

CHCRUS

This is the accepted manuscript made available via CHORUS. The article has been published as:

N=3 solution in dyonic ISO(7) gauged maximal supergravity and its uplift to massive type IIA supergravity Yi Pang and Junchen Rong Phys. Rev. D **92**, 085037 — Published 22 October 2015 DOI: 10.1103/PhysRevD.92.085037

MI-TH-1531

An $\mathcal{N} = 3$ Solution in Dyonic ISO(7) Gauged Maximal Supergravity and Its Uplift to Massive Type IIA

Yi Pang¹ and Junchen Rong¹

¹ George P. & Cynthia Woods Mitchell Institute for Fundamental Physics and Astronomy, Texas A&M University, College Station, TX 77843, USA

ABSTRACT

We consider a certain $\mathcal{N} = 1$ supersymmetric, $SO(3) \times SO(3)$ invariant, subsector of the dyonic ISO(7)-gauged maximal supergravity in four-dimensions. The theory contains two scalar fields and two pseudoscalar fields. We look for stationary points of the scalar potential, especially the one preserving $\mathcal{N} = 3$ supersymmetry of the original ISO(7) gauged theory. The $\mathcal{N} = 3$ stationary point corresponding to the AdS vacuum in the D = 4 theory is lifted to a warped $AdS_4 \times X_6$ type solution in massive type IIA supergravity. This D = 10 background should be the dual of a certain $\mathcal{N} = 3$ Chern-Simons matter theory in three dimensions.

1 Introduction

The four-dimensional SO(8) gauged maximally supersymmetric $\mathcal{N} = 8$ supergravity was widely considered to be a unique theory, since its construction thirty years ago [1]. Interestingly, using embedding tensor formulation [2], it was recently discovered that the SO(8) gauging can be realized in a mixed electric/magnetic frame compatible with the $\mathcal{N} = 8$ supersymmetry [3]. A new parameter commonly called ω was introduced to parameterize the choice of the electric/magnetic frame. Inequivalent $\mathcal{N} = 8$ theories are defined by values of ω in the interval $0 \leq \omega \leq \pi/8$ [3–5]. This development has stimulated numerous studies, including its critical points [3,7–14] and the consequences of the deformation for the holographic dual theory [14–18]. Through the work of [4, 19, 20], it is proven that the purely electric SO(8) gauged theory can be obtained from M-theory via a consistent reduction on S^7 . However, a no-go theorem [21] shows that the ω -deformed SO(8) gauged supergravities cannot be realized via a compactification that is locally described by ten or eleven dimensional supergravity. This means we are not able to determine the dual CFTs by applying the standard AdS/CFT correspondence, whose formulation is usually facilitated by the existence of a weakly coupled local supergravity theory on the bulk side ¹.

Besides SO(8) gauged theory, there exist also 4D maximal supergravities with other subgroups of SL(8) being gauged. In this paper, we will consider the ISO(7) gauged theory². The ISO(7) gauging can be derived from the SO(8) gauging via an Inönü-Wigner contraction [27]. Similar to the SO(8) case, the ISO(7) group can be dyonically gauged, giving rise to two inequivalent theories: the theory in the purely electric frame and the theory in a mixed electric/magnetic frame [5]. The former is known to be derivable from M-theory by a consistent reduction on $S^6 \times T^1$, or from type IIA supergravity via a consistent S^6 reduction [28], the latter can be embedded to massive type IIA, with the Romans mass mbeing identified as the parameter of the magnetic gauging upon reduction [29]. By virtue of the consistency of the reduction, solutions of the ISO(7) gauged theory can be lifted to solutions in (massive) type IIA. Some previous work on supersymmetric $AdS_4 \times X_6$ type solutions in (massive) type IIA supergravity can be found in [29,32–39]. All these solutions preserve at most eight supercharges, or in other words, $\mathcal{N} = 2$ supersymmetry in four

¹It is possible that the ω -deformed SO(8) gauged theories can still be embedded into string theory via a framework more general than supergravity, for example, double field theory with the strong section constraint being relaxed, following the idea of [22–24].

²Recent work studying maximal supergravities with other dyonic non-compact gauging can be seen in [25, 26].

dimensions.

In this paper we present explicitly, the first analytic $AdS_4 \times X_6$ solution in massive type IIA preserving $\mathcal{N} = 3$ supersymmetry in four dimensions. The D = 10 solution is obtained by lifting the $\mathcal{N} = 3 \ AdS_4$ solution of the D = 4 dyonic ISO(7) gauged supergravity. It was shown in [13] that the critical points in the ISO(7) gauged theory can have at most $\mathcal{N} = 3$ supersymmetry, and in addition to the SO(3) R-symmetry the $\mathcal{N} = 3$ point preserves an extra SO(3) symmetry. We shall denote the full symmetry group by $SO(3)_R \times SO(3)_D$ ³. This subgroup is characterized by starting from $SO(3)_1 \times SO(3)_2 \times SO(3)_3 \times SO(3)_4 \in SO(8)$. The factor $SO(3)_D$ is then the diagonal in $SO(3)_1 \times SO(3)_2 \times SO(3)_3$, and the factor $SO(3)_R$ is SO(3)₄⁴. In fact, the $\mathcal{N} = 3$ critical point can be captured by a subsector of the full theory which is invariant under the $SO(3)_R \times SO(3)_D$ subgroup of ISO(7) [14, 40]. To obtain the $\mathcal{N} = 3$ critical point, we truncate the ISO(7) gauged maximal supergravity to its scalar subsector invariant under $SO(3)_R \times SO(3)_D$. The truncated theory preserves $\mathcal{N} = 1$ supersymmetry, and it encompasses the scalar sectors invariant under SO(7) and G_2 as special cases. The $\mathcal{N} = 3$ critical point appears to be a non-supersymmetric solution of the truncated theory, however preserves $\mathcal{N} = 3$ supersymmetry of the original $\mathcal{N} = 8$ theory. Besides the $\mathcal{N} = 3$ critical point, we also found a single $\mathcal{N} = 1$, G₂-invariant point and three $\mathcal{N} = 0$ critical points, of which the G₂-and SO(3)_R × SO(3)_D-invariant points are stable, while the SO(7) invariant point is unstable. The G_2 and SO(7)-invariant point is previously known in [8], where the G_2 and SO(7)-invariant points in D = 4 maximal supergravities with all gaugings are listed. We then uplift the $\mathcal{N} = 3$ critical point as well as the G₂-invariant, $\mathcal{N} = 1,0$ critical points to massive type IIA supergravity using the uplift formulae given in [29].

This paper is organized as follows. In section 2, we briefly review the ingredients of the dyonic ISO(7) gauged supergravity. In section 3, we truncate the ISO(7) gauged supergravity to the scalar subsector invariant under the SO(3)_R × SO(3)_D, list the critical points in this sector, and study the mass spectra of the scalar fluctuations around these AdS vacua. In section 4, we uplift the D = 4 critical points to D = 10 and obtain the first $\mathcal{N} = 3$ solution in massive type IIA explicitly. We discuss possible CFT dual of the $\mathcal{N} = 3$ solution

³The notation $SO(3)_R$ should not be confused with the SO(3) R-symmetry, which is in fact associated with $SO(3)_D$.

⁴The SO(3)_R×SO(3)_D subgroup used here is different from the commonly used SO(3)×SO(3) subgroup, in which the first SO(3) refers to the diagonal in SO(3)₁×SO(3)₂ and the second SO(3) is the diagonal in SO(3)₃×SO(3)₄. Stationary points of the SO(8) and SO(4,4) gauged maximal D = 4 supergravities which preserve the commonly used SO(3)×SO(3) subgroup have been uplifted to D = 11 supergravity in [30,31].

and conclude in section 5.

Note Added: While we were preparing this draft, the work [40] appeared in arXiv and it seems to overlap with our paper on the critical points and their spectra analysis in the $SO(3)_D \times SO(3)_R$ -invariant sector.

2 Dyonic ISO(7) gauged $\mathcal{N} = 8$ supergravity

The D = 4 gauged maximal supergravity is characterized by the embedding tensor $\Theta_M{}^{\alpha}$ which completely specifies the gauging [2]. Embedding tensor enters the Lagrangian through the quantity

$$X_M = \Theta_M{}^\alpha t_\alpha, \tag{2.1}$$

which satisfies

$$[X_M, X_N] = -X_{MN}{}^P X_P, (2.2)$$

implying a closed algebra. Indices $M, N \dots$ transform as the **56** of $E_{7(7)}$ which decomposes into $\mathbf{28} \oplus \mathbf{\overline{28}}$ of SU(8) or $\mathbf{28} \oplus \mathbf{28'}$ of SL(8). The decomposition of the **56** of $E_{7(7)}$ under SU(8) or SL(8) suggests two different bases in which $E_{7(7)}$ covariant quantities can be formulated. In the SL(8) basis, $V_M = \{V_{[AB]}, V^{[AB]}\}$, and the pure scalar sector of the 4D gauged maximal supergravity is given as

$$e^{-1}\mathcal{L} = \frac{1}{8} \operatorname{Tr}(\partial_{\mu}\mathcal{M}\partial^{\mu}\mathcal{M}^{-1}) - \frac{1}{672} (X_{MN}{}^{R}X_{PQ}{}^{S}\mathcal{M}^{MP}\mathcal{M}^{NQ}\mathcal{M}_{RS} + 7X_{MN}{}^{Q}X_{PQ}{}^{N}\mathcal{M}^{MP}),$$
(2.3)

where \mathcal{M}^{MN} is the inverse of \mathcal{M}_{MN} . The latter is constructed from the bilinear of the 56-bein

$$\mathcal{M}_{MN} = (LL^{\dagger})_{MN}, \quad L(\phi)_M \underline{}^N = S_M^{\dagger} \underline{}^P \hat{L}(\phi) \underline{}_P \underline{}^N, \tag{2.4}$$

with

$$S_{\underline{M}}{}^{N} = \frac{1}{4\sqrt{2}} \begin{pmatrix} \Gamma_{ij}{}^{AB} & \mathrm{i}\Gamma_{ijAB} \\ \Gamma^{ijAB} & -\mathrm{i}\Gamma^{ij}{}_{AB} \end{pmatrix}, \quad \hat{L}_{\underline{M}}{}^{\underline{N}} = \exp \begin{pmatrix} 0 & \phi_{ijkl} \\ \phi^{ijkl} & 0 \end{pmatrix}.$$
(2.5)

The indices $\underline{M}, \underline{N}$ label the **56** irreps realized in the SU(8) basis, and therefore the unitary matrix S^{\dagger} converts the SU(8) basis to the SL(8) basis [7]. Indices A, B, \ldots are associated with the fundamental representation of SL(8) while i, j, \ldots transform as the **8** of SU(8). The 2-gamma matrices Γ_{ij}^{AB} comprise the generators of SO(8) in the chiral spinor representation. The position of the index is not crucial. Here, we use the same notation as [14] where the convention of the 2-gamma matrices is given explicitly⁵.

⁵Slightly different from [14], here we interchanged the definition of Γ_2 and Γ_7 .

Dyonic ISO(7) gauged maximal supergravity arises when the electric and magnetic parts of the embedding tensor are both non-vanishing. Specifically, X_{MN}^{P} takes the form

$$X_{ABM}{}^{N} = \begin{pmatrix} -f_{ABCD}{}^{EF} & \\ & f_{ABEF}{}^{CD} \end{pmatrix}, \quad X^{AB}{}_{M}{}^{N} = \begin{pmatrix} -g^{AB}{}_{CD}{}^{EF} & \\ & g^{AB}{}_{EF}{}^{CD} \end{pmatrix},$$
(2.6)

where $f_{ABCD}{}^{EF} = 2\sqrt{2}\delta^{[E}_{[A}\theta_{B][C}\delta^{F]}_{D]}$ and $g^{AB}{}_{EF}{}^{CD} = 2\sqrt{2}\delta^{[A}_{[E}\xi^{B][C}\delta^{D]}_{F]}$ in which $\theta = q \cdot \operatorname{diag}(\mathbb{I}_{7}, 0), \quad \xi = m \cdot \operatorname{diag}(0_{7}, 1).$

The inequivalent theories correspond to m = 0 or $m \neq 0$. The reason [5] is that one can use the SL(8) Cartan generator

$$\Lambda_{\rm red} = \begin{pmatrix} \mathbb{I}_7 & \\ & -7 \end{pmatrix}, \tag{2.8}$$

(2.7)

to rescale the electric and magnetic coupling constant (g, m) separately, and the effects of Λ_{red} on the scalar matrix \mathcal{M}_{MN} can be absorbed by a non-linear field redefinition.

3 Truncation to the $SO(3)_R \times SO(3)_D$ invariant sector

In this section, we truncate the $ISO(7) = SO(7) \times \mathbb{R}^7$ gauged theory to its subsector invariant under an $SO(3)_R \times SO(3)_D$ group, which is embedded in ISO(7) by the following series

$$ISO(7) \supset SO(7) \supset SO(3)_R \times SO(3)_L \times SO(3) \supset SO(3)_R \times \left[SO(3)_L \times SO(3)\right]_D$$
(3.1)

Here $[SO(3)_L \times SO(3)]_D$ means the diagonal subgroup of $SO(3)_L \times SO(3)$, which we denote by $SO(3)_D$ in later discussion for brevity. The scalar coset invariant under the $SO(3)_R \times$ $SO(3)_D$ subgroup can be parameterized by the following SU(8) complex self-dual 4-form⁶

$$\phi_{ijkl} = (\phi_1 \cos \sigma_1 \Phi_1 + \phi_2 \cos \sigma_2 \Phi_2)_{AB} (\Gamma^{ijkl})_{AB}$$
$$+ i(\phi_1 \sin \sigma_1 \Phi_1 + \phi_2 \sin \sigma_2 \Phi_2)_{\dot{A}\dot{B}} (\Gamma^{ijkl})_{\dot{A}\dot{B}}, \qquad (3.2)$$

where

$$\Phi_1 = \frac{1}{2} \operatorname{diag}(-1, -1, -1, -1, 1, 1, 1, 1), \quad \Phi_2 = \operatorname{diag}(0, 0, 0, 0, -1, -1, -1, 3), \quad (3.3)$$

⁶One way of understanding this formula is to treat $SO(3)_R \times SO(3)_D$ as a subgroup of SO(8). In the of notation the gamma matrices, $\{i, A, \dot{A}\}$ label the indices of $\{8_s, 8_v, 8_c\}$ representations respectively. Φ_{1AB} and Φ_{2AB} break SO(8) into $[SO(3) \times SO(4)]_v$, while $\Phi_{1\dot{A}\dot{B}}$ and $\Phi_{2\dot{A}\dot{B}}$ break SO(8) into $[SO(3) \times SO(4)]_c$. When Φ_{1AB} , Φ_{2AB} , $\Phi_{1\dot{A}\dot{B}}$, $\Phi_{2\dot{A}\dot{B}}$ are present at the same time, SO(8) is broken into the SO(3)_R \times SO(3)_D subgroup, which is the common subgroup of $[SO(3) \times SO(4)]_v$ and $[SO(3) \times SO(4)]_c$.

and $(\Gamma^{ijkl})_{AB}$, $(\Gamma^{ijkl})_{\dot{A}\dot{B}}$ are the upper and lower 8 by 8 diagonal blocks of the 16 by 16 4-gamma matrix respectively

$$\Gamma^{ijkl} = \begin{pmatrix} (\Gamma^{ijkl})_{AB} \\ (\Gamma^{ijkl})_{\dot{A}\dot{B}} \end{pmatrix}.$$
(3.4)

Substitute the scalar 4-form ansatz into the Lagrangian (2.3), we obtain the scalar potential for the $SO(3)_D \times SO(3)_R$ invariant sector of the dyonic ISO(7) gauged maximal supergravity as

$$V = \frac{1}{64}g^{2}(\cosh\phi_{2} - \cos\sigma_{2}\sinh\phi_{2})^{2}(\cos\sigma_{2}\sinh\phi_{2} + \cosh\phi_{2}) \\\times \left(196\cos(\sigma_{1} + 3\sigma_{2})\sinh\phi_{1}\sinh^{3}\phi_{2} + 4\cos(\sigma_{1} - 3\sigma_{2})\sinh\phi_{1}\sinh^{3}\phi_{2} \\+ 35\cos(\sigma_{1} - \sigma_{2})\sinh\phi_{1}\sinh\phi_{2} + 467\cos(\sigma_{1} + \sigma_{2})\sinh\phi_{1}\sinh\phi_{2} \\+ 47\cos(\sigma_{1} - \sigma_{2})\sinh\phi_{1}\sinh^{3}\phi_{2} - 97\cos(\sigma_{1} + \sigma_{2})\sinh\phi_{1}\sinh^{3}\phi_{2} \\+ 4\cos\sigma_{1}\sinh\phi_{1}\cosh\phi_{2}(28\cos2\sigma_{2}\sinh^{2}\phi_{2} + 47) \\- 2\cosh\phi_{1}(28\cos3\sigma_{2}\sinh^{3}\phi_{2} + 101\cos\sigma_{2}\sinh\phi_{2} - 7\cos\sigma_{2}\sinh^{3}\phi_{2} \\+ \cosh\phi_{2}(200\cos2\sigma_{2}\sinh^{2}\phi_{2} + 226) - 50\cosh^{3}\phi_{2}) \\- 28\cos\sigma_{1}\sinh\phi_{1}\cosh^{3}\phi_{2} - 768\right) \\+ gm\sin^{2}\sigma_{2}\sinh^{2}\phi_{2}(\cos\sigma_{2}\sinh\phi_{2} - \cosh\phi_{2})^{3} \times \left(3\sin\sigma_{1}\sinh\phi_{1}\cosh\phi_{2} \\+ \sinh\phi_{2}(4\sin\sigma_{2}\cosh\phi_{1} - \sinh\phi_{1}(3\sin\sigma_{1}\cos\sigma_{2} + 4\sin\sigma_{2}\cos\sigma_{1}))\right) \\+ \frac{1}{2}m^{2}(\cosh\phi_{1} - \cos\sigma_{1}\sinh\phi_{1})(\cosh\phi_{2} - \cos\sigma_{2}\sinh\phi_{2})^{6}.$$
(3.5)

So far we have been focused only on the scalar sector. In fact, the full fledged $SO(3)_D \times SO(3)_R$ invariant sector preserves $\mathcal{N} = 1$ supersymmetry which implies the potential can be expressed in terms of a superpotential W, with

$$V = 2\left(4\left|\frac{\partial W}{\partial \phi_1}\right|^2 + \frac{2}{3}\left|\frac{\partial W}{\partial \phi_2}\right|^2 - 3|W|^2\right),\tag{3.6}$$

where

$$W = \frac{g}{4} (\cosh\frac{\phi_2}{2} + e^{i\sigma_2}\sinh\frac{\phi_2}{2})^2 (\cosh\frac{\phi_2}{2} - e^{i\sigma_2}\sinh\frac{\phi_2}{2})^3 \\ \times \left(\cosh\frac{\phi_2}{2} (7\cosh\frac{\phi_1}{2} - e^{i\sigma_1}\sinh\frac{\phi_1}{2}) - e^{i\sigma_2}(\cosh\frac{\phi_1}{2} - 7e^{i\sigma_1}\sinh\frac{\phi_1}{2})\sinh\frac{\phi_2}{2}\right) \\ + \frac{im}{4} (\cosh\frac{\phi_1}{2} - e^{i\sigma_1}\sinh\frac{\phi_1}{2})(\cosh\frac{\phi_2}{2} - e^{i\sigma_2}\sinh\frac{\phi_2}{2})^6.$$
(3.7)

The location of the critical point depends on m/g, however, as mentioned before, the effects due to different m/g can be compensated by a nonlinear field redefinition. Therefore we can choose this ratio to be the one most convenient for our purpose. In finding the locations of the critical points, we choose

$$\frac{m}{q} = 2. \tag{3.8}$$

The scalar potential (3.5) possesses an $SO(3)_R \times SO(3)_D$ invariant stationary point preserving $\mathcal{N} = 3$ supersymmetry of the original $\mathcal{N} = 8$ theory which lies outside the residual $\mathcal{N} = 1$ supersymmetry of the truncated theory. In terms of the complexified fields

$$\xi_1 = \tanh \phi_1 e^{\mathbf{i}\sigma_1}, \quad \xi_2 = \tanh \phi_2 e^{\mathbf{i}\sigma_2}, \tag{3.9}$$

the $\mathcal{N} = 3$ point is given by

$$\xi_1 = \frac{3}{5} - \frac{2i}{5}, \quad \xi_2 = \frac{i}{2},$$
 (3.10)

The mass spectrum of the fluctuations around this vacuum has been obtained previously in [13] by a group theoretic method without referring to the detailed position of the critical point.

$$m^{2}L_{0}^{2}: \quad 1 \times (3(1+\sqrt{3})); \quad 6 \times (1+\sqrt{3}); \quad 1 \times (3(1-\sqrt{3})); \quad 6 \times (1-\sqrt{3}); \\ 4 \times (-\frac{9}{4}); \quad 18 \times (-2); \quad 12 \times (-\frac{5}{4}); \quad 22 \times 0,$$
(3.11)

where the AdS radius squared $L_0^2 = -6/V$, $V(g = 1, m = 2) = -32/\sqrt{3}$. (The integer to the left of the multiplication sign indicates the degeneracy of the mass eigenvalue, while the number to the right indicates the corresponding mass-squared.) There is another $\mathcal{N} = 1$ critical point with G₂ global symmetry

$$\xi_1 = \xi_2 = -\frac{\mathbf{i}}{4}.\tag{3.12}$$

To obtain this critical point, we have combined the Newton-Raphson method with the "Inverse Symbolic Calculator" technique. The mass spectrum of the scalar fluctuations around this vacuum is given by

$$m^2 L_0^2$$
: 1 × (4 ± $\sqrt{6}$); 14 × 0; 27 × ($-\frac{1}{6}(11 - \sqrt{6})$), (3.13)

where the AdS radius squared $L_0^2 = -6/V$, $V(g = 1, m = 2) = -512\sqrt{3}/(25\sqrt{5})$. Besides the supersymmetric critical points, there are three more non-supersymmetric critical points, preserving SO(3)_R × SO(3)_D, G₂ and SO(7)_v global symmetries respectively. The SO(3)_R × SO(3)_D invariant critical point is located at

$$\xi_1 = 0.353669 + 0.0552267i, \quad \xi_2 = -0.0293804 + 0.534729i,$$
 (3.14)

with the mass spectrum given by

$$m^{2}L_{0}^{2}: 1 \times (6.72740); 1 \times (5.28662); 4 \times (-1.96422); 9 \times (-1.75110); 9 \times (-1.58816); 1 \times (-1.58552); 8 \times (-1.17591); 5 \times (-0.98271); 4 \times (-0.72962); 5 \times (0.62977); 1 \times (0.58436); 22 \times 0,$$
(3.15)

where the AdS radius squared $L_0^2 = -6/V$, V(g = 1, m = 2) = -18.662034. The G₂ invariant point is given by

$$\xi_1 = \xi_2 = \frac{\mathbf{i}}{2}.\tag{3.16}$$

The associated mass spectrum has the simple structure

$$m^2 L_0^2: \quad 2 \times 6; \quad 14 \times 0; \quad 54 \times (-1),$$
 (3.17)

where the AdS radius squared $L_0^2 = -6/V$, $V(g = 1, m = 2) = -32/\sqrt{3}$. It should be emphasized that the spectra associated with the non-supersymmetric $SO(3)_R \times SO(3)_D$ and G₂-invariant critical points lie above the Breitenlohner-Freedman (BF) bound. The last $SO(7)_v$ -invariant critical point is located at

$$\phi_1 = \phi_2 = -\frac{1}{6}\log\frac{5}{4}, \quad \sigma_1 = \sigma_2 = 0.$$
 (3.18)

The mass spectrum of the $SO(7)_v$ -invariant critical point reads

$$m^2 L_0^2$$
: 1×6; 7×0; 35×($-\frac{6}{5}$); 27×($-\frac{12}{5}$), (3.19)

where the AdS radius squared $L_0^2 = -6/V$, $V(g = 1, m = 2) = -15 \times 5^{\frac{1}{6}}/2^{\frac{1}{3}}$. This point is unstable against fluctuations as the mass squared of some scalar modes is below the BF bound. The scalar mass spectra associated with the two G₂-invariant critical points and the single SO(7)_v-invariant point coincide with those given in [8], where the G₂-and SO(7)_vinvariant critical points in D = 4 maximal supergravities with all gaugings are analyzed.

Since different choices of m/g are related by highly nonlinear field redefinition, we expect that if m/g had been different from 2, the expressions of the critical points will not be as neat as the ones given above.

4 Lifting the 4D critical points to 10D massive type IIA

Previously, in order to find the exact positions of the critical points, we have set m = 2gfor convenience. Thus a direct application of the uplift formulae will lead to solutions in massive type IIA supergravity with m = 2g. To recover the general *m*-dependence of the solution, we utilize two scaling symmetries of the equations of motion of the massive type IIA supergravity. In the type IIA Einstein frame conventions of [42], these scalings act on the fields and the parameter of the theory as follows

$$\hat{A}_{(1)} \to \tau \hat{A}_{(1)}, \quad \hat{A}_{(2)} \to \tau \hat{A}_{(2)}, \quad \hat{A}_{(3)} \to \tau^2 \hat{A}_{(3)}, \quad d\hat{s}_{10}^2 \to \tau^{\frac{5}{4}} d\hat{s}_{10}^2, \quad e^{\hat{\phi}} \to \tau^{-\frac{1}{2}} e^{\hat{\phi}}, \\ \hat{A}_{(1)} \to \kappa \hat{A}_{(1)}, \quad \hat{A}_{(2)} \to \kappa^2 \hat{A}_{(2)}, \quad \hat{A}_{(3)} \to \kappa^3 \hat{A}_{(3)}, \quad d\hat{s}_{10}^2 \to \kappa^2 d\hat{s}_{10}^2, \quad m \to \frac{m}{\kappa}.$$
(4.1)

Notice that the second scaling is merely based on the dimensionality. The scaling symmetry of the 4D theory (2.8) reflects itself in the 10D theory as a combination of the above two scalings with $\tau = \lambda^{20}$, $\kappa = \lambda^{-14}$, where λ is identified as the parameter of $\Lambda_{\rm red}$ (2.8).

4.1 Supersymmetric $SO(3)_R \times SO(3)_L$ invariant solution in massive IIA

In terms of the auxiliary coordinates on S^6

$$\mu^{1} = \sin \xi \cos \theta_{1} \cos \chi_{1}, \quad \mu^{2} = \sin \xi \cos \theta_{1} \sin \chi_{1},$$

$$\mu^{3} = \sin \xi \sin \theta_{1} \cos \psi, \quad \mu^{4} = \sin \xi \sin \theta_{1} \sin \psi,$$

$$\nu^{1} = \cos \xi \cos \theta_{2}, \quad \nu^{2} = \cos \xi \sin \theta_{2} \cos \chi_{2},$$

$$\nu^{3} = \cos \xi \sin \theta_{2} \sin \chi_{2},$$

(4.2)

which satisfy $\sum_{A=1}^{4} \mu^{A} \mu^{A} + \sum_{i=1}^{3} \nu^{i} \nu^{i} = 1$, the metric on the round S^{6} takes the form

$$ds_{S^6}^2 = d\xi^2 + \sin^2 \xi d\Omega_3^2 + \cos^2 \xi d\Omega_2^2$$

= $d\xi^2 + \sin^2 \xi \left(d\theta_1^2 + \cos^2 \theta_1 d\chi_1^2 + \sin^2 \theta_1 d\psi^2 \right) + \cos^2 \xi (d\theta_2^2 + \cos^2 \theta_2 d\chi_2^2).$ (4.3)

To lift the solution of the 4D dyonic ISO(7) gauged supergravity to that in the 10D massive type IIA supergravity, we utilize the uplift formulae given in [29], in which the internal components of the 10D metric, the dilaton, and various form fields are constructed in terms of the SL(7)-covariant blocks of the D = 4 scalar matrix \mathcal{M}_{MN}

$$g^{mn} = \frac{1}{4}g^2 \Delta K_{IJ}^m K_{KL}^n \mathcal{M}^{IJ,KL},$$

$$e^{-\frac{3}{2}\hat{\phi}} = -g^{mn}\hat{A}_m\hat{A}_n + \Delta x_I x_J \mathcal{M}^{I8J8},$$

$$\hat{A}_m = \frac{1}{2}g \Delta g_{mn} K_{IJ}^n x_K \mathcal{M}^{IJK8},$$

$$\hat{A}_{mn} = -\frac{1}{2}\Delta g_{pm} K_{IJ}^p \partial_n x^K \mathcal{M}^{IJ}_{K8},$$

$$\hat{A}_{mnp} = \hat{A}_m \hat{A}_{np} + \frac{1}{8}g \Delta g_{mq} K_{IJ}^q K_{np}^{KL} \mathcal{M}^{IJ}_{KL},$$
(4.4)

where $K_{IJ}^m = 2g^{-2}\mathring{g}^{mn}x_{[I}\partial_n x_{J]}, K_{mn}^{IJ} = 4g^{-2}\partial_{[m}x_I\partial_{n]}x_J$. Due to our gamma matrix notation, $\{x^I\}$ are related to $\{\mu^A, \nu^i\}$ by the similarity transformation

$$S = \operatorname{diag}(1, 1, 1, 1, -1, -1, 1), \tag{4.5}$$

which brings the SL(7)-covariant blocks of the D = 4 scalar matrix \mathcal{M}_{MN} into a form invariant under the standard SO(3)_R × SO(3)_D transformation given in [14].

The 10D solution corresponding to the $\mathcal{N} = 3$ critical point is then obtained as follows

$$L^{-2}d\hat{s}_{10}^2 = \Delta^{-1}\left(\frac{3\sqrt{3}}{16}ds_{AdS_4}^2\right) + g_{mn}dy^m dy^n, \tag{4.6}$$

where

$$\Delta = 3^{\frac{9}{8}} 2^{-\frac{3}{4}} (\cos 2\xi + 3)^{-\frac{1}{8}} \Xi^{-\frac{1}{4}}, \qquad \Xi = (24\cos 2\xi + 3\cos 4\xi + 37), \tag{4.7}$$

and the internal metric on the deformed S^6 is given as

$$g_{mn}dy^{m}dy^{n} = \frac{3\sqrt{3}}{4}(\Delta\Xi)^{-1} \left[-\sin^{2}2\xi d\xi^{2} + 8(\cos 2\xi + 3)d\mu \cdot d\mu + 4(\cos 2\xi + 3)d\nu \cdot d\nu + 16\mu^{A}\eta^{i}_{AB}d\mu^{B}\epsilon^{ijk}\nu^{j}d\nu^{k} - \frac{16}{\cos 2\xi + 3}(d\mu^{A}\eta^{i}_{AB}\mu^{B}\nu^{i})^{2} \right], (4.8)$$

where η^i 's are the generators of SO(3)_L embedded in SO(4) \simeq SO(3)_R × SO(3)_L,

$$\eta^{1} = \begin{pmatrix} 0 & 0 & 0 & -1 \\ 0 & 0 & -1 & 0 \\ 0 & 1 & 0 & 0 \\ 1 & 0 & 0 & 0 \end{pmatrix}, \quad \eta^{2} = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix}, \quad \eta^{3} = \begin{pmatrix} 0 & -1 & 0 & 0 \\ 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & -1 \\ 0 & 0 & 1 & 0 \end{pmatrix}.$$
(4.9)

Denote $\mathcal{K}^i \equiv \mu^A \eta^i_{AB} d\mu^B$, various p-form fields can then be expressed as

$$L^{-1}e^{\frac{3}{4}\phi_0}\hat{A}_{(1)} = \frac{2}{\cos 2\xi + 3}\nu^i \mathcal{K}^i, \tag{4.10}$$

$$L^{-2}e^{-\frac{1}{2}\phi_{0}}\hat{A}_{(2)} = -\Xi^{-1} \bigg[-8d\nu^{i} \wedge \mathcal{K}^{i} + 6\sin 2\xi d\xi \wedge \nu^{i}\mathcal{K}^{i} + 2(3\cos 2\xi + 5)\nu^{i}\eta^{i}_{AB}d\mu^{A} \wedge d\mu^{B} - (3\cos 2\xi + 5)\epsilon^{ijk}\nu^{i}d\nu^{j} \wedge d\nu^{k} \bigg],$$

$$(4.11)$$

$$L^{-3}e^{\frac{1}{4}\phi_0}\hat{A}_{(3)} = -\Xi^{-1} \bigg[6\sin 2\xi \epsilon^{ijk} d\xi \wedge \mathcal{K}^i \nu^j \wedge d\nu^k + 2(3\cos 2\xi + 5)\epsilon^{ijk} \nu^i d\nu^j \wedge d\mu^A \wedge \eta^k_{AB} d\mu^B + 2(3\cos 2\xi + 5)\epsilon^{ijk} \nu^i d\nu^j \wedge d\mu^A \wedge \eta^k_{AB} d\mu^B + 2\epsilon^{ijk} \partial_{\mu} \partial_$$

$$+4\epsilon^{ijk}\mathcal{K}^{i}\wedge d\nu^{j}\wedge d\nu^{k} + \frac{8}{3}\csc^{2}\xi\epsilon^{ijk}\mathcal{K}^{i}\wedge\mathcal{K}^{j}\wedge\mathcal{K}^{k}\right] + \frac{3\sqrt{3}}{8}\Omega_{(3)}.$$
 (4.12)

where $d\Omega_{(3)} = \text{vol}(AdS_4)$ which is the volume element of the "unit" AdS_4 . Finally, the 10D dilaton is given by

$$e^{-\frac{3}{2}\hat{\phi}} = e^{-\frac{3}{2}\phi_0} \frac{\Delta\Xi}{3\sqrt{3}(\cos 2\xi + 3)}.$$
(4.13)

Notice that everything is written in terms of $SO(3)_R \times SO(3)_D$ invariant quantities (any function of ξ is invariant as $\mu \cdot \mu = \sin^2 \xi$ is an invariant quantity). Here we have introduced two constants $L^2 = 2^{-\frac{1}{12}}g^{-25/12}m^{1/12}$ and $e^{\phi_0} = 2^{\frac{5}{6}}g^{\frac{5}{6}}m^{-\frac{5}{6}}$ using the scaling symmetries (4.1). We have checked that our solution satisfies all the equations of motion of massive type IIA supergravity in the convention of [42].

4.2 G₂-invariant solutions in massive type IIA

In our notation, we can write down the almost complex structure on unit S^6 as

$$J_{(2)} = \frac{1}{2!} J_{mn} dy^m \wedge dy^n = \mathcal{K}^i \wedge d\nu^i + \frac{1}{2} \nu^i \eta^i_{AB} d\mu^A \wedge d\mu^B + \frac{1}{2} \epsilon^{ijk} \nu^i d\nu^j \wedge d\nu^k, \qquad (4.14)$$

which satisfies $J_{mn}J^{nl} = -\delta_m^l$, and also $-\frac{1}{2}J_{(2)} \wedge J_{(2)} = *_6J_{(2)}$. The parallel torsion of $J_{(2)}$ is

$$G_{(3)} = -\frac{1}{3}dJ_{(2)}.$$
(4.15)

Then $H_{(3)} \equiv *_6 G_{(3)}$ satisfies the relation

$$dH_{(3)} = 2J_{(2)} \wedge J_{(2)}, \tag{4.16}$$

where " $*_6$ " is the Hodge dual defined with respect to the S^6 metric. The uplift of the $\mathcal{N} = 1$ G₂-invariant critical point gives rise to the 10D solution

$$L^{-2}d\hat{s}_{10}^{2} = \alpha^{-3} \left(\frac{25\sqrt{15}}{256}ds_{AdS_{4}}^{2}\right) + \alpha ds_{S^{6}}^{2}, \quad \alpha = \frac{15^{3/8}}{2\sqrt{2}},$$
$$e^{-\frac{3}{2}\hat{\phi}} = e^{-\frac{3}{2}\phi_{0}}\alpha^{-1}, \qquad L^{-1}e^{\frac{3}{4}\phi_{0}}\hat{A}_{(1)} = 0,$$
$$L^{-2}e^{-\frac{1}{2}\phi_{0}}\hat{A}_{(2)} = \frac{1}{4}J_{(2)}, \qquad L^{-3}e^{\frac{1}{4}\phi_{0}}\hat{A}_{(3)} = \frac{1}{4}H_{(3)} + \frac{25\sqrt{15}}{128}\Omega_{(3)}, \qquad (4.17)$$

where $ds_{S^6}^2$ is the metric of the unit S^6 given in (4.3). Again, we introduced $L^2 = 2^{-\frac{1}{12}}g^{-25/12}m^{1/12}$ and $e^{\phi_0} = 2^{\frac{5}{6}}g^{\frac{5}{6}}m^{-\frac{5}{6}}$ using the scaling symmetries. It is recalled that in [33], general $\mathcal{N} = 1^7$ flux compactification of massive type IIA string has been analyzed. Therefore, our $\mathcal{N} = 1$, G₂-invariant solution should be contained in the discussion of [33].

The stable non-supersymmetric G_2 -invariant solution is obtained by uplifting the nonsupersymmetric G_2 -invariant critical point of the D = 4 theory and the result is given by

$$L^{-2}d\hat{s}_{10}^{2} = \alpha^{-3} \left(\frac{3\sqrt{3}}{16}ds_{AdS_{4}}^{2}\right) + \alpha ds_{S^{6}}^{2}, \quad \alpha = \frac{3^{3/8}}{2^{3/4}},$$
$$e^{-\frac{3}{2}\hat{\phi}} = e^{-\frac{3}{2}\phi_{0}}\alpha^{-1}, \qquad L^{-1}e^{\frac{3}{4}\phi_{0}}\hat{A}_{(1)} = 0,$$
$$L^{-2}e^{-\frac{1}{2}\phi_{0}}\hat{A}_{(2)} = -\frac{1}{2}J_{(2)}, \qquad L^{-3}e^{\frac{1}{4}\phi_{0}}\hat{A}_{(3)} = -\frac{1}{2}H_{(3)} + \frac{3\sqrt{3}}{8}\Omega_{(3)}.$$
(4.18)

⁷Here $\mathcal{N} = 1$ means four real supercharges.

5 Discussions and conclusions

In the work [29], it has been proposed that the holographic duals of the $AdS_4 \times \widetilde{S}^6$ (\widetilde{S}^6 means a deformed 6-sphere) backgrounds in massive type IIA are the simplest superconformal field theory first considered by Schwarz [43], with a simple gauge group SU(N), adjoint matter and level k. The Romans mass parameter m which gives rise to the magnetic coupling of the D = 4 theory has been argued to be related to the Chern-Simons level by [29, 44]

$$m = k/(2\pi\ell_s),\tag{5.1}$$

where ℓ_s is the string length. The $\mathcal{N} = 2$ solutions found in [29,39] are conjectured to be dual to certain $\mathcal{N} = 2$ Chern-Simons matter theories. Some evidence has been given by matching the partitions of the supergravity solutions with those of the conjectured CFT's [29,39].

This proposal seems to be natural if one can think of this theory as the IR fixed point of the world volume theory on D2-branes probing a flat transverse space, deformed by a Chern-Simons term. The Chern-Simons term is induced onto D2-brane by the Romans mass. The theory has a superpotential of the form $\mathcal{W} = \text{Tr}X[Y, Z]$, which respects the same global symmetry as the supergravity solution. An $\mathcal{N} = 3$ theory can be obtained by an RG flow from the $\mathcal{N} = 2$ Chern-Simons matter theory triggered by a relevant deformation $\frac{1}{2}\epsilon \text{Tr}Z^2$. In the IR, the coefficient ϵ runs to k/π , and the dimension of Z becomes 1. Since the kinetic term of Z becomes irrelevant, one can then integrate out Z, yielding a superpotential of the form $\mathcal{W} = \frac{\pi}{2k} \text{Tr}[X, Y]^2$ for the remaining two massless chiral multiplets. This is the scenario studied in [45]. The resulting IR theory consists of a Chern-Simon gauge field and a hypermultiplet encompassing an $\text{SU}(2)_{\text{R}} \times \text{SU}(2)_{\text{f}}$ global symmetry. The cosmological constant of the $\mathcal{N} = 3$ solution is smaller than that of the $\mathcal{N} = 2$ solution, therefore it is possible that in the gravity side, the $\mathcal{N} = 3$ solution can be attained from the $\mathcal{N} = 2$ solution via a holographic RG flow. If this is true, it is likely that the holographic dual of our $\mathcal{N} = 3$ solution is the $\mathcal{N} = 3$ Chern-Simons matter theory mentioned before.

Some evidence of this proposal can be seen from matching the bulk states of the 4D supergravity theory with the first few chiral primary multiplets in the dual CFT. A complete match between the short supergravity multiplets and the short CFT chiral primary multiplets requires the knowledge of the full Kaluza-Klein spectrum, nonetheless, we can make the first attempt by comparing the spectrum of the 4D supergravity theory with the lowest-level chiral primaries. Recall that [13] in the $\mathcal{N} = 3 \ AdS_4$ vacuum, the $\mathcal{N} = 8$ supergravity multiplet decomposes into one short supergravity multiplet DS(2,3/2,0)_S, two short gravitino multiplets DS(3/2,3/2,1/2)_S, one long gravitino multiplet DS(3/2, $\sqrt{3}$,0)_L

and three massless vector multiplets DS(1,1). On the CFT side, the chiral primary multiplets (labeled by the quantum number of the chiral primary operator) of the previously mentioned $\mathcal{N} = 3$ CFT was studied in [46]. In their notation, the short supergravity multiplet corresponds to the short chiral primary multiplet $(\frac{3}{2}, \frac{1}{2}, 0, 0)$, the two short gravitino multiplets correspond to $(\frac{3}{2}, 0, \frac{1}{2}, \frac{1}{2})$, and the three massless vector multiplets correspond to (1, 0, 1, 1).

As possible future research directions, it should be interesting to look for a D2 brane background whose near horizon limit smoothly approaches the $\mathcal{N} = 3$ solution found in the paper. Remember that besides the supersymmetric solutions, we have also found two stable non-supersymmetric critical points which are invariant under $SO(3)_R \times SO(3)_D$ and G_2 respectively. The stability of these solutions suggests their CFT duals may exist. It is therefore very intriguing to ask what kind of CFT could possibly be dual to these nonsupersymmetric AdS backgrounds.

The following diagram exhibits the heights of the cosmological constants for various critical points, from which one can guess whether two critical points can be connected by a holographic RG flow. It should be interesting to see if one can actually construct domain wall solutions interpolating these critical points, which can be useful for studying the 3D CFT.



Figure 1: List of critical points according to the heights of their cosmological constants. The $\mathcal{N} = 2$ point found in [29] is included here. This suggests possible holographic RG flows among these critical points.

Acknowledgements

We are grateful to Chris Pope for useful comments. This work is partially supported by DOE grant DE-FG02-13ER42020.

References

- [1] B. de Wit and H. Nicolai, N=8 Supergravity, Nucl. Phys. B 208, 323 (1982).
- B. de Wit, H. Samtleben and M. Trigiante, The Maximal D=4 supergravities, JHEP 0706 (2007) 049, arXiv:0705.2101 [hep-th].
- [3] G. Dall'Agata, G. Inverso and M. Trigiante, Evidence for a family of SO(8) gauged supergravity theories, Phys. Rev. Lett. 109, 201301 (2012), arXiv:1209.0760 [hep-th].
- [4] B. de Wit and H. Nicolai, Deformations of gauged SO(8) supergravity and supergravity in eleven dimensions, JHEP 1305 (2013) 077, arXiv:1302.6219 [hep-th].
- [5] G. Dall'Agata, G. Inverso and A. Marrani, Symplectic Deformations of Gauged Maximal Supergravity, JHEP 1407, 133 (2014), arXiv:1405.2437 [hep-th].
- [6] B. de Wit and H. Nicolai, The Consistency of the S**7 Truncation in D=11 Supergravity, Nucl. Phys. B 281, 211 (1987).
- [7] G. Dall'Agata and G. Inverso, On the Vacua of N = 8 Gauged Supergravity in 4 Dimensions, Nucl. Phys. B 859, 70 (2012), arXiv:1112.3345 [hep-th].
- [8] A. Borghese, A. Guarino and D. Roest, All G₂ invariant critical points of maximal supergravity, JHEP **1212** (2012) 108, arXiv:1209.3003 [hep-th].
- [9] A. Borghese, G. Dibitetto, A. Guarino, D. Roest and O. Varela, The SU(3)-invariant sector of new maximal supergravity, JHEP 1303, 082 (2013), arXiv:1211.5335 [hep-th].
- [10] A. Borghese, A. Guarino and D. Roest, Triality, Periodicity and Stability of SO(8) Gauged Supergravity, JHEP 1305, 107 (2013), arXiv:1302.6057 [hep-th].
- [11] A. Anabalon and D. Astefanesei, Black holes in ω-deformed gauged N = 8 supergravity, Phys. Lett. B 732, 137 (2014), arXiv:1311.7459 [hep-th].
- [12] H. Lü, Y. Pang and C. N. Pope, An ω deformation of gauged STU supergravity, JHEP 1404, 175 (2014), arXiv:1402.1994 [hep-th].

- [13] A. Gallerati, H. Samtleben and M. Trigiante, The N > 2 supersymmetric AdS vacua in maximal supergravity, JHEP 1412, 174 (2014), arXiv:1410.0711 [hep-th].
- [14] Y. Pang, C. N. Pope and J. Rong, Holographic RG Flow in a New SO(3) × SO(3) Sector of ω-Deformed SO(8) Gauged N = 8 Supergravity, JHEP 1508 (2015) 122, arXiv:1506.04270 [hep-th].
- [15] J. Tarrío and O. Varela, *Electric/magnetic duality and RG flows in AdS*₄/*CFT*₃, JHEP 1401, 071 (2014), arXiv:1311.2933 [hep-th].
- [16] A. Guarino, On new maximal supergravity and its BPS domain-walls, JHEP 1402, 026 (2014), arXiv:1311.0785 [hep-th].
- [17] S. Cremonini, Y. Pang, C. N. Pope and J. Rong, Superfluid and metamagnetic phase transitions in ω-deformed gauged supergravity, JHEP 1504, 074 (2015), arXiv:1411.0010 [hep-th].
- [18] A. Borghese, Y. Pang, C. N. Pope and E. Sezgin, Correlation Functions in ω-Deformed N=6 Supergravity, JHEP 1502 (2015) 112, arXiv:1411.6020 [hep-th].
- [19] B. de Wit, H. Nicolai and N. P. Warner, The Embedding of Gauged N = 8 Supergravity Into d = 11 Supergravity, Nucl. Phys. B 255 (1985) 29.
- [20] H. Godazgar, M. Godazgar, O. Krueger and H. Nicolai, Consistent 4-form fluxes for maximal supergravity, arXiv:1507.07684 [hep-th].
- [21] K. Lee, C. Strickland-Constable and D. Waldram, New gaugings and non-geometry, arXiv:1506.03457 [hep-th].
- [22] F. Hassler and D. Lst, Consistent Compactification of Double Field Theory on Nongeometric Flux Backgrounds, JHEP 1405 (2014) 085, arXiv:1401.5068 [hep-th].
- [23] R. Blumenhagen, F. Hassler and D. Lst, Double Field Theory on Group Manifolds, JHEP 1502 (2015) 001, arXiv:1410.6374 [hep-th].
- [24] R. Blumenhagen, P. d. Bosque, F. Hassler and D. Lust, Generalized Metric Formulation of Double Field Theory on Group Manifolds, JHEP 1508 (2015) 056, arXiv:1502.02428 [hep-th].
- [25] H. Kodama and M. Nozawa, Inflation in maximal gauged supergravities, JCAP 1505 (2015) 05, 028, arXiv:1502.01378 [hep-th].

- [26] A. Guarino, CSO_c superpotentials, arXiv:1508.05055 [hep-th].
- [27] C. M. Hull, A New Gauging of N = 8 Supergravity, Phys. Rev. D 30, 760 (1984).
- [28] M. Pernici, K. Pilch, P. van Nieuwenhuizen and N. P. Warner, Noncompact Gaugings and Critical Points of Maximal Supergravity in Seven-dimensions, Nucl. Phys. B 249, 381 (1985).
- [29] A. Guarino, D. L. Jafferis and O. Varela, The String Origin of Dyonic $\mathcal{N} = 8$ Supergravity and its Simple Chern-Simons Duals, arXiv:1504.08009 [hep-th].
- [30] H. Godazgar, M. Godazgar, O. Krger, H. Nicolai and K. Pilch, An SO(3)×SO(3) invariant solution of D = 11 supergravity, JHEP 1501 (2015) 056, arXiv:1410.5090 [hep-th].
- [31] W. H. Baron and G. Dall'Agata, Uplifting non-compact gauged supergravities, JHEP 1502 (2015) 003, arXiv:1410.8823 [hep-th].
- [32] K. Behrndt and M. Cvetic, General N = 1 supersymmetric flux vacua of (massive) type IIA string theory, Phys. Rev. Lett. 95, 021601 (2005), hep-th/0403049.
- [33] K. Behrndt and M. Cvetic, General N=1 supersymmetric fluxes in massive type IIA string theory, Nucl. Phys. B 708, 45 (2005), hep-th/0407263.
- [34] C. Nunez, I. Y. Park, M. Schvellinger and T. A. Tran, Supergravity duals of gauge theories from F(4) gauged supergravity in six-dimensions, JHEP 0104, 025 (2001), hep-th/0103080.
- [35] D. Lust and D. Tsimpis, Supersymmetric AdS(4) compactifications of IIA supergravity, JHEP 0502, 027 (2005), hep-th/0412250.
- [36] A. Tomasiello, New string vacua from twistor spaces, Phys. Rev. D 78, 046007 (2008), arXiv:0712.1396 [hep-th].
- [37] D. Lust and D. Tsimpis, New supersymmetric AdS(4) type II vacua, JHEP 0909, 098 (2009), arXiv:0906.2561 [hep-th].
- [38] A. Rota and A. Tomasiello, AdS₄ compactifications of AdS₇ solutions in type II supergravity, JHEP 1507 (2015) 076, arXiv:1502.06622 [hep-th].
- [39] M. Fluder and J. Sparks, D2-brane Chern-Simons theories: F-maximization = amaximization, arXiv:1507.05817 [hep-th].

- [40] A. Guarino and O. Varela, *Dyonic ISO(7) supergravity and the duality hierarchy*, arXiv:1508.04432 [hep-th].
- [41] O. Aharony, O. Bergman, D. L. Jafferis and J. Maldacena, N=6 superconformal Chern-Simons-matter theories, M2-branes and their gravity duals, JHEP 0810, 091 (2008), arXiv:0806.1218 [hep-th].
- [42] M. Cvetic, H. Lü and C. N. Pope, Gauged six-dimensional supergravity from massive type IIA, Phys. Rev. Lett. 83, 5226 (1999), hep-th/9906221.
- [43] J. H. Schwarz, Superconformal Chern-Simons theories, JHEP 0411 (2004) 078, hepth/0411077.
- [44] D. Gaiotto and A. Tomasiello, *The gauge dual of Romans mass*, JHEP **1001** (2010) 015, arXiv:0901.0969 [hep-th].
- [45] D. Gaiotto and X. Yin, Notes on superconformal Chern-Simons-Matter theories, JHEP 0708, 056 (2007), arXiv:0704.3740 [hep-th].
- [46] S. Minwalla, P. Narayan, T. Sharma, V. Umesh and X. Yin, Supersymmetric States in Large N Chern-Simons-Matter Theories, JHEP 1202, 022 (2012), arXiv:1104.0680 [hep-th].