

CHCRUS

This is the accepted manuscript made available via CHORUS. The article has been published as:

Spin-wave analysis of the low-temperature thermal Hall effect in the candidate Kitaev spin liquid α -RuCl_{3}

Jonathan Cookmeyer and Joel E. Moore Phys. Rev. B **98**, 060412 — Published 27 August 2018 DOI: 10.1103/PhysRevB.98.060412

Spin Wave Analysis of Low-Temperature Thermal Hall Effect in the Candidate Kitaev Spin Liquid α -RuCl₃

Jonathan Cookmeyer¹ and Joel E. $Moore^{1,2}$

¹Department of Physics, University of California, Berkeley, CA, 94720, USA and

²Materials Sciences Division, Lawrence Berkeley National Laboratory, Berkeley, California, 94720, USA

(Dated: August 15, 2018)

Proposed effective Hamiltonians from the literature for the material α -RuCl₃ are used to compute the magnon thermal Hall conductivity, κ_{xy} , using linear spin wave theory for the magnetically ordered state. No model previously proposed that was tested explains published experimental data. Models with Kitaev interaction K > 0 are seen to predict $\kappa_{xy} \gtrsim 0$, which is inconsistent with the data. Fluctuations toward a Kitaev-type spin liquid would have the wrong sign to explain the data. However, a slight variant of a previously proposed model predicts a large κ_{xy} , demonstrating that the low-temperature thermal Hall effect could be generated exclusively by the Berry curvature of the magnon bands. The experimental data of κ_{xy} can therefore serve as another method to constrain a proposed effective Hamiltonian.

I. INTRODUCTION

The Kitaev spin-spin interaction on a honeycomb lattice provides a solvable model with abelian and nonabelian anyonic excitations, as well as Majorana edge modes.¹ Renewed interest in finding a material that realizes the Kitaev interaction was, in part, sparked by the observation that it can arise as an effective Hamiltonian for some transition metals with strong spin-orbit coupling.²

The material α -RuCl₃ has generated much excitement due to it presenting some promise as such a Kitaevmagnet (see Ref. 3 and references therein). The Ru atoms compose an ideal honeycomb lattice, and the material itself is a Mott insulator with the requisite spinorbit coupling.^{4,5} At low temperatures, the spins enter into a zigzag ordering, but a large in-honeycomb-plane field may induce a spin liquid state.^{6–10}

Raman scattering results on α -RuCl₃ reveal a continuum of excitations that exists above and below the magnetic transition temperature;¹¹ after subtracting a bosonic background, the excitations appear to be fermionic, which suggests proximity to a spin liquid state.¹² The continuum of excitations is also seen in inelastic neutron scattering data and seem to be qualitatively similar to that expected from the pure Kitaev model,^{7,13} and linear spin wave theory fits to inelastic neutron scattering data suggest a significant Kitaev interaction.^{8,13} Experiments with terahertz spectroscopy^{14–16} and electron spin resonance¹⁷ above and below the field induced transition further demonstrate the existence of interesting features in the excitation spectrum that might also be a sign of the Kitaev spin liquid.

Numerous theoretical studies have proposed effective Hamiltonians for α -RuCl₃, which have all revealed a Kitaev term,^{18–23} and the symmetric off-diagonal exchange term $\Gamma^{24,25}$ plays an important role.^{18,19} As summarized in Ref. 19, the differences in model proposals comes from the fact that there are many different parameters allowed by symmetry for an effective Hamiltonian, different crystal structures have been proposed (with C2/m being preferred from more recent X-ray experiments^{26,27}), and the first-principle calculations depend heavily on interaction parameters. Much of the theoretical analysis of the above experiments has relied on spin wave theory (for below $T_N \approx 7$ K) and/or on calculations within the pure Kitaev model as it is accessible to quantum Monte Carlo simulations (see Ref. 28, for example).

Thermal conductivity measurements also provide a unique probe of this material and have generally seen "unusual" results.^{9,10,29–31} A subset of the thermal conductivity data is an observation of a sizable thermal Hall conductivity κ_{xy} ,^{10,29,30} which, when compared with the theoretical predictions of the thermal hall conductivity of a pure Kitaev model at non-zero temperature,²⁸ perhaps suggests the Kitaev-magnet nature of α -RuCl₃. Additionally, in the case of a large in-plane magnetic field, κ_{xy}/T is reportedly quantized¹⁰ at the same value that would be expected from the pure Kitaev model for a Majorana edge mode, $\pi k_B^2/(12\hbar)$.¹

Though the data is not perfectly quantized, that may be explained due to interaction with phonons.^{32,33} Theoretical investigation has tried to explain the quantized value by finding the spin-liquid ground state of proposed Hamiltonians through a variational Monte Carlo method.³⁴ It was shown that a \mathbb{Z}_2 spin-liquid, as in the Kitaev model, is not preferred within the proposed $K-\Gamma$ model (see Eq. (1) and Table I), and the authors of Ref. 34 claim that such a result will be true for a $J_1 - K - \Gamma - J_3$ model too.

Interestingly, relative to the quantized value, the data of Ref. 29 also suggests a large magnitude κ_{xy}/T even in the magnetically ordered state of the material below $T_N \approx 7$ K, with sign reversed relative to the putative spin liquid state. Although the diagonal thermal conductivity κ_{xx} receives a large contribution from phonons, the phonon thermal Hall conductivity at $T \leq 10$ K and $\mu_0 H \leq 15$ T has been measured for multiple other materials to be small ($\kappa_{xy} \leq 10^{-4}$ W/K/m).^{35,36} It is well known that Berry curvature of the magnon bands can induce a finite thermal Hall effect,^{37,38} and it is well



Figure 1. (Color online). The conventions we use. The honeycomb lattice is shown on the right with the X, Y, and Z bond types labeled. The dashed line shows a unit cell containing one point in each of the four sub-lattices, $\{A, B, C, D\}$. The first Brillouin zone (1BZ) is shown in the upper left with four points labeled. The dashed line represents the 1BZ corresponding to the dashed box on the honeycomb lattice where the solid line represents the 1BZ of the triangular lattice. There are two coordinate systems xyz and abc that are shown in the bottom left. Note c is in the [111] direction and a is in the [112] direction. We fix the lattice spacing to be 1 so that $|\alpha_i| = 1$.

studied that honeycomb Hamiltonians can have thermal Hall effects.^{39–41} Recent studies have also found that *ferromagnetically* ordered Kitaev materials have magnon bands with non-zero Chern number⁴² and thermal Hall effect.⁴³

To our knowledge, however, no theoretical calculation of the thermal Hall conductivity due to magnons for low magnetic fields for α -RuCl₃ has been carried out. Taking the low-temperature data of Ref. 29 at face value, it provides a new test of any proposed effective Hamiltonian. In this work, we carry out such a calculation via linear spin wave theory (SWT). Though SWT may be unable to capture some features of the aforementioned experiments, it has been seen to well explain the THz spectroscopy data.¹⁶ Indeed, in this work, we find that no previously proposed model predicts a large enough magnon thermal Hall effect, but, by decreasing the strength of a third nearest-neighbor Heisenberg interaction, such a model can be found. The conclusion is that the Berry phase of magnon bands could explain the experimental observations in the ordered phase, but only if the effective Hamiltonian is somewhat different from previous proposals.

II. SPIN WAVE THEORY

Many effective Hamiltonian models have been proposed for α -RuCl₃ (see Table I), and most of them can

be captured by a $J_1 - K - \Gamma - \Gamma' - J_3$ Hamiltonian:

$$H = \sum_{\langle ij\rangle} J_1 \mathbf{S}_i \cdot \mathbf{S}_j + K S_i^{\gamma} S_j^{\gamma} + \Gamma(S_i^{\alpha} S_j^{\beta} + S_i^{\beta} S_j^{\alpha}) + \Gamma' \left[S_i^{\gamma} (S_j^{\alpha} + S_j^{\beta}) + S_j^{\gamma} (S_i^{\alpha} + S_i^{\beta}) \right]$$
(1)
$$+ \sum_{\langle \langle \langle ij \rangle \rangle \rangle} J_3 \mathbf{S}_i \cdot \mathbf{S}_j - \sum_i \mathbf{h} \cdot \mathbf{S}_i$$

where $\mathbf{h} = g\mu_B\mu_0\mathbf{H}$ with μ_B the Bohr magneton and g the g-factor. The index $\gamma(i, j)$ refers to the bond type of (i, j) as indicated in Fig. 1, and α, β refer to the other two coordinates (e.g. $(\alpha, \beta) = (x, y)$ if $\gamma = z$). Note that the axes are arranged such that the, e.g., X bond type is perpendicular to the x-axis.

Proceeding with a standard spin wave theory (SWT) analysis following Ref. 44, we can arrive at a Hamiltonian of free magnons.⁴⁵ We allow the spin moments to be pointed in arbitrary directions along the four sublattices indicated in Fig. 1. We introduce four types of Holstein-Primakov bosons b_i^X on the $X \in \{A, B, C, D\}$ sublattice at point r_i . We fix the angles by insisting on having no linear term in the boson creation/annihilation operators, and we always look for a zigzag solution (i.e. $\theta_A = \theta_B, \theta_C = \theta_D, \phi_A = \phi_B$ and $\phi_C = \phi_D$). Defining $\psi_k^{\dagger} = (b_k^{A,\dagger}, b_k^{B,\dagger}, b_k^{C,\dagger}, b_k^{D,\dagger}, b_{-k}^A, b_{-k}^B, b_{-k}^C)$, we can write the Hamiltonian, up to a constant, as

$$H = \frac{1}{2} \sum_{k} \psi_{k}^{\dagger} \mathcal{H} \psi_{k} = \sum_{k} \sum_{n=1}^{4} \omega_{n} \left(\gamma_{k,n}^{\dagger} \gamma_{k,n} + \frac{1}{2} \right), \quad (2)$$

where, in the last step, we perform a Bogoliubov transformation $(\vec{\gamma}_k, \vec{\gamma}_{-k}^{\dagger})^T = \phi_k = T_k^{-1}\psi_k$ to diagonalize the Hamiltonian. To satisfy the boson commutation relations $\sigma_3 T_k^{\dagger} \sigma_3 = T_k^{-1}$ with $\sigma_3 = \text{diag}(1, 1, 1, 1, -1, -1, -1, -1)$. To ensure the validity of SWT, we compute the spin

To ensure the validity of SWT, we compute the spin reduction for the ith boson

$$\Delta S_{0,i} = \frac{1}{V_{1BZ}} \int_{1BZ} d^2k \left(\sum_{j=1}^8 |T_{ij}|^2 n_{BE}(|\omega_j|) + \sum_{j=4}^8 |T_{ij}|^2 \right)$$
(3)

where $n_{\rm BE}(\omega) = 1/(e^{\omega/(k_BT)} - 1)$ is the Bose-Einstein distribution, and $V_{\rm 1BZ}$ is the volume of the (2D) first Brillouin Zone (1BZ). The reduction is often significant, but is always less than 70%, unless otherwise mentioned.

We can use SWT to find the magnon thermal Hall conductivity using an expression derived in Ref. 37 through linear response theory:

$$\frac{\kappa_{xy}}{k_B^2 T/\hbar} = -\frac{1/d_c}{(2\pi)^2} \int_{1\text{BZ}} d^2k \sum_{n=1}^4 \left(c_2(n_{\text{BE}}(\omega_n)) - \frac{\pi^2}{3} \right) \Omega_{nn}$$
$$\Omega_{nn} = \left(i\epsilon_{\mu\nu c} \sigma_3 \frac{\partial T_k^{\dagger}}{\partial k_{\mu}} \sigma_3 \frac{\partial T_k}{\partial k_{\nu}} \right)_{nn} \tag{4}$$

Name	J_1	K	Γ	Γ'	J_2	J_3	K_3	Ref.
HK	-4.6	7.0	_	_	_	_	_	13
КΓ	-	-6.8	9.5	_	_	_	_	8
$1(HK\Gamma J3)$	-1.7	-6.7	6.6	_	_	2.7	_	18
$2(HK\Gamma J3)$	-5.5	7.6	8.4	_	_	2.3	_	18
$3(HK\Gamma J3)$	-0.5	-5.0	2.5	_	_	0.5	_	19
$4(HK\Gamma J3)$	0.1	-5.5	7.6	_	_	0.1	_	22
$5(HK\Gamma J3)$	-0.3	-10.9	6.1	_	_	0.03	_	22
$6(HK\Gamma J3)$	-3.5	4.6	6.4	_	_	0.8	_	20
$(HK\Gamma\Gamma')$	-1	$^{-8}$	4	-0.95	_	_	_	20
ΗКΓ	-12	17	12	_	_	_	_	21
НКГКЗ	-1.8	-10.6	3.8	_	_	1.25	0.65	23
$HK\Gamma J2$	1.2	-5.6	1.0	_	0.3	0.3	_	46
7(HKFJ3)	-0.5	-5.0	2.5	_	_	0.1125	_	this paper

Table I. This information is primarily drawn from Table 1 in Ref. 50 with some models added. All values are in meV. For this analysis, we ignore the K_3 , J_2 values, which are there for completeness. Some groups propose different models within the same paper depending on the space group symmetry. For the HK and K Γ model, we add a small Γ and J_3 term, respectively to help the numerics.

where ϵ_{ijk} is the Levi-Civita symbol, d_c is the distance between 2D planes (assuming well separated layers), Ω_{nn} is the Berry curvature of the *n*th magnon band, and, with $\text{Li}_2(z)$ being the dilogarithm, $c_2(x) = (1+x)(\ln(1+x) - \ln(x))^2 - (\ln x)^2 - 2\text{Li}_2(-x)$.

We now specify the parameters in Eq. (1) to perform the calculation. We consider many different parameter sets as have been proposed in the literature, which are summarized in Table I. Though there is some experimental disagreement in the g-factor,^{27,46} susceptibility measurements^{3,47,48} give the paramagnetic moment to be $\gtrsim 2\mu_B$, with S = 1/2. We therefore fix g = 2.3 to get the correct order of magnitude as in Refs. 16, 18, and 19, and S = 1/2. According to Refs. 29 and 49, the interplanar distance is $d_c = 5.72$ Å.

The results of our SWT calculation are presented in Figs. 2. We have compared our code with the results of Refs. 39 and 43 to verify correctness. We have also plotted the data from Ref. 29. We do not plot $5(\text{HK}\Gamma\text{J}3)$ or HK $\Gamma\text{J}2$ since the spin wave solution is not stable (i.e. there are complex eigenvalues) above some critical field $\mu_0 H_c < 10$ T.

We see rather poor agreement between the models and the theory. Although most models do predict κ_{xy} of the correct sign, models HK, K Γ , 4(HK Γ J3) do not. Further notice that all models with K > 0 predict $\kappa_{xy} \gtrsim 0$.

To investigate why there is such a large discrepancy between the theoretical κ_{xy} and the data of Ref. 29, we try to find a large κ_{xy} in a minimal $J_1 - K - \Gamma - J_3$ model. It is worth noting that there is not much freedom. From Curie-Weiss temperature data $|K| \sim 100$ $K = 8.6 \text{ meV},^{3,51}$ which is similar to the estimate of Ref. 29 and is commonly seen in almost all of the models in Table I. Furthermore, it has been observed that the magnetic moments lie in the *ac* plane and make an angle of approximately 35° ,²⁶ which requires a particular Γ/K . Minimizing the classical energy assuming the moments are in the [xxz] direction, in the $J_1 - K - \Gamma - J_3$ model, we obtain an expression equivalent to one in Ref. 50:

$$\frac{\Gamma}{K} = \frac{2}{\sqrt{2}\tan(\theta) + 1 - \sqrt{2}\cot(\theta)},\tag{5}$$

where $z = \cos(\theta)$ and we assume x > 0. Two minima of the classical energy can be found with K < 0; $\Gamma/K \approx -0.82$ and K > 0; $\Gamma/K \approx 0.0065$.

Therefore, for two values of Γ/K with differing signs of K, we have only freedom in J_1 and J_3 . $J_1 < 0$ and $J_3 > 0$ help stabilize the zigzag order, so we place these constraints. In the K > 0 case with large enough $|J_1|$ to stabilize the zigzag order, we always found $\kappa_{xy} \gtrsim 0$, though a more thorough search of the parameter space might be needed.

In the K < 0 case, we instead start with the results of the meta-analysis of *ab initio* models from Ref. 19: $\Gamma/|K| \approx 0.5$ and $J_1/|K| \approx 0.1$. Fixing K = -5 meV as in their proposed model, we scan possible values of J_3 . We find that sufficiently low J_3 leads to large enough κ_{xy} to explain all but the lowest temperature point of Ref. 29. As a representative model, we find that $J_3 = 0.1125$ does well to reproduce the temperature data, as is shown in Fig. 2. Note that this model is representative not unique. Generically, with $-5 \text{ meV} \gtrsim K \gtrsim -8 \text{ meV}$, there is a value of $J_3 \gtrsim |K|/200$ that provides an orderof-magnitude fit to the data. For small J_3 , though, we find a large spin reduction with $\Delta S_0/S \sim 0.9$. Further, these models predict much larger κ_{xy} at T = 7 K than is measured in Ref. 29 as is seen in Fig. 2(b). Because of the proximity to the temperature at which long range order is lost, it is perhaps expected that whatever process is creating a large positive κ_{xy} above $T_N \sim 7$ K is beginning to affect the conductivity at T = 7 K. We are not claiming that our model explains the considerable existing data from other experiments (e.g. inelastic neutron scattering, electron spin resonance, etc.); the model is to demonstrate that, theoretically, the thermal Hall effect data could be described solely through SWT while still satisfying the conditions from Ref. 19.

To see why the κ_{xy} increased, we plot in Fig. 3 an example of the linear SWT bands and Berry curvature for a particular path through the 1BZ for our model 7(HKГJ3) vs. the similar model 3(HKГJ3) on which it is based. It is worth noting that even though the bands are qualitatively similar, the Berry curvature of the two models is quite different. For example, there is a large enhancement of the Berry curvature around the Y point.

The difference in magnitude of κ_{xy} can be understood as follows. The function $f(\omega_n) = -(c_2(n_{\text{BE}}(\omega_n)) - \pi^2/3)$ scaling the Berry curvature in the κ_{xy} integrand essentially serves as a high-pass filter with frequency $\omega_{Hi} = k_B T$. That is, if $\omega_n \ll k_B T$, $f(\omega_n) \sim T/\omega_n$ and if $\omega_n \gg k_B T$, $f(\omega_n) \sim \pi^2/3$. Since the sum of the Berry curvature integrated over the 1BZ is zero,³⁷ then



Figure 2. (Color online). We plot κ_{xy} as computed from Eq. (4) for the various models in Table I as a function of (a) temperature and (b) magnetic field. We also plot the data from Ref. 29 as blue dots. The inset of (a) shows a zoomed-out version of the same graph. In (b), models with $\kappa_{xy} \gtrsim 0$ were removed. Our model, which agrees well with the data in (a), does not agree with the data in (b). Since the data of Ref. 29 shows $\kappa_{xy} > 0$ at $T > T_N$ and excitations in the pure Kitaev model contribute to $\kappa_{xy} > 0$,²⁸ it is expected that at $T \approx T_N \approx 7$ K the contribution from just the magnons should be *below* the experimental data, as is true for our model. We do not plot 5(HK\GammaJ3) or HK\GammaJ2 since the zigzag spin wave solution becomes unstable for some critical magnetic field $\mu_0 H < 10$ T. Our proposed model, 7(HK\GammaJ3) has a large spin reduction $\Delta S_0/S \sim 0.9$ at T = 7 K.



Figure 3. (Color online). See Fig. 1 for the naming of the 1BZ points. We plot the SWT bands, ω_n , and the Berry curvature, Ω_{nn} , for the 3(HKГJ3) model (solid lines) and the 7(HKГJ3) model (dashed lines) at $\mu_0 H = 12$ T. The latter model has a much larger κ_{xy} as seen in Fig. 2 due to the fact that a) the gap between the lower two bands around the Y point is smaller and at lower energy and b) the gap between the green and red band is at a lower energy. Note that the Berry curvature is largest where the band gap is smallest, and the lower energy means that the effect of $-c_2(n_{\rm BE}(\omega_n))$ is more significant. (For Ω_{nn} , the path is changed to be slightly inside the 1BZ as opposed to being on the boundary, when applicable.)

 $k_BT \gtrsim \omega_n$ for κ_{xy} to be significant. Furthermore, we can see from Fig. 3 that the Berry curvature is largest when there is a small gap between the bands. To make the largest possible κ_{xy} , there must be small gaps in the bands at energies $\omega_n \lesssim k_BT$.

Our analysis tends to favor J_3 smaller than has been

proposed. As can be seen in plots in Refs. 18 and 46, decreasing J_3 tends to move closer to a transition out of the zigzag order. Since smaller band gaps lead to larger Berry curvature, this result makes sense as SWT would predict a magnetic ordering phase transition when a gap in two bands close: assuming the energies of the bands have the form $\omega_{\pm} = a \pm \sqrt{b}$, a transition occurs at b = 0 since b < 0 leads to instability of the spin wave solution.

These observations, however, call into question whether the magnon thermal Hall effect is indeed responsible or if the low temperature data is accurate. Denoting $x \equiv \omega_n/T$, we see that $f(\omega_n) = \pi^2/3 - e^{-x}(2+2x+x^2) + \mathcal{O}(e^{-2x})$ for $x \gg 1$. Since the mass of the magnons has been estimated to be ~ 2 meV from inelastic neutron scattering,⁸ κ_{xy} , as predicted from linear SWT, should be exponentially decreasing in 1/T at $T \ll 2$ meV ≈ 23 K. All the predictions of magnon thermal Hall effect shown in Fig. 2 show such a dependence but the data does not.

Another interpretation of our results is that the models in Table I are consistent with the data if the magnons are not the dominant source of the κ_{xy} at low temperatures. Phonons could in principle give a larger contribution than ordinarily observed. To elaborate, in past experiments, although $\kappa_{xy} \lesssim 10^{-4}$ W/K/m, the Hall angle was measured at $\mu_0 H \sim 10$ T to be $\kappa_{xy}/\kappa_{xx} \sim 1-5 \times 10^{-4}.^{35,36}$ Since $\kappa_{xx} \sim 2-6$ W/K/m for α -RuCl₃ at $T \lesssim 10$ K,^{9,29-31} we would then estimate $\kappa_{xy} \sim 3 \times 10^{-4} \kappa_{xx} \approx$ 1.2×10^{-3} W/K/m, which is of the correct order. Regarding contributions in the ordered phase by Kitaev-like excitations from fluctuations, note that since the pure Kitaev model predicts $\kappa_{xy} > 0$,²⁸ and the experimental data shows that κ_{xy} switches sign at $T \sim T_N$,²⁹ these observations would be inconsistent with having fluctuations into the Kitaev model explain the discrepancy between the predicted and observed low temperature κ_{xy} .

Finally, we should emphasize that our results focus on linear SWT. Since the condition $\Delta S_0/S \ll 1$ is not met in a rigorous sense, our results should be treated as a first-order estimate. A more detailed calculation using non-linear SWT keeping interaction terms between the magnons would be an important addition in the future, though such a calculation would be difficult with known techniques.

III. CONCLUSIONS

We have investigated whether the thermal hall conductivity data reported in Ref. 29 at low temperatures for the material α -RuCl₃ can be explained through linear SWT as an example of the magnon thermal Hall effect. Although for many of the effective Hamiltonians

- ¹ A. Kitaev, Annals of Physics **321**, 2 (2006), january Special Issue.
- ² G. Jackeli and G. Khaliullin, Phys. Rev. Lett. **102**, 017205 (2009).
- ³ S. Trebst, arXiv preprint arXiv:1701.07056 (2017).
- ⁴ K. W. Plumb, J. P. Clancy, L. J. Sandilands, V. V. Shankar, Y. F. Hu, K. S. Burch, H.-Y. Kee, and Y.-J. Kim, Phys. Rev. B **90**, 041112 (2014).
- ⁵ A. Koitzsch, C. Habenicht, E. Müller, M. Knupfer, B. Büchner, H. C. Kandpal, J. van den Brink, D. Nowak, A. Isaeva, and T. Doert, Phys. Rev. Lett. **117**, 126403 (2016).
- ⁶ S.-H. Baek, S.-H. Do, K.-Y. Choi, Y. S. Kwon, A. U. B. Wolter, S. Nishimoto, J. van den Brink, and B. Büchner, Phys. Rev. Lett. **119**, 037201 (2017).
- ⁷ A. Banerjee, J. Yan, J. Knolle, C. A. Bridges, M. B. Stone, M. D. Lumsden, D. G. Mandrus, D. A. Tennant, R. Moessner, and S. E. Nagler, Science **356**, 1055 (2017).
- ⁸ K. Ran, J. Wang, W. Wang, Z.-Y. Dong, X. Ren, S. Bao, S. Li, Z. Ma, Y. Gan, Y. Zhang, J. T. Park, G. Deng, S. Danilkin, S.-L. Yu, J.-X. Li, and J. Wen, Phys. Rev. Lett. **118**, 107203 (2017).
- ⁹ I. A. Leahy, C. A. Pocs, P. E. Siegfried, D. Graf, S.-H. Do, K.-Y. Choi, B. Normand, and M. Lee, Phys. Rev. Lett. **118**, 187203 (2017).
- ¹⁰ Y. Kasahara, T. Ohnishi, Y. Mizukami, O. Tanaka, S. Ma, K. Sugii, N. Kurita, H. Tanaka, J. Nasu, Y. Motome, T. Shibauchi, and Y. Matsuda, Nature **559**, 227 (2018).
- ¹¹ L. J. Sandilands, Y. Tian, K. W. Plumb, Y.-J. Kim, and K. S. Burch, Phys. Rev. Lett. **114**, 147201 (2015).
- ¹² J. Nasu, J. Knolle, D. L. Kovrizhin, Y. Motome, and R. Moessner, Nature Physics **12**, 912 (2016).
- ¹³ A. Banerjee, C. Bridges, J.-Q. Yan, A. Aczel, L. Li, M. Stone, G. Granroth, M. Lumsden, Y. Yiu, J. Knolle, *et al.*, Nature materials **15**, 733 (2016).
- ¹⁴ A. Little, L. Wu, P. Lampen-Kelley, A. Banerjee, S. Patankar, D. Rees, C. A. Bridges, J.-Q. Yan, D. Mandrus, S. E. Nagler, and J. Orenstein, Phys. Rev. Lett. **119**, 227201 (2017).

proposed in the literature, we find a non-zero κ_{xy} of the correct sign, none could satisfactorily explain the data. By modifying the model of Ref. 19, we were able to find large enough κ_{xy} to explain the 7 K $\gtrsim T \gtrsim 3$ K data showing that it is possible to explain the data via linear SWT. It also appears that K > 0 is not favored solely based on the sign of κ_{xy} . Taking the data at face value, these measurements provide a novel way to constrain a proposed effective Hamiltonian. In the future, more experimental data, better theoretical methods to incorporate both spin waves and Kitaev-model quasiparticles, and more careful treatment of the phonon thermal Hall effect would be useful in constraining proposed effective Hamiltonians.

Acknowledgements: The authors thank J. W. Orenstein for useful conversations. We acknowledge support from the Quantum Materials Program at LBNL, funded by the US Department of Energy under Contract No. DE-AC02-05CH11231.

- ¹⁵ Z. Wang, S. Reschke, D. Hüvonen, S.-H. Do, K.-Y. Choi, M. Gensch, U. Nagel, T. Rõ om, and A. Loidl, Phys. Rev. Lett. **119**, 227202 (2017).
- ¹⁶ L. Wu, A. Little, E. E. Aldape, D. Rees, E. Thewalt, P. Lampen-Kelley, A. Banerjee, C. A. Bridges, J. Yan, S. Patankar, *et al.*, arXiv preprint arXiv:1806.00855 (2018).
- ¹⁷ A. N. Ponomaryov, E. Schulze, J. Wosnitza, P. Lampen-Kelley, A. Banerjee, J.-Q. Yan, C. A. Bridges, D. G. Mandrus, S. E. Nagler, A. K. Kolezhuk, and S. A. Zvyagin, Phys. Rev. B **96**, 241107 (2017).
- ¹⁸ S. M. Winter, Y. Li, H. O. Jeschke, and R. Valentí, Phys. Rev. B **93**, 214431 (2016).
- ¹⁹ S. M. Winter, K. Riedl, P. A. Maksimov, A. L. Chernyshev, A. Honecker, and R. Valentí, Nature Communications 8, 1152 (2017).
- ²⁰ H.-S. Kim and H.-Y. Kee, Phys. Rev. B **93**, 155143 (2016).
- ²¹ H.-S. Kim, V. S. V., A. Catuneanu, and H.-Y. Kee, Phys. Rev. B **91**, 241110 (2015).
- ²² W. Wang, Z.-Y. Dong, S.-L. Yu, and J.-X. Li, Phys. Rev. B 96, 115103 (2017).
- ²³ Y. S. Hou, H. J. Xiang, and X. G. Gong, Phys. Rev. B 96, 054410 (2017).
- ²⁴ J. G. Rau, E. K.-H. Lee, and H.-Y. Kee, Phys. Rev. Lett. 112, 077204 (2014).
- ²⁵ V. M. Katukuri, S. Nishimoto, V. Yushankhai, A. Stoyanova, H. Kandpal, S. Choi, R. Coldea, I. Rousochatzakis, L. Hozoi, and J. van den Brink, New Journal of Physics **16**, 013056 (2014).
- ²⁶ H. B. Cao, A. Banerjee, J.-Q. Yan, C. A. Bridges, M. D. Lumsden, D. G. Mandrus, D. A. Tennant, B. C. Chakoumakos, and S. E. Nagler, Phys. Rev. B **93**, 134423 (2016).
- ²⁷ Y. Kubota, H. Tanaka, T. Ono, Y. Narumi, and K. Kindo, Phys. Rev. B **91**, 094422 (2015).
- ²⁸ J. Nasu, J. Yoshitake, and Y. Motome, Phys. Rev. Lett. 119, 127204 (2017).
- ²⁹ Y. Kasahara, K. Sugii, T. Ohnishi, M. Shimozawa, M. Yamashita, N. Kurita, H. Tanaka, J. Nasu, Y. Motome,

T. Shibauchi, and Y. Matsuda, Phys. Rev. Lett. **120**, 217205 (2018).

- ³⁰ R. Hentrich, M. Roslova, A. Isaeva, T. Doert, W. Brenig, B. Büchner, and C. Hess, arXiv preprint arXiv:1803.08162 (2018).
- ³¹ R. Hentrich, A. U. B. Wolter, X. Zotos, W. Brenig, D. Nowak, A. Isaeva, T. Doert, A. Banerjee, P. Lampen-Kelley, D. G. Mandrus, S. E. Nagler, J. Sears, Y.-J. Kim, B. Büchner, and C. Hess, Phys. Rev. Lett. **120**, 117204 (2018).
- ³² Y. Vinkler-Aviv and A. Rosch, Phys. Rev. X 8, 031032 (2018).
- ³³ M. Ye, G. B. Halász, L. Savary, and L. Balents, arXiv preprint arXiv:1805.10532 (2018).
- ³⁴ Z.-X. Liu and B. Normand, Phys. Rev. Lett. **120**, 187201 (2018).
- ³⁵ A. V. Inyushkin and A. N. Taldenkov, JETP Letters 86, 379 (2007).
- ³⁶ K. Sugii, M. Shimozawa, D. Watanabe, Y. Suzuki, M. Halim, M. Kimata, Y. Matsumoto, S. Nakatsuji, and M. Yamashita, Phys. Rev. Lett. **118**, 145902 (2017).
- ³⁷ R. Matsumoto, R. Shindou, and S. Murakami, Phys. Rev. B **89**, 054420 (2014).
- ³⁸ H. Katsura, N. Nagaosa, and P. A. Lee, Phys. Rev. Lett. 104, 066403 (2010).
- ³⁹ S. A. Owerre, Journal of Applied Physics **120**, 043903 (2016), https://doi.org/10.1063/1.4959815.
- ⁴⁰ S. A. Owerre, Phys. Rev. B **94**, 094405 (2016).

- ⁴¹ K. H. Lee, S. B. Chung, K. Park, and J.-G. Park, Phys. Rev. B **97**, 180401 (2018).
- ⁴² D. G. Joshi, Phys. Rev. B **98**, 060405 (2018).
- ⁴³ P. A. McClarty, X.-Y. Dong, M. Gohlke, J. G. Rau, F. Pollmann, R. Moessner, and K. Penc, Phys. Rev. B 98, 060404 (2018).
- ⁴⁴ D. H. Jones, Q. A. Pankhurst, and C. E. Johnson, Journal of Physics C: Solid State Physics **20**, 5149 (1987).
- ⁴⁵ See Supplemental Material at [URL will be inserted by publisher] for the derivation.
- ⁴⁶ R. Yadav, N. A. Bogdanov, V. M. Katukuri, S. Nishimoto, J. Van Den Brink, and L. Hozoi, Scientific reports 6, 37925 (2016).
- ⁴⁷ J. A. Sears, M. Songvilay, K. W. Plumb, J. P. Clancy, Y. Qiu, Y. Zhao, D. Parshall, and Y.-J. Kim, Phys. Rev. B **91**, 144420 (2015).
- ⁴⁸ M. Majumder, M. Schmidt, H. Rosner, A. A. Tsirlin, H. Yasuoka, and M. Baenitz, Phys. Rev. B **91**, 180401 (2015).
- ⁴⁹ R. D. Johnson, S. C. Williams, A. A. Haghighirad, J. Singleton, V. Zapf, P. Manuel, I. I. Mazin, Y. Li, H. O. Jeschke, R. Valentí, and R. Coldea, Phys. Rev. B **92**, 235119 (2015).
- ⁵⁰ L. Janssen, É. C. Andrade, and M. Vojta, Phys. Rev. B 96, 064430 (2017).
- ⁵¹ D. Hirobe, M. Sato, Y. Shiomi, H. Tanaka, and E. Saitoh, Phys. Rev. B **95**, 241112 (2017)