



This is the accepted manuscript made available via CHORUS. The article has been published as:

Initial eccentricity fluctuations and their relation to higherorder flow harmonics

Roy A. Lacey, Rui Wei, J. Jia, N. N. Ajitanand, J. M. Alexander, and A. Taranenko Phys. Rev. C **83**, 044902 — Published 8 April 2011

DOI: 10.1103/PhysRevC.83.044902

Initial eccentricity fluctuations and their relation to higher-order flow harmonics

Roy A. Lacey, ^{1,2,*} Rui Wei, ¹ J. Jia, ^{1,2} N. N. Ajitanand, ¹ J. M. Alexander, ¹ and A. Taranenko¹ ¹Department of Chemistry, Stony Brook University, Stony Brook, NY, 11794-3400, USA ²Physics Department, Bookhaven National Laboratory, Upton, New York 11973-5000, USA (Dated: March 21, 2011)

Monte Carlo simulations are used to compute the centrality dependence of the participant eccentricities (ε_n) in Au+Au collisions, for the two primary models currently employed for eccentricity estimates - the Glauber and the factorized Kharzeev-Levin-Nardi (fKLN) models. They suggest specific testable predictions for the magnitude and centrality dependence of the flow coefficients v_n , respectively measured relative to the event planes Ψ_n . They also indicate that the ratios of several of these coefficients may provide an additional constraint for distinguishing between the models. Such a constraint could be important for a more precise determination of the specific viscosity of the matter produced in heavy ion collisions.

PACS numbers: 25.75.-q, 25.75.Dw, 25.75.Ld

[1–16]. An experimental manifestation of this flow is the anisotropic emission of particles in the plane transverse to the beam direction [17, 18]. This anisotropy can be characterized by the even order Fourier coefficients;

$$v_{\rm n} = \left\langle e^{in(\phi_p - \Psi_{RP})} \right\rangle, \ n = 2, 4, .., \tag{1}$$

where ϕ_p is the azimuthal angle of an emitted particle, Ψ_{RP} is the azimuth of the reaction plane and the brackets denote averaging over particles and events [19]. Characterization has also been made via the pair-wise distribution in the azimuthal angle difference ($\Delta \phi = \phi_1 - \phi_2$) between particles [17, 20, 21];

$$\frac{dN^{\text{pairs}}}{d\Delta\phi} \propto \left(1 + \sum_{n=1}^{\infty} 2v_n^2 \cos(n\Delta\phi)\right). \tag{2}$$

Anisotropic flow is understood to result from an asym-11 metric hydrodynamic-like expansion of the medium pro-12 duced by the two colliding nuclei. That is, the spacial 13 asymmetry of the produced medium drives uneven pres-14 sure gradients in- and out of the reaction plane and hence, 15 a momentum anisotropy of the particles emitted about 16 this plane. This mechanistic picture is well supported by 17 the observation that the measured anisotropy for hadron $_{18}~p_T \lesssim 2~{\rm GeV/c},$ can be described by relativistic hydrody-19 namics [5, 10, 12, 14, 15, 22–31].

The differential Fourier coefficients $v_2(N_{part})$ and $v_2(p_T)$ have been extensively studied in Au+Au collisions

Collective flow continues to play a central role in on- 22 at RHIC [20, 32–38]. One reason for this has been the going efforts to characterize the transport properties of 23 realization that these elliptic flow coefficients are sensithe strongly interacting matter produced in heavy ion 24 tive to various transport properties of the expanding hot collisions at the Relativistic Heavy Ion Collider (RHIC) 25 medium [5-7, 9, 11, 13, 23, 39-41]. Indeed, considerable 26 effort has been, and is being devoted to the quantita-27 tive extraction of the specific shear viscosity η/s (i.e. the 28 ratio of shear viscosity η to entropy density s) via com-²⁹ parisons to viscous relativistic hydrodynamic simulations 30 [9-12, 14, 15, 30], transport model calculations [6, 13, 42] (1) 31 and hybrid approaches which involve the parametriza-32 tion of scaling violations to ideal hydrodynamic behavior ³³ [7, 16, 40, 43, 44]. The initial eccentricity of the collision 34 zone and its associated fluctuations, has proven to be an essential ingredient for these extractions.

> Experimental measurements of the eccentricity have 37 not been possible to date. Consequently, much reliance 38 has been placed on the theoretical estimates obtained 39 from the overlap geometry of the collision zone, speci-40 fied by the impact parameter b or the number of par-41 ticipants N_{part} [31, 34, 43, 45–52]. For these estimates, 42 the geometric fluctuations associated with the positions 43 of the nucleons in the collision zone, serve as the under-44 lying cause of the initial eccentricity fluctuations. That 45 is, the fluctuations of the positions of the nucleons lead 46 to fluctuations of the so-called participant plane (from 47 one event to another) which result in larger values for the eccentricities (ε) referenced to this plane.

> The magnitude of these fluctuations are of course 50 model dependent, and this leads to different predictions 51 for the magnitude of the eccentricity. More specifically, ₅₂ the ε_2 values obtained from the Glauber [34, 53] and the 53 factorized Kharzeev-Levin-Nardi (fKLN) [54, 55] models, 54 (the two primary models currently employed for eccen-55 tricity estimates) give results which differ by as much $_{56}$ as $\sim 25\%$ [56, 57] – a difference which leads to an ap- $_{57}$ proximate factor of two uncertainty in the extracted η/s ₁ value [9, 16]. Thus, a more precise extraction of η/s re-

^{*} E-mail: Roy.Lacey@Stonybrook.edu

32

2 quires a clever experimental technique which can measure 3 the eccentricity and/or the development of experimental 4 constraints which can facilitate the requisite distinction 5 between the models used to calculate eccentricity.

Recently, significant attention has been given to the 7 study of the influence of initial geometry fluctuations on $_8$ higher order eccentricities $\varepsilon_{n,n\geq 3}$ [30, 31, 47, 50–52, 58– 9 60, with an eye toward a better understanding of how 10 such fluctuations manifest into the harmonic flow corre-11 lations characterized by v_n (for odd and even n), and 12 whether they can yield constraints that could serve to 13 pin down the "correct" model for eccentricity determina-14 tion. For the latter, the magnitude of ε_n and its detailed 15 centrality dependence is critical. Therefore, it is essen-16 tial to resolve the substantial differences in the ε_n values 17 reported and used by different authors [30, 31, 47, 50-

Here, we argue that the magnitudes and trends for 20 the eccentricities ε_n imply specific testable predictions 21 for the magnitude and centrality dependence of the flow v_n , measured relative to their respective event 23 planes Ψ_n . We also show that the values for ε_n obtained 24 for the Glauber [34, 53] and fKLN [54, 55] models, in-25 dicate sizable model dependent differences which could ²⁶ manifest into experimentally detectable differences in the 27 centrality dependence of the ratios $v_3/(v_2)^{3/2}$, $v_4/(v_2)^2$ 28 and $v_2/v_{n,n>3}$. Such a constraint could be important for 29 a more precise determination of the specific viscosity of 30 the hot and dense matter produced in heavy ion colli-31 sions.

ECCENTRICITY SIMULATIONS

Monte Carlo (MC) simulations were used to calculate event averaged eccentricities (denoted here as ε_n) in Au+Au collisions, within the framework of the Glauber (MC-Glauber) and fKLN (MC-KLN) models. For each event, the spatial distribution of nucleons in the colliding nuclei were generated according to the Woods-Saxon function:

$$\rho(\mathbf{r}) = \frac{\rho_0}{1 + e^{(\mathbf{r} - R_0)/d}},\tag{3}$$

where $R_0 = 6.38 \, \text{fm}$ is the radius of the Au nucleus and $_{34}$ $d=0.53\,\mathrm{fm}$ is the diffuseness parameter.

For each collision, the values for N_{part} and the num- $_{36}$ ber of binary collisions $N_{\rm coll}$ were determined within the 37 Glauber ansatz [53]. The associated ε_n values were then 38 evaluated from the two-dimensional profile of the density of sources in the transverse plane $\rho_s(\mathbf{r}_\perp)$, using modified 40 versions of MC-Glauber [53] and MC-KLN [55] respec-41 tively.

harmonic of the shape profile [47, 50];

$$S_{nx} \equiv S_n \cos(n\Psi_n) = \int d\mathbf{r}_{\perp} \rho_s(\mathbf{r}_{\perp}) \omega(\mathbf{r}_{\perp}) \cos(n\phi)$$

$$S_{ny} \equiv S_n \sin(n\Psi_n) = \int d\mathbf{r}_{\perp} \rho_s(\mathbf{r}_{\perp}) \omega(\mathbf{r}_{\perp}) \sin(n\phi)$$

$$\Psi_n = \frac{1}{n} \tan^{-1} \left(\frac{S_{ny}}{S_{nx}}\right), \tag{6}$$

42 where ϕ is the azimuthal angle of each source and the 43 weight $\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{2}$ and $\widetilde{\omega}(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{n}$ are used in 44 respective calculations. Here, it is important to note 45 that the substantial differences reported for ε_n in Refs. 46 [30, 31, 47, 50–52, 58–60] is largely due to the value of 47 $\omega(\mathbf{r}_{\perp})$ employed.

The eccentricities were calculated as:

$$\varepsilon_n = \langle \cos n(\phi - \Psi_n) \rangle \tag{7}$$

and

$$\varepsilon_n^* = \langle \cos n(\phi - \Psi_m) \rangle, \ n \neq m.$$
 (8)

48 where the brackets denote averaging over sources and 49 events belonging to a particular centrality or impact pa-50 rameter range; the starred notation is used here to dis- $_{51}$ tinguish the n-th order moments obtained relative to an 52 event plane of a different order Ψ_m .

For the MC-Glauber calculations, an additional entropy density weight was applied reflecting the combination of spatial coordinates of participating nucleons and binary collisions [48, 56];

$$\rho_s(\mathbf{r}_\perp) \propto \left[\frac{(1-\alpha)}{2} \frac{dN_{\text{part}}}{d^2 \mathbf{r}_\perp} + \alpha \frac{dN_{\text{coll}}}{d^2 \mathbf{r}_\perp} \right], \tag{9}$$

where $\alpha = 0.14$ was constrained by multiplicity measurements as a function of N_{part} for Au+Au collisions [61]. 55 These procedures take account of the eccentricity fluctu-56 ations which stem from the event-by-event misalignment 57 between the short axis of the "almond-shaped" collision ₅₈ zone and the impact parameter. Note that ε_n (cf. Eq. 7) v_n measurements relative to the so-called 60 participant planes [34, 53]. That is, each harmonic ε_n 61 is evaluated relative to the principal axis determined by $_{62}$ maximizing the n-th moment. This is analogous to the ₆₃ measurement of v_n with respect to the n-th order event-64 plane in actual experiments [62]. It however, contrasts 65 recent experimental measurements in which a higher or-66 der coefficient (v_4) has been measured with respect to a ₆₇ lower order event plane (Ψ_2) [38, 63]. Note as well that 68 we have established that the angles Ψ_n for the odd and 69 even harmonics are essentially uncorrelated for the $N_{\rm part}$ 70 range of interest to this study.

A. Results for
$$\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{2}$$
 and $\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{n}$

For each event, we compute an event shape vector S_n 73 Figure 1 shows a comparison of $\varepsilon_{n,n\leq 6}$ vs. $N_{\rm part}$ for and the azimuth of the the rotation angle Ψ_n for n-th 74 $\omega(\mathbf{r}_\perp) = \mathbf{r}_\perp^2$, for MC-Glauber (a) and MC-KLN (b) for

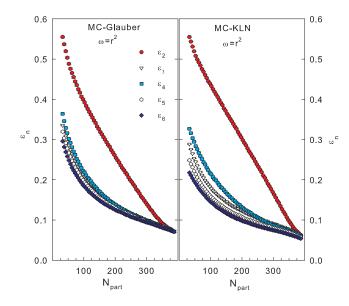


FIG. 1. Calculated values of $\varepsilon_{n,n\leq 6}$ vs. N_{part} for $\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^2$ for MC-Glauber (a) and MC-KLN (b) for Au+Au collisions. The open and filled symbols indicate the results for odd and even harmonics respectively.

75 Au+Au collisions. The filled and open symbols indi-1 cate the results for the even and odd harmonics respec-² tively. For this weighting scheme, ε_n is essentially the $_3$ same for $n \geq 3$, and have magnitudes which are signifi-4 cantly less than that for ε_2 , except in very central colli-5 sions where the effects of fluctuation dominate the mag-6 nitude of $\varepsilon_{n,n\geq 2}$. Note the approximate $1/\sqrt(N_{\text{part}})$ de-7 pendence for $\overline{\varepsilon}_{n,n\geq 3}$. The smaller magnitudes for $\varepsilon_{n,n\geq 3}$ 8 (with larger spread) apparent in Fig. 1(b), can be at-9 tributed to the sharper transverse density distributions 10 for MC-KLN.

Figure 2 shows a similar comparison of $\varepsilon_{n,n\leq 6}$ vs. N_{part} 12 for calculations performed with the weight $\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{n}$. 13 This weighting results in an increase in the sensitivity 14 to the outer regions of the transverse density distribu-15 tions. Consequently, the overall magnitudes for $\varepsilon_{n,n>3}$ ₁₆ are larger than those shown in Fig. 1. This weighting 17 also lead to a striking difference in the relative magni-¹⁸ tudes of $\varepsilon_{n,n\geq 2}$ for MC-Glauber (a), MC-KLN (b) and ¹⁹ the results for $\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{2}$ shown in Fig. 1.

ECCENTRICITY RATIOS II.

20

The magnitudes and trends of the calculated eccentric-22 ities shown in Figs. 1 and 2 are expected to influence the 23 measured values of v_n . To estimate this influence, we 24 first assume that the resulting anisotropic flow is directly ₂₅ proportional to the initial eccentricity, as predicted by 26 perfect fluid hydrodynamics. Here, our tacit assumption 27 is that a possible influence from the effects of a finite $\varepsilon_{n,n\geq 3}$ on the values for $v_{n,n\geq 3}$ should be similar irre-

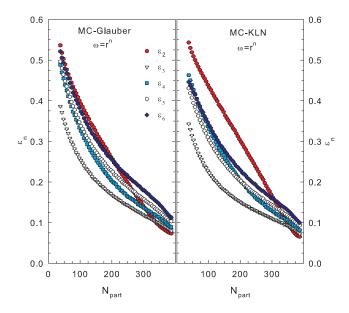


FIG. 2. Same as Fig. 1 for $\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{n}$.

29 cate that η/s is small [4, 6, 7, 9–16, 30, 40, 43, 44] – of $_{30}$ the same magnitude as for the conjectured KSS bound $\eta/s = 1/(4\pi)$ [64].

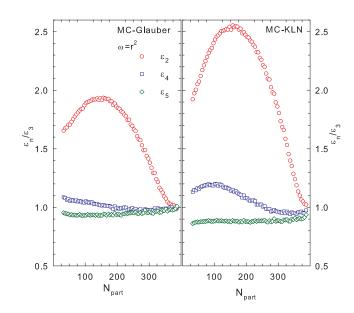


FIG. 3. Comparison of $\varepsilon_{2,4,5}/\varepsilon_3$ vs. $N_{\rm part}$ for Au+Au collisions. Results are shown for MC-Glauber (a) and MC-KLN (b) calculations.

Figure 1 indicates specific testable predictions for the 33 relative influence of $\varepsilon_{n,n\geq 2}$ on the magnitudes of $v_{n,n\geq 2}$. That is, (i) ε_2 should have a greater influence than $\varepsilon_{n,n>3}$ 35 in non-central collisions, (ii) the respective influence of 28 viscosity (η/s) is small because current estimates indi- 37 spective of centrality and (iii) the ratios $v_{4,5,6}/v_3$ should

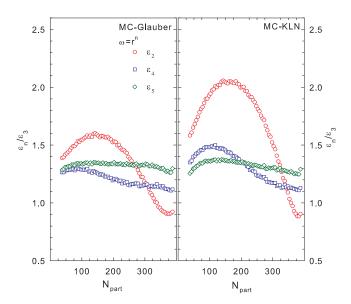


FIG. 4. Same as Fig. 3 for $\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{n}$.

follow a specific centrality dependence due to the influence of $\varepsilon_{4,5,6}/\varepsilon_3$. Such a dependence is illustrated in Fig. 3 where we show the centrality dependence of the ratios $\varepsilon_{2,4,5}/\varepsilon_3$, obtained for MC-Glauber (a) and MC-KLN (b) 4 calculations. They suggest that, if MC-Glauber-like eccentricities, with weight $\omega(\mathbf{r}_\perp) = \mathbf{r}_\perp^2$, are the relevant 6 eccentricities for Au+Au collisions, then the measured 7 ratio v_2/v_3 should increase by a factor ≈ 2 , from central to mid-central collisions ($N_{\rm part} \sim 350-150$). For 9 $N_{\rm part} \lesssim 150$, Fig. 2(a) shows that the ratio v_2/v_3 could 10 even show a modest decrease. The eccentricity ratios in 11 volving the higher harmonics suggest that, if they are 12 valid, the measured values of $v_{4,5,6}/v_3$ should show little, if any, dependence on centrality, irrespective of their 14 magnitudes.

The ratios $\varepsilon_{2,4,5}/\varepsilon_3$ obtained for MC-KLN calculations are shown in Fig. 3 (b). While they indicate qualitative trends which are similar to the ones observed in Fig. 3 (a), their magnitudes and their detailed dependence on centrality are different. Therefore, if the qualitative trends discussed earlier were indeed found in data, then these differences suggest that precision measurements of the centrality dependence of the relative ratios for v_2/v_3 , v_4/v_3 , v_5/v_3 , ... for several p_T selections, could provide a constraint for aiding the distinction between fKLN-like and Glauber-like initial collision geometries. Specifically, smaller (larger) values of the relative ratios are to be expected for v_2/v_3 and v_4/v_3 for Glauber-like (fKLN-like) initial geometries. Note the differences in the expected centrality dependencies as well.

Figure 4 compares the eccentricity ratios $\varepsilon_{2,4,5}/\varepsilon_3$ ob- $\varepsilon_{2,4,5}/\varepsilon_{3,0}$ to stantially larger than those in Fig. 5. The latter dif- $\varepsilon_{2,2}$ tained for MC-Glauber (a) and MC-KLN (b) calcula- $\varepsilon_{3,2}$ tions with the weight $\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{n}$. The magnitudes $\varepsilon_{4,2}$ of these ratios and their centrality dependencies are dis- $\varepsilon_{4,2}$ tinct for MC-Glaber and MC-KLN. They are also quite $\varepsilon_{4,2}$ ratios for $\varepsilon_{4,2}/(\varepsilon_{2})^{2}$ imply much larger measured $\varepsilon_{4,2}$ ratios for $\varepsilon_{4,2}/(\varepsilon_{2})^{2}$ than the value of 0.5 predicted by per-

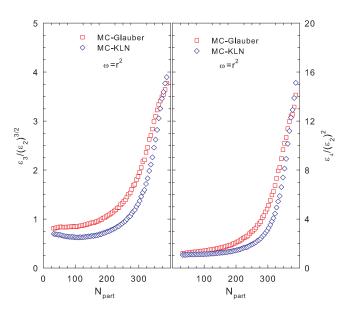


FIG. 5. Comparison of $\varepsilon_3/(\varepsilon_2)^{3/2}$ vs. $N_{\rm part}$ (a) and $\varepsilon_4/(\varepsilon_2)^2$ vs. $N_{\rm part}$ (b) for MC-Glauber and MC-KLN initial geometries (as indicated) for Au+Au collisions.

 $_{36}$ different from the ratios shown in Fig. 3. This suggests $_{37}$ that precision measurements of the centrality dependence $_{38}$ of the relative ratios $v_2/v_3,\,v_4/v_3,\,v_5/v_3,...$ (for several $_{39}$ p_T selections) should not only allow a clear distinction between MC-Glauber and MC-KLN initial geometries, $_{41}$ but also a distinction between the the $\omega(\mathbf{r}_\perp)=\mathbf{r}_\perp{}^2$ and $_{42}$ $\omega(\mathbf{r}_\perp)=\mathbf{r}_\perp{}^n$ weighting methods.

A finite viscosity will influence the magnitudes of v_n . Thus, for a given p_T selection, the measured ratios for $v_2/v_3, \ v_4/v_3, \ v_5/v_3, \dots$ will be different from the eccentricity ratios shown in Figs. 3 and 4. Note as well that, even for ideal hydrodynamics, the predicted magnitude of v_4/ε_4 is only a half of that for v_2/ε_2 [59]. Nonetheless, the rather distinct centrality dependent eccentricity patterns exhibited in Figs. 3 and 4 suggests that measurements of the ratios of these flow harmonics should still allow a distinction between MC-Glauber and MC-mathematical KLN initial geometries, as well as a distinction between the two weighting methods.

The ratios $v_3/(v_2)^{3/2}$ and $v_4/(v_2)^2$ have been recently found to scale with p_T [65], suggesting a reduction in the influence of viscosity on them. Thus, the measured ratios $v_n/(v_2)^{n/2}$ could give a more direct indication of the centrality dependent influence of $\varepsilon_n/(\varepsilon_2)^{n/2}$ on $v_n/(v_2)^{n/2}$. The open symbols in Figs. 5 and 6 indicate a substantial difference between the ratios $\varepsilon_3/(\varepsilon_2)^{3/2}$ (a) and $\varepsilon_4/(\varepsilon_2)^2$ (b) for the MC-Glauber and MC-KLN geometries as instantially larger than those in Fig. 5. The latter difference reflects the different weighting schemes used, i.e. $\omega(\mathbf{r}_\perp) = \mathbf{r}_\perp^n$ and $\omega(\mathbf{r}_\perp) = \mathbf{r}_\perp^2$ respectively. Interestingly, the ratios for $\varepsilon_4/(\varepsilon_2)^2$ imply much larger measured ratios for $v_4/(v_2)^2$ than the value of 0.5 predicted by per-

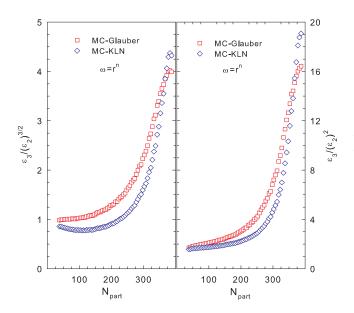


FIG. 6. Same as Fig. 5 for $\omega(\mathbf{r}_{\perp}) = \mathbf{r}_{\perp}^{n}$.

fect fluid hydrodynamics (without fluctuations) [66, 67]. However, they show qualitative trends which are similar to those for the measured ratios $v_4/(v_2)^2$, obtained for v_4 evaluations relative to the Ψ_2 plane[38, 63]. The relatively steep rise of the ratios in Figs. 5 and 6 (albeit steeper for MC-Glauber), can be attributed to the larger influence that fluctuations have on the higher harmonics. Note that these are the same fluctuations which give rise to the "anomalously low" values of ε_4 evaluated with respect to Ψ_2 in central collisions [50].

Figures 3 - 6 suggests that measurements of the centrality dependence of the ratios $v_3/(v_2)^{3/2}$ and $v_4/(v_2)^2$, in conjunction with those for v_2/v_3 , v_4/v_3 , v_5/v_3 ... may provide a robust constraint for the role of initial eccentricity fluctuations, as well as an additional handle for making a distinction between Glauber-like and fKLN-16 like initial geometries. These measurements could also lend insight, as well as place important constraints for the degree to which a small value of η/s and/or the effects of thermal smearing, modulate the higher order flow harmonics [compared to v_2] as has been suggested [31, 52, 60].

III. SUMMARY

23

In summary, we have presented results for the initial eccentricities $\varepsilon_{n,n\leq 6}$ for Au+Au collisions with different weighting schemes, for the two primary models currently employed for eccentricity estimates at RHIC. The calculated values of $\varepsilon_{n,n\leq 6}$, which are expected to influence the measured flow harmonics v_n , suggests that measurements of the centrality dependence of $v_2/(v_3)$, v_4/v_3 , $v_3/(v_2)^{3/2}$, $v_4/(v_2)^2$, etc. could provide stringent constraints for vali-

dating the predicted influence of eccentricity fluctuations on v_n , as well as an important additional handle for making a distinction between Glauber-like and fKLN-like initial geometries. Measurements of v_n and their ratios are now required to exploit these simple tests.

Acknowledgments We thank Wojciech Broniowski for profitable discussions and invaluable model calculation cross checks. This research is supported by the US 40 DOE under contract DE-FG02-87ER40331.A008 and by 41 the NSF under award number PHY-1019387.

18

- [1] M. Gyulassy and L. McLerran, Nucl. Phys. A750, 30 42
- [2] D. Molnar and P. Huovinen, Phys. Rev. Lett. 94, 012302 (2005), arXiv:nucl-th/0404065.
- R. A. Lacey et al., Phys. Rev. Lett. 98, 092301 (2007).
- A. Adare et al., Phys. Rev. Lett. 98, 172301 (2007).
- P. Romatschke and U. Romatschke, Phys. Rev. Lett. 99, 172301 (2007).
- [6] Z. Xu, C. Greiner, and H. Stocker, Phys. Rev. Lett. 101, 082302 (2008).
- H.-J. Drescher, A. Dumitru, C. Gombeaud, and J.-Y. 10 Ollitrault, Phys. Rev. C76, 024905 (2007). 11
- [8] E. Shuryak, Prog. Part. Nucl. Phys. 62, 48 (2009). 12
- [9] M. Luzum and P. Romatschke, Phys. Rev. C78, 034915 13 (2008).14
- 15 [10] H. Song and U. W. Heinz, J. Phys. **G36**, 064033 (2009).
- [11] A. K. Chaudhuri, (2009), arXiv:0910.0979 [nucl-th]. 16
- [12] K. Dusling and D. Teaney, Phys. Rev. C77, 034905 17 (2008), arXiv:0710.5932 [nucl-th].
- V. Greco, M. Colonna, M. Di Toro, and G. Ferini, 19 (2008), arXiv:0811.3170 [hep-ph]. 20
- P. Bozek and I. Wyskiel, PoS EPS-HEP-2009, 039 21 (2009), arXiv:0909.2354 [nucl-th]. 22
- G. S. Denicol, T. Kodama, and T. Koide, 23 arXiv:1002.2394 [nucl-th]. 24
- [16] R. A. Lacey et al., (2010), arXiv:1005.4979 [nucl-ex]. 25
- [17] R. A. Lacey, Nucl. Phys. A698, 559 (2002). 26
- [18] R. J. M. Snellings, Nucl. Phys. A698, 193 (2002). 27
- [19] J.-Y. Ollitrault, Phys. Rev. D46, 229 (1992). 28
- [20] K. Adcox et al., Phys. Rev. Lett. 89, 212301 (2002).
- [21] A. Mocsy and P. Sorensen, (2010), arXiv:1008.3381 [hep-30 31
- U. Heinz and P. Kolb, Nucl. Phys. A702, 269 (2002). 32
- [23] D. Teaney, Phys. Rev. C68, 034913 (2003). 33
- [24] P. Huovinen, P. F. Kolb, U. W. Heinz, P. V. Ruuskanen, 34 and S. A. Voloshin, Phys. Lett. **B503**, 58 (2001). 35
- T. Hirano and K. Tsuda, Phys. Rev. C66, 054905 (2002), 36 arXiv:nucl-th/0205043. 37
- R. Andrade et al., Eur. Phys. J. **A29**, 23 (2006), 38 arXiv:nucl-th/0511021. 39
- C. Nonaka, N. Sasaki, S. Muroya, and O. Miyamura, 101 40 Nucl. Phys. **A661**, 353 (1999), arXiv:nucl-th/9907046. 41
- H. Niemi, K. J. Eskola, and P. V. Ruuskanen, Phys. Rev. 103 42 C79, 024903 (2009), arXiv:0806.1116 [hep-ph]. 43
- R. Peschanski and E. N. Saridakis, Phys. Rev. C80, 44 024907 (2009), arXiv:0906.0941 [nucl-th]. 45
- H. Holopainen, H. Niemi, and K. J. Eskola, (2010),46 arXiv:1007.0368 [hep-ph]. 47
- B. Schenke, S. Jeon, and C. Gale, (2010),48 arXiv:1009.3244 [hep-ph]. 49
- J. Adams et al., Phys. Rev. Lett. 92, 062301 (2004). 50
- [33] S. S. Adler et al., Phys. Rev. Lett. 91, 182301 (2003). 51
- [34] B. Alver et al., Phys. Rev. Lett. 98, 242302 (2007). 52
- S. Afanasiev et al. (PHENIX), Phys. Rev. Lett. 99, 53 052301 (2007). 54
- B. I. Abelev et al. (STAR), Phys. Rev. C77, 054901 55 (2008), arXiv:0801.3466 [nucl-ex]. 56
- S. Afanasiev et al. (PHENIX), Phys. Rev. C80, 024909 57
- and A. Adare (The PHENIX), (2010), arXiv:1003.5586 59 nucl-ex.

- 61 [39] U. W. Heinz and S. M. H. Wong, Phys. Rev. C66, 014907
- 63 [40] R. A. Lacey and A. Taranenko, PoS CFRNC2006, 021 (2006).
- 65 [41] H. Song and U. W. Heinz, Phys. Rev. C77, 064901 (2008).66
- 67 [42] D. Molnar and M. Gyulassy, Nucl. Phys. A697, 495 (2002), arXiv:nucl-th/0104073.
- [43] R. A. Lacey, A. Taranenko, and R. Wei, (2009).arXiv:0905.4368 [nucl-ex].
- H. Masui, J.-Y. Ollitrault, R. Snellings, and A. Tang, 71 Nucl. Phys. A830, 463c (2009), arXiv:0908.0403 [nuclex]. 73
- [45] M. Miller and R. Snellings, (2003), arXiv:nucl-74 ex/0312008.
- [46] Y. Hama et al., Phys. Atom. Nucl. 71, 1558 (2008). 76
- W. Broniowski, P. Bozek, and M. Rybczynski, Phys. 77 Rev. C76, 054905 (2007).
- ⁷⁹ [48] T. Hirano and Y. Nara, Phys. Rev. **C79**, 064904 (2009).
- [49]Gombeaud and J.-Y. Ollitrault, (2009),80 arXiv:0907.4664 [nucl-th].
- R. A. Lacey et al., Phys. Rev. C81, 061901 (2010), 82 arXiv:1002.0649 [nucl-ex].
- P. Staig and E. Shuryak, (2010), arXiv:1008.3139 [nucl-84 th].
- G.-Y. Qin, H. Petersen, S. A. Bass, 86 (2010), arXiv:1009.1847 [nucl-th]. 87
- 88 M. L. Miller, K. Reygers, S. J. Sanders, and P. Steinberg. Ann. Rev. Nucl. Part. Sci. 57, 205 (2007).
- T. Lappi and R. Venugopalan, Phys. Rev. C74, 054905 90 (2006).
- [55] H.-J. Drescher and Y. Nara, Phys. Rev. C76, 041903 92 (2007).
- T. Hirano, U. W. Heinz, D. Kharzeev, R. Lacey, and Y. Nara, Phys. Lett. **B636**, 299 (2006).
- H.-J. Drescher, A. Dumitru, A. Hayashigaki, 96 and Y. Nara, Phys. Rev. C74, 044905 (2006). 97
- B. Alver and G. Roland, Phys. Rev. C81, 054905 (2010), 98 arXiv:1003.0194 [nucl-th]. 99
- B. H. Alver, C. Gombeaud, M. Luzum, and J.-Y. Olli-100 trault, (2010), arXiv:1007.5469 [nucl-th].
- H. Petersen, G.-Y. Qin, S. A. Bass, and B. Muller, 102 (2010), arXiv:1008.0625 [nucl-th].
- B. B. Back et al. (PHOBOS), Phys. Rev. C70, 021902 104 (2004).105
- [62] Note that the event planes for the eccentricities are spec-106 ified by the initial state coordinate asymmetry, whereas 107 the experimental event planes for v_n^* are specified by the 108 109 final state momentum space anisotropy. The two planes are correlated in ideal hydrodynamics. 110
- J. Adams et al. (STAR), Phys. Rev. C72, 014904 (2005). 111
- P. Kovtun, D. T. Son, and A. O. Starinets, Phys. Rev. 112 Lett. **94**, 111601 (2005), hep-th/0405231. 113
- 114 Roy A. Lacey and A. Taranenko, in: Proc. of the Winter Workshop on nuclear Dynamics, 2011.
- N. Borghini and J.-Y. Ollitrault, Phys. Lett. B642, 227 (2006).
- M. Csanad, T. Csorgo, and B. Lorstad, Nucl. Phys. 118 A742, 80 (2004), arXiv:nucl-th/0310040.