



CHORUS

This is the accepted manuscript made available via CHORUS. The article has been published as:

Reanalysis of radioisotope measurements of the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ cross section

Alan E. Robinson

Phys. Rev. C **94**, 024613 — Published 17 August 2016

DOI: [10.1103/PhysRevC.94.024613](https://doi.org/10.1103/PhysRevC.94.024613)

A reanalysis of radioisotope measurements of the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ cross section

Alan E. Robinson

*Fermi National Accelerator Laboratory, Batavia, Illinois, USA**

The ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ reaction is enhanced by a near threshold $1/2^+$ state. Contradictions between existing measurements of this reaction cross section affect calculations of astrophysical r-process yields, dark matter detector calibrations, and the theory of the nuclear structure of ${}^9\text{Be}$. Select well-documented radioisotope ${}^9\text{Be}(\gamma, n)$ source yield measurements have been reanalyzed, providing a set of high-accuracy independently measured cross sections without the large systematic errors from recent beamline experiments [1, 2]. A single-level Breit-Wigner fit of these corrected measurements yields are $E_R = 1736.8(18)$ keV, $\Gamma_\gamma = 0.742(25)$ eV, and $\Gamma_n = 252(17)$ keV for the $1/2^+$ state, excluding a virtual state solution.

PACS numbers: 26.30.Hj, 25.20.-x, 27.20.+n

The near threshold $1/2^+$ state of the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ reaction is important for several processes in nuclear and astrophysics. This cross section is used to calculate the formation rate of ${}^9\text{Be}$ via the ${}^4\text{He}(\alpha n, \gamma){}^9\text{Be}$ reaction, one of the most important light-element reactions for r-process nucleosynthesis [3]. Neutrons from ${}^{88}\text{Y}/\text{Be}$ and other radioisotope sources using the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ reaction near threshold are being used widely for dark matter detector energy scale and yield calibrations [4, 5] as these low energy neutrons closely mimic dark matter recoils in detectors [6].

Recent experimental [2] and theoretical studies of ${}^9\text{Be}$ suggest that three-cluster dynamics are required to describe its photodissociation. This could imply that the astrophysical ${}^9\text{Be}$ production proceeds via a single step, with a much larger production rate at low energies than that calculated by the two-step process via ${}^8\text{Be}$ [7, 8]. Three-cluster models calculations using the complex scaling method do not find a resonant $1/2^+$ state [9, 10], but a virtual $1/2^+$ state would not be found by these calculations. A virtual state has a complex energy eigenvalue of [11],

$$E_\lambda = (E_R - S_n) + i\Gamma/2 \quad (1)$$
$$|E_\lambda| \begin{cases} < 0 & \text{virtual state} \\ > 0 & \text{resonance state} \end{cases}$$

defined by the real quantities E_R , the resonance energy, S_n , the neutron separation energy, and Γ , the resonance width. These parameters defining the virtual or resonant nature of the $1/2^+$ state can be measured from the position and shape of the near threshold peak of the ${}^9\text{Be}(\gamma, n)$ cross section.

There is little agreement between many measurements of the parameters of this $1/2^+$ state. The yields of radioisotope photoneutron sources provide the simplest method to measure the cross section, a technique that has been used many times since 1935 [12–19]. These measure-

ments use a limited set of (precisely known photon energies and can use homogeneous and isotropic neutron detectors with well understood sensitivity to the low-energy neutrons produced by the reaction. Other measurements using bremsstrahlung photon beams [20, 21], and more recently using inverse Compton photon beams [1, 2, 22], such as the High-Intensity γ Source (HI γ S), provide cross sections over a range of energies inaccessible to photoneutron sources while sacrificing energy resolution and simplicity in experimental design. All these techniques rely on a comparison of absolute quantities. Accurate cross section measurements require knowledge of the absolute photon source strengths, the neutron detection efficiencies, and the photon energy spectra. The energy dependence and systematic uncertainty of the neutron detection efficiencies are improved when simple, homogeneous and isotropic neutron detectors are used. The cross section can also be found using the inelastic scattering of charged particles [23–29]. These measurements require significant background subtraction and extrapolation to low momentum transfer to recover a cross section. A selection of measured and evaluated cross sections in Figure 1 show disagreements of up to 60%, with the strongest disagreements among recent measurements.

Radioisotope measurements of the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ cross section using simple detector designs should have low systematic uncertainties. However, discrepancies among some of these measurements are as severe as those from more complicated experiments at accelerator facilities. Most of the radioisotope measurements were performed in the 1940s, 50s, and 60s, before high precision neutron standards, cross sections, and simulation programs were available. Fortunately, several of these experiments have been well documented with results that are traceable to modern calibrations. This paper will apply corrections to these originally measured cross section values to construct a trusted set of high-accuracy measured radioisotope source neutron yields near the ${}^9\text{Be}(\gamma, n)$ threshold.

* fbfree@fnal.gov

Item	from John and Prosser		Recalculated	
	Correction (%)	Uncertainty (%)	Correction (%)	Uncertainty (%)
1. Gamma-ray source strength				
(a) peak-to-total ratio	—	5.0	+3.88	3.5
(b) subtraction of other γ -ray lines	—	2.0	+0.42	2.0
(c) half-life of ^{124}Sb	—	0.8	-0.8	0.1
2. Neutron source strength				
(a) absolute strength of NBS source	—	3.0	+0.4	0.85
(b) neutron escape from MnSO_4 bath	-0.9	0.3	-0.84	0.05
(c) capture of fast neutrons in bath	-2.8	0.5	-2.34	0.12
(d) neutron counting statistics	—	0.8	—	0.8
3. Finite source size	-0.6	< 0.1	—	—
4. Attenuation of gamma rays in Be shell	+2.5	< 0.1	—	—
5. Neutrons produced by other γ -ray lines ^a	-4.4	1.0	-2.49	—
6. Source geometry effects			+2.65	2.0
Total Correction	-6.2%		1.007	
Overall uncertainty (square root of sum of squares)		6.4%		4.7%

^a Not applied when calculating neutron yields (see text).

Table I. Corrections to the measured $^9\text{Be}(\gamma, n)$ cross section from an $^{124}\text{Sb}/\text{Be}$ source, reproduced from Table I in John and Prosser [18] and extended with recalculated corrections. These corrections reconstruct the measured neutron production cross section compared to that of a 1.69 MeV photon through a thin beryllium target. The finite source size (3) and gamma attenuation (4) effects are combined and reported with other source geometry effects (6) in an MCNPX-Polimi simulation in the reevaluation. See text for descriptions of other recalculated corrections.

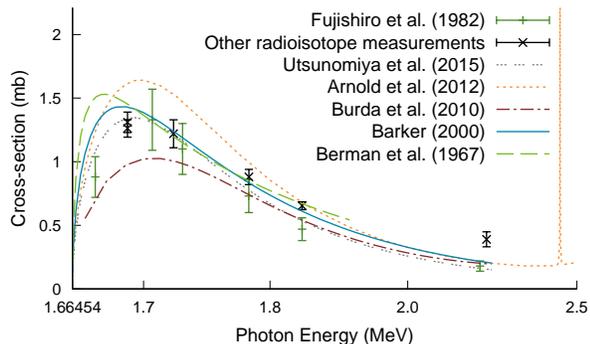


Figure 1. A selection of existing near-threshold cross section measurements of the $^9\text{Be}(\gamma, n)^8\text{Be}$ reaction [1, 2, 15–19, 21, 29]. Lines show the cross section from fitted Breit-Wigner parameters. The parameters from Barker [30] were fit to data from [27].

I. RADIOISOTOPE MEASUREMENTS

A. John and Prosser

John and Prosser [18] used the MnSO_4 bath technique [31] to measure the yield of a $^{124}\text{Sb}/\text{Be}$ source and traced their measured neutron yield to NBS-1, the world’s most precisely calibrated neutron standard. The strength of their ^{124}Sb source was measured using a scintillating NaI crystal.

A list of corrections to the measurement was provided in John and Prosser [18, Table I], and is reproduced with corrections in Table I. The largest correction and uncertainty is to the peak to total ratio of the NaI crystal. John and Prosser cite an unpublished calibration of a a 2” thick 1.75” diameter crystal using radioisotopes with simple decay schemes. Their uncertainty, and an uncertainty common to NaI crystal measurements at the time, was due to their efficiency calibration and its extrapolation in photon energy [32, 33]. Today, photon propagation Monte Carlo methods can accurately predict the peak to Compton ratio of a detector with known dimensions. An MCNP simulation of the detector with a 1.41 MeV threshold gives a ratio of 0.2244 versus the originally used ratio of 0.216. A 3.5% uncertainty ($5\%/\sqrt{2}$) in the photon source strength is retained to account for uncertainties subdominant to those specified by John and Prosser.

In 1962, only the 1.69 MeV and the 2.09 MeV lines were well established in the high energy ^{124}Sb photon spectrum [34]. John and Prosser subtracted from the measured photon yield the contribution from the Compton tail from the 2.09-MeV line. Several percent and sub-percent intensity sub-threshold photon lines have been found since 1962 that lie within the NaI detector resolution of the 1.69-MeV peak [35]. Assuming a NaI(Tl) detector with 5.4% energy resolution at the peak, an additional 0.42% background subtraction is applied. The ^{124}Sb half-life has been revised to 60.20(3) days from 60.4 ± 0.2 days. As the neutron counting significantly pre-

ceded the photon counting, the decay correction was increased.

To calibrate the neutron yield against NBS-1, an intermediate $^{226}\text{Ra}/\text{Be}$ (α, n) source was used. Correction factors for this comparison were recalculated using MCNPX-Polimi simulations, replacing the original analytic corrections made by John and Prosser. Photon neutrons were generated in the simulation by generating photons and simulating the $^9\text{Be}(\gamma, n)^8\text{Be}$ reaction. An MCNP library [36] based on the cross section measured by Arnold *et al.* [1] was used. The $^{226}\text{Ra}/\text{Be}$ neutron spectrum used to generate source particles in the simulation was calculated using the JENDL-AN/05 evaluated $^9\text{Be}(\alpha, n)$ cross section [37] and a modified version of the SOURCES-4C program [38, 39]. Neutron propagation was modelled using the ENDF/B-VII.1 cross section libraries [40]. The ratio of the number of neutrons captured on manganese and the number of neutrons produced was calculated from these simulations and used to compare the $^{226}\text{Ra}/\text{Be}$ and $^{124}\text{Sb}/\text{Be}$ neutron yield measurements. The largest difference in yield is due to the increased probability of neutron capture on sulfur and oxygen for high energy neutrons from the $^{226}\text{Ra}/\text{Be}$ source.

The absolute neutron yield of the NBS-1 neutron source has been revised slightly upward by 0.4% and the yield uncertainty reduced to 0.85% since its 1955 calibration [31, 41].

John and Prosser, and most other neutron yield measurements, present cross sections after correcting for the neutron production from sub-dominant photon energies using other known points of the $^9\text{Be}(\gamma, n)$ cross section, such as with Item 5 in Table I. Radioisotope source neutron yields, the number of neutrons produced per decay, may be alternatively expressed as the sum of the photon branching ratios α_i and cross sections σ_i . This reanalysis will present results in terms of radioisotope neutron yields. For John and Prosser’s $^{124}\text{Sb}/\text{Be}$ measurement, a yield of $\sum_i \alpha_i \sigma_i = 0.678(32)$ mb is found. This expression ignores the contribution of bremsstrahlung photons from high energy β decays, or neutron production from high energy betas that reach the beryllium. The bremsstrahlung and beta contribution to the neutron yield was calculated [42] and found to be negligible for ^{124}Sb and all other isotopes considered by John and Prosser, Gibbons *et al.*, and Snell *et al.*. This and all other reanalyzed yields and yield ratios are given in Table II.

John and Prosser compared their $^{124}\text{Sb}/\text{Be}$ source to $^{28}\text{Al}/\text{Be}$ and $^{206}\text{Bi}/\text{Be}$ sources. As their neutron emission rate was too weak to measure using the MnSO_4 bath method, a ‘Long Counter’ was used for the comparison. MCNPX-Polimi simulations of the Harwell IV Long Counter [43] found that the counter had equal sensitivity (within 1%) to neutrons from each of the three sources. John and Prosser’s 3% correction to the sensitivity of neutrons from $^{28}\text{Al}/\text{Be}$ was reduced to 0.7%. No other corrections to the cross section beyond those in

Measurement	Isotope	Yield or Yield Ratio	Ratio to original
John and Prosser	^{124}Sb	0.678(32) mb	1.079
	$^{28}\text{Al}/^{124}\text{Sb}$	1.311(43)	0.978
	$^{206}\text{Bi}/^{124}\text{Sb}$	0.620(30)	1
Gibbons <i>et al.</i>	^{124}Sb	0.669(29) mb	1.004
	^{88}Y	0.660(29) mb	1.004
Snell <i>et al.</i>	^{24}Na	0.620(19) mb	0.920
	^{72}Ga	0.171(5) mb	0.792
Fujishiro <i>et al.</i>	^{58}Co	0.0083(6) mb	1.83
	^{105}Ru	0.00256(17) mb	1.79
	^{65}Ni	0.0074(5) mb	1.68
	^{28}Al	1.21(9) mb	1.66
	^{88}Y	0.80(8) mb	1.71
	^{38}Cl	< 0.135(14) mb	< 1.69

Table II. Reanalyzed source yields, $\sum_i \alpha_i \sigma_i$, for the given radioisotopes paired with ^9Be . Except for the $^{24}\text{Na}/\text{Be}$ yield, these are used to fit the parameters of the $1/2^+$ state. The values derived from Fujishiro *et al.* [19] shown above are freely scaled in the fit. The yield for ^{38}Cl includes $\leq 10\%$ contributions from inelastic scattering of β -rays on beryllium and bremsstrahlung photons; an upper limit is set for the yield from ^{38}Cl γ -rays.

the original paper were added. The reported cross sections were converted into the neutron yield ratios shown in Table II.

B. Gibbons *et al.*

Gibbons *et al.* [17] measured both $^{124}\text{Sb}/\text{Be}$ and $^{88}\text{Y}/\text{Be}$ neutron sources to high precision using different techniques from those of John and Prosser. The decay rate of their sources were determined using a 4π ionization counter. Their neutron source strengths were measured using a 5-foot diameter graphite moderating sphere and BF_3 thermal neutron detectors. The neutron count rate was calibrated against a source traceable to NBS-1.

Gibbons *et al.* [17, Table I] had calculated several corrections to the their source yield. These corrections were recalculated, and found to be almost entirely in agreement with the original analysis. Only the source strength of the NBS neutron source was changed, increasing the yield by 0.4% and reducing its uncertainty to 0.85%.

C. Snell *et al.*

Snell *et al.* [15] measured the neutron yields from beryllium and deuterium targets using ^{72}Ga and ^{24}Na radioisotopes. Both isotopes produce photons above the 2.2 MeV deuterium dissociation threshold. The absolute neutron yields and the ratio of beryllium to deuterium neutron yields were measured for each radioisotope. The

measured ratios can be compared to the well known modern cross section for the photodissociation of deuterium [40]. Both the reported absolute and the relative yields were reanalysed and found to be consistent within their final 5% and 3% uncertainties [39]. The yield measurement relative to the deuterium target is used in this reanalysis.

Snell *et al.* measured neutrons by sampling epithermal neutrons in large volume of paraffin moderator using indium foil sandwiched between two cadmium foils. The activated indium was counted using a thin-walled Geiger counter. The neutron energy dependence of the neutron yield measurements of Snell *et al.* has been validated using MCNPX-Polimi [44]. One additional correction, for the thermalization of neutrons reentering the deuterium or beryllium from the surrounding moderator, leads to an small additional loss of efficiency. The ratio of the detection efficiency for each source was recalculated. The original uncertainty in the relative source activity is retained, and it dominates the total uncertainty of 3%.

D. Fujishiro *et al.*

The measurements of Fujishiro *et al.* [19] are the most recently published radioisotope measurements of the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ and have heavily influenced modern evaluations of the cross section [29, 45, 46]. Fujishiro *et al.* used a reactor to irradiate a large variety of short-lived radioisotopes that produce photons with energies near the $1/2^+$ state, shown in Table III.

In contrast to the other radioisotope measurements noted above, Fujishiro *et al.* used a non-homogeneous neutron detector: a ring of four BF_3 detectors embedded in a paraffin cylinder. The neutron energy dependence of this detector was originally calculated using a one-dimensional Monte Carlo code and normalized to the flux of a ${}^{24}\text{Na}\text{-D}_2\text{O}$ neutron source. The relative neutron detection efficiency was recalculated using MCNPX-Polimi simulations of the three-dimensional geometry given by Fujishiro *et al.*. The sensitivity, but not the model number, of the BF_3 counters used were given. The new simulations use LND model 2025 BF_3 detectors filled to 500 torr for the 5.0 cps/nv thermal neutron detectors [55]. The energy dependence of the neutron efficiencies calculated by MCNPX-Polimi match that found by Fujishiro *et al.*, but the new calculation finds an absolute efficiency that is much lower. The calculated sensitivity to ${}^{24}\text{Na}$ /deuterium neutrons is 6.2% versus the $9.2 \pm 0.7\%$ measured and used as a calibration by Fujishiro *et al.*

The cross section calculation of Fujishiro *et al.* was reproduced using the original neutron detection efficiency, background subtracted neutron count rates, source intensities, beryllium geometry, and calculation method presented in their paper. Assuming a beryllium density of 1.85 g/cm^3 , the recalculated cross sections are 12% larger than those presented by Fujishiro *et al.*. Given the normalization discrepancies in this calculation and the neu-

tron detection efficiency, only the ratios of the neutron yields measured by Fujishiro *et al.* will be used.

The neutron yields are calculated using the MCNP calculated detector efficiency, and the neutron count rate, source intensities, and uncertainties given by Fujishiro *et al.*. Fujishiro *et al.* included a 6% uncertainty in the absolute normalization of their source intensities that is excluded when calculating relative source intensities in this reanalysis. As the source intensities had been measured after they had decayed to $\sim 10^5 \text{ Bq}$, a correction is made to account for updated half-life values. The source intensities are also corrected to account for updated values of the branching fraction of the photon energy used to measure the source intensities. These corrections are shown in Table III.

Additional contributions to the neutron yield from high-energy β -rays and their bremsstrahlung photons were calculated. The contributions are negligible ($< 0.1\%$ of the total yield) for all measured radioisotopes except ${}^{38}\text{Cl}$. The majority of ${}^{38}\text{Cl}$ decays produce a β with a 4.9 MeV endpoint energy. Depending on the geometry of the source and encapsulation used by Fujishiro *et al.*, $\sim 10\%$ of the measured neutron flux from their ${}^{38}\text{Cl}/\text{Be}$ source could be caused by these high energy electrons. A similarly unaccounted high-energy β contribution had been incorrectly claimed by Fujishiro *et al.* [56] as evidence for the three-body breakup reaction of ${}^9\text{Be}$ below the two-body threshold of $S_n = 1664.54 \text{ MeV}$ [57]. No correction is applied for these additional neutron production mechanisms. Instead, the measured ${}^{38}\text{Cl}$ neutron yield is used an upper limit for the yield produced by the ${}^{38}\text{Cl}$ γ emissions.

E. Other Radioisotope Measurements

Other radioisotope measurements of the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ cross section exist that are not used in this reanalysis. These measurements were either not calibrated against neutron standards, or insufficient information about the experiments were available to correct their measured yields.

Hamermesh and Kimball [16] measured the neutron yield from a ${}^{144}\text{Pr}/\text{Be}$ source with its dominant emission at 2.185 MeV. They used an unspecified NaI detector to measure the photon yield, and compared the strength of the 1.49-MeV γ from ${}^{144}\text{Pr}$ against the 1.38-MeV γ from ${}^{24}\text{Na}$ without correcting for the difference in detector efficiency between the two energies. The change in detection efficiency depends by $\mathcal{O}(10\%)$ on the size and geometry of crystal used. Their measurement is consistent with the ${}^{72}\text{Ga}$ measurement of Snell *et al.* and ${}^{38}\text{Cl}$ measurement of Fujishiro *et al.* within this large uncertainty.

Russell *et al.* [14] measured the neutron yield of a set of six radioisotopes paired with beryllium and a subset of four radioisotopes paired with deuterium. Their measurement of the neutron flux used a 'long counter' calibrated using a ${}^{226}\text{Ra}\text{-Be}(\alpha, n)$ neutron source. This

Isotope	n count rate (s^{-1})	Intensity (MBq)	n Det. Eff. (%)	Branching Fraction Correction	Half-life		Half-life Correction
					Fujishiro et al.	Revised	
^{58}Co	2.38(2)	34.2(32)	6.87	$0.9945(1) \div 0.9944$	70.8 d	70.86(6) d	0.995(5)
^{105}Ru	0.403(10)	18.8(16)	6.79	$0.473(5) \div 0.481$	4.44 h	4.44(2) h	1.00(2)
^{65}Ni	1.150(13)	19.2(17)	6.77	$23.59(14) \div 23.58$	2.52 h	2.51719(26) h	1.0058(5)
^{28}Al	62.0(4)	6.2(6)	6.69	1	2.24 min	2.245(2) min	0.991(4)
^{88}Y	7.01(4)	1.12(13)	6.55	$0.937(3) \div 0.9136$	106.6 d	106.627(21) d	0.9994(5)
^{38}Cl	7.81(4)	8.1(10)	6.09	$33.3(7) \div 31.03$	37.14 min	37.230(14) min	0.989(2)

Table III. Radioisotope sources of > 1.67 MeV photons used in Fujishiro *et al.* [19]. The neutron count rate, neutron detection efficiency (ϵ_n), photon source intensity, and source intensity corrections are shown. Half-lives and branching fractions are from relevant Nuclear Data Sheets [47–52] and from recent experiments [53, 54].

calibration source produces neutrons with far greater energies than the radioisotope sources and is used with a detector design whose sensitivity varies with neutron energy. Their photon source activities were calculated from thermal neutron capture cross sections and the measured neutron flux of the nuclear reactors in which they were activated. While improvements in the analysis of this measurement are possible, very large systematic uncertainties in both the neutron and photon yields would remain, and little information would be gained from a reanalysis.

Earlier measurement of $^9\text{Be}(\gamma, n)^8\text{Be}$ source yields [12, 13, 58] lack detailed information or the precision required for reanalysis.

II. FIT

The $^9\text{Be}(\gamma, n)^8\text{Be}$ cross section was fit using least squares to a sum of isolated Breit-Wigner states. The $1/2^+$ state was fit to the form

$$\sigma_{1/2^+}(E_\gamma) = \frac{\pi(\hbar c)^2}{4E_\gamma^2} \frac{\Gamma_n \sqrt{\frac{E_\gamma - S_n}{E_R - S_n}} \Gamma_\gamma \left(\frac{E_\gamma}{E_R}\right)^3}{(E_\gamma - E_R)^2 + \frac{E_\gamma - S_n}{E_R - S_n} (\Gamma_n/2)^2} \quad (2)$$

where the neutron separation energy $S_n = 1664.54$ keV, Γ_n and Γ_γ are the neutron and photon channel widths, E_R is the resonance energy, and E_γ is the incoming photon energy.

The contribution to the cross section from other higher energy resonances may affect the measured width of the $1/2^+$ resonance. Arnold *et al.* [1] provides the best direct measurement of the shape of the $^9\text{Be}(\gamma, n)$ cross section between 2 MeV and 4 MeV. The contribution from higher energy resonances was fixed using their resonance parameterization shown in Table IV. The widths of the fitted $3/2^+$ and $3/2^-$ resonances given by Arnold *et al.* were multiplied by 1.57 in order to match *their reported experimental data* (see Table IV caption). Ten percent level discrepancies exist in the neutron detector efficiency calibration of Arnold *et al.* [63], and between their measurement and radioisotope measurements at low energy

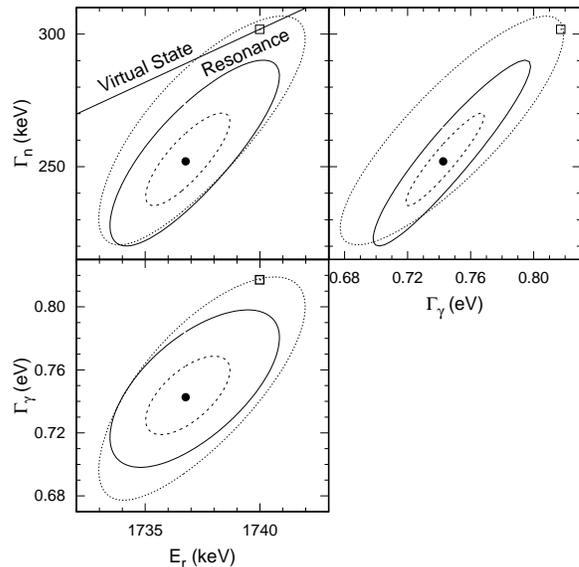


Figure 2. Fit Breit-Wigner parameters to the $1/2^+$ state. Best fit parameters (solid circles) are $E_R = 1736.8(18)$ keV, $\Gamma_\gamma = 0.742(25)$ eV, and $\Gamma_n = 252(17)$ keV with a $\chi^2/\text{d.o.f.}$ of 4.70/7. 68% (dashed) and 95% (solid) confidence regions are shown. A fit excluding the data from Fujishiro *et al.* [19] was also performed, with the 95% confidence region shown by dotted lines. A virtual state would exist if $\Gamma \approx \Gamma_n > E_R - S_n$ (see Equation 1). The best fit virtual state solution for the fit including all reanalyzed radioisotope data is shown (open squares) and is excluded with 99.3% confidence (2.7σ). If the data from Fujishiro *et al.* is excluded, a virtual state is disfavored with 93% confidence (1.8σ). A script to calculate these likelihoods is available in the supplemental materials.

shown in Figure 4. A smaller $5\% \pm 3\%$ discrepancy exists between the $^{24}\text{Na}/\text{Be}$ source yield measured by Snell *et al.* [15] and the adopted contribution from high energy resonances. Given these discrepancies, a 10% scale uncertainty, larger than the 4.6% systematic uncertainty given by Arnold *et al.*, was applied to the summed contribution from the higher energy resonances.

The photon energies, branching ratios, and reanalyzed measured neutron yield for each radioisotope measure-

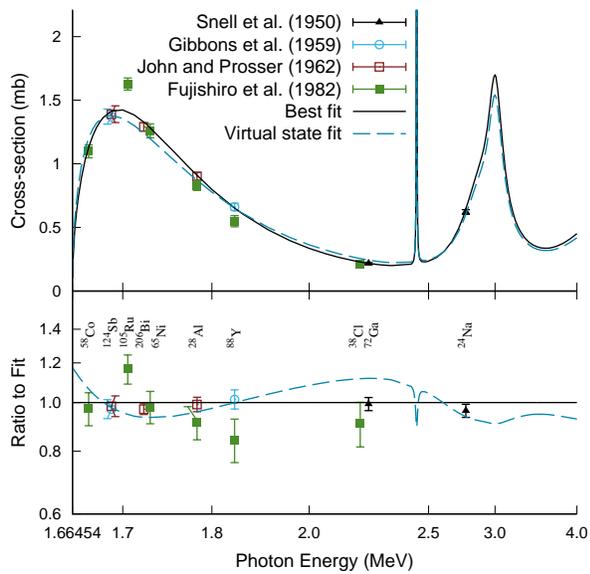


Figure 3. Reanalyzed cross sections from radioisotope measurements of the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ cross section. Only the cross sections for the highest intensity photon energy of each radioisotope are shown, assuming that ratios to the cross sections at other photon energies equal the best fit ratios. The absolute yield of the measurements by Fujishiro *et al.* [19] are floated. The 2.7 MeV measurement by Snell *et al.* [15] of the ${}^{24}\text{Na}/\text{Be}$ source yield was not included in the fit, but it is used to validate the contribution of higher energy resonances as measured by Arnold *et al.* [1].

J^π	E_R (keV)	Γ_γ (eV)	Γ_n (keV)	β_j to ${}^9\text{Be}(\text{g.s.})$ (%)
$5/2^-$	2431	0.098	0.77	6
$1/2^+$	2880	1.8	393	100
$5/2^+$	3008	0.45	168	70
$3/2^+$	4704	7.8	2419	38
$3/2^-$	5590	15.7	1477	38

Table IV. Resonance parameters characterizing the contribution of high-energy resonances to the ${}^9\text{Be}(\gamma, n)$ cross section measured by Arnold *et al.* [1]. These resonances are subtracted from the radioisotope data in order to fit the parameters of the near threshold $1/2^+$ state. Note that the table values are used as a parameterization, and may not be physical. In particular, the relatively narrow $3/2^-$ and $3/2^+$ resonances [59] are being used to fit out the cluster dipole resonance discussed in Section III. The widths of $3/2^+$ and $3/2^-$ resonances are multiplied by 1.57 from the values given by Arnold *et al.* in order to match their experimental data points. A 10% scale uncertainty on the summed cross section from these resonances is applied when fitting the parameters of the $1/2^+$ state. The ground state branching ratios were selected to approximately match the observed ratio of thermal neutrons detected in two detector rings used by Arnold *et al.*. Uncertainties in the ground state branching ratios are not provided and may be large.

ment, shown in Table II, were input to the fit. The yields measured by Fujishiro *et al.* [19] were allowed to float freely. The fit marginalized over the 10% uncertainty in the contribution from the higher energy resonances, and over the 0.85% uncertainty in the NBS-1 neutron source strength. The χ^2 contribution of the ${}^{206}\text{Bi}/\text{Be}$ and ${}^{28}\text{Al}/\text{Be}$ yield measurements of John and Prosser [18] were calculated using their ratio to the ${}^{124}\text{Sb}/\text{Be}$ source yield. The absolute source yields were used for all other measurements. The ${}^{24}\text{Na}/\text{Be}$ source yield from Snell *et al.* [15] was excluded as its energy is well above the $1/2^+$ state. Results from the fit are shown in Figures 2 and 3. At the best fit point, the measurements of Fujishiro *et al.* and the higher resonance contribution are scaled down by 31.7%, and 1.1% respectively.

A virtual state is disfavored by the fit with 99.5% confidence. As the measurements of Fujishiro *et al.* have large systematic uncertainties, require rescaling to match other radioisotope measurements, and have residuals in Figure 3 that hint at potential bias, a second fit was performed excluding their data. This second fit still disfavors a virtual state interpretation, but only with 95% confidence.

Recent measurements [2] and theoretical studies [10] suggest interference between positive parity states and a cluster dipole resonance near $E_\gamma = 8$ MeV. The effect of the interference should be small given the wide separation of these resonance energies, but even a small effect may reduce confidence in the exclusion of a virtual $1/2^+$ state. Interference effects are neglected in this analysis.

For use in dark matter detector calibrations, the yields and uncertainties of individual commercially available photoneutron sources are given in Table V.

III. COMPARISON TO PHOTON BEAM EXPERIMENTS

Figure 4 and Table VI compare recent beamline measurements of the ${}^9\text{Be}(\gamma, n)$ reaction to the present reanalysis of radioisotope measurements. Except for the measurements of Fujishiro *et al.* and the ${}^{28}\text{Al}/\text{Be}$ and ${}^{206}\text{Bi}/\text{Be}$ measurements of John and Prosser, all of the reanalyzed photoneutron measurements used homogeneous and isotropic neutron detectors with well modelled and slowly varying sensitivities for low-energy neutrons. Such detector designs are not suitable for use in a beamline environment.

Arnold *et al.* [1] and Utsunomiya *et al.* [2] detected neutrons using hollow cylindrical moderators with thermal neutron detectors positioned in concentric rings within the moderator. The ratio of the neutron detection rate in each ring allows the measurement of a weighed average neutron energy. Arnold *et al.* [63] calibrated their detector using several well known neutron production processes, and found a 4% discrepancy between the simulated and measured ratio of the count rate in the inner versus the outer ring, as shown in their Figure 13. Utsunomiya *et al.* claim to have measured a significant num-

Isotope	$t_{1/2}$	E_γ (keV)	α_i (%)	$E_n^{c.m.}$ to ${}^8\text{Be}$ g.s. (keV)	Neutron yield ($\sum_i \alpha_i \sigma_i$) (mb)	${}^8\text{Be}$ g.s. branching (%)
${}^{58}\text{Co}$	70.86(6) d	1674.73	0.517(10)	9.05	0.00584(24)	100.0
${}^{124}\text{Sb}$	60.20(3) d	1690.97	47.57(18)	23.47	0.672(18)	100.0
		2090.93	5.49(3)	378.68	0.0145(6)	99.5
		Others			0.00254(4)	99.6
		Total			0.687(18)	100.0
${}^{105}\text{Ru}$	4.44(2) h	1698.17	0.0766(9)	29.86	0.00109(3)	100.0
		1721.15	0.0299(3)	50.27	$3.93(14) \times 10^{-4}$	100.0
		Others			$1.6(5) \times 10^{-5}$	100.0
		Total			0.00156(5)	100.0
${}^{206}\text{Bi}$	6.243(3) d	1718.7	31.9(5)	48.10	0.424(16)	100.0
		1878.65	2.01(4)	190.15	0.01062(26)	99.9
		Others			0.00580(22)	99.4
		Total			0.441(16)	100.0
${}^{65}\text{Ni}$	2.51719(26) h	1724.92	0.399(12)	53.62	0.00513(23)	100.0
${}^{28}\text{Al}$	2.245(2) min	1778.99	100	101.64	0.909(19)	100.0
${}^{88}\text{Y}$	106.627(21) d	1836.06	99.2(3)	152.33	0.646(7)	100.0
		2734.0	0.71(7)	949.8	0.0042(5)	95.3
		3219.7	0.0070(20)	1381.1	$3.6(11) \times 10^{-5}$	81.5
		Total			0.651(7)	99.9
${}^{38}\text{Cl}$	37.230(14) min	2167.40	44.4(9)	446.58	0.101(5)	99.1
		Others			$1.38(27) \times 10^{-4}$	63.2
		Total			0.102(5)	99.0
${}^{72}\text{Ga}$	14.10(1) h	1862.00	5.410(18)	175.36	0.0309(4)	99.9
		2201.59	26.87(12)	476.95	0.0585(25)	98.9
		2491.03	7.73(3)	734.00	0.0179(8)	94.8
		2507.72	13.33(6)	748.82	0.0318(14)	95.3
		Others			0.033(4)	97.7
Total			0.172(6)	97.8		
${}^{207}\text{Bi}$	31.55(4) y	1770.23	6.87(3)	93.86	0.0660(16)	100.0
${}^{226}\text{Ra}$	1600(7) y	1729.60	2.878(8)	57.78	0.0360(12)	100.0
		1764.49	15.30(3)	88.77	0.153(4)	100.0
		1847.43	2.025(9)	162.42	0.01244(14)	100.0
		2204.06	4.924(18)	479.14	0.0107(5)	98.9
		Others	9.69		0.0183(4)	94.4
Total			0.230(5)	99.5		

Table V. Total neutron yields for reanalyzed and commercially available radioisotopes used in near threshold ${}^9\text{Be}$ photoneutron sources. For each major photon energy, the branching ratio to the ${}^8\text{Be}$ ground state assuming the resonance parameters given in Table IV, and the energy of the outgoing neutron for that branch in the center of mass frame are given. The neutron yield is the product of the ${}^9\text{Be}(\gamma, n)$ cross section σ_i and the photon branching fraction α_i for each photon energy E_γ . Photon yields, energies and half-lives, including those of decay chain daughter isotopes, are from [35, 47, 51, 60–62]. Neutrons with other final states, such as $\alpha + \alpha + n$ or ${}^8\text{Be}(3030 \text{ keV}) + n$ will be produced at lower average energy and greater energy spread than those for the ground state transition. Neutron yield uncertainties are propagated from the fit to the $1/2^+$ state and from photon branching ratios.

ber of neutrons from the ${}^9\text{Be}(\gamma, n)$ reaction at energies well below those expected from the ${}^9\text{Be}(\gamma, n){}^8\text{Be}(\text{g.s.})$ reaction channel. The calibration of their neutron energy scale depends on simulations that may have a similar discrepancy to that found by Arnold *et al.*. Their claim likely would have been reported by Arnold *et al.* [1], and may also be inconsistent with the neutron energy spectrum measurements of ${}^{116m}\text{In}/\text{Be}$ and other photoneutron sources using proton recoil spectrometers [64]. These inconsistencies demonstrate the importance of the

neutron yield measurements from well-modelled detectors used in the present analysis.

Table VI shows a comparison of the position, height, and width of the excitation function peak of the $1/2^+$ resonance for radioisotope, inverse Compton, and electron scattering experiments. The radioisotope experiments find a peak position that agrees well with the measurements of Arnold *et al.* and Utsunomiya *et al.*. The different resonance parameters between these experiments is due to differences in the measured shape of the peak

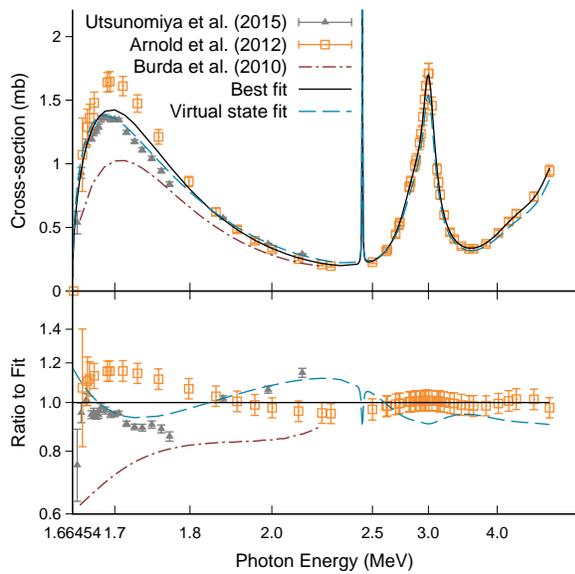


Figure 4. Comparison of the fitted ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ cross section to recent inverse Compton scattered photon beam [1, 2] and electron scattering [29] measurements (see text).

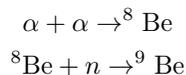
Reference	E_R (keV)	$E_{\gamma, \text{max}}$ (keV)	σ_{max} (mb)	FWHM (keV)
Present	1736.8(18)	1696.8	1.43	152
Arnold <i>et al.</i> [1]	1731(2)	1698.3	1.64	138
Utsunomiya <i>et al.</i> [2]	1728(1)	1694.3	1.35	161
Burda <i>et al.</i> [29]	1748(6)	1705.9	1.03	177

Table VI. The ${}^9\text{Be}(\gamma, n)$ excitation function peak energy, cross section, and width near the $1/2+$ resonance, and the resonance energy for recent experiments. Note that the position of the excitation function peak and the resonance position are not simply related for this near threshold resonance. Charged particle scattering experiments including Burda *et al.* [29] fail to accurately reproduce the cross section measured by direct experiments. The peak positions and areas measured by inverse Compton scatter photon beam experiments [1, 2] matches the present analysis, while the peak widths and heights differ significantly.

that may be caused by miscalibrated neutron detectors.

IV. ASTROPHYSICAL PRODUCTION

The ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ cross section may be used to calculate the dominant two step contribution to the astrophysical production rate of the $\alpha(\alpha n, \gamma){}^9\text{Be}$ reaction. The reaction proceeds via



with both ${}^8\text{Be}$ and ${}^9\text{Be}$ in their ground state. A direct three-body reaction is possible and may dominate the reaction rate at low temperatures [8], but high-sensitivity

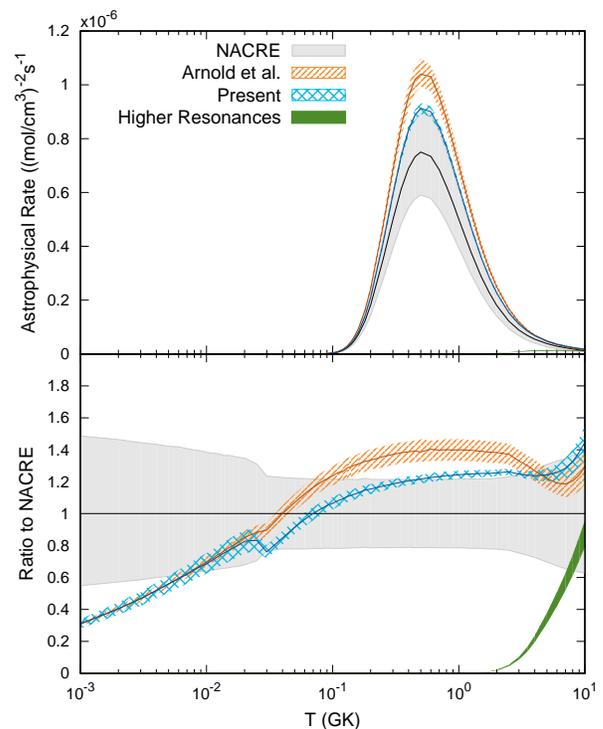


Figure 5. Calculated rate from the two step ${}^4\text{He}(\alpha n, \gamma){}^9\text{Be}$ fusion reaction compared to previous evaluations [1, 46]. Bands show $\pm 1\sigma$ uncertainties.

searches for this process have not demonstrated its existence [65]. The thermal production rate by the two-step process is a double integral over the energy of two Maxwell-Boltzman distributed collision velocities, the cross sections for the two subprocesses, and the mean lifetime of the ${}^8\text{Be}$ nucleus, where the ${}^8\text{Be}$ nucleus may be produced off-shell [46].

The ${}^8\text{Be}(n, \gamma){}^9\text{Be}$ reaction is related by the reciprocity theorem to the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ cross section, as described by Arnold *et al.* [1]. The cross section for elastic $\alpha + \alpha$ scattering via the ${}^8\text{Be}$ compound state is taken from [66], with resonance width $\Gamma_{\alpha\alpha} = 5.57 \pm 0.25$ eV and resonant energy $E_r = 92.03$ keV.

Table VII shows the adopted ${}^9\text{Be}$ astrophysical production rate while Figure 5 shows this rate in comparison with the calculations of NACRE [46] and Arnold *et al.* [1]. At low temperatures, the rate follows that of Arnold *et al.*. The low temperature rate differs from that of NACRE by a factor of $\sqrt{E/92.03 \text{ keV}}$ in the integrand, where E is the center of mass energy of the α particles in the first reaction step. At astrophysically relevant temperatures of 1 GK to 5 GK, the new rate lies largely between those from NACRE and Arnold *et al.*. At high energies, the contribution from higher energy resonances is increased compared to that calculated by Arnold *et al.* because of increases to the resonance widths (see Table IV caption).

T (GK)	Rate	T (GK)	Rate	T (GK)	Rate	T (GK)	Rate
0.001	$1.21(11) \times 10^{-59}$	0.016	$1.04(9) \times 10^{-24}$	0.14	$3.98(12) \times 10^{-8}$	1	$6.22(13) \times 10^{-7}$
0.002	$1.01(9) \times 10^{-47}$	0.018	$8.44(71) \times 10^{-24}$	0.15	$5.87(18) \times 10^{-8}$	1.25	$4.72(9) \times 10^{-7}$
0.003	$6.31(56) \times 10^{-42}$	0.02	$5.20(43) \times 10^{-23}$	0.16	$8.20(24) \times 10^{-8}$	1.5	$3.62(6) \times 10^{-7}$
0.004	$2.89(25) \times 10^{-38}$	0.025	$2.49(19) \times 10^{-21}$	0.18	$1.40(4) \times 10^{-7}$	1.75	$2.83(5) \times 10^{-7}$
0.005	$1.18(10) \times 10^{-35}$	0.03	$3.84(21) \times 10^{-19}$	0.2	$2.10(6) \times 10^{-7}$	2	$2.259(35) \times 10^{-7}$
0.006	$1.16(10) \times 10^{-33}$	0.04	$1.61(8) \times 10^{-15}$	0.25	$4.08(10) \times 10^{-7}$	2.5	$1.527(22) \times 10^{-7}$
0.007	$4.58(40) \times 10^{-32}$	0.05	$2.36(11) \times 10^{-13}$	0.3	$5.92(14) \times 10^{-7}$	3	$1.107(16) \times 10^{-7}$
0.008	$9.58(83) \times 10^{-31}$	0.06	$6.20(28) \times 10^{-12}$	0.35	$7.35(17) \times 10^{-7}$	3.5	$8.47(15) \times 10^{-8}$
0.009	$1.26(11) \times 10^{-29}$	0.07	$6.15(26) \times 10^{-11}$	0.4	$8.32(19) \times 10^{-7}$	4	$6.78(14) \times 10^{-8}$
0.01	$1.16(10) \times 10^{-28}$	0.08	$3.33(13) \times 10^{-10}$	0.45	$8.88(20) \times 10^{-7}$	5	$4.76(15) \times 10^{-8}$
0.011	$8.19(71) \times 10^{-28}$	0.09	$1.21(5) \times 10^{-9}$	0.5	$9.13(20) \times 10^{-7}$	6	$3.62(15) \times 10^{-8}$
0.012	$4.63(40) \times 10^{-27}$	0.1	$3.32(12) \times 10^{-9}$	0.6	$9.00(20) \times 10^{-7}$	7	$2.90(14) \times 10^{-8}$
0.013	$2.19(19) \times 10^{-26}$	0.11	$7.48(26) \times 10^{-9}$	0.7	$8.43(18) \times 10^{-7}$	8	$2.40(13) \times 10^{-8}$
0.014	$8.92(76) \times 10^{-26}$	0.12	$1.45(5) \times 10^{-8}$	0.8	$7.70(16) \times 10^{-7}$	9	$2.05(13) \times 10^{-8}$
0.015	$3.21(27) \times 10^{-25}$	0.13	$2.51(8) \times 10^{-8}$	0.9	$6.94(14) \times 10^{-7}$	10	$1.78(12) \times 10^{-8}$

Table VII. Calculated rate from the two step ${}^4\text{He}(\alpha n, \gamma){}^9\text{Be}$ fusion reaction versus temperature. By convention, rates are in units of $(\text{mol}/\text{cm}^3)^{-2}\text{s}^{-1}$ [46].

V. SUMMARY

Radioisotope source yield measurements of the near threshold $1/2^+$ state of the ${}^9\text{Be}(\gamma, n){}^8\text{Be}$ reaction cross section have been reanalyzed. After reanalysis, these measurements are self-consistent and provide precise experimental bounds of the Breit-Wigner parameters, with absolute cross section uncertainties under 2%. Twenty percent level inconsistencies between the results of Arnold *et al.* [1], Utsonomiya *et al.* [2], and this reanalysis are likely due to systematic errors in the low-energy neutron absolute detection efficiency of the two beamline experiments. The best fit to the cross section indicates that the $1/2^+$ state is a resonance.

ACKNOWLEDGMENTS

This work was derived from my thesis under the guidance of Juan Collar at the University of Chicago. I thank Juan Collar, Charles Arnold and Hiroaki Utsonomiya for discussions about recent measurements of this cross section. This work was completed with the support of the Fermi Research Alliance, LLC under Contract No. De-AC0207CH11359 with the United States Department of Energy.

-
- [1] C. W. Arnold, T. B. Clegg, C. Iliadis, H. J. Karwowski, G. C. Rich, J. R. Tompkins, and C. R. Howell, *Phys. Rev. C* **85**, 044605 (2012).
- [2] H. Utsunomiya, S. Katayama, I. Gheroghe, S. Imai, H. Yamaguchi, D. Kahl, Y. Sakaguchi, T. Shima, K. Takahisa, and S. Miyamoto, *Phys. Rev. C* **92**, 064323 (2015).
- [3] T. Sasaqui, T. Kajino, G. J. Mathews, K. Otsuki, and T. Nakamura, *Astrophys. J* **634**, 1173 (2005).
- [4] L. Hsu, At Calibration of Low Energy Particle Detectors Workshop, Chicago, 2015 (unpublished).
- [5] C. Amole *et al.* (PICO Collaboration), ArXiv e-prints (2015), arXiv:1510.07754.
- [6] J. I. Collar, *Phys. Rev. Lett.* **110**, 211101 (2013).
- [7] V. D. Efros, P. von Neumann-Cosel, and A. Richter, *Phys. Rev. C* **89**, 027301 (2014).
- [8] J. Casal, M. Rodríguez-Gallardo, J. M. Arias, and I. J. Thompson, *Phys. Rev. C* **90**, 044304 (2014).
- [9] M. Odsuren, Y. Kikuchi, T. Myo, M. Aikawa, and K. Katō, *Phys. Rev. C* **92**, 014322 (2015).
- [10] Y. Kikuchi, M. Odsuren, T. Myo, and K. Katō, *Phys. Rev. C* **93**, 054605 (2016).
- [11] For a detailed review of nuclear reaction resonances, see A. M. Lane and R. G. Thomas, *Rev. Mod. Phys.* **30**, 257 (1958).
- [12] J. Chadwick and M. Goldhaber, *Proc. R. Soc. A* **151**, 479 (1935).
- [13] H. von Halban, *C. R. Acad. Sci.* **1938/01**, 1170 (1938).
- [14] B. Russell, D. Sachs, A. Wattenberg, and R. Fields, *Phys. Rev.* **73**, 545 (1948).
- [15] A. H. Snell, E. C. Barker, and R. L. Sternberg, *Phys. Rev.* **80**, 637 (1950).
- [16] B. Hamermesh and C. Kimball, *Phys. Rev.* **90**, 1063 (1953).
- [17] J. H. Gibbons, R. L. Macklin, J. B. Marion, and H. W. Schmitt, *Phys. Rev.* **114**, 1319 (1959).
- [18] W. John and J. M. Prosser, *Phys. Rev.* **127**, 231 (1962).
- [19] M. Fujishiro, T. Tabata, K. Okamoto, and T. Tsujimoto, *Can. J. Phys.* **60**, 1672 (1982).
- [20] M. J. Jakobson, *Phys. Rev.* **123**, 229 (1961).
- [21] B. L. Berman, R. L. V. Hemert, and C. D. Bowman, *Phys. Rev.* **163**, 958 (1967).
- [22] H. Utsunomiya, Y. Yonezawa, H. Akimune, T. Yamagata, M. Ohta, M. Fujishiro, H. Toyokawa, and H. Ohgaki, *Phys. Rev. C* **63**, 018801 (2000).
- [23] K. V. Tucker SN, Treacy PB, *Aust. J. Phys.* **23**, 651 (1970).
- [24] R. R. Spencer, G. C. Phillips, and T. E. Young, *Nucl. Phys.* **21**, 310 (1960).
- [25] H. N. Ngoc, M. Hors, and J. P. Y. Jorba, *Nucl. Phys.* **42**, 62 (1963).
- [26] H.-G. Clerc, K. J. Wetzels, and E. Spamer, *Nucl. Phys. A* **120**, 441 (1968).
- [27] G. Kuechler, A. Richter, and W. von Witsch, *Z. Phys. A* **326**, 447 (1987).
- [28] S. Dixit *et al.*, *Phys. Rev. C* **43**, 1758 (1991).
- [29] O. Burda, P. von Neumann-Cosel, A. Richter, C. Forssén, and B. A. Brown, *Phys. Rev. C* **82**, 015808 (2010).
- [30] F. C. Barker, *Aust. J. Phys.* **53**, 247 (2000).
- [31] E. D. McGarry and E. W. Boswell, National Bureau of Standards Special Publication No. 250-18, 1988.
- [32] W. F. Miller and W. J. Snow, *Nucleonics* **19**, 174 (1961).
- [33] R. L. Heath, Idaho National Laboratory Report No. IDO-16408, 1957.
- [34] R. Girgis and R. V. Lieshout, *Physica* **25**, 133 (1959).
- [35] J. Katakura and Z. D. Wu, *Nuclear Data Sheets* **109**, 1655 (2008).
- [36] A. E. Robinson, *Phys. Rev. C* **89**, 032801 (2014).
- [37] T. Murata, H. Matsunobu, and K. Shibata, JAEA-Research Report No. 2006-052, 2006.
- [38] W. B. Wilson *et al.*, “Sources 4c: a code for calculating ([alpha],n), spontaneous fission, and delayed neutron sources and spectra,” Los Alamos National Laboratory Report No. LA-UR-02-1839, 2002.
- [39] A. E. Robinson, Ph.D. thesis, University of Chicago (2015).
- [40] M. B. Chadwick, M. Herman, P. Obložinský, M. E. Dunn, Y. Danon, A. C. Kahler, D. L. Smith, *et al.*, *Nuclear Data Sheets* **112**, 2887 (2011).
- [41] J. W. De Juren and J. Chin, *J. Res. Nat. Bur. Stand.* **55**, 311 (1955).
- [42] W. C. Barber, *Annu. Rev. Nucl. Sci.* **12**, 1 (1962).
- [43] W. D. Allen, Atomic Energy Research Establishment Report No. NP/R 1667, 1955.
- [44] E. Padovani, S. A. Pozzi, S. D. Clarke, and E. C. Miller, Documentation for Radiation Safety Informational Computer Center code package CCC-791, 2012 (unpublished).
- [45] F. C. Barker, *Can. J. Phys.* **61**, 1371 (1983).
- [46] C. Angulo *et al.*, *Nucl. Phys. A* **656**, 3 (1999).
- [47] C. D. Nesaraja, S. D. Geraedts, and B. Singh, *Nuclear Data Sheets* **111**, 897 (2010).
- [48] K. S. Krane, *Phys. Rev. C* **81**, 044310 (2010).
- [49] E. Browne and J. K. Tuli, *Nuclear Data Sheets* **111**, 2425 (2010).
- [50] M. S. Basunia, *Nuclear Data Sheets* **114**, 1189 (2013).
- [51] E. A. McCutchan and A. A. Sonzogni, *Nuclear Data Sheets* **115**, 135 (2014).
- [52] J. A. Cameron and B. Singh, *Nuclear Data Sheets* **109**, 1 (2008).
- [53] D. D. Frenne and E. Jacobs, *Nuclear Data Sheets* **105**, 775 (2005).
- [54] K. Krane, M. Keck, E. Norman, and A. Shivprasad, *Applied Radiation and Isotopes* **70**, 740 (2012).
- [55] LND Inc, *2052 Cylindrical BF3 Neutron Detector* (LND Inc, Oceanside, New York, USA).
- [56] M. Fujishiro, K. Okamoto, and T. Tsujimoto, *Can. J. Phys.* **61**, 1579 (1983).
- [57] D. E. Alburger, R. E. Chrien, R. J. Sutter, and J. F. Wishart, *Phys. Rev. C* **70**, 064611 (2004).
- [58] Table III of A. Wattenberg, Argonne National Laboratory Report No. NP-1100, 1949 and references therein.
- [59] D. Tilley, J. Kelley, J. Godwin, D. Millener, J. Purcell, C. Sheu, and H. Weller, *Nucl. Phys. A* **745**, 155 (2004).
- [60] F. G. Kondev and S. Lalkovski, *Nuclear Data Sheets* **112**, 707 (2011).
- [61] S. Singh, A. K. Jain, and J. K. Tuli, *Nuclear Data Sheets* **112**, 2851 (2011).
- [62] S.-C. Wu, *Nuclear Data Sheets* **110**, 681 (2009).
- [63] C. Arnold, T. Clegg, H. Karwowski, G. Rich, J. Tompkins, and C. Howell, *Nucl. Instrum. Methods A* **647**, 55 (2011).
- [64] K. Mueck and F. Bensch, *J. Nucl. Energy* **27**, 857 (1973).

- [65] D. E. Alburger, R. E. Chrien, R. J. Sutter, and J. F. Wishart, *Phys. Rev. C* **70**, 064611 (2004).
- [66] S. Wüstenbecker, H. Becker, H. Ebbing, W. Schulte, M. Berheide, M. Buschmann, C. Rolfs, G. Mitchell, and J. Schweitzer, *Z. Phys. A* **344**, 205 (1992).