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Beam-Energy Dependence of Charge Balance Functions from Au+Au Collisions at RHIC

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102	Balance functions have been measured in terms of relative pseudorapidity $(\Delta \eta)$ for charged particle
102	pairs at the Relativistic Heavy-Ion Collider (RHIC) from Au+Au collisions at $\sqrt{s_{\rm NN}} = 7.7$ GeV to
104	200 GeV using the STAR detector. These results are compared with balance functions measured
105	at the Large Hadron Collider (LHC) from Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 2.76$ TeV by the ALICE
106	Collaboration. The width of the balance function decreases as the collisions become more central
107	and as the beam energy is increased. In contrast, the widths of the balance functions calculated

ed using shuffled events show little dependence on centrality or beam energy and are larger than the observed widths. Balance function widths calculated using events generated by UrQMD are wider than the measured widths in central collisions and show little centrality dependence. The measured widths of the balance functions in central collisions are consistent with the delayed hadronization of a deconfined quark gluon plasma (QGP). The narrowing of the balance function in central collisions at $\sqrt{s_{\rm NN}}$ = 7.7 GeV implies that a QGP is still being created at this relatively low energy.

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Event-by-event charge correlations and fluctuations 115 can be used as a tool to study the dynamics of hadroniza-116 tion in relativistic heavy-ion collisions [1–33]. One such 117 observable, the balance function [27-30], is sensitive to 118 the correlation of balancing charges. The basic idea of 119 the balance function is that charge is created in balanc-120 ing pairs that originate from the same point in space and 121 time. By means of a like-sign subtraction, the balance 122 123 function can yield the distribution of relative momentum between the balancing charges. Balance functions 124 are sensitive to the mechanisms of charge formation and 125 the subsequent relative diffusion of the balancing charges 126 [27] and are also affected by the freeze-out temperature₁₆₈ 127 and radial flow [28]. Model calculations show that collec-128 tive flow is not sufficient to explain the balance-function 170129 widths measured in central Au+Au collisions at $\sqrt{s_{\rm NN}}_{171}$ 130 = 200 GeV [28, 30, 34]. Balance functions for central¹⁷² 131 collisions have been shown to be consistent with $blast_{173}$ 132 wave models where the balancing charges are required to 133 come from regions with similar collective flow [30]. The $_{_{\rm 175}}$ 134 inferred high degree of correlation in coordinate space has 135 been postulated as a signal for delayed hadronization [27]. 136 In central collisions, a deconfined system of quarks and 137 gluons is created, which cools and expands [35]. Most, 138 of the observed balancing charges are then created when 139 the deconfined system hadronizes, which limits the time $_{^{181}}$ 140 available for the balancing charges to diffuse away from $_{182}$ 141 one another. This leads to tighter correlations in coor-142 dinate space of balancing charges, and due to collective 184143 motion, results in tighter correlations in relative momen-144 tum and relative rapidity. Alternatively, if the charges $_{186}$ 145 are created early (on the order of 1 fm/c), the balancing₁₈₇ 146 charges are less correlated in the final state because the 1_{188} 147 balancing charges have more time to move apart from 148 one another. Thus, a narrow balance function in terms 149 of relative pseudorapidity or relative rapidity in central 150 collisions compared with peripheral collisions implies de-151 layed hadronization. 152 193

The balance function is a conditional distribution $[27]_{,19}^{,19}$ which can be written as:

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$$B(\Delta \eta) = \frac{1}{2} \left\{ \frac{N_{+-}(\Delta \eta) - N_{++}(\Delta \eta)}{N_{+}} \right.$$

$$+ \frac{N_{-+}(\Delta \eta) - N_{--}(\Delta \eta)}{N_{-}} \left. \right\} .$$

$$(1)_{19}^{19}$$

$$(2)$$

The balance function in terms of $\Delta \eta$ ($B(\Delta \eta)$) represents²⁰¹ 155 the probability of seeing a particle which has a relative²⁰² 156 pseudorapidity $\Delta \eta$ with respect to its opposite sign part-²⁰³ 157 ner, given the condition that its opposite sign partner²⁰⁴ 158 has already been seen inside the detector. Specifically,²⁰⁵ 159 $N_{+-}(\Delta \eta)$ is calculated by taking in turn each positive²⁰⁶ 160 particle in an event and incrementing a histogram of²⁰⁷ 161 $\Delta \eta = |\eta(+) - \eta(-)|$ with respect to all negative par-208 162 ticles in that event. $N_{+-}(\Delta \eta)$ is then summed over all₂₀₉ 163 events. A similar procedure is followed for N_{++} , $N_{--,^{210}}$ 164 and N_{-+} . For the denominators, $N_{+(-)}$ is the number₂₁₁ 165 of positive (negative) particles integrated over all events.212 166 The balance function is calculated for all events in a given₂₁₃ 167

$\sqrt{s_{\rm NN}}$ (GeV)	Year	Events (M)
200	2010	32
62.4	2010	15
39	2010	10
27	2011	28
19.6	2011	15
11.5	2010	7.7
7.7	2010	2.2

TABLE I. Summary of the data used in this analysis.

centrality bin at each incident energy.

The system size and centrality dependence of the balance function for all charged particles has been studied by the NA49 Collaboration at $\sqrt{s_{\rm NN}} = 17.3$ GeV for p+p, C+C, Si+Si, and Pb+Pb collisions [36]. The balance function for all charged particles narrows in central Pb+Pb collisions at 17.3 GeV and the widths of the balance functions for p+p, C+C, Si+Si, and Pb+Pb collisions scale with the number of participating nucleons. The NA49 Collaboration has also published results [37] for the rapidity dependence and beam energy dependence of the balance function for all charged particles for Pb+Pb collisions from $\sqrt{s_{\rm NN}} = 6.3$ GeV to 17.3 GeV. The balance function was observed to narrow in central collisions for midrapidity, but did not narrow at forward rapidity. The authors of Ref. [37] showed that the narrowing of the balance function in terms of $\Delta \eta$ in central collisions was explained with the AMPT (a multiphase transport) model [38] incorporating delayed hadronization, while models such as HIJING (heavy-ion jet interaction generator, version 1.38, default parameters) [39] and UrQMD (ultra relativistic quantum molecular dynamics, version 3.3, with default parameters) [40] failed to reproduce the observed narrowing.

The STAR Collaboration has presented a study of the longitudinal scaling of the balance function in Au+Au collisions at $\sqrt{s_{\rm NN}} = 200$ GeV [41]. STAR has published results for balance functions from p+p, d+Au, and Au+Au collisions at $\sqrt{s_{\rm NN}} = 130$ and 200 GeV in terms of $\Delta \eta$, relative rapidity (Δy), relative azimuthal angle $(\Delta \phi)$, and invariant relative momentum (q_{inv}) for all charged particles, for charged pions, and for charged kaons [42, 43]. The balance functions for all charged particles and for charged pions narrow as the events become more central while balance functions calculated using HIJING and UrQMD showed no centrality dependence. The ALICE Collaboration has recently published measurements [44] from Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 2.76$ TeV that also show that the balance functions in terms of $\Delta \eta$ narrow in central collisions.

In this paper, we report measurements of balance functions for all charged particles with $0.2 < p_{\rm T} < 2.0 \text{ GeV}/c$ in terms of relative pseudorapidity ($\Delta \eta$) from Au+Au collisions at $\sqrt{s_{\rm NN}} = 7.7$, 11.5, 19.6, 27, 39, 62.4, and 200 GeV. We observe that the balance functions narrow in central collisions and narrow as the beam energy is



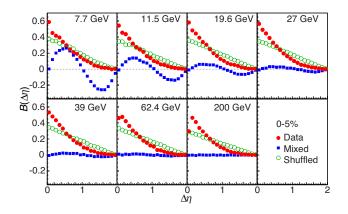


FIG. 1. (Color online) The balance function in terms of $\Delta \eta$ for all charged particles with $0.2 < p_{\rm T} < 2.0 \text{ GeV}/c$ from central Au+Au collisions (0-5%) for $\sqrt{s_{\rm NN}}$ from 7.7 to 200 GeV. The data are the measured balance functions corrected by subtracting balance functions calculated using mixed events. Also shown are balance functions calculated using shuffled events.

increased. We compare with the results from ALICE 214 for Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 2.76$ TeV by restricting 215 STAR's acceptance to $\Delta\eta \leq 1.6$ to match the acceptance 216 of ALICE, correcting for the acceptance of STAR in η , 217 and calculating the width of the balance function over the 218 range $0.1 < \Delta \eta < 1.6$, where the lower limit of 0.1 is cho-219 sen to suppress effects from interpair correlations [e.g., 220 Hanbury-Brown and Twiss (HBT) and final-state inter-221 actions]. We observe that the balance function in terms 222 of $\Delta \eta$ narrows as the beam energy is raised from $\sqrt{s_{\rm NN}} =$ 223 7.7 GeV to 2.76 TeV. When the observed balance func-224 tion widths are scaled by the width observed in the most 225 peripheral bin, the relative widths still decrease as the 226 events become more central and as the beam energy is 227 increased. These results contrast with those presented 228 in Ref. [44], where the scaled balance function widths 229 are shown to be nearly the same at RHIC and LHC en-230 ergies. The present observations are consistent with the 231 concept of delayed hadronization of a deconfined quark 232 gluon plasma (QGP) with the deconfined system having 233 a longer lifetime at the highest energy. The narrowing of 234 the balance function in central collisions at $\sqrt{s_{\rm NN}} = 7.7$ 235 GeV implies that a QGP might still be created at this 236 relatively low energy. 237

The data were taken with the STAR detector [45] dur-238 ing the years 2010 and 2011. Table I shows a summary of 239 the data sets used in this analysis. Au+Au collisions were 240 studied at seven beam energies ranging from 7.7 GeV to₂₅₀ 241 200 GeV. The centrality of each collision was determined₂₅₁ 242 according to the measured charged hadron multiplicity₂₅₂ 243 within the pseudorapidity range $|\eta| < 0.5$. Nine central-253 244 ity bins were used: 0-5% (most central), 5-10%, 10-20%, 254 245 20-30%, 30-40%, 40-50%, 50-60%, 60-70%, and 70-80%255 246 (most peripheral). At each of the seven beam energies,256 247 the average number of participating nucleons, N_{part} , is₂₅₇ 248 calculated for each of the nine centrality bins using a258 249

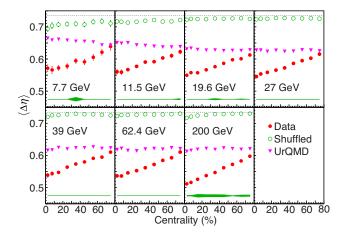


FIG. 2. (Color online) Energy dependence of the balance function widths compared with the widths of the balance functions calculated using shuffled events. Also shown are the balance function widths calculated using UrQMD. The dashed line represents the width of the balance function calculated using shuffled events for a constant $dN/d\eta$ distribution. Error bars represent the statistical error and the shaded bands represent the systematic error.

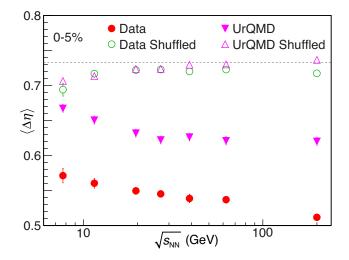


FIG. 3. (Color online) Balance function widths for the most central events (0-5%) compared with balance function widths calculated using shuffled events. Also shown are balance function widths calculated using UrQMD and shuffled UrQMD events. The dashed line represents the width of the balance function calculated using shuffled events for a constant $dN/d\eta$ distribution.

Glauber model. To ensure a more uniform detector acceptance, events were accepted only when the position of the reconstructed primary vertex was within 30 cm of the center of STAR ($|z_{vertex}| < 30$ cm). In addition, the radial position of the primary vertex was required to be less than 2 cm from the center of the beam line to avoid beam pipe events. All events were required to have at least one matched track with the STAR Time-of-Flight (TOF) system [46] to suppress pile-up events.

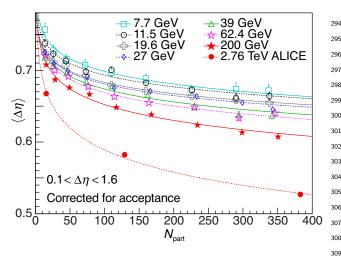


FIG. 4. (Color online) Acceptance-corrected balance function³¹⁰ widths for Au+Au measured over the range $0.1 < \Delta \eta < 1.6_{311}$ compared with similar results from Pb+Pb collisions from₃₁₂ ALICE [44]. Only statistical errors are shown. Lines repre-313 sent fits of the form $a + b(N_{\text{part}})^{0.01}$. 314

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All tracks in the Time Projection Chamber (TPC)³¹⁷ 259 were required to have more than 15 measured space³¹⁸ 260 points along the trajectory. The ratio of the number³¹⁹ 261 of reconstructed space points to possible space points³²⁰ 262 along the track was required to be greater than 0.52 to³²¹ 263 avoid track splitting. Tracks in the TPC were character-322 264 ized by the distance of closest approach (DCA), which is³²³ 265 the smallest distance between the projection of the track³²⁴ 266 and the measured event vertex. To suppress decay ef-325 267 fects and background, all tracks were required to have326 268 a DCA less than 3 cm. A transverse momentum cut³²⁷ 269 of $0.2 < p_{\rm T} < 2.0 \text{ GeV}/c$ and a pseudorapidity cut of³²⁸ 270 $|\eta| < 1.0$ were applied. 329 271

In addition to real data, mixed events and shuffled³³⁰ 272 events were also used in this analysis. Mixed events are³³¹ 273 created by grouping the events according to bins in cen- $^{\scriptscriptstyle 332}$ 274 trality and bins in the position of the reconstructed ver-³³³ 275 tex of the event along the beam direction. Ten centrality³³⁴ 276 bins and five bins in z_{vertex} were used. A set of mixed³³⁵ 277 events is created by taking one track chosen at random³³⁶ 278 from an event, which is selected according to the bin in^{337} 279 centrality and the bin in event vertex position. A mixed³³⁸ 280 event includes no more than one track from any observed³³⁹ 281 event. This mixed-event data set has the same number³⁴⁰ 282 of events with the same multiplicity distribution as the $^{\scriptscriptstyle 341}$ 283 original data set but all correlations are removed. The³⁴² 284 mixed-event subtraction was important, especially at low³⁴³ 285 energies, to account for the effects caused by unbalanced³⁴⁴ 286 positive charges in each event. 287

Shuffled events are produced by randomly shuffling the 288 charges of the particles in each event, which removes the 289 charge correlations while retaining global charge conser-290 vation. Because shuffling uniformly distributes a parti-291 cle's balancing partner across the measured phase space, 292 balance functions calculated using shuffled events can be₃₄₅ Here i is the bin number and $B(\Delta \eta_i)$ is the value of 203

used to gauge the widest balance functions that one can measure within the experimental acceptance of STAR.

Fig. 1 shows the balance functions in terms of $\Delta \eta$ for all charged particles for Au+Au collisions at $\sqrt{s_{\rm NN}} = 7.7$, 11.5, 19.6, 27, 39, 62.4, and 200 GeV for the most central events (0.5%) along with balance functions calculated using shuffled events and balance functions calculated using mixed events. The data shown in the figure are the measured balance functions corrected by subtracting balance functions calculated using mixed events. These data have not been corrected for efficiency or acceptance. The conclusions of this paper involve the width of the balance function in which the efficiency cancels out. The model calculations shown in this paper use the STAR acceptance. When comparisons are made with the width of the balance functions reported by ALICE [44], the STAR data are corrected for acceptance.

At the lower energies, the balance functions calculated using mixed events exhibit an oscillatory distribution that is 0 at $\Delta \eta = 0$, has a positive value at $\Delta \eta = 0.5$, is 0 at $\Delta \eta = 1$, has a negative value at $\Delta \eta = 1.5$ and is 0 again at $\Delta \eta = 2$. This oscillatory behavior lessens as the events become more peripheral and as the beam energy is increased. This effect is due to unbalanced positive charge that is not subtracted by the same sign subtraction inherent in the balance function. The additional positive charges are dominantly protons and have a different $dN/d\eta$ distribution than the negative charges that are dominantly pions. The $dN/d\eta$ distributions for the difference between the positive and negative charges have minima at $\eta = -1$, $\eta = 0$, and $\eta = 1$. Thus, when the balance function in terms of $\Delta \eta$ is calculated for mixed events at the lower energies and in more central collisions, the oscillatory distribution is obtained. At $\sqrt{s_{\rm NN}}$ = 200 GeV, the balance functions calculated using mixed events are zero for all centralities, which indicates that the amount of unbalanced positive charge is small. As the beam energy is decreased, the unbalanced positive charge increases and the balance functions calculated using mixed events become significant.

The corrected balance functions are narrower than the balance functions calculated using shuffled events and the balance functions narrow as the events become more central (see below). Also visible are the effects of interpair correlations [HBT and final-state interactions] that model calculations have shown to be significant for $\Delta \eta \lesssim 0.1$ [29]. Specifically, $B(\Delta \eta)$ for $\Delta \eta < 0.1$ is noticably higher than the trend of the remaining points at 7.7 GeV while $B(\Delta \eta)$ for $\langle \Delta \eta \rangle < 0.1$ is lower than the trend at 200 GeV. The width of the balance function is characterized in terms of a weighted average:

$$\langle \Delta \eta \rangle = \frac{\sum_{i=i_{\text{lower}}}^{i_{\text{upper}}} B(\Delta \eta_i) \Delta \eta_i}{\sum_{i=i_{\text{lower}}}^{i_{\text{upper}}} B(\Delta \eta_i)}.$$
 (2)

the balance function for the relative pseudorapidity bin $\Delta \eta_i$. The weighted average is calculated over a range in $\Delta \eta$ chosen to minimize contributions from HBT and Coulomb effects ($\Delta \eta \geq 0.1$) and maximize the acceptance of STAR ($\Delta \eta \leq 2.0$).

Fig. 2 shows the balance function widths for Au+Au 351 collisions from $\sqrt{s_{\rm NN}} = 7.7 \text{ GeV}$ to 200 GeV for nine cen-352 trality bins. The widths are calculated for $0.1 < \Delta \eta <$ 353 2.0 to remove the distortions caused by interpair corre-354 lations for $\Delta \eta < 0.1$ [29]. The widths of the balance 355 functions calculated using shuffled events are larger than 356 the widths of the balance functions calculated using data. 357 The widths of the balance functions using shuffled events 358 shown in Fig. 2 are close to the value 0.733, which one 359 would expect for shuffled events from a flat $dN/d\eta$ dis-360 tribution over the range $-1 < \eta < 1$. The data show a 361 smoothly decreasing width with increasing beam energy 362 and as the collisions become more central. Fig. 2 also 363 shows the widths of balance functions calculated using 364 UrQMD. The UrQMD calculations are analyzed in the 365 same way as the data with the balance functions calcu-366 lated from mixed UrQMD events being subtracted from 367 the balance functions calculated using UrQMD. For beam 368 energies below 20 GeV, the balance function widths from⁴⁰⁴ 369 UrQMD increase as events become more central whereas⁴⁰⁵ 370 the measured widths decrease. Above 20 GeV, the bal-406 371 ance function widths from UrQMD show little centrality⁴⁰⁷ 372 dependence. In peripheral collisions, the balance func-408 373 tions widths from UrQMD approach the value of the⁴⁰⁹ 374 measured balance function widths. The UrQMD model⁴¹⁰ 375 is a hadronic model that does not have a deconfined⁴¹¹ 376 phase and has little flow. The early hadronization time of⁴¹² 377 the particles calculated using UrQMD combined with the413 378 strong interaction between final state particles causes the414 379 larger balance function widths in central collisions while⁴¹⁵ 380 the balance function widths calculated using UrQMD are⁴¹⁶ 381 close to the measured balance function widths in periph-417 382 418 eral collisions. 383

One source of systematic errors was estimated by 384 studying the difference between the 200 GeV results 385 from three different runs (in 2007, 2010, and 2011) that 386 used different tracking software and incorporated differ-419 387 ent hardware configurations in STAR. A second $source_{420}$ 388 of systematic errors was estimated by varying the DCA₄₂₁ 389 used to select tracks. A third source of systematic er_{422} 390 ror was estimated by varying the range of the z_{vertex} of₄₂₃ 391 events accepted in STAR. The systematic errors in the $_{a_{24}}$ 392 extracted widths are shown as a shaded band in Fig. 2_{425} 393 Note that the systematic error in the width for the $most_{426}$ 394 central bin at all energies was of the same order or $less_{427}$ 395 than the statistical errors. 396 428

Fig. 3 shows the width of the balance function in terms⁴²⁹ of $\Delta \eta$ for central collisions (0-5%) as a function of beam⁴³⁰ smoothly with increasing beam energy. Also shown are⁴³² the widths of the balance function calculated using events⁴³³ generated with the UrQMD model. Although the energy⁴³⁴ trends for the width of the balance function in UrQMD⁴³⁵

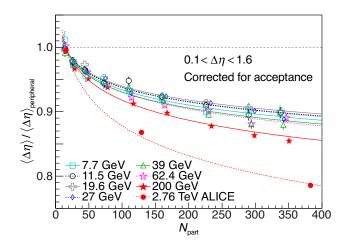


FIG. 5. (Color online) Acceptance-corrected balance function widths for Au+Au measured over the range $0.1 < \Delta \eta < 1.6$ normalized to the most peripheral centrality bin compared with similar results from Pb+Pb collisions from ALICE [44]. Only statistical errors are shown. Lines represent fits of the form $a + b(N_{\text{part}})^{0.01}$.

and data appear similar, the data are much narrower, and as shown in Fig. 2, UrQMD predicts the wrong centrality dependence. The widths of the balance function calculated from shuffled events from both the data and UrQMD are much larger than the widths calculated using the data. The decrease of the shuffled widths at the lower beam energies reflects the fact that the $dN/d\eta$ distributions at the lower beam energies are not completely flat. The fact that the measured balance function widths decrease smoothly with increasing beam energy and are much smaller than the widths predicted by UrQMD is consistent with the idea of delayed hadronization.

To compare with the balance functions measured by ALICE, we correct our measured balance functions for the acceptance of STAR in η using the expression [28]:

$$B_{\Delta\eta_{\max}=2}\left(\Delta\eta\right) = B_{\infty}\left(\Delta\eta\right)\left(1 - \frac{\Delta\eta}{2}\right),\qquad(3)$$

where $B_{\infty}(\Delta \eta)$ is the STAR balance function corrected for acceptance in η assuming that STAR's acceptance is constant for $-1 < \eta < 1$, and $B_{\Delta \eta_{\text{max}}=2}$ is the measured STAR balance function that is not corrected for acceptance in η . For the comparison with the results from ALICE [44], we calculate the widths of the acceptance-corrected balance functions over the range $0.1 \leq \Delta \eta \leq 1.6$ to suppress effects from interpair correlations and to match the acceptance of ALICE in η . Fig. 4 shows these widths as a function of centrality and beam energy for Au+Au collisions. In the same figure we show the width of the balance function from Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 2.76$ GeV calculated for three published centralities [44] over the same range in $\Delta \eta$ using the reported statistical errors. Without knowing the correlations between systematic errors in [44], combining the systematic errors in quadrature appears to lead to a gross

⁴³⁶ overestimate of the uncertainties at 2.76 TeV.

The balance functions narrow as the beam energy is₄₇₁ raised from $\sqrt{s_{\rm NN}} = 7.7$ GeV up to 2.76 TeV and the₄₇₂ balance functions narrow as the collisions become more₄₇₃ central. These observations are consistent with the con-₄₇₄ cept of delayed hadronization.

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The authors of Ref. [44] assert that the relative de-476 442 crease of $\Delta \eta$ with centrality does not change appreciably⁴⁷⁷ 443 with beam energy. To address this point, we calculate₄₇₈ 444 the ratio of the width of the balance function at each cen-479 445 trality to the width of the balance function in the most₄₈₀ 446 peripheral bin at each beam energy, $\langle \Delta \eta \rangle / \langle \Delta \eta \rangle_{\rm peripheral^{481}}$ 447 Because the peripheral bin at the lower energies has low₄₈₂ 448 statistics, we first fit the measured widths at each beam483 449 energy with a function of the form $a + b(N_{\text{part}})^{0.01}$ and 484 450 then take the ratio of the measured widths to the width485 451 of the fitted distribution at the most peripheral centrality₄₈₆ 452 bin. These results are shown in Fig. 5. 487 453

The relative decrease of the balance function width is₄₈₈ 454 much larger at $\sqrt{s_{\rm NN}} = 2.76$ TeV. The relative decrease₄₈₉ 455 then gets smaller as the beam energy is lowered. Thus, we₄₉₀ 456 observe that the relative decrease of the balance function491 457 width clearly changes with beam energy. The difference₄₉₂ 458 between the present analysis and the one presented by the₄₉₃ 459 authors of Ref. [44] is that we calculate the widths over494 460 the range $0.1 \leq \Delta \eta \leq 1.6$ for both experiments, which₄₉₅ 461 minimizes the contributions from interpair correlations.496 462 In contrast, the authors of Ref. [44] calculated the widths497 463 over the range $0.0 \leq \Delta \eta \leq 1.6$ for both the ALICE and 498 464 STAR balance functions. We do not compare with the499 465 results from NA49 [37] here because the acceptance of 500 466 NA49 in η is relatively small. 501 467

468 Model calculations [28, 29] show that a part of the⁵⁰² 469 narrowing of the balance function in central collisions⁵⁰³ is due to radial flow. Thus, the fact that the balance functions in central Pb+Pb collisions at 2.76 TeV are narrower than those in central Au+Au collisions at 200 GeV may be due to an increase in radial flow. One would expect that the balance function in terms of $\Delta \eta$ would be narrower for a longer-lived deconfined QGP, which implies that these results are consistent with the concept of delayed hadronization.

In conclusion, we observe that the balance function in terms of $\Delta \eta$ is narrow in central collisions of Au+Au. At higher beam energies, the balance function in terms of $\Delta \eta$ in central collisions of Pb+Pb is even narrower. This observed narrowing is consistent with the concept of the delayed hadronization of a deconfined QGP produced in these collisions. We observe that the balance functions in Au+Au events at $\sqrt{s_{\rm NN}} = 7.7$ GeV still narrow as the collisions become more central, which suggests that a deconfined QGP might still be produced at this relatively low beam energy.

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