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First spectroscopic study of odd-odd ⁷⁸Cu

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Nuclei in the vicinity of 78 Ni are important benchmarks for nuclear structure, which can reveal changes in the shell structure far from stability. Spectroscopy of the odd-odd isotope 78 Cu was performed for the first time in an experiment with the EURICA setup at the Radioactive Isotope Beam Factory at RIKEN Nishina Center. Excited states in the neutron-rich isotope were populated following the β decay of 78 Ni, produced by in-flight fission and separated by the BigRIPS separator. A level scheme based on the analysis of γ - γ coincidences is presented. Tentative spin and parity

assignments were made when possible based on the β -decay feeding intensities and γ -decay properties of the excited states. Time correlations between β and γ decay show clear indications of an isomeric state with a half-life of 3.8(4) ms. Large-scale Monte Carlo shell-model calculations were performed using the A3DA-m interaction and a valence space comprising the full fp shell and the $1g_{9/2}$ and $2d_{5/2}$ orbitals for both protons and neutrons. The comparison of the experimental results with the shell-model calculations allows interpreting the excited states in terms of spin multiplets arising from the proton-neutron interaction. The results provide further insight into the evolution of the proton single-particle orbitals as a function of neutron number, and quantitative information about the proton-neutron interaction outside the doubly-magic 78 Ni core.

I. INTRODUCTION

One of the fundamental questions in nuclear physics is to understand how nuclear structure changes when moving away from well-known stable nuclei towards exotic nuclei with large proton-neutron asymmetry. Doubly magic nuclei and their neighbors play a crucial role for understanding the mechanisms that affect the energies and ordering of single-particle orbitals and the size of shell gaps [1, 2]. The nucleus ⁷⁸Ni is of particular interest for studies of shell evolution. With 28 protons and 50 neutrons, it has the largest neutron-to-proton ratio of all closed-shell nuclei with traditional magic numbers. The doubly magic character of ⁷⁸Ni was recently confirmed in a spectroscopic study that identified the first excited 2⁺ state at a high excitation energy of 2.6 MeV [3]. Robust shell closures for both Z=28 and N=50 are consistent with β -decay half-lives of nuclei in the region [4] and with the masses of neutron-rich Cu isotopes [5]. Nuclei with few valence particles or holes outside a ⁷⁸Ni core represent therefore important benchmarks for theoretical models. The nucleus 78 Cu with Z=29 and N=49is ideally suited to obtain information about the protonneutron interaction outside the ⁷⁸Ni core. Properties of nuclei in the vicinity of ⁷⁸Ni, in particular masses and β -decay half-lives, but also the occurrence of isomeric states, are furthermore important for modelling the nucleosynthesis in the region of the first r-process abundance peak [6, 7].

Earlier experiments in the 78 Ni region have seen evidence for an inversion of the proton $\pi 2p_{3/2}$ and $\pi 1f_{5/2}$ orbitals [8, 9]. The ordering of the two states becomes inverted in 75 Cu, where the ground state was found to have $I^{\pi} = 5/2^-$ [8], and an isomeric $3/2^-$ state was found at very low excitation energy [10]. The excitation energy of the $3/2^-$ state continues to increase relative to the $5/2^-$ ground state in 77 Cu [9] and 79 Cu [11], consistent with the crossing of the $\pi 1f_{5/2}$ and $\pi 2p_{3/2}$ orbitals. The change in single-particle energies was explained by the monopole component of the tensor interaction, which is attractive between the $\nu 1g_{9/2}$ and $\pi 1f_{5/2}$ orbitals, but repulsive between the $\nu 1g_{9/2}$ and $\pi 2p_{3/2}$ orbitals [12]. In 78 Cu, with only one proton and one neutron hole outside the doubly magic core, relatively pure configurations are

expected. The ground state and excited states at low excitation energy are expected to be dominated by the negative-parity multiplet arising from the coupling of an odd proton in the $\pi 1 f_{5/2}$ orbital with a neutron hole in the $\nu 1 g_{9/2}$ orbital.

The easiest way to couple the valence particles to positive-parity states is by neutron excitation from the $\nu 2p_{1/2}$ into the $\nu 1g_{9/2}$ orbital, leaving an unpaired neutron in the $\nu 2p_{1/2}$ orbital. Positive-parity states are expected to be found at higher excitation energy, and those with low spin are expected to be strongly fed in β decay by allowed transitions.

Shell-model calculations for the heavy odd-odd Cu isotopes were performed earlier by Van Roosbroeck et~al. using schematic δ and quadrupole-quadrupole (QQ) interactions for single proton and neutron shells outside a 68 Ni core, as well as using a larger valence space comprising the pf and $1g_{9/2}$ orbitals for both protons and neutrons with a more realistic interaction [13]. The results of the calculations reflected the transition from particle-particle to particle-hole coupling as neutrons fill the $\nu 1g_{9/2}$ orbital, consistent with expectations from the parabolic rule [14].

Monte Carlo shell-model (MCSM) calculations [15] based on a larger valence space outside a 40 Ca core with the A3DA-m interaction [16] were able to reproduce detailed spectroscopic data for both 77 Cu [9] and 79 Cu [11]. Extending the experimental spectroscopic information to heavier odd-odd Cu isotopes is crucial for understanding the interaction between proton particles and neutron holes outside the 78 Ni core, and to provide additional benchmarks for the MCSM calculations. It was furthermore shown that residual proton-neutron interactions between the pf and sdg shells have implications for calculating electron capture rates during core collapse supernovae [17].

Before the present experiment, no excited states in 78 Cu were known. Its half-life has previously been measured to be 330.7 ± 2.0 ms [4]. Magnetic dipole and electric quadrupole moments have been measured for the ground states up to A=78 [18], which found a tentative assignment of (6^-) for 78 Cu. The present work provides the first spectroscopic information on the odd-odd isotope 78 Cu. Experimental details and the data analysis are described in sections II and III, respectively. Results including spectra, level schemes, and spin-parity assignments, are presented in section IV. The results are discussed and compared to MCSM calculations in section

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V, followed by a summary and conclusions in section VI.

II. EXPERIMENTAL SETUP

The data presented in this article were obtained in experiments carried out at the Radioactive Isotope Beam Factory (RIBF) of the RIKEN Nishina Centre for Accelerator-based Science outside Tokyo, Japan, during two separate beam times as part of the EURICA campaign [19]. A primary beam of ²³⁸U was accelerated subsequently by four cyclotrons to an energy of 345 MeV per nucleon with an average intensity of 10 pnA. In-flight fission reactions of the incident ²³⁸U projectiles were induced on a ⁹Be target of 555 mg/cm² areal density, which was located at the F0 focal point at the entrance of the BigRIPS fragment separator [20]. The fission fragments were separated in the first stage of the BigRIPS separator by using the $B\rho - \Delta E - B\rho$ method [21]. Particle identification (PID) was performed in the second stage of the fragment separator by combining information on the time-of-flight through the separator with the magnetic rigidity $B\rho$ and the characteristic energy loss ΔE of the fragments. The ions of interest were further transmitted through the ZeroDegree spectrometer [21] to the focal point F11, where their β decay and subsequent γ ray emission were detected. A resulting PID plot can be found in Ref. [4]. The settings of the BigRIPS separator were optimized for the transmission of ⁷⁸Ni and ⁸¹Zn, respectively, during the two experiments.

The separated fission fragments were implanted into the Wide-range Active Silicon-Strip Stopper Array for Beta and ion detection (WAS3ABi) [22], which consisted of a stack of 8 double-sided silicon strip detectors (DSSSDs). Each detector had 60 horizontal and 40 vertical strips of 1 mm pitch, resulting in a total of 2400 pixels of size 1×1 mm² per detector. Each detector had a thickness of 1 mm, with 0.5 mm separation in depth between the detectors. To ensure that the ions were stopped in the center of WAS3ABi, a thin Al degrader was located in front of the detectors. The WAS3ABi array was surrounded by the EURICA array of 12 Euroball Cluster detectors. Each Cluster detector consisted of 7 HPGe detectors, yielding a total of 84 Ge crystals with an absolute photo-peak efficiency of $\sim 6.5\%$ for 1.3 MeV γ rays. Ion implantation, β decay, and γ decay events were recorded in time-stamped list mode, allowing the correlation of γ -decay events with the β decays of specific fission fragments that were identified in mass and atomic number. More details on the experimental setup can be found in Refs. [19, 22].

III. DATA ANALYSIS

As a first step, subsets of data were generated according to the atomic number Z and mass number A of the ions that were identified in BigRIPS and implanted into

WAS3ABi. Ion implantation events were correlated in time and position with subsequent β decays. A total of 7.2×10^3 ⁷⁸Ni ions were implanted, and 3.0×10^3 correlated β -decay events were detected. The β -decay half-life of ⁷⁸Ni was found to be $T_{1/2}=122.2(51)$ ms in a separate analysis of the same data [4], while the Q_{β} value is 9910 (400) keV [23]. Finally, γ rays detected in the Ge detectors were correlated in time with β -decay events in the Si detectors. The data from the two experiments were analyzed separately, and the resulting γ -ray spectra and γ - γ coincidence matrices were combined afterwards. The individual steps of the data analysis are described in more detail in the following.

Signals from heavy ion implantation in the DSSSDs are easily distinguished from the detection of β -decay electrons by their signal amplitude. For each implantation event of a $^{78}{\rm Ni}$ ion, the data were scanned for the subsequent β decays within a given time window of 2 s. If more than one β -decay event were registered within the correlation time window, only the first one was considered. To reduce the number of random coincidences between implantation and β -decay events, it was required that the implanted ion and β -decay electron were detected in the same, a neighboring, or next to neighboring pixel of the same DSSSD layer.

Finally, correlated events between implanted ions and β decays were used to select γ rays that were promptly following the β decay of ⁷⁸Ni, within a time window of approximately 200 ns. The information on the time difference between β decays and detected γ rays was furthermore used to search for isomeric decays. In the case that two neighboring crystals within the same Ge cluster detector gave coincident signals, their energies were summed to account for Compton scattering and to increase the detection efficiency for γ rays with high energy.

Gamma-ray singles spectra were sorted for different correlation time windows between the ion implantation and β -decay events. Because the detection efficiency for electrons in the DSSSD is less than 100 %, the β decay can remain undetected. In case a subsequent β decay (or β -delayed neutron decay) occurs within the correlation time window, γ rays from the decay daughter or even granddaughter can appear in the spectrum. Limiting the correlation time to short intervals of the order of the half-life of 78 Ni strongly suppresses γ rays originating from subsequent decays, but also removes γ rays occurring within 78 Cu. The relative suppression of γ rays as a function of correlation time was used to assign unknown γ rays to ⁷⁸Cu. Known γ rays following the decay of ⁷⁸Cu into ⁷⁸Zn, [13] were used to validate the procedure. After the assignment of the strongest γ rays to ⁷⁸Cu, the strict time constraint between implantation and β decay was relaxed to search for $\gamma-\gamma$ coincidences and to construct the level scheme in a compromise between high statistics for the γ rays of interest and suppressing γ rays from daughter decays.

Fig. 1 shows a γ -ray singles spectrum for ⁷⁸Cu. To

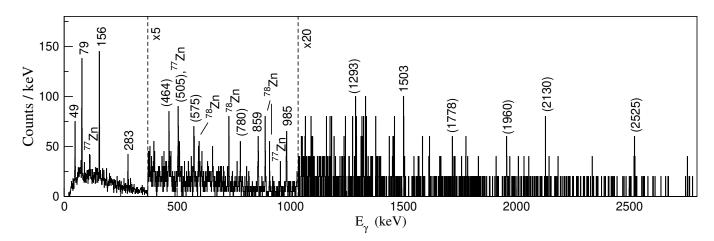


FIG. 1. Gamma-ray singles spectrum for 78 Cu. Note that the higher-energy regions of the spectrum were scaled by the indicated factors. An ion- β correlation window of 2 s was used. Peaks marked by their energy (in keV) are transitions assigned to 78 Cu. Energies are given in parentheses for transitions that were assigned to 78 Cu, but could not be placed in the level scheme. Transitions following the subsequent β and β -n decay of 78 Cu are labelled as 78 Zn and 77 Zn, respectively. The 505 keV transition appears in both 78 Cu and the β -n daughter 77 Zn.

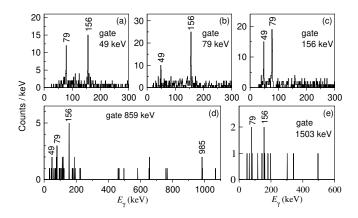


FIG. 2. γ - γ coincidence spectra for ⁷⁸Cu gated on the 49 keV (a), 79 keV (b), 156 keV (c), 859 keV (d), and 1503 keV transition (e).

maximize the level of statistics, a relatively long correlation time window of 2 s was used, resulting in stronger contributions from the daughter decays in $^{77,78}{\rm Zn}$. All peaks that are labeled by their energy were assigned to $^{78}{\rm Cu}$. Some transitions, however, could not be placed in the level scheme because of lacking $\gamma\text{-}\gamma$ coincidence relationships. These transitions, which are labeled by their energies in parentheses, were assigned to $^{78}{\rm Cu}$ based only on their time correlation with the $\beta\text{-decay}$ detection. Where possible, coincidence relationships between γ rays were used to validate their assignment to $^{78}{\rm Cu}$. Examples for gated coincidence spectra are shown in Fig. 2. Only γ rays that could be placed unambiguously were included in the level scheme.

The absolute intensity of β -decay feeding was determined from the intensity balance of γ rays feeding and depopulating a given state, which was corrected for de-

tection efficiency and internal conversion and normalized to the number of implanted ions. However, because of the incomplete level scheme and missing γ -ray feeding from above, this apparent β feeding is only a limit, and a conversion into $\log ft$ values is not meaningful. The probability for β -delayed neutron emission was measured to be $P_n = 25.8(38)\%$ [24]. The observed apparent β feeding accounts for less than 58% of the decays of implanted ions. Taking into account P_n , less than 78% of β -decay feeding was observed. The values for the total observed feeding represents only upper limits and a significant fraction of feeding strength could therefore remain unobserved.

IV. RESULTS

The decay scheme for $^{78}\mathrm{Cu}$ is shown in Fig. 3. The information presented in the decay scheme is furthermore summarized in Table I, together with information on γ -ray intensities and uncertainties for all quantities. It should be noted that no excited states were known prior to the present experiment. The analysis of time correlations allowed associating 16 γ -ray transitions with $^{78}\mathrm{Cu}$, as indicated in Fig. 1. Of these, seven could be placed in the decay scheme based on their coincidence relationships.

The ground-state spin of ⁷⁸Cu was previously assigned as $(4,5)^-$ [13] based on the feeding of states in the β -decay daughter ⁷⁸Zn, and, in later works, as (6^-) [25] and (5^-) [26]. A laser spectroscopy experiment showed best agreement with I=6, suggesting a ground-state spin-parity of (6^-) [18].

The three strongest transitions, with energies of 49, 79, and 156 keV are in mutual coincidence, as shown in Fig. 2. Because they are much stronger than any other

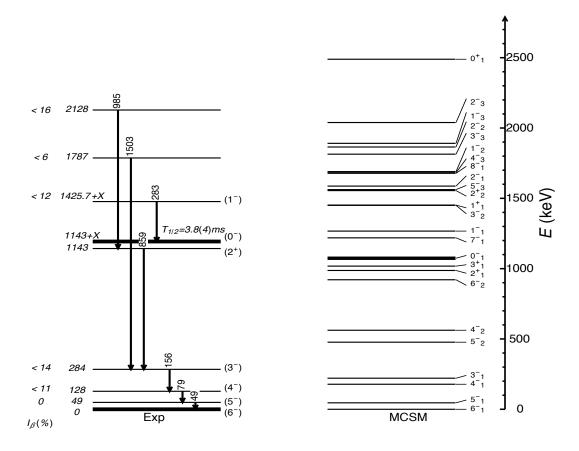


FIG. 3. Experimental level scheme for ⁷⁸Cu and comparison to MCSM calculations (see text for details).

transition, it is reasonable to assume that they form a cascade built on the ground state. It can be furthermore assumed that the low-energy transitions connect members of the negative-parity $\pi 1 f_{5/2} \times \nu 1 g_{9/2}^{-1}$ multiplet, because other configurations are only expected at higher excitation energy. The intensities of the three transitions are approximately the same when taking electron conversion into account and assuming M1 multipolarity. Other multipolarities for transitions between negativeparity states would result in significantly higher conversion coefficients. As an example, the conversion coefficient for a 49 keV transition of M1 multipolarity is 0.385, whereas it is 8.98 for E2 multipolarity [27]. Other multipolarities than M1 would therefore either require strong β feeding, which would be highly forbidden, or strong γ feeding that is only in coincidence with one or two transitions within the cascade, for which there is no evidence in the data. Higher multipolarity for transitions with such low energy would in addition result in relatively long lifetimes and delayed coincidences, whereas all three transitions are in prompt coincidence. It can therefore be concluded that the three strongest transitions form a cascade of M1 transitions feeding a (6^-) ground state from a (3⁻) state at 284 keV. However, the data does not provide any indication for the ordering of the three transitions. According to the empiric rule established by Paar [14], the energies in proton-neutron multiplets with particle-hole character are expected to follow a parabolic trend with I(I+1) and a minimum at spin $I=j_\pi+j_\nu-1$, which is equal to 6 in the case of the $\pi 1 f_{5/2} \times \nu 1 g_{9/2}^{-1}$ configuration. With the ordering of the three transitions within the cascade as shown in the level scheme of Fig. 3, the states fit the expected parabolic trend well, which seems to justify this choice.

The singles spectrum in Fig. 1 shows a transition of 283 keV, and one would be tempted to place this transition as the decay from the 284 keV state. However, the 283 keV transition is not in coincidence with any of the transitions feeding the 284 keV state, as can be seen in Fig. 2(d) and (e). Furthermore, the transition would have M3 multipolarity if it was depopulating the 3⁻ state, and would therefore unlikely be prompt. Any other ordering of spins for the low-energy states that would allow a prompt 283 keV transition would be in conflict with the observed intensities for the low-energy transitions. It is therefore concluded that the 283 keV transition is originating from a state at higher excitation energy.

The 859, 985, and 1503 keV transitions were clearly seen in the γ -ray singles spectrum (see Fig. 1). Based on their time correlations with β decay, they can be identified as belonging to 78 Cu. All three transitions are in coincidence with the cascade of three low-energy transi-

TABLE I. γ -ray energies and intensities, along with their initial state excitation energies, spin and parity, and β -feeding.

	E_{γ} [keV]	<i>I</i> [07.]	E_i	I^{π}	β -feeding [%]
	/ L J	$I\gamma$ [/0]			ρ-reeding [70]
	X		1143 + X	'	
	49(1)	$43(4)^{a}$	49(1)	(5^{-})	0
	79(1)	$44(1)^{a}$	128(2)	(4^{-})	< 11
	156.1(2)	$33(1)^{a}$	284(2)	(3^{-})	< 14
	282.6(7)	12(1)	1425.7 + X	(1^{-})	< 12
	464.5(5)	12(1)			
	505(2)	9(1)			
	575(3)	10(1)			
	779.9(6)	7(1)			
	859.0(6)	13(1)	1143(2)	(2^{+})	
	984.5(6)	16(1)	2128(2)		< 16
	1293(3)	2(2)			
	1503(4)	6(2)	1787(4)		< 6
	1778(4)	3(2)			
	1960(2)	2(1)			
	2130(3)	2(1)			
	2525(3)	2(1)			
- 3					

^a Corrected for internal conversion assuming pure M1 character.

tions, as can be seen in Fig. 2. In addition, the 859 and 985 keV transitions are in mutual coincidence. Consequently, the 859 and 1503 keV transitions are placed on top of the (3⁻) state at 284 keV excitation energy, and the 985 keV transition on top of the 859 keV transition, feeding a state at 1143 keV excitation energy. With only few transitions placed in the level scheme, the observed β feeding is incomplete and cannot be used for spin assignments. Because the higher-lying states are less likely to be affected by unobserved γ feeding, they are likely to have low spin. The proposed decay scheme is therefore consistent with bridging the large spin gap between states that are fed by allowed β transitions and a (6⁻) ground state.

The analysis of time correlations in the decay of ⁷⁸Cu revealed clear evidence for an isomeric state, as is illustrated in Fig. 4. The spectrum in Fig. 4(a) shows γ rays following the implantation of ions identified as 78 Cu $(T_{1/2} = 330.7(20) \text{ ms})$ within 200 ms. As expected, the spectrum shows the known transitions in ⁷⁸Zn and ⁷⁷Zn [26]. The spectrum also shows hints of the 156 and 859 keV transitions originating from excited states in ⁷⁸Cu. When selecting a short correlation time of 5 ms, as shown in Fig. 4(b), the cascade of low-energy transitions and the 859 keV transition of ⁷⁸Cu appears in the spectrum. The spectrum of Fig. 4(b) was further cleaned by selecting events where a low-energy signal was detected in the same DSSSD pixel as the ion implantation, reducing the efficiency for β -decay events and enhancing events originating from conversion electrons. The delayed γ ray spectrum shows the 859 keV transition, but not the one at 985 keV. This confirms the ordering of the cascade, and clearly shows that the isomeric state is located above the state at 1143 keV. The absence of the 283 keV transition in the delayed spectrum furthermore confirms that it does not originate from the (3^-) state. Because the 283 keV transition is relatively strong, but not seen in coincidence with any other transition, it seems likely that it is feeding the isomeric state.

The time spectrum of the isomeric decay is shown in Fig. 4(c), from which a half-life of $T_{1/2}=3.8(4)$ ms can be extracted. The spectrum shows the time difference between the implantation of a ⁷⁸Cu ion and the detection of a low-energy signal in the same pixel of the DSSSD, with an additional condition that one of the four γ -ray transitions following the decay of the isomer was detected. The fact that the decay from the isomer to the state at 1143 keV remained unobserved could be explained by a small energy difference and consequently a large conversion coefficient, consistent with the conversion-electron signal observed in the DSSSD.

Any spin assignment with I > 1 for the isomer would be incompatible with the observed half-life of 3.8(4) ms. For a 1⁺ state, for example, an M2 transition of at least 859 keV to the (3⁻) state would likely result in a half-life of nanoseconds rather than milliseconds. A 1⁻ assignment or any higher spin would result in prompt decay to one of the low-lying states. Also a 0⁺ assignment seems highly unlikely for the isomeric state. An E3 transition of at least 859 keV and a strength of 1 Weisskopf unit would result in a half-life of less than 10 μ s, 500 times shorter than the observed value. The most likely assignment for the isomeric state is therefore 0⁻. A possible M3 decay to the state at 284 keV would be sufficiently hindered for it to be unobserved. Instead, a low-energy decay to the state at 1143 keV would become competitive, consistent with the conversion electron signal in the DSSSD. The state at 1143 keV would in this case most likely have spin-parity $I^{\pi} = 2^{+}$. The half-life of 3.8(4) ms can be explained by an energy difference of $\sim 50 \text{ keV}$ between the two states and an M2 transition of 1 Weisskopf unit. Such a transition would have a conversion coefficient of ~ 6 , consistent with the DSSSD signal and the non-observation of a γ ray. Although it is not possible to determine the precise excitation energy of the isomeric state in this way, a 0⁻ assignment and a low-energy M2 transition of a few tens of keV to a 2⁺ state at 1143 keV is the only scenario that can explain all observations.

The data is insufficient to determine whether the isomeric state is directly fed by β decay. It is likely that some of the prompt γ -ray transitions that were observed in the singles spectrum for ⁷⁸Cu feed the isomer. The fact that prompt coincidence relationships are lacking for the relatively strong 283 keV transition suggests that this transition is feeding the isomeric state directly. The resulting state has an excitation energy of (1427+X) keV, with X being the energy difference between the isomer and the state at 1143 keV.

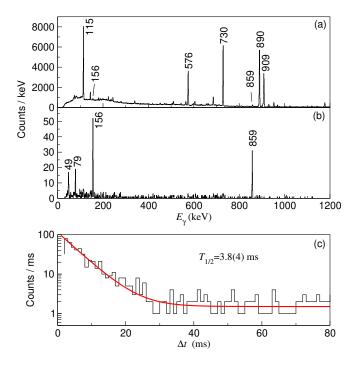


FIG. 4. (a) Decay spectrum of ⁷⁸Cu ions within 200 ms from implantation in the DSSSDs, revealing known transitions in ⁷⁸Zn and the 115 keV transition in ⁷⁷Zn. The small peaks at 156 keV and 859 keV, originating from excited states in ⁷⁸Cu, indicate the presence of an isomeric state. (b) Decay spectrum of ⁷⁸Cu within 5 ms, with an additional condition that a low-energy signal was detected in the same DSSSD pixel as the ion implantation. (c) Time difference between implantation of ⁷⁸Cu ions and detection of low-energy signals in the DSSSD, with an additional condition that one of the four γ -ray transitions following the decay of the isomer (as seen in spectrum b) was detected in EURICA. The exponential fit yields a half-life of $T_{1/2}=3.8(4)$ ms.

V. DISCUSSION

In order to understand the nature of the experimentally observed states, shell-model calculations are necessary. In the present work, we compare the experimental results with Monte Carlo shell-model (MCSM) calculations [15] using the A3DA-m effective interaction, which has successfully described the structure of nuclei in the 78 Ni region [9, 11, 16, 28]. The valence space comprised the full pf shell together with the $1g_{9/2}$ and $2d_{5/2}$ orbitals for both protons and neutrons without restrictions. Calculated states up to an excitation energy of 2.5 MeV, for both negative and positive parity, are compared to the experimental level scheme in Fig. 3. Table II shows the occupation numbers for protons and neutrons found in the MCSM calculations for the ten lowest states.

The calculations reproduce the sequence of negativeparity states that was established by the cascade of M1 transitions very well, including $I^{\pi}=6^{-}$ for the ground state. The occupation numbers illustrate that these states have a relatively pure particle-hole character based on the $\pi 1 f_{5/2} \times \nu 1 g_{9/2}^{-1}$ configuration. This is consistent with the inversion of the $\pi 1 f_{5/2}$ and $\pi 2 p_{3/2}$ orbitals near N=48, which was observed previously [9]. The calculations predict the remaining 2_1^- and 7_1^- members of the multiplet to be at much higher excitation energy, with the negative-parity states based on the $\pi 2 p_{3/2} \times \nu 1 g_{9/2}^{-1}$ configuration in between. The latter comprises the 5_2^- , 4_2^- , 6_2^- , and 3_2^- states.

The MCSM calculations predict a 0⁻ state at 1074 keV, close to the excitation energy of the observed isomeric state, which lends further support to the (0^{-}) assignment for the isomer. The calculations find a relatively pure $\pi 1 f_{5/2} \nu 2 d_{5/2}$ configuration for the 0⁻ state. The excitation energy of the isomeric state contains therefore not only information on the interaction energy between the $\pi 1 f_{5/2}$ and $\nu 2 d_{5/2}$ orbitals, but also on the size of the N=50 shell gap. The good agreement between the experimental and theoretical excitation energy indicates that both quantities are well described by the A3DA-m interaction. The $\pi 1 f_{5/2} \nu 2 d_{5/2}$ configuration gives rise to a multiplet of negative-parity states comprising the $0_1^-, 1_1^-, 5_3^-, 4_3^-, 3_3^-,$ and 2_2^- states. The experimental state at 1427 + X keV excitation energy is a potential candidate for the 1_1^- state, as strong M1 transitions are expected between the states of the multiplet, although such an assignment remains speculative.

All calculated low-lying positive-parity states are based on a neutron excitation from the $\nu 2p_{1/2}$ to the $\nu 1g_{9/2}$ orbital. The coupling between an odd neutron in the $\nu 2p_{1/2}$ orbital and an odd proton in the $\pi 1f_{5/2}$ orbital results in a doublet of the 2⁺ and 3⁺ states. The occupation numbers of Table II show that the 2_1^+ and 3_1^+ states are indeed dominated by the $\pi 1 f_{5/2} \nu 2 p_{1/2}$ configuration. The experimental state at 1143 keV, which is tentatively assigned as (2^+) , agrees reasonably well with the calculated 2_1^+ state at 987 keV, and is consequently a candidate for a member of the $\pi 1 f_{5/2} \nu 2 p_{1/2}$ doublet. The coupling of a $\nu 2p_{1/2}$ neutron with a $\pi 2p_{3/2}$ proton results in a doublet of the 1_1^+ and 2_2^+ states, which are calculated to be at higher energy around 1.5 MeV. The calculated 0_1^+ state at approximately 2.5 MeV, finally, is found to be based on the $\pi 2p_{1/2}\nu 2p_{1/2}$ configuration. It would be speculative to associate any of the higher-lying experimental states with any of the calculated states.

Figure 5(a) shows the excitation energies of the negative-parity states from the MCSM calculations as a function of the squared angular momentum I(I+1). The calculated occupation numbers were used to assign the states to the multiplets with predominant $\pi 1 f_{5/2} \nu 1 g_{9/2}^{-1}$, $\pi 2 p_{3/2} \nu 1 g_{9/2}^{-1}$, and $\pi 1 f_{5/2} \nu 2 d_{5/2}$ configuration. Experimental excitation energies of states with a tentative spin assignment are also included in Fig. 5(a), which illustrates again the rather good agreement between the MCSM calculations and experiment. The multiplets involving a hole in the $\nu 1 g_{9/2}$ orbital show a parabolic de-

E (MeV)	J^{π}	$\pi 1 f_{7/2}$	$\pi 2p_{3/2}$	$\pi 1 f_{5/2}$	$\pi 2p_{1/2}$	$\pi 1g_{9/2}$	$\pi 1d_{5/2}$	$\nu 1 f_{7/2}$	$\nu 2p_{3/2}$	$\nu 1 f_{5/2}$	$\nu 2p_{1/2}$	$\nu 1g_{9/2}$	$\nu 1d_{5/2}$
0.000	6-	7.73	0.25	0.96	0.02	0.04	0.01	7.99	3.99	5.99	1.99	8.86	0.19
0.046	5^{-}	7.71	0.33	0.87	0.03	0.04	0.01	7.99	3.99	5.99	1.99	8.85	0.20
0.178	4^{-}	7.71	0.35	0.84	0.05	0.04	0.01	7.99	3.99	5.99	1.98	8.83	0.21
0.221	3^{-}	7.74	0.46	0.73	0.02	0.04	0.01	7.99	3.99	5.99	1.99	8.86	0.18
0.477	5^{-}	7.70	0.87	0.30	0.07	0.04	0.01	7.99	3.99	5.99	1.98	8.84	0.21
0.562	4^{-}	7.74	0.84	0.34	0.03	0.04	0.01	7.99	3.99	5.99	1.98	8.85	0.19
0.921	6^{-}	7.71	0.95	0.25	0.03	0.05	0.01	7.99	3.98	5.99	1.98	8.88	0.18
0.987	2^{+}	7.63	0.26	1.04	0.02	0.04	0.01	7.99	3.92	5.92	1.15	9.78	0.24
1.018	3^{+}	7.60	0.26	1.07	0.02	0.04	0.01	7.99	3.95	5.88	1.15	9.76	0.26
1.074	0-	7.58	0.29	1.03	0.05	0.04	0.01	7.98	3.92	5.92	1.89	8.13	1.15

TABLE II. Occupation of proton and neutron orbitals in $fpg_{9/2}d_{5/2}$ spaces

pendence on angular momentum, with the extreme couplings $I=j_{\nu}\pm j_{\pi}$ having the highest energy, as expected for particle-hole coupling [14]. The $\pi 1 f_{5/2} \nu 2 d_{5/2}$ multiplet, on the other hand, has particle-particle character, which favors (anti-)parallel coupling.

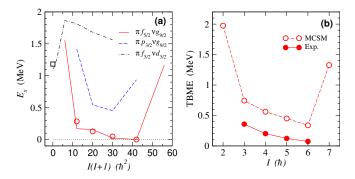


FIG. 5. (a) Excitation energies of the negative-parity multiplets with predominant $\pi 1 f_{5/2} \nu 1 g_{9/2}^{-1}$, $\pi 2 p_{3/2} \nu 1 g_{9/2}^{-1}$, and $\pi 1 f_{5/2} \nu 2 d_{5/2}$ configurations from the MCSM calculations as a function of I(I+1). The symbols indicate the experimental excitation energies for those states for which a tentative spin assignment and association with the multiplet was possible. (b) Two-body matrix elements (TBME) extracted for the $\pi f_{5/2} \nu g_{9/2}^{-1}$ spin multiplet from both the experimental and calculated excited sates (see text for details).

The observed excited states in the odd-odd Cu nuclei allow determining the proton-neutron monopole interaction, which is responsible for changes in single-particle energies, *i.e.* for the shell evolution far from stability. A description of the procedure to extract two-body matrix elements (TBME) from experimental excitation energies of proton-neutron multiplets can be found in Ref. [1]. In the following we apply this procedure for the case of ⁷⁸Cu. As can be seen in Table II, the negative-parity states of the $\pi 1f_{5/2}\nu 1g_{9/2}^{-1}$ multiplet have relatively pure single-particle (hole) configurations. The results are therefore well suited to extract experimental TBME for the interaction between a $1f_{5/2}$ proton with a $1g_{9/2}^{-1}$ neutron hole. Starting from a ⁷⁸Ni core, the contributions of a non-interacting proton particle and neutron hole are obtained from the binding energies of ⁷⁹Cu [5] and ⁷⁷Ni [23] for the

 $1f_{5/2}$ proton and the $1g_{9/2}$ neutron hole, respectively. The resulting value is found to be 74 keV lower than the experimental binding energy for the ground state of ⁷⁸Cu [29], which includes the repulsive residual interaction between the $1f_{5/2}$ proton and the $1g_{9/2}$ neutron hole when coupled to spin 6^- . The TBME for the various spin couplings are consequently shifted by 74 keV compared to the excitation energies of the corresponding states. It should be noted that the masses of ⁷⁷Ni and ⁷⁸Ni are not known experimentally, and that the extrapolated values have an uncertainty of 400 keV [23]. The absolute values of the experimental TBME depend therefore strongly on the extrapolated masses of the Ni isotopes. The experimental values are compared to TBME of the A3DA-m interaction in Fig. 5(b), where the theoretical TBME for the $\pi f_{5/2} \nu g_{9/2}^{-1}$ particle-hole interaction were obtained by applying the Pandya transformation [30] to the corresponding particle-particle TBME. The comparison shows that the A3DA-m interaction describes the relative size of the TBME well. The experimental results on the excited states in ⁷⁸Cu can be used to refine the shell model interaction once more precise mass values for ⁷⁷Ni and ⁷⁸Ni become available.

VI. SUMMARY AND CONCLUSIONS

Excited states in $^{78}\mathrm{Cu}$ have been observed for the first time following the β decay of ⁷⁸Ni. The neutron-rich ⁷⁸Ni isotopes were produced at the Radioactive Isotope Beam Factory at RIKEN Nishina Center, Japan, by in-flight fission induced by a primary beam of ²³⁸U at 345 MeV per nucleon incident on a ⁹Be target. The secondary beams were separated by the BigRIPS separator and transported to the decay station, where they were implanted into the WAS3ABi detector. The HPGe detectors of the EURICA array were used to detect γ rays following the β decay of ⁷⁸Ni. An isomeric state with a half-life of 3.8(4) ms was discovered in ⁷⁸Cu and tentatively assigned as (0^{-}) . The combination of information from $\gamma\gamma$ -coincidence data and the decay of the isomeric state allowed building a partial level scheme for ⁷⁸Cu. Spins and parities could be tentatively assigned for some

of the states.

The experimental results were compared to large-scale MCSM calculations using the A3DA-m interaction and a valence space comprising the $1f_{7/2}$, $2p_{3/2}$, $1f_{5/2}$, $2p_{1/2}$, $1g_{9/2}$, and $2d_{5/2}$ orbitals for both protons and neutrons. The shell-model calculations show a remarkable agreement with the experimental results. Combining the experimental results with the calculations, it was possible to interpret the low-lying states in terms of spin multiplets arising from the coupling of an odd proton in either the $\pi 1 f_{5/2}$ or $\pi 2 p_{3/2}$ orbital with an odd neutron in the $\nu 1g_{9/2}, \ \nu 2p_{1/2}, \ {\rm or} \ \nu 2d_{5/2}$ orbital. The results confirm the previously observed crossing between the $\pi 2p_{3/2}$ and $\pi 1 f_{5/2}$ orbitals. The interpretation of the isomeric state as based on the $\pi 1 f_{5/2} \nu 2 d_{5/2}$ configuration provides information on the N=50 shell gap. Because configurations are pure, it was possible to extract experimental two-body matrix elements for the $\pi 1 f_{5/2} - \nu 1 g_{9/2}^{-1}$ interaction, which represent important input for future shellmodel calculations in the ⁷⁸Ni region. Extending the work to ⁸⁰Cu would represent an important step for investigating the proton-neutron interaction beyond $^{78}\mathrm{Ni}.$

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^[1] O. Sorlinand M.-G. Porquet, Progress in Particle and Nuclear Physics **61**, 602 (2008).

^[2] T. Otsuka, A. Gade, O. Sorlin, T. Suzuki, and Y. Utsuno, Rev. Mod. Phys. 92, 015002 (2020).

^[3] R. Taniuchi, C. Santamaria, P. Doornenbal, A. Obertelli, K. Yoneda, G. Authelet, H. Baba, D. Calvet, F. Château, A. Corsi, A. Delbart, J.-M. Gheller, A. Gillibert, J. D. Holt, T. Isobe, V. Lapoux, M. Matsushita, J. Menéndez, S. Momiyama, T. Motobayashi, M. Niikura, F. Nowacki, K. Ogata, H. Otsu, T. Otsuka, C. Péron, S. Péru, A. Peyaud, E. C. Pollacco, A. Poves, J.-Y. Roussé, H. Sakurai, A. Schwenk, Y. Shiga, J. Simonis, S. R. Stroberg, S. Takeuchi, Y. Tsunoda, T. Uesaka, H. Wang, F. Browne, L. X. Chung, Z. Dombradi, S. Franchoo, F. Giacoppo, A. Gottardo, K. Hadynska-Klek, Z. Korkulu, S. Koyama, Y. Kubota, J. Lee, M. Lettmann, C. Louchart, R. Lozeva, K. Matsui, T. Miyazaki, S. Nishimura, L. Olivier, S. Ota, Z. Patel, E. Sahin, C. Shand, P.-A. Söderström, I. Stefan, D. Steppenbeck, T. Sumikama, D. Suzuki, Z. Vajta, V. Werner, J. Wu, and Z. Y. Xu, Nature **569**, 53 (2019).

^[4] Z. Y. Xu, S. Nishimura, G. Lorusso, F. Browne, P. Doornenbal, G. Gey, H.-S. Jung, Z. Li, M. Niikura, P.-A. Söderström, T. Sumikama, J. Taprogge, Z. Vajta,

H. Watanabe, J. Wu, A. Yagi, K. Yoshinaga, H. Baba, S. Franchoo, T. Isobe, P. R. John, I. Kojouharov, S. Kubono, N. Kurz, I. Matea, K. Matsui, D. Mengoni, P. Morfouace, D. R. Napoli, F. Naqvi, H. Nishibata, A. Odahara, E. Sahin, H. Sakurai, H. Schaffner, I. G. Stefan, D. Suzuki, R. Taniuchi, and V. Werner, Phys. Rev. Lett. 113, 032505 (2014).

^[5] A. Welker, N. A. S. Althubiti, D. Atanasov, K. Blaum, T. E. Cocolios, F. Herfurth, S. Kreim, D. Lunney, V. Manea, M. Mougeot, D. Neidherr, F. Nowacki, A. Poves, M. Rosenbusch, L. Schweikhard, F. Wienholtz, R. N. Wolf, and K. Zuber, Phys. Rev. Lett. 119, 192502 (2017).

^[6] K. L. Kratz, H. Gabelmann, P. Mller, B. Pfeiffer, H. L. Ravn, A. Whr, and T. I. Collaboration, Zeitschrift fr Physik A Hadrons and Nuclei 10.1007/BF01290331 (1991).

^[7] S. Nikas, G. M. Pinedo, and re Sieverding, Journal of Physics: Conference Series 1668, 012029 (2020).

^[8] K. T. Flanagan, P. Vingerhoets, M. Avgoulea, J. Billowes, M. L. Bissell, K. Blaum, B. Cheal, M. De Rydt, V. N. Fedosseev, D. H. Forest, C. Geppert, U. Köster, M. Kowalska, J. Krämer, K. L. Kratz, A. Krieger, E. Mané, B. A. Marsh, T. Materna, L. Mathieu, P. L.

- Molkanov, R. Neugart, G. Neyens, W. Nörtershäuser, M. D. Seliverstov, O. Serot, M. Schug, M. A. Sjoedin, J. R. Stone, N. J. Stone, H. H. Stroke, G. Tungate, D. T. Yordanov, and Y. M. Volkov, Phys. Rev. Lett. **103**, 142501 (2009).
- [9] E. Sahin, F. L. Bello Garrote, Y. Tsunoda, T. Otsuka, G. de Angelis, A. Görgen, M. Niikura, S. Nishimura, Z. Y. Xu, H. Baba, F. Browne, M.-C. Delattre, P. Doornenbal, S. Franchoo, G. Gey, K. Hadyńska-Klęk, T. Isobe, P. R. John, H. S. Jung, I. Kojouharov, T. Kubo, N. Kurz, Z. Li, G. Lorusso, I. Matea, K. Matsui, D. Mengoni, P. Morfouace, D. R. Napoli, F. Naqvi, H. Nishibata, A. Odahara, H. Sakurai, H. Schaffner, P.-A. Söderström, D. Sohler, I. G. Stefan, T. Sumikama, D. Suzuki, R. Taniuchi, J. Taprogge, Z. Vajta, H. Watanabe, V. Werner, J. Wu, A. Yagi, M. Yalcinkaya, and K. Yoshinaga, Phys. Rev. Lett. 118, 242502 (2017).
- [10] C. Petrone, J. M. Daugas, G. S. Simpson, M. Stanoiu, C. Plaisir, T. Faul, C. Borcea, R. Borcea, L. Cáceres, S. Calinescu, R. Chevrier, L. Gaudefroy, G. Georgiev, G. Gey, O. Kamalou, F. Negoita, F. Rotaru, O. Sorlin, and J. C. Thomas, Phys. Rev. C 94, 024319 (2016).
- [11] L. Olivier, S. Franchoo, M. Niikura, Z. Vajta, D. Sohler, P. Doornenbal, A. Obertelli, Y. Tsunoda, T. Otsuka, G. Authelet, H. Baba, D. Calvet, F. Château, A. Corsi, A. Delbart, J.-M. Gheller, A. Gillibert, T. Isobe, V. Lapoux, M. Matsushita, S. Momiyama, T. Motobayashi, H. Otsu, C. Péron, A. Peyaud, E. C. Pollacco, J.-Y. Roussé, H. Sakurai, C. Santamaria, M. Sasano, Y. Shiga, S. Takeuchi, R. Taniuchi, T. Uesaka, H. Wang, K. Yoneda, F. Browne, L. X. Chung, Z. Dombradi, F. Flavigny, F. Giacoppo, A. Gottardo, K. Hadyska-Klek, Z. Korkulu, S. Koyama, Y. Kubota, J. Lee, M. Lettmann, C. Louchart, R. Lozeva, K. Matsui, T. Miyazaki, S. Nishimura, K. Ogata, S. Ota, Z. Patel, E. Sahin, C. Shand, P.-A. Söderström, I. Stefan, D. Steppenbeck, T. Sumikama, D. Suzuki, V. Werner, J. Wu, and Z. Xu, Phys. Rev. Lett. 119, 192501 (2017).
- [12] T. Otsuka, T. Suzuki, R. Fujimoto, H. Grawe, and Y. Akaishi, Phys. Rev. Lett. 95, 232502 (2005).
- [13] J. Van Roosbroeck, H. De Witte, M. Gorska, M. Huyse, K. Kruglov, D. Pauwels, J.-C. Thomas, K. Van de Vel, P. Van Duppen, S. Franchoo, J. Cederkall, V. N. Fedoseyev, H. Fynbo, U. Georg, O. Jonsson, U. Köster, L. Weissman, W. F. Mueller, V. I. Mishin, D. Fedorov, A. De Maesschalck, N. A. Smirnova, and K. Heyde, Phys. Rev. C 71, 054307 (2005).
- [14] V. Paar, Nuclear Physics A **331**, 16 (1979).
- [15] N. Shimizu, T. Abe, Y. Tsunoda, Y. Utsuno, T. Yoshida, T. Mizusaki, M. Honma, and T. Otsuka, Progress of Theoretical and Experimental Physics 2012, 10.1093/ptep/pts012 (2012), 01A205, https://academic.oup.com/ptep/article-pdf/2012/1/01A205/11585951/pts012.pdf.
- [16] Y. Tsunoda, T. Otsuka, N. Shimizu, M. Honma, and Y. Utsuno, Phys. Rev. C 89, 031301 (2014).
- [17] K. Langanke, G. Martínez-Pinedo, J. M. Sampaio, D. J. Dean, W. R. Hix, O. E. B. Messer, A. Mezzacappa, M. Liebendörfer, H.-T. Janka, and M. Rampp, Phys. Rev. Lett. 90, 241102 (2003).
- [18] R. P. de Groote, J. Billowes, C. L. Binnersley, M. L. Bissell, T. E. Cocolios, T. Day Goodacre, G. J. Farooq-Smith, D. V. Fedorov, K. T. Flanagan, S. Franchoo, R. F. Garcia Ruiz, A. Koszorús, K. M. Lynch, G. Neyens,

- F. Nowacki, T. Otsuka, S. Rothe, H. H. Stroke, Y. Tsunoda, A. R. Vernon, K. D. A. Wendt, S. G. Wilkins, Z. Y. Xu, and X. F. Yang, Phys. Rev. C **96**, 041302 (2017).
- [19] P.-A. Sderstrm, S. Nishimura, P. Doornenbal, G. Lorusso, T. Sumikama, H. Watanabe, Z. Xu, H. Baba, F. Browne, S. Go, G. Gey, T. Isobe, H.-S. Jung, G. Kim, Y.-K. Kim, I. Kojouharov, N. Kurz, Y. Kwon, Z. Li, K. Moschner, T. Nakao, H. Nishibata, M. Nishimura, A. Odahara, H. Sakurai, H. Schaffner, T. Shimoda, J. Taprogge, Z. Vajta, V. Werner, J. Wu, A. Yagi, and K. Yoshinaga, Nuclear Instruments and Methods in Physics Research Section B: Beam Interactions with Materials and Atoms 317, 649 (2013), xVIth International Conference on ElectroMagnetic Isotope Separators and Techniques Related to their Applications, December 27, 2012 at Matsue, Japan.
- [20] N. Fukuda, T. Kubo, T. Ohnishi, N. Inabe, H. Takeda, D. Kameda, and H. Suzuki, Nuclear Instruments and Methods in Physics Research Section B: Beam Interactions with Materials and Atoms 317, 323 (2013), xVIth International Conference on ElectroMagnetic Isotope Separators and Techniques Related to their Applications, December 27, 2012 at Matsue, Japan.
- [21] T. Kubo, D. Kameda, H. Suzuki, N. Fukuda, H. Takeda, Y. Yanagisawa, M. Ohtake, K. Kusaka, K. Yoshida, N. Inabe, T. Ohnishi, A. Yoshida, K. Tanaka, and Y. Mizoi, Progress of Theoretical and Experimental Physics 2012, 10.1093/ptep/pts064 (2012), 03C003, http://oup.prod.sis.lan/ptep/articlepdf/2012/1/03C003/11595011/pts064.pdf.
- [22] S. Nishimura, Progress of Theoretical and Experimental Physics 2012, 03C006 (2012).
- [23] M. Wang, W. Huang, F. Kondev, G. Audi, and S. Naimi, Chinese Physics C 45, 030003 (2021).
- [24] A. Tolosa Delgado, Study of beta-delayed neutron emitters in the region of 78ni and its impact on r-process nucleosynthesis, Ph.D. thesis, U. Valencia (main) (2020).
- [25] C. J. Gross, J. A. Winger, S. V. Ilyushkin, K. P. Rykaczewski, S. N. Liddick, I. G. Darby, R. K. Grzywacz, C. R. Bingham, D. Shapira, C. Mazzocchi, S. Padgett, M. M. Rajabali, L. Cartegni, E. F. Zganjar, A. Piechaczek, J. C. Batchelder, J. H. Hamilton, C. T. Goodin, A. Korgul, and W. Krolas, Acta Phys. Pol. B 40, 447 (2009).
- [26] A. Korgul, K. P. Rykaczewski, J. A. Winger, S. V. Ilyushkin, C. J. Gross, J. C. Batchelder, C. R. Bingham, I. N. Borzov, C. Goodin, R. Grzywacz, J. H. Hamilton, W. Królas, S. N. Liddick, C. Mazzocchi, C. Nelson, F. Nowacki, S. Padgett, A. Piechaczek, M. M. Rajabali, D. Shapira, K. Sieja, and E. F. Zganjar, Phys. Rev. C 86, 024307 (2012).
- [27] T. Kibdi, T. Burrows, M. Trzhaskovskaya, P. Davidson, and C. Nestor, Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment 589, 202 (2008).
- [28] F. L. Bello Garrote, E. Sahin, Y. Tsunoda, T. Otsuka, A. Görgen, M. Niikura, S. Nishimura, G. de Angelis, G. Benzoni, A. I. Morales, V. Modamio, Z. Y. Xu, H. Baba, F. Browne, A. M. Bruce, S. Ceruti, F. C. L. Crespi, R. Daido, M.-C. Delattre, P. Doornenbal, Z. Dombradi, Y. Fang, S. Franchoo, G. Gey, A. Gottardo, K. Hadyska-Klek, T. Isobe, P. R. John, H. S. Jung, I. Kojouharov, T. Kubo, N. Kurz, I. Kuti, Z. Li, G. Lorusso, I. Matea, K. Matsui, D. Men-

goni, T. Miyazaki, S. Momiyama, P. Morfouace, D. R. Napoli, F. Naqvi, H. Nishibata, A. Odahara, R. Orlandi, Z. Patel, S. Rice, H. Sakurai, H. Schaffner, L. Sinclair, P.-A. Söderström, D. Sohler, I. G. Stefan, T. Sumikama, D. Suzuki, R. Taniuchi, J. Taprogge, Z. Vajta, J. J. Valiente-Dobón, H. Watanabe, V. Werner, J. Wu, A. Yagi, M. Yalcinkaya, R. Yokoyama, and K. Yoshinaga, Phys. Rev. C 102, 034314 (2020).

- [29] S. Giraud, L. Canete, B. Bastin, A. Kankainen, A. Fantina, F. Gulminelli, P. Ascher, T. Eronen, V. Girard-Alcindor, A. Jokinen, A. Khanam, I. Moore, D. Nesterenko, F. de Oliveira Santos, H. Penttil, C. Petrone, I. Pohjalainen, A. De Roubin, V. Rubchenya, M. Vilen, and J. yst, Physics Letters B 833, 137309 (2022).
- [30] S. P. Pandya, Phys. Rev. **103**, 956 (1956).