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Phys. Rev. D **83**, 046012 — Published 25 February 2011

DOI: [10.1103/PhysRevD.83.046012](https://doi.org/10.1103/PhysRevD.83.046012)

Dynamical Gap and Cuprate-like Physics from Holography

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We study the properties of fermion correlators in a boundary theory dual to the Reissner-Nordström AdS_{d+1} background in the presence of a bulk dipole (Pauli) interaction term with strength p . We show that by simply changing the value of the parameter p we can tune continuously from a Fermi liquid (small p), to a marginal Fermi liquid behavior at a critical value of p , to a generic non-Fermi liquid at intermediate values of p , and finally to a Mott insulator at large values of the bulk Pauli coupling. As all of these phases are seen in the cuprate phase diagram, the holographic model we study has the key elements of the strong coupling physics typified by Mott systems. In addition, we extend our analysis to finite temperature and show that the Mott gap closes. Of particular interest is that it closes when the ratio of the gap to the critical temperature is of the order of ten. This behavior is very much similar to that observed in the classic Mott insulator VO_2 . We then analyze the non-analyticities of the boundary theory fermion correlators for generic values of frequency and momentum by calculating the quasi-normal modes of the bulk fermions. Not surprisingly, we find no evidence for the dipole interaction inducing an instability in the boundary theory. Finally, we briefly consider the introduction of superconducting condensates, and find that in that case, the fermion gap is driven by scalar-fermion couplings rather than by the Pauli coupling.

PACS numbers:

I. INTRODUCTION

Holography can offer unprecedented insight into the dynamics of strongly coupled systems. In the recent past, it has become clear that the domain of applicability of holography goes beyond high energy physics and includes strongly-correlated systems in condensed matter physics, as well (see [1–4] for reviews). Since in most cases of interest, we do not possess a microscopic understanding of the field theory dynamics under study, a phenomenological point of view (the so-called “bottom-up” approach) is taken where a minimal gravitational setup is devised for analyzing a specific strong-coupling feature of a system. In some situations where string or M-theory completion of a bottom-up construction is known, one may wonder how a bottom-up result is modified in a top-down approach. For example, a quantity of interest in strongly-correlated condensed matter systems, which can easily be computed using holography, is the fermion spectral function (which is proportional to the imaginary part of the fermion retarded two-point function). There has recently been much discussion about this quantity in the holographic literature [5–9] where analyzing the Dirac equation for a charged probe fermion propagating in a gravitational background (usually a charged black hole), one can show that the retarded two-point function of the dual fermionic operator in the boundary theory shows a variety of unexpected emergent phenomena. In top-down approaches to holographic systems, the fermions are generically coupled to gravity and gauge fields in a variety of ways, beyond minimal coupling. It is certainly desirable to analyze how such non-minimal bulk couplings modify the fermion spectral function in the boundary theory, or may lead to new interesting emergent phenomena.

Recently, it was shown [10–13] that there are consistent truncations of ten and eleven dimensional supergravities to five and four dimensional bulk theories that possess an interesting class of gauge interactions and charged matter, allowing for novel condensed matter physics phenomena, such as superconductivity, to be explored in a consistent top-down approach [14–16]. The fermionic sector of these truncations has also been worked out in [17, 18] where a number of generic couplings for the fermions (with possible applications to strongly-correlated condensed matter systems) have been realized.

Motivated by these studies, we considered in [19] a generic non-minimal fermion coupling in which a spin-1/2 fermion couples to the gauge field through a dipole (Pauli) interaction of the form $F_{ab}\bar{\psi}\Gamma^{ab}\psi$. In fact, we just considered the simplest possible setup, in which a spin-1/2 fermion propagates in the background of a Reissner-Nordström AdS_{d+1} black hole. We found that as one changes the strength of this interaction, spectral weight of the dual fermionic operator is transferred between bands, and beyond a critical value of the dipole coupling, a gap emerges in the fermion density of states. We then concluded that a possible interpretation of this interaction is that it drives the dynamical formation of a (Mott) gap, in the absence of continuous symmetry breaking. In the Hubbard model, the Mott gap forms in $d > 1$ once the on-site interaction U exceeds a critical value in the half-filled system. Upon doping, spectral weight shifts from high to low energies. Consequently, we argued that the strength of the dipole interaction mimics the combined effects of doping and the on-site interaction strength U .

In this paper, we continue our study of the dipole interaction in more detail. We investigate the existence of Fermi surfaces as the dipole coupling p is varied. For

the range of parameters considered in this paper, we find that there is no Fermi surface above a certain value of p . This is the range of parameters considered in [19], where the boundary theory exhibits two main features of Mott insulators, a dynamically generated gap (in the absence of continuous symmetry breaking) and spectral weight transfer. In addition, for the values of the dipole coupling p for which there is a Fermi surface in the boundary theory, we find that at a critical value of this coupling, marginal-Fermi liquid behavior ensues. Consequently, we are able with our model (see Figure 1) to describe at one extreme, the Mott insulating state with a dynamically generated gap, a transition to a marginal Fermi liquid (MFL) and at the other, a Fermi-liquid-like regime in which the excitation spectrum scales linearly with the frequency. As all of these regimes are accessed simply by changing the value of the dipole coupling p , this suggests a direct parallel between p and the hole-doping level in the high-temperature copper-oxide superconductors (hereafter cuprates). In the cuprates, the strong electron correlation physics ends at a value of doping (typically optimal doping) where marginal-Fermi liquid [21] behavior ensues. Perhaps the occurrence of the MFL state of the probe fermions in our holographic setup is an indication that this phenomenological model is ultimately a robust feature of the transition from strong to weakly interacting physics in doped Mott systems.

Having discussed the pole structure of the boundary theory fermion (retarded) correlators for non-zero values of p at small frequency, we analyze the non-analyticities of those correlators for generic values of frequency and momentum. We do this by numerically calculating the quasi-normal modes of the bulk fermion in the (extremal) Reissner-Nordström AdS_{d+1} background. Following the motion of the poles in the complex frequency plane as a function of momentum, we compute their dispersion relations and, for the range of parameters considered, confirm that all of the poles stay in the lower half of the complex frequency plane, for all momenta. Hence, as expected, turning on a non-zero bulk dipole coupling in our set up does not cause an instability in the boundary theory.

We consider our setup at finite temperature and find that as the temperature increases the gap closes and, moreover, the critical temperature for which this happens is much less than the value of the gap. In this regard, the closing of the gap induced by temperature parallels what one obtains in the classic Mott system VO_2 [20], namely, the gap closes at a temperature much lower than the gap.

The paper is organized as follows. In section II we consider a bulk fermion in the Reissner-Nordström AdS_{d+1} black hole background and couple it to the $U(1)$ gauge field non-minimally through a dipole interaction with strength p . We then derive the Dirac equations and rewrite them as flow equations which will be more convenient for numerically calculating the boundary theory fermion correlators. In section III we investigate the ex-

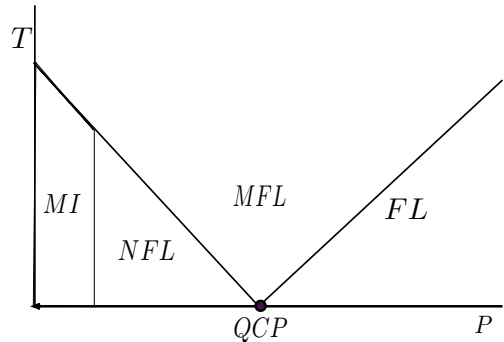


FIG. 1: A cartoon of the phase diagram of the boundary theory considered here. MI indicates a Mott insulator, a phase with a gap in the absence of symmetry breaking. NFL denotes non-Fermi liquid behavior which is distinct from the gapped spectrum of a Mott insulator. MFL indicates marginal-Fermi liquid behavior in which the “electron” self energy scales as $\omega \log \omega$ at $T = 0$, and a FL (Fermi liquid) regime in which the dispersion is linear in frequency. The tuning parameter in this model is the Pauli (dipole) coupling. Similar behavior is obtained in the non-superconducting features of the cuprate materials by tuning the hole-doping level x .

istence of Fermi surfaces as a function of p by solving the Dirac equations at zero frequency. In section IV, we first discuss the small-frequency behavior of the poles of the boundary theory fermion (retarded) correlators when p is non-zero. We then analyze the non-analyticities of those correlators for generic values of frequency and momenta. In section V we study the effects of temperature in our holographic setup. Finally, in section VI, we discuss the relevance of our work to cuprate phase diagram and conclude with open questions as well as extensions for future work. In particular, we contemplate the extension of our results to superconducting backgrounds, which also do not possess a finite ground state degeneracy at zero temperature.

II. BULK ANALYSIS

As we alluded to above, we consider just one form of non-minimal coupling, in which a spin-1/2 fermion is coupled to the gauge field through a dipole interaction¹ of the form $F_{ab}\bar{\psi}\Gamma^{ab}\psi$, and propagates in the background of a Reissner-Nordström AdS_{d+1} black-hole (hereafter, denoted by RN- AdS_{d+1}). Thus, we consider the bulk Lagrangian

$$\sqrt{-g}i\bar{\psi}(\not{D} - m - ip\not{F})\psi, \quad (1)$$

¹ In even bulk dimensions, there is a similar interaction which includes a Γ_5 . We will not consider this interaction in this paper.

in $d + 1 \geq 4$ dimensions where

$$\begin{aligned}\bar{\psi} &= \psi \Gamma^t, \\ \not{D} &= e_c^M \Gamma^c \left(\partial_M + \frac{1}{4} \omega_M^{ab} \Gamma_{ab} - i q A_M \right), \\ \not{F} &= \frac{1}{2} \Gamma^{ab} e_a^M e_b^N F_{MN},\end{aligned}\quad (2)$$

with e_a^M and ω_M^{ab} being the (inverse) vielbein and the spin connection, respectively. We denote the bulk coordinate indices by capital letters $M, N, \dots = \{t, x^i, r\}$ while the tangent space indices are denoted by $a, b, \dots = \{t, \underline{x}^i, \underline{r}\}$. We will reserve the Greek indices μ, ν, \dots to denote boundary coordinate directions. We use Dirac matrices $\Gamma^t, \Gamma^1, \dots, \Gamma^r$ satisfying the Clifford algebra $\{\Gamma^a, \Gamma^b\} = 2\eta^{ab}$. Also, $\Gamma_{ab} = \frac{1}{2}[\Gamma_a, \Gamma_b]$. In what follows, we will rescale $p \rightarrow pL/(d-2)$ for convenience.

The RN-AdS $_{d+1}$ background has a metric and a gauge connection which can be written

$$ds^2 = \frac{r^2}{L^2} [-f(r)dt^2 + d\vec{x}^2] + \frac{L^2}{r^2} \frac{dr^2}{f(r)}, \quad (3)$$

$$A = \mu \left[1 - \left(\frac{r_0}{r} \right)^{d-2} \right] dt, \quad (4)$$

where

$$\begin{aligned}f(r) &= 1 - M \left(\frac{r_0}{r} \right)^d + Q^2 \left(\frac{r_0}{r} \right)^{2(d-1)}, \\ \mu &= \left(\frac{d-1}{2d-4} \right)^{1/2} \frac{Q r_0}{L^2}, \quad M = 1 + Q^2,\end{aligned}\quad (5)$$

with r_0 being the horizon, given by the largest real root of $f(r_0) = 0$. The temperature T of this (black hole) background is given by

$$T = (d-2) \frac{r_0}{4\pi L^2} \left(\frac{d}{d-2} - Q^2 \right). \quad (6)$$

From the above equation, one notes that the RN-AdS $_{d+1}$ black hole is extremal when $Q^2 = d/(d-2)$ while the density and entropy remain finite. Since the background is invariant under $A_t \rightarrow -A_t$, without loss of generality, we can choose μ , or equivalently Q , to be positive. Thus, we can take $0 < Q \leq \sqrt{d/(d-2)}$, where the equality corresponds to extremality.

To analyze the Dirac equations of the bulk fermion, we find it more convenient to go to momentum space by Fourier transforming $\psi(r, x^\mu) \sim e^{ik \cdot x} \psi(r, k^\mu)$, where $k^\mu = (\omega, \vec{k})$. The Fourier transform of the Dirac operator \not{D} is of the form

$$\begin{aligned}\not{D} &= \frac{r}{L} \sqrt{f(r)} \Gamma^r \left[\partial_r + \frac{f'(r)}{4f(r)} + \frac{d}{2r} \right] \\ &\quad - i \frac{L}{r \sqrt{f(r)}} \Gamma^t [\omega + q A_t(r)] + i \frac{L}{r} \vec{k} \cdot \vec{\Gamma},\end{aligned}\quad (7)$$

while

$$\not{F} = (d-2) \frac{\mu}{r_0} \left(\frac{r_0}{r} \right)^{d-1} \Gamma^{rt}. \quad (8)$$

To decouple the Dirac equations, we introduce projectors $\Gamma_\pm = \frac{1}{2}(1 \pm \Gamma^r \Gamma^t \vec{k} \cdot \vec{\Gamma})$ and write $\psi_\pm(r) = r^{d/2} f(r)^{1/4} \Gamma_\pm \psi(r)$. Without loss of generality, we set $k_1 = k$ and $k_{i \neq 1} = 0$, and take the basis

$$\begin{aligned}\Gamma^r &= \begin{pmatrix} -\sigma_3 \otimes \mathbb{1} & 0 \\ 0 & -\sigma_3 \otimes \mathbb{1} \end{pmatrix}, \quad \Gamma^t = \begin{pmatrix} i\sigma_1 \otimes \mathbb{1} & 0 \\ 0 & i\sigma_1 \otimes \mathbb{1} \end{pmatrix}, \\ \Gamma^1 &= \begin{pmatrix} -\sigma_2 \otimes \mathbb{1} & 0 \\ 0 & \sigma_2 \otimes \mathbb{1} \end{pmatrix}.\end{aligned}\quad (9)$$

where σ_j 's are the Pauli matrices, and $\mathbb{1}$ is a $2^{\frac{d-3}{2}}$ -dimensional identity matrix for odd values of d , and $2^{\frac{d-4}{2}}$ -dimensional for d even. Note that by choosing $k_1 = k$ and $k_{i \neq 1} = 0$ the rest of the gamma matrices do not appear in the Dirac equations. So, we did not bother to include those in (9). One then finds

$$\begin{aligned}\frac{r^2}{L^2} \sqrt{f(r)} \partial_r \psi_\pm &= \frac{i\sigma_2}{\sqrt{f(r)}} \left[\omega + \mu q \left(1 - \frac{r_0^{d-2}}{r^{d-2}} \right) \right] \psi_\pm \\ &\quad - \sigma_1 \left(\mu p \frac{r_0^{d-2}}{r^{d-2}} \pm k \right) \psi_\pm - \sigma_3 \frac{r}{L} m \psi_\pm.\end{aligned}\quad (10)$$

We see that the Pauli coupling modifies the appearance of k in the above Dirac equations. To see the effects of p more clearly, consider the solutions of the Dirac equations (10) in the asymptotic and near horizon regimes. Asymptotically, the solutions behave as

$$\begin{aligned}\psi_\pm(r, \omega, k) &= a_\pm(\omega, k) r^{mL} \begin{pmatrix} 0 \\ 1 \end{pmatrix} [1 + \dots] \\ &\quad + b_\pm(\omega, k) r^{-mL} \begin{pmatrix} 1 \\ 0 \end{pmatrix} [1 + \dots].\end{aligned}\quad (11)$$

The effect of p asymptotically is to modify the subleading terms. For $m \in [0, \frac{1}{2})$ both terms in (11) are normalizable and one can choose either a_\pm or b_\pm to be the sources for the dual fermion operator in the boundary theory. In this paper, we take $m \in [0, \frac{1}{2})$ and consider the conventional quantization where a_\pm are the sources. Thus, the dual fermion operator has dimension $\Delta = \frac{3}{2} + m$. Choosing in-falling boundary conditions near the horizon results in a retarded correlator of the form

$$G_R(\omega, k) = \begin{pmatrix} G_+(\omega, k) \mathbb{1} & 0 \\ 0 & G_-(\omega, k) \mathbb{1} \end{pmatrix}, \quad (12)$$

with $G_\pm(\omega, k) = b_\pm(\omega, k)/a_\pm(\omega, k)$. Note that the Dirac equations (10) imply $G_+(\omega, k) = G_-(\omega, -k)$.

When the background is extremal, $f(r)$ has a double zero at the horizon, $f(r) \sim d(d-1)(1 - r_0/r)^2 + \dots$, and this fact makes taking the limit of $\omega \rightarrow 0$ of the equations (10) near the horizon subtle. To take care of the subtlety, one realizes [8, 22] that near the horizon (in which the geometry approaches AdS $_2 \times \mathbb{R}^{d-1}$ for $T = 0$) the equations for ψ_\pm in (10) organize themselves as functions of $\zeta = \omega L_2^2/(r - r_0)$ with $L_2 = L/\sqrt{d(d-1)}$

being the radius of AdS_2 . The coordinate ζ is the suitable radial coordinate for the AdS_2 part of the near horizon region, and in this region, we can write ψ_{\pm} in terms of ζ and expand in powers of ω as follows

$$\psi_{I\pm}(\zeta) = \psi_{I\pm}^{(0)}(\zeta) + \omega \psi_{I\pm}^{(1)}(\zeta) + \omega^2 \psi_{I\pm}^{(2)}(\zeta) + \dots \quad (13)$$

Now, substituting (13) into (10), we find that to leading order

$$\begin{aligned} \psi_{I\pm}^{(0)'}(\zeta) &= \frac{L_2}{\zeta} \left[m\sigma_3 + \left(c_d \frac{p}{L} \pm \frac{kL}{r_0} \right) \sigma_1 \right] \psi_{I\pm}^{(0)}(\zeta) \\ &\quad - i\sigma_2 \left(1 + \frac{qe_d}{\zeta} \right) \psi_{I\pm}^{(0)}(\zeta), \end{aligned} \quad (14)$$

where $e_d = 1/\sqrt{2d(d-1)}$, and $c_d = 1/[(2d-4)e_d]$. Equations (14) are identical to the equations of motion for massive spinor fields [8] with masses (m, \tilde{m}_+) and (m, \tilde{m}_-) in AdS_2 , where \tilde{m}_{\pm} are time-reversal violating mass terms, with the identification

$$\tilde{m}_{\pm} = c_d \frac{p}{L} \pm \frac{kL}{r_0}. \quad (15)$$

Thus, $\psi_{I\pm}^{(0)}(\zeta)$ are dual to spinor operators \mathcal{O}_{\pm} in the IR CFT with conformal dimensions $\delta_{\pm} = \nu_{\pm}^{\pm} + \frac{1}{2}$ where

$$\begin{aligned} \nu_k^{\pm} &= \sqrt{m_{k\pm}^2 L_2^2 - q^2 e_d^2 - i\epsilon}, \\ m_{k\pm}^2 &= m^2 + \left(c_d \frac{p}{L} \pm \frac{kL}{r_0} \right)^2. \end{aligned} \quad (16)$$

We see that turning on p modifies the scaling in the infrared in an important way – effectively, the momentum is pushed up and down by p . We will explore the details of this in what follows.

One can write a formal expression for the fermion retarded correlator (12) at low frequency in terms of the retarded Green functions of the IR CFT spinor operators \mathcal{O}_{\pm} . As shown in [8], such a formal expression is extremely useful in analyzing the small ω behavior of the boundary theory Green functions. This is done by matching the inner AdS_2 and outer AdS_4 solutions in the so-called “matching region” where the $\zeta \rightarrow 0$ and $\omega/\zeta \rightarrow 0$ limits are taken. In so doing, one finds that the coefficients $a_{\pm}(\omega, k)$ and $b_{\pm}(\omega, k)$ in (12) are given by

$$\begin{aligned} a_{\pm}(\omega, k) &= \left[a_{\pm}^{(0)} + \omega a_{\pm}^{(1)} + \mathcal{O}(\omega^2) \right] \\ &\quad + \left[\tilde{a}_{\pm}^{(0)} + \omega \tilde{a}_{\pm}^{(1)} + \mathcal{O}(\omega^2) \right] \mathcal{G}_k^{\pm}(\omega), \end{aligned} \quad (17)$$

$$\begin{aligned} b_{\pm}(\omega, k) &= \left[b_{\pm}^{(0)} + \omega b_{\pm}^{(1)} + \mathcal{O}(\omega^2) \right] \\ &\quad + \left[\tilde{b}_{\pm}^{(0)} + \omega \tilde{b}_{\pm}^{(1)} + \mathcal{O}(\omega^2) \right] \mathcal{G}_k^{\pm}(\omega), \end{aligned} \quad (18)$$

where $a_{\pm}^{(n)}, \tilde{a}_{\pm}^{(n)}, b_{\pm}^{(n)}$ and $\tilde{b}_{\pm}^{(n)}$ are all functions of k and can, in principle, be determined numerically. Also, in the above expressions, $\mathcal{G}_k^{\pm}(\omega)$ denote the retarded Green

functions of the dual IR CFT operators \mathcal{O}_{\pm} which are given by [8]

$$\mathcal{G}_k^{\pm}(\omega) = c_{\pm}(k) \omega^{2\nu_{\pm}}, \quad (19)$$

with

$$\begin{aligned} c_{\pm}(k) &= e^{-i\pi\nu_{\pm}} \frac{\Gamma(-2\nu_{\pm})\Gamma(1+\nu_{\pm}-iqe_d)}{\Gamma(2\nu_{\pm})\Gamma(1-\nu_{\pm}-iqe_d)} \times \\ &\quad \frac{(m+i\tilde{m}_{\pm})L_2 - iqe_d - \nu_{\pm}}{(m+i\tilde{m}_{\pm})L_2 - iqe_d + \nu_{\pm}}. \end{aligned} \quad (20)$$

Note that the expressions (17) are (18) are not valid when $2\nu_{\pm}$ is an integer. In such cases there would be additional terms like $\omega^n \log \omega$ (with n being a positive integer) on the right hand sides of (17) and (18).

In order to obtain $G_{\pm}(\omega, k)$ for generic values of ω and k , one must solve the Dirac equations (10) numerically. For numerical purposes, it is convenient to work with dimensionless quantities. So, we rescale r, ω and k in the Dirac equations (10) by defining

$$r \rightarrow r_0 u, \quad \omega \rightarrow \frac{r_0}{L^2} \omega, \quad k \rightarrow \frac{r_0}{L^2} k. \quad (21)$$

It is also more convenient to convert the Dirac equations (10) into the so-called flow equations [6, 23]. For that, we first write $\psi_{\pm}^T = (\beta_{\pm}, \alpha_{\pm})$ and define $\xi_{\pm} = \beta_{\pm}/\alpha_{\pm}$, in terms of which the Dirac equations (10) then reduce to the non-linear flow equations

$$\begin{aligned} u^2 \sqrt{f(u)} \partial_u \xi_{\pm} &= -2(mL)u \xi_{\pm} \\ &\quad + [v_-(u) \mp k] + [v_+(u) \pm k] \xi_{\pm}^2, \end{aligned} \quad (22)$$

where

$$v_{\pm}(u) = \frac{1}{\sqrt{f(u)}} [\omega + Qq(1 - u^{2-d})] \pm Qp u^{2-d}. \quad (23)$$

To obtain the retarded Green functions of the boundary theory operators, one has to choose the infalling boundary condition at the horizon for the (dual) bulk fields [24, 25]. Expressed in terms of ξ_{\pm} , the infalling boundary condition for $\psi_{\pm}^T = (\beta_{\pm}, \alpha_{\pm})$ at the horizon translates into

$$\xi_{\pm}(u=1) = \begin{cases} i & \omega \neq 0, \\ (mL_2 - \nu_{\pm})/(qe_d + \tilde{m}_{\pm}L_2) & \omega = 0. \end{cases} \quad (24)$$

The matrix of Green functions (12) then takes the form

$$G_R(\omega, k) = \lim_{\epsilon \rightarrow 0} \epsilon^{-2mL} \begin{pmatrix} \xi_+ \mathbf{1} & 0 \\ 0 & \xi_- \mathbf{1} \end{pmatrix} \Big|_{u=\frac{1}{\epsilon}}, \quad (25)$$

where one picks the finite terms as $\epsilon \rightarrow 0$. Up to normalization, the fermion spectral function is defined by

$$A(\omega, k) \equiv \text{Tr Im } G_R(\omega, k). \quad (26)$$

III. CONTINUUM AND BOUND STATES

In this section, we will study the effects of small and negative values of p . As we will see, in this regime, there is a (non-Fermi-liquid) Fermi peak whose properties change as we vary p . In this regime, there is some similarity to the properties of Fermi surfaces studied in [6, 8]. As we change parameters, the scaling dimensions change, and we can pass from a non-Fermi liquid (NFL) to a marginal Fermi liquid (MFL) and on to fermions which have some resemblance to Landau Fermi liquids (FL). However, we will see clearly that there is a positive value of p beyond which the Fermi peak ceases to exist. In what follows, we will mostly set $m = 0$ and $q = 1$ (as we vary p), although similar results hold for a range of these parameters.

We begin by focussing on the regime that has been called log-oscillatory in [6, 8], in which Fermi peaks do not occur. There are some important changes when $p \neq 0$ that we will explain below. When $q^2 > 2m^2L^2$, there exists a range of momenta $k \in \mathcal{I}_+$ for which the dimension of the IR CFT operator \mathcal{O}_+ becomes imaginary. Similarly, for $k \in \mathcal{I}_-$ the dimension of \mathcal{O}_- becomes imaginary. Here, we have defined $\mathcal{I}_\pm = (\mp c_d p - k_o, \mp c_d p + k_o)$ with $k_o = \sqrt{(q^2/2) - m^2L^2}$. (Figure 2 shows plots of \mathcal{I}_\pm versus p .) Consequently, $\text{Im } G_\pm(0, k)$ is generically non-vanishing for $k \in \mathcal{I}_\pm$, respectively; see Figure 3 for plots of $\text{Im } G_\pm(0, k)$ as a function of p . At $p = 0$, one has $\mathcal{I}_+ = \mathcal{I}_- \equiv \mathcal{I}$. This case was analyzed in [6, 8] where it was found that for $k \in \mathcal{I}$ and for small ω , both $\text{Im } G_\pm(\omega, k)$ are periodic in $\log \omega$, with the same period.

At $p = 0$, the range of momenta for which $\text{Im } G_\pm(\omega, k)$ become log-oscillatory at small ω is the same for each, namely $k \in \mathcal{I}$. This degeneracy does not persist for non-zero p , hence the fermion spectral function $A(\omega, k)$ will also have non-oscillatory components. For $p \in [-k_o/c_d, k_o/c_d] - \{0\}$, both $\text{Im } G_\pm(\omega, k)$ show log-oscillatory behavior (with different periods, though) only for $k \in \mathcal{I}_+ \cap \mathcal{I}_-$. For $|p| > k_o/c_d$ where $\mathcal{I}_+ \cap \mathcal{I}_- = \emptyset$, one finds that in the regime where $\text{Im } G_-(\omega, k)$ shows log-oscillatory behavior, $\text{Im } G_+(\omega, k)$ is not oscillatory and vice versa.

For real ν_k^\pm , the boundary conditions for $\xi_\pm(u = 1)$ at $\omega = 0$ are real. Since the equations (22) are real, one deduces that $\text{Im } G_\pm(0, k) = 0$. Thus, $\text{Re } G_\pm(0, k) = G_\pm(0, k) = b_\pm^{(0)}/a_\pm^{(0)}$. $\text{Re } G_\pm(0, k)$ may have poles which would be given generically by the zeros of $a_\pm^{(0)}$. Each zero of $a_\pm^{(0)}$ defines a Fermi momentum k_F , given that $b_\pm^{(0)}$ do not vanish as $k \rightarrow k_F$. Since $G_-(\omega, k) = G_+(\omega, -k)$, vanishing of $a_-^{(0)}$ at some $k = k_F$ implies that $a_+^{(0)}$ vanishes at $k = -k_F$. So, in order to find k_F , we can just analyze the zeros of $a_-^{(0)}$. From the asymptotic behavior of ψ_- which is given in (11), together with (17) and the definition of $\psi_-^T = (\beta_-, \alpha_-)$, it is easy to see that, at $\omega = 0$, $\alpha_-(u, k) = a_-^{(0)} u^{mL} + \dots$ as $u \rightarrow \infty$. (Also, note that at

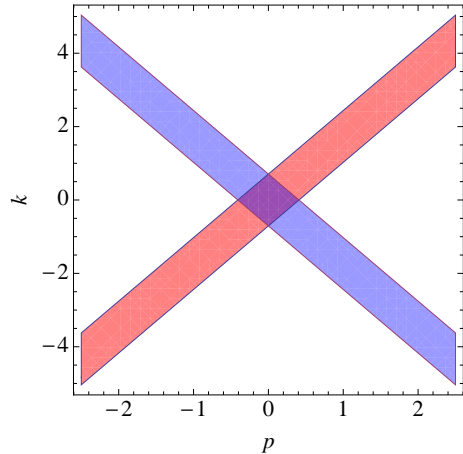


FIG. 2: Plots of \mathcal{I}_\pm versus p (for $d = 3$, $\Delta_\psi = 3/2$ and $q = 1$). The red band depicting \mathcal{I}_- is where $\text{Im } G_-(\omega, k)$ becomes oscillatory at small ω . The blue band (\mathcal{I}_+) shows the region where $\text{Im } G_+(\omega, k)$ is oscillatory (at small ω).

$\omega = 0$, $\beta_-(u, k) = b_-^{(0)} u^{-mL} + \dots$ as $u \rightarrow \infty$.) So the k_F 's define a set of momenta for which, at $\omega = 0$, $\psi_-(u, k)$ becomes normalizable (a 'bound state') as $u \rightarrow \infty$. To find the k_F 's, we analyze the equation for $\alpha_-(u, k)$ as follows.

Plugging $\psi_-^T = (\beta_-, \alpha_-)$ into the Dirac equations (10), one obtains a set of two coupled linear differential equations for α_- and β_- . Setting $\omega = 0$ and decoupling these two equations, we obtain

$$-\frac{u^2 \sqrt{f(u)}}{v_-^0(u) + k} \partial_u \left(\frac{u^2 \sqrt{f(u)}}{v_+^0(u) - k} \partial_u \right) \alpha_- = \alpha_-, \quad (27)$$

$$-\frac{u^2 \sqrt{f(u)}}{v_+^0(u) - k} \partial_u \left(\frac{u^2 \sqrt{f(u)}}{v_-^0(u) + k} \partial_u \right) \beta_- = \beta_-, \quad (28)$$

where we have set $m = 0$ for convenience. In (27), the superscript "0" on $v_\pm(u)$ indicates that we have set $\omega = 0$ in (23). Once again, we set $d = 3$ and $q = 1$ in what follows. As we vary p , we look (numerically) for the momenta k_F for which $\alpha_-(u \rightarrow \infty, k_F) = 0$, given an appropriate boundary condition for $\alpha_-(u, \omega = 0, k)$ at the horizon. Indeed, solving (27) near the horizon, one easily obtains that $\alpha_-(u, k) \sim f(u)^{\pm \nu_k^-/2}$ as $u \rightarrow 1$. Because by assumption we are in a regime where ν_k^- is real and positive² (and, in fact, generically irrational), $f(u)^{-\nu_k^-/2}$ blows up as $u \rightarrow 1$. Thus, $\alpha_-(u \rightarrow 1, k) \sim f(u)^{\nu_k^-/2}$ is the regular horizon boundary condition that should be chosen.

In Figure 4 we have plotted such values of k_F as a function of p . Starting with negative values of p (while

² For $\nu_k^- = 0$, one finds that $\alpha_-(u \rightarrow 1, k) = a(1 + \dots) + b \log(u - 1)(1 + \dots)$ where the dots represent terms which vanish as $u \rightarrow 1$, and a and b are some constants. In order for $\alpha_-(u, k)$ not to blow up at the horizon, one should then choose $b = 0$.

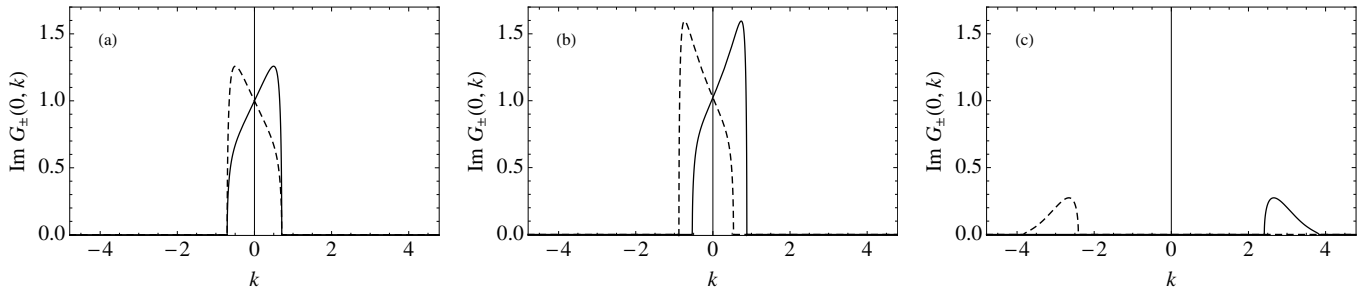


FIG. 3: Plots of $\text{Im } G_{-}(0, k)$ (solid line) and $\text{Im } G_{+}(0, k)$ (dashed line) for (a) $p = 0$, (b) $p = 0.1$ and (c) $p = 1.8$. We set $d = 3$, $\Delta_{\psi} = 3/2$ and $q = 1$. Similar plots can be obtained for negative values of p by switching the solid lines with the dashed lines. Focusing on positive p , we see that the maximum value of $\text{Im } G_{\pm}(0, k)$ increases as $p \rightarrow 1/\sqrt{6}$, after which (namely, for $p > 1/\sqrt{6}$) it rapidly decreases.

keeping $q = 1$ fixed), k_F increases as we raise p causing it to move towards the boundary of the oscillatory region \mathcal{I}_{-} . As p approaches $1/\sqrt{6}$ from below, k_F approaches $\sqrt{2}$ (in units of r_0/L^2 , from below). The blue dots in Figure 4 show the location of k_F 's versus p and the red band depicts the oscillatory region. There is a single Fermi surface for each p as we increase p up to $p = 1/\sqrt{6}$. We have explicitly checked that $\beta_{-}(u = \infty, k = k_F)$ does not vanish, so $k = k_F$ are genuine poles of $\text{Re } G_{-}(0, k)$. At $p = 1/\sqrt{6}$, $k_F = \sqrt{2}$, and as a result $\nu_{k_F}^{-}$ vanishes (recall that $d = 3$, $m = 0$ and $q = 1$). (At this point, and in fact at any point in which $2\nu_{k_F}^{\pm} \in \mathbb{Z}$, the analysis should be more carefully done, as logarithms must be included.) For $p > 1/\sqrt{6}$, we do not see a Fermi surface as $\alpha_{-}(u = \infty, k)$ does not vanish outside the oscillatory region. We have checked this numerically up to $p = 10$, and, given the observed behavior of $\alpha_{-}(u = \infty, k)$, we do not expect it to change as we increase p further. Indeed, Figure 5 shows plots of $\alpha_{-}(u = \infty, k)$ versus k for sample values of p . We have also plotted $\beta_{-}(u = \infty, k)$, shown by the red curves in Figure 5.

Following [8], the excitations around these Fermi surfaces can be analyzed. Using (17) and (18), near $k = k_F$ and at small ω , $G_{-}(\omega, k)$ takes the form

$$G_{-}(\omega, k) \approx \frac{b_{-}^{(0)}(k_F)}{\partial_k a_{-}^{(0)}(k_F)k_{\perp} + \omega a_{-}^{(1)}(k_F) + \tilde{a}_{-}^{(0)}(k_F)\mathcal{G}_{k_F}^{-}(\omega)}, \quad (29)$$

where $k_{\perp} = k - k_F$. Suppose the denominator in (29) vanishes at some $\omega_{*}(k) = \text{Re } \omega_{*}(k) - i \text{Im } \omega_{*}(k)$. For $p = 0$, the dispersion relation, the width and the residue of the pole were worked out in detail in [8]. Parts of the data in these quantities (such as the scaling of the dispersion relation) come from the IR CFT (or, equivalently, the AdS_2 part of the near horizon geometry), and other parts (such as Fermi velocity) from the UV physics. For non-zero p , the data which come from the IR CFT will be slightly modified according to (16) whereas the data coming from the UV physics could be substantially modified.

For $-0.53 < p < 1/\sqrt{6}$, we find that $1/2 > \nu_{k_F}^{-} > 0$. As a result, the small ω excitations around $k = k_F$ will have a (non-Fermi liquid) dispersion relation $\text{Re } \omega_{*}(k) \propto k_{\perp}^z$

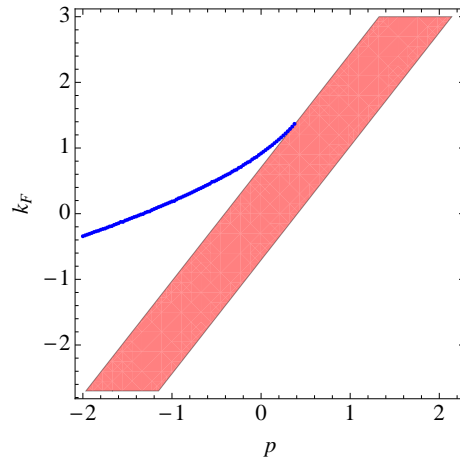


FIG. 4: k_F 's (shown by blue dots) versus p . For $p \leq 1/\sqrt{6}$ there is a single Fermi surface for each p . For $p > 1/\sqrt{6}$ we do not find Fermi surfaces. The orange band shows the oscillatory region \mathcal{I}_{-} .

and a width $\text{Im } \omega_{*}(k) \propto k_{\perp}^z$ where $z = 1/(2\nu_{k_F}^{-})$. Thus, for this range of p , $\text{Im } \omega_{*}(k)/\text{Re } \omega_{*}(k)$ does not vanish as $\omega \rightarrow 0$, implying that these excitations are not stable. Note that the residue at the pole is given by $Z \propto k_{\perp}^{z-1}$. At $p = -0.53$, $\nu_{k_F}^{-} = 1/2$ and the excitations near the Fermi surface are of the marginal Fermi liquid type. For $-1.54 < p < -0.53$, $1 > \nu_{k_F}^{-} > 1/2$, hence, the small ω excitations around $k = k_F$ will have a linear dispersion relation $\text{Re } \omega_{*}(k) \propto k_{\perp}$ and a width $\text{Im } \omega_{*}(k) \propto k_{\perp}^{2\nu_{k_F}^{-}}$. So, for this range of p , these excitations are stable as $\text{Im } \omega_{*}(k)/\text{Re } \omega_{*}(k) \rightarrow 0$ as $\omega \rightarrow 0$. It is in this sense that we refer to this region as the Fermi liquid. Also, we found that k_F goes through zero at $p = -1.317$, signifying that the excitations change over from ‘particle-like’ to ‘hole-like’. Consequently, we find that simply by varying p , we can tune from a Fermi liquid, $p < -0.53$ to a marginal Fermi liquid at $p = -0.53$ to a generic non-Fermi liquid for $-0.53 < p < 1/\sqrt{6}$ and finally to a Mott insulator for p sufficiently large. Precisely how the system behaves for large values of p will now be addressed in the quasi-normal mode analysis.

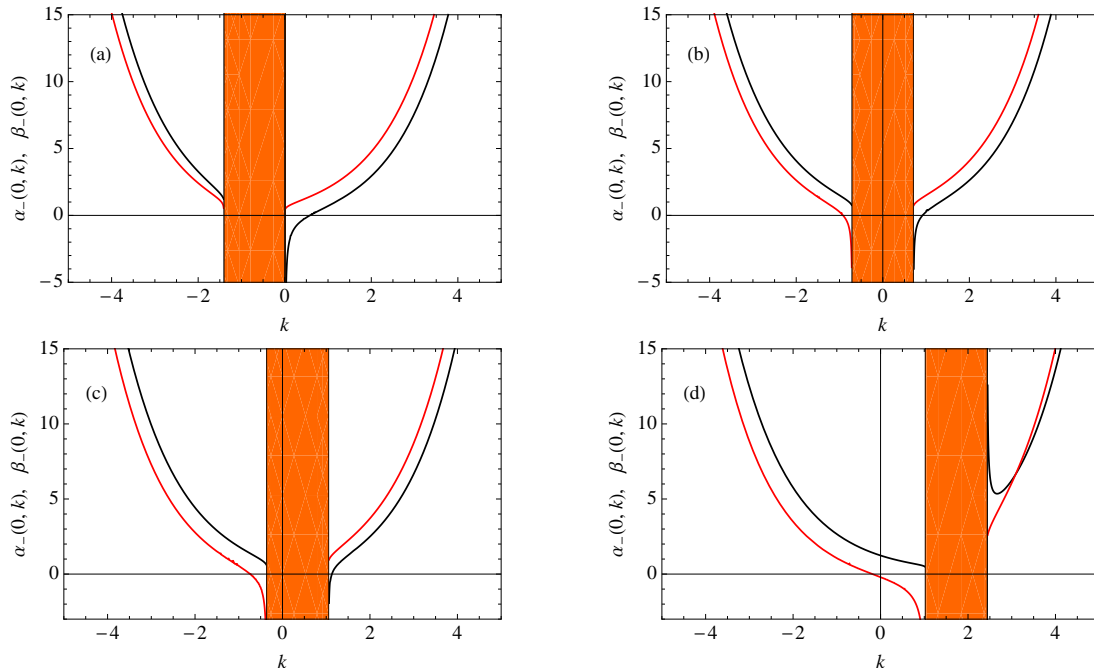


FIG. 5: Plots of $\alpha_-(u = \infty, k)$ (black curves) and $\beta_-(u = \infty, k)$ (red curves) versus k for (a) $p = -0.4$, (b) $p = 0$, (c) $p = 0.2$, (d) $p = 1$. The orange strip in each plot shows the oscillatory region \mathcal{I}_- . The plots are generated for $d = 3$, $m = 0$ and $q = 1$. By $k \rightarrow -k$, similar plots could be obtained for $\alpha_+(u = \infty, k)$ and $\beta_+(u = \infty, k)$.

IV. QUASI-NORMAL MODES AND STABILITY

In this section we analyze the poles of $G_{\pm}(\omega, k)$ and, in particular, discuss how they move in the complex ω -plane as we vary k . Since $G_+(\omega, -k) = G_-(\omega, k)$, without loss of generality, we can just focus on the poles of $G_-(\omega, k)$. We denote the poles of $G_-(\omega, k)$ by $\omega_*(k)$. At small ω , the poles of $G_-(\omega, k)$ can be worked out semi-analytically. Indeed, for $p = 0$, the small ω poles were worked out in detail in [8] for $k \in \mathcal{I}$ as well as k close to k_F where it was argued that such poles are all located in the lower half of the complex ω -plane. Since the arguments of [8] are mainly based on the data coming from the IR region (the near horizon AdS_2 region of the background), they can easily be extended to non-zero values of p , where results similar to those in the case of $p = 0$ are obtained. For example, at small ω and for $k \in \mathcal{I}_-$, where $\text{Im } G_-(\omega, k)$ is oscillatory, the poles (for a fixed k) are exponentially separated on a straight line which is in the lower half ω -plane. (The line is diagonally oriented, *i.e.* it ends on the $\omega = 0$ branch point.) Equivalently, for those values of p for which there exists a Fermi surface, the small ω poles near $k = k_F$ are all located in the lower half ω -plane. In particular, if $\nu_{k_F}^- < 1/2$, then $\omega_*(k \rightarrow k_F)$ as a function of k follows a straight line in the lower half of the complex ω -plane. The semi-analytic arguments of [8], and their generalizations to non-zero p , are applicable only for small ω , and for $k \in \mathcal{I}_{\pm}$, or when k is near k_F . Nevertheless, on general grounds, one expects the poles to be located in the lower half ω -plane

beyond the small ω regime (and, of course, for all values of k). To find the poles beyond the small ω regime, one is usually forced to do numerics which, in the context of the AdS/CFT correspondence, involves performing some quasi-normal mode analyses in the bulk.

Besides isolated poles, $G_{\pm}(\omega, k)$ at zero temperature will have a branch cut (at least for small ω) which could be understood without doing the numerics. Note that since $\mathcal{G}_k^{\pm}(\omega) \sim \omega^{2\nu_k^{\pm}}$ appears in the expressions for $a_{\pm}(\omega, k)$ and $b_{\pm}(\omega, k)$ in (17) and (18), $G_{\pm}(\omega, k)$ will have a branch point at $\omega = 0$ for generic values of k (where $2\nu_k^{\pm}$ are irrational), and a branch cut, which we take to be extended in the negative imaginary axis. For those values of k for which $2\nu_k^{\pm} \in \mathbb{Z}$, there is still a branch cut which is due to the appearance of logarithmic terms of the form $\omega^n \log \omega$ ($n \in \mathbb{Z}$) in the expressions for $a_{\pm}(\omega, k)$ and $b_{\pm}(\omega, k)$. The branch cut seems to be a distinctive feature of the two-point retarded correlators of operators in the zero temperature d -dimensional boundary theory dual to the extremal RN- AdS_{d+1} background. Indeed, the branch cut was observed explicitly in the correlators of scalar and spinor operators in [8, 27] as well as the conserved currents in the shear and sound channels in [28, 29]. As we will see below, this branch cut appears in our quasi-normal mode analysis. At finite temperature, however, the branch cut dissolves into a series of isolated poles on the negative imaginary axis.

Generically, $G_R(\omega, k)$ will have poles whenever $a_{\pm}(\omega, k) = 0$. In the context of the AdS/CFT correspondence, this problem could be addressed by comput-

ing the quasi-normal modes of ψ_{\pm} in the RN-AdS $_{d+1}$ background, which are solutions to the Dirac equations (10) subject to the boundary conditions that they are infalling at the horizon and normalizable asymptotically. Except in very special cases, the generic values of the quasi-normal frequencies are usually computed numerically. We use the so-called Leaver's method [26] for this purpose. For concreteness, we take the boundary theory to be (2+1)-dimensional, *i.e.* $d = 3$. Our analysis can straightforwardly be extended to larger values of d . Also, as in the previous discussions, we consider $m = 0$ and $q = 1$.

Substituting $\psi_{\pm}^T = (\beta_{\pm}, \alpha_{\pm})$ in the Dirac equations (10), and setting $m = 0$, one finds

$$u^2 \sqrt{f(u)} \partial_u \beta_{\pm} = [v_{-}(u) \mp k] \alpha_{\pm}, \quad (30)$$

$$u^2 \sqrt{f(u)} \partial_u \alpha_{\pm} = -[v_{+}(u) \pm k] \beta_{\pm}. \quad (31)$$

The equations for α_{\pm} are the relevant equations for obtaining the quasi-normal frequencies of ψ_{\pm} . Squaring the above equations, the decoupled equations for α_{\pm} are easily obtained

$$\frac{u^2 \sqrt{f(u)}}{v_{-}(u) \mp k} \partial_u \left(\frac{u^2 \sqrt{f(u)}}{v_{+}(u) \pm k} \partial_u \right) \alpha_{\pm} = -\alpha_{\pm}. \quad (32)$$

As we alluded to above, without loss of generality, we can focus on the quasi-normal frequencies of ψ_{-} and just analyze the equation for α_{-} in (32).

In what follows, we switch to a new radial coordinate $z = 1/u$ which is more convenient for doing the numerics in this section. In terms of the new radial coordinate, the horizon is at $z = 1$ and the asymptotic boundary at $z = 0$. The equation for α_{-} in (32) then becomes

$$\frac{\sqrt{f(z)}}{v_{-}(z) + k} \partial_z \left(\frac{\sqrt{f(z)}}{v_{+}(z) - k} \partial_z \right) \alpha_{-} = -\alpha_{-}. \quad (33)$$

To compute the quasi-normal modes of ψ_{-} , the behavior of α_{-} should be infalling at the horizon and normalizable at the boundary. As mentioned above, we use Leaver's method [26] to compute the quasi-normal frequencies. For that, we first pull out the leading behavior of α_{-} at the horizon as well as the boundary and write

$$\alpha_{-}(z) = e^{i \frac{\omega}{6(1-z)}} f(z)^{-i \left(\frac{\omega}{9} + \frac{q}{4\sqrt{3}} \right)} z \tilde{\alpha}_{-}(z). \quad (34)$$

Note that $\tilde{\alpha}_{-}(z = 1)$ is a constant which could be set equal to unity as the equation for α_{-} is homogeneous. Next, we write $\tilde{\alpha}_{-}(z)$ as a power series in z around a point $z_0 = 1/2$ (so that the radius of convergence of the series covers both the horizon and the boundary)

$$\tilde{\alpha}_{-}(z) = \sum_{m=0}^M \tilde{\alpha}_m^{-}(\omega, k) \left(z - \frac{1}{2} \right)^m. \quad (35)$$

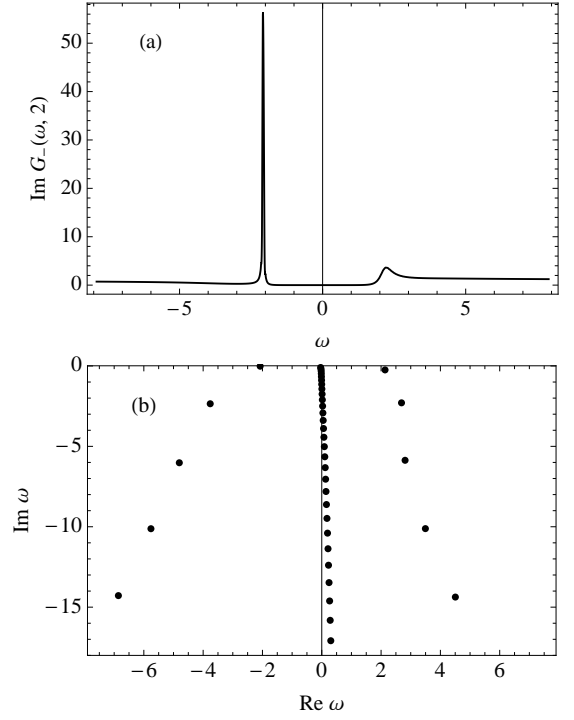


FIG. 6: (a) $\text{Im } G_{-}(\omega, k)$ as a function of ω for $k = 2$. (b) The quasi-normal frequencies of α_{-} for $k = 2$, $d = 3$, $p = 5$, $q = 1$ and $m = 0$ in both plots. Also, $M = 250$.

Substituting (35) and (34) into (33), one obtains

$$\sum_{m=0}^M A_{mp}^{-}(\omega, k) \tilde{\alpha}_m^{-}(\omega, k) = 0, \quad (36)$$

where $A_{mp}^{-}(\omega, k)$ are the elements of a $(M+1)$ by $(M+1)$ matrix $A^{-}(\omega, k)$. The quasi-normal frequencies (for a fixed k) are then the solutions to

$$\det A^{-}(\omega, k) = 0. \quad (37)$$

The bottom plot in Figure 6 shows the quasi-normal frequencies of α_{-} for $p = 5$ and $k = 2$. To generate this plot we set $M = 250$. Due to space limitations, the plot only shows a handful of the quasi-normal frequencies. As M is increased, the poles located along the negative imaginary axis become closer to one another, suggesting that their existence is due to taking M to be finite, and in the limit of $M \rightarrow \infty$ they should indeed form the branch cut we mentioned earlier. On the other hand, increasing M does not seem to change the qualitative behavior of the poles which are oriented almost diagonally on each side of the negative imaginary axis. Notice that the poles are all located on the lower half ω -plane. As may be seen from the plot, the branch cut bends to the right for large negative values of $\text{Im } \omega$. This behavior is different from the cases studied in [28, 29] where the unbroken parity symmetry of the boundary theory forces the branch cut of

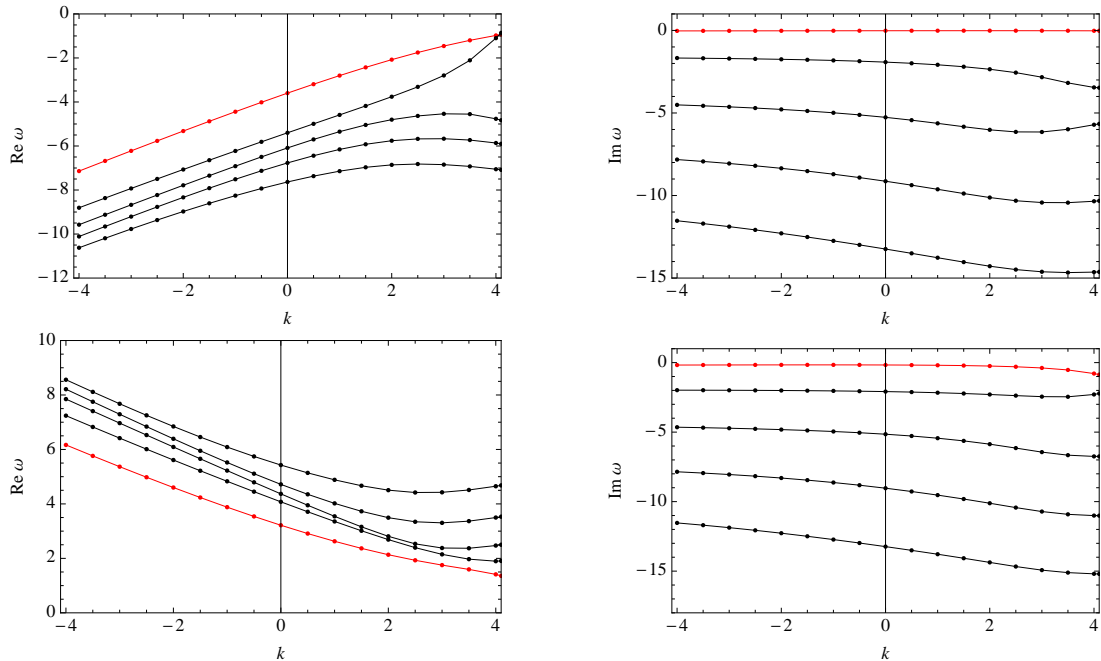


FIG. 7: The top two plots show the dependence on k (dispersion relation) of the real and imaginary parts of the first five quasi-normal modes (depicted in Figure 6(b)) on the left hand side of the negative imaginary axis. The bottom two plots show the dispersion relation of the real and imaginary parts of the first five quasi-normal modes of Figure 6(b) which are on the right hand side of the negative imaginary axis. The plots are generated for $d = 3$, $p = 5$, $m = 0$ and $q = 1$ and $M = 250$. The red data corresponds to the mode closest to the real axis in the complex ω -plane.

the retarded correlators to stay on the negative imaginary axis. Indeed, the bending of the branch cut is similar to what was observed in [27] for the retarded correlators of charged scalar operators in the presence of a magnetic field. The top plot in Figure 6 shows $\text{Im } G_-(\omega, k = 2)$ as a function of ω . The location of the peak on the left hand side and the bump on the right hand side match quite well with the two quasi-normal frequencies which are closest to the real axis. It is apparent that all of the other quasi-normal modes are relatively wide and individually have small residue.

The dispersion relation, $\omega_*(k)$, of the quasi-normal frequencies shown in Figure 6(b) can be computed numerically by following their motion in the complex ω -plane as k is varied. As shown in Figure 7, the two quasi-normal frequencies which are closest to the real axis have markedly different dispersion relations than the rest (higher resonances). The effects of these higher resonances become important as one goes beyond the small frequency and momentum approximation. For large $|k|$, the imaginary part of the quasi-normal frequencies is approximately constant (and negligible compared to the real part). In contrast, for large $|k|$, the real part is proportional to k , with the constant of proportionality being equal to ± 1 (within our numerical precision). This behavior is expected because at large frequency and momentum, the vacuum of the boundary theory is effectively Lorentz-invariant. So, the dispersion relation of the excitations should effectively be relativistic at large frequency

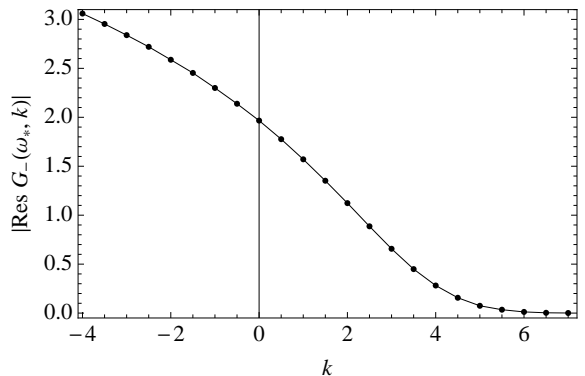


FIG. 8: $|\text{Res } G_-(\omega_*, k)|$ as a function of k for the leading negative-frequency pole in Figure 6(b) which is closest to the real axis and located to the left of the negative imaginary axis. We set $d = 3$, $p = 5$, $q = 1$ and $m = 0$.

and momentum where the effect of charge density is negligible.

To have a better understanding of the spectrum, it is important to know how the residues (at the poles) behave as a function of k . In order to numerically compute the residues of $G_-(\omega, k)$ at $\omega = \omega_*(k)$, we first developed series expansions for $\alpha_-(z; \omega, k)$ and $\beta_-(z; \omega, k)$ both near the horizon around $z = 1 - \epsilon$ and near the boundary around $z = \delta$. We then numerically integrated the (decoupled) differential equation for $\alpha_-(z; \omega, k)$ and $\beta_-(z; \omega, k)$ from $z = 1 - \epsilon$ to $z = \delta$ and matched the nu-

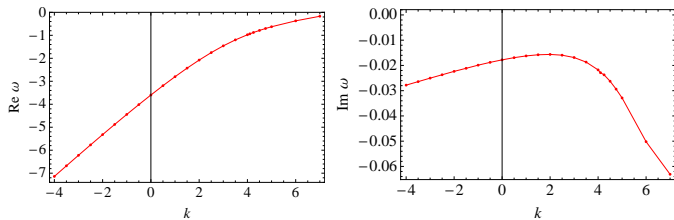


FIG. 9: A close-up of the real (left plot) and imaginary (right plot) parts of the dispersion relation of the leading pole shown in Figure 6(b) which is closest to the real axis and located to the left of the negative imaginary axis.

merically integrated solutions and their (first) derivatives with their boundary series expansions at $z = \delta$. In so doing, we were able to compute the residues of $G_-(\omega, k)$ at $\omega = \omega_*(k)$ for a fixed k , denoted by $\text{Res } G_-(\omega_*, k)$. Repeating the same steps for different k 's, one can numerically obtain the dependence of the residues on k . We computed the k -dependence of the residues of $G_-(\omega, k)$ for the leading negative-frequency pole (closest to the real axis). Shown in Figure 8 is the absolute value of $\text{Res } G_-(\omega_*, k)$ as a function of k for this pole. Also, the plots in Figure 9 show a close up of the real and imaginary parts of the dispersion relation of this pole. Comparing the plot of the residue to Figure 6(a), it is now apparent why the gap forms: as the leading negative-frequency quasi-normal mode approaches $\omega = 0$, its residue dies off quickly. Since no other mode has appreciable spectral weight, this accounts for the suppression of the spectral weight near $\omega = 0$. For larger values of p , the residue falls off more rapidly, and consequently the gap widens.

V. FINITE TEMPERATURE

So far, our analysis has been at zero temperature. However, there are important aspects of Mott insulators that transpire at finite temperature. In particular, there are Mott insulators [20] exhibiting a transition to a conducting state as the temperature is increased. The classic example of this is VO_2 . Below $T_* = 340\text{K}$, VO_2 becomes insulating with a gap of $\Delta = 0.6\text{ eV}$. This ratio of the gap to the critical temperature Δ/T_* is approximately 20. This behavior should be contrasted with systems such as superconductors in which $U(1)$ symmetry is broken and $\Delta/T_c \approx 1 - 2$. That Δ/T_* well exceeds unity is one of the unresolved puzzles with VO_2 . It points to strong correlations being the source of the gap rather than the breaking of some spontaneous symmetry as in the case of superconductivity. Optical conductivity studies [30] reveal that spectral weight as far away as 6 eV contributes to the formation of the Drude peak at zero frequency once the Mott gap closes. Such UV-IR mixing is a ubiquitous feature of Mott systems. While we have argued that our holographic setup can capture the high-low energy spectral weight transfer, we have not yet addressed

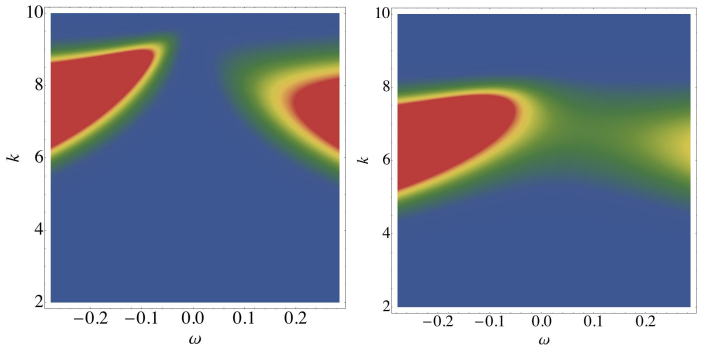


FIG. 10: A close-up of the density plots of $\text{Im } G_-(\omega, k)$ for $p = 6$ and $T/\mu \simeq 5.15 \times 10^{-3}$ (left) and $T/\mu \simeq 3.98 \times 10^{-2}$ (right). A gap is still seen in the plot on the left while it is closed in the plot on the right.

the finite temperature aspects of the Mott problem.

The boundary theory we are investigating here can easily be studied at finite temperature by considering the RN-AdS₄ background away from extremality, namely for $0 < Q < \sqrt{3}$. Using the same procedures outlined above, we obtained the spectral function and studied the density of states as a function of temperature. As Figure 10 reveals, the Mott gap observed here does in fact close as the temperature increases. Further, the transition is sharp. To estimate the ratio of the zero-temperature gap to the temperature at which the gap closes, T_* , we take a close-up of the density of states and study its evolution as a function of temperature, see Figure 11. Indeed for $p = 6$ (or $p = 7$), we find that $\Delta/T_* \sim 10$. Though smaller than Δ/T_* in vanadium oxide, it does illustrate that dynamically generated gap we have found here does possess non-trivial temperature dynamics.

For the record, we show in Figure 12 the quasi-normal frequencies of α_- (bottom plot) as well as $\text{Im } G_-(\omega, k)$ (top plot) for $k = 2$, $p = 5$ and $T/\mu = 0.16$. Notice that the branch cut at zero temperature has dissolved at finite temperature into a series of isolated poles on the negative imaginary axis³. Also, as it can easily be seen from the plots in Figure 12, the (real part of the) two quasi-normal frequencies of α_- which are closest to the real axis match quite well with the location of the peak on the left hand side and the bump on the right hand side in $\text{Im } G_-(\omega, k = 2)$. The other quasi-normal frequencies represent poles in $\text{Im } G_-(\omega, k = 2)$ which are relatively wide and have small residue. Although not shown, to the extent that we have checked, all the poles stay on the lower half ω -plane as k is varied.

³ Note that in order to generate the bottom plot in Figure 12, the expression in (34) should be appropriately modified to reflect the fact that the system is at finite temperature. This is tied to the fact that at finite temperature $f(r)$ has a single zero at the horizon.

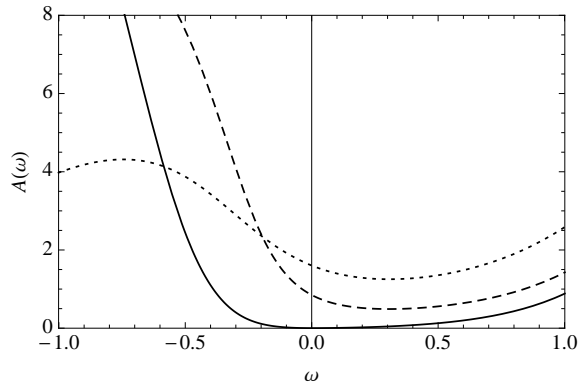


FIG. 11: A close-up of the density of states $A(\omega)$ at $p = 6$ for $T/\mu \simeq 0.44$ (dotted), 0.16 (dashed) and 5.15×10^{-3} (solid).

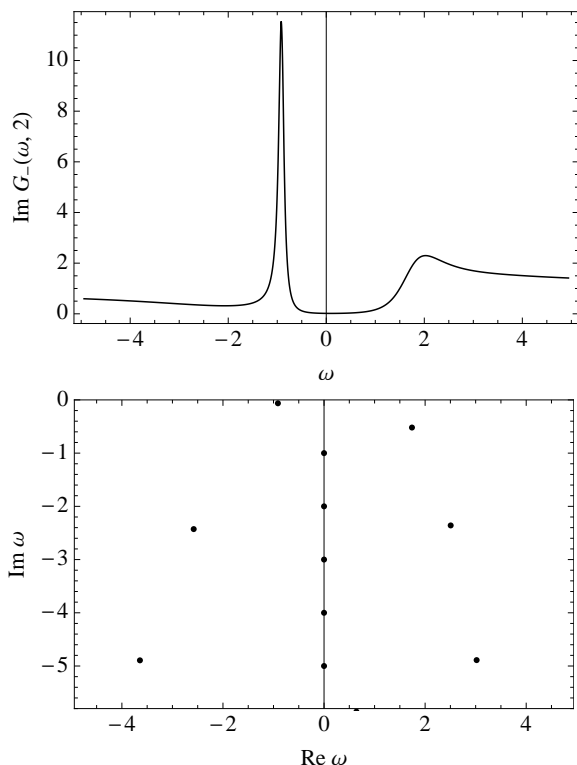


FIG. 12: The plots in (a) and (b) show, for $k = 2$ and $T/\mu = 0.16$, $\text{Im } G_-(\omega, k)$ as a function of ω and the quasi-normal frequencies of α_- , respectively. Here, $d = 3$, $p = 5$, $q = 1$, $m = 0$, and $M = 250$.

VI. DISCUSSION

We have studied extensively the dichotomous behavior of the boundary theory fermion correlators in the presence of a bulk Pauli coupling in our holographic set up. For the boundary theory dual to the extremal RN-AdS₄ background, we showed that as we vary p from large negative values up to a small positive value of $p = 1/\sqrt{6}$ (while keeping $m = 0$ and $q = 1$ fixed), the behavior of the excitations change from Fermi liquid like (for $p < -0.53$), though not in the precise Landau sense in which the

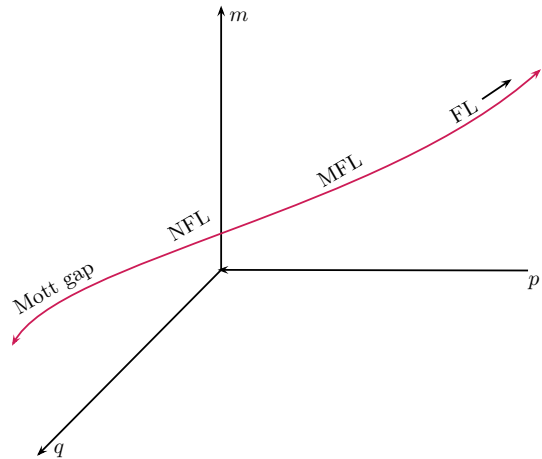


FIG. 13: A cartoon of the zero temperature “phase diagram” in the m, q, p parameter space. Different regions of the phase diagram correspond to each of the principal structures in the cuprate phase diagram (compare to Figure 1).

width of the excitations is quadratic in frequency, to a marginal Fermi liquid at $p = -0.53$ and on to a non-Fermi liquid for $-0.53 < p \leq 1/\sqrt{6}$. In the context of the earlier work [8] in which such behavior was observed by changing the scaling dimension, as well as the charge, of the boundary theory fermion operator, the Pauli coupling offers a more direct connection with Mott physics. Our argument here is based on the fact that for large positive values of p a Mott gap arises (as evidenced by a vanishing of the quasiparticle residue shown in Figure 8) in the spectrum of the boundary theory fermion operator without the apparent breaking of a continuous symmetry. This is Mott physics. We have seen these features by holding m and q fixed while varying p . It is clear that the basic properties that we have seen will persist throughout a domain in the m, q, p parameter space. In Figure 13, we suggest that some locus through the parameter space can be identified with doping in the cuprate phase diagram: each of the principle features in the normal state of the cuprates is present. It is interesting to compare this heuristic phase diagram with that of the cuprates in which there is a continuous evolution from a Mott insulator in the undoped state to a Fermi liquid in the overdoped regime. In between these extremes lie non-superconducting non-Fermi liquid states characterized by a pseudogap (a suppression of the density of states without any long-range superconductivity) and a strange metal in which the resistivity is a linear function of temperature. Our work suggests that the Pauli coupling mimics the role of the electron filling.

It is natural to investigate how the introduction of a superconducting condensate would complement the physics that we have discussed here. A suitable charged background at zero-temperature was studied in [31], following [32, 33]. For a range of parameters (namely, for $m_\phi^2 - 2q_\phi^2 < -3/2$ where m_ϕ and q_ϕ are the mass and the charge of the bulk scalar field, respectively), a bulk

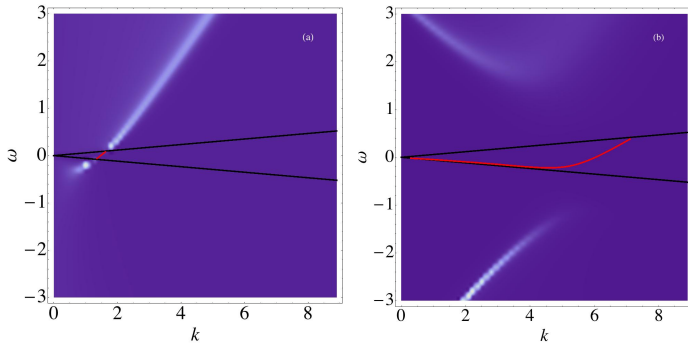


FIG. 14: Density plot of the boundary theory fermion spectral function for (a) $p = 0$ and (b) $p = 3$. Here, $q_\phi = 1.5$, $L = 1$, and $\mu = 2\sqrt{3}$. The black lines depict the IR light-like region and the red curves represent the bound states.

solution with a non-zero charged scalar condensate is preferred over the Reissner-Nordström solution. (Indeed, for this range of parameters, the Reissner-Nordström solution is unstable against turning on the scalar field in the bulk.) This solution is asymptotically AdS_4 , and $m_\phi = 0$ has a near horizon geometry that is also AdS_4 with a finite speed of light, c_{IR} , and a finite dynamical exponent. The horizon of the zero-temperature solution (which is the Poincaré horizon of AdS_4) is at $r = 0$ and there is no residual entropy at zero temperature. Although we will discuss this elsewhere [34], preliminary studies of the effect of the Pauli coupling on the boundary theory fermion correlators⁴ indicate that the main contributor to the gap in the fermion spectral density in the superconducting phase is the so-called Majorana scalar-fermion coupling (denoted by η_5 in [36]). As there are limited tools available for the study of “electrons” at strong coupling, the model proposed here could offer key insight into how superconductivity emerges from a background in which all energy scales are coupled. As a first step, we set $\eta_5 = 0$, and briefly discuss here the effect of the Pauli coupling on fermion correlators in a boundary theory dual to the superconducting background of [31]. Suppose $m_\phi = 0$, and $q_\phi > \sqrt{3}/2$, so that the near horizon geometry is AdS_4 (with a characteristic radius L_{IR}) and assume there exists a $(2+1)$ -dimensional IR CFT dual to this AdS_4 near-horizon geometry. For definiteness, we set $q_\phi = 2q$, where q is the charge of the bulk fermion⁵. The Dirac equation for ψ_\pm (as well as the corresponding flow equations for ξ_\pm) and the IR boundary conditions can easily be worked

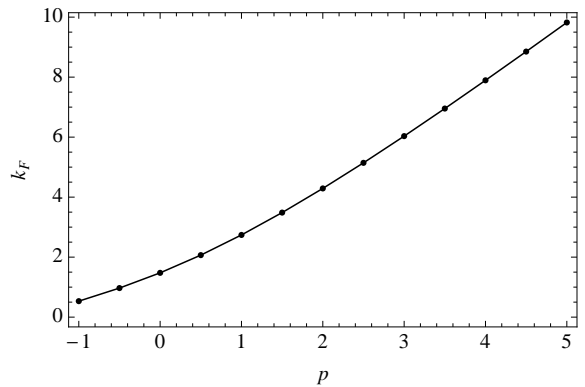


FIG. 15: k_F as a function of p . Here, $q_\phi = 1.5$, $L = 1$, and $\mu = 2\sqrt{3}$.

out. A crucial difference compared to the case of the RN- AdS_4 is that here the dimension of the IR CFT operators dual to $\psi_\pm(r \rightarrow 0)$ does not depend on p . Depending on the sign of $s^2 \equiv -\omega^2/c_{\text{IR}}^2 + k^2$, the Green functions $G_\pm(\omega, k)$ exhibit different behaviors. (Note that one still has $G_-(\omega, k) = G_+(\omega, -k)$.) For $s^2 > 0$ (IR space-like region), the horizon boundary conditions for ξ_\pm are real and since the flow equations are also real, one concludes that the boundary theory fermion spectral density is zero in this region, except when there are bound states (by which we mean poles of $\text{Re } G_\pm(\omega, k)$ in the $s^2 > 0$ region) of the Dirac equation. For $s^2 < 0$ (IR time-like region), on the other hand, the horizon boundary conditions are complex resulting generically in a non-vanishing fermion spectral density. The IR space-like region is the primary feature of the finite c_{IR} theory that distinguishes it from the RN- AdS_4 theory. In particular, there is no analogue here of the log-oscillatory region (at $\omega = 0$), and thus no analogous mechanism for the disappearance of zero-frequency poles as p is increased. Indeed as we describe below, as p is increased, the zero-frequency pole moves but persists. Thus in the presence of superconductivity where the $U(1)$ symmetry is spontaneously broken, the gap in the fermion spectral density is determined by the Majorana scalar-fermion coupling η_5 . This implies that the parameter responsible for Mottness is distinct from those involved in superconductivity.

Figure 14 shows a density plot of the fermion spectral function for $p = 0$ (left) and $p = 3$ (right), where $q_\phi = 1.5$ and $\mu = 2\sqrt{3}$. The density plot for $p = 0$, which has been previously obtained in [36], is also shown for the purpose of comparison with the density plot for a non-zero value of p such as $p = 3$. As the plot in Figure 14(b) shows, turning on a non-zero value of p suppresses the spectral density of the incoherent excitations (those in the IR time-like region) and pushes them away from the boundary (the IR light-like region, depicted by solid black lines). But, since at $\omega = 0$ there is a Fermi peak (and indeed an infinite number of long-lived bound states for other values of ω in the IR space-like region), turning

⁴ See [35–40] where the authors analyze fermion correlators (in the absence of the Pauli interaction) in some superconducting backgrounds.

⁵ This condition is not required when $\eta_5 = 0$. We consider this condition so that our analysis here can be generalized to the case where $\eta_5 \neq 0$ [34]. Also, note that the convention of charge in [31, 36] is different than our convention in previous sections by a factor of two, namely $q_{\text{here}} = 2q_{\text{there}}$. In this discussion, we use the convention of [31, 36] for q and q_ϕ .

on p does not result in the formation of a gap. In fact, for the above-mentioned parameters, we find that, for $p = 3$, $k_F \approx 6.0$. For $p = 0$, the Fermi peak, for the same parameters, is at $k_F \approx 1.5$ [36]. Turning on p moves around the location of this Fermi peak as shown in Figure 15.

We note again that there are regions of parameter space (in particular, q_ϕ and m_ϕ) where the preferred geometry is either Reissner-Nordström or the superconducting geometry. Our results indicate that while the fermion gap in the superconducting geometry is controlled by the Majorana scalar-fermion coupling [31], a Mott gap can still form in the Reissner-Nordström regime. The decoupling of these two effects is promising in the context of the cuprate phase diagram.

Acknowledgments

We would like to thank E. Fradkin, and S. Hartnoll for discussions. M.E. and P.W.P. acknowledge financial support from the NSF DMR-0940992 and the Center for Emergent Superconductivity, a DOE Energy Frontier Research Center, Award Number DE-AC0298CH1088. R.G.L. is supported by DOE grant FG02-91-ER40709 and would like to thank the Galileo Galilei Institute for Theoretical Physics for support and the participants of the program ‘AdS/CFT and the Holographic States of Matter’ for many useful discussions.

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