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# Short Baseline Reactor $\bar{\nu} - e$ Scattering Experiments and Nonstandard Neutrino Interactions at Source and Detector

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We investigate nonstandard interaction effects in antineutrino–electron scattering experiments with baselines short enough to ignore standard oscillation phenomena. The setup is free of ambiguities from the interference between new physics and oscillation effects and is sensitive to both semileptonic new physics at the source and purely leptonic new physics in the weak interaction scattering at the detector. We draw on the TEXONO experiment as the model system, extending its analysis of nonstandard interaction effects at the detector to include the generally allowed nonstandard interaction phase at the detector and both non-universal and flavor changing new physics at the reactor source. We confirm that the current data allows for new physics constraints at the detector of the same order as those currently published, but we find that constraints on the source new physics are at least an order of magnitude weaker. The new physics phase effects are at the 5% level, noticeable in the 90% C.L. contour plots but not significantly affecting the conclusions. Based on projected increase in sensitivity with an upgraded TEXONO experiment, we estimate the improvement of sensitivity to both source and detector nonstandard interactions. We find that the bounds on source parameters improve by an order of magnitude, but do not reach parameter space beyond current limits. On the other hand, the detector new physics sensitivity would push current limits by factors 5 to 10 smaller.

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## I. INTRODUCTION

In the past several years, reactor neutrino experiments [1–3] and long baseline accelerator experiments [4, 5] have produced important advances in our understanding of neutrino mixing by measuring the key mixing parameter  $\theta_{13}$  by two completely independent processes. The reactor experiments measure  $\bar{\nu}_e$  disappearance in the flux of  $\bar{\nu}_e$ s, indicating oscillation into other neutrino flavors during the one or two kilometer trip from reactor core to detector. The accelerator experiment measures the appearance of a  $\nu_e$  component in the  $\nu_\mu$  beam from an accelerator during the hundreds of kilometers trip from the accelerator laboratory to the detection site. Together, the results already constrain the CP-violating phase angle in the mixing matrix [4, 5]. Moreover, the data provide a potentially powerful probe of non-standard interactions (NSI) [6] in the neutrino sector involving some combination of neutrino source, propagation, and detection [7–9].

In this paper, we explore the constraints on semileptonic, charged-current (CC), non-universal (NU) and flavor-changing (FC) NSI parameters and likewise for both NU and FC purely leptonic NSI parameters. The former appear in effective Lagrangians for neutrino production from reactors and from accelerators and for neutrino detection by inverse beta decay. The latter appear in neutrino production from muon decay and from neutrino detection by  $\nu - e$  or  $\bar{\nu} - e$  scattering. We will focus on the case of *very* short baseline reactor  $\bar{\nu}_e$  source and detection of the recoil electron from  $\bar{\nu}_e + e \rightarrow \bar{\nu}_e + e$  scattering at the detector. We rely heavily on the example provided by the TEXONO experiment [10, 11], which measures the recoil electron spectrum from reactor anti-neutrinos interacting with electrons in a CsI(Tl) detector. The baseline is less than 30m, and the oscillation of the beam can be ignored, thus providing an especially clean test of FC “wrong flavor”  $\bar{\nu}_\mu$  or  $\bar{\nu}_\tau$  or NU “right flavor”  $\bar{\nu}_e$  from the semileptonic nuclear decays in the reactor. Baselines this short avoid the degeneracies between NSI parameters and standard

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neutrino mixing parameters that occur in analysis of data from reactor experiments with kilometer [1–3] or tens of kilometer baselines [12], degeneracies that are touched on in several recent studies [13–15].

We extend work in Ref. [11] by incorporating the effects of NSI produced at the source and by including the phase dependence of the FC NSI at the detector using the data from Texono’s experiment. In Ref. [11], only the NSI at the detector in the single channel of  $\bar{\nu}_e - e$  scattering are considered. With NSI at the source, there is a modification of the  $\bar{\nu}_e$  component and an addition of  $\bar{\nu}_\mu$  and  $\bar{\nu}_\tau$  components, so  $\bar{\nu}_\mu - e$  scattering and  $\bar{\nu}_\tau - e$  scattering must be incorporated by including NSI in the elastic, purely neutral current (NC),  $\bar{\nu}_\mu - e$  and  $\bar{\nu}_\tau - e$  cross sections, applicable for analyzing data from any short baseline neutrino scattering experiment where the oscillation effects are ignorable. Our “no NSI propagation effects” study complements those that probe NSI with solar neutrino, accelerator neutrino and other reactor neutrino experiments, which involve different combinations of NSI at source, propagation and detection effects [16–22].

In Section II, III and IV, we define our notation, specify our cross sections and define the flux factors that go with each cross section to unify all the standard model (SM) plus NSI contributions to the rate at source and detector in a single framework. Our formalism allows us to make joint confidence level (C.L.) contours with NSI parameters at source and at detector or at source alone and at detector alone. In Sec. IV, we apply this formalism to the TEXONO data and check key results from Refs. [10] and [11], while in Sec. V we apply the formalism to the modeled data based on the realistically achievable sensitivity proposed for an upgrade of the TEXONO experiment [23]. We recap and conclude in Sec. VI. The Appendix A in Sec. VII briefly summarizes the reactor flux and target density input to the recoil electron spectrum in the TEXONO experiment. Appendix B provides a table summarizing relevant model independent NSI parameter bounds from Ref. [9].

## II. FORMALISM OF SOURCE AND DETECTOR NSI

### A. NSI effective Lagrangians at source and detector

In the problem we address here, the source of antineutrinos is the semileptonic, CC decays of reactor nuclei. At the level of the quark content of the nucleons, the transition  $d \rightarrow u + e + \bar{\nu}$  provides the antineutrinos for the elastic  $\bar{\nu} - e$  scattering process at the detector. To allow for lepton-flavor-violating decays at the source, we adopt the semileptonic, CC, effective Lagrangian [13, 24, 25]

$$\mathcal{L}^S = -2\sqrt{2}G_F(\delta_{\alpha\beta} + K_{\alpha\beta})(\bar{L}_\alpha \gamma_\lambda P_L U_{\beta a} \nu_a)(\bar{d} \gamma^\lambda P_L u)^\dagger + h.c., \quad (1)$$

where repeated flavor-basis indices “ $\alpha$ ” and “ $\beta$ ” and mass-basis indices “ $a$ ” are summed over. We confine ourselves to the left-handed quark helicity projection case for simplicity. The inclusion of the right handed terms adds nothing essential to our discussion. Since we consider neutrino-propagation baselines are only a few tens of meters and energies that are in the few MeV range, oscillations play no role and we can effectively replace  $U_{\beta a} \bar{\nu}_a \rightarrow \bar{\nu}_\beta$  in making the rate calculations we present here. The complex coefficients  $K_{\alpha\beta}$  represent the relative coupling strengths of the flavor combinations in the presence of new physics, while in the SM,  $K_{\alpha\beta} = 0$ .

To represent the NSI effects in the purely leptonic sector [24–28] for the simplified elastic  $\bar{\nu} - e$  scattering case of interest, we write the effective Lagrangian as

$$\begin{aligned} \mathcal{L}^\ell &= \mathcal{L}_{NU}^\ell + \mathcal{L}_{FC}^\ell \\ &= -2\sqrt{2}G_F \sum_\alpha (\bar{e} \gamma_\mu (\tilde{g}_{\alpha R} P_R + (\tilde{g}_{\alpha L} + 1) P_L) e) (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\alpha) \\ &\quad - 2\sqrt{2}G_F \sum_{\alpha \neq \beta} \epsilon_{\alpha\beta}^P (\bar{e} \gamma_\lambda P e) (\bar{\nu}_\alpha \gamma^\lambda P_L \nu_\beta). \end{aligned} \quad (2)$$

The first term in Eq. (2) is the NU case and the second term is the FC case. The coefficients  $\tilde{g}_{\alpha R}$  and  $\tilde{g}_{\alpha L}$  are

$$\tilde{g}_{\alpha R} = \sin^2 \theta_w + \epsilon_{\alpha\alpha}^{eR} \text{ and } \tilde{g}_{\alpha L} = \sin^2 \theta_w - \frac{1}{2} + \epsilon_{\alpha\alpha}^{eL}. \quad (3)$$

Hermiticity of  $\mathcal{L}^\ell$  requires that the NSI matrix of parameters be Hermitian:  $\epsilon_{\alpha\beta}^{eR,L} = (\epsilon_{\beta\alpha}^{eR,L})^*$ , so the FC NSI parameters are complex in general. Adopting the commonly used “ $\epsilon$ ” notation for the leptonic sector makes the distinction between source (Ks) and detector ( $\epsilon$ s) clear. With the effective Lagrangians defined, we are now ready to summarize the cross sections and flux factors we need for the study of the NSI effects at source and detector.

### B. $\bar{\nu}_e - e$ , $\bar{\nu}_\mu - e$ and $\bar{\nu}_\tau - e$ differential scattering cross sections in lab frame

In the notation for the NSI terms defined in Eq. (2) above, the differential cross section for the  $\bar{\nu}_e - e$  scattering with neutrino lab energy  $E_\nu$  and recoil electron kinetic energy  $T$  can be summarized by the expression

$$\begin{aligned} \left[ \frac{d\sigma(\bar{\nu}_e e)}{dT} \right]_{SM+NSI} &= \frac{2G_F^2 m_e}{\pi} [\tilde{g}_{eR}^2 + \sum_{\alpha \neq e} |\epsilon_{\alpha e}^{eR}|^2 \\ &\quad + \left( (\tilde{g}_{eL} + 1)^2 + \sum_{\alpha \neq e} |\epsilon_{\alpha e}^{eL}|^2 \right) \left( 1 - \frac{T}{E_\nu} \right)^2 \\ &\quad - \left( \tilde{g}_{eR}(\tilde{g}_{eL} + 1) + \sum_{\alpha \neq e} \Re[(\epsilon_{\alpha e}^{eR})^* \epsilon_{\alpha e}^{eL}] \right) \frac{m_e T}{E_\nu^2}], \end{aligned} \quad (4)$$

which is the sum of the scattering cross sections for the three, incoherent processes  $\bar{\nu}_e + e \rightarrow \bar{\nu}_e + e$ ,  $\bar{\nu}_e + e \rightarrow \bar{\nu}_\mu + e$  and  $\bar{\nu}_e + e \rightarrow \bar{\nu}_\tau + e$ . The  $\bar{\nu}_e + e \rightarrow \bar{\nu}_e + e$  cross section is represented by the terms containing the  $\tilde{g}_{eL}$  and  $\tilde{g}_{eR}$  parameters. It is the coherent sum of the NC and CC contributions. The complex parameters  $\epsilon_{\alpha e}^{eL}$ , where  $\alpha \neq e$ , can be written either as  $\epsilon_{\alpha e}^{eL} = \Re[\epsilon_{\alpha e}^{eL}] + i\Im[\epsilon_{\alpha e}^{eL}]$  or as  $|\epsilon_{\alpha e}^{eL}| \exp(i\phi_{\alpha e}^{eL})$ , where  $\phi_{\alpha e}^{eL}$  is the phase angle of the complex quantity. Written out in more detail, the NSI contributions are  $|\epsilon_{\alpha e}^{eR}|^2 = (\Re[\epsilon_{\alpha e}^{eR}])^2 + (\Im[\epsilon_{\alpha e}^{eR}])^2$ , and similarly for  $R \rightarrow L$ . In the last term,  $\Re[(\epsilon_{\alpha e}^{eR})^* \epsilon_{\alpha e}^{eL}] = \Re[\epsilon_{\alpha e}^{eR}] \Re[\epsilon_{\alpha e}^{eL}] + \Im[\epsilon_{\alpha e}^{eR}] \Im[\epsilon_{\alpha e}^{eL}]$ . This notation makes it clear that when the  $\epsilon$  parameters are taken as real positive or negative, then the "Re" and "Im" notation can be dropped and one can drop the absolute magnitude signs everywhere. All of the NSI studies with  $\nu - e$  or  $\bar{\nu} - e$  scattering at the detector tacitly make this assumption [11, 16–19, 22]. If the parameters are written as  $|\epsilon_{\alpha e}^{eL}| \exp(i\phi_{\alpha e}^{eL})$  and  $|\epsilon_{\alpha e}^{eR}| \exp(i\phi_{\alpha e}^{eR})$ , then the coefficient in the last term can be expressed as

$$\Re[(\epsilon_{\alpha e}^{eR})^* \epsilon_{\alpha e}^{eL}] = |\epsilon_{\alpha e}^{eR}| |\epsilon_{\alpha e}^{eL}| \cos(\phi_{\alpha e}^{eL} - \phi_{\alpha e}^{eR}). \quad (5)$$

With this parameterization, the values of  $|\epsilon_{\alpha e}^{eR}|$  and  $|\epsilon_{\alpha e}^{eL}|$  are always positive and the sign of the term is controlled by  $\cos(\phi_{\alpha e}^{eL} - \phi_{\alpha e}^{eR})$ .

To include the NSI at the reactor source, using the notation from [13], one multiplies the contribution to the rate by  $|1 + K_{ee}|^2$ . Though Ref. [13] works only to first order in NSI parameters and drops the highly constrained linear term  $2\Re[K_{ee}]$  [9], in the present calculation we must work to second order to assess the impact of the NSI, so both  $\Im[K_{ee}]$  and  $\Re[K_{ee}]$  will be included in the NU case  $\bar{\nu}_e + e \rightarrow \bar{\nu}_e + e$ .

For the other incoming neutrino flavors, we multiply the  $\bar{\nu}_\mu - e$  cross section by the factor  $|K_{e\mu}|^2$  for the  $\bar{\nu}_\mu$  component of the flux and by  $|K_{e\tau}|^2$  for the  $\bar{\nu}_\tau$  component. The  $\bar{\nu}_\mu - e$  differential cross section is

$$\begin{aligned} \left[ \frac{d\sigma(\bar{\nu}_\mu e)}{dT} \right]_{SM+NSI} &= \frac{2G_F^2 m_e}{\pi} [\tilde{g}_{\mu R}^2 + \sum_{\alpha \neq \mu} |\epsilon_{\alpha \mu}^{eR}|^2 \\ &\quad + \left( \tilde{g}_{\mu L}^2 + \sum_{\alpha \neq \mu} |\epsilon_{\alpha \mu}^{eL}|^2 \right) \left( 1 - \frac{T}{E_\nu} \right)^2 \\ &\quad - \left( \tilde{g}_{\mu R} \tilde{g}_{\mu L} + \sum_{\alpha \neq \mu} \Re[(\epsilon_{\alpha \mu}^{eR})^* \epsilon_{\alpha \mu}^{eL}] \right) \frac{m_e T}{E_\nu^2}]. \end{aligned} \quad (6)$$

The cross section for  $\bar{\nu}_\tau - e$  scattering is obtained by replacing  $\mu$  by  $\tau$  everywhere in the above equation. The definitions of  $\tilde{g}_{\mu R, \mu L}$  and  $\tilde{g}_{\tau R, \tau L}$  are obvious counterparts to the definition of  $\tilde{g}_{eR, eL}$  in Eq. (3).

### C. Discussion of NSI at the source and the full NSI effects

The distance between the source and detector in the TEXONO experiment is less than 30m, so we will use the fact that the oscillation effects, proportional to  $\sin^2(m_i^2 - m_j^2)L/4E_\nu$ , are ignorable for the range of interest,  $3 \text{ MeV} \leq E_\nu \leq 8 \text{ MeV}$ . In effect, this means that the flavor of neutrino that is produced at the source is the same as the flavor that reaches the detector. The factors that control the flux of each flavor in the incoming beam produced at the source are the  $K_{\alpha\beta}$ . The TEXONO flux model is the result of a large number of independent nuclear reactions. In the presence of NSI, the emitted flux can be thought of as an incoherent sum of  $\bar{\nu}_e, \bar{\nu}_\mu$  and  $\bar{\nu}_\tau$  with weights  $|1 + K_{ee}|^2, |K_{e\mu}|^2$  and

$|K_{e\tau}|^2$ . The source and detector NSI effects on the rate are then expressed through the following factor, denoted by  $\mathcal{F}$ , that will multiply the reactor flux and the target electron number density to get the differential rate  $\frac{dR_X}{dT}$ , as described in Appendix A:

$$\mathcal{F} = |1 + K_{ee}|^2 \left[ \frac{d\sigma(\bar{\nu}_e e)}{dT} \right] + |K_{e\mu}|^2 \left[ \frac{d\sigma(\bar{\nu}_\mu e)}{dT} \right] + |K_{e\tau}|^2 \left[ \frac{d\sigma(\bar{\nu}_\tau e)}{dT} \right], \quad (7)$$

where the cross section formulas are as given in Eqs. (4) and (6) and the SM plus NSI designation is understood.

### III. PROBING MODEL PARAMETERS WITH RECOIL ELECTRON ENERGY SPECTRUM DATA: THE TEXONO EXPERIMENT

We reproduce and recap the TEXONO experiment [10] and its related analyses [11, 23] that are directly relevant to our NSI parameters study. The neutrino flux spectrum and the event rate data and its theoretical representation are briefly summarized in the Appendix A. In Ref. [10], the primary goal was an independent determination of the weak mixing parameter  $\sin^2 \theta_W$ , determined strictly from low energy, purely leptonic recoil spectrum data in the  $\bar{\nu}_e + e \rightarrow \bar{\nu}_e + e$  elastic scattering process. The paper stresses that this data is more sensitive to the right-handed NC component in Eq. (4) than is the corresponding  $\nu_e + e \rightarrow \nu_e + e$  scattering case, where the roles of  $g_L$  and  $g_R$  are reversed. The  $\bar{\nu}_e - e$  scattering is consequently more sensitive to  $g_R = \sin^2 \theta_W$ . Using their flux and binned rate spectrum [29], we show the result of a  $\chi^2$  analysis with statistical errors only in Fig. 1(a). The  $1\sigma$  and 90% C.L. lines are included for guidance. We find a best fit of  $\sin^2 \theta_W = 0.251 \pm 0.030$  in agreement with TEXONO's result.

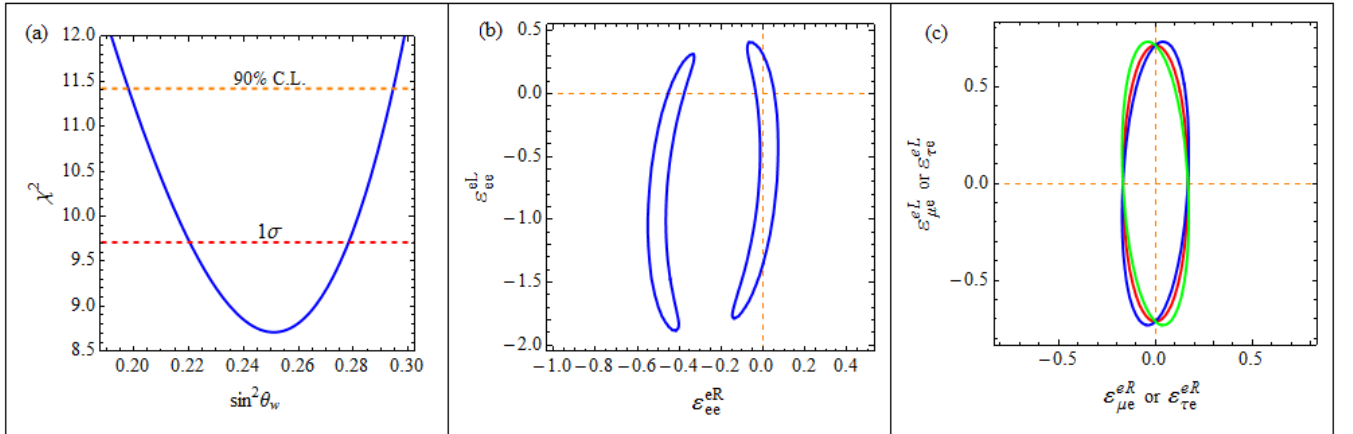


FIG. 1: SM  $\sin^2 \theta_W$  vs.  $\chi^2$ , 1(a), and our calculation of the 90% C.L. limits of Fig. 4(a) and 4(b) of Ref. [13] in Figs. 1(b) and 1(c). In Fig. 1(c), we show the 90% C.L. boundary for the fit to TEXONO rate data using Eq. (5) in the scattering cross section Eq. (2). The blue(B), red(R) and green(G) curves, right-to-left at the top, are for  $\cos(\phi_{\alpha e}^{eL} - \phi_{\alpha e}^{eR}) = 1, 0$  and  $-1$ , respectively. The blue curve, with  $\cos(\phi_{\alpha e}^{eL} - \phi_{\alpha e}^{eR}) = 1$ , corresponds to that shown in Fig. 4c of Ref. [13].

Following publication of their experimental results [10], detailing the experiment and the results on  $\sin^2 \theta_W$  and on an upper limit of the neutrino magnetic moment, the collaboration presented limits on NSI parameters and on couplings of unparticles to neutrinos and electrons [11]. Since we are pursuing an extension of the NSI bounds to include the possibility of semileptonic NSI modifications to the reactor source of  $\bar{\nu}_e$ s and the interplay with the purely leptonic detection NSI, we are primarily interested in C.L. boundaries in two parameter fits to the data and the joint limits obtained from these analyses. For illustration, we check our evaluation of the 90% C.L. boundaries in the  $\epsilon_{ee}^{eR} - \epsilon_{ee}^{eL}$  plane and, alternatively, the  $\epsilon_{e\tau}^{eR} - \epsilon_{e\tau}^{eL}$  plane, Figs. 4a and 4b in [11]. We show the result of this exercise in Figs. 1(b) and the blue boundary, rightmost at the top, in Fig. 1(c). In both cases we find that our results and TEXONO's agree within the ability to read off values along the contours. We show the 90% C.L. projections of these plots on the individual axes for the two cases in Table I. The red and green curves, center and leftmost in Fig. 1(c) are examples of other phase choices, as we discuss in subsection IIIA. For the NU case of  $\epsilon_{ee}^{eR} - \epsilon_{ee}^{eL}$  plane, we quote the right-hand solution values, since both the R and L limits are the most stringent for this solution. The FC case assumes the NSI parameters are purely real. There is no degeneracy in this case, and the projected individual two

parameter limits are straightforward. The weak correlation between the R and L NSI parameters is due to the small R-L NSI interference term. Though our contour agrees with that as obtained in Ref. [11] and our  $\epsilon_{\tau e}^{eL}$  bounds agree with the ones quoted in their Table I, our limits on  $\epsilon_{\tau e}^{eL}$  are somewhat smaller.

Figure No.	R-Parameter Bounds	L-Parameter Bounds
1(b)	$-0.15 < \epsilon_{ee}^{eR} < 0.08$	$-1.79 < \epsilon_{ee}^{eL} < 0.41$
1(c)	$-0.18 < \epsilon_{\alpha e}^{eR} < 0.18$	$-0.76 < \epsilon_{\alpha e}^{eL} < 0.76$

TABLE I: Bounds at 90% C.L. obtained from Fig. 1(b) and 1(c) in the absence of any source NSI where  $\alpha = \mu$  or  $\tau$ .

#### A. Role of the detector NSI phases in determining the C.L. boundaries

The R-L interference term in the differential cross sections depends on the FC NSI parameter phases, as displayed for the case  $\bar{\nu}_e + e \rightarrow \bar{\nu} + e$  in Eq. (5). From the point of view of this general formula, the blue boundary, rightmost at the top, in Fig. 1(c) can be interpreted as the composite of the cases  $\phi_{\mu e}^{eR} = \phi_{\mu e}^{eL} = 0$ , where  $\epsilon_{\mu e}^{eR}$  and  $\epsilon_{\mu e}^{eL}$  are both real and positive,  $\phi_{\mu e}^{eR} = \pi$  and  $\phi_{\mu e}^{eL} = 0$ , where  $\epsilon_{\mu e}^{eR}$  is real and negative and  $\epsilon_{\mu e}^{eL}$  is real and positive,  $\phi_{\mu e}^{eR} = \phi_{\mu e}^{eL} = \pi$ , where  $\epsilon_{\mu e}^{eR}$  and  $\epsilon_{\mu e}^{eL}$  are both real and negative, and, finally,  $\phi_{\mu e}^{eR} = 0$  and  $\phi_{\mu e}^{eL} = \pi$ , where  $\epsilon_{\mu e}^{eR}$  is real and positive and  $\epsilon_{\mu e}^{eL}$  is real and negative. Alternatively, it can be interpreted as the composite of cases where 0 and  $\pi$  are replaced with  $\pi/2$  and  $3\pi/2$  and real replaced with imaginary. Because the R-L interference term is suppressed by the factor  $m_e T / E_\nu^2$  and  $E_\nu \geq 3$  MeV, the changes in the parameter boundaries as the phase differences range from 0 to  $\pi$  are small, as shown in Fig. 1(c). Conclusions about allowed boundaries for NSI parameters for the range of energies of interest in reactor experiments are affected very little in this analysis, but for experiments with significantly lower energy radioactive sources or for low energy solar neutrino experiments such as Borexino [30], the R-L correlation term can be relatively larger and the phase effects may be important. For present purposes, we illustrate the range of effects that change of phases can make on the C.L. boundary in Fig. 1(c). The small changes in boundaries is shown in the figure by the difference between the blue, red and green curves, corresponding to  $\cos(\phi_{\alpha e}^{eL} - \phi_{\alpha e}^{eR}) = 1, 0$  and  $-1$ , reading from right to left at the top of the figure. As one sees, the correlation disappears for the case that  $\cos(\phi_{\alpha e}^{eL} - \phi_{\alpha e}^{eR}) = 0$ , the red, middle curve. The R-L correlation term vanishes in this case because the R and L parameters are  $\pi/2$  out of phase; one can be real and the other imaginary, for example.

#### IV. INTERPLAY BETWEEN $K_{\alpha\beta}$ (SOURCE) AND $\epsilon_{\alpha\beta}^{eR,L}$ (DETECTOR) NSI PARAMETERS

In this section we take pairs of source and detector NSI coefficients to survey the 90% C.L. boundaries in the various two-parameter spaces. We focus on the bounds on the source parameters and assess the strength of the bounds found to the bounds currently available in the literature. At the same time, we check for consistency of the bounds on the detector NSI parameters with those found by TEXONO [11], which we checked in the preceding section.

Since the current bounds on the real part of  $K_{ee}$  are of the order  $10^{-3}$ , as given in Ref. [9] and found independently from Daya Bay data in Ref. [15], and these are much tighter than we can imagine providing with the current analysis based on the TEXONO data, we assume  $K_{ee}$  is purely imaginary in this section. Consequently, the source coefficient in the case of incident  $\bar{\nu}_e$  in Eq. (7) is  $K_{ee}^2 = 1 + (\Im[K_{ee}])^2$ . Bounds found will then refer to  $\Im[K_{ee}]$ . Fig. 2 shows the 90% C.L. boundaries for the fits to the TEXONO data as parameterized by one source NSI coefficient and one detector coefficient with all of the other NSI coefficients set to zero. From the 90% C.L. contours shown in Fig. 2, we can determine the 90% C.L. bounds on the source NU  $K_{ee}$  parameter and any of the  $\epsilon_{\alpha e}^{eR,L}$  at the detector by projecting onto the parameter axes for each contour. We find the limits on the NU parameters quoted in Table II. In all of the cases involving the source FC semileptonic NSI parameters  $K_{\alpha\beta}$ , there is no bound on any of the leptonic, detector NSI parameters  $\epsilon_{\alpha\beta}^{eR,L}$  as  $K_{\alpha\beta} \rightarrow 0$ , because the source is receiving only  $\bar{\nu}_e$  flux in this limit. In this sense, the parameters  $K_{\alpha\beta}$  and  $\epsilon_{\alpha\beta}^{eR,L}$  are highly correlated. There is still the possibility for placing upper bounds on the  $K_{e\alpha}$  parameters in this case if the detector NSI parameters are constrained to be smaller than their current bounds [9], which are near zero on the scale of Fig. 2. We can then place upper 90% C.L. bounds on the values of  $K_{e\mu}$  or  $K_{e\tau}$  for the special case where detector NSI  $\epsilon_{\mu\mu}^{eR,L} = \epsilon_{\mu\mu}^{eR} = 0$ , and likewise for  $\mu \rightarrow \tau$ . These one-parameter-at-a time

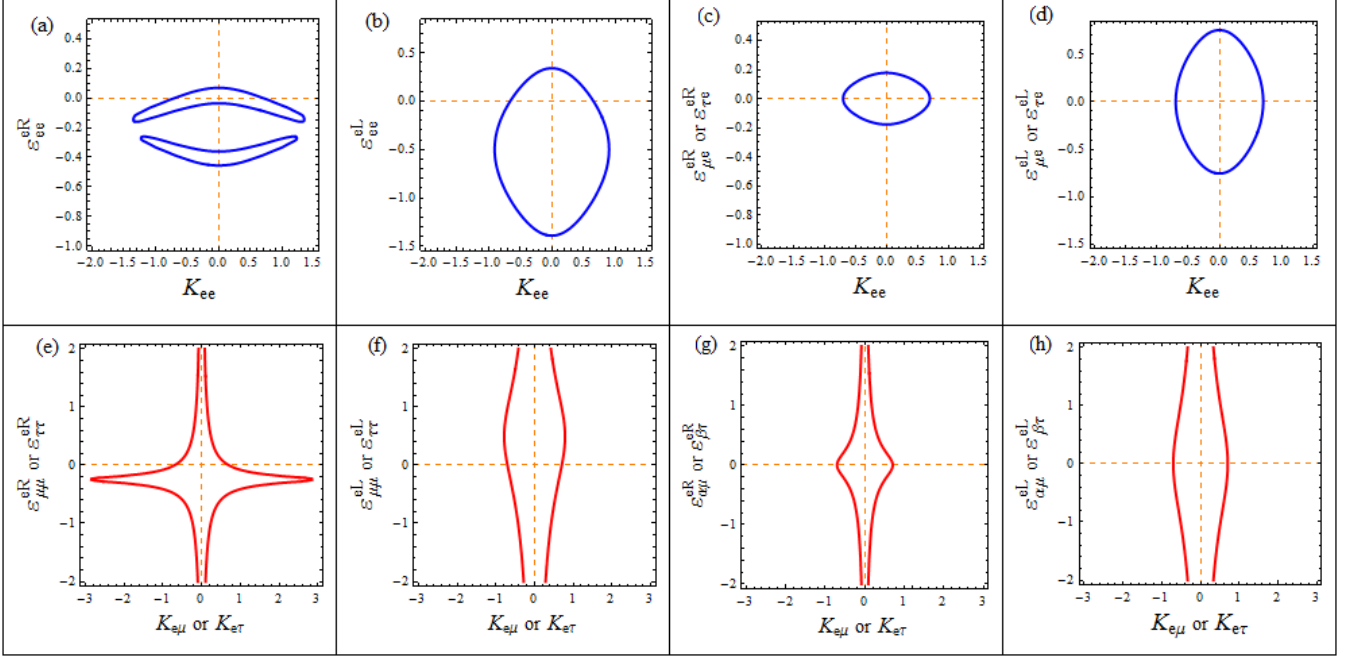


FIG. 2: C.L. boundary regions for TEXONO *realistic* data. **Upper Panel:** Correlation between the source NSI parameter ( $K_{ee}$ ) and the corresponding detector NSI parameters ( $\epsilon_{ee}^{R,L}$  and  $\epsilon_{\alpha e}^{R,L}$ , where  $\alpha = \mu$  or  $\tau$ ) at 90% C.L.. See the text for details. **Lower Panel:** Correlation between the source NSI parameter ( $K_{e\alpha}$ ) and the corresponding detector NSI parameters ( $\epsilon_{\mu\mu}^{R,L}$ ,  $\epsilon_{\tau\tau}^{R,L}$  and  $\epsilon_{\alpha\mu}^{R,L}$ ,  $\epsilon_{\beta\tau}^{R,L}$ , where  $\alpha = e$  or  $\tau$  and  $\beta = e$  or  $\mu$ ) at 90% C.L.. See the text for details.

upper bounds on  $K_{e\mu}$  or  $K_{e\tau}$ , the type commonly reported in the literature, are the bounds we quote in Table II [31]. Because only  $\epsilon_{\mu e}^R$  is taken to be non-zero in the two parameter analysis yielding Fig. 2(c), there is no dependence on its phase, and similarly for  $\epsilon_{\mu e}^L$  in Fig. 2(d). As indicated in Eq. (5) and illustrated in Fig. 1(c), both must be included in a fitting analysis for the relative phase to play a role.

Figure No.	NSI Parameters at Source	NSI Parameters at Detector
2(a)	$-1.35 < \text{Im } K_{ee} < 1.35$	$-0.17 < \epsilon_{ee}^R < 0.07$
2(b)	$-0.9 < \text{Im } K_{ee} < 0.9$	$-1.4 < \epsilon_{ee}^L < 0.34$
2(c)	$-0.72 < \text{Im } K_{ee} < 0.72$	$-0.18 < \epsilon_{\alpha e}^R < 0.18$
2(d)	$-0.72 < \text{Im } K_{ee} < 0.72$	$-0.76 < \epsilon_{\alpha e}^L < 0.76$
2(e)	$-0.72 < \text{Im } K_{e\mu} \text{ or } \text{Im } K_{e\tau} < 0.72$	$\epsilon_{\mu\mu}^R$ and $\epsilon_{\tau\tau}^R$ are unbounded
2(f)	$-0.72 < \text{Im } K_{e\mu} \text{ or } \text{Im } K_{e\tau} < 0.72$	$\epsilon_{\mu\mu}^L$ and $\epsilon_{\tau\tau}^L$ are unbounded
2(g)	$-0.72 < \text{Im } K_{e\mu} \text{ or } \text{Im } K_{e\tau} < 0.72$	$\epsilon_{\alpha\mu}^R$ and $\epsilon_{\beta\tau}^R$ are unbounded
2(h)	$-0.72 < \text{Im } K_{e\mu} \text{ or } \text{Im } K_{e\tau} < 0.72$	$\epsilon_{\alpha\mu}^L$ and $\epsilon_{\beta\tau}^L$ are unbounded

TABLE II: Bounds at 90% C.L. obtained from Fig. 2 where  $\alpha = e$  or  $\tau$  and  $\beta = e$  or  $\mu$ .

Briefly summarized, the results of this study based on the published TEXONO data show that the sensitivity to reactor *source* NSI,  $K_{\alpha\beta}$ , is at least an order of magnitude less than the sensitivity of the data used to establish the currently available bounds. On the other hand, the sensitivity to *detector* NSI is of the same order of magnitude as the current bounds for the right handed NSI couplings, though much less for the left-handed couplings. The future improvements in sensitivity, as envisioned by the TEXONO collaboration [11, 23], should change this situation considerably, and we turn to this consideration in the next section.

## V. FUTURE PROSPECTS

In this section we study the future prospects for tightening the source and detector NSI parameter bounds by adopting the projected “realistic and feasible” improvements in statistical sensitivities reported in Table 2 and the related text in Ref. [23]. Their essential point is that statistical uncertainty of the measured value of  $\sin^2 \theta_W$  can realistically be reduced to  $\pm 0.0013$ . We follow the experimental setup from Ref. [10, 11] and generate our data in 10 energy bins, each of step 0.5 MeV. We generate our “data model” by assuming that the best fit value turns out to be  $\sin^2 \theta_W = 0.2387$  [32], the value cited for comparison to their experimental fit value  $\sin^2 \theta_W = 0.251$  by Ref. [10]. We define our model  $\chi^2$  distribution by forming

$$\chi^2 = \sum_i \left( \frac{R_{NSI} - R_{SM}}{\Delta_{stat}} \right)_i^2 \quad (8)$$

where  $R_{SM}$  is the data model rate,  $R_{NSI}$  is the predicted event rate with all unknown NSI parameters and  $\Delta_{stat}$  is the statistical uncertainty over each bin. We define  $\Delta_{stat}$  as the deviation from the central value  $R_{SM}$  within  $1\sigma$  statistical uncertainty, obtained by evaluating the rate with  $\sin^2 \theta_W = 0.2387$ . To achieve a fit to the 10 bins of rate data that yields the projected uncertainty of  $\pm 0.0013$  for  $\sin^2 \theta_W$ , we find that evaluating the rates in each bin with  $\sin^2 \theta_W$  roughly  $(\sqrt{10} \simeq 3) \times \pm 0.0013$  per bin and taking the average deviation from the central value yields a data set whose uncertainties are consistent with expectations [23]. We take this model set as the basis for estimated future sensitivity to NSI [22]. The results shown in Fig. 1 are redone using future prospects data in Fig. 3 and the bounds obtained at 90% C.L. are given in Table III [22].

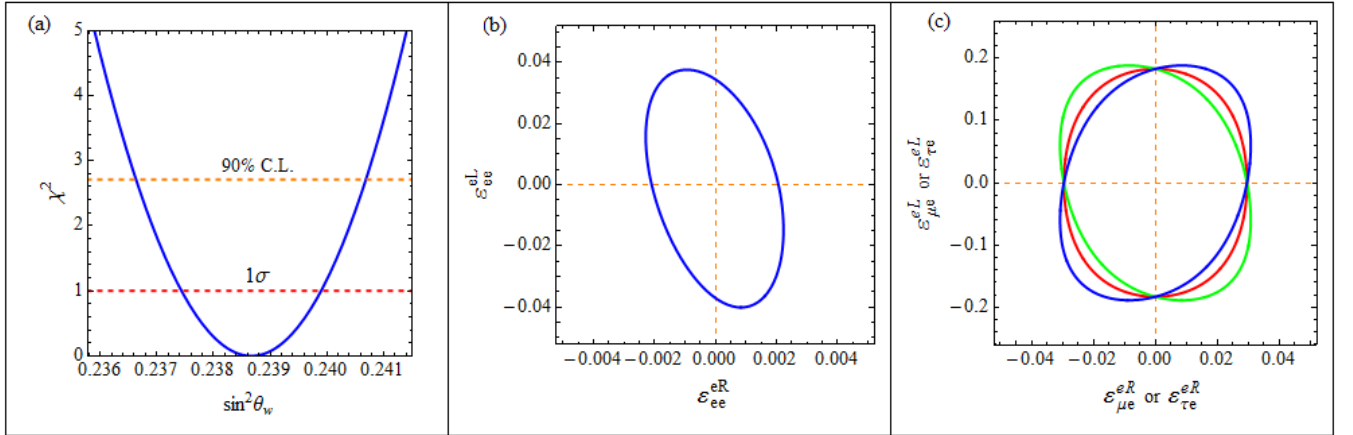


FIG. 3: The 90% C.L. contours for NU and FC target NSI parameters for the model future prospects data, where the source NSI parameters are all set to zero.

Figure No.	R-Parameter Bounds	L-Parameter Bounds
3(b)	$-0.0023 < \epsilon_{ee}^R < 0.0023$	$-0.04 < \epsilon_{ee}^L < 0.04$
3(c)	$-0.03 < \epsilon_{\alpha e}^R < 0.03$	$-0.19 < \epsilon_{\alpha e}^L < 0.19$

TABLE III: Bounds at 90% C.L. obtained from Fig. 3(b) and 3(c) in the absence of any source NSI where  $\alpha = \mu$  or  $\tau$ .

From Fig. 3 and the bounds summarized in Table III, we see immediately the impact of improved sensitivity to the presence of NSI at the detector in the removal of the degeneracy in the  $\epsilon_{ee}^R$  vs.  $\epsilon_{ee}^L$  plot when compared to Fig. 1. The purely leptonic NU and FC new physics effects can be probed with up to two orders of magnitude higher refinement in the right-handed lepton sector and up to an order of magnitude more refinement in the left-handed sector. With comparable experimental sensitivity in a  $\nu_e + e \rightarrow \nu_e + e$  experiment, a complimentary result with the *left-handed* sector being favored could be achieved [33, 34].



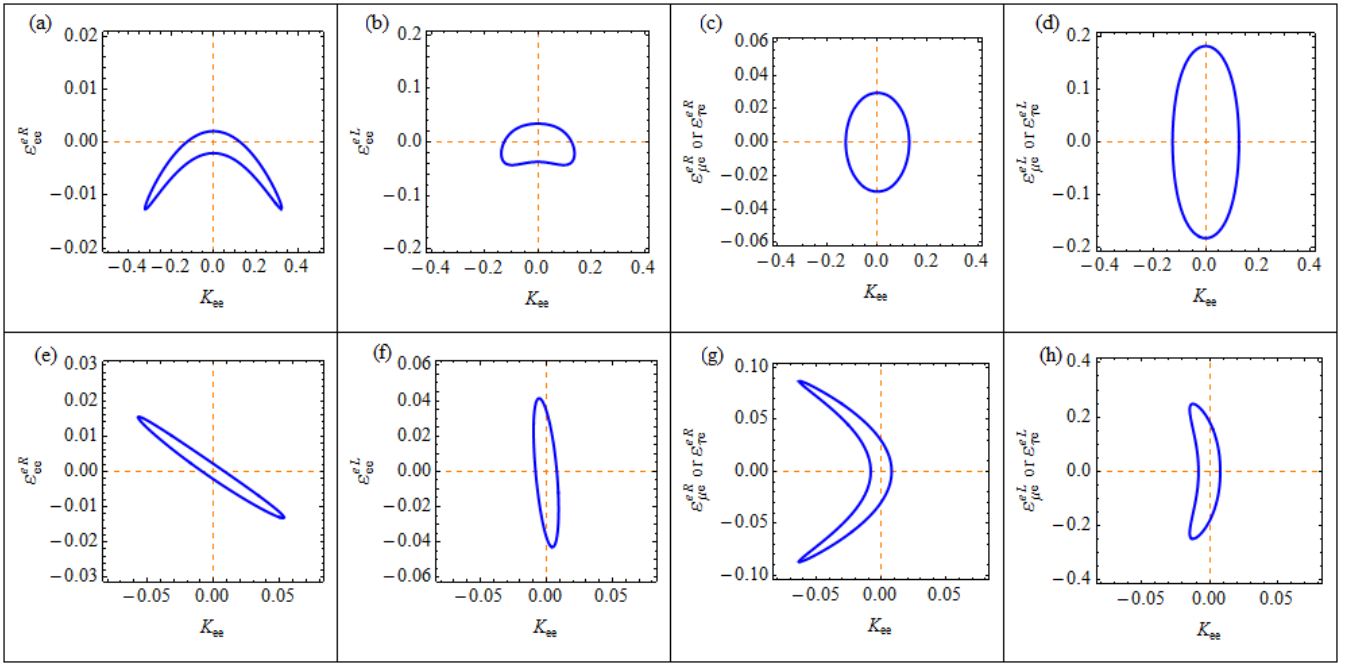


FIG. 4: C.L. boundary regions for *future prospects* data. **Upper Panel:** Correlation between the source NSI parameter ( $\text{Im } K_{ee}$ ) and the corresponding detector NSI parameters ( $\epsilon_{ee}^{R,L}$  and  $\epsilon_{\alpha e}^{R,L}$ , where  $\alpha = \mu$  or  $\tau$ ) at 90% C.L.. See the text for details. **Lower Panel:** Correlation between the source NSI parameter ( $\text{Re } K_{ee}$ ) and the corresponding detector NSI parameters ( $\epsilon_{\mu\mu}^{R,L}$ ,  $\epsilon_{\tau\tau}^{R,L}$  and  $\epsilon_{\alpha\mu}^{R,L}$ ,  $\epsilon_{\beta\tau}^{R,L}$  where  $\alpha = e$  or  $\tau$  and  $\beta = e$  or  $\mu$ ) at 90% C.L.. See the text for details.

Figure No.	NSI Parameters at Source	NSI Parameters at Detector
4(a)	$-0.33 < \text{Im } K_{ee} < 0.33$	$-0.013 < \epsilon_{ee}^{eR} < 0.002$
4(b)	$-0.14 < \text{Im } K_{ee} < 0.14$	$-0.045 < \epsilon_{ee}^{eL} < 0.036$
4(c)	$-0.13 < \text{Im } K_{ee} < 0.13$	$-0.03 < \epsilon_{\alpha e}^{eR} < 0.03$
4(d)	$-0.13 < \text{Im } K_{ee} < 0.13$	$-0.18 < \epsilon_{\alpha e}^{eL} < 0.18$
4(e)	$-0.057 < \text{Re } K_{ee} < 0.054$	$-0.013 < \epsilon_{ee}^{\mu R} < 0.016$
4(f)	$-0.01 < \text{Re } K_{ee} < 0.01$	$-0.043 < \epsilon_{ee}^{\mu L} < 0.042$
4(g)	$-0.064 < \text{Re } K_{ee} < 0.007$	$-0.086 < \epsilon_{\alpha e}^{\mu R} < 0.086$
4(h)	$-0.015 < \text{Re } K_{ee} < 0.008$	$-0.25 < \epsilon_{\alpha e}^{\mu L} < 0.25$

TABLE IV: Bounds at 90% C. L. obtained from Fig. (4) where  $\alpha = e$  or  $\tau$  and  $\beta = e$  or  $\mu$ .

Turning to the cases where the NSI can be active at *both* the source and detector, we study the parameter spaces of combined source-detector pairs in Fig. 4 and in accompanying Table IV. The sensitivity to the combinations improves typically by factors of 5 to 10 in both source and detector probes compared to the bounds shown in Fig. 2 and Table II. Comparing to current bounds in our Appendix B, for example, we find that the bound on  $\epsilon_{ee}^{eR}$  in entry 4(e) is a factor 10 below the bound given there, while the bound on  $\epsilon_{\tau e}^{eR}$  given in entry 4(g) is a factor of 5 below its bound quoted in [9]. In the case of NU  $K_{ee}$  couplings, the constraints are becoming competitive with those published [9], being within about a factor of 3 for both the imaginary part (top 4 rows) and the real part (bottom 4 rows) of  $K_{ee}$ . Looking at entry 4(c) or 4(d) in Table IV, we find that  $|\text{Im } K_{ee}| < 0.13$ , compared to the current best bound of 0.041, which is also the best bound for  $|\text{Im } K_{e\tau}|$ , compared to our bound of 0.1 shown in Table V. Thus, an upgraded TEXONO experiment could provide independent confirmation of the bounds on these parameters, but would not probe new parameter space in the search for new physics. Similarly, as shown in Fig. 5 and Table V, the bounds on the FC semileptonic parameters  $K_{e\mu}$  and  $K_{e\tau}$  achievable by an upgraded TEXONO experiment are within a factor 2 or 3 of the current bounds and possibly provide independent support, but not reach new regions in their parameter space.

Though the FC  $K_{e\alpha}$  vs.  $\epsilon_{\alpha\mu}$  or  $\epsilon_{\alpha\tau}$  studies, Fig. 5, provide no bounds on the  $\epsilon$ s because the “wrong flavor” source

neutrinos are zero in the  $K_{e\mu}$  or  $K_{e\tau} \rightarrow 0$  limit, the  $|\epsilon_{\alpha e}^{eR,L}|$  limits in Table IV apply to  $|\epsilon_{e\alpha}^{eR,L}|$  because of the Hermiticity constraint  $\epsilon_{\alpha\beta}^{eR,L} = (\epsilon_{\beta\alpha}^{eR,L})^*$ , as noted after Eq. (3).

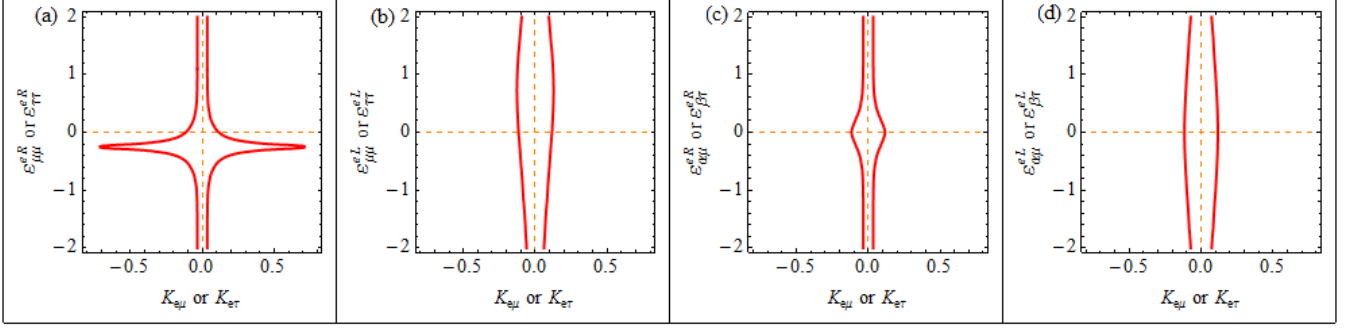


FIG. 5: Correlation between the source NSI parameter ( $\text{Re } K_{e\mu}$  and  $\text{Re } K_{e\tau}$ ) and the corresponding detector NSI parameters ( $\epsilon_{\mu\mu}^{R,L}$ ,  $\epsilon_{\tau\tau}^{R,L}$  and  $\epsilon_{\alpha\mu}^{R,L}$ ,  $\epsilon_{\beta\tau}^{R,L}$  where  $\alpha = e$  or  $\tau$  and  $\beta = e$  or  $\mu$ ) at 90% C.L.. See the text for details.

Figure No.	NSI Parameters at Source	NSI Parameters at Detector
5(a)	$-0.1 < K_{e\mu} \text{ or } K_{e\tau} < 0.1$	$\epsilon_{\mu\mu}^{eR}$ and $\epsilon_{\tau\tau}^{eR}$ are unbounded
5(b)	$-0.1 < K_{e\mu} \text{ or } K_{e\tau} < 0.1$	$\epsilon_{\mu\mu}^{eL}$ and $\epsilon_{\tau\tau}^{eL}$ are unbounded
5(c)	$-0.1 < K_{e\mu} \text{ or } K_{e\tau} < 0.1$	$\epsilon_{\alpha\mu}^{eR}$ and $\epsilon_{\beta\tau}^{eR}$ are unbounded
5(d)	$-0.1 < K_{e\mu} \text{ or } K_{e\tau} < 0.1$	$\epsilon_{\alpha\mu}^{eL}$ and $\epsilon_{\beta\tau}^{eL}$ are unbounded

TABLE V: Bounds obtained from Fig. (5) at 90% C. L. where  $\alpha = e$  or  $\tau$  and  $\beta = e$  or  $\mu$ . All the source NSI parameters  $K_{\alpha\beta}$  are either pure real or imaginary.

## VI. SUMMARY AND CONCLUSIONS

We have explored the consequences of adding new physics effects at the reactor source in a  $\bar{\nu} + e \rightarrow \bar{\nu} + e$  scattering experiment. We have used the data from the TEXONO experiment and also a model data based on their projected improved sensitivity in a future upgrade. This experiment has the virtue that its 30m baseline does not allow for oscillation effects at the detector, so that any new physics at the source is not degenerate with oscillation effects during propagation. After developing the needed framework in Secs. II and III, where we explicitly include the NSI phases in the FC leptonic, detector parameters, we reviewed the 90% C.L. boundaries presented in Ref. [11] in Sec IV, but including the phase effects on the boundary in the FC,  $\epsilon_{e\mu}^{eR} - \epsilon_{e\mu}^{eL}$  parameter space. We checked that we properly reproduce the boundaries and the value, and statistical error of the TEXONO examples, but adding the small but noticeable dependence on the choice of phases for the FC detector NSI parameters, filling a gap in the literature. The effects on the bounds one derives is at the 5% level. In lower energy experiments with sufficient statistics, this phase effect may be more striking as the coefficient of the correlation term becomes larger relative to the other terms contributing to the rate.

Including the NSI at the reactor source, we surveyed examples of the interplay between the source and detector effects with a series of source vs. detector 90% C.L. boundaries based on the TEXONO data. We find that the right (R-) parameter bounds on the detector NSI parameters  $\epsilon_{\alpha e}^{eR}$ ,  $\alpha = e, \mu$  and  $\tau$ , are about the same as the current best bounds, as summarized from Ref. [9] in our Appendix B, but the corresponding left (L-) parameter bounds are factors 5 - 10 larger. All of the bounds on the source,  $K_{\alpha\beta}$  parameters are one-to-two orders of magnitude larger than best current bounds. Because the source FC parameters must be non-zero for a bound on the detector parameters  $\epsilon_{\alpha e}^{eL,R}$  to exist, no meaningful bounds can be placed independently on the latter, but they differ only by a phase from the  $\epsilon_{\alpha e}^{eL,R}$  parameters, as noted after Eq. (3), so the bounds on detector parameters listed in rows 2(c), and 2(d) in Table 2 apply as well to the detector parameters in rows 2(g) and 2(h) when  $\alpha$  and  $\beta = e$ .

Turning to the companion study of our model data based on the estimated future improvements in an upgraded TEXONO experiment, we basically repeated the exercises of Secs. IV and V to survey the parameter spaces in

anticipation of this upgrade. Compared to the bounds based on current data our estimates of future, high sensitivity data show that an order of magnitude increase in the level of sensitivity to source and detector NSI parameters is achievable compared to the sensitivity with the current TEXONO data. This brings the bounds on detector NSI parameters well below current bounds in all but the case of  $\epsilon_{e\mu}^{eL}$ , which is the same as the current bound. Our new approach to bounding the CC, semileptonic NSI at the source results in projected bounds that are comparable to the current ones. The very feature that makes this class of ultra-short-baseline experiments especially clean for probing the source NSI parameters, namely the lack of interference with neutrino mixing amplitudes, makes it less sensitive. The parameters of interest appear as the modulus *squared* in the FC case, while in the NU case, the interference with the SM contribution gives a boost to the sensitivity to the real part of  $K_{ee}$ , which has a very tight bound already, coming from CKM unitarity or lepton universality and for the same reason.

To conclude, we see that the currently envisaged upgrade to the TEXONO experiment promises to probe an order of magnitude deeper into the right-handed, leptonic NSI parameter space. To improve the sensitivity to the left-handed, leptonic NSI couplings, high intensity, short baseline  $\nu_e$  experiments with large targets, along the lines of the LENA project [33, 34] will be needed. To delve deeper into the semileptonic, CC parameter space with a reactor, antineutrino source, a third generation of the TEXONO type of experiment would be needed, since we find that the current plans would only bring bounds to the level of those currently available. Otherwise, oscillation experiments with interference between the relevant NSI parameters and oscillation amplitudes involving standard oscillation parameters, independently measured and known to high accuracy would be needed, as remarked in Ref. [13].

## VII. APPENDIX

### A. Reactor neutrino spectrum and event rate: the TEXONO Experiment at Kuo-Sheng

The reactor antineutrinos spectrum produced at Kuo-Sheng Nuclear Reactor is given in Fig. 6. We find the follow-

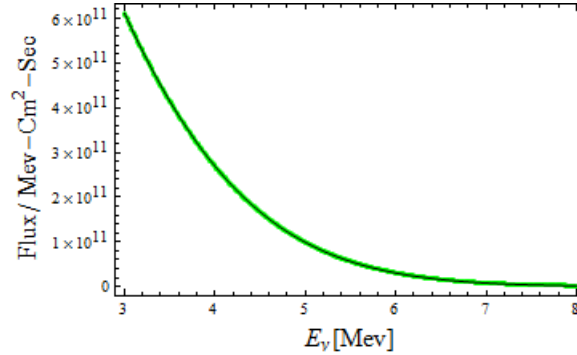


FIG. 6: Typical antineutrino spectrum at 28m from core at Kuo-Sheng. The green curve is the data and the black curve inside it is the fit.

ing fit function for the reactor neutrino spectrum between 3 Mev to 8 Mev,

$$\frac{d\phi(\bar{\nu}_e)}{dE_\nu} = \sum_0^6 \frac{a_n}{(E_\nu)^n} \quad (9)$$

where the fit parameters  $a_0, a_1 \dots a_6$  have the values given in Table VI,

$a_0$	$a_1$	$a_2$	$a_3$	$a_4$	$a_5$	$a_6$
$-1.23779 \cdot 10^{12}$	$3.72889 \cdot 10^{13}$	$-4.38337 \cdot 10^{14}$	$2.52571 \cdot 10^{15}$	$-7.4559 \cdot 10^{15}$	$1.11498 \cdot 10^{16}$	$-6.74817 \cdot 10^{15}$

TABLE VI: Values of the fit parameters for the neutrino spectrum.

The experimentally observed event rate ( $R_E$ ) is then compared with the theoretically modeled or expected event rate ( $R_X$ ). The differential rate with respect to  $T$ , kinetic energy of the recoil electron, is

$$\frac{dR_X}{dT} = \rho_e \int_T^{E_\nu^{max}} \mathcal{F}(E_\nu) \frac{d\phi(E_\nu)}{dE_\nu} dE_\nu, \quad (10)$$

so the rate integrated over the  $i^{th}$  bin in  $T$  is

$$R_X^i = \int_{T(i)}^{T(i+1)} \frac{dR_X}{dT}. \quad (11)$$

Here  $\rho_e$  is the electron number density per kg of target mass of CsI(Tl), and  $\frac{d\phi(E_\nu)}{dE_\nu}$  is the neutrino spectrum as given in Eq. (9) and  $\mathcal{F}(E_\nu)$  is the factor containing the NSI detector cross sections and the corresponding NSI source parameter coefficients, as given in Eq. (7).

We use the following definition of  $\chi^2$  from Ref. [11] to perform the minimum- $\chi^2$  fit,

$$\chi^2 = \sum_i \left( \frac{R_E^i - R_X^i}{\Delta_{stat}^i} \right)^2, \quad (12)$$

where  $R_E^i$  and  $R_X^i$  are the experimental and expected event rates over the  $i^{th}$  data bin and  $\Delta_{stat}^i$  is the corresponding statistical uncertainty of the measurement.

### B. Bounds of Ref. [9]

NSI Parameters at Source	NSI Parameters at Detector	
$ K_{ee}  < 0.041$	$ \epsilon_{ee}^{eR}  < 0.14,$	$ \epsilon_{ee}^{eL}  < 0.06$
—	$ \epsilon_{e\mu}^{eR}  < 0.10,$	$ \epsilon_{e\mu}^{eL}  < 0.10$
—	$ \epsilon_{e\tau}^{eR}  < 0.27,$	$ \epsilon_{e\tau}^{eL}  < 0.4$
$ K_{e\mu}  < 0.025$	$ \epsilon_{\mu e}^{eR}  < 0.10,$	$ \epsilon_{\mu e}^{eL}  < 0.10$
—	$ \epsilon_{\mu\mu}^{eR}  < 0.03,$	$ \epsilon_{\mu\mu}^{eL}  < 0.03$
—	$ \epsilon_{\mu\tau}^{eR}  < 0.10,$	$ \epsilon_{\mu\tau}^{eL}  < 0.10$
$ K_{e\tau}  < 0.041$	$ \epsilon_{\tau e}^{eR}  < 0.27,$	$ \epsilon_{\tau e}^{eL}  < 0.4$
—	$ \epsilon_{\tau\mu}^{eR}  < 0.10,$	$ \epsilon_{\tau\mu}^{eL}  < 0.10$
—	$ \epsilon_{\tau\tau}^{eR}  < 0.4,$	$ \epsilon_{\tau\tau}^{eL}  < 0.16$

TABLE VII: Bounds at 90% C.L. taken from Eq. (44) and Eq. (45) of Ref. [9] for comparison. Notice that we have used our notation for their bounds for convenience. It should be noted that there is a separate upper bound  $\text{Re } K_{ee} \sim 10^{-3}$  from the CKM unitarity and lepton universality constraints.

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