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# Novel high-pressure calcium carbonates

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#### Abstract

Calcium and magnesium carbonates are believed to be the host compounds for most of the oxidized carbon in the Earth's mantle. Here, using evolutionary crystal structure prediction method USPEX, we systematically explore the MgO-CO<sub>2</sub> and CaO-CO<sub>2</sub> systems at pressures ranging from 0 to 160 GPa to search for thermodynamically stable magnesium and calcium carbonates. While MgCO<sub>3</sub> is the only stable magnesium carbonate, three calcium carbonates are stable under pressure: well-known CaCO<sub>3</sub>, and newly predicted Ca<sub>3</sub>CO<sub>5</sub> and CaC<sub>2</sub>O<sub>5</sub>. Ca<sub>3</sub>CO<sub>5</sub> polymorphs are found to contain isolated orthocarbonate CO<sub>4</sub><sup>4-</sup> tetrahedra, and are stable at relatively low pressures (>11 GPa), whereas CaC<sub>2</sub>O<sub>5</sub> is stable above 33 GPa and its polymorphs feature polymeric motifs made of CO<sub>4</sub>-tetrahedra. Detailed analysis of chemical stability of CaCO<sub>3</sub>, Ca<sub>3</sub>CO<sub>5</sub> and CaC<sub>2</sub>O<sub>5</sub> in the environment typical of the Earth's lower mantle reveals that none of these compounds can exist in the Earth's lower mantle. We conclude that MgCO<sub>3</sub> is the main host of oxidized carbon throughout the lower mantle.

## Introduction

Behaviour of carbon in the Earth's mantle is important for the global carbon cycle. The generally accepted view is that mantle carbon exists in a number of reduced, neutral, and oxidized forms (i.e., Fe<sub>3</sub>C cementite, diamond, carbonates) <sup>[1]</sup>. Over the past few decades, magnesium and calcium carbonates (MgCO<sub>3</sub> and CaCO<sub>3</sub>) have received considerable attention because they are believed to be the host compounds for most of the oxidized carbon in the mantle <sup>[2-6]</sup>. Zero-temperature calculations of Pickard and Needs predicted that CaCO<sub>3</sub> will become more favorable than MgCO<sub>3</sub> at pressures above 100 GPa at mantle chemistry and therefore should be present in the lowermost matle <sup>[6]</sup>, but thermal effects could change this conclusion (and this is indeed the case). Moreover, all previous previous works assumed that compositions known at atmospheric pressure (CaCO<sub>3</sub>, MgCO<sub>3</sub>) as the only possibilities. Recent works <sup>[7-9]</sup> proved that chemistry is greatly altered by pressure: new unexpected compound appear so often that they are more of a rule than exception. It is, therefore, necessary to check for additional possible carbonates.

At mantle pressures, a series of phase transitions occur in MgCO<sub>3</sub> and CaCO<sub>3</sub>. Tables S1 and S2 list high-pressure forms of MgCO<sub>3</sub> and CaCO<sub>3</sub> predicted in previous works <sup>[3,5,6]</sup>. From zero pressure up to the pressure of the core-mantle boundary (136 GPa), both MgCO<sub>3</sub> and CaCO<sub>3</sub> will experience several interesting phase transitions. For example, it was predicted that polymorphs of CaCO<sub>3</sub> stable below 76 GPa <sup>[6]</sup> (or 75 GPa according to our calculations) feature CO<sub>3</sub>-triangles, while chains of CO<sub>4</sub> tetrahedra are present in the higher-pressure form of CaCO<sub>3</sub>.

At ambient pressure and temperature, carbon atom in CO<sub>2</sub> has *sp*-hybridization with linear geometry and 2-fold coordination, while in  $CO_3^{2-}$  it has *sp*<sup>2</sup>-hybridization resulting in planar triangular geometry and 3-fold coordination. sp<sup>3</sup>-hybridization (resulting in CO<sub>4</sub> tetrahedra) is unfavorable due to very small size of C<sup>4+</sup> cation compared to O<sup>2-</sup> atom: at realistic (very short) C-O distances, steric O-O repulsion would be too high. However, at high pressure, carbon prefers to be in the *sp*<sup>3</sup> state and behaves in many ways akin to silicon at normal pressure. High coordination gives volume advantage, which offsets the steric effects.

In this work, we assess the traditional assumption of CaCO<sub>3</sub> and MgCO<sub>3</sub> stoichiometries of calcium and magnesium carbonates. As will be detailed below, new carbonates are indeed predicted, and we examine their stability at different pressures and temperatures, and in the chemical environment of the lower mantle of the Earth.

#### Computational methodology

To search for stable magnesium and calcium carbonates at mantle pressures, we have explored the MgO-CO<sub>2</sub>, CaO-CO<sub>2</sub> Mg-C-O, Ca-C-O and MgO-CaO-CO<sub>2</sub> systems using the variable-composition evolutionary algorithm (EA) technique, as implemented in the USPEX code [10-13]. Here we performed EA crystal structure predictions in the pressure range from 0 to 160 GPa with up to 40 atoms in the primitive unit cell. The first generation of structures was created randomly. In all subsequent generations, structures were produced by heredity (40%), symmetric random generator (20%), softmutation (20%) and transmutation (20%) operators, and the best 60% of previous generation were used as parents to generate the next generation of structures. For all structures generated by USPEX, structure relaxations and total energy calculations were performed using the VASP code <sup>[14]</sup> in the framework of density functional theory <sup>[15]</sup>. In these calculations, we used the Perdew-Burke-Ernzerhof generalized gradient approximation functional (PBE-GGA <sup>[16]</sup>) to treat exchange-correlation, and the all-electron projector augmented wave (PAW  $^{[17]}$ ) method to describe core-valence interactions -  $3s^23p^64s^2$ ,  $3s^2$ ,  $2s^22p^2$  and 2s<sup>2</sup>2p<sup>4</sup> shells were treated as valence for Ca, Mg, C and O, respectively. The plane-wave kinetic energy cutoff of 600 eV and uniform k-point meshes for sampling the Brillouin zone with reciprocal-space resolution of  $2\pi \times 0.05$  Å<sup>-1</sup> were employed. Once stable compounds and structures were found, their properties were computed with denser *k*-points meshes, which had reciprocal-space resolution of  $2\pi \times 0.03$  Å<sup>-1</sup>.

#### **Results and discussions**

1. Phase stability at mantle pressures

We have performed crystal structure predictions at 0 GPa, 15 GPa, 20 GPa, 40 GPa, 60 GPa, 80 GPa 100 GPa and 160 GPa for the CaO-CO<sub>2</sub> system, at 0 GPa, 60 GPa, 100 GPa and 160 GPa for the MgO-CO<sub>2</sub> system, and 25 GPa, 50 GPa, 100 GPa, 130 GPa for ternary systems: Mg-C-O, Ca-C-O and MgO-CaO-CO<sub>2</sub>. At a given pressure, stable compounds were determined by the thermodynamic convex hull construction. As shown in Fig. 1(a), among all possible magnesium carbonates, only  $MgCO_3$  was found to be on the convex hull. This indicates that, except  $MgCO_3$ , no other magnesium carbonates can be thermodynamically stable at mantle pressures. At the same time, in the CaO-CO<sub>2</sub> system, besides the well-known CaCO<sub>3</sub>, we have discovered two hitherto unreported thermodynamically stable calcium carbonates  $Ca_3CO_5$  and  $CaC_2O_5$  (and one near-ground state compound  $Ca_2CO_4$ ), as shown in Fig. 1(b). Phonon calculations show that the two new stable calcium carbonate phases are dynamically stable, i.e. have no imaginary phonon frequencies, as shown in Fig. 2. Unexpectedly, we also found that three high-pressure forms of  $CaC_2O_5(Pc, Fdd2)$  and C2) and Cmcm-Ca<sub>3</sub>CO<sub>5</sub> can maintain dynamical stability at 0 GPa, see Fig. S1 in Supplementary Materials. This means that  $Ca_3CO_5$  and  $CaC_2O_5$  can be quenchable to ambient conditions at low temperature. Lattice parameters and enthalpies of formation of stable and metastable calcium carbonates were listed in Table S3 in Supplementary Materials. For the well-known  $CaCO_3$ , we have also listed available experimental and theoretical values for comparison <sup>[18-20]</sup>.

By calculating enthalpy-pressure curves for all stable compounds in the CaO-CO<sub>2</sub> and MgO-CO<sub>2</sub> systems (see Fig. S2 in Supplementary Materials), we have obtained their pressure-composition phase diagrams at pressures up to 160 GPa (see Fig. 3). As shown in Fig. 3, stable phases and phase transition pressures in MgO, CaO, CO<sub>2</sub>, MgCO<sub>3</sub> and CaCO<sub>3</sub> are in good agreement with previous studies <sup>[3,5,6, 21-23]</sup>. We note that, for CaCO<sub>3</sub>, Smith *et al.* recently proposed a new  $P2_1/c$ -II phase which is stable between 27.2 GPa and 37.5 GPa<sup>[24]</sup>. Unfortunately, we missed this structure in the present structural predictions; moreover, we are unable to include this structure in the present study because of the lack of its detailed structural parameters. Considering the slight energy difference between  $P2_1/c$ -II phase and  $P2_1/c$ -I of CaCO<sub>3</sub>, we should

say that  $P2_1/c$ -II phase of CaCO<sub>3</sub> will not affect our main conclusion too much. Phase transformations of Ca<sub>3</sub>CO<sub>5</sub> and CaC<sub>2</sub>O<sub>5</sub> are as follows: (1) For Ca<sub>3</sub>CO<sub>5</sub>, the orthorhombic *Cmcm* phase is predicted to become stable at 11 GPa, and to transform to tetragonal *I4/mcm* phase at 55 GPa; (2) For CaC<sub>2</sub>O<sub>5</sub>, above 33 GPa, four stable phases (low-pressure *Pc*, *Fdd2*, high-pressure *Pc* and *C2*) are predicted. Structural transformation from low-pressure monoclinic *Pc*-phase to orthorhombic *Fdd2*-phase occurs at 38 GPa, and from *Fdd2*-phase to high-pressure monoclinic *Pc*-phase at 72 GPa, and from high-pressure *Pc*-phase to monoclinic *C2*-phase at 82 GPa. Fig.4 shows the computed equations of state (EOS) of all stable calcium carbonates. One can see that pressure-induced phase transitions in CaCO<sub>3</sub> and Ca<sub>3</sub>CO<sub>5</sub>, but not in CaC<sub>2</sub>O<sub>5</sub> are accompanied by large volume discontinuities.

It should be possible to synthesize the newly predicted calcium carbonates  $(Ca_3CO_5 \text{ and } CaC_2O_5)$ . Several phases of MgCO<sub>3</sub> and CaCO<sub>3</sub>, previously predicted by our method and similar techniques, have already been confirmed by experiment, such as C2/m <sup>[5]</sup> and  $P2_1$  <sup>[5]</sup> phases in MgCO<sub>3</sub>, and  $P2_1/c$  <sup>[6,24]</sup>, *Pmmn* <sup>[3,20]</sup> and thermodynamically metastable *P*-1 <sup>[3,25]</sup> in CaCO<sub>3</sub>. Considering hundreds of papers where it was assumed that CaCO<sub>3</sub> is the only possible calcium carbonate, and the importance of calcium carbonate for fundamental chemistry and physics and the hot ongoing quest for sp<sup>3</sup> (tetrahedral) carbonates, we believe that our newly predicted calcium carbonates will stimulate experiments.

#### 2. Crystal structures of stable calcium carbonates

Crystal structures of the predicted stable and metastable calcium carbonates, visualized by VESTA package <sup>[26]</sup>, are shown in Fig. 5. At mantle pressures, crystal structures of CaCO<sub>3</sub> have been carefully studied before <sup>[3,6]</sup>. With the increase of pressure, CaCO<sub>3</sub> successively adopts five phases (calcite *R*-3*c*, aragonite *Pnma*, low-pressure *P*2<sub>1</sub>/*c*, post-aragonite *Pmmn*, and high-pressure *P*2<sub>1</sub>/*c*). As shown in Fig. 5(a)-(e), we can find that the former four phases contain triangular  $CO_3^{2^-}$  ions with  $sp^2$ -hybridization, while the fifth phase adopts pyroxene-type structure with chains of corner-linked  $CO_4^{4^-}$  tetrahedra above 75 GPa.

High-pressure phases of CaCO<sub>3</sub> (> 75 GPa) and MgCO<sub>3</sub> (> 83 GPa) contain  $CO_4^{4-}$  tetrahedra. In CaCO<sub>3</sub> above 75 GPa, we see chains of corner-sharing tetrahedra. In MgCO<sub>3</sub> above 83 GPa,  $CO_4^{4-}$  tetrahedra form  $C_3O_9^{6-}$  rings, and above 180 GPa form chains <sup>[5,6]</sup>. The differences between CaCO<sub>3</sub> and MgCO<sub>3</sub>, come from different sites of Ca and Mg. Ca<sup>2+</sup> is much larger than Mg<sup>2+</sup>, and requires the anion sublattice to have more open space to fit it.

On the Ca-rich side, our newly predicted phases include stable Cmcm-Ca<sub>3</sub>CO<sub>5</sub>, *I4/mcm*-Ca<sub>3</sub>CO<sub>5</sub> and metastable *P*2<sub>1</sub>/*m*-Ca<sub>2</sub>CO<sub>4</sub>, all of which contain isolated CO<sub>4</sub><sup>4-</sup> tetrahedra, as shown in Fig. 5(f)-(h). Our calculations prove that calcium orthocarbonate Ca<sub>3</sub>CO<sub>5</sub> can be stable at very low pressure (11 GPa), which is much lower than the formation pressure of orthocarbonic acid (314 GPa) <sup>[27]</sup>. Chemically, Ca<sub>3</sub>CO<sub>5</sub> can be represented as CaO•Ca<sub>2</sub>CO<sub>4</sub> with coexistence of both O<sup>2-</sup> and CO<sub>4</sub><sup>4-</sup> anions, while metastable Ca<sub>2</sub>CO<sub>4</sub> is a typical orthocarbonate. Stability of Ca<sub>3</sub>CO<sub>5</sub> at a surprisingly low pressure of 11 GPa means that the CO<sub>4</sub><sup>4-</sup> units may also be present at such pressures in carbonate melts. Phonon calculations show that Ca<sub>3</sub>CO<sub>5</sub> can be quenchable to ambient conditions at low temperatures.

Unlike Ca<sub>3</sub>CO<sub>5</sub> and Ca<sub>2</sub>CO<sub>4</sub>, with higher CO<sub>2</sub> content in CaC<sub>2</sub>O<sub>5</sub>, CO<sub>4</sub><sup>4-</sup> tetrahedra are connected into 2D-sheets in *Pc*-CaC<sub>2</sub>O<sub>5</sub> and *C*2-CaC<sub>2</sub>O<sub>5</sub>, and into a 3D-framework in *Fdd*2-CaC<sub>2</sub>O<sub>5</sub>, as shown in Fig. 5(i)-(k). Polymerization of CO<sub>3</sub><sup>2-</sup> can be described as a transformation from carbonyl (C=O) functional groups to ether bonds (C-O-C), as shown in Fig. 6(a). The charged oxygen atom bonded to carbon is a nucleophilic site, whereas carbon atoms in CO<sub>2</sub> molecules are positively charged. This makes an electrophilic reaction possible, with polymerized CO<sub>3</sub><sup>2-</sup> sharing one electron pair with CO<sub>2</sub> upon formation of a polymeric framework C<sub>2</sub>O<sub>5</sub><sup>2-</sup> anion, as shown in Fig. 6(b).

As discussed in previous work <sup>[28]</sup>, the oxygen sharing by  $CO_3^{2-}$  and  $CO_2$  can be described as oxo-Grotthuss mechanism. Here the formation of  $CaC_2O_5$  is its enhanced version. The combination of  $CO_3^{2-}$  and  $CO_2$  offsets the electrostatic repulsion between  $CO_3^{2-}$  anions. This is why the participation of  $CO_2$  greatly decreases the polymerization pressure (compared with CaCO<sub>3</sub>) from 75 to 33 GPa.

#### 3. Are $Ca_3CO_5$ and $CaC_2O_5$ possible in the Earth's lower mantle?

By means of the quasi-harmonic approximation (QHA), we first explored thermodynamic stability of all calcium carbonates in the pressure range from 80 GPa to 160 GPa and temperature at 2000 K. We note that, under such pressure and temperature conditions,  $CO_2$  is expected to be solid <sup>[29]</sup>; its Gibbs free energy can thus be accurately computed based on the crystalline structure. As shown in Fig. 7, temperature has a tiny effect on Gibbs free energy of formation of each calcium carbonate and all three calcium carbonates (Ca<sub>3</sub>CO<sub>5</sub>, CaCO<sub>3</sub> and CaC<sub>2</sub>O<sub>5</sub>) that are stable at 0 K are still stable at 2000 K, indicating that they will not decompose at the Earth's lower mantle conditions.

Then, we studied the chemical stability of stable Ca<sub>3</sub>CO<sub>5</sub>, CaC<sub>2</sub>O<sub>5</sub> and CaCO<sub>3</sub> and metastable Ca<sub>2</sub>CO<sub>4</sub> by exploring their possible reactions with compounds MgSiO<sub>3</sub> <sup>[30,31]</sup>, CaSiO<sub>3</sub> <sup>[32,33]</sup>, SiO<sub>2</sub> <sup>[34,35]</sup> and MgO - we remind that (Mg,Fe)SiO<sub>3</sub>, CaSiO<sub>3</sub> and (Mg,Fe)O are the dominant compounds of the Earth's lower mantle. These reactions are listed below:

$Ca_3CO_5+3SiO_2=3CaSiO_3+CO_2$	(1)
$Ca_3CO_5 + MgO = MgCO_3 + 3CaO$	(2)
$Ca_3CO_5+3MgSiO_3=3CaSiO_3+MgCO_3+2MgO$	(3)
Ca <sub>3</sub> CO <sub>5</sub> +MgSiO <sub>3</sub> +2SiO <sub>2</sub> =3CaSiO <sub>3</sub> +MgCO <sub>3</sub>	(4)
$Ca_2CO_4 + 2SiO_2 = 2CaSiO_3 + CO_2$	(5)
$Ca_2CO_4+MgO=MgCO_3+2CaO$	(6)
$Ca_2CO_4+2MgSiO_3=2CaSiO_3+MgCO_3+MgO$	(7)
$Ca_2CO_4 + MgSiO_3 + SiO_2 = 2CaSiO_3 + MgCO_3$	(8)
CaCO <sub>3</sub> +SiO <sub>2</sub> =CaSiO <sub>3</sub> +CO <sub>2</sub>	(9)
CaCO <sub>3</sub> +MgO=MgCO <sub>3</sub> +CaO	(10)
CaCO <sub>3</sub> +MgSiO <sub>3</sub> =CaSiO <sub>3</sub> +MgCO <sub>3</sub>	(11)
$CaC_2O_5 + SiO_2 = CaSiO_3 + 2CO_2$	(12)
$CaC_2O_5+2MgO=CaO+2MgCO_3$	(13)
CaC <sub>2</sub> O <sub>5</sub> +MgSiO <sub>3</sub> =CaSiO <sub>3</sub> +MgCO <sub>3</sub> +CO <sub>2</sub>	(14)
CaC <sub>2</sub> O <sub>5</sub> +MgSiO <sub>3</sub> +MgO=CaSiO <sub>3</sub> +2MgCO <sub>3</sub>	(15)
$CaC_2O_5+MgO=CaCO_3+MgCO_3$	(16)
CaC <sub>2</sub> O <sub>5</sub> +CaSiO <sub>3</sub> +MgO=2CaCO <sub>3</sub> +MgSiO <sub>3</sub>	(17)

For each chemical reaction, we used the most stable structures of each compound at relevant pressures and temperatures. It should be noted that there are some other phases suggested for some major mantle compounds <sup>[36]</sup>, but we would prefer not to take them as references in view of their thermodynamic metastabilities.

Fig. 8 shows the computed Gibbs free energy of each reaction in the pressure range from 80 GPa to 160 GPa and temperature of 2000 K. For the newly predicted Ca<sub>3</sub>CO<sub>5</sub> and CaC<sub>2</sub>O<sub>5</sub>, we found that they both cannot exist in the Earth's lower mantle. As shown in Fig. 8(a)&(b), Ca<sub>3</sub>CO<sub>5</sub> will always react with MgSiO<sub>3</sub> and SiO<sub>2</sub>; CaC<sub>2</sub>O<sub>5</sub> does not react with the main mantle compound MgSiO<sub>3</sub>, but will react with MgO, and also with a mixture of MgO and CaSiO<sub>3</sub>. For the well-known CaCO<sub>3</sub>, we found that CaCO<sub>3</sub> will not react with MgO, SiO<sub>2</sub>, and MgSiO<sub>3</sub> at zero temperature and pressure above 90 GPa (see Fig. S3 in Supplementary Materials), in agreement with previous results <sup>[5,6]</sup>. However, at 2000 K, there is a big change in the behavior of reaction (11) - CaCO<sub>3</sub> will always react with MgSiO<sub>3</sub> at pressures below 140 GPa, as shown in Fig. 8(c). Fig. 9 shows the phase diagram for reaction (11). It shows that CaCO<sub>3</sub> will never exist in the Earth's lower mantle, with large excess of MgSiO<sub>3</sub>. Therefore, we conclude that throughout the Earth's lower mantle polymorphs of MgCO<sub>3</sub> are the main hosts of oxidized carbon.

## Conclusions

In summary, evolutionary crystal structure predictions have been performed for  $MgO-CO_2$  and  $CaO-CO_2$  systems with the aim of exploring stable magnesium and calcium carbonates at pressures ranging from 0 GPa to 160 GPa. For the  $MgO-CO_2$  system, we found that there is only one stable magnesium carbonate  $MgCO_3$ . For the  $CaO-CO_2$  system, in addition to  $CaCO_3$ , we also discovered two hitherto unknown stable calcium carbonates  $Ca_3CO_5$ ,  $CaC_2O_5$  and one near-ground-state compound  $Ca_2CO_4$ .

Ca<sub>3</sub>CO<sub>5</sub> can be represented as CaO•Ca<sub>2</sub>CO<sub>4</sub>, and is a calcium orthocarbonate, and is stable at a remarkably low pressure of 11 GPa. This is the lowest-pressure material with CO<sub>4</sub>-tetrahedra. CaC<sub>2</sub>O<sub>5</sub> is the product of electrophilic reaction:  $CO_3^{2^2}+CO_2$  and an enhanced version of oxo-Grotthuss mechanism, which greatly decreases the polymerization pressure of  $CO_3^{2^2}$ : 33 GPa, compared to 75 GPa in CaCO<sub>3</sub>.

We have checked chemical stability of  $Ca_3CO_5$  and  $CaC_2O_5$  in the Earth's lower mantle environments by investigating possible chemical reactions involving MgCO<sub>3</sub>, CO<sub>2</sub>, MgSiO<sub>3</sub>, CaSiO<sub>3</sub>, SiO<sub>2</sub>, CaO, and MgO. Our results indicate that, while chemically very interesting, none of these new carbonates can be present in the Earth's mantle. Ca<sub>3</sub>CO<sub>5</sub> is so far the lowest pressure stable compound with CO<sub>4</sub> tetrahedra and suggests that already at pressures of ~11 GPa, carbonate melts can have a large concentration of CO<sub>4</sub> groups.

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#### Figure legend:

Fig. 1 Thermodynamic convex hulls for MgO-CO<sub>2</sub> and CaO-CO<sub>2</sub> systems at zero temperature and high pressure (with zero-point energy correction).

Fig. 2 Phonon dispersion curves of the newly predicted calcium carbonates at high pressures and zero temperature.

Fig. 3 Pressure-composition phase diagrams of MgO-CO<sub>2</sub> and CaO-CO<sub>2</sub> system at zero temperature.

Fig. 4 Equations of state of all stable calcium carbonates at zero temperature.

Fig. 5 Crystal structures of predicted stable and metastable calcium carbonates.

Fig. 6 Mechanism of (a) the polymerization of  $CO_3^{2-}$  and (b) the formation of  $CaC_2O_5$ .

Fig. 7 Thermodynamic convex hulls for the CaO-CO<sub>2</sub> system at 2000 K and various pressures.

Fig. 8 Gibbs free energies of mantle-relevant reactions as a function of pressure (at 2000 K).

Fig. 9 Relative stability of the  $MgCO_3 + CaSiO_3$  assemblage versus  $CaCO_3 + MgSiO_3$ .

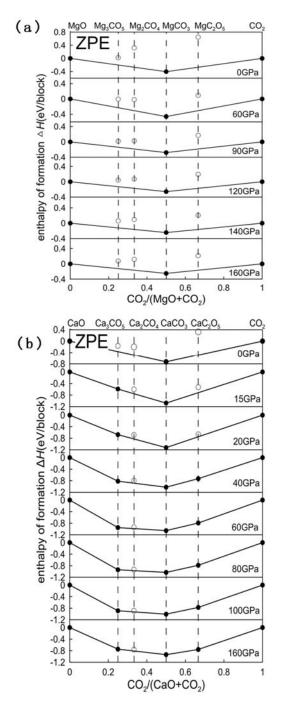


Fig. 1 Thermodynamic convex hulls for MgO-CO<sub>2</sub> and CaO-CO<sub>2</sub> systems at zero temperature and high pressure (with zero-point energy correction). Filled circles denote stable structures and open circles denote metastable structures. Enthalpies of formation from oxides are normalized to one

oxide unit.

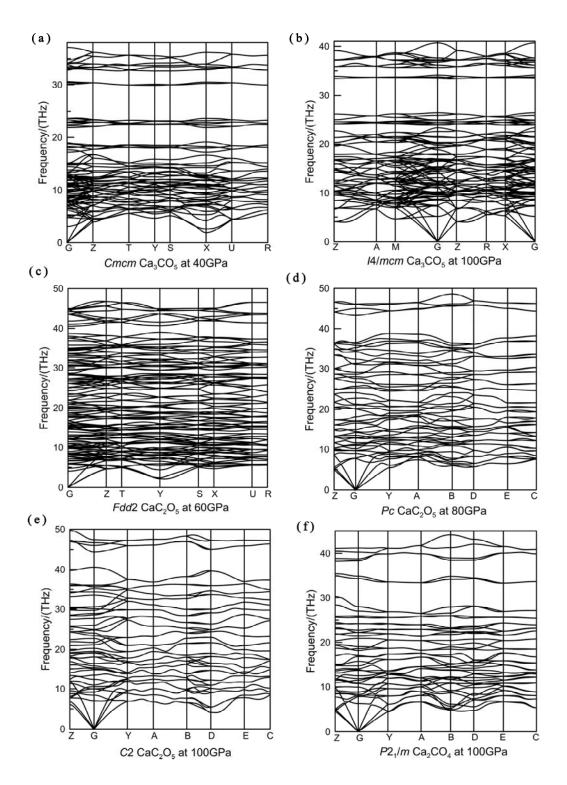


Fig. 2 Phonon dispersion curves of the newly predicted calcium carbonates at high pressures and zero temperature.

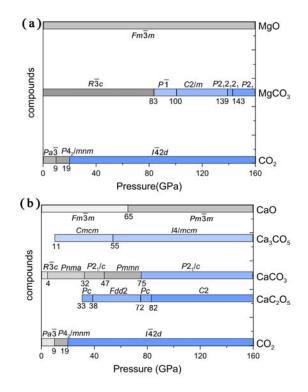


Fig. 3 Pressure-composition phase diagrams of MgO-CO2 and CaO-CO2 system at zero

temperature.

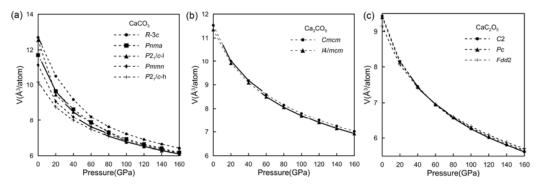


Fig. 4 Equations of state of all stable calcium carbonates at zero temperature. Solid lines denote equations of state of each phase in its region of stability.

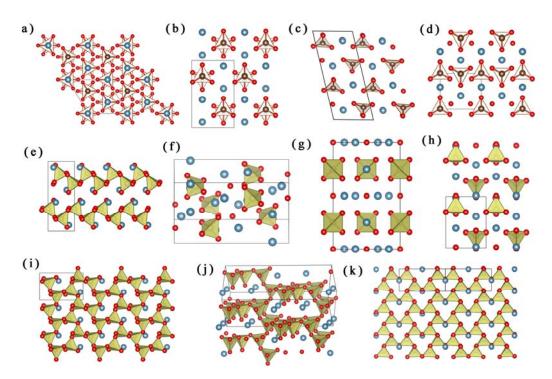


Fig. 5 Crystal structures of predicted stable and metastable calcium carbonates. (a) *R*-3*c* (calcite);
(b) *Pnma* (aragonite); (c) *P*2<sub>1</sub>/*c*-l; (d) *Pmmn* (post-aragonite); (e) *P*2<sub>1</sub>/*c*-h; (f) *Cmcm* Ca<sub>3</sub>CO<sub>5</sub>; (g) *I*4/*mcm* Ca<sub>3</sub>CO<sub>5</sub>; (h) *P*2<sub>1</sub>/*m* Ca<sub>2</sub>CO<sub>4</sub>; (i) *Pc* CaC<sub>2</sub>O<sub>5</sub>; (j) *Fdd*2 CaC<sub>2</sub>O<sub>5</sub>; (k) *C*2 CaC<sub>2</sub>O<sub>5</sub>

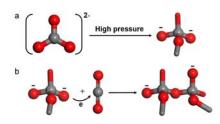


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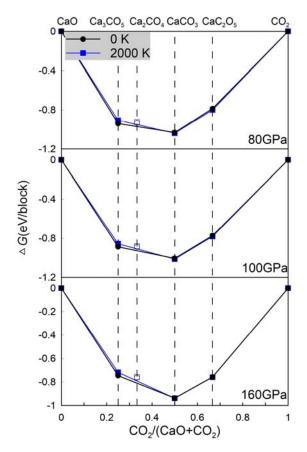


Fig. 7 Thermodynamic convex hulls for the CaO-CO<sub>2</sub> system at 2000 K and various pressures.Filled symbols denote stable structures, open symbols - metastable structures. Gibbs free energies of formation from oxides are normalized to one oxide unit.

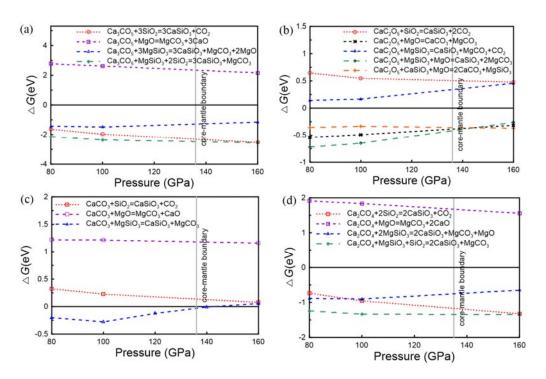


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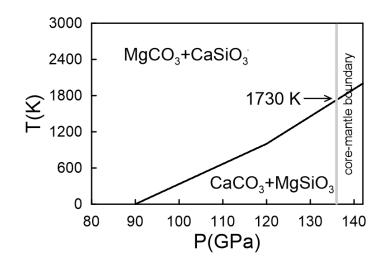


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