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# Partial restoration of spin-isospin $S U(4)$ Symmetry and the One-QRPA method in Double Beta Decay 

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#### Abstract

The one-QRPA method is used to describe simultaneously both double decay beta modes, giving special attention to the partial restoration of spin-isospin $S U(4)$ symmetry. To implement this restoration and to fix the model parameters, we resort to the energetics of Gamow-Teller resonances and to the minima of the single $\beta^{+}$-decay strengths. This makes the theory predictive regarding the $\beta \beta_{2 \nu}$-decay, producing the $2 \nu$ moments in ${ }^{48} \mathrm{Ca},{ }^{76} \mathrm{Ge},{ }^{82} \mathrm{Se},{ }^{96} \mathrm{Zr},{ }^{100} \mathrm{Mo},{ }^{128,130} \mathrm{Te}$, and ${ }^{150} \mathrm{Nd}$, that are of the same order of magnitude as the experimental ones; however, the agreement with $\beta \beta_{2 \nu}$ data is only modest. To include contributions coming from induced nuclear weak currents, we extend the $\beta \beta_{0 \nu}$-decay formalism employed previously in C. Barbero et al., Nuc. Phys. A628, 170 (1998), which is based on the Fourier-Bessel expansion. The numerical results for the $\beta \beta_{0 \nu}$ moments in the above mentioned nuclei are similar to those obtained in other theoretical studies although smaller on average by $\sim 40 \%$. We attribute this difference basically to the one-QRPA-method, employed here for the first time, instead of the currently used two-QRPA-method. The difference is partially due also to the way of carrying out the restoration of the spin-isospin symmetry. It is hard to say which is the best way to make this restoration, since the $\beta \beta_{0 \nu}$ moments are not experimentally measurable. The recipe proposed here is based on physically robust arguments. The numerical uncertainties in the $\beta \beta$ moments, related with: i) their strong dependence on the residual interaction in the particleparticle channel when evaluated within the QRPA, and ii) lack of proper knowledge of single-particle energies, have been quantified. It is concluded that the partial restoration of the $S U(4)$ symmetry, generated by the residual interaction, is crucial in the description of the $\beta \beta$-decays, regardless of the nuclear model used.


## I. INTRODUCTION

Due to the nuclear pairing force, there exists in nature about 50 "anomalous" nuclear structure systems where the odd-odd isobar, within the isobaric triplet $(N, Z)$, $(N-1, Z+1),(N-2, Z+2)$, has a higher mass than the even-even neighbors. As a consequence, the single $\beta$ decay is energetically forbidden and $\beta \beta$-decay turns out to be the only possible mode of disintegration. This is a second-order weak process whose electromagnetic analogies are the atomic Raman scattering and nuclear $\gamma \gamma-$ decay [1]. It is the slowest physical process observed so far, and can be used to learn about neutrino physics, provided we know how to deal with the nuclear structure.

The usual modes of $\beta \beta$ disintegrations are: (i) the twoneutrino double beta $\left(\beta \beta_{2 \nu}\right)$ decay, that can occur by two successive $\beta$ decays, passing through the intermediate virtual states of the $(N-1, Z+1)$ nucleus, and (ii) the neutrinoless $\beta \beta$ ( $\beta \beta_{0 \nu}$ ) decay, where there are no neutrinos in the final state. There is consensus in the scientific community that we shall not understand the $\beta \beta_{0 \nu}$-decay unless we understand the $\beta \beta_{2 \nu}$-decay. Our goal is to describe the two $\beta \beta$-decay modes consistently.

The neutrino massiveness was definitively established at the end of the 20th century through experimental observation of neutrino oscillations [2]. Nevertheless, despite this great progress, some fundamental properties are still unknown in neutrino physics, such as the Dirac or Majorana nature of neutrinos (whether they are their
own antiparticle), or the absolute neutrino mass-scale and hierarchy. The first question would be answered with the detection of the $\beta \beta_{0 \nu}$-decay. The atomic nuclei are used as the detectors of the elusive neutrinos and the next generation of experiments for many different nuclei is searching for this rare decay mode, including ${ }^{48} \mathrm{Ca},{ }^{76} \mathrm{Ge},{ }^{100} \mathrm{Mo},{ }^{116} \mathrm{Cd},{ }^{128,130} \mathrm{Te},{ }^{124,126,134} \mathrm{Xe},{ }^{136} \mathrm{Ce}$, ${ }^{150} \mathrm{Nd}$, and ${ }^{160} \mathrm{Gd}$. A summary of the experiments with the above nuclei is well explained in recent reviews, such as in Barabash [3], or Tosi [4].

A realistic quantum many-body system is generally characterized by a generic microscopic Hamiltonian, which is accessible only through approximate methods. In this regard, the mean-field theories commonly serve as an appropriate starting point but, unfortunately, they often violate the symmetries of the Hamiltonian. Such is the case for conventional BCS theory, which is an excellent zero-order approximation. However, it violates both the conservation of particle number and the spin-isospin $S U(4)$ symmetry. The first of these disadvantages does not play a very important role, but the second is crucial in the description of the $\beta \beta$-decay ${ }^{1}$. There is a general

[^0]consensus regarding this issue that: i) the $S U(4)$ symmetry is to be restored by the residual interaction, and ii) that this restoration must not be complete as this would inhibit both $\beta \beta$-decays [8]. Therefore, we speak of Partial $S U(4)$ Symmetry Restoration (PSU4SR). The question is: how to do it in a proper way? Here, we make an attempt to answer this question.

The symmetries broken by the BCS are restored by the residual interaction via the Quasiparticle RandomPhase Approximation (QRPA) and, in recent years, significant attention has been devoted to the restoration of the isospin symmetry in the evaluation of the $\beta \beta$-decay nuclear moments (NM) within this framework [9-11]. This was accomplished by separating the renormalization parameter $g_{p p}$ of the particle-particle proton-neutron interaction into isovector $g_{p p}^{T=1}$, and isoscalar $g_{p p}^{T=0}$ parts, and by choosing the first one to be essentially equal to the average pairing constant. In this way, the requirement that the Fermi $(F) \beta \beta_{2 \nu}$ matrix element $M_{F}^{2 \nu}$ vanishes is fulfilled (see Eq. (A.9)), while the corresponding vector $(V) \beta \beta_{0 \nu}$ matrix element $M_{V}^{0 \nu}$ is substantially reduced, and the full matrix element $M^{0 \nu}$, which mainly comes from the axial-vector $(A)$ moment $M_{A}^{0 \nu}$, is reduced by $\approx 10 \%$. The parameter $g_{p p}^{T=0}$ is fitted in the usual way with the requirement that the calculated values of the full $\beta \beta_{2 \nu}$ matrix elements $M^{2 \nu}$ agree with their experimental values. On the other hand, the PSU4SR has been also studied recently in the framework of schematic models [12, 13].

To implement the PSU4SR, we use a recipe based on energetics of $F$ and $G T$ resonances (in the particle-hole ( $p h$ ) channel), and on the minima of $F$ and $G T \beta^{+}$strengths (in the particle-particle ( $p p$ ) channel). Thus, the physical substratum is the same as in our previous QRPA work on the same issue [14-23, 25-28], and here we just bring up to date those studies. To implement this, we have to take into account the pseudoscalar $(P)$ and weak-magnetism $(M)$ matrix elements $M_{P}^{0 \nu}$, and $M_{M}^{0 \nu}$, as suggested by Šimkovic et al. [29] (see also[30, Appendix A]), which we have not done before, i.e., we consider now the full nuclear weak current

$$
\begin{align*}
J^{\mu \dagger}(\mathbf{x}) & =\bar{\Psi}(\mathbf{x}) \tau^{+}\left[g_{V} \gamma^{\mu}-g_{A} \gamma^{\mu} \gamma_{5}\right. \\
& \left.-i g_{M} \frac{\sigma^{\mu \nu} q_{\nu}}{2 M_{N}}-g_{P} q^{\mu} \gamma_{5}\right] \Psi(\mathbf{x}), \tag{1.1}
\end{align*}
$$

and not only the usual $V$ and $A$ terms, which we have discussed so far. We use the standard notation $[25,27$, 29, 30].

The main features of our formalism are:

[^1]1. We solve the RPA equations only once for the intermediate $(N-1, Z+1)$ nucleus [23-27], while it is usually solved twice for $(N, Z)$ and $(N-2, Z+2)$ nuclei, followed by some kind of averaging procedure. This is an outstanding difference since, as shown bellow, the one-QRPA method yields significantly smaller $\beta \beta_{0 \nu}$ moments than the currently used two-QRPA-method.
2. The residual interaction is described by the $\delta$-force (in units of $\mathrm{MeV} \cdot \mathrm{fm}^{3}$ )

$$
\begin{equation*}
V=-4 \pi\left(v^{s} P_{s}+v^{t} P_{t}\right) \delta(r) \tag{1.2}
\end{equation*}
$$

where the spin-singlet and spin-triplet parameters in the $p p$ channel, i.e., $v_{p p}^{s}$ and $v_{p p}^{t}$ correspond, respectively, to $g_{p p}^{T=1}$ and $g_{p p}^{T=0}$.
3. In essence, $v_{p p}^{s}$ is fixed in the same way as $g_{p p}^{T=1}$. Namely, we require that the vector $\beta^{+}$-strength $S_{F}^{+}$(defined in (2.31)) becomes minimal, which is achieved when the ratio $s=v_{p p}^{s} / \bar{v}_{\text {pair }}^{s}$ becomes $s_{\text {sym }}=1$, with $\bar{v}_{\text {pair }}^{s}=\left(v_{\text {pair }}^{s}(\mathrm{p})+v_{\text {pair }}^{s}(\mathrm{n})\right) / 2$. This is a strong sign that the isospin symmetry is restored within the QRPA, leading to

$$
\begin{equation*}
S_{F}^{+} \cong 0, M_{F}^{2 \nu} \cong 0, M_{V}^{0 \nu}\left(J^{\pi}=0^{+}\right) \cong 0 \tag{1.3}
\end{equation*}
$$

as well as to the concentration of the vector $\beta^{-}$ strength $S_{F}^{-}$in the Isobaric Analog State (IAS); see Fig. 1 in Ref. [21]; $M_{V}^{0 \nu}\left(J^{\pi}=0^{+}\right)$stands for the contribution of the intermediate states $J^{\pi}=0^{+}$ to the NM $M_{V}^{0 \nu}$.
4. To fix the parameter $v_{p p}^{t}$, we follow the same recipe as in the case of $v_{p p}^{s}$, i.e., we require that the $G T \beta^{+}$ strength $S_{G T}^{+}$(defined in (2.32)) becomes minimal, which indicates the PSU4SR, as was shown in Figs. 2 and 3 in Ref. [21], and Figs. 4.5 and 4.6 in Ref. [28]. For the corresponding $p p$ ratio $t=v_{p p}^{t} / \bar{v}_{\text {pair }}^{s}$, we obtain now $t_{\text {sym }} \neq 1$, with

$$
\begin{equation*}
S_{G T}^{+} \neq 0, M_{G T}^{2 \nu} \neq 0, M_{A}^{0 \nu}\left(J^{\pi}=1^{+}\right) \neq 0 \tag{1.4}
\end{equation*}
$$

and not all $G T \beta^{-}$strength $S_{G T}^{-}$is concentrated in the $G T$ resonance (GTR); $M_{A}^{0 \nu}\left(J^{\pi}=1^{+}\right)$has similar meaning to $M_{V}^{0 \nu}\left(J^{\pi}=0^{+}\right)$in (1.3).
5. The important difference with other studies is that the experimental $\beta \beta_{2 \nu}$ moments are not used for gauging the isoscalar $p p$ parameter $t$. In this way, the QRPA model turns out to be predictive regarding $M^{2 \nu}$. As a matter of fact, in Ref. [9], and in most of the QRPA calculations, the condition imposed on $g_{p p}^{T=0}$ is to reproduce the value of $\left|\mathcal{M}_{\text {exp }}^{2 \nu}\right|$, with the justification that $\mathcal{M}^{0 \nu}$ and $\mathcal{M}^{2 \nu}$ are similar. It is true that they have in common the fact of connecting the same nuclear states, and transforming two neutrons into two protons, but dynamically they are quite different: while in the $\beta \beta_{2 \nu}$-decay
two on-shell Dirac neutrinos are emitted, in the $\beta \beta_{0 \nu}$-decay an off-shell Majorana neutrino is exchanged. As a consequence, in the first case the momentum transfer is of the order of a few MeV , which makes the long wavelength approximation valid, and only the allowed ( $F$ and $G T$ ) operators need to be considered. Instead, in the second case, the momentum transfer is $\sim 100-200 \mathrm{MeV}$ and many $V$ and $A$ multipoles contribute; more still, the induced $P$ and $M$ currents, whose effects are very small in the $\beta \beta_{2 \nu}$-decay [27], also contribute quite significantly.

6 . The restoration of the isospin and $S U(4)$ symmetries, broken in the mean field approximations, are manifested not only in the $p p$ channel but also in the particle-hole $p h$ channel. In fact, we have monitored the $p h$ parameters $v_{s}^{p h}$ and $v_{t}^{p h}$ from the experimental energetics of the IAS and the GTR (in units of MeV ) [31]:

$$
\begin{equation*}
E_{G T R}-E_{I A S}=26 A^{-1 / 3}-18.5(N-Z) / A \tag{1.5}
\end{equation*}
$$

where the first term on the rhs comes from the $S U(4)$ symmetry-breaking caused by the spinorbital coupling, while the second term may be interpreted as the symmetry-restoration effect induced by the residual interaction [31-33], which displaces the $G T$ towards the IAS with increasing $N-Z$ [31-33].
In short, we can say that in our nuclear model there are no free parameters.

This article is organized as follows: in Sec. II we elaborate a formalism, based on the Fourier-Bessel expansion introduced previously [21, 22, 25-27], which allows us to evaluate in a rather simple way the pseudoscalar and weak magnetism operators, such as they appear in weak current (1.1). In Sec. III we discuss the different QRPA methods that are employed in the evaluation of the $\beta \beta$-decay NM, pointing out the advantages of using just one QRPA equation instead of two, as is often done. In Sec. IV A we explain the determination of the model parameters, both in particle-hole $(p h)$ and particle-particle ( $p p$ ) channels, which restore the $S U(4)$ symmetry and are used in the evaluation of the $\beta \beta$-decay moments. In this section, extensive numerical evaluations of the NM are presented as well, by solving only one QRPA equation. Those for the $2 \nu$-decays are confronted with the experimental data, while the predicted $0 \nu$ values are compared with some recent calculations. In Sec IV B we perform the calculations of the $0 \nu$ NM in the standard way, i.e., by solving two QRPA equations, and by adjusting the isoscalar strength to the measured $\beta \beta_{2 \nu}$ half-life. This allows us to directly compare our results with the recent QRPA calculations performed with realistic nucleon-nucleon (NN) forces, and thus discern and clarify the size of the following effects: a) two QRPA diagonalizations, b) the chosen type of NN interaction, and c)
of how to set the parameters of the nuclear Hamiltonian. Different calculations of the $\beta \beta_{0 \nu}$ NM are confronted in Sec. V and a few final remarks are made. Finally, in the Appendix the QRPA quenching mechanism in the Single-Mode Model (SMM), which is the simplest version of the $\beta \beta$-QRPA with only one intermediate state for each $J^{\pi}[16,22]$, is discussed.

## II. FORMALISM

## A. $\beta \beta_{0 \nu}$ Nuclear Moments

The $\beta \beta_{0 \nu}$ nuclear moments for the decay from the ground state $|I\rangle$ in the $(N, Z)$ nucleus to the ground state $|F\rangle$ in the $(N-2, Z+2)$ nucleus (with energies $E_{I}$ and $E_{F}$ and spin and parity $J^{\pi}=0^{+}$) can be expressed as (see, for instance, Eq. (14) in Ref. [25])

$$
\begin{equation*}
M^{0 \nu}=\frac{\mathrm{R}}{4 \pi} \sum_{N} \int d \mathbf{k} v(k ; N) \mathrm{M}^{0 \nu}(\mathbf{k} ; N) \tag{2.1}
\end{equation*}
$$

with

$$
\begin{align*}
\mathrm{M}^{0 \nu}(\mathbf{k} ; N) & \equiv\langle F| J_{\mu}^{\dagger}(-\mathbf{k})|N\rangle \\
& \times\langle N| J^{\mu \dagger}(\mathbf{k})|I\rangle \tag{2.2}
\end{align*}
$$

and

$$
\begin{equation*}
J^{\mu \dagger}(\mathbf{k})=\int d \mathbf{x} J^{\mu \dagger}(\mathbf{x}) e^{-i \mathbf{k} \cdot \mathbf{x}} \tag{2.3}
\end{equation*}
$$

is the Fourier transform of the hadronic current (1.1) in momentum space. Moreover, $\mathrm{R}=r_{0} A^{1 / 3}$, with $r_{0}=1.2$ fm is introduced to make the $0 \nu \mathrm{NM}$ dimensionless, and

$$
\begin{equation*}
v(k ; N)=\frac{2}{\pi} \frac{1}{k\left(k+\omega_{N}\right)} \tag{2.4}
\end{equation*}
$$

with

$$
\begin{equation*}
\omega_{N}=E_{N}-\frac{1}{2}\left(E_{I}+E_{F}\right) \tag{2.5}
\end{equation*}
$$

as the neutrino potential, where $k=|\mathbf{k}|$ is the modulus of the spatial part of the four transfer momentum, and the summation goes over all intermediate states $N$.

Within the impulse Non-Relativistic Approximation (NRA), and when the velocity terms are omitted, the hadronic currents read [5, 25-27, 34],

$$
\begin{equation*}
J_{N R A}^{\mu}(\mathbf{x})=(\rho(\mathbf{x}), \mathbf{j}(\mathbf{x})) \tag{2.6}
\end{equation*}
$$

where

$$
\begin{align*}
\rho(\mathbf{x}) & =g_{V} \sum_{n} \tau_{n}^{+} \delta\left(\mathbf{x}-\mathbf{r}_{n}\right), \\
\mathbf{j}(\mathbf{x}) & =\sum_{n} \tau_{n}^{+} \delta\left(\mathbf{x}-\mathbf{r}_{n}\right)\left[-g_{A} \boldsymbol{\sigma}_{n}\right. \\
& \left.+f_{M}^{\prime} \boldsymbol{\nabla} \times \boldsymbol{\sigma}_{n}-g_{P}^{\prime} \boldsymbol{\nabla} \boldsymbol{\nabla} \cdot \boldsymbol{\sigma}_{n}\right] \tag{2.7}
\end{align*}
$$

are the one-body densities and currents, with $f_{M}=g_{V}+$ $g_{M}$, and $f_{M}^{\prime}=f_{M} /\left(2 M_{N}\right), g_{P}^{\prime}=g_{P} /\left(2 M_{N}\right)$.

The intermediate-energy-dependent moments (2.2) are expressed in the form:

$$
\begin{equation*}
\mathrm{M}^{0 \nu}(\mathbf{k} ; N)=\sum_{X} \mathrm{M}_{X}^{0 \nu}(\mathbf{k} ; N) \tag{2.8}
\end{equation*}
$$

with $X=V, A, P, M$, and ${ }^{2}$

$$
\begin{align*}
\mathbf{M}_{V}^{0 \nu}(\mathbf{k} ; N)= & g_{V}^{2}\langle F| \sum_{n} \tau_{n}^{+} e^{i \mathbf{k} \cdot \mathbf{r}_{n}}|N\rangle \\
& \langle N| \sum_{m} \tau_{m}^{+} e^{-i \mathbf{k} \cdot \mathbf{r}_{m}}|I\rangle, \\
\mathbf{M}_{A}^{0 \nu}(\mathbf{k} ; N)= & -g_{A}^{2}\langle F| \sum_{n} \tau_{n}^{+} \boldsymbol{\sigma}_{n} e^{i \mathbf{k} \cdot \mathbf{r}_{n}}|N\rangle \\
\cdot & \langle N| \sum_{m} \tau_{m}^{+} \boldsymbol{\sigma}_{m} e^{-i \mathbf{k} \cdot \mathbf{r}_{m}}|I\rangle \\
\mathbf{M}_{P}^{0 \nu}(\mathbf{k} ; N)= & -g_{P}^{\prime}\left(g_{P}^{\prime} k^{2}-2 g_{A}\right) \\
& \langle F| \sum_{n} \tau_{n}^{+} \boldsymbol{\sigma}_{n} \cdot \mathbf{k} e^{i \mathbf{k} \cdot \mathbf{r}_{n}}|N\rangle \\
& \langle N| \sum_{m} \tau_{m}^{+} \boldsymbol{\sigma}_{m} \cdot \mathbf{k} e^{-i \mathbf{k} \cdot \mathbf{r}_{m}}|I\rangle \\
\mathrm{M}_{M}^{0 \nu}(\mathbf{k} ; N)= & f_{M}^{\prime 2}\langle F| \sum_{n} \tau_{n}^{+} \boldsymbol{\sigma}_{n} \times \mathbf{k} e^{i \mathbf{k} \cdot \mathbf{r}_{n}}|N\rangle \\
\cdot & \langle N| \sum_{m} \tau_{m}^{+} \boldsymbol{\sigma}_{m} \times \mathbf{k} e^{-i \mathbf{k} \cdot \mathbf{r}_{m}}|I\rangle \tag{2.9}
\end{align*}
$$

The multipole expansion of NM is performed here, using the Fourier-Bessel relationship

$$
\begin{equation*}
e^{i \mathbf{k} \cdot \mathbf{r}}=4 \pi \sum_{L} i^{L} j_{L}(k r)\left(Y_{L}(\hat{\mathbf{k}}) \cdot Y_{L}(\hat{\mathbf{r}})\right) \tag{2.10}
\end{equation*}
$$

to express them in terms of spherical tensor operators

$$
\begin{align*}
\mathrm{Y}_{J M}(k) & =\sum_{n} \tau_{n}^{+} j_{J}\left(k r_{n}\right) Y_{J M}\left(\hat{\mathbf{r}}_{n}\right)  \tag{2.11}\\
\mathrm{S}_{L J M}(k) & =\sum_{n} \tau_{n}^{+} j_{L}\left(k r_{n}\right)\left[\boldsymbol{\sigma}_{n} \otimes Y_{L}\left(\hat{\mathbf{r}}_{n}\right)\right]_{J M}
\end{align*}
$$

In this way, and after performing the angular integration on $\Omega_{k}$,

$$
\begin{equation*}
\int d \Omega_{k} \mathrm{M}_{X}^{0 \nu}(\mathbf{k}, N) \equiv \mathrm{M}_{X}^{0 \nu}(k, N) \tag{2.12}
\end{equation*}
$$

one obtains

$$
\mathrm{M}_{V}^{0 \nu}(k, N)=g_{V}^{2}(4 \pi)^{2} \sum_{J}\langle F| \mathrm{Y}_{J}(k)|N\rangle\langle N| \mathrm{Y}_{J}(k)|I\rangle
$$

[^2]\[

$$
\begin{align*}
\mathrm{M}_{A}^{0 \nu}(k, N) & =g_{A}^{2}(4 \pi)^{2} \sum_{L J}(-1)^{L+J}\langle F| \mathrm{S}_{L J}(k)|N\rangle \\
& \cdot\langle N| \mathrm{S}_{L J}(k)|I\rangle, \\
\mathrm{M}_{P}^{0 \nu}(k, N) & =g_{P}^{\prime}\left(g_{P}^{\prime} k^{2}-2 g_{A}\right)(4 \pi)^{2} \sum_{L L^{\prime} J l}(-)^{\left(L+L^{\prime}\right) / 2} \\
& \times(11 \mid l)\left(L L^{\prime} \mid l\right) \hat{L} \hat{L}^{\prime}\left\{\begin{array}{ccc}
1 & 1 & l \\
L & L^{\prime} & J
\end{array}\right\} \\
& \times\langle F| \mathrm{S}_{L J}(k)|N\rangle \cdot\langle N| \mathrm{S}_{L^{\prime} J}(k)|I\rangle \\
\mathrm{M}_{M}^{0 \nu}(k, N) & =-f_{M}^{\prime 2}\left(k^{2}\right)(4 \pi)^{2} \sum_{L L^{\prime} J l}(-)^{\left(L+L^{\prime}\right) / 2} \hat{L} \hat{L}^{\prime} \\
& \times(11 \mid l)\left(L L^{\prime} \mid l\right)\left\{\begin{array}{cc}
1 & 1 \\
L & l \\
L & J
\end{array}\right\}[2-l(l+1) / 2] \\
& \times\langle F| \mathrm{S}_{L J}(k)|N\rangle \cdot\langle N| \mathrm{S}_{L^{\prime} J}(k)|I\rangle . \tag{2.13}
\end{align*}
$$
\]

The expression (2.1) is written again in the form (2.8), i.e., as $M^{0 \nu}=\sum_{X} M_{X}^{0 \nu}$, where the moments $M_{X}^{0 \nu}$ are derived from the moments (2.13) after multiplying them by the factor $\mathrm{R} k^{2} v(k ; N) / 4 \pi$, and integrating over $k$. For instance,

$$
\begin{align*}
M_{A}^{0 \nu} & =4 \pi \mathrm{R} g_{A}^{2} \sum_{L J N}(-1)^{L+J} \int v(k, N) k^{2} d k \\
& \times\langle F| \mathrm{S}_{L J}(k)|N\rangle \cdot\langle N| \mathrm{S}_{L J}(k)|I\rangle \tag{2.14}
\end{align*}
$$

To incorporate the nuclear structure, we employ the relation (see [25, Eq. 36)])

$$
\begin{align*}
& \sum_{N}\langle F| \mathrm{T}_{J}(k)|N\rangle \cdot\langle N| \mathrm{T}_{J}(k)|I\rangle \\
= & (-)^{J} \sum_{\alpha \pi p n p^{\prime} n^{\prime}} \rho^{p h}\left(p n p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right) \\
\times & \left\langle p\left\|\mathrm{~T}_{J}(k)\right\| n\right\rangle\left\langle p^{\prime}\left\|\mathrm{T}_{J}(k)\right\| n^{\prime}\right\rangle, \tag{2.15}
\end{align*}
$$

where

$$
\begin{equation*}
\rho^{p h}\left(p n p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right)=\rho^{-}\left(p n ; J_{\alpha}^{\pi}\right) \rho^{+}\left(p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right), \tag{2.16}
\end{equation*}
$$

and

$$
\begin{align*}
\rho^{-}\left(p n J_{\alpha}^{\pi}\right) & =\hat{J}^{-1}\left\langle 0_{F}^{+}\left\|\left(a_{p}^{\dagger} a_{\bar{n}}\right)_{J^{\pi}}\right\| J_{\alpha}^{\pi}\right\rangle \\
\rho^{+}\left(p n J_{\alpha}^{\pi}\right) & =\hat{J}^{-1}\left\langle J_{\alpha}^{\pi}\left\|\left(a_{p}^{\dagger} a_{\bar{n}}\right)_{J^{\pi}}\right\| 0_{I}^{+}\right\rangle \tag{2.17}
\end{align*}
$$

are the $\beta^{\mp}$ one-body state dependent $p h$ density matrices, the index $\alpha$ labels different intermediate states with the same spin $J$ and parity $\pi$, and $\hat{J} \equiv \sqrt{2 J+1}$. For convenience, we made the substitution

$$
\begin{equation*}
|I\rangle,|F\rangle,|N\rangle \rightarrow\left|0_{I}^{+}\right\rangle,\left|0_{F}^{+}\right\rangle,\left|J_{\alpha}^{\pi}\right\rangle . \tag{2.18}
\end{equation*}
$$

For example, Eq. (2.14) reads now

$$
\begin{align*}
M_{A}^{0 \nu} & =4 \pi \mathrm{R} g_{A}^{2} \sum_{L J}(-1)^{1+L} \sum_{\alpha \pi p n p^{\prime} n^{\prime}} \rho^{p h}\left(p n p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right) \\
& \times \int v\left(k, \omega_{J_{\alpha}^{\pi}}\right) k^{2} d k\left\langle p\left\|\mathrm{~S}_{L J}(k)\right\| n\right\rangle\left\langle p^{\prime}\left\|\mathrm{S}_{L J}(k)\right\| n^{\prime}\right\rangle \tag{2.19}
\end{align*}
$$

Thus, the final results for the $\beta \beta_{0 \nu}$ NM are:

$$
\begin{align*}
M_{V}^{0 \nu} & =\sum_{J_{\alpha}^{\pi}}(-)^{J} \sum_{p p^{\prime} n n^{\prime}} \rho^{p h}\left(p n p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right) \\
& \times W_{J 0 J}(p n) W_{J 0 J}\left(p^{\prime} n^{\prime}\right) \mathcal{R}_{J J}^{V}\left(p n p^{\prime} n^{\prime} ; \omega_{J_{\alpha}^{\pi}}\right), \\
M_{A}^{0 \nu} & =\sum_{L J_{\alpha}^{\pi}}(-)^{L+1} \sum_{p p^{\prime} n n^{\prime}} \rho^{p h}\left(p n p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right) \\
& \times W_{L 1 J}(p n) W_{L 1 J}\left(p^{\prime} n^{\prime}\right) \mathcal{R}_{L L}^{A}\left(p n p^{\prime} n^{\prime} ; \omega_{J_{\alpha}^{\pi}}\right), \\
M_{P}^{0 \nu} & =-\sum_{J_{\alpha}^{\pi} L L^{\prime} l}(-)^{J+\left(L+L^{\prime}\right) / 2} \hat{L} \hat{L}^{\prime}\left(L L^{\prime} \mid l\right)(11 \mid l) \\
& \times \sum_{p p^{\prime} n n^{\prime}} \rho^{p h}\left(p n p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right) W_{L 1 J}(p n) W_{L^{\prime} 1 J}\left(p^{\prime} n^{\prime}\right) \\
& \times\left\{\begin{array}{ll}
L & L^{\prime} \\
1 & l \\
1
\end{array}\right\} \mathcal{R}_{L L^{\prime}}^{P}\left(p n p^{\prime} n^{\prime} ; \omega_{J_{\alpha}^{\pi}}\right), \\
M_{M}^{0 \nu} & =-\sum_{J_{\alpha}^{\pi} L L^{\prime} l}(-)^{J+\left(L+L^{\prime}\right) / 2 \hat{L} \hat{L}^{\prime}\left(L L^{\prime} \mid l\right)(11 \mid l)} \\
& \times \sum_{p p^{\prime} n n^{\prime}} \rho^{p h}\left(p n p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right) W_{L 1 J}(p n) W_{L^{\prime} 1 J}\left(p^{\prime} n^{\prime}\right) \\
& \times\left\{\begin{array}{ll}
L & L^{\prime} \\
1 & l \\
1 & J
\end{array}\right\}[2-l(l+1) / 2] \mathcal{R}_{L L^{\prime}}^{M}\left(p n p^{\prime} n^{\prime} ; \omega_{J_{\alpha}^{\pi}}\right), \tag{2.20}
\end{align*}
$$

with the angular parts: ${ }^{3}$

$$
W_{L S J}(p n)=\sqrt{2} \hat{S} \hat{J} \hat{L} \hat{l}_{n} \hat{j}_{n} \hat{j}_{p}\left(l_{n} L \mid l_{p}\right)\left\{\begin{array}{ccc}
l_{p} & \frac{1}{2} & j_{p}  \tag{2.21}\\
L & S & J \\
l_{n} & \frac{1}{2} & j_{n}
\end{array}\right\}
$$

while the two-body radial integrals are defined as

$$
\begin{align*}
\mathcal{R}_{L L^{\prime}}^{X}\left(p n p^{\prime} n^{\prime} ; \omega_{J_{\alpha}^{\pi}}\right) & =\mathrm{R} \int d k k^{2} v_{X}\left(k ; \omega_{J_{\alpha}^{\pi}}\right) \\
& \times R_{L}(p n ; k) R_{L^{\prime}}\left(p^{\prime} n^{\prime} ; k\right), \tag{2.22}
\end{align*}
$$

with

$$
\begin{equation*}
R_{L}(p n ; k)=\int_{0}^{\infty} u_{n_{p} l_{p}}(r) u_{n_{n} l_{n}}(r) j_{L}(k r) r^{2} d r \tag{2.23}
\end{equation*}
$$

being one-body radial integrals, and $u_{n l}$ the radial singleparticle functions. Finally, the effective neutrino potentials in $v_{X}\left(k ; \omega_{J_{\alpha}^{\pi}}\right)(2.22)$ become

$$
\begin{align*}
v_{V}\left(k ; \omega_{J_{\alpha}^{\pi}}\right) & =g_{V}^{2}\left(k^{2}\right) v\left(k ; \omega_{J_{\alpha}^{\pi}}\right), \\
v_{A}\left(k ; \omega_{J_{\alpha}^{\pi}}\right) & =g_{A}^{2}\left(k^{2}\right) v\left(k ; \omega_{J_{\alpha}^{\pi}}\right),  \tag{2.24}\\
v_{M}\left(k ; \omega_{J_{\alpha}^{\pi}}\right) & =k^{2} f_{M}^{\prime 2}\left(k^{2}\right) v\left(k ; \omega_{J_{\alpha}^{\pi}}\right), \\
v_{P}\left(k ; \omega_{J_{\alpha}^{\pi}}\right) & =k^{2} g_{P}^{\prime}\left(k^{2}\right)\left[2 g_{A}\left(k^{2}\right)-k^{2} g_{P}^{\prime}\left(k^{2}\right)\right] v\left(k ; \omega_{J_{\alpha}^{\pi}}\right),
\end{align*}
$$

[^3]where, within the new notation put forward in (2.18), Eqs. (2.4) and (2.5) are now expressed as
\[

$$
\begin{equation*}
v\left(k ; \omega_{J_{\alpha}^{\pi}}\right)=\frac{2}{\pi} \frac{1}{k\left(k+\omega_{J_{\alpha}^{\pi}}\right)} \tag{2.25}
\end{equation*}
$$

\]

with

$$
\begin{align*}
\omega_{J_{\alpha}^{\pi}} & =E_{J_{\alpha}^{+}}-\frac{E_{0+{ }_{I}}+E_{0_{F}^{+}}}{2} \\
& =E_{J_{\alpha}^{+}}-E_{0_{I}^{+}}+\frac{1}{2} Q_{\beta \beta} \tag{2.26}
\end{align*}
$$

where $Q_{\beta \beta}=E_{0_{I}^{+}}-E_{0_{F}^{+}}$is the Q -value of the $\beta \beta$-decay.
Finite Nucleon Size (FNS) effects are introduced through the usual dipole form factors

$$
\begin{align*}
g_{V} & \rightarrow g_{V}\left(k^{2}\right) \equiv g_{V}\left(\frac{\Lambda_{V}^{2}}{\Lambda_{V}^{2}+k^{2}}\right)^{2} \\
g_{A} & \rightarrow g_{A}\left(k^{2}\right) \equiv g_{A}\left(\frac{\Lambda_{A}^{2}}{\Lambda_{A}^{2}+k^{2}}\right)^{2} \\
f_{M}^{\prime} & \rightarrow f_{M}^{\prime}\left(k^{2}\right) \equiv f_{M}^{\prime}\left(\frac{\Lambda_{V}^{2}}{\Lambda_{V}^{2}+k^{2}}\right)^{2} \\
g_{P}^{\prime} & \rightarrow g_{P}^{\prime}\left(k^{2}\right) \equiv g_{P}^{\prime}\left(\frac{\Lambda_{A}^{2}}{\Lambda_{A}^{2}+k^{2}}\right)^{2} \tag{2.27}
\end{align*}
$$

as in Refs. $[18,29,35]$, and $\Lambda_{V}=0.85 \mathrm{GeV}$ and $\Lambda_{A}=$ 1.086 GeV are the vectorial and axial-vectorial cut-off parameters, respectively.

The weak coupling constants in (1.1) are fixed as follows: $g_{V}=1$ and $g_{M}=3.7$ from Conservation of Vector Current (CVC), $g_{A}=1.27$ from the experimental data [36], and $g_{P}=2 M_{N} g_{A} /\left(q^{2}+m_{\pi}^{2}\right)$ from the assumption of Partially Conserved Axial Current (PCAC) [34].

The Short Range Correlations (SRC) between the two nucleons are taken into account in the standard way via the correlation function [37, 38]

$$
\begin{equation*}
f^{S R C}(r)=1-j_{0}\left(k_{c} r\right) \tag{2.28}
\end{equation*}
$$

where $k_{c}=3.93 \mathrm{fm}^{-1}$ is the Compton wavelength of the $\omega$-meson. This leads to the following modification of the potentials $v_{X}\left(k ; \omega_{J_{\alpha}^{\pi}}\right)$ in the momentum space (see Eqs. (2.14) and (2.15) in Ref. [18], as well as Refs. [21, 2528, 39]. )

$$
\begin{equation*}
v_{X}\left(q ; \omega_{J}\right) \rightarrow v_{X}\left(q ; \omega_{J}\right)-\Delta v_{X}\left(q ; \omega_{J}\right) \tag{2.29}
\end{equation*}
$$

with

$$
\begin{align*}
\Delta v_{X}\left(q ; \omega_{J}\right) & =\frac{1}{2 q_{c}^{2}} \int_{-1}^{1} d x \int d k k^{2} v_{X}\left(k ; \omega_{J}\right) \\
& \times \delta\left(\sqrt{q^{2}+k^{2}+2 x q k}-q_{c}\right) \tag{2.30}
\end{align*}
$$

It is not difficult to show that

$$
\begin{align*}
\Delta v_{X}\left(q ; \omega_{J}\right) & =\frac{1}{2} \int_{-1}^{1} d x \\
& \times v_{X}\left(\sqrt{q^{2}+q_{c}^{2}+2 x q q_{c}} ; \omega_{J}\right) \tag{2.31}
\end{align*}
$$

and this is the expression used to evaluate the SRC.

## B. $\beta \beta_{2 \nu}$ Matrix Element and Charge-Exchange Transition Strengths

Independently of the nuclear model used and only considering the allowed transitions, the $\beta \beta_{2 \nu}$ moment reads

$$
\begin{equation*}
M^{2 \nu}=M_{F}^{2 \nu}+M_{G T}^{2 \nu} \tag{2.32}
\end{equation*}
$$

with [21]

$$
\begin{align*}
& M_{F}^{2 \nu}=g_{V}^{2} \sum_{p n p^{\prime} n^{\prime} \alpha} \rho^{p h}\left(p n p^{\prime} n^{\prime} ; 0_{\alpha}^{\pi}\right) \frac{W_{000}(p n) W_{000}\left(p^{\prime} n^{\prime}\right)}{\omega_{0_{\alpha}^{+}}} \\
& M_{G T}^{2 \nu}=-g_{A}^{2} \sum_{p n p^{\prime} n^{\prime} \alpha} \rho^{p h}\left(p n p^{\prime} n^{\prime} ; 1_{\alpha}^{\pi}\right) \frac{W_{011}(p n) W_{011}\left(p^{\prime} n^{\prime}\right)}{\omega_{1_{\alpha}^{+}}} \tag{2.33}
\end{align*}
$$

The single charge-exchange $\beta^{\mp}$ strengths are also discussed here. They are

$$
\begin{equation*}
S_{F}^{\mp}=\sum_{p n \alpha}\left|\rho^{\mp}\left(p n 0_{\alpha}^{+}\right) W_{000}(p n)\right|^{2} \tag{2.34}
\end{equation*}
$$

and

$$
\begin{equation*}
S_{G T}^{\mp}=\sum_{p n \alpha}\left|\rho^{\mp}\left(p n 1_{\alpha}^{+}\right) W_{011}(p n)\right|^{2} \tag{2.35}
\end{equation*}
$$

## III. CHARGE-EXCHANGE QRPA

All of the formalism presented in the previous section is valid in general, and any nuclear model can be used to evaluate the one-body density matrices (2.17). The most frequently used model is charge-exchange QRPA. It was formulated, and applied to allowed $\beta^{ \pm}$-decays and collective $G T$ resonance, by Halbleib and Sorensen (HS) in 1967 [40], as follows:

1) BCS equations for the initial even-even nucleus $(N, Z)$ are solved to obtain the occupation coefficients $\left(v_{n}, v_{p}\right),\left(u_{n}=\sqrt{1-v_{n}^{2}}, u_{p}=\sqrt{1-v_{p}^{2}}\right)$, the quasiparticle energies $\left(\epsilon_{n}, \epsilon_{p}\right)$ and the chemical potentials $\left(\lambda_{n}, \lambda_{p}\right)$ for neutrons and protons, as well as the ground state energy $E_{0_{i}}$, and the BCS vacuum

$$
\begin{equation*}
\left|0_{I}\right\rangle=\prod_{p}\left(u_{p}+v_{p} c_{p}^{\dagger} c_{\bar{p}}^{\dagger}\right) \prod_{n}\left(u_{n}+v_{n} c_{n}^{\dagger} c_{\bar{n}}^{\dagger}\right)| \rangle \tag{3.1}
\end{equation*}
$$

where $\rangle$ stands for the particle vacuum. The $u$ 's and $v$ 's in the parent nucleus are determined under the constraints

$$
\begin{equation*}
\sum_{j_{p}}\left(2 j_{p}+1\right) v_{p}^{2}=Z, \quad \sum_{n}\left(2 j_{n}+1\right) v_{n}^{2}=N \tag{3.2}
\end{equation*}
$$

where $Z$ and $N$ are the number of protons and neutrons, respectively, in the parent nucleus.
2) Transition $\beta^{\mp}$-densities

$$
\begin{align*}
\rho^{-}\left(p n J_{\alpha}^{\pi}\right) & =u_{p} v_{n} X_{J_{\alpha}^{\pi}}(p n)+u_{n} v_{p} Y_{J_{\alpha}^{\pi}}(p n), \\
\rho^{+}\left(p n J_{\alpha}^{\pi}\right) & =u_{n} v_{p} X_{J_{\alpha}^{\pi}}(p n)+u_{p} v_{n} Y_{J_{\alpha}^{\pi}}(p n) \tag{3.3}
\end{align*}
$$

and excitation energies

$$
\begin{equation*}
E_{J_{\alpha}^{\pi}}^{N \pm 1, Z \mp 1}=E_{0_{I}^{+}}+\Omega_{J_{\alpha}^{\pi}} \pm \lambda_{n} \mp \lambda_{p}, \tag{3.4}
\end{equation*}
$$

in neighboring odd-odd, $(N \pm 1, Z \mp 1)$, nuclei (see, for instance, [41, Sec. 6.3.4]), are obtained by solving the pn-QRPA equation

$$
\left(\begin{array}{ll}
A_{J^{\pi}} & B_{J^{\pi}}  \tag{3.5}\\
B_{J^{\pi}} & A_{J^{\pi}}
\end{array}\right)\binom{X_{J_{\alpha}^{\pi}}}{Y_{J_{\alpha}^{\pi}}}=\Omega_{J_{\alpha}^{\pi}}\binom{X_{J_{\alpha}^{\pi}}}{-Y_{J_{\alpha}^{\pi}}}
$$

for forward and backward going amplitudes, $X_{J_{\alpha}^{\pi}}(p n)$ and $Y_{J_{\alpha}^{\pi}}(p n)$, and QRPA excitation energies $\Omega_{J_{\alpha}^{\pi}}$ on vacuum (3.1). Both $F$ and $G T$ strengths, given by (2.34) and (2.35), fulfill the well known Ikeda sum rule

$$
\begin{equation*}
S^{\beta}=S^{-}-S^{+}=N-Z \tag{3.6}
\end{equation*}
$$

It is important to mention that the ground state correlations (GSC) for the charge-changes decay $(N, Z) \xrightarrow{\beta^{-}}(N-1, Z+1)$ is the decay $(N, Z) \xrightarrow{\beta^{+}}(N+$ $1, Z-1)$, and viceversa. In effect, the exchange $\rho^{-}\left(p n J_{\alpha}^{\pi}\right) \leftrightarrow \rho^{+}\left(p n J_{\alpha}^{\pi}\right)$ in (3.3) is obtained from the exchange $X_{J_{\alpha}^{\pi}}(p n) \leftrightarrow Y_{J_{\alpha}^{\pi}}(p n)$.

When the QRPA is applied to $\beta \beta$-decay, one has to deal simultaneously with two ground states $E_{0_{I}}$ and $E_{0_{F}}$, which requires further steps in modeling the theory in order to end up with some sort of averaging. This is inevitable, even in the case of particle number projected QRPA [19].

## A. Method I

Intensive implementations of QRPA to $\beta \beta$-decay began only about 20 years after the HS work [40], when Vogel and Zirnbauer [42] discovered that the GSC play an essential role in suppressing the $\beta \beta_{2 \nu}$ rates. Their QRPA calculations of $M_{G T}^{2 \nu}$ are carried out for both the initial and final nuclei and the resulting matrix elements are averaged. That is, they repeat the steps 1) and 2) for the $(N, Z)$ and $(N-2, Z+2)$ ground states, and for intermediate states $1_{\alpha}^{+}$and $\overline{1}_{\alpha}^{+}$in the nucleus $(N-1, Z+1)$. In the second case the BCS vacuum is

$$
\begin{equation*}
\left|0_{F}\right\rangle=\prod_{p}\left(\bar{u}_{p}+\bar{v}_{p} c_{p}^{\dagger} c_{\bar{p}}^{\dagger}\right) \prod_{n}\left(\bar{u}_{n}+\bar{v}_{n} c_{n}^{\dagger} c_{\bar{n}}^{\dagger}\right)| \rangle \tag{3.7}
\end{equation*}
$$

derived under the constraints

$$
\begin{equation*}
\sum_{p}\left(2 j_{p}+1\right) \bar{v}_{p}^{2}=Z+2, \quad \sum_{n}\left(2 j_{n}+1\right) \bar{v}_{n}^{2}=N-2 \tag{3.8}
\end{equation*}
$$

which fulfill the sum rule

$$
\begin{equation*}
\bar{S}^{\beta}=N-Z-4 . \tag{3.9}
\end{equation*}
$$

The energy denominator $\omega_{1_{\alpha}^{+}}$in (2.33) can be evaluated from experimental data or ${ }^{\alpha}$ self-consistently within the BCS-QRPA framework. When the latter is performed, one finds that [41, Sec. 6.3.4]

$$
\begin{equation*}
Q_{\beta \beta}=E_{0_{I}^{+}}-E_{0_{F}^{+}}=2\left(\lambda_{n}-\lambda_{p}\right) \tag{3.10}
\end{equation*}
$$

and, therefore, from (3.5)

$$
\begin{equation*}
E_{1_{\alpha}^{+}}-E_{0_{I}^{+}}=\Omega_{1_{\alpha}^{+}}-\frac{1}{2} Q_{\beta \beta} \tag{3.11}
\end{equation*}
$$

which from (2.26) yields

$$
\begin{equation*}
\omega_{1_{\alpha}^{+}}=\Omega_{1_{\alpha}^{+}} \tag{3.12}
\end{equation*}
$$

Proceeding in the same way for the final state $0_{F}$, one finds that the averaged $G T$ moment is:

$$
\begin{align*}
M_{G T}^{2 \nu} & =-\frac{g_{A}^{2}}{2} \sum_{p n p^{\prime} n^{\prime} \alpha} W_{011}(p n) W_{011}\left(p^{\prime} n^{\prime}\right) \\
& \times\left[\frac{\rho^{p h}\left(p n p^{\prime} n^{\prime} ; 1_{\alpha}^{+}\right)}{\Omega_{1_{\alpha}^{+}}}+\frac{\rho^{p h}\left(p n p^{\prime} n^{\prime} ; \overline{1}_{\alpha}^{+}\right)}{\Omega_{\overline{1}_{\alpha}^{+}}}\right] \tag{3.13}
\end{align*}
$$

## B. Method II

Shortly after the discovery of the importance of the GSC in Ref. [42], Civitarese, Faessler and Tomoda [43] made their calculations, repeating the steps 1) and 2) for the ground states of $(N, Z)$ and $(N-2, Z+2)$ nuclei, and for intermediate states $1_{\alpha}^{+}$and $1_{\alpha^{\prime}}^{+}$in the $(N-1, Z+1)$ nucleus, and arrived at the same conclusion about the importance of the GSC (see also Ref. [44]). In our notation, their $\beta \beta_{2 \nu}$ moment reads

$$
\begin{align*}
M_{G T}^{2 \nu} & =-g_{A}^{2} \sum_{p n p^{\prime} n^{\prime} \alpha \alpha^{\prime}} W_{011}(p n) W_{011}\left(p^{\prime} n^{\prime}\right) \\
& \times \frac{\rho^{p h}\left(p n ; 1_{\alpha}^{+}\right)\left\langle 1_{\alpha}^{+} \mid \overline{1}_{\alpha^{\prime}}^{+}\right\rangle \rho^{p h}\left(p^{\prime} n^{\prime} ; \overline{1}_{\alpha^{\prime}}^{+}\right)}{m_{e} c^{2}+\frac{1}{2} Q_{\beta \beta}+E_{1_{\alpha}^{+}}-E_{0_{i}}} \tag{3.14}
\end{align*}
$$

where the overlap is given by

$$
\begin{equation*}
\left\langle 1_{\alpha}^{+} \mid \overline{1}_{\alpha^{\prime}}^{+}\right\rangle=\sum_{p n}\left[X_{1_{\alpha}^{+}}(p n) X_{\overline{1}_{\alpha^{\prime}}^{+}}(p n)-Y_{1_{\alpha}^{+}}(p n) Y_{\overline{1}_{\alpha^{\prime}}^{+}}(p n)\right] \tag{3.15}
\end{equation*}
$$

This overlap is introduced since the intermediate states $\left|1_{\alpha}^{+}\right\rangle$and $\left|\overline{1}_{\alpha^{\prime}}^{+}\right\rangle$, being generated from different ground states, are not orthogonal to each other. When this nonorthogonality is very pronounced the numerical results
could be eventually unreliable. Making use of (3.11) the energy denominators in (3.14) become $m_{e} c^{2}+\Omega_{1_{\alpha}^{+}}$.
In more recent applications of Method II [45, 46, 72] this denominator was replaced by $\left(\Omega_{1_{\alpha}^{+}}+\Omega_{\overline{1}_{\alpha^{\prime}}}\right) / 2$. Moreover, the BCS overlap factor $\left\langle 0_{I}^{+} \mid 0_{F}^{+}\right\rangle$, which is about 0.8 , has been incorporated in these last studies. Thus, the $\beta \beta_{2 \nu}$ moment reads

$$
\begin{align*}
M_{G T}^{2 \nu} & =-2 g_{A}^{2}\left\langle 0_{I}^{+} \mid 0_{F}^{+}\right\rangle \sum_{p n p^{\prime} n^{\prime} \alpha \alpha^{\prime}} W_{011}(p n) W_{011}\left(p^{\prime} n^{\prime}\right) \\
& \times \frac{\rho^{+}\left(p^{\prime} n^{\prime} ; \overline{1}_{\alpha^{\prime}}^{+}\right)\left\langle\overline{1}_{\alpha^{\prime}}^{+} \mid 1_{\alpha}^{+}\right\rangle \rho^{-}\left(p n ; 1_{\alpha}^{+}\right)}{\Omega_{1_{\alpha}^{+}}+\Omega_{\overline{1}_{\alpha^{\prime}}^{+}}} \tag{3.16}
\end{align*}
$$

where

$$
\begin{equation*}
\left\langle 0_{I}^{+} \mid 0_{F}^{+}\right\rangle=\prod_{p}\left(u_{p} \bar{u}_{p}+v_{p} \bar{v}_{p}\right) \prod_{n}\left(u_{n} \bar{u}_{n}+v_{n} \bar{v}_{n}\right) \tag{3.17}
\end{equation*}
$$

The following substitutions have to be made in the evaluation of the $\beta \beta_{0 \nu}$ moments:

$$
\begin{align*}
\sum_{J_{\alpha}^{\pi}} & \rho^{p h}\left(p n p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right) \rightarrow\left\langle 0_{I}^{+} \mid 0_{F}^{+}\right\rangle \times \\
& \sum_{J^{\pi} \alpha \alpha^{\prime}} \rho^{+}\left(p^{\prime} n^{\prime} ; \bar{J}_{\alpha^{\prime}}^{+}\right)\left\langle\bar{J}_{\alpha^{\prime}}^{+} \mid J_{\alpha}^{+}\right\rangle \rho^{-}\left(p n ; J_{\alpha}^{+}\right), \tag{3.18}
\end{align*}
$$

in (2.20), together with

$$
\begin{equation*}
\omega_{J_{\alpha}^{\pi}} \rightarrow\left(\Omega_{J_{\alpha}^{+}}+\Omega_{\bar{J}_{\alpha^{\prime}}^{+}}\right) / 2 \tag{3.19}
\end{equation*}
$$

in (2.25).

## C. Method III

The above equations (3.13) and (3.14) for $M_{G T}^{2 \nu}$ cannot be derived mathematically, but they are physically sound ansatz for the HS equations (3.1)-(3.6) originally designed for the single $\beta$-decay, to make possible the calculations of $\beta \beta$-decay. In view of this, a new pn-QRPA, specially tailored for the last processes, was derived analytically in Ref. [16]. It is based on appropriate canonical quasi-particle transformations for which the GSC for the $\beta^{\mp}$ transitions in the intermediate $(N-1, Z+1)$ nucleus are the $\beta^{ \pm}$transitions in the same nucleus. ${ }^{4}$ Only one QRPA equation

$$
\left(\begin{array}{cc}
\tilde{A}_{J^{\pi}} & \tilde{B}_{J^{\pi}}  \tag{3.20}\\
\tilde{B}_{J^{\pi}} & \tilde{A}_{J^{\pi}}
\end{array}\right)\binom{\tilde{X}_{J_{\alpha}^{\pi}}}{\tilde{Y}_{J_{\alpha}^{\pi}}}=\tilde{\Omega}_{J_{\alpha}^{\pi}}\binom{\tilde{X}_{J_{\alpha}^{\pi}}}{-\tilde{Y}_{J_{\alpha}^{\pi}}},
$$

[^4]is solved for the vacuum
\[

$$
\begin{equation*}
|\tilde{0}\rangle=\prod_{p}\left(u_{p}+\bar{v}_{p} c_{p}^{\dagger} c_{\bar{p}}^{\dagger}\right) \prod_{n}\left(\bar{u}_{n}+v_{n} c_{n}^{\dagger} c_{\bar{n}}^{\dagger}\right)| \rangle \tag{3.21}
\end{equation*}
$$

\]

which contains information on both initial and final ground states. Unbarred and barred quantities are derived as before, i.e., by solving the BCS equations for initial and final nuclei with constraints (3.2) and (3.8), respectively. The matrices $\tilde{A}_{J^{\pi}}$ and $\tilde{B}_{J^{\pi}}$ are given by [16, Eq. (3)], and the $G T$ moment is:

$$
\begin{align*}
M_{G T}^{2 \nu} & =-g_{A}^{2} \sum_{p n p^{\prime} n^{\prime} \alpha} W_{011}(p n) W_{011}\left(p^{\prime} n^{\prime}\right) \\
& \times \frac{\tilde{\rho}^{-}\left(p n 1_{\alpha}^{+}\right) \tilde{\rho}^{+}\left(p^{\prime} n^{\prime} 1_{\alpha}^{+}\right)}{\tilde{\Omega}_{1_{\alpha}^{+}}} \tag{3.22}
\end{align*}
$$

where

$$
\begin{align*}
& \tilde{\rho}^{-}\left(p n J_{\alpha}^{\pi}\right)=\sqrt{\sigma_{p} \sigma_{n}}\left(u_{p} v_{n} \tilde{X}_{J_{\alpha}^{\pi}}(p n)+\bar{u}_{n} \bar{v}_{p} \tilde{Y}_{J_{\alpha}^{\pi}}(p n)\right), \\
& \tilde{\rho}^{+}\left(p n J_{\alpha}^{\pi}\right)=\sqrt{\sigma_{p} \sigma_{n}}\left(\bar{v}_{p} \bar{u}_{n} \tilde{X}_{J_{\alpha}^{\pi}}(p n)+u_{p} v_{n} \tilde{Y}_{J_{\alpha}^{\pi}}(p n)\right), \tag{3.23}
\end{align*}
$$

and

$$
\begin{equation*}
\sigma_{p}^{-1}=u_{p}^{2}+\bar{v}_{p}^{2}, \quad \sigma_{n}^{-1}=u_{n}^{2}+\bar{v}_{n}^{2} \tag{3.24}
\end{equation*}
$$

The $G T$ strengths

$$
\begin{equation*}
\tilde{S}_{G T}^{\mp}=\sum_{p n \alpha}\left|\tilde{\rho}^{\mp}\left(p n 1_{\alpha}^{+}\right) W_{011}(p n)\right|^{2} \tag{3.25}
\end{equation*}
$$

fulfill now the sum rule

$$
\begin{equation*}
\tilde{S}^{\beta}=\tilde{S}^{-}-\tilde{S}^{+} \cong N-Z-2 \tag{3.26}
\end{equation*}
$$

Note that here the averaging is no longer carried out at the level of the QRPA but within the BCS approximations.

In addition to being mathematically and physically justified, Method III has several advantages over Methods I and II, namely: i) only one QRPA equation is solved instead of two, ii) it is not necessary to deal with troublesome overlaps (3.15), iii) $M_{G T}^{2 \nu}$ can be evaluated without diagonalizing the QRPA matrix (3.16); it is enough to invert this matrix [16], and iv) it allows us to formulate the SMM, which illustrates several aspects of the PSU4SR and the role played by the GSC [18], as shown in the Appendix.

## D. Method IV

In his studies of single $\beta$-decay, Cha [48] has argued that "because the intersection between two-qp's takes place in a residual nucleus, we should calculate $\epsilon$ 's, $u$ 's, and $v$ 's in the daughter nucleus." Motivated by this argument, and in order to make the QRPA calculation as simple as possible without losing the physical content of
the model, a further step was taken in Ref. [23] in the evaluation of the moments $M_{G T}^{2 \nu}$. There, instead of dealing with the two-vacua QRPA, as done in Ref. [16], BCS equations are solved only for the intermediate nucleus, obtaining the vacuum

$$
\begin{equation*}
\left|0_{i n t}^{\prime}\right\rangle=\prod_{p}\left(u_{p}^{\prime}+v_{p}^{\prime} c_{p}^{\dagger} c_{\bar{p}}^{\dagger}\right) \prod_{n}\left(u_{n}^{\prime}+v_{n}^{\prime} c_{n}^{\dagger} c_{\bar{n}}^{\dagger}\right)| \rangle \tag{3.27}
\end{equation*}
$$

where the $u^{\prime \prime}$ s and $v^{\prime}$ 's are determined under the constraints

$$
\begin{equation*}
\sum_{p}\left(2 j_{p}+1\right) v_{p}^{\prime 2}=Z+1, \quad \sum_{n}\left(2 j_{n}+1\right) v_{n}^{\prime 2}=N-1 \tag{3.28}
\end{equation*}
$$

satisfying the sum rule

$$
\begin{equation*}
S^{\beta^{-}}=N-Z-2 \tag{3.29}
\end{equation*}
$$

for $\beta^{-}$-decay, which is very similar to (3.26).
In a manner similar to (3.22), the $G T$ moment is

$$
\begin{align*}
M_{G T}^{2 \nu} & =-g_{A}^{2} \sum_{p n p^{\prime} n^{\prime} \alpha} W_{011}(p n) W_{011}\left(p^{\prime} n^{\prime}\right) \\
& \times \frac{\rho^{\prime-}\left(p n 1_{\alpha}^{+}\right) \rho^{\prime+}\left(p^{\prime} n^{\prime} 1_{\alpha}^{+}\right)}{\Omega^{\prime} 1_{\alpha}^{+}} \tag{3.30}
\end{align*}
$$

where the primed quantities have the same meaning as the corresponding unprimed ones in (3.3) and (3.4). The $\beta \beta_{0 \nu}$-moments are evaluated in the same way. That is, Eq. (2.16) is evaluated as

$$
\begin{equation*}
\rho^{p h}\left(p n p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right)=\rho^{\prime-}\left(p n ; J_{\alpha}^{\pi}\right) \rho^{\prime+}\left(p^{\prime} n^{\prime} ; J_{\alpha}^{\pi}\right) \tag{3.31}
\end{equation*}
$$

and Eq. (2.25) as

$$
\begin{equation*}
v\left(k ; \omega_{J_{\alpha}^{\pi}}\right)=v\left(k ; \Omega_{J_{\alpha}^{\pi}}\right) \tag{3.32}
\end{equation*}
$$

Finally, the unperturbed (BCS) one body densities are

$$
\begin{align*}
\rho_{B C S}^{\prime-}\left(p n J_{\alpha}^{\pi}\right) & =u_{p}^{\prime} v_{n}^{\prime} \\
\rho_{B C S}^{\prime+}\left(p n J_{\alpha}^{\pi}\right) & =u_{n}^{\prime} v_{p}^{\prime} \tag{3.33}
\end{align*}
$$

As already pointed out in Ref. [49], the two-QRPA Methods I and II involve also the nuclei $(N+1, Z-1)$ and $(N-3, Z+3)$. This is so because the GSC for the transitions $(N, Z) \xrightarrow{\beta^{-}}(N-1, Z+1)$ and $(N-1, Z+$ $1) \xrightarrow{\beta^{-}}(N-2, Z+2) \cong(N-2, Z+2) \xrightarrow{\beta^{+}}(N-1, Z+1)$ correspond, respectively, to transitions $(N, Z) \xrightarrow{\beta^{+}}(N+$ $1, Z-1)$ and $(N-2, Z+2) \xrightarrow{\beta^{-}}(N-3, Z+3)$. On the contrary, the one-QRPA Methods III and IV only involve the nuclei within the isobaric triplet $(N, Z),(N-1, Z+1)$, $(N-2, Z+2)$ where the $\beta \beta$-decay occurs. Also, in the last methods, similarly to what happens in the single $\beta$ decay, the GSC for the decay $(N, Z) \xrightarrow{\beta^{-}}(N-1, Z+1)$ is the decay $(N-1, Z+1) \xrightarrow{\beta^{-}}(N-2, Z+2) \cong(N-$ $2, Z+2) \xrightarrow{\beta^{+}}(N-1, Z+1)$, and vice versa.

Shortly after having been formulated, all four methods were extended to the $\beta \beta_{0 \nu}$ moments $[21,22,24-26,50$,

51], where the importance of the GSC was evidenced once again. However, Method IV is being used here for the first time in a simultaneous study of both double decay beta modes. This was precisely the main motivation to present the numerical results that follow, based on this one-QRPA method, and by fixing the isoscalar strength $t$ from the PSU4SR.

## IV. NUMERICAL RESULTS

## A. Method IV with $t$ from PSU4SR

As explained in Sec. I, within our modus operandi all nuclear model parameters are fixed. To set them in the $p h$ channel we use the energetics of the IAS and GTR [31], with the result $v_{p h}^{s}=55$ and $v_{p h}^{t}=92$ in units of MeV . $\mathrm{fm}^{3}$. These values are used for all nuclei, with exception of ${ }^{48} \mathrm{Ca}$ where $v_{p h}^{s}=27$ and $v_{p h}^{t}=64$ were employed. Within the $p p$ channel, $v_{p p}^{s}$ and $v_{p p}^{t}$, or more precisely, the ratios $s$ and $t$, are determined from the condition that the strengths $S_{F}^{+}$and $S_{G T}^{+}$become minimal.

TABLE I: Values of the parameters $s_{s y m}$ and $t_{s y m}$, and experimental and calculated energies of the IAS and GTR in the initial nucleus. The energies are given in units of MeV .

|  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| ${ }^{A} Z$ | $s_{s y m}$ | $t_{\text {sym }}$ | $E_{I A S}^{c a l}$ | $E_{I A S}^{e x p}$ | $E_{G T R}^{c a l}$ | $E_{G T R}^{e x p}$ |
|  |  |  |  |  |  |  |
| ${ }^{48} \mathrm{Ca}$ | 1.00 | 1.20 | 8.70 | 7.36 | 13.66 | 11.43 |
| ${ }^{76} \mathrm{Ge}$ | 1.00 | 1.23 | 11.47 | 10.21 | 13.92 | 13.42 |
| ${ }^{82} \mathrm{Se}$ | 1.00 | 1.30 | 12.25 | 10.59 | 15.59 | 13.41 |
| ${ }^{96} \mathrm{Zr}$ | 1.00 | 1.55 | 14.18 | 11.85 | 16.10 | 14.45 |
| ${ }^{100} \mathrm{Mo}$ | 1.00 | 1.49 | 13.70 | 12.29 | 15.83 | 14.93 |
| ${ }^{128} \mathrm{Te}$ | 1.00 | 1.41 | 13.74 | 14.06 | 14.36 | 15.75 |
| ${ }^{130} \mathrm{Te}$ | 1.00 | 1.45 | 14.71 | 13.98 | 14.95 | 15.42 |
| ${ }^{150} \mathrm{Nd}$ | 1.00 | 1.29 | 20.21 | 15.42 | 18.46 | 16.61 |

Fig. 1 shows the behavior of vector observables as a function of $s$ (on the left side), and axial-vector observable as a function of $t$ (on the right side) for ${ }^{48} \mathrm{Ca},{ }^{76} \mathrm{Ge}$, ${ }^{82} \mathrm{Se}$ and ${ }^{96} \mathrm{Zr}$ nuclei. Their behaviors are very similar, which led us to propose our recipe to set the isoscalar strength. One sees that $s_{\text {sym }}=1.0$, while values of $t_{\text {sym }}$ (indicated by points on the axis $t$ ) are $>1$, and vary with the mass number $A$. Exactly the same happens for the remaining four nuclei, and their values of $s_{s y m}$ and $t_{\text {sym }}$ are listed in Table I. The values exhibited for the latter are very close to those obtained previously in [21, Table 4], and in [28, Table 4.9], where Method III was used to calculate the NM. The above similarity is the main reason for associating PSU4SR to isospin symmetry restoration in $\beta \beta$-decay.

In the literature, the isoscalar strength is usually adjusted by employing the measured $\beta \beta_{2 \nu}$ half-lives, from which the absolute value of $M^{2 \nu}$ can be extracted but not its sign. In doing so, two different values of $t$ are obtained: one $\left(t=t_{\uparrow}\right)$ when $M^{2 \nu}$ is assumed to be positive $\left(M^{2 \nu} \equiv M_{\uparrow}^{2 \nu}>0\right)$, and another one $\left(t=t_{\downarrow}\right)$ when $M^{2 \nu}$ is assumed to be negative $\left(M^{2 \nu} \equiv M_{\downarrow}^{2 \nu}<0\right) .{ }^{5}$ In addition to this disadvantage, it is clear that in this case the model is no longer predictive. We have done such a calculation in Ref. [21] within Method III, finding that in all cases $t_{\text {sym }} \approx t_{\uparrow}$.

Also in Table I, the theoretical energies of $F$ and $G T$ resonances are displayed. They are defined, respectively, as

$$
\begin{equation*}
E_{I A S}=\frac{\sum_{p n \alpha}\left|\rho^{\prime-}\left(p n 0_{\alpha}^{+}\right)\right|^{2} \Omega_{0_{\alpha}^{+}}^{\prime}}{\sum_{p n \alpha}\left|\rho^{\prime-}\left(p n 0_{\alpha}^{+}\right)\right|^{2}} \tag{4.1}
\end{equation*}
$$

and

$$
\begin{equation*}
E_{G T R}=\left.\frac{\sum_{p n \alpha}\left|\rho^{\prime-}\left(p n 1_{\alpha}^{+}\right)\right|^{2} \Omega_{1_{\alpha}^{+}}^{\prime}}{\sum_{p n \alpha}\left|\rho^{\prime-}\left(p n 1_{\alpha}^{+}\right)\right|^{2}}\right|_{\substack{\Omega_{1}^{\prime}>10 \mathrm{MeV}}} \tag{4.2}
\end{equation*}
$$

where the constraint on $G T$ energies has been imposed since significant $G T$ strength is always observed at low energy far from the location of the GTR. This is particularly so in the case of ${ }^{96} \mathrm{Zr}$, which causes the calculated energy of the GTR to be relatively low.

It is worth noting that, for $s=s_{\text {sym }}$ and $t=t_{\text {sym }}$, the total strengths $S_{F}^{+}$and $S_{G T}^{+}$not only are minimum, but the concentrations of transition intensities $S_{F}^{-}$and $S_{G T}^{-}$ in resonant states are also maximum.

The corresponding experimental energies of the IAS are evaluated as

$$
\begin{equation*}
E_{I A S}^{E x p}=\mathcal{E}_{\text {Coul }}(Z+1, A)-\mathcal{E}_{\text {Coul }}(Z, A) \tag{4.3}
\end{equation*}
$$

where (see Eq. (69) in Ref. [5])

$$
\begin{equation*}
\mathcal{E}_{\text {Coul }}(Z, A)=0.70 \frac{Z^{2}}{A^{1 / 3}}\left[1-0.76 Z^{-2 / 3}\right] \mathrm{MeV} \tag{4.4}
\end{equation*}
$$

The energy difference $E_{G T R}^{E x p}-E_{I A S}^{E x p}$ is estimated from (1.5). A relatively good agreement between the calculations and the experimental estimates indicates that: 1) our choice of the coupling constants in the $p h$ channel is reasonable, and 2) in the closure approximation for the $\beta \beta_{0 \nu}$-decay it is proper to replace the intermediate energies $\omega_{J_{\alpha}^{+}}$in (2.23) by an average value $\bar{\omega}_{J_{\alpha}^{+}}=12 \mathrm{MeV}$. It should be said, however, that the calculated energies $E_{I A S}^{c a l}$ differ appreciably from the "experimental" energies

[^5]

FIG. 1: (Color online) (a) \& ( $\mathrm{a}^{\prime}$ ): $0 \nu$ NM normalized to $g_{A}^{2} ;(\mathrm{b}) \&\left(\mathrm{~b}^{\prime}\right): 2 \nu$ NM given in natural units; and (c) $\&\left(\mathrm{c}^{\prime}\right): \beta^{+}-$decay transition strengths. We show the vector observables, as a function of the ratio $s=v_{p p}^{s} / \bar{v}_{p a i r}^{s}$, on the left side, and axial-vector ones, as a function of the ratio $t=v_{p p}^{t} / \bar{v}_{\text {pair }}^{s}$, on the right side. The values of $t_{s y m}$ on the axis $t$ are indicated by points.

TABLE II: $\beta \beta_{2 \nu}$-decay moments evaluated within the BCS (unperturbed) and QRPA (perturbed) approximations are compared with the experimental results $\left|M_{e x p}^{2 \nu}\right|$ recommended by Barabash [3]. All the quantities are given in natural units. As explained in the text, the upper and lower theoretical errors on $M^{2 \nu}$ were evaluated with $t=1.03 \times t_{\text {sym }}$, and $t=0.97 \times t_{\text {sym }}$, respectively, where the values of $t_{\text {sym }}$ are those listed in Table I.

|  | BCS |  |  | QRPA |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| ${ }^{A} Z$ | $M_{F}^{2 \nu}$ | $M_{G T}^{2 \nu}$ | $M^{2 \nu}$ | $M_{F}^{2 \nu}$ | $M_{G T}^{2 \nu}$ | $\left\|M^{2 \nu}\right\|$ | $\left\|M_{\text {exp }}^{2 \nu}\right\|$ |
| ${ }^{48} \mathrm{Ca}$ | -0.148 | -0.545 | -0.693 | -0.004 | 0.022 | $0.018_{-0.035}^{+0.110}$ | $0.038 \pm 0.003$ |
| ${ }^{76} \mathrm{Ge}$ | -0.193 | -0.693 | -0.886 | -0.000 | 0.051 | $0.051_{-0.030}^{+0.035}$ | $0.113 \pm 0.006$ |
| ${ }^{82} \mathrm{Se}$ | -0.217 | -0.686 | -0.903 | -0.001 | 0.062 | $0.062_{-0.033}^{+0.033}$ | $0.083 \pm 0.004$ |
| ${ }^{96} \mathrm{Zr}$ | -0.107 | -0.878 | -0.985 | -0.001 | 0.024 | $0.023_{-0.036}^{+0.0157}$ | $0.080 \pm 0.004$ |
| ${ }^{100} \mathrm{Mo}$ | -0.126 | -1.213 | -1.339 | -0.001 | 0.035 | $0.034_{-0.182}^{+0.185}$ | $0.185 \pm 0.005$ |
| ${ }^{128} \mathrm{Te}$ | -0.296 | -1.174 | -1.470 | -0.003 | 0.086 | $0.083_{-0.026}^{+0.029}$ | $0.046 \pm 0.006$ |
| ${ }^{130} \mathrm{Te}$ | -0.263 | -1.025 | -1.288 | -0.002 | 0.083 | $0.081_{-0.022}^{+0.022}$ | $0.031 \pm 0.004$ |
| ${ }^{150} \mathrm{Nd}$ | -0.057 | -0.887 | -0.944 | -0.001 | 0.067 | $0.067_{-0.011}^{+0.011}$ | $0.058 \pm 0.004$ |

$E_{I A S}^{e x p}$ in the case of ${ }^{48} \mathrm{Ca}$ and ${ }^{150} \mathrm{Nd}$ nuclei. The difference between energies $E_{G T R}^{c a l}$ and $E_{G T R}^{e x p}$ in ${ }^{96} \mathrm{Zr}$ is quite significant also. But, we have not found any satisfactory explanation for these discrepancies.

To appreciate the effect of the residual interaction, and hence of the PSU4SR, we will compare the QRPA with
the BCS, which is its mean field approximation in this case. We show our calculated values of the $\beta \beta_{2 \nu}$-decay moments in natural units as the corresponding experimental moments, recommended by Barabash [3], are given in these units. Our BCS (unperturbed) and QRPA (perturbed) results are listed in Table II, together with

TABLE III: $\beta \beta_{0 \nu}$-decay moments $M_{X}^{0 \nu}$, as well the total moments $M^{0 \nu}=\sum_{X} M_{X}^{0 \nu}$ (normalized to $g_{A}^{2}$, with $g_{A}=1.27$ ), evaluated within the BCS (unperturbed) and QRPA (perturbed) approximations, are shown. In both cases the FNS and SRC effects are included. Same as in Table II, the upper and lower theoretical g on $M^{0 \nu}$ were evaluated with $t=1.03 \times t_{s y m}$, and $t=0.97 \times t_{s y m}$, respectively, where the values of $t_{\text {sym }}$ are those listed in Table I. At the bottom of the table are shown the ${ }^{76}$ Ge results: i) without SRC , in the row labeled as ${ }^{76} \mathrm{Ge}{ }^{*}$, ii) the bare values of moments, i.e., without the FNS and SRC effects, in the row labeled as ${ }^{76} \mathrm{Ge}{ }^{* *}$, and iii) the moments obtained in Ref. [11] and derived from relations (4.5).

|  | BCS |  |  |  |  | QRPA |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| ${ }^{A} Z$ | M | $M_{A}^{0 \nu}$ | $M_{P}^{0 \nu}$ | $M_{M}^{0 \nu}$ | $M^{0 \nu}$ | $M_{V}^{0 \nu}$ | $M_{A}^{0 \nu}$ | $M_{P}^{0 \nu}$ | $M_{M}^{0 \nu}$ | $M^{0 \nu}$ |
| ${ }^{48} \mathrm{Ca}$ | 1.91 | 9.10 | -1.54 | 0.49 | 9.96 | 0.58 | 2.57 | -0.76 | 0.33 | $2.72_{+0.32}^{-0.40}$ |
| ${ }^{76} \mathrm{Ge}$ | 2.52 | 12.35 | -2.15 | 0.71 | 13.42 | 0.64 | 3.02 | -0.86 | 0.40 | $3.19_{+0.46}^{-0.24}$ |
| ${ }^{82} \mathrm{Se}$ | 2.61 | 12.58 | -2.21 | 0.72 | 13.70 | 0.65 | 2.76 | -0.84 | 0.39 | $2.96_{+0.22}^{-0.23}$ |
| ${ }^{96} \mathrm{Zr}$ | 2.43 | 12.70 | -2.15 | 0.71 | 13.70 | 0.70 | 1.89 | -0.74 | 0.38 | $2.22_{+0.35}^{-0.42}$ |
| ${ }^{100} \mathrm{Mo}$ | 2.85 | 15.17 | -2.51 | 0.84 | 16.35 | 0.82 | 2.48 | -0.90 | 0.45 | $2.85{ }_{+0.42}^{-0.43}$ |
| ${ }^{128} \mathrm{Te}$ | 2.78 | 13.55 | -2.13 | 0.66 | 14.87 | 0.84 | 3.31 | -0.97 | 0.41 | $3.59_{+0.19}^{-0.19}$ |
| ${ }^{130} \mathrm{Te}$ | 2.48 | 12.12 | -1.91 | 0.60 | 13.29 | 0.75 | 2.81 | -0.84 | 0.36 | $3.07_{+0.16}^{-0.16}$ |
| ${ }^{150} \mathrm{Nd}$ | 2.02 | 10.94 | -1.75 | 0.57 | 11.77 | 0.77 | 3.95 | -0.93 | 0.37 | $4.16_{+0.11}^{-0.12}$ |
| ${ }^{76} \mathrm{Ge}^{*}$ | 2.5 | 12.54 | -2.21 | 0.71 | 13.57 | 0.65 | 3.14 | -0.90 | 0.40 | 3.29 |
| ${ }^{76} \mathrm{Ge}^{* *}$ | 2.90 | 13.72 | -2.55 | 1.08 | 15.14 | 0.85 | 3.83 | -1.11 | 0.65 | 4.22 |
| ${ }^{76} \mathrm{Ge}$ [11] |  |  |  |  |  | 1.74 | 5.48 | -1.60 | 0.29 | 5.26 |

the so-called effective moments $\left|M_{e x p}^{2 \nu}\right|=g_{A}^{2}\left|M^{2 \nu}\right|$ from Ref. [3]. For the axial-vector coupling constant we have used its free nucleon value of $g_{A}=1.27$ [36], instead of the quenched value $g_{A}^{\text {eff }} \cong 1$ since: 1 ) in this way we obtain a better agreement with data, and 2) although $g_{A}^{\text {eff }}$ is often used in the description of simple $\beta$-decays, there is no compelling evidence for using it in the $\beta \beta$-decays.

Even though the recipe introduced above to set the parameters in the $p p$ channel makes the theory predictive, this does not necessarily mean that the theoretical predictions have to match with the experimental data. But still, it is not possible to completely suppress the well known sensitivity of $M^{2 \nu}$ on the parameter $t$ in the neighborhood of $t_{\text {sym }}[21,23]$. This, in turn, means that a relatively small variation of $t$ causes large variations in $M^{2 \nu}$, being particularly pronounced for ${ }^{48} \mathrm{Ca}$ and ${ }^{100} \mathrm{Mo}$. Here one should keep in mind that ${ }^{48} \mathrm{Ca}$ is a double-closed shell nucleus, while ${ }^{100} \mathrm{Mo}$ has the $Z=40$ subshell closed. Therefore, the QRPA may not be the fully appropriate model in these two cases.

The above behavior can be seen clearly by appealing to the SMM discussed in the Appendix, where $t_{\text {sym }}$ is well defined $\left(t_{\text {sym }} \equiv t_{0}\right)$ and $t_{0}$ depends on the dominant $\left|p n ; 1^{+}\right\rangle$intermediate state. This, however, does not happen in full numerical calculations where $t_{\text {sym }}$ depends in a significant way on the mean field, through the pairing coupling $\bar{v}_{s}^{\text {pair }}$, and the single-particle energies (spe). These quantities are determined in a phenomenological way, and, therefore, are not well established ${ }^{6}$. We have

[^6]assumed that the resulting uncertainties can be quantified by attributing errors of $\pm 3 \%$ to $t_{\text {sym }}$. Thus, the upper and lower theoretical errors on $M^{2 \nu}$ in Table II were evaluated with $t=1.03 \times t_{\text {sym }}$, and $t=0.97 \times t_{\text {sym }}$, respectively, where the values of $t_{s y m}$ are those listed in Table I. It turns out, however, that the value of $\left|M_{e x p}^{2 \nu}\right|$ in ${ }^{76} \mathrm{Ge},{ }^{128} \mathrm{Te}$ and ${ }^{130} \mathrm{Te}$ fall outside the theoretical errors.

Even if the agreement between theory and data is not as good as one might wish, it is interesting to note that the BCS results differ from $\left|M_{\text {exp }}^{2 \nu}\right|$ in a manner similar to the differences given by the QRPA results. In fact, while $\left|M_{e x p}^{2 \nu}\right|$ are smaller by a factor ranging from 7.2 in ${ }^{100} \mathrm{Mo}$ to 42 in ${ }^{130} \mathrm{Te}$, the QRPA results are smaller by a factor going from 14.0 in ${ }^{150} \mathrm{Nd}$ to 42.8 in ${ }^{96} \mathrm{Zr}$. Large differences (roughly of one order of magnitude) between BCS and QRPA moments come from the PSU4SR, which is crucial to make the theory consistent with experimental data. The conservation of the number of particles is, by far, less significant [20]. It is worth noting that, while in the BCS approximation the moments $M_{F}^{2 \nu}$ contribute significantly to the total moments $M^{2 \nu}$, in the QRPA approach they can be neglected for all practical purposes. Moreover, given that there are no free parameters in the nuclear model, the agreement between the theory and the data, as seen from the last two columns in Table II, could be considered to be fairly good.

We must take some care when comparing our four $\beta \beta_{0 \nu}$-decay moments $M_{V}^{0 \nu}, M_{A}^{0 \nu}, M_{P}^{0 \nu}, M_{M}^{0 \nu}$, with those defined by other groups, since we do not separate the

TABLE IV: Fine structure of $M^{0 \nu}$ moments (normalized to $g_{A}^{2}$, with $g_{A}=1.27$ ) for ${ }^{76} \mathrm{Ge}$. The contributions of different intermediate-state angular momenta $J^{\pi}$ are listed for both parities $\pi= \pm$.

|  | BCS |  |  |  |  |  | QRPA |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $J^{\pi}$ | $M_{V}^{0 \nu}$ | $M_{A}^{0 \nu}$ | $M_{P}^{0 \nu}$ | $M_{M}^{0 \nu}$ | $M^{0 \nu}$ | $M_{V}^{0 \nu}$ | $M_{A}^{0 \nu}$ | $M_{P}^{0 \nu}$ | $M_{M}^{0 \nu}$ | $M^{0 \nu}$ |  |
| $0^{+}$ | 1.06 | 0.00 | 0.00 | 0.00 | 1.06 | 0.02 | 0.00 | 0.00 | 0.00 | 0.02 |  |
| $1^{+}$ | 0.00 | 4.75 | -0.48 | 0.05 | 4.33 | 0.00 | -0.39 | -0.05 | 0.01 | -0.43 |  |
| $2^{+}$ | 0.36 | 0.54 | 0.00 | 0.05 | 0.95 | 0.14 | 0.24 | 0.00 | 0.03 | 0.40 |  |
| $3^{+}$ | 0.00 | 1.01 | -0.35 | 0.06 | 0.72 | 0.00 | 0.45 | -0.16 | 0.03 | 0.32 |  |
| $4^{+}$ | 0.14 | 0.23 | 0.00 | 0.04 | 0.42 | 0.08 | 0.14 | 0.00 | 0.03 | 0.24 |  |
| $5^{+}$ | 0.00 | 0.40 | -0.18 | 0.04 | 0.27 | 0.00 | 0.24 | -0.11 | 0.03 | 0.16 |  |
| $6^{+}$ | 0.06 | 0.10 | 0.00 | 0.03 | 0.18 | 0.04 | 0.07 | 0.00 | 0.02 | 0.13 |  |
| $7^{+}$ | 0.00 | 0.15 | -0.07 | 0.02 | 0.11 | 0.00 | 0.11 | -0.05 | 0.02 | 0.08 |  |
| $8^{+}$ | 0.02 | 0.03 | 0.00 | 0.01 | 0.06 | 0.01 | 0.02 | 0.00 | 0.01 | 0.05 |  |
| $9^{+}$ | 0.00 | 0.06 | -0.03 | 0.01 | 0.05 | 0.00 | 0.04 | -0.02 | 0.01 | 0.03 |  |
| $10^{+}$ | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |  |
| $\pi=+$ | 1.64 | 7.27 | -1.11 | 0.33 | 8.15 | 0.29 | 0.92 | -0.39 | 0.19 | 1.00 |  |
| $0^{-}$ | 0.00 | 0.15 | -0.07 | 0.00 | 0.08 | 0.00 | 0.07 | -0.04 | 0.00 | 0.03 |  |
| $1^{-}$ | 0.47 | 0.62 | 0.00 | 0.03 | 1.12 | 0.15 | 0.24 | 0.00 | 0.01 | 0.40 |  |
| $2^{-}$ | 0.00 | 2.26 | -0.47 | 0.06 | 1.85 | 0.00 | 0.66 | -0.16 | 0.02 | 0.52 |  |
| $3^{-}$ | 0.24 | 0.43 | 0.00 | 0.06 | 0.72 | 0.11 | 0.23 | 0.00 | 0.03 | 0.37 |  |
| $4^{-}$ | 0.00 | 0.80 | -0.29 | 0.06 | 0.57 | 0.00 | 0.40 | -0.15 | 0.03 | 0.28 |  |
| $5^{-}$ | 0.12 | 0.21 | 0.00 | 0.05 | 0.38 | 0.06 | 0.13 | 0.00 | 0.03 | 0.23 |  |
| $6^{-}$ | 0.00 | 0.36 | -0.15 | 0.05 | 0.26 | 0.00 | 0.21 | -0.09 | 0.03 | 0.15 |  |
| $7^{-}$ | 0.05 | 0.10 | 0.00 | 0.04 | 0.19 | 0.03 | 0.07 | 0.00 | 0.03 | 0.12 |  |
| $8^{-}$ | 0.00 | 0.10 | -0.05 | 0.02 | 0.08 | 0.00 | 0.07 | -0.03 | 0.01 | 0.05 |  |
| $9^{-}$ | 0.00 | 0.01 | 0.00 | 0.00 | 0.02 | 0.00 | 0.01 | 0.00 | 0.00 | 0.02 |  |
| $10^{-}$ | 0.00 | 0.02 | -0.01 | 0.00 | 0.01 | 0.00 | 0.02 | -0.01 | 0.00 | 0.01 |  |
| $\pi=-$ | 0.88 | 5.06 | -1.04 | 0.37 | 5.28 | 0.35 | 2.11 | -0.48 | 0.19 | 2.18 |  |

tensor contribution from the $G T$ contribution, nor do we separate $M_{P}^{0 \nu}$ into its $P P$ and $A P$ pieces. For instance, when confronted with the results of Ref. [11] the following correspondence is valid:

$$
\begin{align*}
& M_{V}^{0 \nu} \rightarrow M_{F}^{V V} \\
& M_{A}^{0 \nu} \rightarrow M_{G T}^{A A} \\
& M_{M}^{0 \nu} \rightarrow M_{G T}^{M M}+M_{T}^{M M} \\
& M_{P}^{0 \nu} \rightarrow M_{G T}^{P P}+M_{T}^{P P}+M_{G T}^{A P}+M_{T}^{A P} . \tag{4.5}
\end{align*}
$$

The moments labeled as $G T$ and $T$ on the right side are, respectively, the $m=0$ and $m=2$ parts of the moments $M_{M}^{0 \nu}$ and $M_{P}$ in (2.19). This expression also permits an easy visualization of the meaning of moments labeled as $A P$ and $P P$.

Our four $\beta \beta_{0 \nu}$-decay moments and their sums $M^{0 \nu}$, evaluated within the BCS (unperturbed) and QRPA (perturbed) approximations are shown in Table III. In both cases, the FNS and SRC effects are included, and the summations over $J_{\alpha}^{\pi}$ in (2.19) go from $J=0$ to $J=10$ for both parities. The numerical results are normalized to $g_{A}^{2}$ in order to compare with other calculations.

Some additional results for the $0 \nu \mathrm{NM}$ in ${ }^{76} \mathrm{Ge}$ are
also shown in Table III, namely i) row ${ }^{76} \mathrm{Ge}$ *: without the effect of SRC, ii) row ${ }^{76} \mathrm{Ge}{ }^{* *}$ : the bare values, i.e., without the FNS and SRC effects, and iii) row Ref. [11]: results obtained in this paper by Hyvärinen and Suhonen for $g_{A}=1.26$, and related to our calculations by means of equations (4.5).

It is worth mentioning that moment $M_{V}^{0 \nu}$ in Ref. [11] is significantly greater than ours, which makes the corresponding total moment $M^{0 \nu}$ also much greater than ours. Something similar can be observed from the comparison of the results for most of the other nuclei, as well as when comparing the results of the Refs. [9, 10] with the present results. Moreover, the moments $M^{0 \nu}$ in [11] are not always greater than ours, as, for example, is the case of ${ }^{96} \mathrm{Zr}$. This makes it very difficult to find the reason for the disagreements between different calculations.

The QRPA moments $M^{0 \nu}$ are also sensitive to the parameter $t$ in the neighborhood of $t_{\text {sym }}$, although not in such a pronounced way as $M^{2 \nu}$. The resulting theoretical QRPA uncertainties, quantified in the way described before, are also shown. These come basically from the uncertainties in $M_{A}^{0 \nu}$, a little bit from $M_{P}^{0 \nu}$ and $M_{M}^{0 \nu}$, and nothing from $M_{V}^{0 \nu}$. Again, the most affected are the
${ }^{48} \mathrm{Ca}$ and ${ }^{100} \mathrm{Mo}$ moments.
The following conclusions can be drawn from the results for the moments $M^{0 \nu}$ :
i) The role of the residual interaction, through the PSU4SR, is critical in reducing the nuclear moments. The reduction for the neutrinoless decay is, however, less pronounced than in the case of two-neutrino decay, as the perturbed (QRPA) moments $M^{0 \nu}$ are only $\sim 5-7$ times smaller than the unperturbed (BCS)moments.
ii) This quenching effect is smaller on induced current moments $M_{P}^{0 \nu}$ and $M_{M}^{0 \nu}$ than on $M_{V}^{0 \nu}$ and $M_{A}^{0 \nu}$, which results from the standard V-A weak current.
iii) Our $M_{M}^{0 \nu}$ are, in principle, larger than in other calculations by the factor $\left(f_{M} / g_{M}\right)^{2}=1.61$, since we include the term $g_{V} / 2 M_{N}$ in the NRA of the weak Hamiltonian as is usually done in studies of single $\beta$-decays. This can be clearly seen from Table III, where all ${ }^{76} \mathrm{Ge}$ moments $M_{X}^{0 \nu}$ from Ref. [11] are higher than ours, except $M_{M}^{0 \nu}$. Note that the differences between both calculations are by far larger than our numerical uncertainties.
iv) Compared to the role played by the residual interaction in the $p p$ channel, the FNS and SRC effects are relatively small. Indeed, the FNS effects cause the bare elements to decrease by $\sim 15-20 \%$, and when the SRC are added an additional decrease of $\sim 3-5 \%$ is produced. These findings are fully consistent with the results exhibited in Table I of Ref. [53], when the SRC are evaluated in the framework of the coupled-cluster method. Moreover, according to the recent studies based on the unitary correlation operator method (UCOM) [54, 55], the SRC have a marginal reduction effect $(<10 \%)$ on the $\beta \beta_{0 \nu}$-decay moments. Due to this fact, as well as because of computational difficulties, their contributions were omitted directly in a recent paper [35]. Our method to evaluate the SRC, given by (2.28), does not guarantee the correct normalization of the two-body wave function. But, this is a small correction on an effect, which by itself is small, and, therefore, it cannot be relevant. We also note that the effects of the SRC are smaller than our estimate of the theoretical uncertainties.

Fine structure of $M^{0 \nu}$ in the case ${ }^{76} G e$ is exhibited in Table IV, where contributions of different intermediatestate angular momenta $J^{\pi}$ are listed for both parities $\pi= \pm$. Most notable issues in this table are:

1) For $M_{V}^{0 \nu}$, only the natural parity intermediate states $\pi=(-)^{J}$, i.e., for $J^{\pi}=0^{+}, 1^{-}, 2^{+}, \cdots$, contribute.
2) For $M_{P}^{0 \nu}$, only the unnatural parity intermediate states $\pi=-(-)^{J}$, i.e., for $J^{\pi}=0^{-}, 1^{+}, 2^{-}, \cdots$, contribute.
3) The residual interaction in the $p p$ channel mainly affects the moment $M_{V}^{0 \nu}$ for $J^{\pi}=0^{+}$and the moment $M_{A}^{0 \nu}$ for $J^{\pi}=1^{+}$.
4) In the QRPA, the negative parity states dominate the positive parity states.

TABLE V: $M^{2 \nu}$ and $M^{0 \nu}$ NM within Method II and Method IV, with three different approximations (App) for the parameter $t$; namely, values obtained from the PSU4SR $\left(t_{\text {sym }}\right)$ and from $\left|M_{e x p}^{2 \nu}\right|\left(t_{\uparrow}\right.$ and $\left.t_{\downarrow}\right)$. Results from Ref. [11] for $g_{A}=1.26$, which should be compared with ours for $t_{\downarrow}$ when is used the Method II, are also shown.


## B. Comparison between Methods II and IV

Relevant results for the comparison of the $0 \nu \mathrm{NM}$, obtained within the two-QRPA-method (Method II) and the one-QRPA-method (Method IV), are presented in Table V. This is done for both manners of fixing the isoscalar parameter $t$, i.e., through the PSU4SR $\left(t_{\text {sym }}\right)$ and from experimental $\beta \beta_{2 \nu}$ NM $\left|M_{e x p}^{2 \nu}\right|\left(t_{\uparrow}\right.$ and $\left.t_{\downarrow}\right)$. A few details for ${ }^{76} \mathrm{Ge}$ are shown in Fig. 2, for different approximations for the parameter $t$. In comparing the results with Methods II and IV, for the same values of $t_{\text {sym }}$ that are given in Table I, one notices quite signif-


FIG. 2: (Color online) Isoscalar parameters $t$ in ${ }^{76} \mathrm{Ge}$ within the Method II for $\left|M_{e x p}^{2 \nu}\right|=0.113$. In (a) the NM $M^{2 \nu}$ is given in natural units; while in (b) the $M^{0 \nu}$ is dimensionless. It should be noted that $M^{2 \nu}$ is negative at $t=0$.
icant differences for both $2 \nu$ and $0 \nu$ NM. It was to be expected that the $2 \nu$ moments differ significantly, since they are very sensitive to $t$. But, it is somewhat surprising that, except in the case of ${ }^{128} \mathrm{Te}$, the $0 \nu$ moments were appreciably larger using Method II instead of Method IV.

As already observed in Ref. [21], the isoscalar strength can not be determined univocally by adjusting the calculation to the measured half-life, since only the module of $M_{\text {exp }}^{2 \nu}$ is obtained from experimental data. As a consequence, two values of $t$ result for the calculated NM: one when $M_{e x p}^{2 \nu}$ is assumed to be positive $\left(t=t_{\uparrow}\right)$, and one when it is assumed to be negative $\left(t=t_{\downarrow}\right)$. Due to a smooth variation of calculated $M^{0 \nu}$ in the neighborhood where $M^{2 \nu}$ passes through zero, the values of $M_{\uparrow}^{0 \nu}$ and $M_{\downarrow}^{0 \nu}$ are not severely different from each other and in most of the cases $M_{\uparrow}^{0 \nu} \sim M_{s y m}^{0 \nu}$.

The statement quoted above, on the difference in the values of $M^{0 \nu}$ obtained for $t_{\text {sym }}$ when using Method II or Method IV, attracts attention. To fully convince ourselves of this, we also compare in Table V the results obtained for $M^{0 \nu}$ from $t_{\uparrow}$ and $t_{\downarrow}$, arriving at nearly the same conclusion. As both comparisons are consistent, our statement has greater reliability.

To calculate the $0 \nu$ moments in the literature $t_{\downarrow}$ is usually chosen, or its equivalent $[9,11]$, (see, for instance, [56, Fig. 1], and [57, Fig. 6]), even though there is no reason a priori to disregard results obtained with $t_{\uparrow}$. Therefore, it is appropriate to compare the $M^{0 \nu}$ values reported in these works for $g_{A}=1.27$ and $g_{A}=1.26$, respectively, with our $M_{\downarrow}^{0 \nu}$. The main difference is that
in Refs. [9] and [11], two-nucleon interactions, based on the Bonn one-boson-exchange G matrix, have been used as the residual interaction, instead of the simple $\delta$-force given by (1.2). Despite this important difference, the discrepancy between our results and those of Ref. [11] are not drastically large, as seen in Table V. In fact, the differences between the present results and those from [11, Table III] are of the same order as those between the results in Refs. [9] and [11].

More explicitly, if we proceed in the same way as in Ref. [11] and quantify the differences by the relative differences $\left(\left|M^{0 \nu}[10]-M_{\downarrow}^{0 \nu}\right| / M_{\downarrow}^{0 \nu}\right)$, we get $(9,0,41,38,22,16) \%$ for $\left({ }^{76} \mathrm{Ge},{ }^{82} \mathrm{Se},{ }^{96} \mathrm{Zr},{ }^{100} \mathrm{Mo},{ }^{128} \mathrm{Te}\right.$, ${ }^{130} \mathrm{Te}$ ), which should be confronted with the differences $(6,35,6,50,3,9) \%$, listed in the last column of [11, Table III] for the same set of nuclei. In our calculation the differences go up $\sim 40 \%$ in ${ }^{96} \mathrm{Zr}$ and ${ }^{100} \mathrm{Mo}$ because the $t$ value is in the QRPA breakdown region, which is reflected in the theoretical errors shown in Table II.

This behavior of the QRPA is a well-known puzzle in the $\beta \beta_{2 \nu}$-decay, and it has not yet been fully disentangled, despite much effort being invested in doing so, through the renormalized QRPA (RQRPA) [23, 58, 59]. However, the results for ${ }^{100}$ Mo shown in [23, Fig. 3] and [58, Fig. 3] could be considered auspicious, since within the RQRPA the moment $M^{2 \nu}$ behaves smoothly in the region of $t$ where the ordinary QRPA collapses.

Rodin et al. [60] justify the procedure of fixing $t$ from $\beta \beta_{2 \nu}$ data, since in this way the $M^{0 \nu}$ values become essentially independent of the size of the single-particle basis. The same thing happens, however, when this parameter is fixed by the PSU4SR procedure. This can be seen, for instance, from Fig. 2 in Ref. [21], where three different single-particle bases for ${ }^{48} \mathrm{Ca}$ have been used in the framework of one-QRPA Method III.

The above authors [60] also argue that: "It follows from the study of Ref. [61] that choosing the negative sign of $\left|M_{e x p}^{2 \nu}\right|$ would lead to a complete disagreement with the systematics of single beta decays." This assertion is based mostly on a work performed within the deformed QRPA [61], where the $\beta \beta_{2 \nu}$-decays suppression mechanism is attributed to nuclear deformation. Such a view is obviously in total opposition to ours, in which the decisive player is the restoration of the $\mathrm{SU}(4)$ symmetry. In addition, the two-QRPA method is used in [61].

Stating in greater detail: while our model is formulated to describe spherical nuclei, a deformed mean-field is used in Ref. [61], complemented with a schematic spinisospin separable residual interaction that contains two parts, an attractive ph and a repulsive pp , with coupling strengths $\chi^{p h}$ and $\kappa^{p p}$, respectively. By performing a detail calculation of the $\beta \beta_{2 \nu}$-decay of ${ }^{76} \mathrm{Ge}$, it is found that: i) the positive value of $\left|M_{e x p}^{2 \nu}\right|$ is reproduced well for $\kappa_{\downarrow}^{p p}=0.028 \mathrm{MeV}$, which was deduced by Homma et al.[62] from a systematic study of the single $\beta^{+}$-decays, and ii) the negative value of $\left|M_{e x p}^{2 \nu}\right|$ is disfavored due to a complete disagreement with this study, since a large value for $\kappa_{\uparrow}^{p p}(\cong 0.07-0.075 \mathrm{MeV})$ is required, that is
above the critical value, $\kappa^{p p} \equiv \kappa_{c}^{p p} \cong 0.06 \mathrm{MeV}$, where the deformed QRPA collapses [62].

This large difference by a factor of $\sim 3$ between $\kappa_{\uparrow}^{p p}$ and $\kappa_{\downarrow}^{p p}$ should be compared with the small difference of $\sim$ $25 \%$ between $t_{\text {sym }}$ and $t_{\downarrow}$, found here for ${ }^{76} \mathrm{Ge}$. Therefore, we will not necessarily encounter the same difficulties in reproducing simple $\beta$-decay with PSU4SR as faced with $\kappa_{\uparrow}^{p p}$ in Ref. [61]. Our preliminary calculations of the GT $\beta^{-}$-strength confirm this fact, but detailed study is still necessary.

Finally, we compare the relative differences between the $M^{0 \nu}$ obtained with our proposal, i.e., with $t_{\text {sym }}$ and Method IV, and with the usual QRPA calculations, based on $t_{\downarrow}$ and Method II. That is, we evaluate the quantity $\mid M_{s y m}^{0 \nu}(\mathrm{IV})-\mathrm{M}_{\downarrow}^{0 \nu}$ (II) $\mid / \mathrm{M}_{\downarrow}^{0 \nu}$ (II), from where we get the differences of $(40,45,37,58,55,11,36,7) \%$ for $\left({ }^{48} \mathrm{Ca},{ }^{76} \mathrm{Ge}\right.$, ${ }^{82} \mathrm{Se},{ }^{96} \mathrm{Zr},{ }^{100} \mathrm{Mo},{ }^{128} \mathrm{Te},{ }^{130} \mathrm{Te},{ }^{150} \mathrm{Nd}$ ). Therefore, the cases our procedure leads to smaller matrix elements by $\sim 40 \%$ compared with standard evaluations. As seen from Table V this difference basically arises from the one-QRPA-method, employed here instead of the currently used two-QRPA-method. The difference is partially due also to the way of carrying out the restoration of the spin-isospin symmetry.

## V. FINAL REMARKS

This study was motivated by the interest shown recently by several groups [9-13] in the relationship between the restoration of $S U(4)$ symmetry and the $\beta \beta$ decay moments, which was addressed by some of the present authors earlier [14-22]. Therefore, we thought it appropriate to update those studies and stress once again the strong bonding between the residual interaction, the GSC, the PSU4SR and the quenching of the $\beta \beta$-decay NM. All this we do in the framework of the QRPA, for which we have provided a review in Sec. III of different approximations used in the literature. In addition, we make a thorough and updated discussion of $\beta \beta_{2 \nu}$ moments, and consider contributions of the induced weak currents to the $\beta \beta_{0 \nu}$ moments.

From the comparison, done in Tables II and III between the mean field results, described here within the BCS approximation, and the full QRPA calculations using Method IV, it is evident that the residual protonneutron interaction plays a fundamental role in the PSU4SR, not only in the $p p$ channel but also in the $p h$ channel. The results shown in Table V testify that this also occurs within the framework of the commonly used Method II, and by adjusting the isoscalar strength to the measured $\beta \beta_{2 \nu}$ half-life.

As explained in the previous section, Method IV only involves the nuclei in which the process takes place, while Method II also implies the neighboring nuclei through the GSC. In addition, Method IV is simpler, and, like Method III, allows a discussion of the consequences of


FIG. 3: (Color online) $\beta \beta_{0 \nu}$ nuclear moments evaluated with several nuclear structure model calculations: i) QRPA by Tübingen (QRPA Tü) [9] ( $g_{A}=1.27$ ), Jyväskylä (QRPA Jy) [11] $\left(g_{A}=1.26\right)$ groups, and our results from Table III (QRPA Ferr) $\left(g_{A}=1.27\right)$, ii) interacting shell model (ISM) [54] ( $g_{A}=$ 1.25), Large-scale shell model (SM (SDPFMU)) [63], iii) interacting boson model (IBM-2) [64] ( $g_{A}=1.269$ ), vi) energy density functional method (EDF) $[65]\left(g_{A}=1.25\right)$, and covariant density functional theory (CDFT) $[35,66]$ ( $g_{A}=1.254$ ). All results are normalized to $g_{A}^{2}$. The theoretical "errors" presented in our calculations are evaluated as described in the text. The present figure is similar to [35, Fig. 7], [57, Fig. 5][67, Fig. 4], and [68, Fig. 1].

GSC within the SMM, and the calculation of NM by a simple matrix inversion, without resolving the equation QRPA [15]. Because of all of this, we find that Method IV is preferable to Method II.

Moreover, the above statement on the role of the residual interaction is valid for any other QRPA calculation, as well as for the shell-model evaluations of the chargeexchange matrix elements and resonances. In other words, in all cases the residual interaction in this way quenches the $\beta \beta$-decays mean-field results. Therefore, it is not surprising that all theoretical studies, shown in Fig. 3 , yield similar results for $\beta \beta_{0 \nu}$ moments, when compared with the mean field results. However, it is worth noting that our results are lower on average by $\sim 40 \%$. The theoretical uncertainties in Fig. 3 were increased, relative to $3 \%$ used in Table III, in such a way so that they also cover the $M_{\uparrow}^{0 \nu}(\mathrm{IV})$ and $M_{\downarrow}^{0 \nu}(\mathrm{IV})$ results of Table V, obtained via fits to the $\beta \beta_{2 \nu}$ data.

The SMM presented in the Appendix clearly shows that the PSU4SR within the QRPA is manifested through a very strong cancelation effect between the for-
ward and backward going contributions in the particleparticle channel. Within the Quasiparticle TammDancoff Approximation [49] and the Shell Model [63], the equivalent quenching effect is induced by the cancelation between seniority zero and seniority four contributions to the $\beta \beta$-moments.

In short, it can be stated that the central achievement of the present work is the realization that PSU4SR, driven by the residual interaction, is the principal actor in shaping the $\beta \beta$-decays, independently of the underlying nuclear model that is used. Being aware of this fact, we have tried to exploit this relevant property of the residual interaction as much as possible. Perhaps it would not be an exaggeration to say that the differences between different theoretical studies are mainly due to the different ways to restore spin-isospin symmetry.

Strictly speaking, the partial $\mathrm{SU}(4)$ symmetry restoration is present in all QRPA calculations, since all of them involve a residual interaction. The advantages of performing it via the minima of the $\beta^{+}$-strength over the fit to $\left|\mathcal{M}_{\text {exp }}^{2 \nu}\right|$, has been disclosed in point 5) in Sec. I. We add here that the proposed recipe to carry out this restoration is based on physically robust arguments, which makes the theory predictive, producing the $2 \nu$ moments that are of the same order of magnitude as the experimental ones. This is done without resorting to any free parameter, which is a non-minor achievement, when compared with the mean field results which are one order magnitude larger, and is one more reason for preferring our way of setting the isovector $p p$ parameter instead of the standard form. The agreement with $\beta \beta_{2 \nu}$ data is only modest, and it is somewhat disconcerting that the estimates of theoretical uncertainties are greater than the experimental ones. Again, the reason for this is that the $t$ value is in the QRPA breakdown region for $\beta \beta_{2 \nu}$ decay. But this is open to further study, and it is possible that in the not too distant future more precise results will be obtained for $2 \nu$-moments.

Moreover, given the widespread use of the $\left|M_{e x p}^{2 \nu}\right|$ fitting method, based on the justification that $M^{0 \nu}$ and $\mathcal{M}^{2 \nu}$ are similar, it is difficult to say which of the two procedures is preferable. To discern between them, it might be useful to simultaneously analyze the single and double GT decays in the framework of the PSU4SR. A step in this direction was given a long time ago by our group [17], which would now have to be updated in the light of recent developments in that direction, such as Refs. [69-73]. In the same sense, it would be interesting to study the first-forbidden beta transitions [74] and their respective giant resonances $[75,76]$, in the context of $\operatorname{SU}(4)$ symmetry.

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## Appendix: QRPA within the Single Mode Model

In the SMM there is only one intermediate state $\alpha$ and the RPA matrix elements in (3.3) become (see [49, Eqs.(28-31)]):

$$
\begin{align*}
A_{J^{\pi}} & =\omega_{0}+\left[\left(u_{\mathrm{p}}^{2} v_{\mathrm{n}}^{2}+v_{\mathrm{p}}^{2} u_{\mathrm{n}}^{2}\right) \mathrm{F}_{J^{\pi}}(p n)\right. \\
& \left.+\left(u_{\mathrm{p}}^{2} u_{\mathrm{n}}^{2}+v_{\mathrm{p}}^{2} v_{\mathrm{n}}^{2}\right) \mathrm{G}_{J^{\pi}}(p n)\right] \\
B_{J^{\pi}} & =2 v_{\mathrm{p}} u_{\mathrm{n}} u_{\mathrm{p}} v_{\mathrm{n}}\left[\mathrm{~F}_{J^{\pi}}(p n)-\mathrm{G}_{J^{\pi}}(p n)\right] . \tag{A.1}
\end{align*}
$$

where

$$
\begin{equation*}
\omega_{0}=-\frac{1}{4}\left[\mathrm{G}_{0^{+}}(p p)+\mathrm{G}_{0^{+}}(n n)\right] \tag{A.2}
\end{equation*}
$$

is the unperturbed energy, and $\mathrm{G}_{J^{\pi}}(p n)=\mathrm{G}\left(\mathrm{pn}, \mathrm{pn} ; \mathrm{J}^{\pi}\right)$, etc. Moreover, from (3.2) one obtains

$$
\begin{align*}
\omega_{J^{\pi}} & =\sqrt{A_{J^{\pi}}^{2}-B_{J^{\pi}}^{2}} \\
X_{J^{\pi}} & =\frac{A_{J^{\pi}}+\omega_{J^{\pi}}}{\sqrt{\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right)^{2}-B_{J^{\pi}}^{2}}} \\
Y_{J^{\pi}} & =\frac{-B_{J^{\pi}}}{\sqrt{\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right)^{2}-B_{J^{\pi}}^{2}}} \tag{A.3}
\end{align*}
$$

and

$$
\begin{align*}
\left(X_{J^{\pi}}^{2}+Y_{J^{\pi}}^{2}\right) & =\frac{\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right)^{2}+B_{J^{\pi}}^{2}}{\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right)^{2}-B_{J^{\pi}}^{2}} \\
X_{J^{\pi}} Y_{J^{\pi}} & =-\frac{\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right) B_{J^{\pi}}}{\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right)^{2}-B_{J^{\pi}}^{2}} \tag{A.4}
\end{align*}
$$

This yields

$$
\begin{equation*}
\rho_{J^{\pi}}=\frac{\left[\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right)^{2}+B_{J^{\pi}}^{2}\right] u_{\mathrm{p}} v_{\mathrm{n}} v_{\mathrm{p}} u_{\mathrm{n}}-\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right) B_{J^{\pi}}\left(u_{\mathrm{p}}^{2} v_{\mathrm{n}}^{2}+v_{\mathrm{p}}^{2} u_{\mathrm{n}}^{2}\right)}{\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right)^{2}-B_{J^{\pi}}^{2}} . \tag{A.5}
\end{equation*}
$$

Since

$$
\begin{align*}
\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right)^{2}+B_{J^{\pi}}^{2} & =2 A_{J^{\pi}}\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right) \\
\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right)^{2}-B_{J^{\pi}}^{2} & =2 \omega_{J^{\pi}}\left(A_{J^{\pi}}+\omega_{J^{\pi}}\right) \tag{A.6}
\end{align*}
$$

and employing (A.4) we arrive at a very simple expression

$$
\begin{equation*}
\rho_{J^{\pi}}=\rho_{0} \frac{\omega_{0}}{\omega_{J^{\pi}}}\left(1+\frac{\mathrm{G}_{J^{\pi}}}{\omega_{0}}\right) \tag{A.7}
\end{equation*}
$$

where $\rho_{0}=u_{\mathrm{p}} v_{\mathrm{n}} v_{\mathrm{p}} u_{\mathrm{n}}$ is the unperturbed BCS two-body particle-hole density matrix. Therefore, the RPA correlations in the SMM, besides modifying the unperturbed energy $\omega_{0}$ into perturbed energies $\omega_{J \pi}$, they introduce the renormalization factors (effective $\beta \beta$-decay charge)

$$
\begin{equation*}
\mathcal{E}_{J^{\pi}}=\frac{\rho_{J^{\pi}}}{\rho_{0}}=\frac{\omega_{0}}{\omega_{J^{\pi}}}\left(1+\frac{\mathrm{G}_{J^{\pi}}}{\omega_{0}}\right) \tag{A.8}
\end{equation*}
$$

which quench all $\beta \beta_{2 \nu}$, and $\beta \beta_{0 \nu}$ moments. The factor $\left(1+\mathrm{G}_{J \pi} / \omega_{0}\right)$ comes from the interference between the forward and backward going RPA contributions, which are coherent in the $p p$ channel and totally out of phase in the particle-hole ( $p h$ ) channel. As a consequence, the $p h$ matrix elements $\mathrm{F}_{J^{\pi}}(p n)$ do not appear in this factor, with only the $p p$ matrix elements $\mathrm{G}_{J^{\pi}}(p n)$ surviving. It is worth noting that the above result is valid in general, i.e., , for any type of residual interaction, and not only for (1.2).

Moreover, using the same interaction between identical and nonidentical particles for $J^{\pi}=0^{+}$one has: $\mathrm{G}_{0^{+}}(p p)=\mathrm{G}_{0^{+}}(n n)=2 \mathrm{G}_{0^{+}}(p n)$, which implies $\omega_{0}=$ $-\mathrm{G}_{0^{+}}(p p) / 2=-\mathrm{G}_{0^{+}}(p n)$, which is the condition for the restoration of the isospin symmetry. For the force de-
scribed by Eq. (1.2), this condition is expressed as

$$
\begin{equation*}
1+\frac{\mathrm{G}_{0^{+}}(p n)}{\omega_{0}}=1-s \tag{A.9}
\end{equation*}
$$

and for $s=s_{\text {sym }}=1$ is $\mathcal{F}_{0^{+}}=0$. This leads to the condition (see (1.3))

$$
\begin{equation*}
S_{F}^{+}=M_{F}^{2 \nu}=M_{F}^{0 \nu}\left(J^{\pi}=0^{+}\right)=0 \tag{A.10}
\end{equation*}
$$

which is well fulfilled in full calculations as shown in Table I. Therefore, the SMM nicely explains the restoration of the isospin symmetry.

The SMM is also appropriate for explaining the maximal restoration of the $S U(4)$ symmetry. In fact,

$$
\begin{equation*}
1+\frac{\mathrm{G}_{1+}(p n)}{\omega_{0}}=1-t / t_{0} \tag{A.11}
\end{equation*}
$$

and $\mathcal{E}_{1+}=0$ for $t=t_{0}$, the value of which depends on the $p n$ single particle state, and $t_{s y m} \equiv t_{0}$. For instance, the dominant single pair configurations in ${ }^{48} \mathrm{Ca}$ and ${ }^{100}$ Mo are, respectively, $\left[0 f_{7 / 2}(n) 0 f_{7 / 2}(p)\right]_{J^{+}}$and $\left[0 g_{7 / 2}(n) 0 g_{9 / 2}(p)\right]_{J^{+}}$, and the corresponding values of $t_{0}$ are $21 / 11$ and $27 / 20$ (see Ref. [22]). Keep in mind that the restoration of symmetry $S U(4)$ should lead to relations (1.3) and (1.4), but in no way should it be total, since in this case there would be no $\beta \beta$-decay [12].

Finally, it should be stressed that, at variance with the $\beta \beta$-decay moments, the energies $\omega_{J \pi}$ in (A.8), as well as $E_{I A S}$ and $E_{G T R}$ given by Eqs. (3.4) and (3.5), do not behave as the factor $\left(1+\mathrm{G}_{J^{\pi}} / \omega_{0}\right)$, but strongly depend on the $p h$ matrix elements $F_{J^{\pi}}$ (see Eqs. (7) and (8) in Ref. [22]).
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[^0]:    ${ }^{1}$ As a matter of fact, a long time ago Bohr and Mottelson pointed out [5]: "The supermultiplet symmetry has approximate validity for the light nuclei spectra, but it is badly broken in heavier nuclei as a consequence of the strong spin-orbit coupling, which leads to the (jj) coupling. The correlations responsible for the

[^1]:    renormalization effect for the $G T$ moments and for the spinmagnetic moments may be viewed as a trend away from the (jj) coupling scheme toward the LS coupling." Equivalently, it can be stated that the residual interaction "restores" the $S U(4)$ symmetry. See also Refs. [6, 7].

[^2]:    ${ }^{2}$ For the sake of convenience, the standard $F$ and $G T \beta \beta_{0 \nu}$ moments will be labeled, respectively, as $V$ and $A$ moments.

[^3]:    ${ }^{3}$ We use here the angular momentum coupling scheme $\left|\left(\frac{1}{2}, l\right) j\right\rangle$.

[^4]:    ${ }^{4}$ The intermediate $(N-1, Z+1)$ nucleus is now represented as a combination of proton-particle neutron-hole excitations on the initial $(N, Z)$ nucleus, and of proton-hole neutron-particle excitations on the final $(N-2, Z+2)$ nucleus.

[^5]:    ${ }^{5}$ In our numerical calculations $M^{2 \nu}$ is negative at $s=t=0$, as seen from Figs. 1 and 2. See also Eq. 3.30.

[^6]:    ${ }^{6}$ Only for ${ }^{150} \mathrm{Nd}$ we have used the spe evaluated theoretically

