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Responses of the CME-sensitive sine observable to resonance backgrounds in heavy-ion collisions

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A new sine observable, $R_{\Psi_2}(\Delta S)$, has been proposed to measure the Chiral Magnetic Effect (CME) in heavy-ion collisions; $\Delta S = \langle \sin \varphi_+ \rangle - \langle \sin \varphi_- \rangle$, where φ_{\pm} are azimuthal angles of positively and negatively charged particles relative to the reaction plane and averages are event-wise, and $R_{\Psi_2}(\Delta S)$ is a normalized event probability distribution. Preliminary STAR data reveal concave $R_{\Psi_2}(\Delta S)$ distributions in 200 GeV Au+Au collisions. Studies with A Multi-Phase Transport (AMPT) and Anomalous Viscous Fluid Dynamics (AVFD) models show concave $R_{\Psi_2}(\Delta S)$ distributions for CME signals and convex ones for typical resonance backgrounds. A recent hydrodynamic study, however, indicates concave shapes for backgrounds as well. To better understand these results, we report a systematic study of the elliptic flow (v_2) and transverse momentum (p_T) dependences of resonance backgrounds with toy-model simulations and Central Limit Theorem (CLT) calculations. It is found that the concavity or convexity of $R_{\Psi_2}(\Delta S)$ depends sensitively on the resonance v_2 (which yields different numbers of decay $\pi^+\pi^-$ pairs in the in-plane and out-of-plane directions) and p_T (which affects the opening angle of the decay $\pi^+\pi^-$ pair). Qualitatively, low p_T resonances decay into large opening-angle pairs and result in more "back-to-back" pairs out-of-plane, mimicking a CME signal, or a concave $R_{\Psi_2}(\Delta S)$. Supplemental studies of $R_{\Psi_3}(\Delta S)$ in terms of the triangular flow (v_3) , where only backgrounds exist but any CME would average to zero, are also presented.

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1. INTRODUCTION

Non-trivial topological gluon fields can form in quantum chromodynamics (QCD) from vacuum fluctuations [1]. Interactions with those gluon fields can change the chirality of quarks in local domains where the approximate chiral symmetry is restored [1–4]. Quarks of the same chirality in a local domain immersed in a strong magnetic field will move in opposite directions along the magnetic field if they bear opposite charges. This charge separation phenomenon is called the Chiral Magnetic Effect (CME) [4, 5].

Heavy-ion collisons provide a suitable environment for the CME to occur: the relativistic spectator protons can create an intense, transient magnetic field [6–9] roughly perpendicular to the reaction plane (spanned by the impact parameter and beam directions); high energy density can be created in the collision zone and the approximate chiral symmetry may be restored [10–14]; and topological gluon fields can emerge from the QCD vacuum [1]. Because the observation of the CME will simultaneously support the above pictures, the detection of such charge separations in heavy-ion collisions is of critical importance.

The common variable that has been used to search for the CME-induced charge separation is the so-called $\Delta\gamma$ variable [15]. Positive charge-dependent signals have been observed in heavy-ion collisions, qualitatively consistent with the CME [16–20]. However, the $\Delta\gamma$ variable is strongly contaminated by elliptic flow induced correlation backgrounds [21–25]. In fact, $\Delta\gamma$ measurements in small systems of p+Pb collisions at the LHC [26] and d+Au collisions at RHIC [27, 28], where only backgrounds are expected, reveal large signals comparable to those measured in heavy-ion collisions. With suppression of backgrounds by event-by-event and event-shapeengineering techniques, experimental data [29–31] show significantly reduced, consistent-with-zero signals for the CME.

Another variable that has been proposed to detect charge separation is the $R_{\Psi_2}(\Delta S)$ variable [32, 33]. We call it the *sine* observable. It is defined as follows. In each event, let

$$\varphi = \phi - \Psi_2, \tag{1}$$

$$\langle S_p \rangle = \frac{1}{N_p} \sum_{1}^{N_p} \sin(\varphi_+), \quad \langle S_n \rangle = \frac{1}{N_n} \sum_{1}^{N_n} \sin(\varphi_-), \quad (2)$$

$$\Delta S_{sep} = \langle S_p \rangle - \langle S_n \rangle , \qquad (3)$$

where ϕ is the particle azimuthal angle in the laboratory frame and φ is therefore the azimuthal angle relative to the second-order harmonic plane Ψ_2 (as a proxy for the unmeasured reaction plane). Subscripts (+, -) indicate the charge sign, and N_p, N_n are the number of particles with positive and negative charge, respectively. A

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parallel set of variables is constructed by randomizing the charges of all particles in the event, respecting the relative multiplicities of positive and negative particles. Then, according to the randomized charges,

$$\langle S_{p'} \rangle = \frac{1}{N'_p} \sum_{1}^{N'_p} \sin(\varphi'_+), \quad \langle S_{n'} \rangle = \frac{1}{N'_n} \sum_{1}^{N'_n} \sin(\varphi'_-), \quad (4)$$

$$\Delta S_{mix} = \langle S_{p'} \rangle - \langle S_{n'} \rangle , \qquad (5)$$

where the primes denote quantities for this so-called shuffled event. The ratio is formed from the event probability distributions of real events in ΔS_{sep} and shuffled events in ΔS_{mix} ,

$$C_{\Psi_2}(\Delta S) = \frac{N(\Delta S_{sep})}{N(\Delta S_{mix})}.$$
(6)

For events with CME signals, charge separation along the magnetic field gives $|\sin \varphi_{\pm}| \approx 1$ and a maximal difference $\sin \varphi_{+} - \sin \varphi_{-} \approx \pm 2$. The distribution of $N(\Delta S_{sep})$ would therefore become wider than its reference distribution. Here, the shuffled event $N(\Delta S_{mix})$ serves as the reference distribution. The ratio of $C_{\Psi_{2}}$ is therefore concave for CME [32, 33].

There can be background sources that change the shape of $C_{\Psi_2}(\Delta S)$. In order to eliminate reaction-plane (RP) independent backgrounds, an analogous variable $C_{\Psi_2}^{\perp}$ is constructed in the way identical to C_{Ψ_2} except changing each φ into $\varphi - \pi/2$. The R_{Ψ_2} variable is defined to be the ratio of C_{Ψ_2} to $C_{\Psi_2}^{\perp}$,

$$R_{\Psi_2}(\Delta S) = \frac{C_{\Psi_2}(\Delta S)}{C_{\Psi_2}^{\perp}(\Delta S)}.$$
(7)

The RP-independent backgrounds would cancel in $R_{\Psi_2}(\Delta S)$. Since the CME signal does not affect $C_{\Psi_2}^{\perp}$ significantly because $\sin(\varphi_{\pm} - \pi/2) \approx 0$, the CME in C_{Ψ_2} would survive in $R_{\Psi_2}(\Delta S)$, making it concave. The RP-dependent backgrounds, such as resonance decays with finite v_2 , can still affect $R_{\Psi_2}(\Delta S)$. However, it was shown to make $R_{\Psi_2}(\Delta S)$ convex [32, 33].

Preliminary STAR data reveal concave $R_{\Psi_2}(\Delta S)$ distributions in 200 GeV Au+Au collisions [34]. Previous studies using A Multi-Phase Transport (AMPT) model where resonance decay background is present but no CME, suggest that $R_{\Psi_2}(\Delta S)$ is convex [33]. Anomalous-Viscous Fluid Dynamics (AVFD) model shows concave $R_{\Psi_2}(\Delta S)$ distributions for CME signals and convex ones for typical resonance backgrounds [33]. A recent hydrodynamic study, however, indicates concave shapes for backgrounds as well [35].

To better understand these results, we present a systematic study of resonance backgrounds as functions of the resonance elliptic flow (v_2) and transverse momentum (p_T) with toy-model simulations and Central Limit Theorem (CLT) calculations. It is found that the concavity or convexity of $R_{\Psi_2}(\Delta S)$ depends sensitively on the resonance v_2 (which yields different numbers of decay $\pi^+\pi^-$ pairs in the in-plane and out-of-plane directions) and p_T (which affects the opening angle of the decay $\pi^+\pi^-$ pair).

Supplemental studies in terms of the triangular flow (v_3) , where only backgrounds exist but any CME would average to zero, are also presented.

2. TOY-MODEL SIMULATION OF RESONANCE BACKGROUNDS

We use a toy model of ρ meson decays to study the behavior of $R_{\Psi_2}(\Delta S)$ as functions of the ρ kinematic variables. The toy model has been used for CME background studies in Ref. [36]. It generates events to be composed of primordial pions and ρ -decay pions. Their input p_T distributions and $v_2(p_T)$ are obtained from data measurements [36–46]. For simplicity, we use the input harmonic plane Ψ_2 (as well as Ψ_3 discussed in Sec. 3) in our analysis.

In order to study the v_2 dependence, we scale $v_{2,\rho}$ (v_2 of ρ) up or down by a p_T -independent factor to investigate how $R_{\Psi_2}(\Delta S)$ responds. Figure 1 shows the results; the curve of C_{Ψ_2} becomes more concave when $v_{2,\rho}$ is increased, and $C_{\Psi_2}^{\perp}$ behaves in the opposite way. Subsequently, $R_{\Psi_2}(\Delta S)$ becomes more concave. This behavior can be qualitatively understood as follows. At the typical resonance p_T in the simulation, the decay daughters are close to each other in azimuthal angle. The numerator of C_{Ψ_2} has the term: $\sin \varphi_+ - \sin \varphi_- \approx \cos \bar{\varphi} \delta \varphi$, where $\bar{\varphi} =$ $(\varphi_+ + \varphi_-)/2, \, \delta \varphi = \varphi_+ - \varphi_-$ are the average and difference of the π^{\pm} azimuths, respectively. When $v_{2,\rho}$ is large, $\bar{\varphi}$ will be relatively close to 0 or π , and $|\cos \bar{\varphi}|$ will be relatively big. Hence, the ΔS in the numerator of C_{Ψ_2} has a wider distribution, and accordingly C_{Ψ_2} becomes more concave (see Fig. 1a). Similarly, the numerator of $C_{\Psi_2}^{\perp}$ has the term: $\sin(\varphi_+ - \pi/2) - \sin(\varphi_- - \pi/2) \approx \sin \bar{\varphi} \delta \varphi$. When $v_{2,\rho}$ is large, $|\sin \bar{\varphi}|$ will be relatively small and close to 0, so the ΔS in the numerator of $C_{\Psi_2}^{\perp}$ has a narrower distribution, and accordingly $C_{\Psi_2}^{\perp}$ becomes more convex (Fig. 1b). Because of the opposite behaviors of C_{Ψ_2} and $C_{\Psi_2}^{\perp}$, we can easily get the dependence of their ratio R_{Ψ_2} on $v_{2,\rho}$: its concavity increases with increasing $v_{2,\rho}$ (Fig. 1c).

Note that the curves in Fig. 1c with zero $v_{2,\rho}$ is counterintuitively nonflat. This is due to the finite $v_{2,\pi}$ (primordial pion v_2). The ρ decays alter the pion multiplicities which affect C_{Ψ_2} and $C_{\Psi_2}^{\perp}$. The finite $v_{2,\pi}$ breaks the symmetry between C_{Ψ_2} and $C_{\Psi_2}^{\perp}$, resulting in the slightly nonflat $R_{\Psi_2}(\Delta S)$. Figure 2 shows R_{Ψ_2} curves with zero $v_{2,\rho}$ for various values of $v_{2,\pi}$. Only weak dependences on $v_{2,\pi}$ are observed for $R_{\Psi_2}(\Delta S)$ (and also $C_{\Psi_2}, C_{\Psi_2}^{\perp}$). When both $v_{2,\rho}$ and $v_{2,\pi}$ are set to zero, then R_{Ψ_2} is indeed flat.

To scan $p_{T,\rho}$ (the p_T of ρ), we fix $v_{2,\rho}$ to a specific value 0.06, because otherwise the value of $v_{2,\rho}$ would be affected by the changing $p_{T,\rho}$. The $v_{2,\pi}$ and p_T of the pri-



FIG. 1. (color online) Observable distributions for various values of $v_{2,\rho}$ (with $v_{2,\pi}$ fixed to its default distribution). Here, $v_{2,\rho}^{def}$ is the default distribution of $v_{2,\rho}$ obtained from data [36–46].



FIG. 2. (color online) $R_{\Psi_2}(\Delta S)$ for various values of $v_{2,\pi}$ (with $v_{2,\rho}$ fixed to 0). Here, $v_{2,\pi}^{def}$ is the default distribution of $v_{2,\pi}$ obtained from data [36–46].

mordial pions are given by default. We find the curves of C_{Ψ_2} , $C_{\Psi_2}^{\perp}$, and R_{Ψ_2} to become more convex when $p_{T,\rho}$ increases (Fig. 3). This is because of the following. When $p_{T,\rho}$ is large, the decay opening angle $\delta\varphi$ is small. The $\cos \bar{\varphi} \delta \varphi$ contribution to ΔS in C_{Ψ_2} and the $\sin \bar{\varphi} \delta \varphi$ contribution to ΔS in $C_{\Psi_2}^{\perp}$ both become small in magnitude, so the distributions of ΔS in both C_{Ψ_2} and $C_{\Psi_2}^{\perp}$ become narrower. The reshuffled ΔS in the denominators of C_{Ψ_2} and $C_{\Psi_2}^{\perp}$ are not as sensitive to the $\delta\varphi$ change as the numerators. Thus, the shapes of C_{Ψ_2} and $C_{\Psi_2}^{\perp}$ both become more convex. Since the change in $\cos \bar{\varphi} \delta \varphi$ is larger than in $\sin \bar{\varphi} \delta \varphi$ with increasing p_T for $\bar{\varphi}$ close to the reaction plane, the narrowing in C_{Ψ_2} is more significant, so R_{Ψ_2} becomes more convex.

Another way to explain the R_{Ψ_2} change is as follows. When $p_{T,\rho}$ is high, the two decay daughters are close to each other and preferentially close to the reaction plane because of the finite $v_{2,\rho}$. This is characteristic of the CME background. At low $p_{T,\rho}$, the two daughters are preferentially more perpendicular to the RP because of the large decay opening angle. This case resembles the CME signal, so the R_{Ψ_2} curves with lower $p_{T,\rho}$ becomes more concave, just like how CME signal would behave. For our typical p_T distribution from data, the high $p_{T,\rho}$ case wins over the case with low $p_{T,\rho}$.

The behaviors of C_{Ψ_2} and $C_{\Psi_2}^{\perp}$ are recapitulated in

Fig. 4 by the RMS (Root Mean Square) of C_{Ψ_2} and $C_{\Psi_2}^{\perp}$.

We summerize our main findings as follow:

- The curve of C_{Ψ_2} becomes more concave when $v_{2,\rho}$ increases, and $C_{\Psi_2}^{\perp}$ more convex, rendering a more concave R_{Ψ_2} .
- The shapes of the observables $(C_{\Psi_2}, C_{\Psi_2}^{\perp}, \text{ and } R_{\Psi_2})$ are only weakly dependent on $v_{2,\pi}$.
- The curves of C_{Ψ_2} and $C_{\Psi_2}^{\perp}$ become more convex when $p_{T,\rho}$ increases. The effect is more significant in C_{Ψ_2} , rendering a more convex R_{Ψ_2} .

3. SUPPLEMENTAL STUDIES USING v₃

The CME is a charge separation with respect to the RP (or the v_2 harmonic plane Ψ_2). The CME-induced charge separation must be zero with respect to the third order harmonic plane because of its random orientation relative to Ψ_2 . Resonance backgrounds, on the other hand, should be still finite with respect to Ψ_3 . In this section, we verify this with our toy model simulation.

In term of v_3 , the reference azimuthal angle is the third harmonic plane:

$$\varphi = \phi - \Psi_3. \tag{8}$$

There have been two different ways to define the sine observables for v_3 , and both are similar to the definition of the observables for v_2 .

A. For the first definition [33], one changes Ψ_2 into Ψ_3 for φ (see Eq. 1, 8) and replaces $-\pi/2$ by $-\pi/3$ for ΔS (both ΔS_{sep} and ΔS_{mix}) in $C_{\Psi_3}^{\perp}$,

$$C_{\Psi_3}^{\perp}: \quad \Delta S = \frac{1}{N_p} \sum_{1}^{N_p} \sin\left(\varphi_{+} - \frac{\pi}{3}\right) - \frac{1}{N_n} \sum_{1}^{N_n} \sin\left(\varphi_{-} - \frac{\pi}{3}\right)$$
(9)

B. For the second definition [35], one still changes Ψ_2 into Ψ_3 for φ . In addition, one adds a factor 3/2 in front



FIG. 3. (color online) Observable distributions for various values of $p_{T,\rho}$ (with $v_{2,\rho}$ fixed to 0.06).

of the azimuths,

$$C_{\Psi_3}: \quad \Delta S = \frac{1}{N_p} \sum_{1}^{N_p} \sin\left(\frac{3}{2}\varphi_+\right) - \frac{1}{N_n} \sum_{1}^{N_n} \sin\left(\frac{3}{2}\varphi_-\right),$$

$$C_{\Psi_3}^{\perp}: \quad \Delta S = \frac{1}{N_p} \sum_{1}^{N_p} \sin\left(\frac{3}{2}\varphi_+ - \frac{\pi}{2}\right) - \frac{1}{N_n} \sum_{1}^{N_n} \sin\left(\frac{3}{2}\varphi_- - \frac{\pi}{2}\right)$$
(10)

We use the Toy Monte Carlo to investigate $R_{\Psi_3}(\Delta S)$ of those two definitions. The toy simulation generates primordial π^+, π^- and ρ with the experimental p_T spectra but with only v_3 of the ρ . Since Ψ_2 and Ψ_3 are uncorrelated, including non-zero v_2 does not change the results. Including a finite v_3 for the primordial pions does not have significant effect. The default function of $v_{3,\rho}(p_T)$ is approximated by that of $v_{2,\rho}(p_T)$ but with half magnitude, i.e.

$$v_{3,\rho}(p_T) = \frac{1}{2} v_{2,\rho}(p_T).$$
(11)

We constrain the azimuthal range to be $[0, 2\pi)$ in our simulation. As will be discussed later in Sec. 4.3, the sine observables of Definition **B** unfortunately depend on which periodic range is used, suggesting Definition **B** is not a physically correct definition. The simulation results are shown in Fig. 5 and Fig. 6.

We make the following observations:

• In Definition **A**, R_{Ψ_3} is always flat.

By Definition **A** itself, R_{Ψ_3} should always be flat, as follows. The Probability Density Function (PDF) is $f(\varphi) = (1 + 2v_3 \cos(3\varphi))/(2\pi)$, whose period is $2\pi/3$. In the definition of $C_{\Psi_3}^{\perp}$, φ is shifted by $\pi/3$ clockwise, $\Delta S(\varphi) \rightarrow \Delta S(\varphi - \pi/3)$. If we keep shifting ΔS by another period in the same direction, we would not change the distribution of ΔS in $C_{\Psi_3}^{\perp}$, which means $\Delta S(\varphi - \pi/3)$ and $\Delta S(\varphi - \pi)$ have the same distribution. From the Definition **A**, we also know that $\Delta S(\varphi - \pi) = -\Delta S(\varphi)$. Because the distribution of $\Delta S(\varphi)$ is symmetric about $\Delta S = 0$, $\Delta S(\varphi)$ and $-\Delta S(\varphi)$ have the same distribution as well. Thus, $\Delta S(\varphi)$ and $\Delta S(\varphi - \pi/3)$ have the same distribution, which means that C_{Ψ_3} and $C_{\Psi_3}^{\perp}$ have the same shape and R_{Ψ_3} must be flat and have the value 1.

This flat R_{Ψ_3} can also be explained by the analysis based on CLT in Sec. 4.

- $\left(\frac{\pi}{2}\right)$ The C_{Ψ_3} and $C_{\Psi_3}^{\perp}$ curves from Definition **A** show a similar dependence on resonance $p_{T,\rho}$ as C_{Ψ_2} and $C_{\Psi_2}^{\perp}$ curves in the v_2 case.
 - The C_{Ψ_3} , $C_{\Psi_3}^{\perp}$, and R_{Ψ_3} curves from Definition **B** are obviously dependent on the $p_{T,\rho}$ and $v_{3,\rho}$. Increasing $p_{T,\rho}$ makes the curves more convex. Increasing $v_{3,\rho}$ makes the C_{Ψ_3} , R_{Ψ_3} curves more concave, and $C_{\Psi_3}^{\perp}$ more convex. Those tendencies are consistent with the scans with respect to v_2 .
 - In Definition **B**, the $C_{\Psi_3}^{\perp}$ and R_{Ψ_3} curves are counterintuitively not flat, even if we set v_3 to zero.

We note that Definition **A** was used only in the early version (version 2) of Ref. [33] where the $R_{\Psi_3}(\Delta S)$ variable was studied with respect to v_3 . In the later version 3 of Ref. [33], Definition **B** was used.

4. ANALYTICAL RESULTS BASED ON CENTRAL LIMIT THEOREM

In this section, we use the Central Limit Theorem (CLT) to analyze the sine observable. This analysis can be applied to all observables discussed in this paper. With a few reasonable approximations, the behavior of the sine observable can be readily understood.

There are many versions of the CLT, and here we use Lindeberg-Levy's expression. Let X_1, X_2, \ldots, X_n be a sequence of independent and identically distributed (i.d.d.) random variables with expectation value $E[X_i] = \mu$ and variance $Var[X_i] = \sigma^2 < \infty$, and

$$S_n = \frac{X_1 + X_2 + \dots + X_n}{n} = \frac{1}{n} \sum_{i=1}^n X_i \qquad (12)$$

denote their mean. As n approaches infinity, the random variable $\sqrt{n}(S_n - \mu)$ converges in distribution to a normal



FIG. 4. (color online) RMS of C_{Ψ_2} and $C_{\Psi_2}^{\perp}$ depending on $v_{2,\rho}$ and $p_{T,\rho}$. (a) RMS of C_{Ψ_2} and $C_{\Psi_2}^{\perp}$ (shown in Fig. 1) depending on $v_{2,\rho}$ (with $v_{2,\pi}$ fixed to its default distribution). (b) RMS of C_{Ψ_2} and $C_{\Psi_2}^{\perp}$ (shown in Fig. 3) depending on $p_{T,\rho}$ (with $v_{2,\rho}$ fixed to 0.06 and $v_{2,\pi}$ fixed to its default distribution).



FIG. 5. (color online) Definition **A**: observable distributions for various values of $v_{3,\rho}$ and $p_{T,\rho}$ (with $v_{3,\pi}$ fixed to 0). The C_{Ψ_3} and $C_{\Psi_3}^{\perp}$ curves, with the same $p_{T,\rho}$ but various $v_{3,\rho}$, are very close to each other in figure (a) and (b) (concave dashed lines for low $p_{T,\rho}$, and convex solid lines for high $p_{T,\rho}$).



FIG. 6. (color online) Definition **B**: The upper plots (a–c) show observable distributions for various values of $v_{3,\rho}$ (with $v_{3,\pi}$ fixed to 0). Here, $v_{3,\rho}^{def}$ is the default distribution of $v_{3,\rho}$ approximated by $0.5v_{2,\rho}^{def}$ (Eq. 11). The lower plots (d–f) show observable distributions for various values of $p_{T,\rho}$ (with $v_{3,\rho}$ fixed to 0.03 and $v_{3,\pi}$ fixed to 0).

 $\mathcal{N}(0, \sigma^2)$. Generally, if X_1, X_2, \ldots, X_n are independent normal distributions,

$$X_i \sim \mathcal{N}(\mu_i, \sigma_i^2),$$
 (13)

then the weighted sum of them is a normal distribution,

$$\sum_{i=1}^{n} a_i X_i \sim \mathcal{N}\left(\sum_{i=1}^{n} a_i \mu_i, \sum_{i=1}^{n} a_i^2 \sigma_i^2\right).$$
(14)

4.1. Analysis of C_{Ψ_2} , $C_{\Psi_2}^{\perp}$, and R_{Ψ_2}

First, we write the PDF of ϕ ,

$$f(\phi) = \frac{1}{2\pi} \left(1 + 2\sum_{m=2}^{\infty} v_m \cos(m(\phi - \Psi_m)) \right), \quad (15)$$

where Ψ_m are normally different and uncorrelated among different m. When we focus only on one specific m, for example m = 2 or 3 in the former discussion, we can just use $\varphi = \phi - \Psi_m$ as the relative azimuth of particles.

4.1.1. Numerator of C_{Ψ_2}

The PDF of ΔS_{sep} can describe $N(\Delta S_{sep})$, the numerator of C_{Ψ_2} . For simplicity, we assume that the number of positive charges is the same as the number of negative charges in the final state. In each event, before any decay, n_{ρ} denotes the number of ρ mesons, and n_{π} denotes the number of primordial pions. Thus,

$$N_n = N_p = n_\rho + 0.5n_\pi.$$
(16)

We rewrite

$$\Delta S_{sep} = \frac{1}{n_{\rho} + 0.5n_{\pi}} \left(\sum_{1}^{n_{\rho}} (\sin \varphi_{+} - \sin \varphi_{-}) \right) + \frac{1}{n_{\rho} + 0.5n_{\pi}} \left(\sum_{1}^{0.5n_{\pi}} (\sin \varphi_{+} - \sin \varphi_{-}) \right).$$
(17)

The first sum is over ρ decay pions, and the second is over primordial pions.

For convenience, we will use the following short-hand notations:

$$c := \cos\varphi, \quad \bar{c} := \cos\bar{\varphi}, \quad s := \sin\varphi, \quad \bar{s} := \sin\bar{\varphi}, \\ \delta := 2\sin(\delta\varphi/2)$$
(18)

where $\bar{\varphi} = (\varphi_+ + \varphi_-)/2$ is related to the ρ angular position and $\delta \varphi = \varphi_+ - \varphi_-$ represents the decay opening angle. We use the indices ρ or π to indicate whether the variables are for ρ or primordial π^{\pm} .

We express the first sum of Eq. 17 into

$$\sum_{1}^{n_{\rho}} (\sin\varphi_{+} - \sin\varphi_{-}) = \sum_{1}^{n_{\rho}} 2\cos(\bar{\varphi})\sin(\delta\varphi/2) = \sum_{1}^{n_{\rho}} \bar{c}_{\rho}\delta.$$
(19)

Because the primordial pions all independently obey the same distribution related to the global harmonic plane, we rewrite the second sum of Eq. 17 into

$$\sum_{1}^{0.5n_{\pi}} (\sin \varphi_{+} - \sin \varphi_{-}) = \sum_{1}^{0.5n_{\pi}} \sin \varphi_{+} - \sum_{1}^{0.5n_{\pi}} \sin \varphi_{-}.$$
 (20)

We make two assumptions: (1) In a resonance decay, $\bar{\varphi}$ could be regarded as an approximation for φ_{ρ} , so the PDF of $\bar{\varphi}$ is the same as the PDF of φ_{ρ} ; (2) For two tracks from one resonance decay, $\cos \bar{\varphi}$ and $2\sin(\delta \varphi/2)$ are independent.

From symmetry, $E[\delta] = E[2\sin(\delta\varphi/2)] = 0$ at any given \bar{c}_{ρ} , so:

$$\mathbf{E}[\bar{c}_{\rho}\delta] = \mathbf{E}[\bar{c}_{\rho}]\mathbf{E}[\delta] = \mathbf{E}[\bar{c}_{\rho}] \times 0 = 0.$$
(21)

We therefore get

$$\operatorname{Var}[\bar{c}_{\rho}\delta] = \operatorname{Var}[\bar{c}_{\rho}]\operatorname{Var}[\delta] + \operatorname{E}[\bar{c}_{\rho}]^{2}\operatorname{Var}[\delta] + \operatorname{Var}[\bar{c}_{\rho}]\operatorname{E}[\delta]^{2}$$
$$= \operatorname{E}\left[\bar{c}_{\rho}^{2}\right]\operatorname{Var}[\delta].$$
(22)

In our simulations, n_{ρ} is a Poisson distribution, so to get the variance of $\sum_{1}^{n_{\rho}} \bar{c}_{\rho} \delta$ is a problem of Compound Poisson Distribution. Thus, we have

$$\operatorname{Var}\left[\sum_{i}^{n_{\rho}} \bar{c}_{\rho} \delta\right] = \operatorname{E}\left[n_{\rho}\right] \operatorname{Var}\left[\bar{c}_{\rho} \delta\right] + \operatorname{E}\left[\bar{c}_{\rho} \delta\right]^{2} \operatorname{Var}\left[n_{\rho}\right]$$

$$= \operatorname{E}\left[n_{\rho}\right] \operatorname{Var}\left[\bar{c}_{\rho} \delta\right].$$
(23)

Eq. 23 indicates that it makes no difference whether n_{ρ} is a single value or a Poisson distribution. For simplicity, we can just use n_{ρ} as if it is fixed to a specific value. According to CLT,

$$\sum_{1}^{n_{\rho}} \bar{c}_{\rho} \delta \sim \mathcal{N}\left(0, n_{\rho} \operatorname{Var}[\delta] \operatorname{E}\left[\bar{c}_{\rho}^{2}\right]\right).$$
(24)

The PDF of φ of the primordial pions has the same form as Eq. 15, so we can readily obtain the variances (Var[c_{π}] and Var[s_{π}]). As for the term about primordial pions, the two terms in the right hand side of Eq. 20 should have the same distribution. The discussion about n_{π} is as same as the discussion of n_{ρ} . According to CLT, we have

$$\sum_{1}^{0.5n_{\pi}} \sin \varphi_{+}, \sum_{1}^{0.5n_{\pi}} \sin \varphi_{-} \sim \mathcal{N}(0.5n_{\pi} \mathrm{E}[s_{\pi}], 0.5n_{\pi} \mathrm{Var}[s_{\pi}]),$$
(25)

so the difference is

$$\sum_{1}^{0.5n_{\pi}} \sin \varphi_{+} - \sum_{1}^{0.5n_{\pi}} \sin \varphi_{-} \sim \mathcal{N}(0, n_{\pi} \operatorname{Var}[s_{\pi}]).$$
(26)

Finally, we write ΔS in our new notation,

$$\Delta S_{sep} = \frac{\sum_{1}^{n_{\rho}} \bar{c}_{\rho} \delta + \sum_{1}^{n_{\pi}/2} s_{\pi^{+}} - \sum_{1}^{n_{\pi}/2} s_{\pi^{-}}}{n_{\rho} + 0.5 n_{\pi}}, \quad (27)$$

where s_{π^+} and s_{π^-} are the sine values for π^+ and π^- from resonance decay, and they obey the same distribution independently, so we just call them both s_{π} . According to CLT,

$$\Delta S_{sep} \sim \mathcal{N}\left(0, \frac{n_{\rho} \operatorname{Var}[\delta] \operatorname{E}\left[\bar{c}_{\rho}^{2}\right] + n_{\pi} \operatorname{Var}[s_{\pi}]}{(n_{\rho} + 0.5 n_{\pi})^{2}}\right) \qquad (28)$$
$$:= \mathcal{N}\left(0, \sigma_{\uparrow}^{2}\right).$$

4.1.2. Denominator of C_{Ψ_2}

The PDF of ΔS_{mix} can describe $N(\Delta S_{mix})$, the denominator of C_{Ψ_2} . The analysis here is very similar to the analysis of ΔS_{sep} . In shuffling, we keep the number of positive charges to be still the same as the number of negative charges.

$$N'_{n} = N'_{p} = n_{\rho} + 0.5n_{\pi} \tag{29}$$

Relaxing this requirement to an average level does not change our results.

After shuffling, all the pions are independent, no matter whether they are primordial or from resonance decays. For pions from resonance decays, the pion azimuth can be written as $\varphi = \bar{\varphi} + \delta \varphi/2 \approx \varphi_{\rho} + \delta \varphi/2$. Because the distribution of $\delta \varphi$ is symmetric about $\delta \varphi = 0$, we just use $+\delta\varphi/2$ here. The expression of ΔS_{mix} can therefore be rewritten as:

$$\Delta S_{mix} = \frac{1}{n_{\rho} + 0.5n_{\pi}} \left(\sum_{1}^{n_{\rho}} (\sin\varphi_{+} - \sin\varphi_{-}) \right) + \frac{1}{n_{\rho} + 0.5n_{\pi}} \left(\sum_{1}^{0.5n_{\pi}} (\sin\varphi_{+} - \sin\varphi_{-}) \right) = \frac{\sum_{1}^{n_{\rho}} \sin(\varphi_{\rho} + \delta\varphi_{+}/2) - \sum_{1}^{n_{\rho}} \sin(\varphi_{\rho} + \delta\varphi_{-}/2)}{n_{\rho} + 0.5n_{\pi}} + \frac{\sum_{1}^{n_{\pi}/2} s_{\pi^{+}} - \sum_{1}^{n_{\pi}/2} s_{\pi^{-}}}{n_{\rho} + 0.5n_{\pi}}.$$
(30)

The second term is already calculated in Eq. 26, and we calculate the distribution of the first term as

$$\sum_{1}^{n_{\rho}} \sin\left(\varphi_{\rho} + \frac{\delta\varphi_{+}}{2}\right) - \sum_{1}^{n_{\rho}} \sin\left(\varphi_{\rho} + \frac{\delta\varphi_{-}}{2}\right)$$

$$\sim \mathcal{N}\left(0, 2n_{\rho} \operatorname{Var}\left[\sin\left(\varphi_{\rho} + \frac{\delta\varphi}{2}\right)\right]\right),$$
(31)

The first and the second moment below are needed in order to complete the calculation of the variance:

$$E\left[\sin\left(\varphi_{\rho} + \frac{\delta\varphi}{2}\right)\right] \\
 =E\left[\sin\varphi_{\rho}\right] E\left[\cos\frac{\delta\varphi}{2}\right] + E\left[\cos\varphi_{\rho}\right] E\left[\sin\frac{\delta\varphi}{2}\right] \quad (32) \\
 =E\left[\sin\varphi_{\rho}\right] E\left[\cos\frac{\delta\varphi}{2}\right],$$

$$E\left[\sin^{2}\left(\varphi_{\rho} + \frac{\delta\varphi}{2}\right)\right]$$

$$= \frac{1}{2} - \frac{1}{2}E\left[\cos(2\varphi_{\rho})\cos(\delta\varphi)\right] + \frac{1}{2}E\left[\sin(2\varphi_{\rho})\sin(\delta\varphi)\right]$$

$$- \frac{1}{2} - \frac{1}{2}E\left[\left(1 - 2\sin^{2}(\varphi_{\rho})\right)\left(1 - \frac{1}{2}\delta^{2}\right)\right] + 0$$

$$= \overline{2} - \overline{2} \operatorname{E} \left[\left(1 - 2 \sin^2 \varphi_\rho \right) \left(1 - \overline{2} \delta^2 \right) \right] + 0$$
$$= \operatorname{E}[s_\rho^2] + \frac{1}{4} \operatorname{Var}[\delta] - \frac{1}{2} \operatorname{E}[s_\rho^2] \operatorname{Var}[\delta].$$
(33)

The last step uses the fact that $E[\delta] = 0$ and therefore $\operatorname{Var}[\delta] = \operatorname{E}[\delta^2].$

Thus, we can get the distribution of ΔS_{mix} ,

1 [/

$$\Delta S_{mix} \sim \mathcal{N}\left(0, \frac{2n_{\rho} \operatorname{Var}\left[\sin\left(\varphi_{\rho} + \frac{\delta\varphi}{2}\right)\right] + n_{\pi} \operatorname{E}[s_{\pi}^{2}]}{(n_{\rho} + 0.5n_{\pi})^{2}}\right)$$
$$=: \mathcal{N}\left(0, \sigma_{\downarrow}^{2}\right).$$
(34)

4.1.3. Shape of C_{Ψ_2}

We use the PDF of normal distribution Gaussian function:

$$f(x|\mu,\sigma) := \frac{1}{\sqrt{2\pi\sigma^2}} \exp\left(-\frac{(x-\mu)^2}{2\sigma^2}\right).$$
(35)

The shape of C_{Ψ_2} is described by the ratio of the PDF of ΔS_{sep} to the PDF of ΔS_{mix} . Using Gaussian functions for those PDFs, the shape of C_{Ψ_2} is

$$C_{\Psi_2}(x) = \frac{f(x|0,\sigma_{\uparrow})}{f(x|0,\sigma_{\downarrow})} = \frac{\sigma_{\downarrow}}{\sigma_{\uparrow}} \exp\left[-\frac{x^2}{2}\left(\frac{1}{\sigma_{\uparrow}^2} - \frac{1}{\sigma_{\downarrow}^2}\right)\right].$$
(36)

Here, x denotes ΔS (representing ΔS_{sep} or ΔS_{mix}).

4.1.4. Shape of $C_{\Psi_2}^{\perp}$

The analysis of $C_{\Psi_2}^{\perp}$ is nearly the same as that of C_{Ψ_2} by shifting the relative azimuth φ by a centain angle: $\varphi' = \varphi - \xi$. Accordingly, we use the parallel short-hand notations as follows:

$$c' := \cos \varphi', \quad \bar{c}' := \cos \bar{\varphi}', \quad s' := \sin \varphi', \\ \bar{s}' := \sin \bar{\varphi}', \quad \delta := 2 \sin(\delta \varphi'/2) = 2 \sin(\delta \varphi/2).$$
(37)

Then, the format of variances here is just like before:

$$\sigma_{\perp\uparrow}^2 = \frac{n_{\rho} \operatorname{Var}[\delta] \operatorname{E}[\bar{c}_{\rho}^{\prime 2}] + n_{\pi} \operatorname{Var}[s_{\pi}^{\prime}]}{(n_{\rho} + 0.5 n_{\pi})^2}, \qquad (38)$$

$$\sigma_{\perp\downarrow}^2 = \frac{2n_{\rho} \operatorname{Var}\left[\sin\left(\varphi_{\rho}' + \frac{\delta\varphi}{2}\right)\right] + n_{\pi} \operatorname{E}[s_{\pi}'^2]}{(n_{\rho} + 0.5n_{\pi})^2}.$$
 (39)

The shape of $C_{\Psi_2}^{\perp}$ is

$$C_{\Psi_2}^{\perp}(x) = \frac{f(x|0,\sigma_{\perp\uparrow})}{f(x|0,\sigma_{\perp\downarrow})} = \frac{\sigma_{\perp\downarrow}}{\sigma_{\perp\uparrow}} \exp\left[-\frac{x^2}{2}\left(\frac{1}{\sigma_{\perp\uparrow}^2} - \frac{1}{\sigma_{\perp\downarrow}^2}\right)\right]. \tag{40}$$

4.1.5. Shape of R_{Ψ_2}

According to the definition of R_{Ψ_2} , the shape of R_{Ψ_2} is given by

$$R_{\Psi_{2}}(x) = \frac{f(x|0,\sigma_{\uparrow})}{f(x|0,\sigma_{\downarrow})} \left/ \frac{f(x|0,\sigma_{\perp\uparrow})}{f(x|0,\sigma_{\perp\downarrow})} \right.$$
$$= \frac{\sigma_{\downarrow}\sigma_{\perp\uparrow}}{\sigma_{\uparrow}\sigma_{\perp\downarrow}} \exp\left[-\frac{x^{2}}{2} \left(\frac{1}{\sigma_{\uparrow}^{2}} - \frac{1}{\sigma_{\downarrow}^{2}} - \frac{1}{\sigma_{\perp\uparrow}^{2}} + \frac{1}{\sigma_{\perp\downarrow}^{2}} \right) \right]. \tag{41}$$

Thus, whether R_{Ψ_2} is convex or concave is determined by the following parameter:

$$\zeta := \frac{1}{\sigma_{\uparrow}^2} - \frac{1}{\sigma_{\downarrow}^2} - \frac{1}{\sigma_{\bot\uparrow}^2} + \frac{1}{\sigma_{\bot\downarrow}^2}.$$
 (42)

- If $\zeta > 0$, then R_{Ψ_2} is convex, and the more positive ζ is, the more convex R_{Ψ_2} will be.
- If ζ < 0, then R_{Ψ2} is concave, and the more negative ζ is, the more concave R_{Ψ2} will be.
- If $\zeta = 0$, then R_{Ψ_2} is flat.

4.2. CLT analysis for v_2

If we only focus on v_2 , the PDF in Eq. 15 can be simplified as

$$f(\varphi) = \frac{1}{2\pi} (1 + 2v_2 \cos(2\varphi)).$$
 (43)

From the definition of $C_{\Psi_2}^{\perp}$ for v_2 , the relative azimuth is shifted by $\xi = \pi/2$. Thus,

$$\begin{split} \mathbf{E}[c_{\rho}^{2}] &= \mathbf{E}[\bar{c}_{\rho}^{2}] = \frac{1 + v_{2,\rho}}{2}, \quad \mathbf{E}[s_{\rho}^{2}] = \frac{1 - v_{2,\rho}}{2}, \\ \mathbf{E}[c_{\rho}^{\prime 2}] &= \mathbf{E}[\bar{c}_{\rho}^{\prime 2}] = \frac{1 - v_{2,\rho}}{2}, \quad \mathbf{E}[s_{\rho}^{\prime 2}] = \frac{1 + v_{2,\rho}}{2}, \quad (44) \\ \mathbf{E}[s_{\pi}^{2}] &= \frac{1 - v_{2,\pi}}{2}, \quad \mathbf{E}[s_{\pi}^{\prime 2}] = \frac{1 + v_{2,\pi}}{2}. \end{split}$$

We can easily get that the first moment of $\sin (\varphi_{\rho} + \delta \varphi/2)$ in Eq. 32 is 0, so its variance is equal to its second moment which can be expressed as Eq. 33 by the terms in Eq. 44. After slightly changing the sequence in the expression of ζ , we have

$$\begin{split} \zeta &= \left(\frac{1}{\sigma_{\uparrow}^{2}} - \frac{1}{\sigma_{\perp\uparrow}^{2}}\right) - \left(\frac{1}{\sigma_{\downarrow}^{2}} - \frac{1}{\sigma_{\perp\downarrow}^{2}}\right) \\ &= \left(\frac{2(n_{\rho} + 0.5n_{\pi})^{2}}{n_{\rho} \operatorname{Var}[\delta](1 + v_{2,\rho}) + n_{\pi}(1 - v_{2,\pi})} \\ &- \frac{2(n_{\rho} + 0.5n_{\pi})^{2}}{n_{\rho} \operatorname{Var}[\delta](1 - v_{2,\rho}) + n_{\pi}(1 + v_{2,\pi})}\right) \\ &- \left(\frac{2(n_{\rho} + 0.5n_{\pi})^{2}}{2n_{\rho}(1 - v_{2,\rho}) + n_{\rho}v_{2,\rho} \operatorname{Var}[\delta] + n_{\pi}(1 - v_{2,\pi})} \\ &- \frac{2(n_{\rho} + 0.5n_{\pi})^{2}}{2n_{\rho}(1 + v_{2,\rho}) - n_{\rho}v_{2,\rho} \operatorname{Var}[\delta] + n_{\pi}(1 + v_{2,\pi})}\right). \end{split}$$
(45)

For further insights, we make two more assumptions (in addition to those in Sec. 4.1.1): (3) The magnitude of v_2 (including $v_{2,\rho}$ and $v_{2,\pi}$) is much smaller than 1. In our simulations, they are around 0.1; (4) In each event, the number of primordial pions are much larger than the number of ρ mesons. In our simulations, $n_{\pi} \approx 10n_{\rho}$.

In our simulations, $v_{2,\rho}$, $v_{2,\pi}$, and n_{ρ}/n_{π} are of the same order of magnitude (~ 0.1). To the leading order of them,

$$\zeta = \frac{n_{\rho}}{n_{\pi}} (2n_{\rho} + n_{\pi})^2 \left(\frac{4v_{2,\pi} - 2v_{2,\rho} - 2v_{2,\pi} \operatorname{Var}[\delta]}{n_{\pi} + n_{\rho} \left(4 + 2\operatorname{Var}[\delta]\right)} \right).$$
(46)

The first derivatives are

$$\frac{\partial \zeta}{\partial v_{2,\rho}} = \frac{n_{\rho}}{n_{\pi}} (2n_{\rho} + n_{\pi})^2 \left(\frac{-2}{n_{\pi} + n_{\rho} \left(4 + 2\operatorname{Var}[\delta]\right)}\right) < 0,$$
(47)

$$\frac{\partial \zeta}{\partial v_{2,\pi}} = \frac{n_{\rho}}{n_{\pi}} (2n_{\rho} + n_{\pi})^2 \left(\frac{4 - 2\operatorname{Var}[\delta]}{n_{\pi} + n_{\rho} \left(4 + 2\operatorname{Var}[\delta]\right)} \right), \quad (48)$$

$$\frac{\partial \zeta}{\partial \operatorname{Var}[\delta]} = \frac{-2n_{\rho}}{n_{\pi}} \left(\frac{2n_{\rho} + n_{\pi}}{n_{\pi} + n_{\rho} \left(4 + 2\operatorname{Var}[\delta]\right)} \right)^{2} \times \left(v_{2,\pi}n_{\pi} + 8v_{2,\pi}n_{\rho} - 2v_{2,\rho}n_{\rho}\right).$$
(49)

When $0 \leq \operatorname{Var}[\delta] < 2$, $\partial \zeta / \partial v_{2,\pi} > 0$; When $2 < \operatorname{Var}[\delta] \leq 4$, $\partial \zeta / \partial v_{2,\pi} < 0$.

Varying the $p_{T,\rho}$ changes $\operatorname{Var}[\delta]$. As long as $v_{2,\rho}$ is no more than $9v_{2,\pi}$ (which is almost always the case), then $\partial \zeta / \partial \operatorname{Var}[\delta] < 0$. In our p_T scan, $v_{2,\rho}$ is a single value 0.06, and the average of $v_{2,\pi}$ is also around this value.

Thus, after suitable approximations, we can see the effects on the shape of R_{Ψ_2} from those variables:

- Increasing $v_{2,\rho}$ makes R_{Ψ_2} more concave.
- Increasing $\operatorname{Var}[\delta]$ makes R_{Ψ_2} more concave. Increasing $p_{T,\rho}$ makes R_{Ψ_2} more convex, because larger $p_{T,\rho}$ makes the two daughter pions closer to each other in angle, yielding a smaller $\operatorname{Var}[\delta]$ (see Fig. 7a).

• Increasing $v_{2,\pi}$ makes R_{Ψ_2} more convex when $\operatorname{Var}[\delta] < 2$, and more concave when $\operatorname{Var}[\delta] > 2$. In our default simulation (Fig. 7b), $\operatorname{Var}[\delta] \approx 1.359^2 < 2$, so R_{Ψ_2} becomes more convex as $v_{2,\pi}$ increases.

The conclusions of the CLT analysis are consistent with the simulation results.



FIG. 7. Var $[\delta]$ in the default simulation and the scan of $p_{T,\rho}$. (a) Var $[\delta]$ depending on $p_{T,\rho}$. (b) δ distribution in the default simulation.

4.3. CLT analysis for v_3

If we only focus on v_3 , the PDF in Eq. 15 could be simplified as follows:

$$f(\varphi) = \frac{1}{2\pi} \left(1 + 2v_3 \cos(3\varphi) \right). \tag{50}$$

4.3.1. Analysis for Definition A

By Definition A, we list the short-hand notations:

$$c := \cos \varphi, \quad s := \sin \varphi, \quad \xi = \pi/3,$$

$$c' := \cos \left(\varphi - \frac{\pi}{3}\right), \quad s' := \sin \left(\varphi - \frac{\pi}{3}\right), \quad (51)$$

$$\delta := 2 \sin \left(\frac{1}{2}\delta\varphi\right).$$

By using the simplified PDF (Eq. 50), we can easily get the second moments needed:

$$E[c_{\rho}^{2}] = E[\bar{c}_{\rho}^{2}] = \frac{1}{2}, \quad E[s_{\rho}^{2}] = \frac{1}{2}, \quad E[s_{\pi}^{2}] = \frac{1}{2},$$

$$E[c_{\rho}^{\prime 2}] = E[\bar{c}_{\rho}^{\prime 2}] = \frac{1}{2}, \quad E[s_{\rho}^{\prime 2}] = \frac{1}{2}, \quad E[s_{\pi}^{\prime 2}] = \frac{1}{2}.$$
(52)

There is no $v_{3,\rho}$ or $v_{3,\pi}$ in any term above, so the shapes of the observables should not change with $v_{3,\rho}$ or $v_{3,\pi}$. We can just utilize the CLT analysis results for v_2 by setting all v_2 values to 0, and then from the expression of ζ in Eq. 45, we see the terms in each bracket cancel each other. Thus, the CLT analysis shows $\zeta = 0$, and accordingly, R_{Ψ_3} should be always flat, as indeed shown in Fig. 5c.

4.3.2. Analysis for Definition B

By Definition \mathbf{B} , we list the short-hand notations:

$$c := \cos\left(\frac{3}{2}\varphi\right) \quad s := \sin\left(\frac{3}{2}\varphi\right), \quad \xi = \frac{\pi}{3},$$

$$c' := \cos\left(\frac{3}{2}\left(\varphi - \frac{\pi}{3}\right)\right), \quad s' := \sin\left(\frac{3}{2}\left(\varphi - \frac{\pi}{3}\right)\right),$$

$$\delta := 2\sin\left(\frac{3}{4}\delta\varphi\right).$$

(53)

From the simplified PDF (Eq. 50), we can get the first and the second moments:

$$E[s_{\rho}] = E[c'_{\rho}] = \frac{6 - 4v_{3,\rho}}{9\pi}, \quad E[c_{\rho}] = E[s'_{\rho}] = 0,$$

$$E[s_{\pi}] = E[c'_{\pi}] = \frac{6 - 4v_{3,\pi}}{9\pi}, \quad E[c_{\pi}] = E[s'_{\pi}] = 0,$$
(54)

$$\begin{split} \mathbf{E}[c_{\rho}^{2}] &= \mathbf{E}[\bar{c}_{\rho}^{2}] = \frac{1+v_{3,\rho}}{2}, \quad \mathbf{E}[s_{\rho}^{2}] = \frac{1-v_{3,\rho}}{2}, \\ \mathbf{E}[c_{\rho}^{\prime 2}] &= \mathbf{E}[\bar{c}_{\rho}^{\prime 2}] = \frac{1-v_{3,\rho}}{2}, \quad \mathbf{E}[s_{\rho}^{\prime 2}] = \frac{1+v_{3,\rho}}{2}, \quad (55) \\ \mathbf{E}[s_{\pi}^{2}] &= \frac{1}{2}, \quad \mathbf{E}[s_{\pi}^{\prime 2}] = \frac{1}{2}, \end{split}$$

where we have a constraint that the azimuthal range must be $0 \leq \varphi < 2\pi$. Because of the non-zero first moments, the R_{Ψ_3} curve is not flat ($\zeta \neq 0$) even if both $v_{3,\rho}$ and $v_{3,\pi}$ are set to 0. This counterintuitive observation is due to the absence of the periodical symmetry in the definition **B**. For the same reason, definition **B** has some disadvantages as follow:

- The R_{Ψ_3} curve is counterintuitively not flat, even if both $v_{3,\rho}$ and $v_{3,\pi}$ are set to 0 which means all azimuths are isotropically distributed.
- The azimuthal range must be set. In the former discussion, we let $\varphi \in [0, 2\pi)$. However, if we let the

azimuthal range be $\varphi \in [-\pi, \pi)$, the first moments will change from Eq. 54 into

$$E[s_{\rho}] = E[c'_{\rho}] = 0, \quad E[c_{\rho}] = E[s'_{\rho}] = \frac{6 - 4v_{3,\rho}}{9\pi},$$

$$E[s_{\pi}] = E[c'_{\pi}] = 0, \quad E[c_{\pi}] = E[s'_{\pi}] = \frac{6 - 4v_{3,\rho}}{9\pi},$$
(56)

which can make obvious differences to the features of the sine observales.

• The azimuthal range is by choice, however, it introduces artificial unphysical differences using Definition **B**. Take Fig. 8 as an example. If we take the azimuthal range $[-\pi, \pi)$, we have $\alpha > 0$ and $\beta < 0$. However, if we take the range $[0, 2\pi)$, β will become $\beta' = \beta + 2\pi$. The contribution of this resonance decay to ΔS_{sep} changes from

$$\sin\left(\frac{3}{2}\alpha\right) - \sin\left(\frac{3}{2}\beta\right)$$

into

$$\sin\left(\frac{3}{2}\alpha\right) - \sin\left(\frac{3}{2}\beta'\right) = \sin\left(\frac{3}{2}\alpha\right) + \sin\left(\frac{3}{2}\beta\right)$$

It seems just like the negative charge becomes a positive one.



FIG. 8. The choice of the azimuthal range affects the physical results using Definition **B**.

We thus conclude that Definition **B** is ill-devised, and should not be used. On the other hand, Definition **A** always yields a flat R_{Ψ_3} distribution and therefore is not sensitive to the CME or background. It therefore appears that the Ψ_3 harmonic plane is not suitable for the sine observables.

SUMMARY

We have presented a systematic study of resonance backgrounds as functions of the resonance v_2 and p_T with toy-model simulations and CLT calculations, in order to better understand the behaviors of the sine observable. It is found that the concavity or convexity of $R_{\Psi_2}(\Delta S)$ depends sensitively on the resonance v_2 (which yields different numbers of decay $\pi^+\pi^-$ pairs in the in-plane and out-of-plane directions) and p_T (which affects the opening angle of the decay $\pi^+\pi^-$ pair). Qualitatively, low p_T resonances decay into large opening-angle pairs and result in more "back-to-back" pairs out-of-plane (because of the positive resonance v_2), mimicking a CME signal, or a concave $R_{\Psi_2}(\Delta S)$. High p_T resonances, on the other hand, result in more close pairs in-plane, constituting a well-known background, or convex $R_{\Psi_2}(\Delta S)$. In other words, resonance backgrounds can yield both concave and convex $R_{\Psi_2}(\Delta S)$ distributions, depending on the resonance kinematics.

We have also conducted a supplemental study using the triangular flow (v_3) and discussed two definitions for the sine variables. For one of the definitions, it is found that $R_{\Psi_3}(\Delta S)$ is always flat due to the inherited symmetry in the definition. For the other definition, $R_{\Psi_3}(\Delta S)$ for v_3 is found to behave similarly as $R_{\Psi_2}(\Delta S)$ for v_2 , if the azimuthal angle is kept in the range $[0, 2\pi)$; R_{Ψ_3} can be concave or convex depending on details. However, R_{Ψ_3} is found to depend on the choice of the azimuthal angle range due to the inconsistency between the periods of R_{Ψ_3} (4 $\pi/3$) and azimuthal position (2 π). If $[-\pi,\pi)$ is chosen to be the range, then the R_{Ψ_3} results are completely different. Therefore, the Ψ_3 may not be suitable for the sine-observable studies. One has to be careful to keep the identical azimuthal angle range in the modeldata comparison studies.

We have verified our toy-model simulation results by analytical CLT calculations.

If the CME is the only source for the RP-dependent and charge-dependent correlations, then the $R_{\Psi_2}(\Delta S)$ would be concave and $R_{\Psi_3}(\Delta S)$ would be convex for the nontrivial definition. However, given the existence of backgrounds, a concave $R_{\Psi_2}(\Delta S)$ and a simultaneous convex $R_{\Psi_3}(\Delta S)$ do not lead to the conclusion of CME. This is because the $R_{\Psi_2}(\Delta S)$ and $R_{\Psi_3}(\Delta S)$ variables do not necessarily have a prior relationship, each individually varying with their respective $v_m(p_T)$ (m = 2, 3)of resonances, and because the R_{Ψ_3} variable depends on what azimuthal range is used. Based on our results, it is clear that the qualitative concavity or convexity of the $R_{\Psi_2}(\Delta S)$ or $R_{\Psi_3}(\Delta S)$ variable, or the comparison between them, cannot conclude on the existence, nor the magnitude, of the CME. Since the $R_{\Psi_2}(\Delta S)$ and $R_{\Psi_3}(\Delta S)$ variables depend on the details of the resonance kinematics and anisotropies, as well as the resonance abundances, a precise knowledge of all resonance distributions is required in order to quantify the CME using the $R_{\Psi_m}(\Delta S)$ (m = 2, 3) observables.

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- [1] T. Lee and G. Wick, Phys.Rev. **D9**, 2291 (1974).
- [2] P. D. Morley and I. A. Schmidt, Z. Phys. C26, 627 (1985).
- [3] D. Kharzeev, R. Pisarski, and M. H. Tytgat, Phys.Rev.Lett. 81, 512 (1998), arXiv:hep-ph/9804221 [hep-ph].
- [4] D. E. Kharzeev, L. D. McLerran, and H. J. Warringa, Nucl.Phys. A803, 227 (2008), arXiv:0711.0950 [hep-ph].
- [5] K. Fukushima, D. E. Kharzeev, and H. J. Warringa, Phys.Rev. **D78**, 074033 (2008), arXiv:0808.3382 [hepph].
- [6] D. Kharzeev, Phys.Lett. B633, 260 (2006), arXiv:hep-ph/0406125 [hep-ph].
- [7] A. Bzdak and V. Skokov, Phys. Lett. B710, 171 (2012), arXiv:1111.1949 [hep-ph].
- [8] W.-T. Deng and X.-G. Huang, Phys. Rev. C85, 044907 (2012), arXiv:1201.5108 [nucl-th].
- [9] J. Bloczynski, X.-G. Huang, X. Zhang, and J. Liao, Phys.Lett. B718, 1529 (2013), arXiv:1209.6594 [nucl-th].
- [10] I. Arsene *et al.* (BRAHMS Collaboration), Nucl.Phys. A757, 1 (2005), arXiv:nucl-ex/0410020 [nucl-ex].
- B. Back *et al.* (PHOBOS Collaboration), Nucl.Phys. A757, 28 (2005), arXiv:nucl-ex/0410022 [nucl-ex].
- [12] J. Adams *et al.* (STAR Collaboration), Nucl.Phys. A757, 102 (2005), arXiv:nucl-ex/0501009 [nucl-ex].
- [13] K. Adcox *et al.* (PHENIX Collaboration), Nucl.Phys. A757, 184 (2005), arXiv:nucl-ex/0410003 [nucl-ex].
- [14] B. Muller, J. Schukraft, and B. Wyslouch, Ann.Rev.Nucl.Part.Sci. 62, 361 (2012), arXiv:1202.3233 [hep-ex].
- [15] S. A. Voloshin, Phys.Rev. C70, 057901 (2004), arXiv:hep-ph/0406311 [hep-ph].
- [16] B. Abelev *et al.* (STAR Collaboration), Phys.Rev.Lett. 103, 251601 (2009), arXiv:0909.1739 [nucl-ex].
- [17] B. Abelev *et al.* (STAR Collaboration), Phys.Rev. C81, 054908 (2010), arXiv:0909.1717 [nucl-ex].
- [18] L. Adamczyk *et al.* (STAR), Phys. Rev. Lett. **113**, 052302 (2014), arXiv:1404.1433 [nucl-ex].
- [19] L. Adamczyk *et al.* (STAR), Phys. Rev. C88, 064911 (2013), arXiv:1302.3802 [nucl-ex].
- [20] B. Abelev *et al.* (ALICE), Phys.Rev.Lett. **110**, 012301 (2013), arXiv:1207.0900 [nucl-ex].
- [21] F. Wang, Phys.Rev. C81, 064902 (2010), arXiv:0911.1482 [nucl-ex].
- [22] A. Bzdak, V. Koch, and J. Liao, Phys.Rev. C81, 031901 (2010), arXiv:0912.5050 [nucl-th].
- [23] S. Schlichting and S. Pratt, Phys.Rev. C83, 014913 (2011), arXiv:1009.4283 [nucl-th].

- [24] J. Zhao, Int. J. Mod. Phys. A33, 1830010 (2018), arXiv:1805.02814 [nucl-ex].
- [25] J. Zhao, Z. Tu, and F. Wang, (2018), arXiv:1807.05083 [nucl-ex].
- [26] V. Khachatryan *et al.* (CMS), Phys. Rev. Lett. **118**, 122301 (2017), arXiv:1610.00263 [nucl-ex].
- [27] J. Zhao (STAR), Proceedings, 47th International Symposium on Multiparticle Dynamics (ISMD2017): Tlaxcala, Tlaxcala, Mexico, September 11-15, 2017, EPJ Web Conf. 172, 01005 (2018), arXiv:1712.00394 [hep-ex].
- [28] J. Zhao (STAR), Proceedings, 46th International Symposium on Multiparticle Dynamics (ISMD 2016): Jeju Island, South Korea, August 29-September 2, 2016, EPJ Web Conf. 141, 01010 (2017).
- [29] L. Adamczyk *et al.* (STAR), Phys. Rev. C89, 044908 (2014), arXiv:1303.0901 [nucl-ex].
- [30] A. M. Sirunyan *et al.* (CMS), (2017), arXiv:1708.01602 [nucl-ex].
- [31] S. Acharya *et al.* (ALICE), (2017), arXiv:1709.04723 [nucl-ex].
- [32] N. Ajitanand, R. A. Lacey, A. Taranenko, and J. Alexander, Phys.Rev. C83, 011901 (2011), arXiv:1009.5624 [nucl-ex].
- [33] N. Magdy, S. Shi, J. Liao, N. Ajitanand, and R. A. Lacey, (2017), arXiv:1710.01717v2 (version 2) [physics.data-an] arXiv:1710.01717v3 (version 3) [physics.data-an].
- [34] R. A. L. (for the STAR Collaboration) (Fudan University, Shanghai China, 2017).
- [35] P. Bozek, (2017), arXiv:1711.02563 [nucl-th].
- [36] F. Wang and J. Zhao, Phys. Rev. C95, 051901 (2017), arXiv:1608.06610 [nucl-th].
- [37] J. Adams *et al.* (STAR), Phys. Rev. Lett. **92**, 092301 (2004), arXiv:nucl-ex/0307023 [nucl-ex].
- [38] S. Adler *et al.* (PHENIX Collaboration), Phys.Rev.Lett. 91, 072301 (2003), arXiv:nucl-ex/0304022 [nucl-ex].
- [39] J. Adams *et al.* (STAR), Phys. Rev. Lett. **92**, 112301 (2004), arXiv:nucl-ex/0310004 [nucl-ex].
- [40] B. Abelev *et al.* (STAR Collaboration), Phys.Rev. C79, 034909 (2009), arXiv:0808.2041 [nucl-ex].
- [41] J. Adams et al. (STAR Collaboration), Phys.Rev. C72, 014904 (2005), arXiv:nucl-ex/0409033 [nucl-ex].
- [42] A. Adare *et al.* (PHENIX), Phys. Rev. Lett. **105**, 142301 (2010), arXiv:1006.3740 [nucl-ex].
- [43] X. Dong, S. Esumi, P. Sorensen, N. Xu, and Z. Xu, Phys. Lett. B597, 328 (2004), arXiv:nucl-th/0403030 [nucl-th].
- [44] L. Adamczyk *et al.* (STAR), Phys. Rev. C92, 024912 (2015), arXiv:1504.01317 [hep-ex].
- [45] K. A. Olive *et al.* (Particle Data Group), Chin. Phys. C38, 090001 (2014).
- [46] B. Abelev *et al.* (STAR Collaboration), Phys.Rev. C82, 034909 (2010), arXiv:0912.1871 [nucl-ex].