



This is the accepted manuscript made available via CHORUS. The article has been published as:

γ-soft  $^{146}$ Ba and the role of nonaxial shapes at N $\approx$ 90 A. J. Mitchell *et al.* 

Phys. Rev. C 93, 014306 — Published 12 January 2016

DOI: 10.1103/PhysRevC.93.014306

# $\gamma$ -soft $^{146}\mathrm{Ba}$ and the role of non-axial shapes at $N\sim90$

A. J. Mitchell, <sup>1, a</sup> C. J. Lister, <sup>1</sup> E. A. McCutchan, <sup>2</sup> M. Albers, <sup>3, b</sup> A. D. Ayangeakaa, <sup>3</sup> P. F. Bertone, <sup>3, c</sup> M. P. Carpenter, C. J. Chiara, 4, 4 P. Chowdhury, J. A. Clark, P. Copp, H. M. David, e A. Y. Deo, 1, f B. DiGiovine,<sup>3</sup> N. D'Olympia,<sup>1, g</sup> R. Dungan,<sup>5</sup> R. D. Harding,<sup>1, 6, h</sup> J. Harker,<sup>3, 4</sup> S. S. Hota,<sup>1, i</sup> R. V. F. Janssens,<sup>3</sup> F. G. Kondev,<sup>7</sup> S. H. Liu,<sup>8,9,j</sup> A. V. Ramayya,<sup>10</sup> J. Rissanen,<sup>11,k</sup> G. Savard,<sup>3,12</sup> D. Seweryniak, R. Shearman, A. A. Sonzogni, S. L. Tabor, W. B. Walters, E. Wang, and S. Zhu<sup>3</sup> <sup>1</sup>Department of Physics and Applied Physics, University of Massachusetts Lowell, Lowell, Massachusetts 01854 <sup>2</sup>National Nuclear Data Center, Brookhaven National Laboratory, Upton, New York 11973 <sup>3</sup>Physics Division, Argonne National Laboratory, Argonne, Illinois 60439 <sup>4</sup>Department of Chemistry and Biochemistry, University of Maryland, College Park, Maryland 20742 <sup>5</sup>Physics Department, Florida State University, Tallahassee, Florida 32306 <sup>6</sup>Department of Physics, University of Surrey, Guildford GU2 7XH, UK <sup>7</sup> Nuclear Engineering Division, Argonne National Laboratory, Argonne, Illinois 60439 <sup>8</sup>Department of Chemistry, University of Kentucky, Lexington, Kentucky 40506 <sup>9</sup>Department of Physics and Astronomy, University of Kentucky, Lexington, Kentucky 40506 <sup>10</sup>Physics Department, Vanderbilt University, Nashville, Tennessee 37235 <sup>11</sup>Lawrence Berkeley National Laboratory, Berkeley, California 94720 <sup>12</sup>Department of Physics, University of Chicago, Chicago, Illinois 60637 (Dated: December 9, 2015)

Low-spin states in the neutron-rich, N=90 nuclide  $^{146}$ Ba were populated following  $\beta$ -decay of  $^{146}$ Cs, with the goal of clarifying the development of deformation in barium isotopes through delineation of their non-yrast structures. Fission fragments of  $^{146}$ Cs were extracted from a 1.7-Ci  $^{252}$ Cf source and mass-selected using the CARIBU facility. Low-energy ions were deposited at the center of a box of thin  $\beta$  detectors, surrounded by a high-efficiency HPGe array. The new  $^{146}$ Ba decay scheme now contains 31 excited levels extending up to  $\sim 2.5$  MeV excitation energy, double what was previously known. These data are compared to predictions from the Interacting Boson Approximation (IBA) model. It appears that the abrupt shape change found at N=90 in samarium and gadolinium is much more gradual in barium and cerium, due to an enhanced role of the  $\gamma$  degree of freedom.

PACS numbers: 23.40.-s, 21.60.Fw, 23.20.Lv

10

11

12

13

14

15

16

17

18

19

20

Keywords:  $\beta$  decay;  $\gamma$ -ray spectroscopy;  $E_{\gamma}$ ,  $I_{\gamma}$ ; N=90 shape transition; Interacting Boson Approximation

## I. INTRODUCTION

<sup>a</sup> Present address: Department of Nuclear Physics, Australian National University, Canberra, ACT 2601, Australia; Email: aj.mitchell@anu.edu.au

b Present address: Ernst & Young GmbH, Wirtschaftspruefungsgesellschaft, Mergenthalerallee 3-5, D-65760 Eschborn, Germany
 c Present address: Marshall Space Flight Center, Huntsville, Alabama 35812

d Present address: U.S. Army Research Laboratory, Adelphi, Marvland 20783

e Present address: GSI Helmholtzzentrum für Schwerionenforschung GmbH, D-64291 Darmstadt, Germany

f Present address: Department of Physics, Indian Institute of Technology Roorkee, Roorkee 247 667, India

g Present address: Passport Systems Inc., 70 Treble Cove Road, 1st Floor, Billerica, Massachusetts 01862

h Present address: Department of Physics, University of York, Heslington, York, YO10 5DD, UK

lington, York, YO<br/>10 5DD, UK  $^{\rm i}$  Present address: Department of Nuclear Physics, Australian Na-

tional University, Canberra, ACT 2601, Australia <sup>j</sup> Present address: West Physics, 3825 Paces Walk SE, Suite 250, Atlanta, Georgia 30339

k Present address: Fennovoima Oy, Salmisaarenaukio 1, 00180 Helsinki, Finland

<sup>1</sup> Present address: National Physical Laboratory, Hampton Road, Teddington, Middlesex TW11 0LW, UK

The transition from spherical, shell-model-like behav-24 ior, to deformed collective motion has always been inter-25 esting, yet controversial, in nuclear structure. Although 26 models exist for each extreme [1, 2], the actual transi-27 tion from one limit to the other remains confused and 28 lacks a ubiquitous description. Stable isotopes of rare-<sub>29</sub> earth elements near Z=64 with N=90 (e.g. <sup>156</sup>Dy 30 (Z = 66) [3], <sup>154</sup>Gd (Z = 64) [4], <sup>152</sup>Sm (Z = 62) [5], and  $^{150}$ Nd (Z=60) [6]) exhibit remarkable similarities in 32 the excitation energies of ground-state bands and excited 33  $J^{\pi} = 0^+$  and  $J^{\pi} = 2^+$  sequences. The abrupt onset of de-34 formation has received particularly intense scrutiny with 35 general discussions often framed in terms of a phase transition [7–9]; in this case, a specific type of phase transition  $_{37}$  encapsulated by the X(5) model [10–13]. However, such 38 an approach is not fully supported by all available ex-39 perimental data and more generalized shape-coexistence 40 models have been proposed [14–16].

A way to clarify this issue is to widen the scope of investigation to both heavier and lighter nuclei. In a general sense, the behavior of transitional nuclei is expected to follow the number of valence particles, as predicted in the

45  $N_p N_n$  scheme of Casten [17]. In practice, the underlying 46 fermionic structure appears to be important, with resid-47 ual interactions between protons and neutrons in specific 48 orbits playing a key role in 'tipping' the nuclear shape 49 from spherical to deformed [18, 19]. In this way, the N = 90 border between shapes retains its significance, 51 although the sharpness of the transition becomes more 52 muted. The nuclei which have been most extensively 53 studied are all stable, but it is relevant to enquire about 54 how the transitional structures evolve as one progresses  $^{55}$  up to  $^{158}{\rm Er}$  (Z = 68), and  $^{160}{\rm Yb}$  (Z = 70) or down to  $^{56}$   $^{148}{\rm Ce}$  (Z = 58), and  $^{146}{\rm Ba}$  (Z = 56). The lighter nuclei 57 in this sequence are quite neutron-rich and cannot be 58 accessed by fusion-evaporation reactions, and so fission-<sub>59</sub> fragment spectroscopy and  $\beta$  decay are the appropriate probes.

Nuclei in this region are also expected to exhibit strong 62 octupole correlations [20]. Polarization of spin-orbit by partners appears to quench the Z=64 sub-shell closure, 64 resulting in strong couplings between  $\Delta J = \Delta L = 3$  nucleon orbitals  $(\pi d_{5/2} - \pi h_{11/2} \text{ and } \nu f_{7/2} - \nu i_{13/2})$ . The onset of octupole collectivity in Gd, Sm, and Nd is observed between N = 88 and N = 90. The Ba isotopes undergo a smoother transition between N=86 $_{69}$  and N=88, two neutrons earlier than expected from <sub>70</sub> the behavior found in the Z=60 to Z=64 range [21]. <sub>100</sub> High-Purity Ge (HPGe) clover detectors. The measure-81 which inform this discussion.

#### EXPERIMENTAL METHOD

The data presented here were obtained at the CAlifornium Rare Ion Breeder Upgrade (CARIBU [28]) facility at Argonne National Laboratory. Spontaneous fission fragments extracted from a 1.7-Ci <sup>252</sup>Cf source were thermalized in a gas catcher, in which interactions with high-purity He gas and with RF and DC fields combine to result in a low-emittance beam. An isotopically-pure <sub>90</sub> beam of singly-charged <sup>146</sup>Cs nuclei was selected by the isobar separator. The beam was cooled to  $\sim 2~{\rm keV}$  and bunched before delivery to the low-energy experimental area. Approximately 300 ions/s were delivered to the new 123 decay-spectroscopy station, where they were implanted on an aluminum foil located at the center of an array of 124  $\gamma$ -ray and  $\beta$ -particle detectors.

(Scintillator And Tape Using Radioactive Nuclei) system 127 146 isobars along the  $\beta^-$  decay chain towards stability coupled to the X-Array, a highly-efficient array of five  $_{128}$  ( $^{146}\mathrm{Ba} \rightarrow ^{146}\mathrm{La} \rightarrow ^{146}\mathrm{Ce} \rightarrow ^{146}\mathrm{Nd}$ ) were recorded in

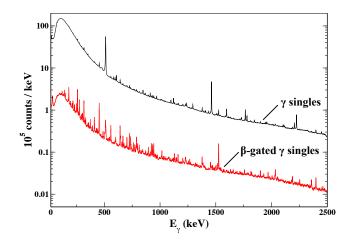


FIG. 1. [Color online] Total HPGe  $\gamma$ -ray energy spectra measured from  $^{146}\mathrm{Cs}$   $\beta$  decay. The black upper spectrum is ungated  $\gamma$ -singles and the red (light gray) lower spectrum is  $\beta$ -gated  $\gamma$ -singles for the same data. By only selecting time-correlated  $\beta$ - $\gamma$  events, overall background is suppressed by up to two orders of magnitude and uncorrelated, roombackground  $\gamma$  rays are removed.

Prompt-fission spectroscopy of <sup>146</sup>Ba (for example [22- 101 ment described in this article utilized the 'Mark-I' detec-24]) has identified the ground-state and negative-parity 102 tor chamber with the 'paddle' scintillator arrangement bands to moderate spins. Octupole deformation in <sup>146</sup>Ba <sub>103</sub> (see Ref. [29] for a description of the experimental set-<sup>74</sup> has been discussed [25–27], with the suggestion that these 104 up). These data were measured as part of the commis-75 effects are weak and disappear at medium to high spins 105 sioning for the new decay station. The catcher foil was <sub>76</sub> [24]. Although the yrast states in <sup>146</sup>Ba are established, <sub>106</sub> located at the geometric center of four symmetrically-77 there is limited information pertaining to the non-yrast, 107 arranged plastic scintillator paddles, and replaced perilow-spin levels. This paper reports on the  $\beta$ -decay of 108 odically to reduce the build up of decay chain activity <sup>146</sup>Cs, with focus on identifying and quantifying proper- <sup>109</sup> over time. Each paddle was positioned in front of a sin-80 ties of the important low-spin, non-yrast states in 146Ba 110 gle clover in the vertical plane of the X-Array, offering 111 large solid angle coverage. Output energy signals from 112 each of the clover crystals and four scintillator preamps were fed directly into a digital data acquisition system.

> The energy spectrum of  $\gamma$  rays detected by the X-Array 115 is presented in Fig. 1. The black upper spectrum represents the ungated  $\gamma$ -ray singles spectrum for all the data. The red (light gray) lower spectrum corresponds to Ge clover events detected in coincidence with an event in 119 a scintillator paddle. Energy and efficiency calibration 120 of the X-Array was determined using standard <sup>152</sup>Eu,  $^{121}$   $^{182}$ Ta, and  $^{243}$ Am sources.

#### RESULTS III.

## $\gamma$ -ray identification

For  $^{146}$ Cs, the decay half-live, 0.321(2) s, and  $\beta$ -125 delayed neutron branching ratio, 14.2(5)%, are well The CARIBU decay station consists of the SATURN  $_{126}$  known [30]. Gamma rays from each of the A = 129 the singles data. Those associated with the de-excitation <sub>130</sub> of <sup>146</sup>Ba were identified using a combination of  $\beta$ - $\gamma$  and  $_{131}$   $\beta$ - $\gamma$ - $\gamma$  coincidence events. Contamination from long-lived 132 activity was strongest in the even-even isobars, <sup>146</sup>Ce <sup>133</sup> and <sup>146</sup>Nd. The odd-odd isobars, <sup>146</sup>La and <sup>146</sup>Pr, were 134 highly fragmented and as such, their relative  $\gamma$ -ray inten-135 sities are weak. A small contribution from <sup>145</sup>Ba, from <sub>136</sub> the  $\beta$ -delayed neutron emission of <sup>146</sup>Cs, was also de-137 tectable. The background-subtracted,  $\beta$ -gated  $\gamma$ -ray sin-138 gles spectrum can be found in Fig. 2. Gamma rays that  $_{139}$  have been identified as transitions in  $^{146}\mathrm{Ba}$  are labelled 140 by their energies. The strongest transitions from the sub-141 sequent decay chain are also labelled. The full range of the energy spectrum in this measurement was  $\sim 3$  MeV. A short test has since been conducted with the range  $_{144}$  extended to  $\sim 10$  MeV; there was no evidence of any further strong, direct decays to the ground state beyond the 146 range of the original experiment.

TABLE I: Observed  $\gamma$ -ray transitions in  $^{146}$ Ba placed in the level scheme of Fig. 5. Relative intensities,  $I_{\gamma}$ , are normalized to the 181-keV  $\gamma$  ray, taken as 100. For absolute intensity per 100 parent decays, the relative intensity should be multiplied by 0.42(5). The method for determining the normalization using the 141-keV  $\gamma$  ray from  $^{146}$ La [30] is described in the text (for reference, the relative intensity of this  $\gamma$  ray is included in the table). Strong transitions were calculated from prompt  $\gamma$ -ray singles data; those marked  $^{\dagger}$  are from coincidence data. Upper limits on  $I_{\gamma}$  for transitions from new levels to the ground state that have not been observed, but may occur if  $J_i^{\pi} \neq 0$ , are marked  $^u$ . Uncertainties are statistical and based on fitting approximations.

$E_{\gamma}$	$I_{\gamma}$	$E_{ m initial}$	$E_{\rm final}$
(keV)		(keV)	(keV)
140.7(1)	41(2)	-	-
181.3(1)	100(3)	181.1(1)	0.0
307.3(1)	$5.3(4)^{\dagger}$	821.6(2)	513.9(2)
332.9(1)	13(2)	513.9(2)	181.1(1)
558.1(1)	23(1)	739.4(1)	181.1(1)
639.9(1)	4.4(3)	821.6(2)	181.1(1)
739.1(2)	5.1(6)	739.4(1)	0.0
743.6(6)	2.7(6)	1566.2(2)	821.6(2)
772.2(1)	5.2(6)	1511.7(2)	739.4(1)
788.9(1)	$0.50(6)^{\dagger}$	1529.1(1)	739.4(1)
795.6(2)	2.0(7)	1309.5(3)	513.9(2)
816.6(6)	1.2(5)	1638.2(3)	821.6(2)
827.3(4)	2.2(9)	1566.2(2)	739.4(1)
871.3(1)	3.4(6)	1052.4(3)	181.1(1)
892.9(4)	$1.24(11)^{\dagger}$	1632.6(2)	739.4(1)
894.1(1)	$0.21(4)^{\dagger}$	1714.9(2)	821.6(2)
918.7(3)	1.4(5)	1657.3(2)	739.4(1)
933.1(1)	$5.2(4)^{\dagger}$	1114.7(2)	181.1(1)
943.6(2)	1.1(1)	1683.1(2)	739.4(1)
976.7(1)	2.6(8)	1714.9(2)	739.4(1)
1052.7(4)	1.5(7)	1566.2(2)	513.9(2)
1073.5(2)	3.5(7)	1255.4(2)	181.1(1)
1115.2(3)	2.9(5)	1114.7(2)	0.0
1128.4(1)	$2.7(2)^{\dagger}$	1309.5(3)	181.1(1)
1160.9(1)	$1.2(1)^{\dagger}$	1342.0(3)	181.1(1)

 $TABLE\ I-continued$ 

	TADLE I	- continued	
$E_{\gamma}$	$I_{\gamma}$	$E_{ m initial}$	$E_{\rm final}$
(keV)		(keV)	(keV)
1217(1)	0.6(5)	1397.8(2)	181.1(1)
1229.5(2)	$0.58(9)^{\dagger}$	1410.8(3)	181.1(1)
1256.1(3)	3.1(6)	1255.4(2)	0.0
1299(1)	0.8(4)	2036.8(2)	739.4(1)
$1310(1)^u$	< 0.19	1309.5(3)	0.0
1330.4(2)	1.6(5)	1511.7(2)	181.1(1)
$1342(2)^{u}$	< 0.19	1342.0(3)	0.0
1348.9(3)	1.6(5)	1529.1(2)	181.1(1)
1385.6(2)	4.3(7)	1566.2(2)	181.1(1)
1397.8(4)	1.1(6)	1397.8(2)	0.0
$1412(1)^{u}$	< 0.20	1410.8(3)	0.0
1451.8(1)	$0.83(12)^{\dagger}$	1632.6(2)	181.1(1)
1457.0(2)	$3.3(7)^{'}$	1638.2(3)	181.1(1)
1487.4(4)	2(1)	1668.5(2)	181.1(1)
1502.5(2)	$2.8(2)^{\dagger}$	1683.1(1)	181.1(1)
1510(1)	0.9(5)	1511.7(2)	0.0
$1529(1)^u$	< 0.21	1529.1(2)	0.0
1533.7(5)	1.5(9)	1714.9(2)	181.1(1)
1566.7(3)	2.6(5)	1566.20(17)	0.0
1598.7(4)	2.3(6)	1780.0(2)	181.1(1)
$1633(1)^{u'}$	< 0.23	1632.6(2)	0.0
$1638(1)^u$	< 0.23	1638.2(3)	0.0
1656.6(4)	3.6(6)	1657.3(2)	0.0
$1669(1)^{u}$	< 0.23	1668.5(3)	0.0
$1684(1)^u$	< 0.23	1683.1(2)	0.0
1715.4(3)	2.7(6)	1714.9(2)	0.0
1751.7(4)	$0.79(14)^{\dagger}$	1932.8(3)	181.1(1)
1780.2(8)	$0.9(6)^{'}$	1780.0(2)	0.0
1787.2(3)	2.3(6)	1968.5(2)	181.1(1)
1798.3(4)	$0.81(15)^{\dagger}$	1979.4(3)	181.1(1)
1814.4(2)	$3.7(6)^{'}$	1995.5(3)	181.1(1)
1856.6(4)	$1.5(2)^{\dagger}$	2036.8(2)	181.1(1)
1878.9(4)	$1.0(2)^{\dagger}$	2060.0(3)	181.1(1)
$1934(1)^{u}$	< 0.28	1932.8(3)	0.0
1953.7(4)	$1.1(2)^{\dagger}$	2134.8(3)	181.1(1)
1968.6(2)	7(1)	1968.5(2)	0.0
$1980(1)^u$	< 0.28	1979.4(3)	0.0
1981.1(9)	2.2(9)	2162.2(3)	181.1(1)
1990.2(5)	$1.0(2)^{\dagger}$	2171.3(2)	181.1(1)
$1996(1)^{u}$	< 0.28	1995.5(3)	0.0
2027.8(4)	$1.3(2)^{\dagger}$	2208.9(3)	181.1(1)
$2037(1)^u$	< 0.29	2036.8(2)	0.0
$2061(1)^u$	< 0.30	2060.0(3)	0.0
$2136(1)^u$	< 0.31	2134.8(3)	0.0
$2162(1)^u$	< 0.32	2162.2(3)	0.0
$2172(1)^u$	< 0.32	2171.3(3)	0.0
$2210(1)^u$	< 0.33	2208.9(3)	0.0
		- (-/	

147

A decay scheme was primarily built upon the 181-  $^{149}$  keV E2 transition connecting the  $2_1^+$  and  $0_1^+$  levels. Al-  $^{150}$  most all other excited states that were identified cas-  $^{151}$  cade through this  $2_1^+$  level and, as such, their associated  $^{152}$   $\gamma$  rays are found to be in coincidence with the strong  $^{153}$  181-keV  $\gamma$  ray. The one exception to this is the 1657-keV  $^{154}$  level;  $\gamma$  transitions from this level to the ground state  $^{155}$  and 739-keV level were observed, but there was no ev-

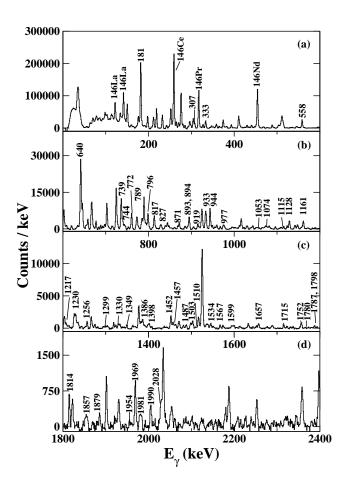


FIG. 2. Portions of the HPGe background-subtracted,  $\beta$ gated  $\gamma$ -singles spectrum from (a) 0 to 600 keV, (b) 600 to 1200 keV, (c) 1200 to 1800 keV, and (d) 1800 to 2400 keV obtained from  $\beta$  decay of  $^{146}\mathrm{Cs.}\,$  The identified  $\gamma$  rays from  $^{146}\mathrm{Ba}\,$ transitions are marked with their measured energies. Gamma rays from the strongest transitions in the long-lived activity of the A = 146 decay-chain sequence are also indicated. Unmarked  $\gamma$  rays were identified as isobaric contaminants in the coincidence data.

156 idence of a transition directly to the  $2_1^+$  state. Twelve 157 levels have decay branches that feed into the 514-, 739-158 or 822-keV levels, which then proceed to decay to the 177 181-keV level or ground state. The remaining 19 lev-160 els were observed to only have decay branches to the 161 181-keV level, and in some cases directly to the ground 162 state. The background-subtracted projection of the  $\beta$ -163 correlated  $\gamma - \gamma$  matrix, gated on the 181-keV transition, 164 is presented in Fig. 3. This was used as the starting point in identifying which  $\gamma$  rays belong to transitions in <sup>146</sup>Ba. There is a small contribution from random coin- $_{\text{167}}$  cidences with strong  $\gamma$  rays of  $^{146}\text{Ce}$  and  $^{146}\text{Nd},~2_{1}^{+}\rightarrow$  $_{168}$   $0_{\rm g.s.}^{+}$  transitions (258 keV and 454 keV).

170 by gating on  $\gamma$  rays above and below the known  $4_1^+$ , 188 decay scheme. This work has identified 19 of these, which 171  $1_1^-$ , and  $3_1^-$  levels. The  $4_1^+$  state decays via a 333-keV 189 were subsequently placed in the new scheme. Thirteen 172 E2 transition to the  $2_1^+$  level, whereas the  $1_1^-$  (558 keV 189 remaining  $\gamma$  rays in the adopted list of Ref. [30] have

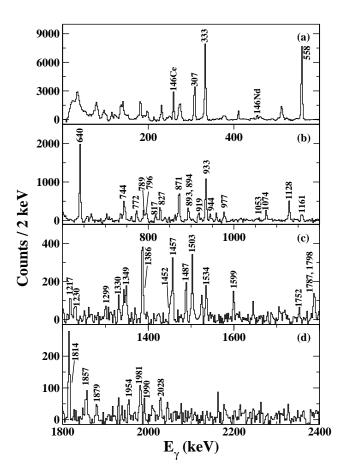


FIG. 3. Background-subtracted projection of the  $\beta$ -gated  $\gamma$ - $\gamma$ coincidence matrix, gated on the 181-keV,  $2_1^+ \rightarrow 0_1^+$  transition from (a) 0 to 600 keV, (b) 600 to 1200 keV, (c) 1200 to 1800 keV, and (d) 1800 to 2400 keV. All other excited states have been observed to possess a decay branch through this level.

 $_{173}$  and 739 keV) and  $3_{1}^{-}$  (307 keV and 640 keV) states each 174 have two decay paths. Figure 4 provides the background-175 subtracted matrix projections with an appropriate gate 176 for each of these levels.

## The decay scheme

The Nuclear Data Sheets list nine confirmed levels,  $_{\mbox{\scriptsize 179}}$  two tentative excited states, and 21  $\gamma\text{-ray}$  transitions for <sup>146</sup>Ba from previous  $\beta$ -decay studies [30]. In this work, we report a total of 31 excited states with 54  $\gamma$ -ray tran-182 sitions, offering a significant increase in the known <sup>146</sup>Ba 183 level structure. Our proposed expansion of the known 184 decay scheme is displayed in Fig. 5.

These data confirm the correct placement of 21  $\gamma$ -ray transitions between the known levels. Thirty-two  $\gamma$  rays Placement of  $\gamma$  rays in the level scheme was confirmed 187 are listed in Ref. [30] without placement in the adopted

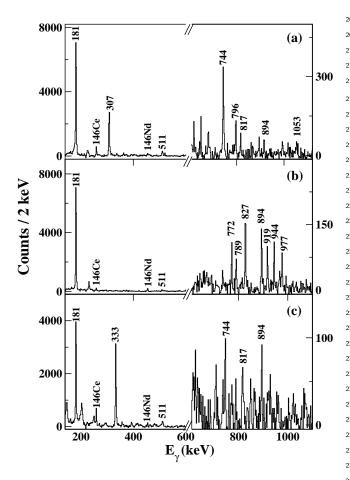


FIG. 4. Background-subtracted projections of the  $\beta$ - $\gamma$ - $\gamma$  coincidence histogram gated on (a) the 333-keV,  $4_1^+ \rightarrow 2_1^+$  tran- $^{146}$ Nd  $2_1^+ \rightarrow 0_1^+$  transitions are labelled.

191 not been observed, suggesting that, in fact, they are not 246 states in 146 Ba will reduce this number. We have examassociated with  $^{146}$ Ba. Furthermore, we have identified  $_{247}$  ined the distribution of  $\beta$ -feeding to the states we have 193 14 new  $\gamma$  rays. Upper limits have been applied to an 248 observed. This was estimated by studying the intensity additional 17 unobserved  $\gamma$  transitions, two of which are 249 balance of  $\gamma$  rays populating and depopulating each level. listed in the ENSDF adopted list of  $\gamma$  rays for <sup>146</sup>Ba.

202 ing this issue will require some other technique, such as 257 in Fig. 5 are shown as upper limits. The key observa-203 Total Absorption Gamma-ray Spectroscopy.

# $\gamma$ -ray intensities

204

The spins of a few low-lying states have been tenta- 263 A summary of the data, including  $\gamma$ -ray energies and 206 tively assigned in the literature. Where possible, these 264 intensities is provided in Table I. For some levels that

208 coefficient for each transition with the BrICC code [31]. The conversion coefficient for the 181-keV transition is 210 0.241(4). As the remainder of observed transitions are 211 greater than 300 keV, conversion coefficients are expected to be negligible.

Relative  $\gamma$ -ray intensities have been determined from the observed number of counts in the  $\beta - \gamma$  singles spectrum, corrected for  $\gamma$ - and  $\beta$ -detection efficiency, such that the  $2_1^+ \rightarrow 0_1^+$  is normalized as 100. The absolute  $\gamma$ -ray normalization was calculated accounting for the known  $\beta$ -delayed neutron emission of <sup>146</sup>Cs (14.2 % [30]), and assuming no delayed neutron emission from <sup>146</sup>Ba. This was achieved from a comparison of the 181-<sub>221</sub> keV <sup>146</sup>Ba  $\gamma$  ray to the 141-keV  $2_3^- \rightarrow 2_1^-$  transition in 222 the daughter, <sup>146</sup>La [30]. In this procedure, it is assumed 223 that the contribution of <sup>146</sup>Ba in the beam was negligible 224 as these ions would be extracted from the gas catcher in 225 a 2<sup>+</sup> charge state, wheras a 1<sup>+</sup> <sup>146</sup>Cs beam was selected 226 through the separator. Any  $\gamma$  decay in <sup>146</sup>La results from  $_{227}$  a  $\beta$  decay of  $^{146}$ Ba in the  $0^+$  ground state. A low-energy, 228 high-spin (6<sup>-</sup>) isomer is reported in <sup>146</sup>La [32]. It is as-229 sumed that this isomer is not populated and the 141-230 keV transition has an absolute intensity  $I_{\gamma}=20.2(20)~\%$ 231 [30]. The number of efficiency-corrected counts observed 232 in both peaks is given in Table I. Using the adopted value <sub>233</sub> of  $I_{\gamma}$  for the 141-keV <sup>146</sup>La  $\gamma$  ray and the  $\beta$ -delayed neu-234 tron branch, the 'total' number of parent <sup>146</sup>Cs decays 235 was determined. The ratio of the 181-keV  $\gamma$ -ray inten-236 sity to this parent population gives the "normalization" 237 for Table I as 0.42(5), that is there are 42 181-keV  $\gamma$  rays per  $100^{146}$ Cs decays.

The intensity balance also allows an estimate of the  $\beta$ -branch of <sup>146</sup>Cs to the ground state of <sup>146</sup>Ba. Even sition, (b) the 558-keV,  $1_1^- \to 2_1^+$  transition, and (c) the 307- 240  $\beta$ -branch of  $^{146}$ Cs to the ground state of  $^{146}$ Ba. Even keV,  $3_1^- \to 2_1^+$  transition. Gamma rays from  $^{146}$ Ba are labelled 241 after correcting for internal conversion, the total identiby their energies. Random coincidence events from <sup>146</sup>Ce and <sub>242</sub> fied decay to the ground state is less than the population <sup>243</sup> of <sup>146</sup>La, so we infer the ground-state feeding in <sup>146</sup>Ba  $_{244}$  to be < 27 %. This is only an estimate, as any extra 245 unobserved feeding to the ground state from high-lying 250 This approach is limited by the completeness of the level While the level scheme has been extended extensively 251 scheme; if low-intensity transitions from high-lying states from what was previously known, the highest level ob- 252 are missed, then this will distort the inferred feeding. An served lies at  $\sim 2.2$  MeV, i.e.,  $\sim 3$  MeV below the neu- 253 indication of the level of "missing"  $\gamma$ -ray strength can be tron separation energy. It is possible that direct  $\beta$  feeding 254 seen through the  $\sim 2$  % population of low-lying 4+and 3to weak states within this energy range occurs which is 255 states which are forbidden decays, so should receive very not measurable with discrete-line spectroscopy. Resolv- 256 little direct  $\beta$ -population. Thus, the  $\beta$ -feeding intensities 258 tion is that the feeding pattern is very widely distributed 259 and no individual high-lying state is strongly populated. 260 Clearly, there is little overlap between the wave function <sup>261</sup> of the <sup>146</sup>Cs ground state and any of the excited levels in <sub>262</sub> <sup>146</sup>Ba.

<sub>207</sub> assignments have been used in calculating the conversion <sub>265</sub> were identified from  $\gamma - \gamma$  coincidence data, the corre-

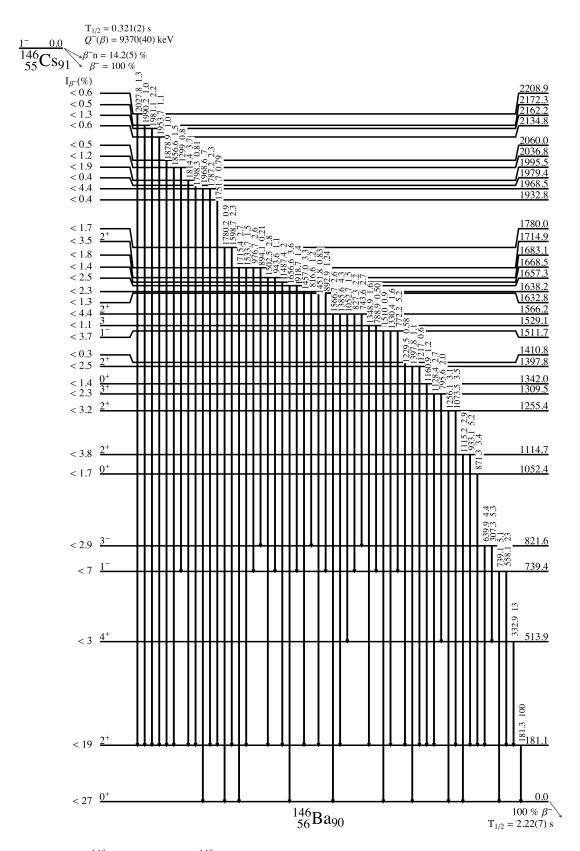


FIG. 5. Decay scheme of  $^{146}$ Ba populated in  $^{146}$ Cs  $\beta^-$  decay. In total, 31 excited states with 54  $\gamma$ -ray transitions have been identified. Labels indicate the energy and relative intensity of each transition. For absolute intensity per 100 decays, multiply by 0.42(5). I $_{\beta^-}$  values were determined by an intensity balance between the  $\gamma$  rays feeding and de-exciting each level, as discussed in the text.

266 sponding transition to the ground state was not observed 319 267 in the singles data. An upper limit on the relative in- 320 268 tensity of such transitions has been determined using the 321 269 intensity of the weakest  $\gamma$  ray that was observable. These 322 270 have not been included in intensity balances or normal- 323 271 ization.

## $J^{\pi}$ assignments

272

293

294

295

296

297

299

300

301

302

303

304

305

306

307

308

309

310

311

312

313

314

315

316

317

318

Spin values can be constrained for many <sup>146</sup>Ba excited 274 states observed from detailed inspection of  $\gamma$ -ray tran-275 sitions to levels with firm spin-parity assignments and  $\beta$ -decay selection rules. The data were not sufficient to 277 confirm these assignments through  $\gamma - \gamma$  directional correlation measurements. The  $J^{\pi} = 1^{-}$  spin and parity 279 of the parent is well known, having been measured via high-resolution laser spectroscopy [33].

It is expected that the observed levels in <sup>146</sup>Ba are 282 mostly populated via allowed  $(1^- \rightarrow 0^-, 1^-, 2^-)$  or firstforbidden decays  $(1^- \rightarrow 0^+, 1^+, 2^+, 3^+)$ . Observation (or non-observation) of  $\gamma$  transitions to the 0<sup>+</sup> ground state can be used to further constrain the spin assignment. Upper limits for relative intensities of unobserved  $\gamma$  transitions have been discussed above. The yrast levels lying below 1-MeV excitation have been reported in angular-correlation measurements [27]. The non-yrast 290 states above 1 MeV typically decay via low-multiplicity <sup>291</sup> cascades through the  $2_1^+$  level. States of J=1 or J=2292 are also seen to decay directly to the ground state.

# • The 1115- and 1255-keV levels

Excited states at 1115 and 1255 keV both  $\gamma$  decay to the  $2_1^+$  and  $0_1^+$  levels, therefore J=1,2 assignments are possible. No transitions to the negative-1101 keV. We assign the 1115-keV and 1255-keV

## • The 1310-keV level

The 1310-keV state feeds the  $4_1^+$  and  $2_1^+$  levels with no observed direct feeding to the  $0_1^+$  level. Given that the  $\gamma$  transitions only involve positive-parity states, we suggest this is the  $3_1^+$  level.

## • The 1342-keV level

Since the only  $\gamma$  transition from the 1342-keV level is to the  $2_1^+$  state, we assign this to be the  $0_2^+$  level. The  $\beta$ -feeding is large enough that, if this were not a  $0^+$  state,  $\gamma$  transitions to other levels would be expected to have intensities above the upper limit for non-observation.

## • The 1398-keV level

ate.

## • The 1512-keV level

We propose a 1<sup>-</sup> assignment to this state since it exhibits strong feeding to other low-spin (J =(0,1,2) states of both positive and negative parity, with an enhanced branch to the  $1_1^-$  state.

## • The 1529-keV level

324

325

326

327

This level decays to the  $1_1^-$  and  $2_1^+$  levels, with no observed direct feeding to the ground state. A  $2_1^$ assignment is allowed, however the strong branch to the  $2_1^+$  level favors a spin assignment of J=3.

## • The 1566- and 1715-keV levels

Strong  $\beta$  feeding and subsequent  $\gamma$  decays to all low-lying vrast states imply a uniquely constrained  $J^{\pi} = 2^{+}$  spin-parity for these levels.

For the remaining states, it has not been possible to 335 draw any solid conclusion pertaining to their appropriate 336 spin-parity assignments. In a few cases, a higher spin assignment is favored since no decay to the ground state was observed. However, the  $\beta$  feeding is weak and so it 339 was not possible to ascertain whether the  $\gamma$  transition 340 does not exist, or lies below the observation limit of the

#### DISCUSSION

With strong octupole correlations prevalent in this re-344 gion, double-octupole vibrations may be observable. The 345 excitation signature of this collective mode would be a 346 two-phonon multiplet  $(0^+, 2^+, 4^+, 6^+)$  located at ap-347 proximately twice the excitation energy of the  $3_1^-$  state. The  $0^+$  and  $6^+$  members decay via two E3 transitions to parity states were observed, suggesting that these 349 the 3- level and then the ground state, and the 2+ and are positive-parity states. The  $\overline{\text{IBA-1}}$  calculations  $^{350}$   $4^+$  members decay via enhanced E1 transitions. While (discussed below) also predict that the  $2^+_2$  lies at  $^{351}$  the  $4^+$  and  $6^+$  members will not be populated in  $\beta$  decay,  $_{352}$  one might expect to find the  $0^+$  and  $2^+$  members at  $\sim$ levels to be the  $2_2^+$  level and  $2_3^+$  level, respectively. 353 (2×822) keV. The 1638-keV level feeds the  $3_1^-$  and  $2_1^+$  $_{354}$  states, and does not  $\gamma$  decay to the ground state; there-355 fore, a case can be made that this state corresponds to  $_{356}$  the  $0^+$  member of the two-octupole phonon multiplet. 357 Similarly, the 1715-keV level also decays through the 3<sub>1</sub> 358 state and may possibly be associated with the 2<sup>+</sup> mem-359 ber of this multiplet. However, additional data are re-360 quired to draw firm conclusions about the observation of double-octupole vibrations in <sup>146</sup>Ba.

Key spectroscopic observables which differentiate between models describing nuclear shape changes are the excitation energies, spins and parities of low-lying nonyrast states, particularly the lowest few  $J^{\pi} = 0^{+}$  and  $J^{\pi} = 2^{+}$  levels, their electromagnetic decay properties, 367 and evidence for collective bands built upon them. Un-368 derstanding the development of collective behavior at the This excited state exhibits  $\gamma$ -decay characteristics  $_{369}$  beginning of the rare-earth region has evolved with our similar to those of the 1115- and 1255-keV levels, 370 capacity to constrain these observables. The focus of therefore a 2<sup>+</sup> spin-parity assignment is appropri- 371 this work is on determining spins and parities of these 372 important levels in <sup>146</sup>Ba. In this respect, the project

373 was only partially successful. The present data set has 374 revealed many new, higher-lying states which do not in-375 form this particular aspect. The data were insufficient <sub>376</sub> for  $\gamma - \gamma$  directional correlation measurements. How-377 ever, the enhanced sensitivity does offer the opportunity 378 for observation of some new low-intensity decays between 379 key low-lying states which constrain their possible spins, 380 sometimes uniquely. There is strong evidence that the 381 181-, 1115-, and 1255-keV levels are the  $J^{\pi} = 2_1^+, 2_2^+,$  $_{382}$  and  $2_3^+$  states. We use these assignments in the following 383 discussion. The remaining uncertainty with these assign-384 ments is the observation of several other low-lying states 385 which are interspersed between these levels and for which a firm spin assignment could not be made. As such, the possibility that these are additional  $J^{\pi}=2^{+}$  levels can-388 not be ruled out.

The decay scheme of <sup>146</sup>Ba was investigated within 390 the framework of the Interacting Boson Approximation (IBA) [34] by Scott et al., [27] and, more recently, by 392 Gupta and Saxena [35]. Both of these studies used a 393  $\chi$ -parameter of  $\chi = -\sqrt{7}/2$ , which corresponds to an 394 axially symmetric potential in the  $\gamma$  degree of freedom 395 centered at  $\gamma = 0^{\circ}$ . A general study of the N = 90 transition region in the IBA [36] indicates that this is unlikely to be the case.

Truncated level schemes of the lowest members of the  $_{399}$  ground-state,  $\beta\text{-}$  and  $\gamma\text{-}\mathrm{vibrational}$  bands for  $^{150}\mathrm{Nd}$  [6], <sup>148</sup>Ce [37], and <sup>146</sup>Ba (this work) are presented in Fig. 6. 401 The key signature of non-axial behavior in <sup>146</sup>Ba lies in 402 the location of the  $2_2^+$  state at 1115 keV with respect to 403 the  $0_2^+$  and  $2_3^+$  levels. Indeed, in line with the systemat-404 ics emerging from Fig. 6 and Fig. 7, we interpret the  $2^+_2$ 405 level as the bandhead of the  $\gamma$ -vibrational sequence and 435 where 406 associate the  $0_2^+$  and  $2_3^+$  with the (quasi- $\beta$ ) band. Hence, 407 the  $\gamma$  and  $\beta$  excitations lie remarkably close in energy, to 408 the extent that the " $2^+_{\beta}$ " and " $2^+_{\gamma}$ " locations are reversed 409 with respect to the heavier isotones. However, these aswould suggest that their unperturbed positions are nearly 441 fit for  $^{146}$ Ba corresponding to a  $\gamma$ -soft shape. 416 degenerate. The relative lowering of the excitation ener-417 gies of the  $J^{\pi}=2^{+}_{\gamma}$  levels in  $^{146}\mathrm{Ba}$  and  $^{148}\mathrm{Ce},$  shown by 418 the ratio  $E(2_{\gamma}^+)/E(2_1^+) = 6.2$  and 6.3, respectively, can  $_{419}$  be compared with 8.2 in  $^{150}\mathrm{Nd}$  and 8.9 in  $^{152}\mathrm{Sm}.$  This is 420 an indication that the triaxial potential energy is soft for 421 146Ba and 148Ce. Such an observation tends to disfavor any interpretation in terms of the X(5) geometric model, 423 which is based on a stiff, axially-symmetric potential in the  $\gamma$  degree of freedom.

In an effort to better understand this evolution of 426 structure in barium, and indeed along N = 90, IBA cal-427 culations were performed. The simplest version of the 428 model was used, which makes no distinction between pro-429 ton and neutron bosons (IBA-1), and employed the Ex-430 tended Consistent-Q Formalism (ECQF) [38]. The entire

[Color online] Truncated level schemes showing the lowest-lying members of the ground-state,  $\beta-$  and  $\gamma$ -vibrational bands in (a) <sup>150</sup>Nd, (b) <sup>148</sup>Ce, and (c) <sup>146</sup>Ba. The individual band sequences are labelled for each N = 90isotope.

431 IBA space can be described with a two-parameter Hamil-432 tonian incorporating a term related to the  $\beta$  deformation,  $\zeta$ , and one associated with the degree of axial asymmetry, 434  $\chi$ . The IBA-1 Hamiltonian is given by [39, 40]:

$$H_{\rm IBA-1}(\zeta) = c \left[ (1 - \zeta)\hat{n}_d - \frac{\zeta}{4N_B} \hat{Q}^{\chi} \cdot \hat{Q}^{\chi} \right], \quad (1)$$

$$\hat{Q}^{\chi} = (s^{\dagger} \tilde{d} + d^{\dagger} s) + \chi (d^{\dagger} \tilde{d})^{(2)}, \tag{2}$$

410 signments, and the association of a projection of angular 436 and  $\hat{n}_d = d^{\dagger} \cdot \tilde{d}$ . The parameters for the fits are included 411 momentum on the axis of deformation, K, are only rig- 437 in Table II. A comparison between the experimental and 412 orously applicable for axially symmetric nuclei. In fact, 438 calculated low-lying level energies is given in Fig. 7. The 413 in any non-axially symmetric case, these states mix, es- 439 calculations are in excellent agreement with the data, 414 pecially in a case like this where the moments of inertia 440 agreeing usually at the 10% level or better, with the best

TABLE II. Parameters  $\zeta$  and  $\chi$  used for each N=90 isotope in the IBA fits of this work.

Isotope	ζ	χ
Ba	0.732	-0.78
Ce	0.653	-0.95
$\operatorname{Nd}$	0.632	-1.03
$\operatorname{Sm}$	0.597	-1.21
$\operatorname{Gd}$	0.595	-1.10
Dy	0.615	-0.85
Er	0.633	-0.61

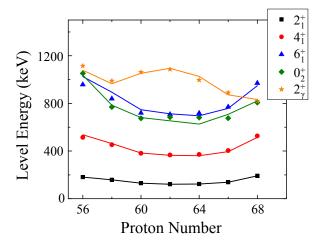


FIG. 7. [Color online] IBA fits (lines) from this work to experimental data (symbols) for the N = 90 isotones.

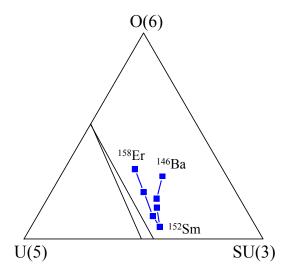


FIG. 8. [Color online] Trajectories within the IBA symmetry triangle for the N=90 isotonic chain, mapped according to the polar coordinate system of [36]. The slanting lines enclose the region of phase coexistence and phase transition.

453 seems to be quite symmetric.

443

455 quite rare and has been discussed as a possible signa- 507 No. 1064819. This research used resources of ANL's 456 ture for nuclei with properties lying along the so-called 508 ATLAS facility, which is a DOE Office of Science User 457 "Alhassid-Whelan Arc of Regularity [42]"; e.g., a small 509 Facility.

458 number of nuclei which have statistically regular spec-459 tra that are found in the mainly chaotic IBA parameter 460 space. An experimental signature of nuclei which may exhibit this regular behavior has been defined as those 462 having  $|E(2_2^+) - E(0_2^+)|/E(2_2^+) \le 0.025$  [43]. In <sup>146</sup>Ba, this quantity is small, 0.055, but just outside the prediction for identifying nuclei on the non-chaotic arc. How-465 ever, this simple experimental signature does not always 466 exactly follow the trajectory of "regular" nuclei which 467 are inferred from a full statistical analysis of the spectra [44]. Interestingly, both  $N = 90^{156}$  Dy and  $^{158}$ Er [45] have been previously identified as nuclei lying close 470 to the regular region [43]. The fact that <sup>146</sup>Ba exhibits 471 a similar degeneracy appears related to  $\gamma$  softness and 472 the symmetry of these shapes above and below axially 473 symmetric Z = 62, <sup>152</sup>Sm.

## CONCLUSIONS

A detailed  $\beta$ -decay spectroscopy measurement has been conducted on the neutron-rich exotic nucleus <sup>146</sup>Ba. This represents the first results from the recently-478 commissioned decay-spectroscopy station for low-energy CARIBU beams at Argonne National Laboratory. The experimental arrangement had a high sensitivity to 481 weak  $\gamma$ -ray transitions and, hence, enabled the study 482 of excited states not strongly populated via  $\beta$  decay. 483 Inspection of these low-intensity transitions has allowed spin constraints for low-lying levels, which have also been considered within the IBA framework. The N=90isotones are situated close to, but slightly to the right, of the phase-transitional region predicted by the IBA. They follow a symmetric behavior about  $^{152}\mathrm{Sm}\ (Z=62)$ which exhibits the highest degree of axial symmetry. Moving away from <sup>152</sup>Sm, isotones of both larger and smaller Z appear to exhibit increasing  $\gamma$  softness.

### ACKNOWLEDGEMENTS

The authors wish to acknowledge the Physics Sup-495 port group at Argonne National Laboratory and Figure 8 highlights the evolution within the so-called 496 engineering work of the Submillimeter-Wave Technol-'Casten triangle' [41] of the N = 90 isotones from <sup>146</sup>Ba <sub>497</sub> ogy Laboratory, University of Massachusetts Lowell. 446 to  $^{158}$ Er. Only with Z=62,64 (samarium and gadolin-498 Figure 5 in this article has been created using the 447 ium) are the N=90 isotones near the axial ( $\chi \sim -1.32$ ) 499 LevelScheme scientific figure preparation system [M. A. 448 route from U(5) to SU(3). Both heavier and lighter iso- 500 Caprio, Comput. Phys. Commun. 171, 107 (2005), 449 tones are best fitted with parameters deep in the interior 501 http://scidraw.nd.edu/levelscheme]. This material 450 of the triangle, that is, they follow the trend to deforma- 502 is based upon work supported by the U.S. Department tion along loci corresponding to non-axial shapes. Both 503 of Energy, Office of Science, Office of Nuclear Physics  $_{452}$  above and below Z=64, the trend of non-axial behavior  $_{504}$  under Grant Nos. DE-FG02-94ER40848 and DE-FG02-505 94ER40834, and Contract No. DE-AC02-06CH11357, The near degeneracy of the  $J^{\pi}=2^{+}_{2}$  and  $0^{+}_{2}$  levels is 506 and the U.S. National Science Foundation under Grant 566

567

571

574

575

577

579

602

603

605

607

609

611

613

614

 M. G. Mayer, Phys. Rev. 74, 235 (1948). 511

512

513

515

516

517

- [2] B. R. Mottelson, Rev. Mod. Phys. 29, 186 (1957).
- [3] C. W. Reich, Nucl. Data Sheets 113, 2537 (2012).
- [4] C. W. Reich, Nucl. Data Sheets 110, 2257 (2009). 514
  - [5] M. J. Martin, Nucl. Data Sheets 114, 1497 (2013).
  - [6] Κ. Basu and Α. Sonzogni, Α. Nucl. Data Sheets 114, 435 (2013).
- F. Iachello, Phys. Rev. Lett. 85, 3580 (2000). 518
- [8] F. Iachello, Phys. Rev. Lett. 87, 052502 (2001). 519
- [9] R. F. Casten, Nat. Phys. 2, 811 (2006). 520
- [10] R. F. Casten Zamfir, and 521 Phys. Rev. Lett. 87, 052503 (2001). 522
- [11] D. Tonev, A. Dewald, T. Klug, P. Petkov, J. Jolie, A. Fit-523 zler, O. Möller, S. Heinze, P. von Brentano, and R. F. 524 Casten, Phys. Rev. C 69, 034334 (2004). 525
- R. Krücken, B. Albanna, C. Bialik, R. F. Casten, 526 J. R. Cooper, A. Dewald, N. V. Zamfir, C. J. Bar-527 ton, C. W. Beausang, M. A. Caprio, A. A. Hecht, 583 528 T. Klug, J. R. Novak, N. Pietralla, and P. von Brentano, 584 529 Phys. Rev. Lett. 88, 232501 (2002). 530
- [13] M. A. Caprio, N. V. Zamfir, R. F. Casten, C. J. Barton, 531 C. W. Beausang, J. R. Cooper, A. A. Hecht, R. Krücken, 532 H. Newman, J. R. Novak, N. Pietralla, A. Wolf, and 588 533 K. E. Zyromski, Phys. Rev. C 66, 054310 (2002). 534
- 535 [14] W. D. Kulp, J. L. Wood, P. E. Garrett, C. Y. Wu, 590 D. Cline, J. M. Allmond, D. Bandyopadhyay, D. Dash- 591 536 dorj, S. N. Choudry, A. B. Hayes, H. Hua, M. G. Mynk, 592 537 M. T. McEllistrem, C. J. McKay, J. N. Orce, R. Teng, 593 538 and S. W. Yates, Phys. Rev. C 77, 061301 (2008). 539
- P. E. Garrett, W. D. Kulp, J. L. Wood, D. Bandyopad-540 hyay, S. Choudry, D. Dashdorj, S. R. Lesher, M. T. 541 McEllistrem, M. Mynk, J. N. Orce, and S. W. Yates, 597 542 Phys. Rev. Lett. 103, 062501 (2009). 543
- J. F. Sharpey-Schafer, S. M. Mullins, R. A. Bark, J. Kau, 599 544 545 F. Komati, E. A. Lawrie, J. J. Lawrie, T. E. Madiba, 600 P. Maine, A. Minkova, S. H. T. Murray, N. J. Ncapayi, 601 546 and P. A. Vymers, Eur. Phys. J. A 47, 5 (2011). 547
- F. Casten and N. Zamfir, |17|548 J. Phys. G Nucl. Partic. 22, 1521 (1996). 549
- T. Nikšić, D. Vretenar, G. A. Lalazissis, and P. Ring, 550 Phys. Rev. Lett. **99**, 092502 (2007). 551
- N. K. Nomura, Shimizu, Otsuka, and 552 Phys. Rev. Lett. 101, 142501 (2008) 553
- A. Butler and W. Nazarewicz, 554 Rev. Mod. Phys. 68, 349 (1996). 555
- 556 [21] S. M. Scott, D. D. Warner, W. D. Hamilton, P. Hungerand B. Pfeiffer, ford, G. Jung, K. D. Wünsch, 557 J. Phys. G: Nucl. Phys. 5, L187 (1979). 558
- J. B. Wilhelmy, S. G. Thompson, R. C. Jared, 559 E. Cheifetz, Phys. Rev. Lett. 25, 1122 (1970). 560
- [23] W. R. Phillips, I. Ahmad, H. Emling, R. Holzmann, 561 R. V. F. Janssens, T. L. Khoo, and M. W. Drigert, 617 [45] R. G. Helmer, Nucl. Data Sheets 101, 325 (2004). 562 Phys. Rev. Lett. 57, 3257 (1986). 563
- [24] W. Urban, M. A. Jones, J. L. Durell, M. Leddy, W. R. 564 Phillips, A. G. Smith, B. J. Varley, I. Ahmad, L. R.

- Morss, M. Bentaleb, E. Lubkiewicz, and N. Schulz, Nucl. Phys. A 613, 107 (1997).
- [25]D. Kusnezov F. Iachello, 568 and Phys. Lett. B **209**, 420 (1988). 569
- [26] H. Mach, W. Nazarewicz, D. Kusnezov, M. Moszyn'ski, 570 B. Fogelberg, M. Hellström, L. Spanier, R. L. Gill, R. F. Casten, and A. Wolf, Phys. Rev. C 41, R2469 (1990).
- S. M. Scott, D. D. Warner, W. D. Hamilton, P. Hunger-573 ford, G. Jung, K. D. Wünsch, and B. Pfeiffer. J. Phys. G: Nucl. Phys. 6, 1291 (1980).
- G. Savard, S. Baker, C. Davids, A. Levand, E. F. Moore, [28]576 R. C. Pardo, R. Vondrasek, B. J. Zabransky, and G. Zinkann, Nucl. Instrum. Meth. B 266, 4086 (2008). 578
  - A. J. Mitchell, P. F. Bertone, B. DiGiovine, C. J. Lister, M. P. Carpenter, P. Chowdhury, J. A. Clark, N. D'Olympia, A. Y. Deo, F. G. Kondev, E. A. Mc-Cutchan, J. Rohrer, G. Savard, D. Seweryniak, and S. Zhu, Nucl. Instrum. Meth. A 763, 232 (2014).
  - L. Peker and J. Tuli, Nucl. Data Sheets 82, 187 (1997).
  - T. Kibédi, T. W. Burrows, M. B. Trzhaskovskaya, P. M. Davidson, and C. W. Nestor Jr., Nucl. Instrum. Meth. A 589, 202 (2008).
  - C. Chung, W. B. Walters, D. S. Brenner, R. L. Gill, M. Shmid, Y. Y. Chu, R. E. Chrien, L.-J. Yuan, F. K. Wohn, and R. A. Meyer, Phys. Rev. C 31, 2199 (1985).
  - A. Coc, C. Thibault, F. Touchard, H. T. Duong, P. Juncar, S. Liberman, J. Pinard, M. Carre, J. Lerme, J. L. Vialle, S. Buttgenbach, A. C. Mueller, and A. Pesnelle, Nucl. Phys. A 468, 1 (1987).
- 595 [34] A. Arima and F. Iachello, Phys. Rev. Lett. 35, 1069 (1975). 596
  - [35]J. В. Gupta Μ. and Saxena. Phys. Rev. C 91, 054312 (2015).
  - [36] E. A. McCutchan, N. V. Zamfir, and R. F. Casten, Phys. Rev. C **69**, 064306 (2004).
  - N. Nica, Nucl. Data Sheets 117, 1 (2014).
  - P. O. Lipas, P. Toivonen, and D. D. Warner, Phys. Lett. B 155, 295 (1985).
- [39] N. V. Zamfir, P. von Brentano, R. F. Casten, and J. Jolie, 604 Phys. Rev. C 66, 021304(R) (2002).
- V. Werner, P. von Brentano, R. F. Casten, and J. Jolie, [40]606 Phys. Lett. B **527**, 55 (2002).
- 608 [41] R. F. Casten and D. D. Warner, Progress in Particle and Nuclear Physics 9, 311 (1983).
- 610 [42] Y. Alhassid and Whelan, Phys. Rev. Lett. 67, 816 (1991).
- J. Jolie, R. F. Casten, Cejnar, S. 612 Ρ. Heinze. Α. McCutchan, N. Zamfir, and Phys. Rev. Lett. **93**, 132501 (2004).
- D. Bonatsos, E. A. McCutchan, and R. F. Casten, 615 Phys. Rev. Lett. **104**, 022502 (2010).