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Phys. Rev. C **106**, 045801 — Published 6 October 2022
DOI: 10.1103/PhysRevC.106.045801

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# Investigation of direct capture in the ${}^{23}Na(p,\gamma){}^{24}Mg$ reaction

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(Dated: September 20, 2022)

The <sup>23</sup>Na(p, $\gamma$ )<sup>24</sup>Mg reaction plays an important role in the nucleosynthesis of elements in the hot bottom burning environment of asymptotic giant branch stars by providing a breakout path from the NeNa to the MgAl cycle. At temperatures above  $\approx 0.06$  GK, the underlying nuclear reaction contributions to the rate are primarily narrow resonances, but at lower temperatures direct and broad resonance tail contributions come to dominate. While there have been recent studies to improve the uncertainties associated with these narrow resonances, little attention has been paid to the non-resonant component. In this work, experimental measurements are reported over the energy range from 0.5 and 1.05 MeV proton beam energy, with a focus on studying the off-resonance region of the cross section. Several transitions were observed where two broad resonances dominate the energy range and whose low energy tails contribute strongly to the low-energy, non-resonant, cross section. In addition, a clear signature of direct capture has been observed for the first time in the <sup>23</sup>Na(p, $\gamma$ )<sup>24</sup>Mg reaction.

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# I. INTRODUCTION

The  ${}^{23}$ Na $(p,\gamma)^{24}$ Mg reaction is an important break-out 15 link from the NeNa-cycle [1] to the MgAl cycle [2]. These 16 cycles require higher temperature environments and may 17 operate in hydrogen shell burning of massive stars [3] or 18 in hot bottom burning of AGB stars [4] or in explosive hy-19 drogen environments such as ONe novae, where the reac-20 tion link may affect the production of long-lived  $^{22}\mathrm{Na}$  and 21 <sup>26</sup>Al nuclei [5]. A recent study of the reaction at the Lab-22 <sup>23</sup> oratory for Underground Nuclear Astrophysics (LUNA) <sup>24</sup> at the Istituto Nazionale di Fisica Nucleare (INFN) Lab-25 oratori Nazionali del Gran Sasso (LNGS) has probed the low energy resonant reaction contributions in the proton 26 energy range between 130 and 400 keV [6]. The strengths 27 of three resonances at 140, 251, and 309 keV were deter-28 mined. These contributions are expected to dominate 29 the reaction rate for the temperature range associated 30 with hot bottom hydrogen burning and explosive hydro-31 gen burning. 32

This work focuses on additional reaction contributions associated with direct radiative capture (DC) to bound states in <sup>24</sup>Mg and the low energy tails of broad resonances at  $E_{\rm c.m.} = 840$  keV ( $E_x = 12.53$  MeV, 1<sup>+</sup>) and

<sup>37</sup> 980 keV ( $E_x = 12.67$  MeV, 2<sup>-</sup>), as well as possible inter-<sup>38</sup> ference patterns tailing into the low energy range of inter-<sup>39</sup> est. The <sup>23</sup>Na( $p, \gamma$ )<sup>24</sup>Mg reaction in this energy range has <sup>40</sup> been studied from 250 to 2500 keV by Switkowski *et al.* <sup>41</sup> [7] and above 500 keV by Baxter *et al.* [8] and Leccia *et al.* <sup>42</sup> [9]. Higher energy resonances between 1 and 2 MeV have <sup>43</sup> been studied by Mourad *et al.* [10], Meyer *et al.* [11] and <sup>44</sup> Endt *et al.* [12]. These earlier measurements focused on <sup>45</sup> the determination of resonance strengths, while in this <sup>46</sup> study the aim is to characterize the off-resonance energy <sup>47</sup> dependence, DC, and interference effects, using *R*-matrix <sup>48</sup> and DC model techniques to estimate their contributions <sup>49</sup> to the stellar reaction rate.

50 The search for strong direct contributions to the  $_{^{51}}$   $^{23}$ Na $(p,\gamma)^{24}$ Mg reaction is greatly aided by previous stud-52 ies of the single particle strength of bound states in <sup>53</sup> the <sup>24</sup>Mg compound system. In particular Refs. [13–16] <sup>54</sup> have used (d, n) and  $({}^{3}\text{He}, d)$  transfer reaction cross sec-<sup>55</sup> tions in combination with Distorted Wave Born Anal-<sup>56</sup> yses (DWBA) to determine proton spectroscopic factors  $_{57}$  (C<sup>2</sup>S). When combined with the DC potential model for-<sup>58</sup> malism of Rolfs [17], the strongest DC transitions can be 59 identified as shown in Fig. 1. While these calculations <sup>60</sup> provide a good starting point for a study, they have large 61 uncertainties associated with them ( $\approx 50\%$ ), necessitating <sup>62</sup> the direct measurement of the  ${}^{23}Na(p,\gamma){}^{24}Mg$  cross sections if an improved level of uncertainty is to be achieved. 63

<sup>66</sup> In this work, low energy *S*-factors are estimated using <sup>67</sup> two different phenomenological models, each associated <sup>68</sup> with the analysis of different types of reaction data. For

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FIG. 1. Fractional contribution of the individual DC transitions to the total DC S-factor. Only those transitions that contribute more than 5% to the total are shown. The calculations used the  $C^2S$  values from Garrett *et al.* [16].

<sup>69</sup> the analysis of the capture cross sections measured in this <sup>124</sup> suppressor. work, phenomenological R-matrix [18] along with the ex- <sup>125</sup> 70 72 73 74 75 76 77 tual resonances in the region of the data, and those from <sup>133</sup> section were taken in close geometry. 78 background levels that mimic the summed contributions 79 from all higher energy states. The largest uncertainties 80 are often from incomplete knowledge of the level struc-81 ture between the experimental data and threshold as well 82 as contributions form the tails of higher lying resonances. 135 83

84 85 86 87 88 89 91 92 93 94 However, the present analysis relies on  $C^2S$  values from  $_{147}$  and total efficiencies from the analysis are also shown. 95 older work where this was not the case. Thus the R- 148 96 97 98 level of the off-resonance S-factors. 99

100 101 102 103 <sup>106</sup> given in Sec. VI with summary statements in Sec. VII.

### II. EXPERIMENTAL SETUP

The experiment was performed at the University of 108 Notre Dame Nuclear Science Laboratory (NSL) using 109 the Stable ion Accelerator for Nuclear Astrophysics (Sta. ANA), a 5 MV single-ended Pelletron accelerator, to pro-111  $_{112}$  duce proton beams between 0.5 and 1 MeV with maximum beam intensity on target of 15  $\mu A$ . Targets were 113 produced by evaporating  $Na_2WO_4$  onto a 0.5 mm thick 114 Ta backing, which also served as a beam stop, which was 115 electrically isolated from the rest of the beam line. The  $_{117}$  target, mounted at  $45^\circ$  degrees relative to the beam di-<sup>118</sup> rection, was water-cooled in order to reduce degradation <sup>119</sup> due to heating. A circular copper tube was mounted inside the beam line, which extended from a cold head to 120 <sup>121</sup> within a few millimeters of the target face. The copper 122 tube was both cooled with liquid nitrogen and biased to 123 -300 V, acting simultaneously as a cold trap and electron

The detector system consisted of a single, high efternal capture model (EC) [19-22] is used to fit the ex- 126 ficiency (120% relative efficiency), high purity, coperimental data and then extrapolate the S-factor to low 127 axial Germanium detector (HPGe). The detector was energies. The strength of the EC is often characterized 128 mounted on a rail system at an angle of 45° relative to by the asymptotic normalization coefficient (ANC). Here 129 the beam. The rail system allowed the detector to be easthe "off-resonance" S-factor is determined by the com- 130 ily moved to different distances from the target in order bination of the EC and tails of resonance contributions. <sup>131</sup> to make summing correction measurements during the These tail contributions are divided into those from ac-  $^{132}$  calibration. Measurements of the  $^{23}$ Na $(p, \gamma)^{24}$ Mg cross

## **III. DATA ANALYSIS**

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The full-energy peak detection efficiency of the HPGe For transfer reactions, DWBA has been traditionally 136 detector was determined with calibration sources of used to extract spectroscopic factors from forward an- $_{137}$  known activity (<sup>60</sup>Co and <sup>137</sup>Cs), and the  $E_{\rm p} = 992 \,\rm keV$ gle angular distributions. A potential model can then be 138 resonance in  ${}^{27}Al(p,\gamma){}^{28}Si$  reaction, which has a wellused, making sure to use the same potential parameters 139 known strength and branching ratios [24]. The efficiency, as the original DWBA analysis, to derive a single-particle 140 as a function of energy and distance, was parametrized direct capture (DC) cross section, which is weighted by 141 empirically [25] and fitted to the measurements at difthe isospin Clebsch-Gordan factor [17]. Here the largest 142 ferent detector distances, accounting for summing-in and uncertainties in the calculated S-factors often come from 143 summing-out. The observed full-energy peak efficiencies uncertainties in the phenomenological potential model 144 (with summing) are compared to those obtained using parameters. It is also now common practice to instead  $_{145}$  the procedure described in Imbriani *et al.* [25] and are derive an ANC directly from the DWBA analysis [23]. 146 shown in Fig. 2. The full-energy peak (without summing)

The yields of the  $\gamma$ -rays corresponding to primary tranmatrix calculations and potential model calculations are  $_{149}$  sitions to the ground state GS ( $\gamma_0$ ), 1.37 ( $\gamma_1$ ), 4.12 ( $\gamma_2$ ), performed separately and the comparison is made at the  $_{150}$  4.24 ( $\gamma_3$ ), 5.23 ( $\gamma_4$ ), 6.01 ( $\gamma_5$ ), 8.44, and 10.73 MeV ex-<sup>151</sup> cited states were determined from the spectra. An exam-In the following, we will first discuss the experimental  $_{152}$  ple off-resonance spectrum at  $E_p = 879$  keV, where the set-up in Sec. II and results in Sec. III before we come to 153 strongest transitions are indicated, is given in Fig. 3. The the detailed *R*-matrix (Sec. IV) and MCMC uncertainty <sup>154</sup> thicknesses of the evaporated targets were determined (Sec. IVA) analysis of the low energy reaction cross sec- 155 through a) scans of the narrow resonances at 309 keV and tion. The DC calculations are then described in Sec. V. 156 510 keV, and b) from the width of the observed primary A discussion of the results and revised reaction rates are <sup>157</sup> gamma peaks at beam energies for which the cross sec-

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FIG. 2. The top panel shows a comparison between close geometry (red) and far geometry (254 mm, green) measurements of the  $\gamma$ -ray full-energy peak detection efficiency. Both data sets include the effects of summing, but in the far geometry they are negligible. The solid lines indicate empirical fits to the data where the summing is either negligible (far geometry) or has been corrected as described in Imbriani et al. [25]. The total efficiency is indicated by the dashed grey line. The bottom panel gives the residuals between the measured and simulated full-energy peak efficiencies (with summing) in close geometry.

158 tion only varies slowly over the beam energy range in the 177 rectly with observable energies and partial widths and <sup>159</sup> target (i.e., off-resonance). Secondary peaks were found <sup>178</sup> to remove the need for boundary conditions. For bound 160 to be heavily affected by diffusion of sodium into deeper 179 states, the strength of the external capture is parameter-161 162 163 164 was determined by fitting the primary  $\gamma$ -ray peaks con- 185 the direct contribution to the cross section. 166 sidering the variation of the yield due to the thickness of  $_{186}$ 167 169 <sup>170</sup> in the Supplemental Material [27].

### IV. **R-MATRIX ANALYSIS**

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172  $_{173}$  perimental data over the energy regions exhibiting broad  $_{195}$  width. Therefore, only the ground state proton and  $\alpha$ -174 resonances or direct capture contributions using the code 196 particle particle particle are included in the *R*-matrix analysis.  $_{175}$  AZURE2 [29, 30]. The alternative *R*-matrix parameter-  $_{197}$  For the  $\gamma$ -ray partitions, only the observed transitions to <sup>176</sup> ization of Brune [31] was utilized in order to work di- <sup>198</sup> the ground state  $(J^{\pi} = 0^+)$ , first  $(E_x = 1.369 \text{ MeV}, 2^+)$ ,

layers of the targets. Although this effect was found to be  $_{180}$  ized in terms of ANCs. Channel radii of  $a_p = 6.0$  fm and smaller in targets with less time between evaporation and  $a_{181} a_{\alpha} = 6.5$  fm were adopted for the proton and  $\alpha$ -particle measurement, due to the difficulties in correcting for this 182 channels, respectively. Masses and separation energies effect, secondary peaks were not evaluated. The differ- 183 were taken from the mass evaluation [32, 33]. For all of ential cross sections for the primary capture transitions  $_{184}$  the transitions observed, E1 multipolarity dominates for

Over the energy range of interest, levels populated by the target using the methods described in Di Leva et al. 187 proton capture in the <sup>24</sup>Mg compound nucleus can par-[26] and are shown in Figs. 4 and 5. The data are given 188 ticle decay to the ground state and first excited state of  $_{189}$  <sup>23</sup>Na and <sup>20</sup>Ne through proton and  $\alpha$ -emission, respec-<sup>190</sup> tively. However, the total widths of the levels populated <sup>191</sup> in the present  ${}^{23}$ Na $(p, \gamma)^{24}$ Mg data are dominated by <sup>192</sup> de-excitation through their ground state proton widths. <sup>193</sup> An additional resonance, excited in the  ${}^{23}$ Na $(p, p_0)^{23}$ Na An R-matrix analysis has been performed for the ex-  $^{194}$  data, also has a substantial ground state  $\alpha$ -particle



FIG. 3. An off-resonance  $\gamma$ -ray spectrum at  $E_p = 879$  keV. The strongest transitions (the full-energy peak and two escape peaks where they are visible) of the  ${}^{23}$ Na $(p, \gamma)^{24}$ Mg reaction are indicated. A strong background peak is observed from the  ${}^{19}$ F $(p, \alpha \gamma)^{16}$ O reaction, owing to fluorine contamination in the Ta backing. Transitions are indicated for both primary resonance (R) or direct capture (DC) to different final states (indicated by the number of the excited state for low lying states) as well as the strong secondary transition from the 1<sup>st</sup> excited state to the ground state (GS). The line at 511 keV due to electron/positron annihilation is also indicated.

<sup>199</sup> second (4.123 MeV, 4<sup>+</sup>), third (4.238 MeV, 2<sup>+</sup>), forth <sup>217</sup> scattering data. An *R*-matrix fit to the <sup>23</sup>Na $(p, p_0)^{23}$ Na 200 201 are included. 202

Previous data from the literature over the energy range 221 203 204 205 206 207 208 209 210 211  $_{212}$   $\alpha$ -scattering, the data of Goldberg *et al.* [34] cover the  $_{230}$  external capture (EC) to model the DC. Only the transi-<sup>213</sup> corresponding excitation energy range, but since the res-<sup>231</sup> tion at 10.731 MeV shows a measurable mixture of reso- $_{214}$  onances observed in the  $^{23}$ Na+p data correspond to levels  $_{222}$  nance and direct contributions, having a resonance that <sup>215</sup> with total widths dominated by their ground state pro-<sup>233</sup> corresponds to the  $E_x = 12.53$  (1<sup>+</sup>) state and flat off- $_{216}$  ton widths, they do not appear as resonances in the  $\alpha$ - $_{234}$  resonance S-factor indicative of direct capture. A break

 $(5.235 \text{ MeV}, 3^+)$ , fifth (6.011 MeV, 4<sup>+</sup>), and two high- <sup>218</sup> data of Baumann *et al.* [28] and to the <sup>23</sup>Na $(p, \gamma)^{24}$ Mg lying  $(8.437 \text{ MeV}, 1^{-} \text{ and } 10.731 \text{ MeV}, 2^{+})$  excited states <sup>219</sup> primary capture data of the present work is shown in 220 Figs. 4 and 5.

The parameters varied in the R-matrix fit are summaunder investigation are quite limited. There are many  $_{222}$  rized in Table I. The experimental data for GS, 1.37 ( $\gamma_1$ ), measurements of the  ${}^{23}$ Na+p reactions at lower ener-  ${}^{223}$  4.24 ( $\gamma_3$ ), and 8.44 MeV transitions could be reproduced gies, characterizing narrow resonance strengths, but most 224 by just resonances arising from the two broad unnatucross section measurements begin at energies above the  $_{225}$  ral parity resonances at  $E_x = 12.53$  (1<sup>+</sup>) and 12.67 (2<sup>-</sup>) current data. The broad resonances observed here corre- 226 MeV as well as interference from background contribuspond to those seen in the  ${}^{23}$ Na $(p, p_0)^{23}$ Na elastic scat-  ${}^{227}$  tions in some cases. On the other hand, the data for the tering data of Baumann et al. [28]. This is the only pro-  $_{228}$  transitions to the excited states at 4.12 ( $\gamma_2$ ), 5.23 ( $\gamma_4$ ), ton scattering data reported over this energy range. In 229 and 6.01 MeV could be described using only hard-sphere



FIG. 4. *R*-matrix fit to the <sup>23</sup>Na( $p, p_0$ )<sup>23</sup>Na data of Baumann *et al.* [28] and the <sup>23</sup>Na( $p, \gamma_{0,1,3}$ )<sup>24</sup>Mg data of the present work. The solid red lines indicate the best fit (50% quantile) while the dashed red lines indicate the 16 and 84% quantiles found from the MCMC analysis. The "n" in the two upmost plots are normalization factors, see Table I.

<sup>235</sup> down of the *R*-matrix reaction components for the fit to <sup>254</sup> tributions alone were sufficient to describe the experi $z_{256}$  the  $E_x = 10.731$  MeV final state transition is shown in  $z_{55}$  mental data. However, to obtain upper limits for the DC 237 Fig. 6. The *R*-matrix parameters for the best fit are given 256 of these transitions, an EC contribution was included in 238 in Table I.

Note that there are two closely spaced bound states  $_{\scriptscriptstyle 258}$  $_{240}$  in  $^{24}$ Mg at  $E_x = 8.35798(13)$  and 8.43731(15) with  $_{259}$  *R*-matrix fit using the Bayesian *R*-matrix Inference Code  $_{241} J^{\pi} = 1^{-}$  and  $4^{+}$ , respectively. It has been assumed  $_{260}$  Kit (BRICK) [35]. BRICK provides communication be-242 that the observed transition yield reported here for the 261 tween the MCMC Python routine emcee [36] and the  $_{243} E_x = 8.44 \text{ MeV}$  transition is dominated by the 1<sup>-</sup> tran- $_{262} \text{ C} + + R$ -matrix code AZURE2 [29, 30]. The MCMC rou-<sup>244</sup> sition. This is suggested by the calculation of the DC <sub>263</sub> tine requires priors for the *R*-matrix fit and normalization  $_{245}$  cross sections for the two transitions from the  $C^2S$  val-  $_{264}$  parameters. For the *R*-matrix fit parameters, uniform 246 ues of Garrett et al. [16]. It is possible that the yield over 265 priors were taken, while for the normalization factors, a  $_{247}$  the resonance could have significant contributions for the  $_{266}$  Gaussian prior with a  $\sigma = 10\%$  was used for both the <sup>248</sup> 4<sup>+</sup> transition, as that can be populated through an E2 <sub>267</sub> scattering and capture data based on the systematic un-249 transition.

### MCMC uncertainty analysis Α. 250

In the *R*-matrix fit described in Sec. IV, the present  $_{273}$ 251 <sup>252</sup> data could be described as a mixture of resonance and <sup>274</sup> ments for many of the observed transitions provided only <sup>253</sup> EC contributions. For some transitions, resonance con-<sup>275</sup> upper limits for the external capture contribution. The

<sup>257</sup> every transition.

An MCMC uncertainty analysis was performed on the <sup>268</sup> certainty information found in the present work and in <sup>269</sup> Baumann et al. [28]. The parameter posteriors and cor-270 relations calculated by emcee are provided as a corner  $_{271}$  plot [37] in the Supplemental Material [27] and the 68%<sup>272</sup> confidence level uncertainties are given in Table I.

As described in Sec. IV, the cross section measure-



FIG. 5. As Fig. 4, except for the  ${}^{23}$ Na $(p, \gamma_{2,4,5,8,44 \text{ MeV},10.37 \text{ MeV}})^{24}$ Mg cross sections.



FIG. 6. *R*-matrix components used to fit the primary transition to the  $E_x = 10.73$  MeV final state in the  ${}^{23}\text{Na}(p,\gamma){}^{24}\text{Mg}$  reaction. The experimental data indicate the presence of both 290 a single resonance, corresponding to the unbound state at 291  $E_x = 12.53$  MeV, and external capture, modeled using *E*1 292 multipolarity only. 293

<sup>276</sup> upper limits (68% confidence) for the ANCs, which were <sup>277</sup> used to parameterize the strength of the external cap-<sup>278</sup> ture, for each of the observed  $\gamma$ -ray transitions are given <sup>279</sup> in Table I. The dimensionless reduced width can be used <sup>280</sup> to provide an approximate measure of the single particle <sup>281</sup> or cluster nature of a state [38] and is given by

$$\theta^2 = \gamma^2 / \gamma_W^2, \tag{1}$$

 $_{^{282}}$  where  $\gamma^2$  is the observable reduced width and  $\gamma^2_W$  is the  $_{^{283}}$  Wigner limit

$$\gamma_W^2 = 3\hbar^2/2\mu a^2. \tag{2}$$

<sup>284</sup> Here  $\mu$  is the reduced mass and a is the channel radius. A <sup>285</sup> pure single particle or cluster state corresponds to  $\theta_{\alpha}^2 \approx 1$ . <sup>286</sup> The upper limits for the dimensionless reduced width am-<sup>287</sup> plitudes for the observed transitions of this work are given <sup>288</sup> in Table II. The largest  $\theta^2$  upper limits are found for the <sup>289</sup> ground state and the 5<sup>th</sup> excited state, but even these <sup>290</sup> upper limits are significantly smaller than one.

The posteriors for the normalization factors of the cap-<sup>292</sup> ture data from the present work returned their priors, <sup>293</sup> indicating, as expected, that there are no other con-<sup>294</sup> straints present in the model that determine the abso-<sup>295</sup> lute scale of the capture cross section. On the other

Ec.m.	$E_x$	$J^{\pi}$	$\Gamma_p$ or ANC	$\Gamma_{\gamma_0}$	$\Gamma_{\gamma_1}$	$\Gamma_{\gamma_3}$	$\Gamma_{8.44 \text{ MeV}}$	$\Gamma_{10.73 \text{ MeV}}$	Γα
(keV)	$(eV)$ (eV) or $(fm^{-1/2})$								
Bound states									
	$0^{\mathbf{a}}$	$0^{+}$	$<\!\!52$						
	$1.3687^{\rm a}$	$2^{+}$	<-25						
	$4.1229^{\rm a}$	$4^{+}$	<8						
	$4.2382^{\rm a}$	$2^{+}$	<22						
	$5.2351^{\rm a}$	$3^{+}$	<11						
	$6.0108^{\rm a}$	$4^{+}$	<11						
	$8.4373^{\rm a}$	$1^{-}$	<7						
	$10.7308^{\rm a}$	$2^{+}$	3.81(37)						
Unbo	und states								
841.2	12.5339(1)	$1^+$	9270(110)	0.292(28)	0.0491(48)	0.113(12)	0.018(3)	$7(1) \times 10^{-3}$	
887.1	12.5798(2)	$2^{+}$	1150(120)						-2900(500)
981.5	12.6745(1)	$2^{-}$	4000(100) / 660(70)		-0.062(6)	0.54(6)	-0.015(6)/0.095(5)		
Background states									
	$13^{\mathrm{a}}$	$1^{+}$	$6.7(2) \times 10^4$	20(10)					
	$13.1^{\rm a}$	$2^{-}$	$(-3.0(2)/-3.7(2)) \times 10^4$						
	$30^{a}$	$1^+$	$5 \times 10^{7a}$	$-2(1) \times 10^4$			$-3.6(5) \times 10^4$		
Norm	alization fa	acto	<u>rs</u>						
$n_{\text{Bauma}}$	$ann, \theta_{lab} = 87.5^{\circ}$	= 0	$.95_{-0.03}^{+0.03}$						
$n_{\mathrm{Baums}}$	$ann, \theta_{lab} = 156.5$	• =	$1.09^{+0.06}_{-0.06}$						

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TABLE I. Best fit parameters for the MCMC analysis using the Python package emcee implemented in the R-matrix code AZURE2 with the Python package BRICK. The signs on the partial widths indicate the sign of the corresponding reduced width amplitude. Upper limits correspond to a 68% confidence limit.

<sup>a</sup> Fixed parameter.

TABLE II. Summary of the upper limits (68% confidence) for the ANCs, reduced width amplitudes, and dimensionless reduced proton widths for the final states in <sup>24</sup>Mg of the of  $\gamma_{(\text{Wigner})} = 1.34 \text{ MeV}^{1/2}$  was used.

$E_x$	$J^{\pi}$	ANC	$\gamma$	$\theta^2$
(MeV)		$({\rm fm}^{-1/2})$	$({\rm MeV}^{1/2})$	
0	$0^+$	<52	< 0.70	< 0.27
1.369	$2^{+}$	$<\!\!25$	< 0.23	< 0.029
4.123	$4^{+}$	< 8	< 0.23	< 0.029
4.238	$2^{+}$	$<\!\!22$	< 0.33	< 0.060
5.235	$3^{+}$	<11	< 0.41	< 0.093
6.011	$4^{+}$	<11	< 0.50	< 0.14
8.437	$1^{-}$	<7	< 0.31	< 0.053
10.731	$2^{+}$	3.81(37)	0.14	0.011

<sup>297</sup> factors are somewhat different from their priors, giving <sup>324</sup> to characterize the total DC.  $_{\rm 298}$  both a different central value and a much smaller un-  $_{\rm 325}$  $_{299}$  certainty. This results from the *R*-matrix model calcu- $_{326}$  the different reaction components that make up the total 300 301 302 cally constrained portion of the cross section dominates 329 by several recent calculations that have used a similar <sup>304</sup> ization factors produce somewhat different cross sections <sup>331</sup> also reported in the NACRE compilation [41], but there 305 to those given by Baumann et al. [28], they are within the 332 the importance of the DC contribution was not realized, 306 10% systematic uncertainty estimated in that work (see 333 thus the non-resonant component was greatly underesti-<sup>307</sup> Table I). It is always possible that there could be a signif-<sup>334</sup> mated. The transfer study of Hale *et al.* [42] was focused

<sup>308</sup> icant modification to the low energy Coulomb scattering <sup>309</sup> cross section from broad resonances or other background <sup>310</sup> sources. Additional scattering data, with comprehensive transitions observed in the present study. A Wigner limit <sup>311</sup> angular coverage, are required in order to make more 312 definitive conclusions.

### DC CALCULATIONS v.

The proton separation energy in <sup>24</sup>Mg is  $_{315} S_p = 11.69$  MeV, with more than 60 known proton- $_{316}$  bound states. Further, proton  $C^2S$  values from transfer 317 studies indicate that the DC de-excitation is rather <sup>318</sup> democratic [13, 15, 16] (see Sec. I) indicating a rather <sup>319</sup> homogeneous distribution of the single particle strength.  $_{320}$  This was reflected in the  $\theta^2$  upper limits found from 321 the *R*-matrix analysis in Sec. IVA, where all the  $_{322} \theta^2 \ll 1$ . Thus the  $^{23}$ Na $(p, \gamma)^{24}$ Mg reaction requires the 296 hand, the posteriors for the scattering data normalization 323 measurement of several individual transitions in order

Fig. 7 summarizes the fractional contributions from lation of the Coulomb scattering cross section and the  $_{327}$   $^{23}$ Na $(p, \gamma)$ <sup>24</sup>Mg reaction rate at low temperatures. Below presence of large amounts of data where this theoreti- 328 ≈0.04 GK, non-resonant capture dominates, as reported in off-resonance regions (see Fig. 4). While the normal- 330 non-resonant S-factor [6, 39, 40]. A rate calculation was



FIG. 7. Relative contributions to the total reaction rate over the low temperature region where the rate is dominated by the DC and resonances at  $E_p = 133, 251$ , and 309 keV.

on levels near the proton separation energy, thus their 335 calculations of the DC also continued to relied on previously measured proton  $C^2S$  values compiled by Endt 337 [43], coming from previous (d, n) and  $({}^{3}\text{He}, d)$  transfer 338 reactions [13–16]. The  $C^2S$  values have uncertainties of  $\approx 30-50\%$  [16], which translate directly to the uncertain-340 ties in the DC S-factors and the DC contribution to the 341 reaction rate below  $\approx 0.04$  GK. 342

Detailed comparisons of the individual transition S-343 344 factors of this work with past calculations cannot be made because previous works only reported the total S-345 factor (sum over all final states). However, these past 346 works all calculated their DC contributions using a po-347 tential model to obtain a single particle S-factor that was 348 then weighted by the  $C^2S$  values from the evaluation [43], 349 which come mainly from Garrett  $et \ al. \ [16]$ . As the po-350 tential model parameters were all well documented, these 351 calculations could be readily repeated for the individual 379 352 states with  $C^2S$  values larger than 0.1 (see Fig. 1). 353

The single particle DC S-factors  $(S_{sp})$  have been calcu-354 <sup>355</sup> lated using the potential model code JEZEBEL [44]. The  $_{\rm 356}$  DC S-factors (S<sub>DC</sub>) were then calculated as

$$(2J+1)S_{\rm DC} = (2J+1)C^2S \times S_{\rm sp},$$
 (3)

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 $_{357}$  using the  $(2J+1)C^2S$  values from Table VIII of Gar- $_{358}$  rett *et al.* [16] and the level spins (J) from the compi-<sup>359</sup> lation [45]. For comparison, the sum of these individual  $_{360}$  transitions is compared with the DC S-factor given by Hale *et al.* [42] and the non-resonant S-factor (which also 361 contains high energy resonant tail contributions) from Il-362 iadis et al. [39] in Fig. 8. 363

The non-resonant S-factors upper limits determined 364  $_{365}$  from the *R*-matrix fit upper limits were found to be, for <sup>366</sup> the most part, consistent with those calculated using the <sup>393</sup>



FIG. 8. Comparison of different total non-resonant S-factor calculations for the  ${}^{23}$ Na $(p, \gamma)^{24}$ Mg reaction. The JEZEBEL calculations from this work (red solid line) and those of Hale et al. [42] (grey dashed line) represent potential model DC Sfactors. That of Iliadis et al. [39] (black dashed-dotted line) is "non-resonant" calculation including both DC and the tails of higher energy broad resonances. The resonant only S-factor from the *R*-matrix calculation is also shown (green double dotted-dashed line), as well as the sum of it and the JEZEBEL calculation (yellow dash-dash dotted line).

 $_{367}$  potential model and  $C^2S$  values. In most transitions ob- $_{368}$  served here, the constraints on the non-resonant S-factor  $_{369}$  from the transfer reaction  $C^2S$  values were still more 370 stringent as illustrated in Figs. 9 and 10. Notably, for  $_{371}$  the  $E_x = 10.73$  MeV transition, it has been found that <sup>372</sup> an external capture component was statistically signifi-373 cant in order to achieve a good fit. The experimental  $_{\rm 374}$  data indicated a non-resonant S-factor smaller than that  $_{375}$  of the transfer reaction  $C^2S$  by  $\approx 50\%$ . Therefore, the 376 uncertainty for the overall non-resonant components of  $_{\rm 377}$  the S-factor have been increased from the 30-40% uncer-<sup>378</sup> tainty estimated by Garrett *et al.* [16] to 50%.

### **REACTION RATE CALCULATIONS** VI.

For the narrow resonance contributions, strengths and <sup>381</sup> energies have been taken from Boeltzig et al. [6], except <sup>382</sup> for the energy of the lowest known resonance that has re-<sup>383</sup> cently been revised by Marshall *et al.* [46]. At tempera- $_{384}$  tures below  $\approx 0.04$  GK, the direct capture and broad resonance tail contributions studied in this work dominate 385 the reaction rate, as shown in Fig. 7. The non-resonant 386 387 portion of the reaction rate and its uncertainties were calculated as follows: 388

- The median rate has been determined using the Sfactors calculated using JEZEBEL (see Sec. V). The present results are found to be about  $\approx 20\%$  lower than previous calculations.
- The lower limit has been calculated using a 50%



FIG. 9. Comparisons of the Hard sphere external capture S-factors calculated from AZURE2 using the upper limit ANCs, determined in this work, given in Table I (red line) to the direct capture S-factors calculated using  $C^2S$  values from the Garrett *et al.* [16] and the potential model code JEZEBEL (blue dashed-dotted line).



FIG. 10. As Fig. 9 but for the transition to the fifth excited state in <sup>24</sup>Mg and the two high lying bound states at  $E_x = 8.44$  and 10.73 MeV.



FIG. 11. Ratio of the present rate (blue solid line) and uncertainty (blue shaded region) to the median of Marshall et al. [46]. The relative uncertainties of Marshall *et al.* [46] (gray shaded region) are also shown for comparison.

uncertainty for the  $C^2S$  values and applying this 394 to the JEZEBEL S-factors. 395

• The upper limit has likewise been calculated using 396 a 50% uncertainty for the  $C^2S$  values and apply-397 ing this to the JEZEBEL S-factors. An additional 398 contribution has also been added for the ground 399 state transition, where it was found in the *R*-matrix 400 fit that the resonance tails of the ground state 401 transition can make a significant contribution (see 402 Sec. IV). 403

404 405 407 408 409 <sup>410</sup> certainties, which can be attributed to the present study's <sup>447</sup> of Edinburgh, UK.

411 inclusion of the interference uncertainty of the ground <sup>412</sup> state transition and a larger uncertainty in the  $C^2S$  val-<sup>413</sup> ues than taken in previous calculations (50% instead of 30%). 414

### VII. SUMMARY

The present study reports cross section measurements 416 for eight individual transitions of the  ${}^{23}Na(p,\gamma){}^{24}Mg$  re-417 action over the laboratory energy range from  $E_p = 0.5$ 418 to 1.05 MeV for the first time. Two broad resonances 419 were observed, whose low energy tails have substantial 420 contributions to the non-resonant S-factor at low en-421 422 ergy. Upper limits were also determined for the exter-<sup>423</sup> nal capture contributions for each transition through an *R*-matrix analyse, which was then compared with DC 424  $_{425}$  S-factors calculated using a potential model and  $C^2S$ 426 values from a transfer measurement. The two methods were found to produce consistent non-resonant S-427 factors and the rather democratic decay through several 428 transitions was observed. Finally, direct capture in the 429  $_{\rm 430}$   $^{23}{\rm Na}(p,\gamma)^{24}{\rm Mg}$  reaction was observed for the first time  $_{431}$  through the  $E_x = 10.73$  MeV transition. Direct measure-<sup>432</sup> ments with greater off-resonance sensitivity are needed <sup>433</sup> in order to improve the constraint of the several other 434 transitions that make significant contributions to the to-435 tal off-resonance capture cross section. A transfer mea-<sup>436</sup> surement to determine bound state ANCs is also highly 437 recommended.

# ACKNOWLEDGMENTS

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This research utilized resources from the Notre Dame 439 440 Center for Research Computing and was funded by the As shown in Fig. 11, the present reaction rate has a 441 National Science Foundation through Grant No. PHYcentral value that is about 20% less than that given 442 2011890 (University of Notre Dame Nuclear Science Labrecently by Marshall et al. [46], but remains within their 443 oratory) and Grant No. PHY-1430152 (the Joint Instiuncertainty range. This lower value is likely the result of 444 tute for Nuclear Astrophysics - Center for the Evolution differences in the potential model codes used and adopted 445 of the Elements). M.W. acknowledges support as a Wolf- $C^2S$  values. The present study finds somewhat larger un- 446 son Fellow of the British Royal Society at the University

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