, we see that the pairing (64) can indeed be reproduced by the thermofield action is we take P to be proportional to γ_5 . This is precisely what we propose to do.

We will now show how this action can be obtained from thermofield dynamics. For this purpose, we can use the formulation (58) in terms of spinor fields, but it is more illuminating to first see how the action (62) or (65) emerges from (A6) in the large N limit. The N-dimensional space \mathcal{H}_1 can be taken to correspond to a representation of SU(2). Since $SU(2)/U(1) = \mathbb{CP}^1$, we can think of these states as describing fuzzy \mathbb{CP}^1 . The states are thus of the form $|\alpha, i\rangle$, where for i = 1, 2, we have an action of SU(2) given by

$$|\alpha, i\rangle \to h_{\alpha\beta}^{(s)} h_{ij} |\beta, j\rangle \tag{66}$$

where $h_{\alpha\beta}^{(s)}$ denotes the spin-*s* representation and h_{ij} is the spin- $\frac{1}{2}$ representation of SU(2). This SU(2) will form the $SU(2)_L \sim SO(3)_L$ in SO(4). A similar choice of states will be made for the tilde sector, with $SU(2)_R \sim SO(3)_R$ instead of $SU(2)_L \sim SO(3)_L$.

Since we do not consider excitations for the matter part, the intermediate states $|k\rangle$ are of the form $|\beta 0\rangle$ and we can take $U_{\beta\alpha} \equiv \langle \beta 0 | U | \alpha 0 \rangle$ to be a unitary matrix. (Elements such as $\langle \beta J | U | \alpha 0 \rangle$, $J \neq 0$, will correspond to transitions from the matter vacuum to excited states of matter.) Further, we must count all states in $\mathcal{H}_1 \otimes \mathcal{H}_2$, i.e., P_+ is the identity. The action (A6) then simplifies as

$$S = -\int dt \left[\operatorname{Tr}(iA_0)_L - \operatorname{Tr}(iA_0)_R \right]$$
(67)

Focusing on the left chirality part first, the states $|\alpha\rangle$ which correspond to a spin-s representation of $SU(2)_L$, with N = 2s + 1, s = n/2, where n is a positive integer, can be taken as arising from the quantization of \mathbb{CP}^1 viewed as a phase space, with the symplectic two-form $n \omega_K$, where ω_K is the Kähler two-form on \mathbb{CP}^1 . In extracting the large N limit, we can replace operators by their symbols which are the classical functions corresponding to them. If \hat{A} is an operator acting on states of the form $|\alpha, i\rangle$, the symbol for \hat{A} is defined as

$$(\hat{A})_{ik} = A_{ik} = \langle -s, i | h^{(s)\dagger} \hat{A} h^{(s)} | -s, k \rangle$$

$$(68)$$

where $|-s\rangle$ is the highest weight state of the spin-s representation. As a 2×2 matrix, h may be explicitly parametrized as

$$h = \frac{1}{\sqrt{1 + \bar{z}z}} \begin{pmatrix} 1 & z \\ -\bar{z} & 1 \end{pmatrix} \begin{pmatrix} e^{i\theta/2} & 0 \\ 0 & e^{-i\theta/2} \end{pmatrix}$$
(69)

where z, \bar{z} are coordinates for one coordinate patch of \mathbb{CP}^1 . $h^{(s)}$ in (68) is the spin-s representative of h in (69). Notice that the symbol is independent of θ and is a function on \mathbb{CP}^1 . It is also a 2×2 matrix. We now define differential operators on SU(2) by

$$R_a h = h t_a \tag{70}$$

where t_a , a = 1, 2, 3, form a basis for the Lie algebra of SU(2), $t_a = \sigma_a/2$ for the 2 × 2 representation. Explicitly in terms of the group parameters φ^i (like \bar{z}, z, θ in (69)),

$$R_a = i \left(E^{-1} \right)^i_a \frac{\partial}{\partial \varphi^i}, \qquad h^{-1} dh = -i t_a E^a_i d\varphi^i \tag{71}$$

Because of the highest weight condition for the states in (68), $h \mid -s, k \rangle$ obeys a holomorphicity condition,

$$(R_1 - iR_2) h | -s, k \rangle = R_- h | -s, k \rangle = h t_- | -s, k \rangle = 0$$
(72)

The symbol or the classical function corresponding to an operator product $\hat{A}\hat{B}$ is given by

$$\begin{aligned} (\hat{A}\hat{B})_{ik} &= \langle -s, i|h^{\dagger} \hat{A}\hat{B} h| - s, k \rangle \\ &= \sum_{a,j} \langle -s, i|h^{\dagger} \hat{A} h|a, j \rangle \langle a, j| h^{\dagger} \hat{B} h| - s, k \rangle \\ &= A_{ij}B_{jk} + \sum_{r=1}^{N-1} \langle -s, i|h^{\dagger} \hat{A} h| - s + r, j \rangle \langle -s + r, j| h^{\dagger} \hat{B} h| - s, k \rangle \\ &= A_{ij}B_{ik} + \sum_{r=1}^{N-1} \left[\frac{(N-1-r)!}{r!(N-1)!} \right] (R_{+}^{r} A)_{ij} (R_{-}^{r} B)_{jk} \end{aligned}$$
(73)
$$= (A * B)_{ik} \end{aligned}$$

The right hand side of this equation defines the *-product which starts off with the matrix product of the functions A, B followed by additional terms involving derivatives. As Nbecomes large, these derivative terms are suppressed by powers of N. Another useful result is that the trace of an operator \hat{A} can be expressed in terms of the integral of its symbol as

$$\operatorname{Tr}(\hat{A}) = \int d\mu \operatorname{Tr} A \tag{74}$$

The remaining trace on the right hand side is just over the matrix elements of the 2×2 matrix A.

Starting from the Hilbert space, the action (A6) has a "gauge invariance" corresponding to the transformation

$$U \to g U, \qquad \mathcal{A} \to g \mathcal{A} g^{-1} + \dot{g} g^{-1}$$
 (75)

For a gauge transformation with g close to the identity, we can write $g \approx 1 - \Phi$ and (75) simplifies as

$$\hat{\mathcal{A}} \to \hat{\mathcal{A}} - \partial_0 \hat{\Phi} - \hat{\mathcal{A}} \, \hat{\Phi} + \hat{\Phi} \, \hat{\mathcal{A}},$$
(76)

where $\partial_0 = \partial/\partial t$ and we use the hat-notation to emphasize that all quantities are still operators. We can represent (76) in terms of the symbols as

$$\mathcal{A} \to \mathcal{A} - \partial_0 \Phi - \mathcal{A} * \Phi + \Phi * \mathcal{A} \tag{77}$$

This transformation still has the full content of the transformation at the level of operators. The symbols \mathcal{A} and Φ are functions of the coordinates of $M = \mathbb{CP}^1 \times \mathbb{R}$ (\mathbb{R} is for the time variable), and are also 2×2 matrices. For an explicit realization of this transformation it is convenient to introduce a set of auxiliary quantities $A_{\mu}dx^{\mu}$ which is a one-form on M such that an ordinary continuum gauge transformation of A_{μ} with parameter Λ induces the transformation (77). In other words we seek functions

$$\mathcal{A}_0 = \mathcal{A}_0(A_0, A_i), \qquad \Phi = \Phi(\Lambda, A_0, A_i) \tag{78}$$

such that

$$\left. \begin{array}{l} A_0 \to A_0 + \partial_0 \Lambda + [A_0, \Lambda] \\ A_i \to A_i + \partial_i \Lambda + [A_i, \Lambda] \end{array} \right\} \Longrightarrow \mathcal{A}_0 + \partial_0 \Phi + \mathcal{A}_0 * \Phi - \Phi * \mathcal{A}_0 \tag{79}$$

The fact that this can be done is the essence of the Seiberg-Witten transformation [29]. A_0 may be considered as the continuum version of \mathcal{A} while A_i , the spatial components, are additional auxiliary variables. The choice of these A_i is part of how the large N limit is taken. Our approach will be to optimize this choice by the equations of motion following from the large N limit of the action.

The solution for (78) is straightforward, although somewhat involved algebraically and reads [30]

$$\mathcal{A}_{0} = A_{0} + \frac{P^{ab}}{2n} \left[\partial_{a}A_{0} A_{b} - A_{a}\partial_{b}A_{0} + F_{a0}A_{b} - A_{a}F_{b0} \right] + \cdots$$

$$\Phi = \Lambda + \frac{P^{ab}}{2n} \left(\partial_{a}\Lambda A_{b} - A_{a}\partial_{b}\Lambda \right) + \cdots$$

$$P^{ab} \equiv \frac{1}{2} \left[\frac{g^{ab}}{2\pi} + i \left(\omega_{K}^{-1} \right)^{ab} \right]$$
(80)

While we do not explicitly display the matrix labels, it should be kept in mind that these are all 2×2 matrices and matrix products are assumed. Taking the trace and using (74) we get

$$\int dt \operatorname{Tr} \mathcal{A} = -\frac{1}{4\pi} \int \operatorname{Tr} \left[(a+A) d(a+A) + \frac{2}{3} (a+A)^3 \right]$$
(81)

where a is the connection for ω_K , i.e., $\omega_K = da$. The result is the integral of a Chern-Simons term, expanded around a as a background. It is useful to think of a + A as the connection of interest, in this case $SU(2)_L$ -valued. The dependence on the choice of \mathbb{CP}^1 with its ω_K as a background to expand around is irrelevant once we get to (81). The background is thus only an auxiliary step in arriving at this result. There are higher terms, with more derivatives and so on, which do retain the dependence on the metrical details of the background. These are negligible for slowly varying A_{μ} ; some of the terms will also cancel out when we include the $SU(2)_R$ sector.

The final result of simplifying (A6) would thus be

$$S = -\frac{1}{4\pi} \int \left[\operatorname{Tr} \left(A \, dA + \frac{2}{3} A^3 \right)_L - \operatorname{Tr} \left(A \, dA + \frac{2}{3} A^3 \right)_R \right] \tag{82}$$

The A's are connections for $SU(2)_L$ and $SU(2)_R$ and this result reproduces the Euclidean gravitational action (62), apart from an overall factor of (l/8 G). This overall degeneracy factor is important and, in this approach, has to come from several copies of the fields we have used.

We have used the action (A6) as it makes more transparent the role of the spatial components of A as part of how the large N limit is taken. We can also obtain the result (82) from the action (58) by integrating out the spinor fields. In the present case, we have $\mathcal{M} = \mathbb{CP}^1$ is two-dimensional. We take the fermions to be in the spin- $\frac{1}{2}$ representation of SU(2), i.e., the indices I, J takes values 1, 2. The fermions Ψ are coupled to the $SU(2)_L$ gauge fields while Φ couple to $SU(2)_R$ fields. We have a field theory of Dirac fermions in 2+1 dimensions and it is well known that the effective cation, upon integrating out the fermions, is the Chern-Simons action. Since the fields Φ are defined on \mathcal{M} with the orientation reversed, the Chern-Simons action generated by these fields will be the negative of the one generated by Ψ . The result, once again, is the action (82). In fact, being a field theory, the action (58) is much easier to use in practice as there are familiar diagrammatic techniques for evaluating the effective action.

We have considered the Euclidean signature so far. The spinor field version of the action, namely, (58), can be easily continued to Minkowski signature, with the gauge fields being valued in the Lie algebra of $SO(2, 1)_L$ and $SO(2, 1)_R$.

4 Discussion

A number of comments are in order at this point.

We have carried out the calculations in section 3 for Euclidean signature. While this is not relevant for the differential form, the trace in the Chern-Simons action over the gauge fields which are written as matrices corresponding to the spin- $\frac{1}{2}$ representation of $SU(2)_L$ and $SU(2)_R$ is sensitive to the signature. A continuation to Minkowski signature would involve SO(2,1) representations, which in the interest of unitarity, should be chosen as infinite dimensional. Defining the trace is then rather tricky. However, the formulation of TFD in terms of the fermionic field theory, as in (57), (58) avoids this problem. Taking the fields in (58) to be SO(2,1) spinors, we have a standard fermionic field theory and we expect that this will be consistent with unitarity.

There are many papers analyzing 2+1 dimensional gravity, starting with the Chern-Simons formulation and considering the evaluation of the partition function [31, 32, 33]. All these lead to strong hints of the underlying string origin of the action. This may very well be the case, but since we are obtaining the Chern-Simons action only in the large N limit, it is not clear how to compare our work with these developments. However, we may note the following. We have used a rank n representation of SU(2) with $n \to \infty$ to approximate the two-dimensional spatial manifold. In Minkowski signature, we should presumably use a representation of SO(2, 1). But we may also consider using a coadjoint orbit of the Virasoro group (which contains a suitable $SO(2, 1) \sim SL(2, \mathbb{R})$) to construct a noncommutative version of the two-manifold [34]. In that case, the analogue of the large N limit would be the limit of large central charge. Indeed this is the limit considered in [32] to compare 2+1 dimensional gravity and string theory. Further elaboration of this connection has to be left to future work at this stage.

Finally, we may note that the spectral action of Connes [10] is not the same as what we have, but there is some similarity. A key result for the spectral action is that the Wodzicki residue of the inverse square of the Dirac operator gives the Einstein action, see in particular Kastler's article in [10]. In our case, we are considering odd dimensional spacetimes. However, the action is basically related to the Dirac action given in (58). The main points of difference are that we have two sets of spinor fields which are naturally obtained in thermofield dynamics, and that they carry opposite chiralities in terms of coupling to gravitational degrees of freedom. Further, we are evaluating the effective action in a limit corresponding to $c \to \infty$ or we are taking the limit where only the lowest modes of the Dirac Hamiltonian of (58) is included. Whether this is related in some fashion to the Wodzicki residue is not clear.

Another point of clarification is the following. There are two ways to think of gravity on noncommutative spaces. We may consider the continuum description as an approximation, with a noncommutative operator version at short distances or at finite N. In this case, one has the full set of dynamical fields for gravity at all levels [10]-[13]. An alternative is to consider gravity as being trivial at the fundamental level only emerging in the continuum limit. This would be more in the spirit of gravity as an entropic phenomenon [17]-[19]. In our case, the spatial components of the gauge field were introduced in (79) to provide a simple realization of the gauge transformation property (77), with their values set to what is given by extremizing the continuum action. Thus there is no dynamics for these components if we stay at the level of the starting action (67) or (A6), making this approach closer to [17]-[19].

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Appendix: A generalization of (37)

The action for the functional integral for thermofield dynamics was given in (37) for a single quantum system. The fermion field version of that was given in (56) and generalized to a multipartite system with identical components in (57, 58). Here we want to consider the direct generalization of (37) and its relation to the Hall effect.

if we consider several systems which are distinct, the generalization is simple. We must have a unitary matrix U for each system and the action is simply a sum of actions of the form (37). The more interesting case is when we have identical subsystems; each subsystem may be a boson or a fermion. The latter is what is relevant for noncommutative geometry. In that case, we have a Hilbert space \mathcal{H} , say, of dimension N. States in this \mathcal{H} replace the notion of points on a manifold. All of \mathcal{H} must be included to account for the total volume of the manifold. For the purpose of time-evolution, it is then useful to think of the space as a state in the N-fold tensor product $\mathcal{H} \otimes \mathcal{H} \cdots \otimes \mathcal{H}$. To avoid double counting, the choice of state in each component \mathcal{H} must be distinct. The simplest way to implement this is to take the totally antisymmetrized state in $\mathcal{H} \otimes \mathcal{H} \cdots \otimes \mathcal{H}$, i.e., an N-fermion state.

An analogy with a quantum Hall system is very useful for this situation. Consider a quantum Hall system in the lowest Landau level with one-particle states $|\alpha\rangle$, $\alpha = 1, 2, \dots, N$. To be general, we assume that the fermion can have multiple internal states, such as spin degrees of freedom. (In the context of noncommutative geometry, this could correspond to degrees of freedom of matter.) We use labels I, J, etc. for the latter, so that states may be represented as $|k\rangle = |\alpha I\rangle$ which span a Hilbert space \mathcal{H} . Let M be the dimension of this space. We consider the dynamics of a droplet of n fermions starting out in one spin state, say, I = 0. The many-body Hilbert space will consist of suitably antisymmetrized states in the n-fold tensor product $\mathcal{H} \otimes \mathcal{H} \otimes \cdots \otimes \mathcal{H}$. Let \mathcal{N} denote the dimension of this Fock space, $\mathcal{N} = M!/((M-n)!n!)$. Let us say that we are considering the first n states, i.e., $|\alpha 0\rangle$, $\alpha = 1, 2, \dots, n$, to be filled. This means that if we write a fermionic field operator $\psi(x) = \sum_k a_k f_k(x)$, where $f_k(x)$ are the single particle wave functions, the droplet is given by the state $a_{10}^{\dagger} a_{20}^{\dagger} \cdots a_{n0}^{\dagger} |0\rangle$ or by the density matrix

$$\rho = a_{10}^{\dagger} a_{20}^{\dagger} \cdots a_{n0}^{\dagger} |0\rangle \langle 0|a_{n0} \cdots a_{20} a_{10}$$
(A1)

The possible transformations of this configuration, including time-evolution, are given by an element of $U(\mathcal{N})$. We have a single pure state chosen by ρ in (A1), so a path integral for the time-evolution of this state would have the action (7), with integration over all $U(t) \in U(\mathcal{N})$ modulo transformations which leave ρ invariant. The relevant action is thus

$$S = \int dt \left[\sum_{\{k\}} i \, (\mathcal{U}^{\dagger})_{10 \, 20 \, \cdots n0, k_1 k_2 \cdots k_n} (\dot{\mathcal{U}})_{k_1 k_2 \cdots k_n, 10 \, 20 \, \cdots n0} \right]$$

$$-\sum_{\{kl\}} (\mathcal{U}^{\dagger})_{10\,20\,\cdots n0, k_1k_2\cdots k_n} H_{k_1k_2\cdots k_n, l_1l_2\cdots l_n} (\mathcal{U})_{l_1l_2\cdots l_n, 10\,20\,\cdots n0} \Big]$$
(A2)

This is the general situation for the case with all sorts of many-body interactions. However, if the Hamiltonian and other observables of interest only involve singe-particle operators, i.e., they are of the form $C = a_k^{\dagger} C_{kl} a_l$, then the action simplifies. The possible transformations will correspond to unitary transformations of the form $U_{kl} \in U(M)$ acting on the single particle states, U being of the form e^{iC} . The states in the *n*-body Hilbert space may be viewed as corresponding to an irreducible representation of the group U(M) obtained as the totally antisymmetrized product of *n* copies of the fundamental representation of U(M). Thus we can simplify (7) by restricting to $U(M) \in U(\mathcal{N})$. In this case the elements of \mathcal{U} are given by the Slater determinant,

$$(\mathcal{U})_{k_1k_2\cdots k_n, 10\,20\cdots n0} = \frac{1}{\sqrt{n!}} \begin{vmatrix} U_{k_1\,10} & U_{k_1\,20} & \cdots & U_{k_1\,n0} \\ U_{k_2\,10} & U_{k_2\,20} & \cdots & U_{k_2\,n0} \\ \cdots & & & & \\ \end{vmatrix}$$
(A3)

It is easily seen that the action (A2) simplifies to

$$S = \int dt \operatorname{Tr} \left[P_{+} \left(i U^{\dagger} \dot{U} - U^{\dagger} H U \right) \right]$$

$$= \int dt \sum_{\alpha=1}^{n} \left[i \langle \alpha 0 | U^{\dagger} | k \rangle \langle k | \dot{U} | \alpha 0 \rangle - \langle \alpha 0 | U^{\dagger} | k \rangle H_{kl} \langle l | U | \alpha 0 \rangle \right]$$
(A4)

$$P_{+} = \sum_{\alpha=1}^{n} |\alpha 0\rangle \langle \alpha 0| \tag{A5}$$

Notice that P is not a density matrix; it is a representation, at the level of the one-particle Hilbert space, of the density matrix for the pure state (A1) of the droplet of n fermions. (It was this version of the action which was used in [22, 23].) In this expression, we do not yet have the part with the negative eigenvalues of P since the tilde system is not included. It is easy to see that the tilde system will have a similar path integral, with the eigenvalues of P being -1. The complete action for the thermofield dynamics of this system is then

$$S = \int dt \sum_{\alpha=1}^{n} \left[i \langle \alpha 0 | U^{\dagger} | k \rangle \langle k | \dot{U} | \alpha 0 \rangle - \langle \alpha 0 | U^{\dagger} | k \rangle H_{kl} \langle l | U | \alpha 0 \rangle \right] - \int dt \sum_{\alpha=1}^{n} \left[i \langle \alpha 0 | \tilde{U}^{\dagger} | k \rangle \langle k | \dot{\tilde{U}} | \alpha 0 \rangle - \langle \alpha 0 | \tilde{U}^{\dagger} | k \rangle H_{kl} \langle l | \tilde{U} | \alpha 0 \rangle \right]$$
(A6)

When applying this to noncommutative geometry, we should take n = N.

Here is a curiosity in this formulation: In the path integral, we also have the functional integration over $\mathbb{CP}^{\mathcal{N}-1}$; this corresponds to the integration over $(\mathcal{U})_{k_1k_2\cdots k_n,10\ 20\cdots n0}$ at each instant of time. In restricting the action to U(M) transformations with P as in (A4), a part of this integration becomes trivial. The left over integration is over the Grassmannian

space $U(M)/U(n) \times U(M-n)$. This gives a factor of V at each instant of time, where V is defined by

$$\int d\mu(\mathbb{CP}^{\mathcal{N}-1}) = V \int d\mu(U(M)/U(n) \times U(M-n))$$
(A7)

Exponentiating these factors we get an additional term in S which is

$$\Delta S = \int \frac{dt}{\epsilon} \log V \tag{A8}$$

Here we are considering dividing the interval of time-integration into segments of length ϵ , with $\epsilon \to 0$ eventually. This extra factor may be thought of as an additional entropy factor arising from the fact that we are restricting the observables to a smaller set, namely, to those of the one-particle type. (It is also similar to the terms with $\delta(0)$ -factor that one encounters in field theories when making a change of field variables upon exponentiating the Jacobian of the transformation.)

After this discussion, it should be clear that (A6), apart from the extra term (A8), is indeed the required generalization once the tilde system is included.

This generalization to several systems can also be phrased as a contour integral, with the action

$$S_C = i \oint_C dt \sum_{\alpha=1}^n \left(U^{\dagger} \dot{U} - U^{\dagger} \mathcal{A} U \right)_{\alpha 0 \ \alpha 0}$$
(A9)

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