

CHCRUS

This is the accepted manuscript made available via CHORUS. The article has been published as:

Search for $2\nu\beta\beta$ decay of ^{136}Xe to the 0_{1}^{+} excited state of ^{136}Ba with the EXO-200 liquid xenon detector

J. B. Albert *et al.* (EXO-200 Collaboration) Phys. Rev. C **93**, 035501 — Published 8 March 2016 DOI: 10.1103/PhysRevC.93.035501

Search for $2\nu\beta\beta$ decay of ¹³⁶Xe to the 0^+_1 excited state of ¹³⁶Ba with the EXO-200 liquid xenon detector

J.B. Albert,¹ D.J. Auty,^{2,*} P.S. Barbeau,³ D. Beck,⁴ V. Belov,⁵ M. Breidenbach,⁶ T. Brunner,^{7,8} A. Burenkov,⁵

G.F. Cao,⁹ C. Chambers,¹⁰ J. Chaves,¹¹ B. Cleveland,^{12,†} M. Coon,⁴ A. Craycraft,¹⁰ T. Daniels,⁶

M. Danilov,⁵ S.J. Daugherty,¹ J. Davis,⁶ S. Delaquis,¹³ A. Der Mesrobian-Kabakian,¹² R. DeVoe,¹¹ T. Didberidze,² J. Dilling,⁸ A. Dolgolenko,⁵ M.J. Dolinski,¹⁴ M. Dunford,¹⁵ W. Fairbank Jr.,¹⁰ J. Farine,¹²

W. Feldmeier,¹⁶ S. Feyzbakhsh,¹⁷ P. Fierlinger,¹⁶ D. Fudenberg,¹¹ R. Gornea,^{15,8} K. Graham,¹⁵ G. Gratta,¹¹

C. Hall,¹⁸ M. Hughes,² M.J. Jewell,¹¹ A. Johnson,⁶ T.N. Johnson,¹ S. Johnston,¹⁷ A. Karelin,⁵ L.J. Kaufman,¹

R. Killick,¹⁵ J. King,¹⁷ T. Koffas,¹⁵ S. Kravitz,¹¹ R. Krücken,⁸ A. Kuchenkov,⁵ K.S. Kumar,¹⁹ D.S. Leonard,²⁰ C. Licciardi,¹⁵ Y.H. Lin,¹⁴ J. Ling,^{4, ‡} R. MacLellan,²¹ M.G. Marino,¹⁶ B. Mong,¹² D. Moore,¹¹ O. Njoya,¹⁹ R. Nelson,²² A. Odian,⁶ I. Ostrovskiy,¹¹ A. Piepke,² A. Pocar,¹⁷ C.Y. Prescott,⁶ F. Retière,⁸ P.C. Rowson,⁶

J.J. Russell,⁶ A. Schubert,¹¹ D. Sinclair,^{15,8} E. Smith,¹⁴ V. Stekhanov,⁵ M. Tarka,¹⁹ T. Tolba,^{13,§} R. Tsang,²

K. Twelker,¹¹ P. Vogel,²³ J.-L. Vuilleumier,¹³ A. Waite,⁶ J. Walton,⁴ T. Walton,¹⁰ M. Weber,¹¹ L.J. Wen,⁹ U. Wichoski,¹² T.A. Winick,¹⁴ J. Wood,²² Q.Y. Xu,¹¹ L. Yang,⁴ Y.-R. Yen,¹⁴, ¶ and O.Ya. Zeldovich⁵

(EXO-200 Collaboration)

¹Physics Department and CEEM, Indiana University, Bloomington, Indiana 47405, USA

²Department of Physics and Astronomy, University of Alabama, Tuscaloosa, Alabama 35487, USA

³Department of Physics, Duke University, and Triangle Universities

Nuclear Laboratory (TUNL), Durham, North Carolina 27708, USA

⁴Physics Department, University of Illinois, Urbana-Champaign, Illinois 61801, USA

⁵Institute for Theoretical and Experimental Physics, Moscow, Russia

⁶SLAC National Accelerator Laboratory, Menlo Park, California 94025, USA

⁷Physics Department, McGill University, Montreal, Quebec H3A 2T8, Canada

⁸ TRIUMF, Vancouver, British Columbia V6T 2A3, Canada

⁹Institute of High Energy Physics, Beijing, China

¹⁰ Physics Department, Colorado State University, Fort Collins, Colorado 80523, USA

¹¹Physics Department, Stanford University, Stanford, California 94305, USA

¹²Department of Physics, Laurentian University, Sudbury, Ontario P3E 2C6, Canada

¹³LHEP, Albert Einstein Center, University of Bern, Bern, Switzerland

¹⁴Department of Physics, Drexel University, Philadelphia, Pennsylvania 19104, USA

¹⁵Physics Department, Carleton University, Ottawa, Ontario K1S 5B6, Canada

¹⁶ Technische Universität München, Physikdepartment and Excellence Cluster Universe, Garching 80805, Germany

¹⁷Amherst Center for Fundamental Interactions and Physics

Department, University of Massachusetts, Amherst, MA 01003, USA

¹⁸Physics Department, University of Maryland, College Park, Maryland 20742, USA

¹⁹ Department of Physics and Astronomy, Stony Brook University, SUNY, Stony Brook, New York 11794, USA

²⁰IBS Center for Underground Physics, Daejeon, Korea

²¹Physics Department, University of South Dakota, Vermillion, South Dakota 57069, USA

²² Waste Isolation Pilot Plant, Carlsbad, New Mexico 88220, USA

²³Kellogg Lab, Caltech, Pasadena, California 91125, USA

EXO-200 is a single phase liquid xenon detector designed to search for neutrinoless double-beta decay of 136 Xe to the ground state of 136 Ba. We report here on a search for the two-neutrino double-beta decay of 136 Xe to the first 0^+ excited state, 0^+_1 , of 136 Ba based on a 100 kg yr exposure of ¹³⁶Xe. Using a specialized analysis employing a machine learning algorithm, we obtain a 90% CL half-life sensitivity of 1.7×10^{24} yr. We find no statistically significant evidence for the $2\nu\beta\beta$ decay to the excited state resulting in a lower limit of $T_{1/2}^{2\nu}$ $(0^+ \rightarrow 0_1^+) > 6.9 \times 10^{23}$ yr at 90% CL. This observed limit is consistent with the estimated half-life of 2.5×10^{25} vr.

I. INTRODUCTION

Nuclear double-beta $(\beta\beta)$ decay with the emission of two neutrinos, first considered by Goeppert-Mayer in 1935 [1], is a second-order weak transition observed in a number of even-even nuclei. The two-neutrino decay mode $(2\nu\beta\beta)$ to the ground state has been directly observed in nine nuclei with half-lives in excess of 10^{18} yr, with ¹³⁶Xe being the longest at 2.2 $\times 10^{21}$ yr [2, 3]. De-

^{*} Now at University of Alberta, Edmonton, Alberta, Canada

[†] Also SNOLAB, Sudbury ON, Canada

[‡] Now at Sun Yat-Sen University, Guangzhou, China

[§] Now at Institute of High Energy Physics, Beijing, China

[¶] Corresponding author: yung-ruey.yen@drexel.edu

cays with half-lives up to 10^{24} yr have been observed using indirect radiochemical and geochemical methods [4].

The Standard Model allows $2\nu\beta\beta$ decays to the first 0⁺ excited state, denoted hereafter as 0⁺₁, of the daughter nucleus if this state is energetically accessible. These decays are suppressed relative to their ground state counterparts and are generally accompanied by the emission of de-excitation γ s, creating a signature that is distinct from typical γ backgrounds and $2\nu\beta\beta$ decays to the ground state.

Measurements of decays to excited states may provide additional constraints on the nuclear matrix elements (NMEs) relevant to $\beta\beta$ decay. Using the NME for the excited state decay in a ratio between it and the ground state decay would allow any shared uncertainties in the NMEs for these transitions to be cancelled. Better knowledge of these NMEs could lead to a more precise determination of the effective Majorana neutrino mass from $0\nu\beta\beta$ half-life measurements [14]. Searches for decays to excited states may also test exotic theories of alternate $\beta\beta$ decay mechanisms.

The first investigation of $2\nu\beta\beta$ decay to excited states was performed by Fiorini et al. in 1977 with ⁷⁶Ge [5]. The first dedicated search was the Milano experiment in 1982, also with ⁷⁶Ge [6], while the first positive signal was obtained in 1995 for the $2\nu\beta\beta$ decay mode of ¹⁰⁰Mo to the 0⁺₁ excited state of ¹⁰⁰Ru [7]. This decay was then precisely measured by the NEMO-3 experiment where all β and γ tracks were observed [8]. $2\nu\beta\beta$ decays to excited states have also been observed in ¹⁵⁰Nd in 2004 [9] and again in 2014 where the γ coincidence was explicitly measured [10]. More recently, current $0\nu\beta\beta$ experiments such as Gerda [11] and KamLAND-Zen [12] have searched for decays to excited states using ⁷⁶Ge and ¹³⁶Xe, respectively.

Fig. 1 shows the energy level scheme of the $\beta\beta$ decay of ¹³⁶Xe. A typical $\beta\beta$ decay with a Q-value of 878.8 keV transitions from the ground state of ¹³⁶Xe to the 0⁺₁ state of ¹³⁶Ba. The de-excitation from the 0⁺₁ state results in the emission of two γ s with energies of 760.5 keV and 818.5 keV. The two de-excitation γ s are emitted with a preferential angular correlation to be aligned or antialigned [14]. Although the 2⁺₁ excited state of ¹³⁶Ba is at lower energy than the 0⁺₁ state, the direct decay to it from the parent 0⁺ ground state of ¹³⁶Xe is highly suppressed due to angular momentum. The intermediate 2⁺₁ excited state has a half-life of 1.930 ps [13], which is not resolved temporally in most practical detectors. EXO-200 has the ability to identify both the $\beta\beta$ and de-excitation γ s in an excited state decay.

For decays to excited states, $2\nu\beta\beta$ decay to the 0^+_1 state is expected to be the dominant decay mode, though $2\nu\beta\beta$ decay to the 2^+_1 state or $0\nu\beta\beta$ decays to either 0^+_1 or 2^+_1 are also possible. While EXO-200 can conduct searches for these other decays, this analysis focuses on the decay to the 0^+_1 excited state.



FIG. 1. Level scheme of the $\beta\beta$ decay of ¹³⁶Xe. Decay to the excited state of 0_1^+ will result in the emission of two γ s during the de-excitation to the ground state. The energy levels of the excited states are taken from [13]. Only the 0^+ and 2^+ levels relevant for this search are shown.

II. THEORETICAL DECAY RATE

The $2\nu\beta\beta$ decay rate for $0^+ \rightarrow 0^+$ transitions, including to the 0^+_1 excited state, can be written in the form:

$$[T_{1/2}^{2\nu}]^{-1} = G^{2\nu}(E_0, Z) \left| M_{GT}^{2\nu} - \frac{g_V^2}{g_A^2} M_F^{2\nu} \right|^2, \qquad (1)$$

where $T_{1/2}^{2\nu}$ is the half-life, $M_{GT}^{2\nu}$ and $M_F^{2\nu}$ are the Gamow-Teller and Fermi nuclear transition matrix elements, respectively, $G^{2\nu}$ is the phase space integral that depends on the Q-value (E_0) and the atomic number of the daughter nucleus (Z), and g_V and g_A are the vector and axialvector weak interaction coupling constants, respectively. The nuclear matrix elements (NMEs), $M^{2\nu}$, and phase space factor (PSF), $G^{2\nu}$, are expected to differ for decays to the ground and excited states. The PSF term results in a suppression of the decay rate due to the smaller Qvalue for the transition to the excited state.

The expected rate for the $2\nu\beta\beta$ decay of 136 Xe to the first excited state is calculated from the measured $2\nu\beta\beta$ decay rate to the ground state and the theoretical values for the PSF and NMEs of both the ground and excited states. For the 136 Xe $2\nu\beta\beta$ decays, the ratio of the PSF between the ground state and the excited state using the Schenter-Vogel Fermi function approximation and the Wilkinson correction for nuclear size is 3915 [15]. Using the PSF from [16], a suppression ratio of 3956 is found. The two calculated values are within 1% of each other, indicating that the theoretical uncertainty on the PSF suppression is small.

The NMEs for $\beta\beta$ decays have been calculated using the microscopic interacting boson model (IBM-2) with a method for isospin restoration and Argonne short range correlations in the closure approximation in [17]. The ratio of the NMEs between the ground state and the excited state for ¹³⁶Xe is 1.7. From Eq. 1, this leads to an additional suppression factor of 2.9 for the decay rate to the excited state. The calculated NME values are model dependent and have larger theoretical uncertainty than the PSF. Including both suppression factors and applying them to the ground state half-life measured by EXO-200 of 2.16×10^{21} yr [3], we estimate the expected $2\nu\beta\beta$ decay half-life of ¹³⁶Xe to the 0_1^+ excited state of ¹³⁶Ba of 2.5×10^{25} yr. While this prediction puts the decay beyond the sensitivity of the present EXO-200 dataset, uncertainties in the NMEs may produce an overestimate of the half-life, as in the cases of ¹⁰⁰Mo and ¹⁵⁰Nd.

III. DETECTOR DESCRIPTION

The EXO-200 detector is a cylindrical, single-phase liquid xenon (LXe) time projection chamber (TPC), 40 cm in diameter and 44 cm in length, filled with xenon enriched to 80.6% in 136 Xe. Two drift regions are separated by a cathode at the center. A detailed description of the detector can be found in [18]. Radioimpurities in the detector components were minimized by a careful screening process [19] and a detector design optimized to use a minimal amount of materials. External radioactivity is reduced by >25 cm thick lead walls on all sides and additional passive shielding of ≥ 50 cm of high purity cryogenic fluid, HFE-7000 [20]. The detector is located inside a clean room at the Waste Isolation Pilot Plant (WIPP) in Carlsbad, NM, USA, under an overburden of 1585^{+11}_{-6} meters water equivalent [21]. The remaining cosmic ray flux is detected by an active muon veto system consisting of plastic scintillation panels that cover the clean room on four sides.

Energy depositions by ionizing radiation create free electrons and scintillation light in the LXe that are registered by anode wire grids and arrays of avalanche photodiodes, respectively. Two sets of wire grids form each anode. From the cathode to the anodes, charges will pass by V-wires (induction) first before being collected on the U-wires (charge collection). The U- and V- wire grids are offset by 60° to allow for two dimensional reconstruction in the plane perpendicular to the axis of the TPC. The time difference between the prompt scintillation light and drifted charge collection allows for the position of the event in the drift direction (Z) to be determined. Charge deposits (clusters) in a given event that are spatially separated by ~ 1 cm or more can be individually resolved. An event can then be classified as singlesite (SS), or multi-site (MS), depending on the number of observed charge clusters. The total energy of an event is determined by combining the charge and scintillation signals, allowing improved energy resolution due to the anticorrelation between these channels [22]. Radioactive γ sources are periodically deployed at several positions near the TPC to characterize the detector response.

IV. DATA SET AND METHODOLOGY

This search uses the same data set ("low-background data") and event selection criteria as the recent searches for $0\nu\beta\beta$ decay [23, 24]. A total of 477.60±0.01 live days of data were collected between September 22, 2011 and September 1, 2013. Events consistent with noise, coincident with the muon veto, with more than one scintillation signal, or within 1 s of other events in the TPC are removed.

The fiducial volume is hexagonal with an apothem of 162 mm. Only regions within this hexagonal volume that are >10 mm from the cathode and anode wire planes are included. This geometry corresponds to a ¹³⁶Xe mass of 76.5 kg, or 3.39×10^{26} atoms of ¹³⁶Xe. The total exposure is 100 kg·yr (736 mol·yr). As in previous analyses, an energy range (summed over all charge clusters in an event) of 980-9800 keV is used. Finally, we require that all events have fully reconstructed U-, V-, and Z-positions.

A Geant4-based [25] Monte Carlo (MC) simulation of the detector and shielding (described in detail in [3]) is used to model the detector response. The simulation of the $2\nu\beta\beta$ decay of ¹³⁶Xe to the 0⁺₁ excited state of ¹³⁶Ba (hereafter referred to simply as "excited state events") accounts for the smaller Q-value and angular correlation between the de-excitation γ s. This MC is used to estimate the detection efficiency, determined by the percentage of these MC excited state events that survive all event selection cuts. The resulting efficiency for the excited state signal is $(23.2 \pm 2.0)\%$, with the dominant losses in efficiency arising from the fiducial volume and full reconstruction cuts. Errors in the estimate of this efficiency are accounted for in Sec. VI.

Based on periodic calibrations using γ sources (²²⁸Th, ⁶⁰Co, ²²⁶Ra, and ¹³⁷Cs), the energy scale and resolution are determined by fitting the full shape of the energy spectra observed in calibration data to MC simulations. In particular, the ⁶⁰Co source produces events with multiple γ s of energies similar to those produced in the excited state decay. These calibration events show good agreement with MC simulations across all energies. The energy scale calibration and resolution are determined separately for SS and MS events.

To search for the $2\nu\beta\beta$ decay to the excited state, a binned maximum-likelihood (ML) fit is performed simultaneously over the SS and MS events using probability density functions (PDFs) in two dimensions: energy and an excited state "discriminator" variable. The PDFs are generated using the MC simulation and smeared with the appropriate energy resolution function. The same background components used in previous analyses [23, 24] are also used here. The discriminator is specifically optimized to search for the unique event structure of the decay to the 0_1^+ excited state using machine learning techniques to be detailed in Sec. V. Using a 2D fit in energy and the machine learning discriminator improves the sensitivity to excited state events by more than a factor of three over the more generic technique in [23, 24], as detailed in Sec. VE.

V. MACHINE LEARNING

This analysis uses machine learning techniques to create a variable ("discriminator") which, for each event, indicates how "signal-like" (+1) or "background-like" (-1) it is. The machine learning software TMVA [26] (part of the ROOT data analysis framework [27]) is used to "train" an algorithm, using simulated signal and background data sets, to construct an optimized discriminator from a set of input variables and characterize the discriminator's performance.

A. Input variables

Prior to applying this algorithm to the low-background data, the input variables used to build the discriminator were finalized by optimizing the expected sensitivity to excited state events (as determined by the method described in Sec. V E), while minimizing systematic errors resulting from disagreement between data and MC (discussed in Sec. VI). By waiting to perform the fit until the method to determine the discriminator variable is finalized, potential biases from tuning the algorithm to the data set are minimized. The number of input variables is limited to reduce the sensitivity of the discriminator to systematic differences between the data and MC simulation.

The majority of the discriminating power is provided by several event variables used in previous EXO-200 analyses [23, 24]. These include the number of charge clusters in the event (multiplicity), the total event energy determined from the ionization and light response (energy), and the minimum separation of the charge deposits in the event from the anode wire plane or cylindrical walls of the TPC (standoff distance). The agreement between data and MC for these variables has been studied in previous analyses [3].

Although energy and multiplicity are included as inputs to the discriminator, the fit procedure also uses energy as the second fit dimension and separates the fit into SS and MS events. Including these additional dimensions allows the fit to better constrain individual background model components, while the discriminator is primarily useful for distinguishing between excited state events and all other backgrounds.

In addition, variables designed to take advantage of the energy deposits from the ¹³⁶Ba de-excitation γ s are used. These variables, γ_1 , γ_2 , and γ_{sum} , are defined as:

$$\gamma_{i} \equiv \min_{i \in S} \left\{ |E_{j} - \epsilon_{i}| \right\}$$
⁽²⁾

where S is the set of charge clusters in an event, E_j is the energy of cluster j, and ϵ_i are the de-excitation energies,

 $\epsilon_1 = 760.5$ keV, $\epsilon_2 = 818.5$ keV, and $\epsilon_{sum} = 1579$ keV. Hence, in the case where an event has a cluster of energy close to ϵ_i , the corresponding variable γ_i will be close to zero. Since only the total scintillation energy of an event is measured by the photodiodes, the γ_i variables are determined using only the energy reconstructed from the charge signal of each cluster. Other variables incorporating charge cluster positions were investigated but showed no improvement in the method's sensitivity.

B. Training data set

The training data consists of a signal class (excited state events) and a background class (all others). The dominant background to this search is $2\nu\beta\beta$ decays to the ground state, since they are the primary component of the low-background data. The background class was drawn from MC simulations, using a list of background components in quantities determined by the best fit model from a prior analysis of this data [24], to ensure that the training circumstances accurately reflect the low-background data set to which the discriminator will be applied. An equal number of excited state events were simulated to produce the signal class. All event selection cuts were applied prior to building this data set. In addition, both MC data classes were split in half, with one half used for training and the other for testing, to ensure that any patterns found in the training data set were not the result of statistical fluctuations. The distributions of each of the input variables used for this training data set are shown in Fig. 2.

C. Machine learning algorithm

Several different machine learning algorithms were investigated, including boosted decision trees (BDTs), multilayer perceptrons (MLPs, a class of neural network), knearest neighbors (KNN), and simple rectangular cuts on the inputs. In preliminary tests of these algorithms using a cut on the discriminator variable to separate signal from background in the testing data set, the BDT provided the greatest discriminating power. Adjustments to the training parameters of the BDT did not indicate significant improvement by this metric. More detailed studies of the sensitivity and systematic errors of the three most promising of these algorithms (MLP and two BDT variants) were also performed, as described in Sec. VE and VI. These studies provided further evidence that the BDT algorithm is optimal among those considered and that changes to its training parameters have a small effect on its performance: the sensitivity for the MLP was roughly 50% that of the BDT, and changing the BDT training parameters led to a 2% increase in sensitivity at the cost of a 5% increase in the systematic error.



FIG. 2. (Color online) Distributions of the input variables to the discriminator based on MC simulations. Signal (excited state events) is shown in solid blue and background (all other events) in hatched red.

D. Trained decision tree results

After training the BDT, the effectiveness of individual variables in deciding whether an event is signal-like or background-like can be determined. The BDT consists of many individual decision trees, each of which performs a series of binary cuts on the input variables, with the final nodes of the cuts each assigned to either signal (+1) or background (-1). The discriminator variable is then given by a sum of the individual trees' assignments, weighted by their classification performance. Further description of the BDT algorithm can be found in [26]. The ranking for any given variable is a measure of the fraction of decision tree cuts which use that variable, and is given in Tab. I. As expected, multiplicity is an effective discriminator between the decays to the excited state, which are largely MS, and $2\nu\beta\beta$ decays to the ground state, which are primarily SS, with most of the additional information contained in the energy variables.

The trained BDT is then applied to low-background data, calibration data, and MC to determine the value of the discriminator variable for each event. In particular, this allows for comparison between data and MC for calibration sources, which can be used to quantify systematic effects. These comparisons are done for 60 Co, 226 Ra, and 228 Th, both for SS and MS events. The primary γ event backgrounds consist of MS events, for which a representative comparison using 228 Th is shown in Fig. 3. While the data and MC distributions typically agree within 10%,

Rank	Variable	Importance
1	multiplicity	0.28
2	energy	0.27
3	γ_{sum}	0.14
4	standoff distance	0.12
5	γ_1	0.10
6	γ_2	0.09

TABLE I. Importance ranking of the input variables in the final boosted decision tree. The "importance" of each variable denotes the fraction of decision tree cuts which use that variable.

the remaining deviations are accounted for as systematic errors, to be discussed in Sec. VI.

Because $2\nu\beta\beta$ decays to the ground state constitute one of the largest backgrounds to this search, it is important to be able to quantify the difference in the shape of the discriminator variable distribution for data and MC for this component. To determine the $2\nu\beta\beta$ spectrum, each non- $2\nu\beta\beta$ component (from MC) is subtracted from the low-background data, in amounts given by a fit from a prior analysis [24]. The final subtracted $2\nu\beta\beta$ distribution is then compared to the MC $2\nu\beta\beta$ component in the discriminator variable (Fig. 3). The results indicate good SS agreement, with some notable differences in the MS spectrum, which are accounted for in Sec. VI.



FIG. 3. Discriminator agreement between data and MC for MS ²²⁸Th (left) and MS $2\nu\beta\beta$ decays to the ground state (right). The $2\nu\beta\beta$ data is generated by subtraction of the non- $2\nu\beta\beta$ components, in amounts given by a prior analysis, from the low-background data. Residuals are shown in the bottom panels, ignoring those bins with insufficient counts.

E. Sensitivity and significance estimates

Before the final fit to the data, the discriminator sensitivity was estimated by performing a two-dimensional profile likelihood fit on toy MC data sets. These MC data sets are created using the best-fit model from a previous analysis of the data [24], with no excited state events included. A fit in energy and discriminator is applied to find the 90% confidence level (CL) upper limit on the number of excited state events, as determined by a profile of the negative log likelihood (NLL). An expected sensitivity of 43 events, corresponding to $T_{1/2}^{2\nu} = 1.7 \times 10^{24}$ yr, is given by the median upper limit obtained from toy MC data sets. This demonstrates an improvement by more than a factor of three over a simple fit to energy and standoff distance (as in [23, 24]), which has a sensitivity of 170 events, corresponding to $T_{1/2}^{2\nu} = 4.2 \times 10^{23}$ yr. The toy MC data sets with zero injected excited state

The toy MC data sets with zero injected excited state events can also be used to calculate the significance of a nonzero final fit to data. The distribution of Δ NLL values for fits with 0 excited state events is calculated; then, the fraction of toy MC data sets with Δ NLL values greater than that found by the fit to data determines the level of compatibility with the null hypothesis. The full sensitivity and significance distributions are shown in Sec. VII.

VI. SYSTEMATIC UNCERTAINTIES

Systematic uncertainties in this analysis can be divided into those deriving from the event selection and background model, previously evaluated in [23], and those unique to this analysis. In all cases, these uncertainties are accounted for by applying Gaussian constraints on the fit parameters.

Unique to this analysis is an excited state event normalization term that accounts for discrepancies between MC and data in the discriminator and energy distributions of background components. To calculate this, toy MC data sets with a non-zero number of excited state events are generated from PDFs that have been skewed by the relative differences between data and MC, as measured in calibration source data (Fig. 3). These toy data sets are then fit to the standard, un-skewed PDFs, and the resulting difference between the fitted and simulated number of excited state events is determined. This fractional difference is accounted for as a systematic error which is applied in the normalization term specific to the excited state component.

While the deviations between calibration data and MC measured for the ²²⁸Th, ²²⁶Ra, and ⁶⁰Co sources are used to skew the corresponding background components, the $2\nu\beta\beta$ ground state decay is skewed by deviations between MC and the background-subtracted data set. Differences between calibration sources are small, and the majority of the skewing error comes from the $2\nu\beta\beta$ component, so the exact choice of which calibration source is most similar to a given background component is inconsequential.

Due to statistical errors resulting from the background subtraction, the skewing cannot be precisely measured in the tail of the $2\nu\beta\beta$ distribution (*i.e.*, values of the discriminator >0.5 or <-0.4 in the MS data). To prevent artificially distorting the distribution by errors resulting from the background subtraction, the skewing is forced to go to zero for bins with too few (< 20) events. Several methods for suppressing the skewing in the low statistics tails were tested, with negligible effects on the resulting error. If an excited state decay signal were present in the data, its incomplete subtraction could lead to a slight overestimation of this systematic; however,



FIG. 4. (Color online) Fit results for the discriminator variable spectrum of MS events. The excited state event distribution is given by the dashed green line, concentrated toward positive discriminator values. Data points are shown in black and residuals between data and the best fit, normalized to the Poisson error, are presented, ignoring bins with 0 events.

while the estimation of this error may be conservative, this background-subtracted data set is only used for this study, so the procedure does not bias the final result. The final error found by this study is 15%.

The additional Gaussian constraints are quantified in Tab. II. The SS fraction constraint is applied to the ratio of the number of SS events to the total number of events (SS/(SS+MS)) for each component, and is determined by the deviations in this ratio between data and MC for the calibration sources. A normalization constraint common to all components, accounting for the uncertainty in detection efficiency, is calculated from studies of the event reconstruction efficiency and the rate of events from calibration sources of known activity. A further normalization constraint is applied to account for uncertainties in the location of degenerate backgrounds. The relative fractions of neutron-capture related backgrounds, coming from cosmic ray muons and radioactive decays in the salt surrounding the experiment, are constrained according to MC studies and data coincident with muon-vetopanel events. The activity of radon in the liquid xenon is measured by observing characteristic ²¹⁴Bi to ²¹⁴Po coincidences in the detector, and the uncertainty on this measurement is translated into a constraint on the normalization of the radon components.

The final systematic accounts for uncertainties in the " β -scale", which describes the possible difference in energy scales of β -like and γ -like events. Since the majority

of the energy in the excited state events is deposited by the de-excitation γ s, these events are calibrated as γ -like components in the MC simulations. The β -scale variable is defined as an energy-independent ratio of γ to β energy scales and is allowed to float as a free parameter in the profile likelihood fit. We find it to be 0.9943 \pm 0.0006.

Among the systematics, the SS fraction error had the largest effect on the final upper limit of the search. This uncertainty allows the largest background, $2\nu\beta\beta$ decays to the ground state, to shift from SS to MS, making them more signal-like.

Constraint	Error (%)
Excited state normalization	15
SS fraction	4
Common normalization	8.6
Background normalization	20
Neutron capture fractions	20
Radon in liquid xenon	10

TABLE II. Gaussian constraints applied to fit parameters to account for systematic uncertainties. These errors are explicitly included as input to the final fit to the low-background data, and are not the systematic uncertainties on the final result.



FIG. 5. (Color online) Fit results for the energy spectrum of MS events. Data points are shown in black and residuals between data and the best fit, normalized to the Poisson error, are presented, ignoring bins with 0 events.



FIG. 6. NLL profile in the number of excited state events. The best fit value is 43 events, while the fit with 0 events has a Δ NLL value of 1.0. The 90% upper limit (dashed line) is 104 events.

VII. RESULTS

Plots of the discriminator variable and energy spectra of MS events for the fit to the data are shown in Fig. 4 and 5. These plots illustrate the relative shapes and best fit quantities for all components, but do not contain all information used in the full fit, such as constraints on the SS fraction. The best fit values for background components are compatible with prior results, with the calculated half-life of the $2\nu\beta\beta$ decay to the ground state agreeing within systematic error [3]. This fit was also checked for robustness against hypothetical backgrounds ⁸⁸Y and ^{110m}Ag, considered due to their mixed β and γ composition that could mimic an excited state signal. Both backgrounds were found to have a $< 1\sigma$ effect on the fit result, and prior analyses found no evidence of either component, so they are not included in the final result. The corresponding profile likelihood scan in the number of excited state events is shown in Fig. 6. The profile finds a best fit value of 43 events, with a 90% CL upper limit of 104 events, assuming Wilks' theorem [28, 29] holds. Taking this result as a limit on the half-life of the decay to the 0_1^+ excited state, we obtain $T_{1/2}^{2\nu} \ge 6.9 \times 10^{23}$ yr at 90% CL.

The upper limit obtained from the data is roughly a factor of two weaker than the median expected sensitivity of $T_{1/2}^{2\nu} = 1.7 \times 10^{24}$ yr calculated in Sec. V.E. The calculation of significance, as described in Sec. V.E., indicates that the result from data is consistent with the null hypothesis at 1.6σ (Fig. 7). Thus, this analysis does not find statistically significant evidence for a non-zero component of $2\nu\beta\beta$ decays to the excited state.



FIG. 7. (Color online) Left: Distribution of 90% CL upper limits on the number of excited state events from toy MC data sets with no excited state events, using fits to energy and either standoff distance (red) or discriminator (blue). Median values are drawn with dashed lines, with the limit from data at 104 events drawn as a solid line. Right: Distribution of Δ NLL values from toy MC data sets with no excited state events. The Δ NLL for 0 events from data is shown as a solid line at 1.0.

VIII. CONCLUSION

a next-generation tonne-scale experiment.

We report here the results from a search for the $2\nu\beta\beta$ decay of ¹³⁶Xe to the 0_1^+ excited state of ¹³⁶Ba with the first two years of EXO-200 data, corresponding to a ¹³⁶Xe exposure of 100 kg·yr. No statistically significant evidence for this process is found. We obtain a limit on the half-life of this process of $T_{1/2}^{2\nu} > 6.9 \times 10^{23}$ yr at 90% CL, which is comparable to the recent results from KamLAND-Zen [12]. However, the ability of this analysis to identify the detailed event structure of excited state decays provides an expected sensitivity of 1.7×10^{24} yr, higher than the observed limit. This limit is consistent with theoretical calculations which predict, with substantial uncertainty, a half-life of $\sim 10^{25}$ yr; hence, sensitivity to this decay may be within reach of future analyses or

IX. ACKNOWLEDGMENTS

The collaboration gratefully acknowledges Thilo Michel for valuable discussions, the KARMEN collaboration for supplying the cosmic-ray veto detectors, and the WIPP for their hospitality. EXO-200 is supported by DOE and NSF in the United States, NSERC in Canada, SNF in Switzerland, IBS in Korea, RFBR-14-02-00675 in Russia, DFG Cluster of Excellence "Universe" in Germany, and CAS-IHEP Fund and ISTCP (2015DFG02000) in China. EXO-200 data analysis and simulation uses resources of the National Energy Research Scientific Computing Center (NERSC), which is supported by the Office of Science of the U.S. Department of Energy under Contract No. DE-AC02-05CH11231.

- [1] M. Goeppert-Mayer, Phys. Rev. 48, 512 (1935).
- [2] K. A. Olive *et al.* (Particle Data Group), Chin. Phys. C38, 090001 (2014).
- [3] J. B. Albert *et al.* (EXO-200), Phys. Rev. C 89, 015502 (2014), arXiv:1306.6106 [nucl-ex].
- [4] A. S. Barabash, Phys. Rev. C 81, 035501 (2010).
- [5] E. Fiorini, Proc. Int. Conf. NEUTRINO '77 2, 315 (1978).
- [6] E. Bellotti, E. Fiorini, C. Liguori, A. Pullia, A. Sarracino, and L. Zanotti, Lett. Nuovo Cim. 33, 273 (1982).
- [7] A. S. Barabash et al., Phys. Lett. B 345, 408 (1995).
- [8] M. F. Kidd, J. H. Esterline, W. Tornow, A. S. Barabash, and V. I. Umatov, Nucl. Phys. A 821, 251 (2009), arXiv:0902.4418 [nucl-ex].

- [9] A. S. Barabash, F. Hubert, P. Hubert, and V. I. Umatov, Phys. Atom. Nucl. 67, 1216 (2004).
- [10] M. F. Kidd, J. H. Esterline, S. W. Finch, and W. Tornow, Phys. Rev. C 90, 055501 (2014), arXiv:1411.3755 [nuclex].
- [11] M. Agostini *et al.* (GERDA), J. Phys. G **42**, 115201 (2015), arXiv:1506.03120 [hep-ex].
- [12] K. Asakura *et al.* (KamLAND-Zen), (2015), arXiv:1509.03724 [hep-ex].
- [13] A. A. Sonzogni, Nuclear Data Sheets 95, 837 (2002), data extracted from the ENSDF database, Oct. 9, 2015, http://www.nndc.bnl.gov/.
- [14] M. Duerr, M. Lindner, and K. Zuber, Phys. Rev. D 84, 093004 (2011), arXiv:1103.4735 [hep-ph].
- [15] P. Vogel, private communication (2012).

- [16] J. Kotila and F. Iachello, Phys. Rev. C 85, 034316 (2012), arXiv:1209.5722 [nucl-th].
- [17] J. Barea, J. Kotila, and F. Iachello, Phys. Rev. C 91, 034304 (2015), arXiv:1506.08530 [nucl-th].
- [18] M. Auger *et al.*, JINST **7**, P05010 (2012), arXiv:1202.2192 [physics.ins-det].
- [19] D. S. Leonard *et al.*, Nucl. Instrum. Meth. A **591**, 490 (2008), arXiv:0709.4524 [physics.ins-det].
- [20] "3M HFE-7000," http://www.3m.com/.
- [21] E.-I. Esch, T. J. Bowles, A. Hime, A. Pichlmaier, R. Reifarth, and H. Wollnik, Nuclear Instruments and Methods in Physics Research A 538, 516 (2005), astroph/0408486.
- [22] E. Conti *et al.*, Phys. Rev. B 68, 054201 (2003), arXiv:hep-ex/0303008 [hep-ex].
- [23] J. B. Albert *et al.* (EXO-200), Nature **510**, 229 (2014), arXiv:1402.6956 [nucl-ex].
- [24] J. B. Albert *et al.* (EXO-200), Phys. Rev. D **90**, 092004 (2014), arXiv:1409.6829 [hep-ex].
- [25] J. Allison et al., IEEE Trans. Nucl. Sci. 53, 270 (2006).
- [26] A. Hocker et al., PoS (ACAT 2007) ACAT, 040 (2007), arXiv:physics/0703039 [physics].
- [27] R. Brun and F. Rademakers, Nucl. Instrum. Meth. A 389, 81 (1997).
- [28] S. S. Wilks, Annals Math. Statist. 9, 60 (1938).
- [29] G. Cowan, *Statistical Data Analysis* (Oxford Science Publications, Clarendon, 1998).