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Phys. Rev. C **84**, 034908 — Published 20 September 2011

DOI: [10.1103/PhysRevC.84.034908](https://doi.org/10.1103/PhysRevC.84.034908)

# Fate of the Initial State Perturbations in Heavy Ion Collisions.

## II Glauber Fluctuations and Sounds

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(Dated: August 30, 2011)

Heavy ion collisions at RHIC are well described by the (nearly ideal) hydrodynamics for average events. In the present paper we study initial state fluctuations appearing on an event-by-event basis, and the propagation of perturbations induced by them. We found that (i) fluctuations of several lowest harmonics have comparable magnitudes, (ii) that at least all odd harmonics are correlated in phase, (iii) thus indicating the local nature of fluctuations. We argue that such local perturbation should be the source of the “Tiny Bang”, a pulse of sound propagating from it. We identify its two fundamental scales as (i) the “sound horizon” (analogous to the absolute ruler in cosmic microwave background and galaxy distribution) and (ii) the “viscous horizon”, separating damped and undamped harmonics. We then qualitatively describe how one can determine them from the data, and thus determine two fundamental parameters of the matter, the (average) *speed of sound* and *viscosity*. The rest of the paper explains how one can study mutual coherence of various harmonics. For that one should go beyond the two-particle correlations, to three (or more) particles. Mutual coherence is important for the picture of propagating sound wave.

PACS numbers:

### I. INTRODUCTION

Starting the introduction, let us note that the issues to be discussed in this paper are somewhat similar in nature to current trends in cosmology of the last decade, which made it a really quantitative science. The sound horizon scale has been seen on the sky, in correlations of cosmic microwave background temperatures and in locations of the galaxies.

Experimental data obtained in heavy ion collisions at the Relativistic Heavy Ion Collider (RHIC) have found the “Little Bang”, a hydrodynamical explosion driven mostly by pressure of the new form of matter, Quark-Gluon Plasma. Their experimental data for radial and elliptic flows has been compiled in the so called “white papers” of RHIC experiments [1–4] in 2004, and compared with predictions of relativistic hydrodynamics. The models which implemented freezeout via hadronic cascades [5–7], as originally suggested in [8], were found to be especially successful. Very recent results from the Large Hadron Collider (LHC) on elliptic flow [9] also turned out to be in agreement with hydrodynamical predictions, suggesting that QGP remains a good liquid even at LHC (see e.g. [10]). Dissipative effects from the QGP viscosity provide only small corrections at the few percent level, see [11–13]. So, by now, we have quantitative overall description of the “Little Bang”.

This paper is about small deviations from this average behavior. Such *perturbations* of the average explosion can come from at least two different sources. The one which we will study in this paper is due to quantum fluctuations in the wave function of the colliding nuclei, which creates “bumpy” distributions of matter, for any collision, which one can decompose into a smooth average plus local perturbations.

The smallness of the perturbation amplitude, with respect to the local density of ambient matter, would suggest the appearance of divergent sound waves, see Fig.1. Similar to the circles from a stone thrown into a pond, hydrodynamics tells us that initial perturbations should become moving waves, with basically nothing left at the original location at later time. This is the picture we are going to work on in this paper.

Another one, to be studied in subsequent papers of this series, are created by the energy deposited by jets propagating through the medium. It has been recently dramatically shown by the ATLAS collaboration [14] that even jets with energy above 100 GeV deposit large part of their energy, and sometimes all of it, into the medium. The first ideas were to look at the resulting perturbations in the form of a Mach cone [15], driven by the view that the energy is deposited more or less homogeneously along the jet path. However more recent developments of the theory, based on AdS/CFT, have lead to the view that the deposited energy grows rapidly with the distance travelled by the jet, with significant deposition at the end point. Thus one may think of the *second kind of the “Tiny Bangs”*, this time occurring in between the beginning and the end of the “Little Bang”. Obviously those should lead to sound circles of smaller absolute and angular sizes.

An alternative idea, of randomly fluctuating shapes of the produced initial fireballs, has resulted in an approach in which different angular harmonics of that distribution are treated separately. The realization that even central  $\vec{b} = 0$  collisions may have some fluctuating ellipticity has lead to the discussion of the elliptic flow event-by-event fluctuations, see [16] and many subsequent works. The so called “triangular flow” related to the fluctuations in the 3-ed harmonic of the flow has been recently studied

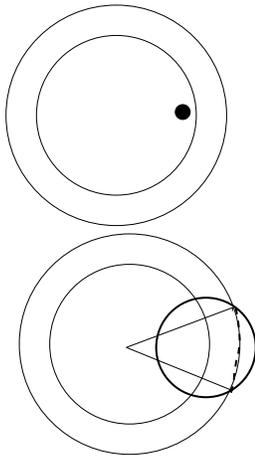


FIG. 1: A sketch of the transverse plane of the colliding system: the two concentric circles are the nuclear radius (inner) and the final radius of the fireball (outer). The black spot in the upper figure is an extra density due to initial density fluctuations. The small perturbation becomes a circle of a sound wave, which will be stopped at freezeout at thick circle (distorted by radial flow). The part outside of the fireball does not exist: the corresponding matter will actually be placed near the edge of the fireball (thick dashed line). The angle between the two thin lines corresponds to the angle between two peaks in the resulting angular distribution.

in an important paper by Alver and Roland [17], with many groups working in this direction now.

The main difference between our approach and that is that we treat such fluctuations not as independent (Gaussian) noise in different harmonics but as certain local perturbations, resulting in certain evolving sound fronts, reaching certain size, shape and diffusivity by the moment of freezeout. In other words, we think that various harmonics are mutually coherent. One of the main goals of this paper is to explain how this coherence can be tested experimentally. We will also provide evidences (based on Glauber model of the perturbations) that such mutual coherence is present in this model.

(On one hand, one may argue that as soon as all the perturbations are small and the equations are linear, it is not important if one expands in harmonics before or after the solution of hydro equations. So the solution to hydro we will discuss in our next companion paper [18] can be used for both “noise” and coherent scenarios.)

Let us also briefly recall the history of the theory efforts. The propagation of sound on top of the fireball has been discussed by J.Casalderrey-Solana and one of us in [19]. In that paper the fireball expansion was modelled by the Big-Bang-like overall expansion of the space, with the same Friedman-Robertson-Walker metric as used for cosmology. The focus of that paper was the effect of time-dependent sound velocity, especially if the phase transition is 1st order and it can vanish at some interval of  $T$ . The interesting finding was the creation of the secondary

– and convergent – sound waves. This idea was further discussed in [20] in connection with the “soft ridge” issue, but with the conclusion that if the current lattice data on the speed of sound is correct, the effect of the reflected wave is too small.

In the same paper [20] it has been found that the usual (unreflected) sound propagation should produce characteristic “two-peak events”, with the angle between the peaks reflecting the sound horizon and numerically being about 1 radian. More specifically, this angle corresponds to the angle at which two intersections of the fireball boundary are seen from the fireball center. The Brazilian group [21] has independently found such peaks with the same angle, induced by a local perturbation. Andrade et al have further pointed out that the *two*-peak events lead to a *three*-peak correlation function, with the side peaks separated by twice larger angle  $\sim 2$  rad. This observation explained what has been found earlier, in the average over many events, in “event-by-event” hydrodynamical studies by the Brazilian group (see [22] and references therein) and Werner et al [23]. Good agreement between the shapes of the correlation function in “event-by-event” pictures and single local perturbation indicate that the former are more or less linear superpositions of the latter ones.

Completing this introduction, let us note that this paper was completed and submitted to archive in Aug.2010, but it was submitted to PRC in June 2011 together with its companion paper [18]. During this time there were obviously many papers devoted to the subject. In particular, Glauber model predictions for several harmonics  $\epsilon_n$  and their root-mean-square fluctuations have been calculated independently several times. Qin et al. [24] paper is an example, which also considered the influence of inelasticity fluctuations of the p+p reaction on the initial density fluctuations. Significant experimental progress resulted now in measurements of higher harmonics till  $n = 9$ , by all five collaborations at RHIC/LHC. Yet the main suggestion of this work – the measurements of the mutual phases of those harmonics – still remains to be done in the future.

## II. SETTING UP THE PROBLEM

### A. The main scales of the problem

Before going to specifics, let us formulate the problem in a more general form. Two generic scales of the hydro approach are (i) the macroscopic scale  $R$  and the microscopic scale  $l$ , being in the relation

$$l \ll R \quad (2.1)$$

which ensures such macroscopic tools as thermo and hydrodynamics should work.

Now let us define two new scales. The first is the *sound*

horizon

$$H_s = \int_{\tau_i}^{\tau_f} d\tau c_s(\tau) \quad (2.2)$$

where the integral is taken from the formation of the hydro system to its freezeout time. While in the Big Bang  $\tau_f \sim 100000 \text{ years}$ , in the Little Bang it is only  $\sim 10 \text{ fm}$ . It is important, that in both cases at the freezeout the collisions cannot support pressure any more, yet the sound wave does not disappear: it just stops, and thus can be observed. For the Big Bang this idea was suggested by Sunyaev and Zeldovich 30 years ago [25]: who called the sound horizon to be the “standard ruler” of the Universe. Its current value,  $H_s \approx 150 \text{ Mps}$ , as seen in galaxy’s distribution and in CMB correlations, is excellent way to measure the Hubble constant and lifetime of the Universe.

In the cosmology such scale corresponds to angles of about 1 degree and thus  $l \sim 200$ . In the Little bang we deal with about 1 radian and  $m \sim 3$ .

The second scale (not important in cosmology), which we would like to call “the viscous horizon scale”  $R_v$ , separates the wavelengths of the sound which *are and are not* dissipated by the viscosity effects. The smooth fireball and fluctuations are described by

$$T_{\mu\nu} = \tilde{T}_{\mu\nu} + \delta T_{\mu\nu} \quad (2.3)$$

The textbook dispersion law for the sound, including the viscosity term, is

$$\omega = c_s k - \frac{i}{2} \frac{4\eta}{3s} \frac{k^2}{T} \quad (2.4)$$

The Fourier transform puts it into momentum form, after which one can solve the time dependence using the momentum-dependent dispersion relation as well as the imaginary part induced by viscosity.

(One may add bulk viscosity to this expression as well, but we keep the shear viscosity for now, assuming it is dominant.)

$$\delta T_{\mu\nu}(t) = e^{-\frac{2}{3} \frac{\eta}{s} \frac{k^2 t}{T} + ik(x-tc_s)} \delta T_{\mu\nu}(0) \quad (2.5)$$

The first term in the exponent defines the new *viscous survival scale*: harmonics with  $k \ll k_v$  would be affected little, and with  $k \gg k_v$  be killed. The definition of this scale is clear from the exponent

$$k_v = \frac{1}{R_v} = \sqrt{\frac{3Ts}{2\tau_f \eta}} \quad (2.6)$$

While the sound horizon determines the size of the sound circles at freezeout, the viscous one determines their width. Note that while the former increases linearly with time  $H_s \sim t$ , this width  $R_v \sim \sqrt{t}$ . So, although the spheres become more diffuse, they are also relatively sharper as time goes by.

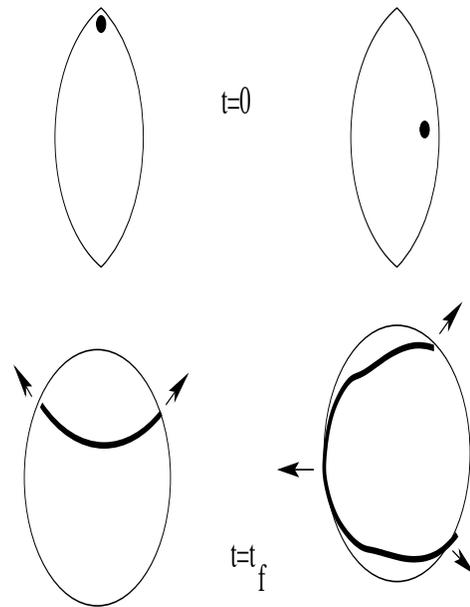


FIG. 2: The two upper pictures correspond to initial time  $t = 0$ : the system has almost shape and contains perturbations (black spots). The two lower pictures show schematically location and diffuseness of the sound fronts at the freezeout time  $t_f$ . The arrows indicate the angular direction of the maxima in the angular distributions, 2 and 3 respectively.

(One may also ask a very good question how early hydrodynamical description of the perturbations may be valid. We will not go into its discussion here, and only state that the applicability limits of Navier-Stokes hydrodynamics is *not* determined by the  $k_v$  scale, but by the magnitude of the higher order terms in gradients, *resummed*. The last word is added because of the issue of alternating sign of the series, raised in “universal resummed hydrodynamics” paper [26]. For the last works on the subject in the ads/cft context see e.g. [27], in which the convergence of large class of non-equilibrium evolutions to such universal behavior is indeed demonstrated and its applicability limits discussed quantitatively.)

How can one measure the two scales,  $H_s$  and  $1/k_v$ , defining the *speed of sound* and *viscosity*, experimentally? One may change the geometry of the collision (by centrality) and the size of the nucleus (by changing the beam A), but an even better way is by observing many angular harmonics of the flow. Note that the amplitudes of the higher harmonics are damped by viscosity more, so, if measured, they would provide a measure of the viscosity.

For central collisions at RHIC/LHC the hierarchy relation between all those four scales is

$$R \sim H_s > R_v > l \quad (2.7)$$

As some representative numbers let us mention 8, 6, 1, 1 fm (at freezeout), respectively. However for mid-central collisions the width (the shorter size) of the “al-

mond”  $R_x$  becomes comparable or smaller than  $H_s$ . As a result, one expects the sound wave to traverse the whole fireball and deposit some amount of (entropy) density to its opposite side. In this case one expects 3-peak events: see Fig.2 for explanation. So, one of our suggestions to experimentalists is to study the centrality at which certain changes in harmonics amplitudes and phases may occur.

### III. THE INITIAL STATE FLUCTUATIONS

#### A. Generalities

Let us start with a comment on what we would call the “initial state”. This term is currently used in at least three different settings:

- (i) The wave function of the colliding nuclei, expressed either in terms of the nucleons (their positions in the transverse plane just prior to the collisions) or in terms of partonic degrees of freedom (positions and longitudinal momenta). Another version of it is the “Color Glass Condensate” (CGC) described as an ensemble of classical gauge fields.
- (ii) The state just after the (Lorentz contracted) nuclei passed each other. It is either the partonic state, including partons newly produced in a collision; or the so called GLASMA, in the classical field description.
- (iii) The state after approximate equilibration is reached, so that macroscopic (hydrodynamical) description can be started.

It is the last one which we mean in this work, as we would apply hydrodynamics as a tool, translating properties of the initial conditions into the final state observed in the experiment. Therefore our “initial state” should correspond to about one unit of the relaxation time after the actual collision, or numerically at a proper time of the order of  $1/2$  fm/c. Thus the inhomogeneity of the initial wave functions should be already smoother than at time zero, by this (so far poorly understood) relaxation process.

As we detail below, this state will be described by some “average” or zeroth-order shape of the fireball (depending of course on the impact parameter, the colliding nucleus and the collision energy), plus “fluctuations” characterized in the first order by an ensemble of small perturbations of the average shape described by Fourier coefficients and phases  $\{\epsilon_n, \psi_n\}$ . Generic expressions would include the zeroth order ensemble-average deformations  $\langle \epsilon_n \rangle$  and deviations which have no average but fluctuations  $\delta\epsilon_n^2 = \langle \epsilon_n^2 \rangle - \langle \epsilon_n \rangle^2$ .

The simplest situation, happening for the second harmonics and sufficiently peripheral collisions, is that the average is much larger than the fluctuations,  $\langle \epsilon_2 \rangle \gg \delta\epsilon_2$ . If so, one may assume Gaussian form of the fluctuations with the width given by  $\delta\epsilon_2$ . But in general the situation is quite different, the odd harmonics are always due to fluctuations and also, for near central collisions,

for all  $n$  both terms in  $\epsilon_n$  come from fluctuations. Till  $n < 10$  they are comparable in magnitude  $\epsilon_n \sim 1/10$  with wide fluctuations. (The distributions are obviously non-Gaussian because they are all positive by definition.) All of those should in principle be provided by some “initial state models”, of which we select Glauber model as the simplest example.

The separation of the initial state fluctuations from all other fluctuations (e.g. fluctuations during the hydrodynamical evolution, hadronization and the freezeout) is possible because of the fundamentally different number of relevant degrees of freedom defining their magnitude. As we will detail in the next section, the so called Glauber fluctuations due to various number of “wounded” (or participant or interacting) nucleons are of the order of

$$\epsilon_n \sim \frac{1}{\sqrt{N_p}} \quad (3.1)$$

where the number of the participant nucleons  $N_p \sim O(100)$ , being limited from above by the total nucleon number  $2A \sim 400$ .

Further fluctuations are determined similarly, but with the number of participants  $N_p$  substituted by the *much larger* number of partons involved, or the total multiplicity  $N_{hadrons} \sim 10^4$  (for RHIC and LHC it is factor 2 different). That is why one may, to certain accuracy, ignore all later-time fluctuations and assume that observable fluctuations in particle spectra and correlation functions are one-to-one translated from the initial state ensemble. Thus we use hydrodynamical equations as a fully deterministic tool, by itself producing no random numbers at all.

Furthermore, for near-central collisions all  $\delta\epsilon_n$  are small, of the order of several percents. So, independently of their possibly complicated distributions and cross correlations, the hydrodynamics applied in linear approximation should be quite reliable tool. Thus hydro equations can be linearized and the linear response coefficients  $\delta v_n / \delta\epsilon_n$  calculated. If so, it does not matter what the actual magnitude of the deformation  $\delta\epsilon_n$  is. Also the linearized perturbations do not interact with each other.

Although we will focus on those calculations in our next paper, let us note here two things. One simple fact is that while angles  $\psi_n$  of the fireball deformations indicate the *maxima* of the distribution (the corners of triangle, square and other polygons), hydro flow goes along their *sides*. Therefore the observed flow angles  $\xi_n$  are rotated from the deformation angle as follows

$$\xi_n = \psi_n + \frac{\pi}{n} \quad (3.2)$$

Our second comment is that higher harmonics  $n$  oscillate with time, displaying acoustic sound properties. At freezeout this leads to their certain signs, and these phases should be added to this relation as well.

## B. Fluctuations in the Glauber model: the amplitudes

Our ‘‘Glauber model’’ is a bit different from that used widely by experimentalists. Both assume that initial state fluctuations originate from the nuclear wave functions. The ‘‘usual Glauber’’ uses randomly placed coordinates of the individual nucleons in the nuclear wave function. However, the nucleons themselves are complicated objects and their interactions are also strongly fluctuating: since there are studies of that we decided to include this source of fluctuations also. This changes numbers a bit, but was found not to be important for any of the qualitative conclusions to be reached.

The nucleon fluctuations we included via the *fluctuating NN cross sections* are to a certain degree known and studied via diffraction, see [28] for the details and earlier references. Naively, from the well known fact of a nucleon being made of quite a large number of partons one might conclude that those fluctuations are small,  $O(1/N_{partons})$ : but this is *not* the case. In our simulation we have assumed the cross section  $\sigma_{NN}$  to be the random Gaussian variable with the variance

$$w_{NN} = \frac{\langle \sigma_{NN}^2 \rangle - \langle \sigma_{NN} \rangle^2}{\langle \sigma_{NN} \rangle^2} \approx 0.25 \quad (3.3)$$

First, like in [17], we simulate a large ensemble of collisions and calculate the magnitude of the  $\epsilon_n$  for several lowest harmonics (up to 6). Their definition is via the Fourier expansion for a single particle distribution

$$f(\phi) = \frac{1}{2\pi} \left( 1 + 2 \sum_n \epsilon_n \cos(n(\phi - \psi_n)) \right) \quad (3.4)$$

where the  $\epsilon_n$  are the participant anisotropies and the  $\psi_n$  are the angles between the x axis and the mayor axis of the participant distribution.

The participant anisotropies are calculated from

$$\epsilon_n = \frac{\sqrt{\langle r^n \cos(n\phi) \rangle^2 + \langle r^n \sin(n\phi) \rangle^2}}{\langle r^n \rangle} \quad (3.5)$$

This expression is calculated in the center of mass of the participant nucleons for each event. Therefore the dipole moment  $n = 1$  made out of the average coordinates

$$\langle x \rangle = \langle r \cos(\phi) \rangle = 0, \quad \langle y \rangle = \langle r \sin(\phi) \rangle = 0 \quad (3.6)$$

are zero by definition.

The 2-d shape of the event can in principle be expanded in the double Taylor series in  $x, y$  or in double series over moments  $r^m \cos(n\phi), r^m \sin(n\phi)$  with integer  $m, n$ . An even better definition would be to follow the customary statistical trick and write the distribution as the *exponent* containing the ‘‘generating function’’ of the angular dependence expended in harmonics

$$P = F_1(r) \exp(F_2), \quad F_2 = \sum_{n>0} r^n \epsilon_n \cos(n(\phi - \psi_n)) \quad (3.7)$$

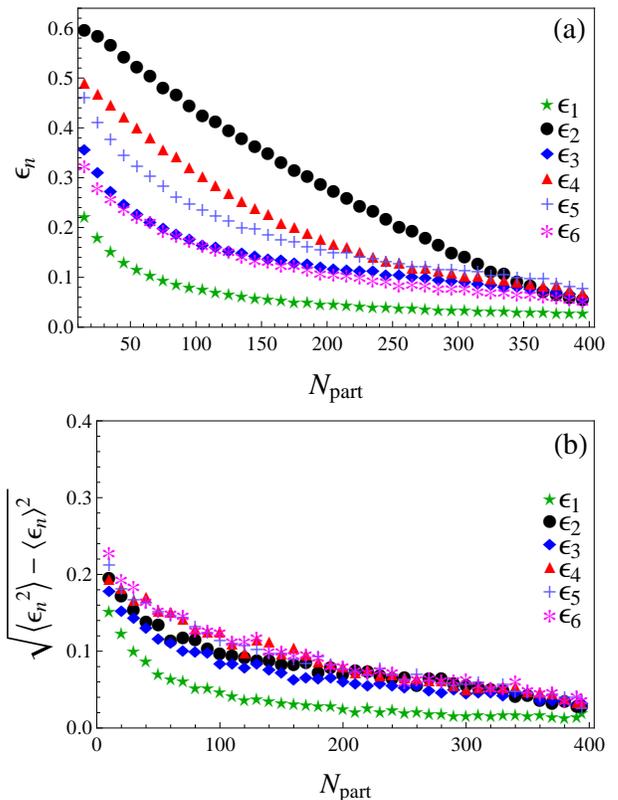


FIG. 3: Average anisotropies (a) and their variations (b), as a function of centrality expressed via the number of participants  $N_{part}$

In this way the positivity of the distribution function, as well as inclusion of trivial higher order effects are ensured.

Since the dipole  $m = n = 1$  term is zero by construction, we define the first odd deformation  $\epsilon_1, \psi_1$  using the term of the expansion  $m = 3, n = 1$  which appears together with the triangular deformation  $m = n = 3$

$$\epsilon_1 = \frac{\sqrt{\langle r^3 \cos(\phi) \rangle^2 + \langle r^3 \sin(\phi) \rangle^2}}{\langle r^3 \rangle} \quad (3.8)$$

The anisotropies calculated in this way are plotted in figure 3 for  $n = 1, 6$ . The plot shows that the eccentricity has the largest value for the well known elliptic deformation  $\epsilon_2$  and a nonzero value of triangularity  $\epsilon_3$ , in agreement with the results reported in [17]. Note that for the near-central collisions  $N_{part} > 300$  the elliptic deformation is no longer dominant, and it is also due to fluctuations. This conclusion becomes evident as one looks at the lower plot in Fig.3, which shows the variations of these  $\epsilon_n$ .

One observation coming from these results is that all other deformations (except for  $\epsilon_1$ , small because the ‘‘true dipole’’ remains zero) are all comparable, ranging from  $O(1/10)$  for central collisions to 0.3-0.5 for most peripheral ones. While in the central bins these perturbations can be considered small and treated as Gaussian random variables, it is clear that for most peripheral bins (when

the number of participants is smaller) the fluctuations are large and thus must be non-Gaussian.

Another consequence is that there is absolutely no ground to single out  $\epsilon_3$ : in fact both  $\epsilon_4$  and  $\epsilon_5$  are larger than  $\epsilon_3$  and  $\epsilon_6$  is about of the same order as  $\epsilon_3$ .

The last point is that their variations (the lower plot) are all comparable to the magnitude. Yet the definition of deformations are such that they are always positive, for each event. This is one more reason why the amplitudes cannot have Gaussian distribution, deviating from it at least for the smallest values.

### C. Fluctuations in the Glauber model: the angles

The angles  $\psi_n$  are defined by:

$$\tan(n\psi_n) = \frac{\langle r^n \sin(n\phi) \rangle}{\langle r^n \cos(n\phi) \rangle} \quad (3.9)$$

and to calculate  $\psi_1$  we use:

$$\tan(\psi_1) = \frac{\langle r^3 \sin(\phi) \rangle}{\langle r^3 \cos(\phi) \rangle} \quad (3.10)$$

Using these expressions we obtain the distribution of the  $\psi_n$ 's for the first six harmonics as shown in Figs 4, 5. In order to better understand the behavior of these angles we will now study their correlation.

(Note that our angle definition is different from the one by Alver-Roland [17]: we do not include extra phase  $\pi/n$  between the flow and deformation directions, see below.)

Let us comment on these distributions, starting from the even ones.

The most obvious one is a distribution of the second (elliptic) harmonic: as seen in Fig.4 the angle  $\psi_2$  is strongly peaked at  $\pi/2$ , corresponding to an elongation of the system in the  $y$  direction, as of course one expects from the overlap ‘‘almond’’ of two nuclei. The distribution of the 4-th angle  $\psi_4$  in Fig.5 shows peaks at angles 0 and  $\pi/2$ : but since quartic symmetry of the 4-th harmonics it simply means that the maxima of the distribution tend to be aligned with the coordinate axes  $x$  and  $y$ . The distribution of the 6-th harmonics is different: it is peaked at the angle  $\pi/6$ . This means that it has no maximum at  $x$  direction but rather in  $y$ . In conclusion, all even harmonics are strongly correlated with the reaction plane, all of them producing maxima along the  $y$  (out-of-plane) direction.

The distribution of the angle  $\psi_1$  is nonzero at all angles, which means it is not very strongly correlated with the reaction plane. It has two maxima, at  $\pm y$  directions, to be called ‘‘tip’’ fluctuations. Although the contribution from angles 0,  $\pi$  or  $x$ -directions is about twice smaller, it also makes an important contribution: we will call it ‘‘waist’’ fluctuations. Note that while the area of the ‘‘waist’’ is larger than ‘‘tips’’: and yet its contribution is smaller.

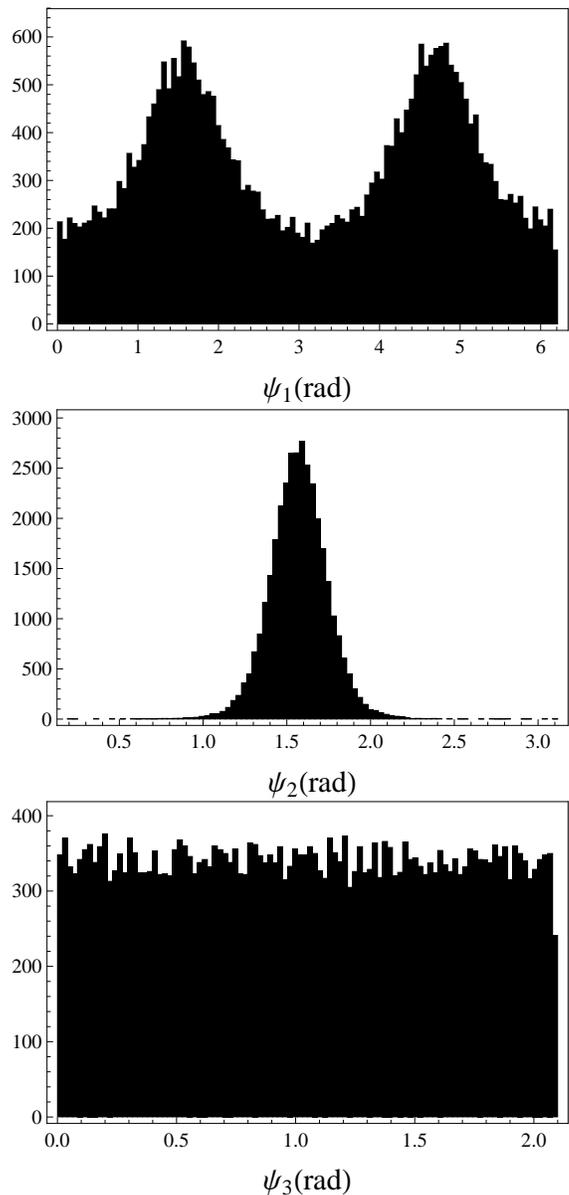


FIG. 4: Distribution of the angles  $\psi_n$  for the first three harmonics, the centrality bin used is  $100 < N_{part} < 300$ .

The distribution over  $\psi_3, \psi_5$  in these figures looks completely uncorrelated with the reaction plane. (This fact has also been noticed in [17] and by others.) However, further scrutiny shows that they are in fact well correlated with  $\psi_1$ , see Fig. 6 (in which we included points repeated by periodicity). The distribution can be crudely characterized by some ‘‘bumps’’ plus ‘‘stripes’’ connecting them.

The interpretation of the ‘‘bumps’’ is that all of them correspond to events with additional ‘‘hot spot’’ at the ‘‘tips’’ of the almond. It is a very natural place for maximal fluctuations, for two reasons. First, this is where the participant density in both nuclei is near zero. Second, because of the factor  $r^3$  they have larger weight than all other places.

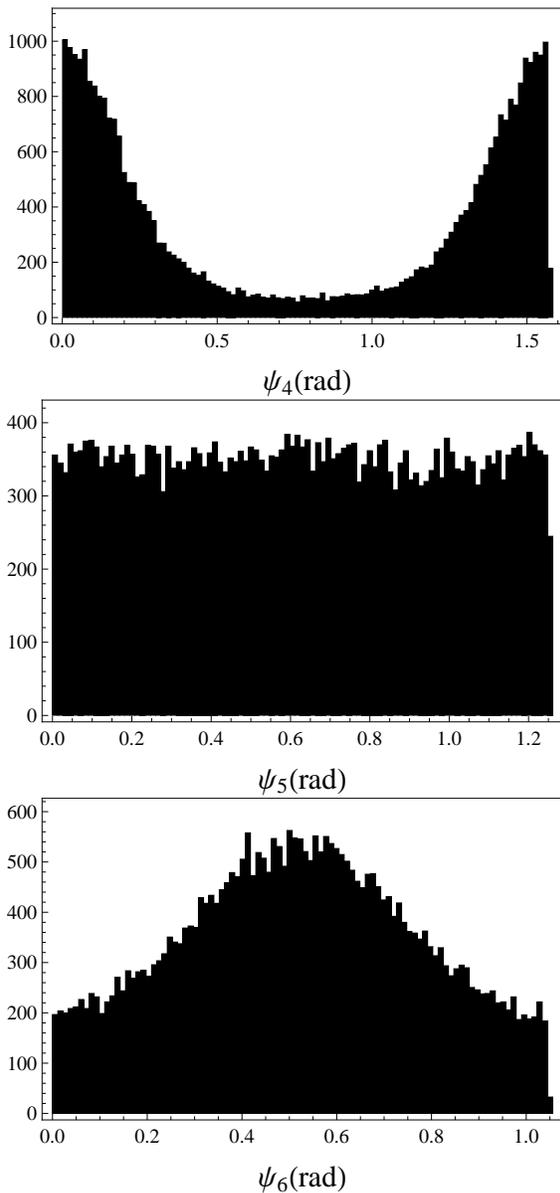


FIG. 5: Distribution of  $\psi_n$  for the harmonics 4-6, the centrality bin used is  $100 < N_{part} < 300$ .

There are two kinds of “stripes”, with positive and negative slope in Fig. 6. The latter ones simply follow from  $\psi_1$  distribution, while the former one is indeed a nontrivial correlation between the angles whose origin we cannot explain. We will continue to discuss its manifestation a bit later. The correlation of  $\psi_5$  with  $\psi_1$  is very similar. The “bumps” at  $\psi_5 - \psi_1 \approx 0$  again mean  $\pm y$  the direction or the “tips”. The plot has similar “stripes”.

Going a bit ahead of ourselves, let us study the “resonant” combinations of angles, as well as angles and amplitudes. As we explain below, those particular combinations of the amplitudes and phases of two harmonics are

$$\langle \epsilon_{n_1} \epsilon_{n_2} \cos(n_1 \psi_{n_1} - n_2 \psi_{n_2}) \rangle \quad (3.11)$$

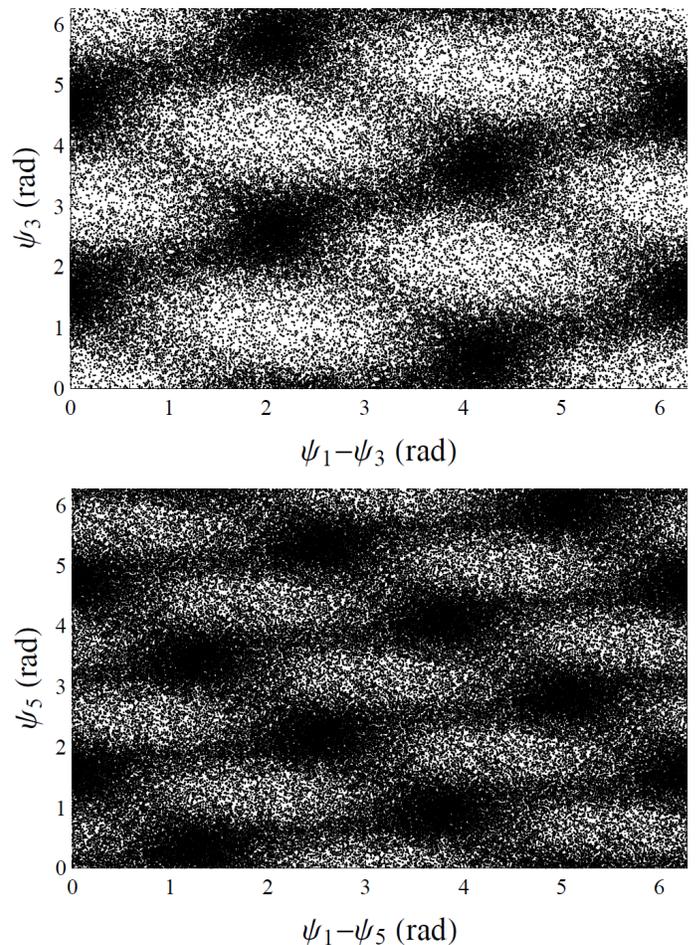


FIG. 6: Scatter plot of the  $\psi_3$  vs  $\psi_3 - \psi_1$  (above), and of the  $\psi_5$  vs  $\psi_5 - \psi_1$  (below), for  $100 < N_{part} < 300$ .

especially in the case when  $n_1, n_2$  differ by two units. We have studied two first examples of the kind, with odd harmonics 1,3,5.

One interesting distribution, shown in Fig.7, is that over the  $\cos$  term itself, for the particular combination of the 1-3 phases. It consists of two clearly different parts: a very narrow peak near  $-1$  and wide flat distributions between  $-1$  and  $1$ . This plot demonstrates a qualitative feature of the phase distribution which was pointed out above. One explanation of the peak near  $-1$  (the angle combination is  $\pi$ ) comes from the fluctuations at the “tips” of the almond, when both  $\psi_1$  and  $\psi_3$  are strongly correlated close to  $\pi/2$ . However, the second interesting correlation seen as “positive slope lines” in Fig.6a because for them  $\psi_1 - 3\psi_3 = \pi$  as well. A similar situation happens for other odd harmonics.

The average value of the combinations (3.11) for 1-3 and 3-5 harmonics as a function of centrality are shown in Fig.8. All values are negative, as the sign is dominated by a peak in  $\cos$  near  $-1$ : the other component more or less averages out. We thus conclude that experimental measurements of the amplitude of such correlations, with the magnitude and the sign, will be especially sensitive

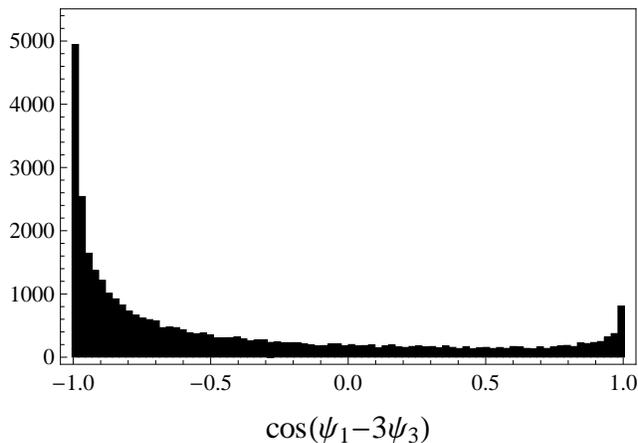


FIG. 7: Scatter plot of  $\cos(3\psi_3 - \psi_1)$

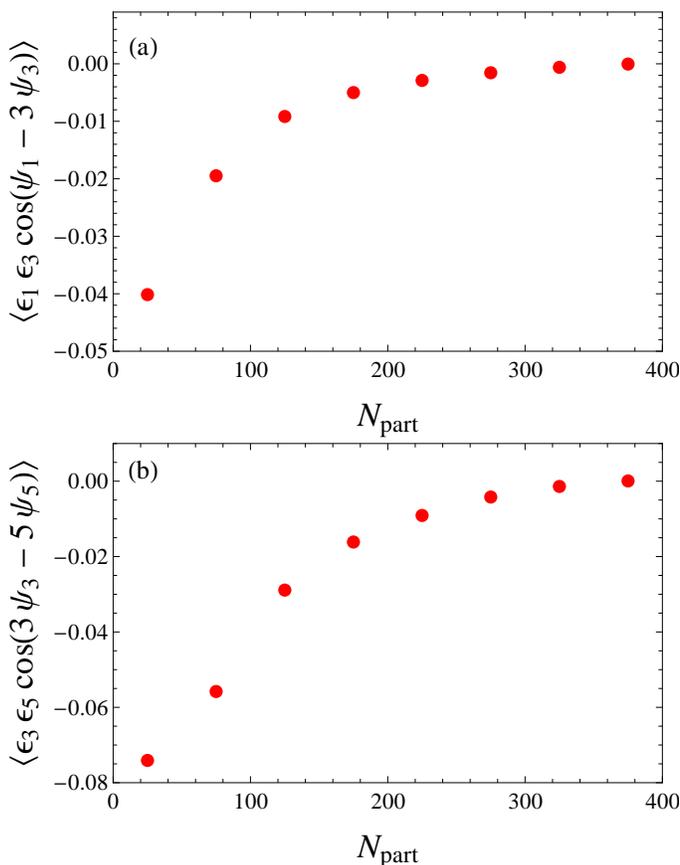


FIG. 8: Correlators  $\langle \epsilon_1 \epsilon_3 \cos(\psi_1 - 3\psi_3) \rangle$  (top) and  $\langle \epsilon_3 \epsilon_5 \cos(3\psi_3 - 5\psi_5) \rangle$  (bottom) as a function of the number of participants. The error bars are omitted since they are smaller than the dots.

to the “almond tip” fluctuations.

Summarizing the observed pattern: we have found that all odd angles are well correlated with each other, forming the “stripes” and “bumps” shown in Fig.6. The fluctuations and correlations seem to be stronger from the “tips” of the almond.

#### D. Comments on other initial state fluctuations

So far the only source of fluctuations included has been (i) the coordinate part of the nuclear wave function prescribing the nucleon positions in the transverse plane, and (ii) the event-by-event fluctuations of the  $NN$  cross section. We found that the former effect dominates and the latter only provides small corrections.

While other sources of fluctuations clearly are subject for future studies, we still provide some comments on those.

One important type of “initial state” fluctuations is of course hard parton scattering events, resulting in jet production. The rate of those very strongly depend on the exact definition of the cutoff beyond which the momentum transfer involved is characterized as “hard”. There are vastly different opinions on where this boundary is theoretically, and experimentally it depends on whether such events are triggered by single large- $p_t$  hadron or by some jet-finding algorithms. Jet production and quenching is of course of high interest, but those should be studied only in a small fraction of all collisions selected by separate triggers. For global fluctuations those can safely be ignored, as the probability of “hard” events is smaller than that of the fluctuations we study.

In the Glauber approach that we used, the local density of produced matter is assumed to be simply proportional to the local density of all participant nucleons,  $N_p(A_1) + N_p(A_2)$ . However, when this density is high enough, it has been argued that the so called “saturation” phenomenon should take place, because of parton absorption processes in the wave function. Other expressions for local matter density have been proposed, e.g.  $\sim \min(N_p) \ln(\max(N_p)/\min(N_p))$  by Kharzeev, Nardi, Levin, where min and max refer to the smaller and the larger of the two. Those are typically the amplitudes of the fluctuations.

A well known approach to their description is the so called “glasma”, which calculates those color fields from random color charges of the leading (larger  $x$ ) partons of the two colliding nuclei. Asymptotically (in a very large nuclei or at very high energy) McLerran and Venugopalan [29] argued that at a particular location in the transverse plane the color charges of partons must be uncorrelated because they come from different nucleons. Therefore it is usually assumed that their color charges fluctuate as random variables. If so, the resulting fluctuations are small, as the total number of partons is very large.

Application of such ideas usually keep the average values of those, such as  $A^{1/3}$  or so. The point of our comment is a warning, that such simplified ideas cannot be used for determining the fluctuations. It has been known for a long time, that while at very small  $x$  the partons, mostly gluons, become numerous, they are still very tightly correlated in the transverse plane. The size of the “gluonic spot” in the nucleon has been known for a long time from diffractive form factors, and in more recent form, from HERA photon diffraction into  $J/\psi$ . This

spot is small, and therefore bright. As documented e.g. in [30] (their Fig.23), the gluon density at the center of the nucleon is about as high as in the center of  $Ca_{40}$ . Therefore, a large number of gluons does not yet imply that all of them merge in the transverse plane into a more or less homogeneous distribution: the positions of the incoming nucleons still remain the dominant source of the initial state fluctuations.

As the “gluonic spots” from single nucleons remain the source of initial state fluctuations, one may ask if numerous partons coming from it may be correlated in their quantum numbers, forming specific large-amplitude color fields. One particular kind of such fields got special attention: those are the topologically nontrivial gluonic configurations called *the QCD sphalerons* [31]. They are gluonic field configurations which originate from excitations of the topologically nontrivial vacuum fields (instantons). While they rapidly explode into multiple gluon state, they strongly violate CP and chiral symmetries locally, producing in particularly  $\pm 2N_f$  ( $\approx 6$ ) units of chirality per sphaleron.

As pointed out in [32], such strong local CP violation induces special event-by-event fluctuations in the CP-odd observables, e.g. they should induce charge asymmetry along the magnetic field. Clearly those should be looked at in special studies.

Another coherent color field configurations which deserve to be specially studied are (colorelectric) flux tubes. In  $pp$  collisions the view that a field created by longitudinally separated charges makes a flux tube is a consequence of confinement, and thus must happen in vacuum. Many popular event generators are based on the Lund model, cleverly parameterizing flux tube production and decay. In AA low energy collisions many flux tubes are produced, and their possible fluctuations into the so called “color ropes” has been studied, initiated by the paper of [33]. If two elementary color charges can be combined, they may either cancel each other or produce higher representations of the SU(3) group, in which case the rope energy (and entropy, after its decay) is proportional to its flux *squared*. Further applications of these ideas for strangeness production in AA collisions can be found in [34].

Studies of the flux tubes lay dormant till recent discovery of the so called “hard ridge”<sup>1</sup> by STAR collaboration at RHIC. One possible origin of it [20] is hydro-carried longitudinal flux tube, created at the hard collision point. This explanation may work provided the flux tubes survive as such till freezeout: as was argued in that paper this indeed should happen at the periphery of the fireball,

where matter is not far from the deconfinement transition, forming a kind of “dual corona” of the QGP fireball similar to Solar corona full of flux tubes.

#### IV. HARMONICS AND THEIR RELATIVE PHASES EXTRACTED FROM CORRELATIONS

##### A. Central collisions: two versus many-particle correlators

Let us for simplicity start with idealized central collisions: if the impact parameter is negligible, one may think of the overall system as completely symmetric in azimuthal angle  $\phi$ .

A particular event has certain perturbation which breaks this symmetry. Its (2-dimensional) distribution over transverse momenta of the secondaries, can be decomposed into Taylor series of the momenta  $p_x, p_y$  or into the angular harmonics

$$\frac{dN}{dp_t^2 d\phi} = f(p_t) \left[ 1 + \sum_{n>0} (2a_n \cos(n\phi) + 2b_n \sin(n\phi)) \right] \\ a_n = \langle \cos(n\phi) \rangle, b_n = \langle \sin(n\phi) \rangle \quad (4.1)$$

(Note that instead of the square bracket one can also use the exponent of the sum, which will include trivial higher order effects and enforces positivity of the distribution: but we would assume all  $v_n$  to be very small, for simplicity.)

Instead of using the  $a, b$  pair, one may also introduce the moduli and the phases writing it as  $2v_n \cos[n(\phi - \xi_n)]$  with positive  $v_n$ . In order to simplify subsequent formulae, we however prefer to write it using the complex exponent

$$\frac{dN}{dp_t^2 d\phi} = f(p_t) \left( \sum_n v_n e^{in\phi - in\xi_n} \right) \quad (4.2)$$

where the sum goes over all integer  $n$ , positive and negative, with  $v_0 = 1$  and  $v_{-n} = v_n$ .

Before going any further, let us give an example of such decomposition for shapes we are interested in. We already mentioned that the sound circles from point perturbations lead to a two-maxima distribution, with angle about 1 rad. To make the points, we select the angle between the peaks to be exactly  $2\pi/3$ , put them at certain positions and give an arbitrary width, for plotting convenience. The result is shown in Fig.9(a) as a (black) solid curve.

We then do the Fourier decomposition and find that for symmetry reasons all coefficient of sinusoidal waves are zero  $b_n = 0$ . To understand the relative magnitude of harmonics one can plot the Fourier “power spectrum”, of the squares  $a_n^2$  versus  $n$ , see Fig.9(b). (This is the information which is included in the two particle correlation function, see below.) The first lesson is that “triangular”

<sup>1</sup> Note that “hard ridge” should not be confused with the “soft ridge”. The latter, as discussed in this and other papers, is naturally explained as a combined effect of “harmonic flows” induced by Glauber fluctuations. Hard ridge, associated with the azimuthal direction of the hard trigger particle, does not yet have a widely accepted explanation.

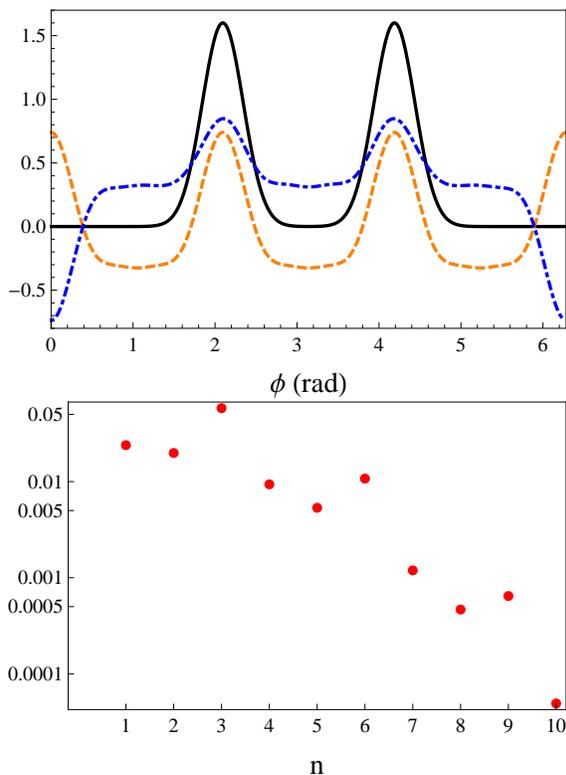


FIG. 9: (color online) (upper) Two-peaked solid (black) curve is the example of single particle angular distribution to be discussed. The (red) dashed curve is the sum of 3,6,9 harmonics, while the (blue) dash-dot ones is the contribution of all other harmonics with  $n < 10$ . (lower) The corresponding “power spectrum” of the harmonics,  $a_n^2$  versus  $n$ .

harmonics 3,6,9... of the type  $3k$ ,  $k$  integer, are significantly enhanced relative to other ones. The reason for that becomes apparent if one compares the sum of such modes (the red dash curve) and the rest of harmonics (the blue dashdotted curve) with the original shape. “Triangular” harmonics get a coherent enhancement from the *three-horn* distribution they describe, while the main role of the rest of the harmonics is to cancel this non-existing third horn and non-existing negative region in between the horns. This interpretation is confirmed by the signs of the harmonics  $a_n$ : all “triangular” ones are positive, all others negative. So, if one would like to reconstruct the distribution from its Fourier series, one may for example find the amplitudes from the power spectrum and the relative signs (phases) of the harmonics.

At the end of chapter 3 we suggested that peripheral enough collisions would result in three-peak distributions, as the sound wave will be able to cross the fireball along its shortest direction. If this happens, and the amplitude of all three would be the same, then the magnitude of “non-triangular” harmonics (other than 3,6,9...) drops further, as there is no need to cancel the non-existing peak.

But why not observe the two-peak event shapes themselves? The reason is that there are multiple perturbations in any event, with random position in the transverse plane. The contribution of an individual fluctuation is very small, so that they can only be studied by finding statistically significant correlations of the particles. Remember that there is  $\sim 10^9$  sample of available events with  $\sim 10^3$  particles per events: two-particle correlators use up to  $\sim 10^6$  pairs of those in an event.

Say, if the elementary perturbation is local (delta-function-like in the transverse plane): its angular position in the transverse plane is the only meaningful azimuthal orientation. The perturbations have random positions in the transverse plane, which we express as

$$\xi_n = \xi_p + \tilde{\xi}_n \quad (4.3)$$

where the tilde indicates the angle *relative* to the perturbation and  $\xi_p$  is the random phase due to location of the perturbation. As  $\xi_p$  is a random variable, all observables should be averaged over it.

What we want to show is that the way the correlations work out is quite different for (i) the two-body and (ii) the many-body (three or more) correlation functions. Indeed, in order to get the two-body correlation function one has to take a the square of the single-body distribution (4.2)

$$\sum_{n_1, n_2} v_{n_1} v_{n_2} \exp[in_1 \phi_1 + in_2 \phi_2 - in_1 \tilde{\xi}_{n_1} - in_2 \tilde{\xi}_{n_2} - i(n_1 + n_2) \xi_p]$$

and average it over  $\xi_p$  randomly distributed over the circle. As a result, in the double sum above only the terms satisfying

$$n_1 + n_2 = 0 \quad (4.4)$$

survive. The double sum collapses into a single sum with the squared amplitude  $e_n^2$ . Second, the sum becomes a function of the angular *difference* between the angles  $\Delta\phi = \phi_1 - \phi_2$ , as expected from the symmetry. And, last but not least, all the *phases*  $\xi_n$  *disappear*. Therefore, two-particle correlators carry the same information as the “power spectrum” of the harmonics (we used above for the example).

This facts are of course well known. The harmonics of the 2-body correlator are

$$C_2(\Delta\phi) = \left\langle \frac{d^2 N}{d\phi_1 d\phi_2} \right\rangle_{|\xi_p} \quad (4.5)$$

$$c_{n\Delta} = \frac{\int d(\Delta\phi) C_2(\Delta\phi) \cos(n\Delta\phi)}{\int d(\Delta\phi) C_2} = \langle v_n^2 \rangle \quad (4.6)$$

or *squared* amplitudes of the original harmonics, *averaged* over the events.

This is e.g. how Alver and Roland [17] and others have obtained their estimates for the “triangular” flow.

However, the situation is different for manybody (three or more) correlation functions. Indeed, if the single-body distribution (4.2) is cubed (or raised into higher power), one finds a *triple* sum in which random phases appears as  $\exp[i(n_1+n_2+n_3)\psi_p]$ , leading now to the “triangular” condition<sup>2</sup>

$$n_1 + n_2 + n_3 = 0 \quad (4.7)$$

Eliminating one of them (e.g.  $n_3$ ) one finds the double sum

$$\sum_{n_1, n_2} v_{n_1} v_{n_2} v_{n_1+n_2} \exp\{i[n_1(\phi_1 - \phi_3) + n_2(\phi_2 - \phi_3) - n_1(\tilde{\xi}_{n_1} - \tilde{\xi}_{n_1+n_2}) - n_2(\tilde{\xi}_{n_1} - \tilde{\xi}_{n_1+n_2})]\}$$

in which the relative phases of different harmonics are still present. Therefore one can still be able to measure the relative phase of harmonics experimentally, by focusing on the corresponding combinations

$$\langle v_{n_1} v_{n_2} v_{n_3} \cos(n_1 \xi_{n_1} + n_2 \xi_{n_2} + n_3 \xi_{n_3}) \rangle \quad (4.8)$$

in which three integers must satisfy the condition (4.7). Obviously the number of solutions to the condition (4.7) is larger than number of harmonics, so all phases can be found. Experimentally, the price to pay is related to the smallness of the harmonics, which in such observables appear in the third (or higher) power. This makes it more difficult to measure, as the values extracted should be larger than the statistical noise.

(It is at this point instructive to get back to our example of the two-peak distribution introduced above and see what those combinations of phase are in this case. One interesting solution to the condition (4.7) is  $1+2=3$ , which relates “enhanced” third harmonic with “subleading” 1 and 2. The sign of this average is defined by  $\cos(\xi_1 + 2\xi_2 - 3\xi_3)$ , and since in our example  $\xi_1 = \xi_2 = \pi, \xi_3 = 0$  this combination of phases leads to  $\cos(\xi_1 + 2\xi_2 - 3\xi_3) = -1$ . But if one wants to measure the “enhanced” harmonics themselves, like in the  $3+3=6, 6+3=9$  etc, the corresponding  $\cos$  would be close to 1 instead. Proceeding in this way one should be able to find such signs and conclude if the two-peak distributions are or are not close to reality.)

Although hydro calculations is a subject of another paper, let us briefly discuss how comparison to theory should be done, assuming that the corresponding averages for some set of  $n_1, n_2, n_3$  are experimentally measured. There are two steps to be done. First, using the

(linearized) hydro one can approximate flow harmonics by initial ones, time the ratios

$$\begin{aligned} & \langle v_{n_1} v_{n_2} v_{n_3} \cos(n_1 \xi_{n_1} + n_2 \xi_{n_2} + n_3 \xi_{n_3}) \rangle \\ &= \left(\frac{v_{n_1}}{\epsilon_{n_1}}\right) \left(\frac{v_{n_2}}{\epsilon_{n_2}}\right) \left(\frac{v_{n_3}}{\epsilon_{n_3}}\right) \langle \epsilon_{n_1} \epsilon_{n_2} \epsilon_{n_3} \cos(n_1 \xi_{n_1} + n_2 \xi_{n_2} + n_3 \xi_{n_3}) \rangle \end{aligned} \quad (4.9)$$

Second, one should change the flow angles we call  $\xi_n$  to the deformation angles  $\psi_n$ . One simple step is phase shift by (3.2), the angle between flow and initial deformation. Note however that in each three terms  $n_i$  in numerator and denominator cancel, leaving only  $3\pi$  or simply the total sign change

$$\begin{aligned} & \cos(n_1 \xi_{n_1} + n_2 \xi_{n_2} + n_3 \xi_{n_3}) = \\ & -\cos(n_1 \psi_{n_1} + n_2 \psi_{n_2} + n_3 \psi_{n_3}) \end{aligned} \quad (4.10)$$

The last step is to check for the direction (sign) of the flow: as we will see in [18] flow direction at freezeout oscillates as a function of  $n$ , so this sign should also be included in the phase. The resulting correlation of the amplitudes and orientations of the initial state fluctuations can be calculated from initial state model, as we have done above for the first three harmonics resonance  $1+2=3$  as well as  $3+2=5$ .

In fact it is not necessary to take Fourier moments of the correlation functions, but plot the correlation functions themselves. (This is especially useful if the experimental angular coverage in azimuth is incomplete.)

Defining the 3-particle correlator as for the 2 particle one, with the averaging over the perturbation angle

$$C_3(\phi_1, \phi_2, \phi_3) = \langle \frac{d^2 N}{d\phi_1 d\phi_2 d\phi_3} \rangle_{\psi_p} \quad (4.11)$$

One finds it for central collisions to be a function of the two angle differences, say  $\phi_1 - \phi_3, \phi_2 - \phi_3$ . The corresponding 2d plots for the 2-peak and 3-peak distributions look quite different, and should be relatively easy to separate.

## B. Mid-central collisions and the two-body correlators relative to the event plane

Nonzero impact parameter violates axial symmetry and creates “directed flows”, e.g, the famous elliptic flow with nonzero  $\langle v_2 \rangle \neq 0$ . By mid-central collisions we mean a centrality region in which  $\langle \epsilon_2 \rangle$  is large itself, and is also large compared to its fluctuations (recall that it is not so for central and very peripheral collisions). For example,  $\epsilon_2$  is 0.5 (0.3) for  $N_p = 100$  (200) participants, with  $\delta\epsilon_2 \approx 0.1$ . Furthermore, as seen in Fig.4b, its angle  $\psi_2$  is very much directed at  $\pm\pi/2$  (the tips of the almond) and therefore (using (3.2 for  $n=2$ ) the flow angle is peaked “in-plane”,  $\xi_2 = 0, \pi$ , as indeed observed.

If so, for one of the harmonics being the second e.g.  $n_3 = \pm 2$  one can approximate a product of three defor-

<sup>2</sup> Long after this paper has been posted we learned that a similar condition was proposed to be used to look for non-Gaussianity of the fluctuations of CMB. Our integers conjugated to azimuthal angle are then promoted to 2-d angular momenta  $l_i$  on the sky, and so the condition means that three vectors can make a closed triangle. Thus the “triangular” name. To our knowledge, non-Gaussianity of the Big Bang is not yet observed.

mations as follows

$$\begin{aligned} & \langle v_{n_1} v_{n_2} v_{n_3} \cos(n_1 \xi_{n_1} + n_2 \xi_{n_2} + n_3 \xi_{n_3}) \rangle \quad (4.12) \\ & \approx \left(\frac{v_{n_1}}{\epsilon_{n_1}}\right) \left(\frac{v_{n_2}}{\epsilon_{n_2}}\right) \left(\frac{v_2}{\epsilon_2}\right) \langle \epsilon_2 \rangle \langle \epsilon_{n_1} \epsilon_{n_2} \cos(n_1 \xi_{n_1} + n_2 \xi_{n_2}) \rangle \end{aligned}$$

by factoring out large and non-fluctuating  $\langle \epsilon_2 \rangle$  from two other harmonics which are small and fluctuating. Note that the resonance condition now means  $n_1 \pm 2 = n_2$ , and that by putting  $\xi_2 = 0$  we have selected a frame in which the (experimentally determined reaction plane) is the x axis.

Basically the lesson here is that for mid-central collisions the “reaction plane” plays the same role as the third body, so we are reduced to two small fluctuating and correlated harmonics. The simplest nontrivial example of resonance condition of the kind is 3-1-2=0 (recently studied by Teaney and Yan [35]), while the next is 5-3-2=0.

We had already calculated the combinations of two fluctuating harmonics with the appropriate cosines above, for these two cases, in the Glauber model. They are by no means small: for example for the centrality bin with [100..300] participants they are

$$\langle \epsilon_1 \epsilon_3 \cos(3\psi_3 - \psi_1) \rangle \sim -0.015 \quad (4.13)$$

$$\langle \epsilon_3 \epsilon_5 \cos(3\psi_3 - 5\psi_5) \rangle \sim -0.05 \quad (4.14)$$

and therefore we expect it to be observable, with about as large statistics as needed for the usual quadratic fluctuations.

## V. SUMMARY

In this work we have (i) discussed the setting, identifying the main scales of the problem. Then (ii) we studied in detail the initial state fluctuations originating from nucleon positions, emphasizing existence of the nontrivial phase relations between different harmonics. Finally, (iii) we discussed correlations functions of 2 and many hadrons, pointed out the principle difference between them, with the latter allowing to measure experimentally the relative phases of these harmonics.

Let us now recapitulate the lessons from this study in a bit more detail.

Unfortunately, the perturbations we speak about are too small to be measured directly, on event-by-event basis, and should be instead reconstructed from the statistically obtained correlation functions. One good thing coming from it is that many independent fluctuations from local perturbations in a single event and in the ensemble are treated by correlation function, in which all trivial uncorrelated effects are statistically subtracted and absent.

Traditionally the initial state perturbations and final state corrections to collective flow are considered in a form of their angular harmonics, which we call  $\epsilon_n, v_n$ ,

respectively. Their relation is calculated by the hydrodynamics: the details of that in the linearized form is the subject of our companion paper [18].

Most papers on the subject consider  $\epsilon_n, v_n$  as independent random variables, which are incoherent fluctuations, added in quadratures, and ignored their phases  $\xi_n, \psi_n$ . We however pointed out that while it can be done so for the two-particle correlations, it is not so in general. Correlators of many (3 and more) particle correlations include those phases of, as do two-body correlations relative to event plane for mid-central collisions. Many different “triplets”  $n_1, n_2, n_3$  related by the “triangular condition” between them can be measured, providing experimental opportunity to find out if and how the perturbations are correlated. This needs to be done, to refine models of the initial state. The magnitude of few such terms we have estimated in the Glauber model, and they are comparable in magnitude to the terms already studied. We also presented theoretical toy models, in which phases between harmonics are especially simple.

Coherence in phases of the deformations imply the interferences between the harmonics of the flow. Only adding them together one can follow how small “hot” (or “cold”) spots created by quantum fluctuations of interacting nucleons propagate away from the point of origin. Only in this way one can understand the role played by the “hydrodynamical causality”, insisting that large part of the fireball should remain completely unaffected by the perturbation since the signal cannot possibly reach it before the freezeout. Only a complete Green function, collecting all hydro harmonics, describes these shapes of propagating waves, as we detail in [18].

There are two basic scales defining those perturbations, the sound horizon  $H_s$  (2.2) and the “the viscous horizon scale”  $R_v$  (2.6). The former gives the size of the perturbation, stemming from a local perturbation, the second its width. We have for example argued that by changing the centrality of the collisions, one can change the relation between the (smaller) fireball size and the sound horizon: this should dramatically change the shape of the event (see Fig.4 for explanation).

It is an important objective of the experimental heavy ion program in general to measure these two scales, extracting experimental values of the speed of sound and viscosity. One specific idea proposed in this work is that by changing centrality one can locate transition when  $H_s$  and the smaller size of the overlap region are equal, observing the change of shape of the underlying event, from those with two peaks to three-peaks ones. If found, it would be a very spectacular confirmation of the view that sound waves can travel large distances  $\sim R$  during heavy ion collisions. It will also put to rest various models which assume significantly shorter freezeout time than predicted by hydrodynamics.

**Acknowledgments.** This work was supported in parts by the US-DOE grant DE-FG-88ER40388. Discussion with Frederique Grassi, explaining their finding

in detail, have been very helpful. We also benefited from multiple discussion with Derek Teaney.

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