This is the accepted manuscript made available via CHORUS. The article has been published as:
$Y$-soft $\wedge\{146\}$ Ba and the role of nonaxial shapes at $N \approx 90$
A. J. Mitchell et al.

Phys. Rev. C 93, 014306 - Published 12 January 2016
DOI: 10.1103/PhysRevC.93.014306

# $\gamma$-soft ${ }^{146} \mathrm{Ba}$ and the role of non-axial shapes at $N \sim 90$ 

A. J. Mitchell,,$^{1, a}$ C. J. Lister, ${ }^{1}$ E. A. McCutchan, ${ }^{2}$ M. Albers, ${ }^{3,}$ b A. D. Ayangeakaa, ${ }^{3}$ P. F. Bertone, ${ }^{3,}$ c<br>M. P. Carpenter, ${ }^{3}$ C. J. Chiara, ${ }^{3,4, ~ d ~ P . ~ C h o w d h u r y, ~}{ }^{1}$ J. A. Clark, ${ }^{3}$ P. Copp, ${ }^{1}$ H. M. David, ${ }^{3, ~ e ~ A . ~ Y . ~ D e o, ~}$, , f B. DiGiovine, ${ }^{3}$ N. D'Olympia, ${ }^{1, g}$ R. Dungan, ${ }^{5}$ R. D. Harding, ${ }^{1,6, h}$ J. Harker, ${ }^{3,4}$ S. S. Hota, ${ }^{1,}$ i R. V. F. Janssens, ${ }^{3}$ F. G. Kondev, ${ }^{7}$ S. H. Liu, ${ }^{8,9, j}$ A. V. Ramayya, ${ }^{10}$ J. Rissanen, ${ }^{11,}$ k G. Savard,,${ }^{3,12}$ D. Seweryniak, ${ }^{3}$ R. Shearman,,$^{1,6,1}$ A. A. Sonzogni, ${ }^{2}$ S. L. Tabor, ${ }^{5}$ W. B. Walters, ${ }^{4}$ E. Wang, ${ }^{10}$ and S. Zhu ${ }^{3}$<br>${ }^{1}$ Department of Physics and Applied Physics, University of Massachusetts Lowell, Lowell, Massachusetts 01854<br>${ }^{2}$ National Nuclear Data Center, Brookhaven National Laboratory, Upton, New York 11973<br>${ }^{3}$ Physics Division, Argonne National Laboratory, Argonne, Illinois 60439<br>${ }^{4}$ Department of Chemistry and Biochemistry, University of Maryland, College Park, Maryland 20742<br>${ }^{5}$ Physics Department, Florida State University, Tallahassee, Florida 32306<br>${ }^{6}$ Department of Physics, University of Surrey, Guildford GU2 7XH, UK<br>${ }^{7}$ Nuclear Engineering Division, Argonne National Laboratory, Argonne, Illinois 60439<br>${ }^{8}$ Department of Chemistry, University of Kentucky, Lexington, Kentucky 40506<br>${ }^{9}$ Department of Physics and Astronomy, University of Kentucky, Lexington, Kentucky 40506<br>${ }^{10}$ Physics Department, Vanderbilt University, Nashville, Tennessee 37235<br>${ }^{11}$ Lawrence Berkeley National Laboratory, Berkeley, California 94720<br>${ }^{12}$ Department of Physics, University of Chicago, Chicago, Illinois 60637

(Dated: December 9, 2015)


#### Abstract

Low-spin states in the neutron-rich, $N=90$ nuclide ${ }^{146} \mathrm{Ba}$ were populated following $\beta$-decay of ${ }^{146} \mathrm{Cs}$, with the goal of clarifying the development of deformation in barium isotopes through delineation of their non-yrast structures. Fission fragments of ${ }^{146} \mathrm{Cs}$ were extracted from a $1.7-\mathrm{Ci}$ ${ }^{252} \mathrm{Cf}$ source and mass-selected using the CARIBU facility. Low-energy ions were deposited at the center of a box of thin $\beta$ detectors, surrounded by a high-efficiency HPGe array. The new ${ }^{146} \mathrm{Ba}$ decay scheme now contains 31 excited levels extending up to $\sim 2.5 \mathrm{MeV}$ excitation energy, double what was previously known. These data are compared to predictions from the Interacting Boson Approximation (IBA) model. It appears that the abrupt shape change found at $N=90$ in samarium and gadolinium is much more gradual in barium and cerium, due to an enhanced role of the $\gamma$ degree of freedom.


PACS numbers: 23.40.-s, 21.60.Fw, 23.20.Lv
Keywords: $\beta$ decay; $\gamma$-ray spectroscopy; $E_{\gamma}, I_{\gamma} ; N=90$ shape transition; Interacting Boson Approximation
${ }^{\text {a }}$ Present address: Department of Nuclear Physics, Australian National University, Canberra, ACT 2601, Australia; Email: aj.mitchell@anu.edu.au
b Present address: Ernst \& Young GmbH, Wirtschaftspruefungsgesellschaft, Mergenthalerallee 3-5, D-65760 Eschborn, Germany
c Present address: Marshall Space Flight Center, Huntsville, Alabama 35812
d Present address: U.S. Army Research Laboratory, Adelphi, Maryland 20783
e Present address: GSI Helmholtzzentrum für Schwerionenforschung GmbH, D-64291 Darmstadt, Germany
${ }^{\mathrm{f}}$ Present address: Department of Physics, Indian Institute of Technology Roorkee, Roorkee 247 667, India
g Present address: Passport Systems Inc., 70 Treble Cove Road, 1st Floor, Billerica, Massachusetts 01862
${ }^{h}$ Present address: Department of Physics, University of York, Heslington, York, YO10 5DD, UK
${ }^{\text {i }}$ Present address: Department of Nuclear Physics, Australian National University, Canberra, ACT 2601, Australia
j Present address: West Physics, 3825 Paces Walk SE, Suite 250, Atlanta, Georgia 30339
${ }^{k}$ Present address: Fennovoima Oy, Salmisaarenaukio 1, 00180 Helsinki, Finland
${ }^{1}$ Present address: National Physical Laboratory, Hampton Road, Teddington, Middlesex TW11 0LW, UK

## I. INTRODUCTION

The transition from spherical, shell-model-like behavior, to deformed collective motion has always been interesting, yet controversial, in nuclear structure. Although models exist for each extreme [1, 2], the actual transition from one limit to the other remains confused and lacks a ubiquitous description. Stable isotopes of rareearth elements near $Z=64$ with $N=90$ (e.g. ${ }^{156}$ Dy $(Z=66)[3],{ }^{154} \mathrm{Gd}(Z=64)[4],{ }^{152} \mathrm{Sm}(Z=62)$ [5], and $\left.{ }^{150} \mathrm{Nd}(Z=60)[6]\right)$ exhibit remarkable similarities in the excitation energies of ground-state bands and excited $J^{\pi}=0^{+}$and $J^{\pi}=2^{+}$sequences. The abrupt onset of deformation has received particularly intense scrutiny with general discussions often framed in terms of a phase transition [7-9]; in this case, a specific type of phase transition encapsulated by the $\mathrm{X}(5)$ model [10-13]. However, such an approach is not fully supported by all available experimental data and more generalized shape-coexistence models have been proposed [14-16].

A way to clarify this issue is to widen the scope of invesigation to both heavier and lighter nuclei. In a general sense, the behavior of transitional nuclei is expected to follow the number of valence particles, as predicted in the
$N_{p} N_{n}$ scheme of Casten [17]. In practice, the underlying fermionic structure appears to be important, with residual interactions between protons and neutrons in specific orbits playing a key role in 'tipping' the nuclear shape from spherical to deformed $[18,19]$. In this way, the $N=90$ border between shapes retains its significance, although the sharpness of the transition becomes more muted. The nuclei which have been most extensively studied are all stable, but it is relevant to enquire about how the transitional structures evolve as one progresses up to ${ }^{158} \operatorname{Er}(Z=68)$, and ${ }^{160} \mathrm{Yb}(Z=70)$ or down to ${ }^{148} \mathrm{Ce}(Z=58)$, and ${ }^{146} \mathrm{Ba}(Z=56)$. The lighter nuclei in this sequence are quite neutron-rich and cannot be accessed by fusion-evaporation reactions, and so fissionfragment spectroscopy and $\beta$ decay are the appropriate probes.
Nuclei in this region are also expected to exhibit strong octupole correlations [20]. Polarization of spin-orbit partners appears to quench the $Z=64$ sub-shell closure, resulting in strong couplings between $\Delta J=\Delta L=3 \mathrm{nu}-$ cleon orbitals $\left(\pi d_{5 / 2}-\pi h_{11 / 2}\right.$ and $\left.\nu f_{7 / 2}-\nu i_{13 / 2}\right)$. The onset of octupole collectivity in Gd, Sm , and Nd is observed between $N=88$ and $N=90$. The Ba isotopes undergo a smoother transition between $N=86$ and $N=88$, two neutrons earlier than expected from the behavior found in the $Z=60$ to $Z=64$ range [21]. Prompt-fission spectroscopy of ${ }^{146} \mathrm{Ba}$ (for example [2224]) has identified the ground-state and negative-parity bands to moderate spins. Octupole deformation in ${ }^{146} \mathrm{Ba}$ has been discussed [25-27], with the suggestion that these effects are weak and disappear at medium to high spins [24]. Although the yrast states in ${ }^{146} \mathrm{Ba}$ are established, there is limited information pertaining to the non-yrast, low-spin levels. This paper reports on the $\beta$-decay of ${ }^{146} \mathrm{Cs}$, with focus on identifying and quantifying properties of the important low-spin, non-yrast states in ${ }^{146} \mathrm{Ba}$ which inform this discussion.

## II. EXPERIMENTAL METHOD

The data presented here were obtained at the CAlifornium Rare Ion Breeder Upgrade (CARIBU [28]) facility at Argonne National Laboratory. Spontaneous fission fragments extracted from a $1.7-\mathrm{Ci}{ }^{252} \mathrm{Cf}$ source were thermalized in a gas catcher, in which interactions with high-purity He gas and with RF and DC fields combine to result in a low-emittance beam. An isotopically-pure beam of singly-charged ${ }^{146} \mathrm{Cs}$ nuclei was selected by the isobar separator. The beam was cooled to $\sim 2 \mathrm{keV}$ and bunched before delivery to the low-energy experimental area. Approximately 300 ions/s were delivered to the new decay-spectroscopy station, where they were implanted on an aluminum foil located at the center of an array of $\gamma$-ray and $\beta$-particle detectors.

The CARIBU decay station consists of the SATURN (Scintillator And Tape Using Radioactive Nuclei) system coupled to the X-Array, a highly-efficient array of five

129

## 130

## 131

132
133
134

TABLE I: Observed $\gamma$-ray transitions in ${ }^{146} \mathrm{Ba}$ placed in the level scheme of Fig. 5. Relative intensities, $I_{\gamma}$, are normalized to the $181-\mathrm{keV} \gamma$ ray, taken as 100 . For absolute intensity per 100 parent decays, the relative intensity should be multiplied by $0.42(5)$. The method for determining the normalization using the $141-\mathrm{keV} \gamma$ ray from ${ }^{146} \mathrm{La}$ [30] is described in the text (for reference, the relative intensity of this $\gamma$ ray is included in the table). Strong transitions were calculated from prompt $\gamma$-ray singles data; those marked ${ }^{\dagger}$ are from coincidence data. Upper limits on $I_{\gamma}$ for transitions from new levels to the ground state that have not been observed, but may occur if $J_{i}^{\pi} \neq 0$, are marked ${ }^{u}$. Uncertainties are statistical and based on fitting approximations.

| $E_{\gamma}$ <br> $(\mathrm{keV})$ | $I_{\gamma}$ | $E_{\text {initial }}$ <br> $(\mathrm{keV})$ | $E_{\text {final }}$ <br> $(\mathrm{keV})$ |
| :---: | :---: | :---: | :---: |
| $140.7(1)$ | $41(2)$ | - | - |
| $181.3(1)$ | $100(3)$ | $181.1(1)$ | 0.0 |
| $307.3(1)$ | $5.3(4)^{\dagger}$ | $821.6(2)$ | $513.9(2)$ |
| $332.9(1)$ | $13(2)$ | $513.9(2)$ | $181.1(1)$ |
| $558.1(1)$ | $23(1)$ | $739.4(1)$ | $181.1(1)$ |
| $639.9(1)$ | $4.4(3)$ | $821.6(2)$ | $181.1(1)$ |
| $739.1(2)$ | $5.1(6)$ | $739.4(1)$ | 0.0 |
| $743.6(6)$ | $2.7(6)$ | $1566.2(2)$ | $821.6(2)$ |
| $772.2(1)$ | $5.2(6)$ | $1511.7(2)$ | $739.4(1)$ |
| $788.9(1)$ | $0.50(6)^{\dagger}$ | $1529.1(1)$ | $739.4(1)$ |
| $795.6(2)$ | $2.0(7)$ | $1309.5(3)$ | $513.9(2)$ |
| $816.6(6)$ | $1.2(5)$ | $1638.2(3)$ | $821.6(2)$ |
| $827.3(4)$ | $2.2(9)$ | $1566.2(2)$ | $739.4(1)$ |
| $871.3(1)$ | $3.4(6)$ | $1052.4(3)$ | $181.1(1)$ |
| $892.9(4)$ | $1.24(11)^{\dagger}$ | $1632.6(2)$ | $739.4(1)$ |
| $894.1(1)$ | $0.21(4)^{\dagger}$ | $1714.9(2)$ | $821.6(2)$ |
| $918.7(3)$ | $1.4(5)$ | $1657.3(2)$ | $739.4(1)$ |
| $933.1(1)$ | $5.2(4)^{\dagger}$ | $1114.7(2)$ | $181.1(1)$ |
| $943.6(2)$ | $1.1(1)$ | $1683.1(2)$ | $739.4(1)$ |
| $976.7(1)$ | $2.6(8)$ | $1714.9(2)$ | $739.4(1)$ |
| $1052.7(4)$ | $1.5(7)$ | $1566.2(2)$ | $513.9(2)$ |
| $1073.5(2)$ | $3.5(7)$ | $1255.4(2)$ | $181.1(1)$ |
| $1115.2(3)$ | $2.9(5)$ | $1114.7(2)$ | 0.0 |
| $1128.4(1)$ | $2.7(2)^{\dagger}$ | $1309.5(3)$ | $181.1(1)$ |
| $1160.9(1)$ | $1.2(1)^{\dagger}$ | $1342.0(3)$ | $181.1(1)$ |

TABLE I - continued

| $E_{\gamma}$ <br> $(\mathrm{keV})$ | $I_{\gamma}$ | $E_{\text {initial }}$ <br> $(\mathrm{keV})$ | $E_{\text {final }}$ <br> $(\mathrm{keV})$ |
| :---: | :---: | :---: | :---: |
| $1217(1)$ | $0.6(5)$ | $1397.8(2)$ | $181.1(1)$ |
| $1229.5(2)$ | $0.58(9)^{\dagger}$ | $1410.8(3)$ | $181.1(1)$ |
| $1256.1(3)$ | $3.1(6)$ | $1255.4(2)$ | 0.0 |
| $1299(1)$ | $0.8(4)$ | $2036.8(2)$ | $739.4(1)$ |
| $1310(1)^{u}$ | $<0.19$ | $1309.5(3)$ | 0.0 |
| $1330.4(2)$ | $1.6(5)$ | $1511.7(2)$ | $181.1(1)$ |
| $1342(2)^{u}$ | $<0.19$ | $1342.0(3)$ | 0.0 |
| $1348.9(3)$ | $1.6(5)$ | $1529.1(2)$ | $181.1(1)$ |
| $1385.6(2)$ | $4.3(7)$ | $1566.2(2)$ | $181.1(1)$ |
| $1397.8(4)$ | $1.1(6)$ | $1397.8(2)$ | 0.0 |
| $1412(1)^{u}$ | $<0.20$ | $1410.8(3)$ | 0.0 |
| $1451.8(1)$ | $0.83(12)^{\dagger}$ | $1632.6(2)$ | $181.1(1)$ |
| $1457.0(2)$ | $3.3(7)$ | $1638.2(3)$ | $181.1(1)$ |
| $1487.4(4)$ | $2(1)$ | $1668.5(2)$ | $181.1(1)$ |
| $1502.5(2)$ | $2.8(2)^{\dagger}$ | $1683.1(1)$ | $181.1(1)$ |
| $1510(1)$ | $0.9(5)$ | $1511.7(2)$ | 0.0 |
| $1529(1)^{u}$ | $<0.21$ | $1529.1(2)$ | 0.0 |
| $1533.7(5)$ | $1.5(9)$ | $1714.9(2)$ | $181.1(1)$ |
| $1566.7(3)$ | $2.6(5)$ | $1566.20(17)$ | 0.0 |
| $1598.7(4)$ | $2.3(6)$ | $1780.0(2)$ | $181.1(1)$ |
| $1633(1)^{u}$ | $<0.23$ | $1632.6(2)$ | 0.0 |
| $1638(1)^{u}$ | $<0.23$ | $1638.2(3)$ | 0.0 |
| $1656.6(4)$ | $3.6(6)$ | $1657.3(2)$ | 0.0 |
| $1669(1)^{u}$ | $<0.23$ | $1668.5(3)$ | 0.0 |
| $1684(1)^{u}$ | $<0.23$ | $1683.1(2)$ | 0.0 |
| $1715.4(3)$ | $2.7(6)$ | $1714.9(2)$ | 0.0 |
| $1751.7(4)$ | $0.79(14)^{\dagger}$ | $1932.8(3)$ | $181.1(1)$ |
| $1780.2(8)$ | $0.9(6)$ | $1780.0(2)$ | 0.0 |
| $1787.2(3)$ | $2.3(6)$ | $1968.5(2)$ | $181.1(1)$ |
| $1798.3(4)$ | $0.81(15)^{\dagger}$ | $1979.4(3)$ | $181.1(1)$ |
| $1814.4(2)$ | $3.7(6)$ | $1995.5(3)$ | $181.1(1)$ |
| $1856.6(4)$ | $1.5(2)^{\dagger}$ | $2036.8(2)$ | $181.1(1)$ |
| $1878.9(4)$ | $1.0(2)^{\dagger}$ | $2060.0(3)$ | $181.1(1)$ |
| $1934(1)^{u}$ | $<0.28$ | $1932.8(3)$ | 0.0 |
| $1953.7(4)$ | $1.1(2)^{\dagger}$ | $2134.8(3)$ | $181.1(1)$ |
| $1968.6(2)$ | $7(1)$ | $1968.5(2)$ | 0.0 |
| $1980(1)^{u}$ | $<0.28$ | $1979.4(3)$ | 0.0 |
| $1981.1(9)$ | $2.2(9)$ | $2162.2(3)$ | $181.1(1)$ |
| $1990.2(5)$ | $1.0(2)^{\dagger}$ | $2171.3(2)$ | $181.1(1)$ |
| $1996(1)^{u}$ | $<0.28$ | $1995.5(3)$ | 0.0 |
| $2027.8(4)$ | $1.3(2)^{\dagger}$ | $2208.9(3)$ | $181.1(1)$ |
| $2037(1)^{u}$ | $<0.29$ | $2036.8(2)$ | 0.0 |
| $2061(1)^{u}$ | $<0.30$ | $2060.0(3)$ | 0.0 |
| $2136(1)^{u}$ | $<0.31$ | $2134.8(3)$ | 0.0 |
| $2162(1)^{u}$ | $<0.32$ | $2162.2(3)$ | 0.0 |
| $2172(1)^{u}$ | $<0.32$ | $2171.3(3)$ | 0.0 |
| $2210(1)^{u}$ | $<0.33$ | $2208.9(3)$ | 0.0 |

${ }_{148}$ A decay scheme was primarily built upon the 181${ }_{149} \mathrm{keV} E 2$ transition connecting the $2_{1}^{+}$and $0_{1}^{+}$levels. Al${ }_{50}$ most all other excited states that were identified cas${ }_{51}$ cade through this $2_{1}^{+}$level and, as such, their associated ${ }_{152} \gamma$ rays are found to be in coincidence with the strong $181-\mathrm{keV} \gamma$ ray. The one exception to this is the $1657-\mathrm{keV}$ level; $\gamma$ transitions from this level to the ground state and $739-\mathrm{keV}$ level were observed, but there was no ev-


FIG. 2. Portions of the HPGe background-subtracted, $\beta$ gated $\gamma$-singles spectrum from (a) 0 to 600 keV , (b) 600 to 1200 keV , (c) 1200 to 1800 keV , and (d) 1800 to 2400 keV obtained from $\beta$ decay of ${ }^{146} \mathrm{Cs}$. The identified $\gamma$ rays from ${ }^{146} \mathrm{Ba}$ transitions are marked with their measured energies. Gamma rays from the strongest transitions in the long-lived activity of the $A=146$ decay-chain sequence are also indicated. Unmarked $\gamma$ rays were identified as isobaric contaminants in the coincidence data. ${ }_{171} 1_{1}^{-}$, and $3_{1}^{-}$levels. The $4_{1}^{+}$state decays via a $333-\mathrm{keV}$ ${ }_{172} E 2$ transition to the $2_{1}^{+}$level, whereas the $1_{1}^{-}(558 \mathrm{keV}$


FIG. 3. Background-subtracted projection of the $\beta$-gated $\gamma-\gamma$ coincidence matrix, gated on the $181-\mathrm{keV}, 2_{1}^{+} \rightarrow 0_{1}^{+}$transition from (a) 0 to 600 keV , (b) 600 to 1200 keV , (c) 1200 to 1800 keV , and (d) 1800 to 2400 keV . All other excited states have been observed to possess a decay branch through this level.




FIG. 4. Background-subtracted projections of the $\beta-\gamma-\gamma$ coincidence histogram gated on (a) the $333-\mathrm{keV}, 4_{1}^{+} \rightarrow 2_{1}^{+}$transition, (b) the $558-\mathrm{keV}, 1_{1}^{-} \rightarrow 2_{1}^{+}$transition, and (c) the 307$\mathrm{keV}, 3_{1}^{-} \rightarrow 2_{1}^{+}$transition. Gamma rays from ${ }^{146} \mathrm{Ba}$ are labelled by their energies. Random coincidence events from ${ }^{146} \mathrm{Ce}$ and ${ }^{146} \mathrm{Nd} 2_{1}^{+} \rightarrow 0_{1}^{+}$transitions are labelled.

## 193

## 194

C. $\gamma$-ray intensities

The spins of a few low-lying states have been tenta206 207
not been observed, suggesting that, in fact, they are not associated with ${ }^{146} \mathrm{Ba}$. Furthermore, we have identified 14 new $\gamma$ rays. Upper limits have been applied to an additional 17 unobserved $\gamma$ transitions, two of which are listed in the ENSDF adopted list of $\gamma$ rays for ${ }^{146} \mathrm{Ba}$.

While the level scheme has been extended extensively from what was previously known, the highest level observed lies at $\sim 2.2 \mathrm{MeV}$, i.e., $\sim 3 \mathrm{MeV}$ below the neutron separation energy. It is possible that direct $\beta$ feeding to weak states within this energy range occurs which is not measurable with discrete-line spectroscopy. Resolving this issue will require some other technique, such as Total Absorption Gamma-ray Spectroscopy.

$$
224
$$

coefficient for each transition with the BrICC code [31]. The conversion coefficient for the $181-\mathrm{keV}$ transition is $0.241(4)$. As the remainder of observed transitions are greater than 300 keV , conversion coefficients are expected to be negligible.

Relative $\gamma$-ray intensities have been determined from the observed number of counts in the $\beta-\gamma$ singles spectrum, corrected for $\gamma$ - and $\beta$-detection efficiency, such that the $2_{1}^{+} \rightarrow 0_{1}^{+}$is normalized as 100 . The absolute $\gamma$-ray normalization was calculated accounting for the known $\beta$-delayed neutron emission of ${ }^{146} \mathrm{Cs}(14.2 \%$ [30]), and assuming no delayed neutron emission from ${ }^{146} \mathrm{Ba}$. This was achieved from a comparison of the 181$\mathrm{keV}{ }^{146} \mathrm{Ba} \gamma$ ray to the $141-\mathrm{keV} 2_{3}^{-} \rightarrow 2_{1}^{-}$transition in the daughter, ${ }^{146} \mathrm{La}$ [30]. In this procedure, it is assumed that the contribution of ${ }^{146} \mathrm{Ba}$ in the beam was negligible as these ions would be extracted from the gas catcher in a $2^{+}$charge state, wheras a $1^{+}{ }^{146} \mathrm{Cs}$ beam was selected through the separator. Any $\gamma$ decay in ${ }^{146}$ La results from a $\beta$ decay of ${ }^{146} \mathrm{Ba}$ in the $0^{+}$ground state. A low-energy, high-spin $\left(6^{-}\right)$isomer is reported in ${ }^{146} \mathrm{La}$ [32]. It is assumed that this isomer is not populated and the 141keV transition has an absolute intensity $I_{\gamma}=20.2(20) \%$ [30]. The number of efficiency-corrected counts observed in both peaks is given in Table I. Using the adopted value of $I_{\gamma}$ for the $141-\mathrm{keV}{ }^{146} \mathrm{La} \gamma$ ray and the $\beta$-delayed neutron branch, the 'total' number of parent ${ }^{146} \mathrm{Cs}$ decays was determined. The ratio of the $181-\mathrm{keV} \gamma$-ray intensity to this parent population gives the "normalization" for Table I as $0.42(5)$, that is there are $42181-\mathrm{keV} \gamma$ rays per $100{ }^{146} \mathrm{Cs}$ decays.

The intensity balance also allows an estimate of the $\beta$-branch of ${ }^{146} \mathrm{Cs}$ to the ground state of ${ }^{146} \mathrm{Ba}$. Even after correcting for internal conversion, the total identified decay to the ground state is less than the population of ${ }^{146} \mathrm{La}$, so we infer the ground-state feeding in ${ }^{146} \mathrm{Ba}$ to be $<27 \%$. This is only an estimate, as any extra unobserved feeding to the ground state from high-lying states in ${ }^{146} \mathrm{Ba}$ will reduce this number. We have examined the distribution of $\beta$-feeding to the states we have observed. This was estimated by studying the intensity balance of $\gamma$ rays populating and depopulating each level. This approach is limited by the completeness of the level scheme; if low-intensity transitions from high-lying states are missed, then this will distort the inferred feeding. An indication of the level of "missing" $\gamma$-ray strength can be seen through the $\sim 2 \%$ population of low-lying $4^{+}$and $3^{-}$ states which are forbidden decays, so should receive very little direct $\beta$-population. Thus, the $\beta$-feeding intensities in Fig. 5 are shown as upper limits. The key observation is that the feeding pattern is very widely distributed and no individual high-lying state is strongly populated. Clearly, there is little overlap between the wave function of the ${ }^{146} \mathrm{Cs}$ ground state and any of the excited levels in ${ }^{146} \mathrm{Ba}$.

A summary of the data, including $\gamma$-ray energies and intensities is provided in Table I. For some levels that were identified from $\gamma-\gamma$ coincidence data, the corre-


FIG. 5. Decay scheme of ${ }^{146} \mathrm{Ba}$ populated in ${ }^{146} \mathrm{Cs} \beta^{-}$decay. In total, 31 excited states with $54 \gamma$-ray transitions have been identified. Labels indicate the energy and relative intensity of each transition. For absolute intensity per 100 decays, multiply by $0.42(5) . \mathrm{I}_{\beta^{-}}$values were determined by an intensity balance between the $\gamma$ rays feeding and de-exciting each level, as discussed in the text.
sponding transition to the ground state was not observed 319 in the singles data. An upper limit on the relative in- ${ }^{320}$ tensity of such transitions has been determined using the ${ }_{321}$ intensity of the weakest $\gamma$ ray that was observable. These ${ }^{322}$ have not been included in intensity balances or normal- ${ }^{323}$ ization.

## D. $J^{\pi}$ assignments

Spin values can be constrained for many ${ }^{146} \mathrm{Ba}$ excited states observed from detailed inspection of $\gamma$-ray transitions to levels with firm spin-parity assignments and $\beta$-decay selection rules. The data were not sufficient to confirm these assignments through $\gamma-\gamma$ directional correlation measurements. The $J^{\pi}=1^{-}$spin and parity of the parent is well known, having been measured via high-resolution laser spectroscopy [33].
It is expected that the observed levels in ${ }^{146} \mathrm{Ba}$ are mostly populated via allowed $\left(1^{-} \rightarrow 0^{-}, 1^{-}, 2^{-}\right)$or firstforbidden decays $\left(1^{-} \rightarrow 0^{+}, 1^{+}, 2^{+}, 3^{+}\right)$. Observation (or non-observation) of $\gamma$ transitions to the $0^{+}$ground state can be used to further constrain the spin assignment. Upper limits for relative intensities of unobserved $\gamma$ transitions have been discussed above. The yrast levels lying below $1-\mathrm{MeV}$ excitation have been reported in angular-correlation measurements [27]. The non-yrast states above 1 MeV typically decay via low-multiplicity cascades through the $2_{1}^{+}$level. States of $J=1$ or $J=2$ are also seen to decay directly to the ground state.

- The 1115- and 1255-keV levels Excited states at 1115 and 1255 keV both $\gamma$ decay to the $2_{1}^{+}$and $0_{1}^{+}$levels, therefore $J=1,2$ assignments are possible. No transitions to the negativeparity states were observed, suggesting that these are positive-parity states. The IBA-1 calculations (discussed below) also predict that the $2_{2}^{+}$lies at 1101 keV . We assign the $1115-\mathrm{keV}$ and $1255-\mathrm{keV}$ levels to be the $2_{2}^{+}$level and $2_{3}^{+}$level, respectively.
- The $1310-k e V$ level

The $1310-\mathrm{keV}$ state feeds the $4_{1}^{+}$and $2_{1}^{+}$levels with no observed direct feeding to the $0_{1}^{+}$level. Given that the $\gamma$ transitions only involve positive-parity states, we suggest this is the $3_{1}^{+}$level.

- The 1342-keV level

Since the only $\gamma$ transition from the $1342-\mathrm{keV}$ level is to the $2_{1}^{+}$state, we assign this to be the $0_{2}^{+}$level. The $\beta$-feeding is large enough that, if this were not a $0^{+}$state, $\gamma$ transitions to other levels would be expected to have intensities above the upper limit for non-observation.

- The 1398-keV level

This excited state exhibits $\gamma$-decay characteristics similar to those of the $1115-$ and $1255-\mathrm{keV}$ levels, therefore a $2^{+}$spin-parity assignment is appropriate.

- The 1512-keV level

We propose a $1^{-}$assignment to this state since it exhibits strong feeding to other low-spin ( $J=$ $0,1,2)$ states of both positive and negative parity, with an enhanced branch to the $1_{1}^{-}$state.

- The 1529-keV level

This level decays to the $1_{1}^{-}$and $2_{1}^{+}$levels, with no observed direct feeding to the ground state. A $2_{1}^{-}$ assignment is allowed, however the strong branch to the $2_{1}^{+}$level favors a spin assignment of $J=3$.

- The 1566- and 1715-keV levels

Strong $\beta$ feeding and subsequent $\gamma$ decays to all low-lying yrast states imply a uniquely constrained $J^{\pi}=2^{+}$spin-parity for these levels.

For the remaining states, it has not been possible to draw any solid conclusion pertaining to their appropriate spin-parity assignments. In a few cases, a higher spin assignment is favored since no decay to the ground state was observed. However, the $\beta$ feeding is weak and so it was not possible to ascertain whether the $\gamma$ transition does not exist, or lies below the observation limit of the data.

## IV. DISCUSSION

With strong octupole correlations prevalent in this re-

$$
428
$$

## 429

was only partially successful. The present data set has revealed many new, higher-lying states which do not inform this particular aspect. The data were insufficient for $\gamma-\gamma$ directional correlation measurements. However, the enhanced sensitivity does offer the opportunity for observation of some new low-intensity decays between key low-lying states which constrain their possible spins, sometimes uniquely. There is strong evidence that the 181-, 1115-, and 1255-keV levels are the $J^{\pi}=2_{1}^{+}, 2_{2}^{+}$, and $2_{3}^{+}$states. We use these assignments in the following discussion. The remaining uncertainty with these assignments is the observation of several other low-lying states which are interspersed between these levels and for which a firm spin assignment could not be made. As such, the possibility that these are additional $J^{\pi}=2^{+}$levels cannot be ruled out.

The decay scheme of ${ }^{146} \mathrm{Ba}$ was investigated within the framework of the Interacting Boson Approximation (IBA) [34] by Scott et al., [27] and, more recently, by Gupta and Saxena [35]. Both of these studies used a $\chi$-parameter of $\chi=-\sqrt{7} / 2$, which corresponds to an axially symmetric potential in the $\gamma$ degree of freedom centered at $\gamma=0^{\circ}$. A general study of the $N=90$ transition region in the IBA [36] indicates that this is unlikely to be the case.
Truncated level schemes of the lowest members of the ground-state, $\beta$ - and $\gamma$-vibrational bands for ${ }^{150} \mathrm{Nd}[6]$, ${ }^{148} \mathrm{Ce}[37]$, and ${ }^{146} \mathrm{Ba}$ (this work) are presented in Fig. 6. The key signature of non-axial behavior in ${ }^{146} \mathrm{Ba}$ lies in the location of the $2_{2}^{+}$state at 1115 keV with respect to the $0_{2}^{+}$and $2_{3}^{+}$levels. Indeed, in line with the systematics emerging from Fig. 6 and Fig. 7, we interpret the $2_{2}^{+}$ level as the bandhead of the $\gamma$-vibrational sequence and associate the $0_{2}^{+}$and $2_{3}^{+}$with the (quasi $-\beta$ ) band. Hence, the $\gamma$ and $\beta$ excitations lie remarkably close in energy, to the extent that the " $2_{\beta}^{+"}$ and " $2_{\gamma}^{+}$" locations are reversed with respect to the heavier isotones. However, these assignments, and the association of a projection of angular momentum on the axis of deformation, $K$, are only rigorously applicable for axially symmetric nuclei. In fact, in any non-axially symmetric case, these states mix, especially in a case like this where the moments of inertia would suggest that their unperturbed positions are nearly degenerate. The relative lowering of the excitation energies of the $J^{\pi}=2_{\gamma}^{+}$levels in ${ }^{146} \mathrm{Ba}$ and ${ }^{148} \mathrm{Ce}$, shown by the ratio $\mathrm{E}\left(2_{\gamma}^{+}\right) / \mathrm{E}\left(2_{1}^{+}\right)=6.2$ and 6.3 , respectively, can be compared with 8.2 in ${ }^{150} \mathrm{Nd}$ and 8.9 in ${ }^{152} \mathrm{Sm}$. This is an indication that the triaxial potential energy is soft for ${ }^{146} \mathrm{Ba}$ and ${ }^{148} \mathrm{Ce}$. Such an observation tends to disfavor any interpretation in terms of the $\mathrm{X}(5)$ geometric model, which is based on a stiff, axially-symmetric potential in the $\gamma$ degree of freedom.
In an effort to better understand this evolution of structure in barium, and indeed along $N=90$, IBA calculations were performed. The simplest version of the model was used, which makes no distinction between proton and neutron bosons (IBA-1), and employed the Extended Consistent-Q Formalism (ECQF) [38]. The entire


FIG. 6. [Color online] Truncated level schemes showing the lowest-lying members of the ground-state, $\beta$ - and $\gamma$-vibrational bands in (a) ${ }^{150} \mathrm{Nd}$, (b) ${ }^{148} \mathrm{Ce}$, and (c) ${ }^{146} \mathrm{Ba}$. The individual band sequences are labelled for each $N=90$ isotope.
 43

## 432

43 434
$\qquad$
IBA space can be described with a two-parameter Hamil32 tonian incorporating a term related to the $\beta$ deformation, 3 , and one associated with the degree of axial asymmetry, ${ }^{34} \chi$. The IBA-1 Hamiltonian is given by $[39,40]$ :

$$
\begin{equation*}
H_{\mathrm{IBA}-1}(\zeta)=c\left[(1-\zeta) \hat{n}_{d}-\frac{\zeta}{4 N_{B}} \hat{Q}^{\chi} \cdot \hat{Q}^{\chi}\right] \tag{1}
\end{equation*}
$$

[^0]\[

$$
\begin{equation*}
\hat{Q}^{\chi}=\left(s^{\dagger} \tilde{d}+d^{\dagger} s\right)+\chi\left(d^{\dagger} \tilde{d}\right)^{(2)} \tag{2}
\end{equation*}
$$

\]

36 and $\hat{n}_{d}=d^{\dagger} \cdot \tilde{d}$. The parameters for the fits are included ${ }_{37}$ in Table II. A comparison between the experimental and ${ }_{38}$ calculated low-lying level energies is given in Fig. 7. The calculations are in excellent agreement with the data, 40 agreeing usually at the $10 \%$ level or better, with the best ${ }^{41}$ fit for ${ }^{146} \mathrm{Ba}$ corresponding to a $\gamma$-soft shape.

TABLE II. Parameters $\zeta$ and $\chi$ used for each $N=90$ isotope in the IBA fits of this work.

| Isotope | $\zeta$ | $\chi$ |
| :---: | :--- | :---: |
| Ba | 0.732 | -0.78 |
| Ce | 0.653 | -0.95 |
| Nd | 0.632 | -1.03 |
| Sm | 0.597 | -1.21 |
| Gd | 0.595 | -1.10 |
| Dy | 0.615 | -0.85 |
| Er | 0.633 | -0.61 |



FIG. 7. [Color online] IBA fits (lines) from this work to experimental data (symbols) for the $N=90$ isotones.


FIG. 8. [Color online] Trajectories within the IBA symmetry triangle for the $N=90$ isotonic chain, mapped according to the polar coordinate system of [36]. The slanting lines enclose the region of phase coexistence and phase transition.

Figure 8 highlights the evolution within the so-called 'Casten triangle' [41] of the $N=90$ isotones from ${ }^{146} \mathrm{Ba}$ to ${ }^{158}$ Er. Only with $Z=62,64$ (samarium and gadolinium) are the $N=90$ isotones near the axial $(\chi \sim-1.32)$ route from $\mathrm{U}(5)$ to $\mathrm{SU}(3)$. Both heavier and lighter isotones are best fitted with parameters deep in the interior of the triangle, that is, they follow the trend to deformation along loci corresponding to non-axial shapes. Both above and below $Z=64$, the trend of non-axial behavior seems to be quite symmetric.

The near degeneracy of the $J^{\pi}=2_{2}^{+}$and $0_{2}^{+}$levels is quite rare and has been discussed as a possible signature for nuclei with properties lying along the so-called 457

$$
460
$$

$$
471
$$

number of nuclei which have statistically regular spectra that are found in the mainly chaotic IBA parameter space. An experimental signature of nuclei which may exhibit this regular behavior has been defined as those having $\left|E\left(2_{2}^{+}\right)-E\left(0_{2}^{+}\right)\right| / E\left(2_{2}^{+}\right) \leq 0.025$ [43]. In ${ }^{146} \mathrm{Ba}$, this quantity is small, 0.055 , but just outside the prediction for identifying nuclei on the non-chaotic arc. However, this simple experimental signature does not always exactly follow the trajectory of "regular" nuclei which are inferred from a full statistical analysis of the spectra [44]. Interestingly, both $N=90{ }^{156} \mathrm{Dy}$ and ${ }^{158} \mathrm{Er}$ [45] have been previously identified as nuclei lying close to the regular region [43]. The fact that ${ }^{146} \mathrm{Ba}$ exhibits a similar degeneracy appears related to $\gamma$ softness and the symmetry of these shapes above and below axially symmetric $Z=62,{ }^{152} \mathrm{Sm}$.

## V. CONCLUSIONS

A detailed $\beta$-decay spectroscopy measurement has 76 been conducted on the neutron-rich exotic nucleus ${ }^{146} \mathrm{Ba}$. This represents the first results from the recentlycommissioned decay-spectroscopy station for low-energy CARIBU beams at Argonne National Laboratory. The experimental arrangement had a high sensitivity to weak $\gamma$-ray transitions and, hence, enabled the study of excited states not strongly populated via $\beta$ decay. Inspection of these low-intensity transitions has allowed spin constraints for low-lying levels, which have also 485 been considered within the IBA framework. The $N=90$ 88 isotones are situated close to, but slightly to the right, 87 of the phase-transitional region predicted by the IBA. ${ }_{88}$ They follow a symmetric behavior about ${ }^{152} \mathrm{Sm}(Z=62)$ 9 which exhibits the highest degree of axial symmetry. Moving away from ${ }^{152} \mathrm{Sm}$, isotones of both larger and 491 smaller $Z$ appear to exhibit increasing $\gamma$ softness.

494 The authors wish to acknowledge the Physics Sup495 port group at Argonne National Laboratory and engineering work of the Submillimeter-Wave Technology Laboratory, University of Massachusetts Lowell. Figure 5 in this article has been created using the LevelScheme scientific figure preparation system [M. A. Caprio, Comput. Phys. Commun. 171, 107 (2005), http://scidraw.nd.edu/levelscheme]. This material is based upon work supported by the U.S. Department of Energy, Office of Science, Office of Nuclear Physics under Grant Nos. DE-FG02-94ER40848 and DE-FG0294ER40834, and Contract No. DE-AC02-06CH11357, and the U.S. National Science Foundation under Grant No. 1064819. This research used resources of ANL's ATLAS facility, which is a DOE Office of Science User Facility.
[1] M. G. Mayer, Phys. Rev. 74, 235 (1948).
[2] B. R. Mottelson, Rev. Mod. Phys. 29, 186 (1957).
[3] C. W. Reich, Nucl. Data Sheets 113, 2537 (2012).
[4] C. W. Reich, Nucl. Data Sheets 110, 2257 (2009).
[5] M. J. Martin, Nucl. Data Sheets 114, 1497 (2013).
[6] S. K. Basu and A. A. Sonzogni, 57 Nucl. Data Sheets 114, 435 (2013).
[7] F. Iachello, Phys. Rev. Lett. 85, 3580 (2000).
[8] F. Iachello, Phys. Rev. Lett. 87, 052502 (2001).
[9] R. F. Casten, Nat. Phys. 2, 811 (2006).
[10] R. F. Casten and N. V. Zamfir, Phys. Rev. Lett. 87, 052503 (2001).
[11] D. Tonev, A. Dewald, T. Klug, P. Petkov, J. Jolie, A. Fitzler, O. Möller, S. Heinze, P. von Brentano, and R. F. Casten, Phys. Rev. C 69, 034334 (2004).
[12] R. Krücken, B. Albanna, C. Bialik, R. F. Casten, ${ }^{58}$ J. R. Cooper, A. Dewald, N. V. Zamfir, C. J. Bar- ${ }_{58}$ ton, C. W. Beausang, M. A. Caprio, A. A. Hecht, 583 T. Klug, J. R. Novak, N. Pietralla, and P. von Brentano, 584 Phys. Rev. Lett. 88, 232501 (2002).
[13] M. A. Caprio, N. V. Zamfir, R. F. Casten, C. J. Barton, ${ }_{586}$ C. W. Beausang, J. R. Cooper, A. A. Hecht, R. Krücken, 587 H. Newman, J. R. Novak, N. Pietralla, A. Wolf, and 588 K. E. Zyromski, Phys. Rev. C 66, 054310 (2002).
[14] W. D. Kulp, J. L. Wood, P. E. Garrett, C. Y. Wu, 59 D. Cline, J. M. Allmond, D. Bandyopadhyay, D. Dash- ${ }^{591}$ dorj, S. N. Choudry, A. B. Hayes, H. Hua, M. G. Mynk, 592 M. T. McEllistrem, C. J. McKay, J. N. Orce, R. Teng, 593 and S. W. Yates, Phys. Rev. C 77, 061301 (2008).
[15] P. E. Garrett, W. D. Kulp, J. L. Wood, D. Bandyopad- ${ }_{595}$ hyay, S. Choudry, D. Dashdorj, S. R. Lesher, M. T. ${ }_{59}$ McEllistrem, M. Mynk, J. N. Orce, and S. W. Yates, ${ }_{597}$ Phys. Rev. Lett. 103, 062501 (2009).
[16] J. F. Sharpey-Schafer, S. M. Mullins, R. A. Bark, J. Kau, 599 F. Komati, E. A. Lawrie, J. J. Lawrie, T. E. Madiba, 600 P. Maine, A. Minkova, S. H. T. Murray, N. J. Ncapayi, 60 and P. A. Vymers, Eur. Phys. J. A 47, 5 (2011).
[17] R. F. Casten and N. V. Zamfir, J. Phys. G Nucl. Partic. 22, 1521 (1996).
[18] T. Nikšić, D. Vretenar, G. A. Lalazissis, and P. Ring, 60 Phys. Rev. Lett. 99, 092502 (2007).
[19] K. Nomura, N. Shimizu, and T. Otsuka, 60 Phys. Rev. Lett. 101, 142501 (2008).
[20] P. A. Butler and W. Nazarewicz, 60 Rev. Mod. Phys. 68, 349 (1996).
[21] S. M. Scott, D. D. Warner, W. D. Hamilton, P. Hungerford, G. Jung, K. D. Wünsch, and B. Pfeiffer, ${ }_{612}$ J. Phys. G: Nucl. Phys. 5, L187 (1979).
[22] J. B. Wilhelmy, S. G. Thompson, R. C. Jared, and 61 E. Cheifetz, Phys. Rev. Lett. 25, 1122 (1970).
[23] W. R. Phillips, I. Ahmad, H. Emling, R. Holzmann, 616 R. V. F. Janssens, T. L. Khoo, and M. W. Drigert, ${ }_{617}$ Phys. Rev. Lett. 57, 3257 (1986).
[24] W. Urban, M. A. Jones, J. L. Durell, M. Leddy, W. R. Phillips, A. G. Smith, B. J. Varley, I. Ahmad, L. R.84[
[31] T. Kibédi, T. W. Burrows, M. B. Trzhaskovskaya, P. M. Davidson, and C. W. Nestor Jr., Nucl. Instrum. Meth. A 589, 202 (2008).
[32] C. Chung, W. B. Walters, D. S. Brenner, R. L. Gill, M. Shmid, Y. Y. Chu, R. E. Chrien, L.-J. Yuan, F. K. Wohn, and R. A. Meyer, Phys. Rev. C 31, 2199 (1985).
[33] A. Coc, C. Thibault, F. Touchard, H. T. Duong, P. Juncar, S. Liberman, J. Pinard, M. Carre, J. Lerme, J. L. Vialle, S. Buttgenbach, A. C. Mueller, and A. Pesnelle, Nucl. Phys. A 468, 1 (1987).
[34] A. Arima and F. Iachello, Phys. Rev. Lett. 35, 1069 (1975).
[35] J. B. Gupta and M. Saxena, Phys. Rev. C 91, 054312 (2015).
[36] E. A. McCutchan, N. V. Zamfir, and R. F. Casten, Phys. Rev. C 69, 064306 (2004).
[37] N. Nica, Nucl. Data Sheets 117, 1 (2014).
[38] P. O. Lipas, P. Toivonen, and D. D. Warner, Phys. Lett. B 155, 295 (1985).
[39] N. V. Zamfir, P. von Brentano, R. F. Casten, and J. Jolie, Phys. Rev. C 66, 021304(R) (2002).
[40] V. Werner, P. von Brentano, R. F. Casten, and J. Jolie, Phys. Lett. B 527, 55 (2002).
[41] R. F. Casten and D. D. Warner, Progress in Particle and Nuclear Physics 9, 311 (1983).
[42] Y. Alhassid and N. Whelan, Phys. Rev. Lett. 67, 816 (1991).
[43] J. Jolie, R. F. Casten, P. Cejnar, S. Heinze, E. A. McCutchan, and N. V. Zamfir, Phys. Rev. Lett. 93, 132501 (2004).
[44] D. Bonatsos, E. A. McCutchan, and R. F. Casten, Phys. Rev. Lett. 104, 022502 (2010).
[45] R. G. Helmer, Nucl. Data Sheets 101, 325 (2004).


[^0]:    ${ }_{35}$ where

