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# Reactor Antineutrino Anomaly with known $\theta_{13}$

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We revisit the reactor antineutrino anomaly using the recent reactor flux independent determination of sizable  $\theta_{13}$  by considering the full set of the absolute reactor  $\bar{\nu}_e$  flux measurements. When normalized to the predicted flux of Mueller *et al.* [1], the new world average, after including results from Palo Verde, Chooz, and Double Chooz, is  $0.959 \pm 0.009$  (experiment uncertainty)  $\pm 0.027$  (flux systematics). Including the data with kilometer baseline, the new world average is only about  $1.4\sigma$  lower than the unity, weakening the significance of the reactor antineutrino anomaly. The upcoming results from Daya Bay, RENO, and the Double Chooz will provide further information about this issue.

## I. INTRODUCTION

The term ‘‘reactor anomaly’’ was coined by Mention *et al.* [2] who noted that the average of the experimentally determined reactor antineutrino flux at reactor-detector distances  $< 100$  m accounts for only  $0.943 \pm 0.023$  of the reevaluated theoretical expectation of Ref. [1]. In addition to the 19 experimental results obtained with detectors distant less than 100 m from the reactor source, we are able now to include in the analysis also the results of the Chooz [3, 4] and Palo Verde [5] as well as of the Double Chooz [6–8] experiments, where the detectors were

further away from the reactor complex. In these cases, the corresponding experimental results need be corrected for the flux loss associated with the known value of the mixing angle  $\theta_{13}$ , which was determined in a model independent way by comparing the count rates in two essentially identical, but separated in distance, detectors. We use the value  $\sin^2 2\theta_{13} = 0.089 \pm 0.011$  obtained in the Daya Bay experiment [9–11], and confirmed by the RENO experiment [12]. The corresponding correction is easy to apply using the formula for the survival probability

$$P_{sur} = 1 - \sin^2 2\theta_{13}(\cos^2 \theta_{12} \sin^2 \Delta_{31} + \sin^2 \theta_{12} \sin^2 \Delta_{32}) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \Delta_{21}, \quad (1)$$

with  $\Delta_{ij} \equiv |\Delta_{ij}| = 1.27|\Delta m_{ij}^2| \frac{L(m)}{E(\text{MeV})}$ . Values of mixing angles and mass-squared differences used in the simulation are taken from Ref. [13] assuming normal mass hierarchy ( $\Delta m_{31}^2 = \Delta m_{32}^2 + \Delta m_{21}^2$ ):

$$\begin{aligned} \sin^2 2\theta_{12} &= 0.857 \pm 0.024 \\ \Delta m_{21}^2 &= (7.50 \pm 0.20) \times 10^{-5} eV^2 \\ \Delta m_{32}^2 &= (2.32 \pm 0.12) \times 10^{-3} eV^2. \end{aligned} \quad (2)$$

## II. ANALYSIS DESCRIPTION

In the following, we will explain in detail the inputs to our analysis as well as the chi-square method. For the 19 experimental results obtained with detectors distant less than 100 m from the reactor source, the measured fluxes normalized to the prediction of Ref. [1] (noted as ‘‘ratio’’) are taken from Ref. [2] and tabulated in Table I, together with the detector technology, fission fractions,

distance to the reactor core, and year of publication. The average survival probability  $P_{sur}$  for each experiment is calculated by integrating over the neutrino antineutrino spectrum [1] convoluted with the inverse beta decay cross section with  $\sin^2 2\theta_{13} = 0.089$ . There are two uncertainties listed in Table I. The  $\sigma_{err}$  represents the total uncertainty on the ratio. The  $\sigma_{corr}$  represents the part of uncertainty correlated among different experiments. In particular, there is a 2.7% uncertainty in  $\sigma_{corr}$  coming from the uncertainty in the predicted reactor flux. (In first versions of Ref. [2], the reactor flux uncertainty was assumed to be 2.7%, which was replaced by 2% in the final version. In this analysis, we chose the more conservative number (2.7%).) Since it will not affect the relative differences among different measurements, we define the reduced uncertainties:

$$\begin{aligned} \sigma_{err}^{reduced} &= \sqrt{\sigma_{err}^2 - 2.7^2} \\ \sigma_{corr}^{reduced} &= \sqrt{\sigma_{corr}^2 - 2.7^2} \end{aligned} \quad (3)$$

after removing the 2.7% overall normalization uncertainty. We change the  $\sigma_{corr}$  between SRP-I and SRP-II [14] from 3.7% (originally quoted in Ref. [2]) to 2.7% (the reactor flux uncertainty only). The original 3.7% assumes almost full correlation between SRP-I and SRP-II, which can not explain the apparent differences be-

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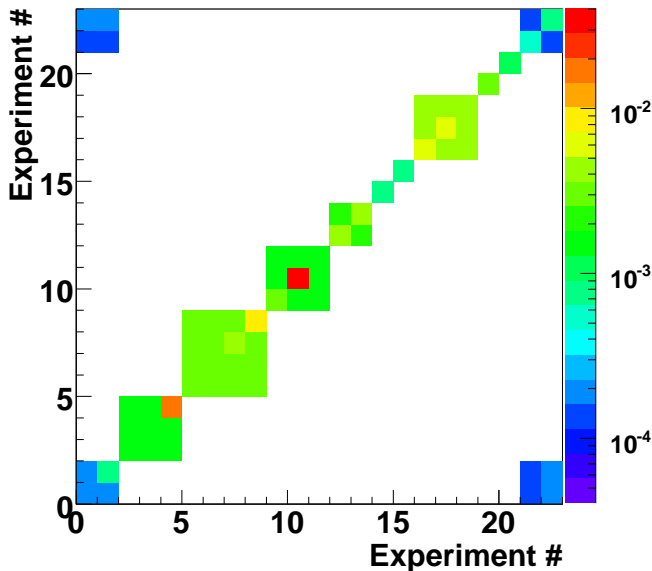


FIG. 1. (color online) The covariance matrix of the reduced experimental uncertainties. The number of experiments can be found in Table. I. The off-diagonal terms show the correlation among different experiments.

tween the two ratios (0.952 vs. 1.018), indicating less correlation between the two experiments. For the same reason, we reduce the  $\sigma_{corr}$  between the ROVNO88-II and ROVNO88-2I from 6.9% to 5.7%, which is the final reported uncertainty for ROVNO88 experiments [15].

There are three reactor cores in the Palo Verde experiment [5]. The distances between detector and each reactor core are 750 m, 890 m, and 890 m. In calcu-

lating the average survival probability  $P_{sur}$ , we assume that all three reactor cores have equal power. The result is compared with  $P_{sur}^{750m}$  assuming full power in only the 750 m reactor and  $P_{sur}^{890m}$  assuming full power in the 890 m reactors. The differences are quoted as an additional uncertainty, which is only about 5% of the total reduced experimental uncertainty.

There are two reactor cores in Chooz experiment [3, 4]. The distances between the detector and each reactor core are 998 m and 1115 m. Similar procedure is applied to calculate the uncertainty for the equal power assumption. The resulting uncertainty is about 6.2% of the total reduced experimental uncertainty. The fission fractions are assumed to be the same as those from Double Chooz [7]. We also calculated the average  $P_{sur}$  by varying these fission fractions. The differences are negligible.

The Double Chooz experiment is conducted at the same location as Chooz. With a single detector, the recent rate-only analyses of the data from delayed neutron capture on Gadolinium (n-Gd) and delayed neutron capture on hydrogen (n-H) reported the value of  $\sin^2 2\theta_{13} = 0.170 \pm 0.052$  [7] and  $\sin^2 2\theta_{13} = 0.044 \pm 0.060$  [8]<sup>1</sup>, by anchoring to the short-baseline Bugey-4 results [16], respectively. Although the measured flux normalized to the prediction of Ref. [1] has not been reported, we can deduce such ratios using the reported fission fractions [7], the reported values of  $\sin^2 2\theta_{13}$  [7, 8], and the Bugey-4 results. The  $\sigma_{err}^{reduced}$  is dominated by the uncertainties of reported  $\sin^2 2\theta_{13}$ , with additional uncertainties coming from the equal power assumption. The  $\sigma_{corr}^{reduced}$  are calculated from the reduced experimental uncertainty  $\sigma_{err}^{reduced}$  from Bugey-4. Furthermore, there are additional correlated uncertainties between the n-H and n-Gd measurements due to the equal power assumption. The final covariance matrix  $W$  using reduced uncertainties is shown in Fig. 1.

The  $\chi^2$  function used in this analysis is constructed as follows:

$$\chi^2(r, \sin^2 2\theta_{13}) = (r \cdot \vec{P}_{sur}(\sin^2 2\theta_{13}) - \vec{R})^T W^{-1} (r \cdot \vec{P}_{sur}(\sin^2 2\theta_{13}) - \vec{R}) + \frac{(\sin^2 2\theta_{13} - 0.089)^2}{0.011^2}. \quad (4)$$

Here,  $W^{-1}$  is the inverted covariance matrix. The vector  $\vec{R}$  contains the reported ratios from all 23 experiments (tabulated in Table. I). The absolute normalization ratio  $r$  is treated as a free parameter. The vector  $\vec{P}_{sur}$  contains the predicted average survival probabilities given a value of  $\sin^2 2\theta_{13}$ . The values of  $\vec{P}_{sur}$  using  $\sin^2 2\theta_{13} = 0.089$  are tabulated in Table. I. The last term in Eq. (4) repre-

sents the constrain on  $\sin^2 2\theta_{13}$  from the latest Daya Bay results [11].

### III. RESULTS

In the following three figures, we show the results of all 23 measurements and the deduced ratios after minimizing the  $\chi^2$  defined in Eq. (4). The global average is determined to be  $0.959 \pm 0.009$ . In Fig. 2, the results are shown in analogous way as in Ref. [2], i.e. as a function of the distance from the corresponding reactor core. We combine results at same baseline together for clarity.

<sup>1</sup> To be consistent with other experiments, we choose the rate-only  $\sin^2 2\theta_{13}$  results, which represent simple measures of the disappearance in the total number of events.

The corresponding  $\chi^2/\text{Dof} = 23.8/22$ . We stress that our error bars do not include the reactor flux uncertainty (2.7%), hence they appear smaller than those in [2].

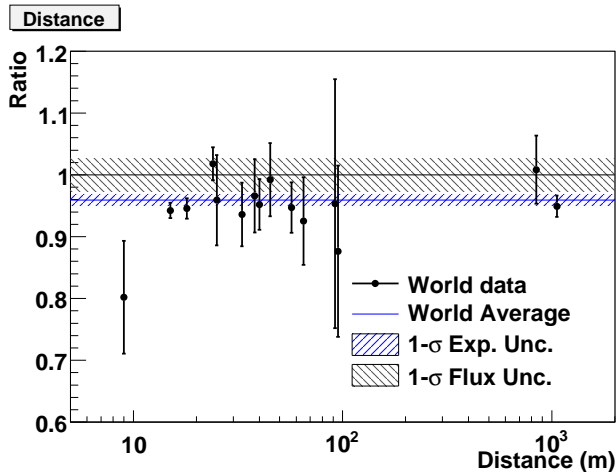


FIG. 2. (color online) The reactor  $\bar{\nu}_e$  capture rate as a function of the distance from the reactor, normalized to the theoretical flux of Ref. [1]. The horizontal bar represents the global average and its  $1\sigma$  error bar. The 2.7% reactor flux uncertainty is shown as a band around unity. We combine results at same baseline (e.g. Chooz, Double Chooz n-H and n-Gd results) together for clarity.

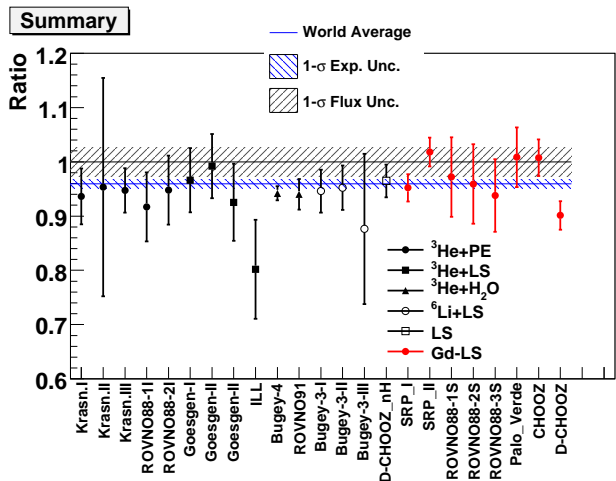


FIG. 3. (color online) The same 23 experimental results as in previous figures are plotted by the detection method employed. See captions of Fig. 2 for details.

The new global average is somewhat larger than the 0.943 value of Ref.[2] quoted earlier, weakening the significance of the reactor antineutrino anomaly. There are two reasons for this difference. First, we include more recent and more distant experiments, of which Palo Verde and Chooz have larger rates. Second, we change the correlated uncertainty between SRP-I and SRP-II from 3.7% to 2.7% (the reactor flux uncertainty only), since the orig-

inal 3.7% assumes almost full correlation between SRP-I and SRP-II, which can not explain the apparent differences between the two ratios (0.952 vs. 1.018). With fixed total experimental uncertainties, this change effectively increases the significance of SRP experiments and leads to about 1% larger world average.

In addition, our results are larger than those reported in Ref. [17] and Ref. [18], which also include the kilometer experiments with known  $\theta_{13}$ . The result reported in Ref. [17] included the Gallium neutrino data [19–25], which was not included in our reactor antineutrino analysis. They also did an analysis by including a RENO preliminary result from the absolute flux analysis. However, such analysis of the RENO experiment has not been, to our knowledge, released and is not finished as yet [26]. In Ref. [18], the measured experimental fluxes are normalized to the predicted flux of Huber [27] with a new neutron lifetime 881.5s (2011 update of PDG [13]). The change in reactor flux model and the neutron lifetime leads to in average 1.6% lower ratios than what we used in this work (tabulated in Table. I). The rest of differences come from the treatment in the correlated uncertainty of SRP experiments and the uncertainty of the reactor flux prediction (2% used in Ref. [18] vs. 2.7% used in this work).

One of the main purposes of this work is to illustrate the impact of kilometer experiments to the results of Mention *et al.* [2]. Therefore, we have adapted the same neutron life time (885.7s) as in Mention *et al.* [2]. The current recommended neutron lifetime from the 2012 Particle Data Group [13] is 880.1s. Using the latest neutron life time would lead to about 0.63% reduction in the average ratio.

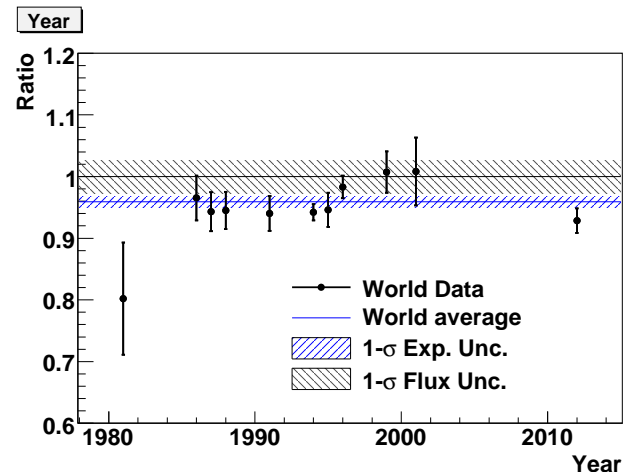


FIG. 4. (color online) The reactor  $\bar{\nu}_e$  capture rate as a function of the year when the measurement was published, normalized to the theoretical flux of Ref. [1]. We combine results at same year (e.g. Double Chooz n-H and n-Gd results) together for clarity. See captions of Fig. 2 for details.

Fig. 3 shows results from all 23 experiments again arranged by the detector technology. Five different tech-

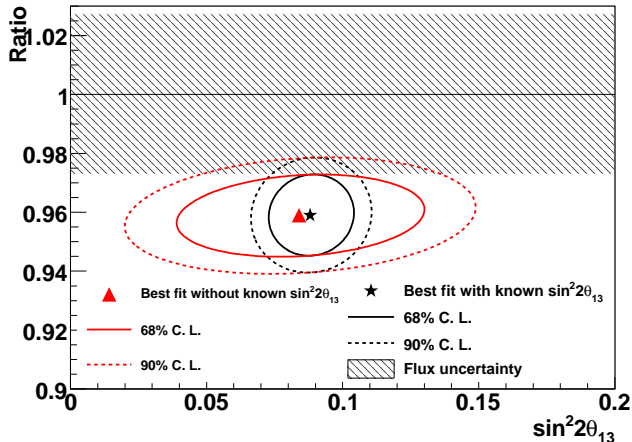


FIG. 5. (color online) Best fits are shown with and without  $\sin^2 2\theta_{13}$  constrains from Daya Bay. The corresponding 68% ( $\Delta\chi^2 < 2.3$ ) and 90% ( $\Delta\chi^2 < 4.6$ ) confidence intervals are shown as well.

nologies were used in the 23 experiments to record the  $\bar{\nu}_e$  capture on protons. In the ILL [28] and Goesgen experiments [29] the liquid scintillator targets cells were interspaced with the  $^3\text{He}$  neutron counters. In the Bugey-3 experiments [30, 31] the liquid scintillator was loaded with  $^6\text{Li}$  to detect the neutron captures. In the Bugey-4 [16] and Rovno91 [32] only the neutron captures were detected. The detector in these two experiments consisted of a water target with embedded  $^3\text{He}$  detectors. In the Krasnoyarsk [33] and Rovno88 [15] experiments again only the total capture rates were measured. The detectors consisted of the polyethylene neutron moderator with  $^3\text{He}$  neutron counters embedded in them. Finally, the Savannah River experiments [14], the Rovno88 [15] and the Chooz, Palo Verde and Double Chooz experiments [7]<sup>2</sup> use the Gd loaded liquid scintillators. The Daya Bay and RENO experiments are employing that technology as well. The results from the Gd loaded scintillator experiments are, with the exception of Double Chooz, in general higher than the new world average.

We should also note that the experiments with Gd loaded scintillator were in general carried out at a later time. In Fig. 4 we show the same data, but now arranged as a function of the year when the corresponding measurements were published. One can see that the more recent experiments, except the Double Chooz, appear to have higher rates than the earlier ones. If such tendency is true, it could either be due to difference in detector technologies or improvement in data acquisition or analysis methods. However, the  $\chi^2$  of the global fit,  $\chi^2/\text{Dof} = 23.8/22$ , signifies that the 23 reactor flux determina-

tions are mutually consistent. Therefore, our observation might be simply due to statistical fluctuations.

In addition, we also use these 23 experiments to extract  $\sin^2 2\theta_{13}$  by minimizing the  $\chi^2$  in Eq. (4) without the penalty term ( $\frac{(\sin^2 2\theta_{13} - 0.089)^2}{0.011^2}$ ). The best-fit  $\sin^2 2\theta_{13}$  is determined to be  $0.084 \pm 0.030$  (one dimension:  $\Delta\chi^2 \equiv \chi^2 - \chi^2_{\text{minimum}} < 1$ ). The non-zero  $\sin^2 2\theta_{13}$  is dominated by the latest Double Chooz results, and is also consistent with the latest Daya Bay results [11]. Fig. 5 shows the best-fit  $\sin^2 2\theta_{13}$  and ratio  $r$  with and without the Daya Bay constrains. The corresponding 68% (two dimensions:  $\Delta\chi^2 < 2.3$ ) and 90% (two dimensions:  $\Delta\chi^2 < 4.6$ ) confidence intervals are shown as well. The correlation between the ratio and  $\sin^2 2\theta_{13}$  is rather weak due to the strong correlation between results from Double Chooz, which dominate the non-zero  $\sin^2 2\theta_{13}$  extraction from these 23 experiments, and result from Bugey-4, which is the most precise short baseline measurement. Therefore, it is not surprising that the precise  $\sin^2 2\theta_{13}$  value from Daya Bay, which is very close to the best fit  $\sin^2 2\theta_{13}$  value from these 23 experiments, does not improve the knowledge on the ratio.

#### IV. SUMMARY AND DISCUSSION

With the known  $\theta_{13}$ , we include results from Palo Verde, Chooz, and Double Chooz, and re-evaluate the reactor antineutrino anomaly. The new world average is determined to be  $0.959 \pm 0.009$  (experiment uncertainty)  $\pm 0.027$  (flux systematics), when normalized to the predicted flux of Mueller *et al.* [1]. The new world average is about  $1.4\sigma$  lower than unity, weakening the significance of the reactor antineutrino anomaly. We also show that the analysis of all 23 experiments yields  $\sin^2 2\theta_{13} = 0.084 \pm 0.030$  in agreement with its independently obtained value from the Daya Bay and RENO experiments.

The forthcoming absolute flux analysis of the Daya Bay, RENO, and Double Chooz experiments will clearly shed more light on this issue. The final answer to the question of agreement between the predicted reactor  $\bar{\nu}_e$  flux and the measured rate clearly depends on the results of these experiments as well as on the careful analysis of the predicted flux and its uncertainties.

The most popular explanation of the anomaly, if it turns out that it is statistically significant, is the existence of additional sterile neutrinos. The consistency of the rate of the short baseline neutrino experiments, demonstrated here and also e.g. in Ref. [2, 18] suggests that such additional neutrinos must have large enough masses that the corresponding oscillation length is at most few meters. Experiments sensitive to such short oscillation length would be able to convincingly prove the existence of the sterile neutrino and allow determination of their masses and mixing angles.

<sup>2</sup> In Double Chooz n-H analysis [8], the neutrino events are largely generated in the liquid scintillator region without the Gd loading.

#	result	Det. type	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	ratio	$\sigma_{err}(\%)$	$\sigma_{corr}(\%)$	L(m)	$P_{sur}$	Year
1	Bugey-4	$^3\text{He} + \text{H}_2\text{O}$	0.538	0.328	0.078	0.056	0.942	3.0	3.0	15	0.999987	1994
2	ROVNO91	$^3\text{He} + \text{H}_2\text{O}$	0.614	0.274	0.074	0.038	0.940	3.9	3.0	18	0.999981	1991
22	Double Chooz	Gd-LS	0.496	0.351	0.087	0.066	0.860	3.7	3.0	998-1115	0.954	2012
23	Double Chooz	LS (n-H)	0.496	0.351	0.087	0.066	0.920	4.0	3.0	998-1115	0.954	2012
3	Bugey-3-I	$^6\text{Li} - \text{LS}$	0.538	0.328	0.078	0.056	0.946	4.8	4.8	15	0.999987	1995
4	Bugey-3-II	$^6\text{Li} - \text{LS}$	0.538	0.328	0.078	0.056	0.952	4.9	4.8	40	0.999907	1995
5	Bugey-3-III	$^6\text{Li} - \text{LS}$	0.538	0.328	0.078	0.056	0.876	14.1	4.8	95	0.999479	1995
6	Goesgen-I	$^3\text{He} + \text{LS}$	0.620	0.274	0.074	0.042	0.966	6.5	6.0	38	0.999916	1986
7	Goesgen-II	$^3\text{He} + \text{LS}$	0.584	0.298	0.068	0.050	0.992	6.5	6.0	45	0.999883	1986
8	Goesgen-III	$^3\text{He} + \text{LS}$	0.543	0.329	0.070	0.058	0.925	7.6	6.0	65	0.999756	1986
9	ILL	$^3\text{He} + \text{LS}$	$\approx 1$	-	-	-	0.802	9.5	6.0	9	0.999995	1981
10	Krasn. I	$^3\text{He} + \text{PE}$	$\approx 1$	-	-	-	0.936	5.8	4.9	33	0.999937	1987
11	Krasn. II	$^3\text{He} + \text{PE}$	$\approx 1$	-	-	-	0.953	20.3	4.9	92	0.999511	1987
12	Krasn. III	$^3\text{He} + \text{PE}$	$\approx 1$	-	-	-	0.947	4.9	4.9	57	0.999812	1987
13	SRP-I	Gd-LS	$\approx 1$	-	-	-	0.952	3.7	2.7	18	0.999981	1996
14	SRP-II	Gd-LS	$\approx 1$	-	-	-	1.018	3.8	2.7	24	0.999967	1996
15	ROVNO88-1I	$^3\text{He} + \text{PE}$	0.607	0.277	0.074	0.042	0.917	6.9	5.7	18	0.999981	1988
16	ROVNO88-2I	$^3\text{He} + \text{PE}$	0.603	0.276	0.076	0.045	0.948	6.9	5.7	18	0.999981	1988
17	ROVNO88-1S	Gd-LS	0.606	0.277	0.074	0.043	0.972	7.8	7.2	18	0.999981	1988
18	ROVNO88-2S	Gd-LS	0.557	0.313	0.076	0.054	0.959	7.8	7.2	25	0.999964	1988
19	ROVNO88-3S	Gd-LS	0.606	0.274	0.074	0.046	0.938	7.2	7.2	18	0.999981	1988
20	Palo Verde	Gd-LS	0.60	0.27	0.07	0.06	0.975	6.0	2.7	750-890	0.967	2001
21	Chooz	Gd-LS	0.496	0.351	0.087	0.066	0.961	4.2	2.7	998-1115	0.954	1999

TABLE I. Tabulated results of all 23 experiments. Experiments are categorized into different groups with horizontal lines. Within each group, the  $\sigma_{corr}$  represent the correlated uncertainties among different experiments. This table is an extension of Table. II of Ref. [2]. There are additional correlated uncertainties, since Double Chooz results were anchored to the Bugey-4. See the text for more explanations.

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- [1] Th. A. Mueller et al. *Phys. Rev.*, **C83**:054615, 2011.  
 [2] G. Mention et al. *Phys. Rev.*, **D83**:073006, 2011.  
 [3] M. Apollonio et al. *Phys. Lett.*, **B466**:415, 1999.  
 [4] M. Apollonio et al. *Eur. Phys. J.*, **C27**:331, 2003.  
 [5] F. Boehm et al. *Phys. Rev.*, **D64**:112001, 2001.  
 [6] Y. Abe et al. *Phys. Rev. Lett.*, **108**:131801, 2012.  
 [7] Y. Abe et al. *Phys. Rev.*, **D86**:052008, 2012.  
 [8] Y. Abe et al. arXiv:1301.2948 (2013).  
 [9] F. P. An et al. *Phys. Rev. Lett.*, **108**:171803, 2012.  
 [10] F. P. An et al. *Nucl. Inst. Method*, **A685**:78, 2012.  
 [11] F. P. An et al. *Chinese Phys.*, **C37**:011001, 2013.  
 [12] J. K. Ahn et al. *Phys. Rev. Lett.*, **108**:191802, 2012.  
 [13] J. Beringer et al. *Phys. Rev.*, **D86**:010001, 2012.  
 [14] Z. D. Greenwood et al. *Phys. Rev.*, **D53**:6054, 1996.  
 [15] A. O. Afonin et al. *Sov. Phys. JETP*, **67**:213, 1988.  
 [16] Y. Declais et al. *Phys. Lett.*, **B338**:383, 1994.  
 [17] E. Ciuffoli, J. Evslin, and H. Li. *JHEP*, **12**:110.  
 [18] M. C. Gonzalez-Garcia, M. Maltoni, J. Salvado, and T. Schwetz. *JHEP*, **1212**:123, 2012.  
 [19] P. Anselmann et al. *Phys. Lett.*, **B342**:440, 1995.  
 [20] W. Hampel et al. *Phys. Lett.*, **B420**:114, 1998.  
 [21] J. N. Abdurashitov et al. *Phys. Rev. Lett.*, **77**:4708, 1996.  
 [22] J. N. Abdurashitov et al. *Phys. Rev.*, **C59**:2246, 1999.  
 [23] J. N. Abdurashitov et al. *Phys. Rev.*, **C73**:045805, 2006.  
 [24] J. N. Abdurashitov et al. *Phys. Rev.*, **C80**:015807, 2009.  
 [25] C. Giunti and M. Laveder. *Phys. Rev.*, **C83**:065504, 2011.  
 [26] Soo-Bong Kim, *private communication*.  
 [27] P. Huber. *Phys. Rev.*, **C84**:024617, 2011.  
 [28] H. Kwon et al. *Phys. Rev.*, **D24**:1097, 1981.  
 [29] G. Zacek et al. *Phys. Rev.*, **D34**:2621, 1986.  
 [30] B. Achkar et al. *Nucl. Phys.*, **B534**:503, 1995.  
 [31] B. Achkar et al. *Phys. Lett.*, **B374**:243, 1996.  
 [32] A. I. Alfonin et al. *JETP*, **67**:213, 1998.  
 [33] G. S. Vidyakin et al. *JETP Lett.*, **59**:390, 1994.