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Measuring the shock impedance mismatch between High Density Carbon and Deuterium at the National Ignition Facility

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Fine-grained diamond, or High Density Carbon (HDC), is being used as an ablator for inertial confinement fusion (ICF) research at the National Ignition Facility (NIF). Accurate equation of state (EOS) knowledge over a wide range of phase space is critical in the design and analysis of integrated ICF experiments. Here, we report shock and release measurements of the shock impedance mismatch between HDC and liquid deuterium conducted during shock-timing experiments having a first shock in the ablator ranging between 8 and 14 Mbar. Using ultrafast Doppler imaging velocimetry to track the leading shock front, we characterize the shock velocity discontinuity upon the arrival of the shock at the HDC/liquid deuterium interface. Comparing the experimental data with tabular EOS models used to simulate integrated ICF experiments indicates the need for an improved multi-phase EOS model for HDC in order to achieve significant increase in neutron yield in indirect-driven ICF implosions with HDC ablators.

12 efficient ablators for laser-driven dynamic compression thanks 13 to their low atomic numbers and relatively high densities. Di-14 15 amond is currently being used as an ablator in High Energy Density (HED) experiments¹⁻³ at the Omega Laser Facility as 16 ell as in integrated experiments for inertial confinement fu-17 sion (ICF) research^{4–8} at the National Ignition Facility (NIF), 18 where its advantages contributed to allow achieving record nu-19 clear fusion yield⁹. An extended knowledge of its equation of state (EOS) and transport properties at extreme densities, pres-21 sures and temperatures is important to ensure the accuracy and 22 redictability of the radiative-hydrodynamic simulations used 23 o model integrated ICF experiments. 24

Diamond, the least compressible material at ambient condi-25 tions, has been extensively studied using various static and dy-26 namic compression techniques. Its shock equation of state has 27 been documented^{10–15} to \sim 25 Mbar along the locus of shock 28 states (Hugoniot) and 50 Mbar with ramp compression^{16,17}. A 29 shock and release experimental study was also performed re-30 ently on single crystal and nanocrystalline diamond (NCD) 31 up to 26 Mbar¹⁵. In addition, the onset of significant opti-32 al reflectivity – *i.e.* metallic-like electrical conductivity – has 33 een evidenced to occur simultaneously with the completion 34 f melting along the Hugoniot^{18,19} near $U_S = 25 km/s$. 35

Due to extremely challenging fabrication constraints, NIF 36 capsule ablators are made of slightly underdense fine grained 37 (lateral crystallite dimension \sim 1-3 μm) polycrystalline 38 diamond^{20,21}, usually referred to – in ICF research – as High 39

Plastic polymers (CH), beryllium and carbon (diamond) are 40 Density Carbon (HDC) with densities ranging from 94 to 99% 41 of the density of single crystal diamond (3.51 q/cm^3). No ex-42 perimental data on the high pressure-high temperature proper-43 ties of this material are available, so that diamond EOS models 44 are used in pre- and post-shot radiation-hydrodynamic simu-45 lations. Note that the targets for the ramp compression ex-⁴⁶ periments of Ref.¹⁷ were not made of full density diamond. 47 Instead, the synthetic diamond material had a layered microstructure with alternating $0.35\mu m$ -thick layers of very fine 48 $_{49}$ 20 nm grains and layers of larger \sim 350 nm grains with 50 an average density of $\sim 3.25g/cm^{321}$. Similarly, the NCD ⁵¹ material studied in Ref.¹⁵ has a significantly lower density $_{52}$ (3.37 q/cm^3) than the HDC currently used in ICF implosion ⁵³ capsules ($\rho \sim 3.44 - 3.48g/cm^3$), because – in similar syn-54 thesis conditions – the microstructure of the obtained HDC 55 material is affected by the curvature of the spherical shells ⁵⁶ and cannot be replicated in planar *foils*.

Here, we report shock and release measurements between 58 HDC and cryogenic liquid deuterium using the shock-timing ⁵⁹ keyhole platform at the NIF^{22,23} (Fig. 1(a)). Our approach is similar to the work of Hamel *et al.* 24 aimed at measuring the 61 shock impedance mismatch between Glow Discharge Poly- $_{62}$ mer (GDP) and D_2 and provides new experimental bench-63 marks for the existing tabular equation of state models for ⁶⁴ HDC and D_2 , in particular on the shock and release behavior 65 in the vicinity of the melting line near 10 Mbar. The experi-66 mental data suggest that the current EOS tabular models need 67 to be improved to accurately describe the shock impedance ⁶⁸ mismatch between HDC and D_2 .



FIG. 1. Experimental concept. (a-c) Keyhole targets^{22,24-26} containing a High density carbon (HDC) capsules filled with liquid deuterium are indirectly driven using up to 192 beams at the National Ignition Facility (NIF) pointed at the top and bottom laser entrance holes. Most recent W-doped capsules have a multi-shell structure with the innermost $5\mu m$ made of undoped HDC with $\rho \sim 3.44 g/cm^3$, a W-doped layer, an undoped HDC outer layer with $\rho \sim 3.48g/cm^3$ and a $3\mu m$ thick nanocrystalline undoped HDC topcoat with $\rho \sim 3.30g/cm^3$. The density of the W-doped layer depends on the doping level, from $3.56g/cm^3$ with 0.16 at.% W to $3.60g/cm^3$ with 0.25 at.% W and $3.64g/cm^3$ with 0.33 at.% W. An accurately shaped, 5-7 ns long pulse launches a series of strong shock waves into the capsule. Using a gold cone inserted in the capsule and sealed by a transparent window, the interferometric Doppler velocimeter VISAR instrument²⁷ tracks the shock velocity history inside the ablator and the deuterium. (d) Observing the shock velocity inside the HDC capsules is challenging because of scattering due to its microstructure, absorption due to photoionization, strong Fresnel reflection at the interface, and weak reflection at the shock front below 25 km/s. (e) Once the shock has reached the fluid D_2 , it is easily tracked with VISAR to measure its shock velocity up to 150 km/s.

Shock timing experiments with cryogenic liquid deu- ${}_{94} n(D_2)|^2 \simeq 13\%$. In addition, the fine-grained structure of 69 70 71 72 73 74 75 with a gold cone inserted in the HDC spherical capsule and 101 the HDC layer. 76 sealed by a transparent window to allow a line of sight for the 77 wo channel, ultrafast, line-imaging Velocimetry Interferom-78 eter System for Any Reflector (VISAR) inside the capsule²⁷. 79 When shocked to sufficiently high pressure, diamond and deu-80 erium both become optically reflecting, allowing VISAR to 81 directly track the velocity of the shock front to establish the 82 ompression sequence. In addition, inserting one or more 83 mirrors in the gold cone allows one to gain insight into the 84 implosion symmetry by comparing the shock sequence along 85 multiple directions^{28,29}. 86

Observing the initial shock before it breaks out of the HDC 87 capsule and into the liquid deuterium fill presents several chal-88 lenges (Fig. 1). Due to the absence of an anti-reflection coat-89 ing, the large refractive index mismatch generates a strong 115 90 91 93

terium as a surrogate for the deuterium-tritium ice fusion ⁹⁵ HDC induces significant incoherent light scattering: thick fuel^{7,8,22,25,26} are essential to the ICF effort at NIF as they al- 96 fine-grained HDC layers appear opaque. HDC can also below us to accurately benchmark integrated implosion simula-97 come opaque due to the absorption of hot electrons³⁰ and/or xtions at the early time. In these experiments, precisely shaped, ⁹⁸ rays generated during the laser plasma interactions inside the 5-7 ns long drive-laser pulses are used to launch a series of ⁹⁹ hohlraum (blanking). These two effects combine to severely shocks into the ablator. We use the keyhole platform (Fig. 1) 100 attenuate the propagation of the VISAR laser probe deep into

> 102 Finally, the shock front reflectivity in HDC can be as low 103 as a few % or less at the lowest pressure range explored in 104 the NIF keyhole experiments. No direct measurement exists 105 but we can estimate that the reflectivity of shocked HDC will 106 rise quickly with increasing shock velocity similarly to the re-107 ported behavior for full density diamond^{18,19} when shocked 108 into the liquid phase. For reference, the reflectivity of di-109 amond increases from $\sim 1\%$ at 21 km/s to saturation near $_{110} \sim 30\%$ above 25 km/s. We can therefore expect to observe a similar rapid onset of shock reflectivity at the same or possibly ¹¹² slightly lower shock velocity, since being 1 to 5% underdense, HDC will reach slightly higher shock temperatures than dia-113 114 mond.

When VISAR is used to monitor the velocity of reflecting Freshel reflection of the VISAR laser probe at the HDC/D_2 ¹¹⁶ shock fronts in a given medium at rest, the observed fringe interface. With $n(HDC) \simeq 2.40$ and $n(D_2) \simeq 1.14$, the in- 117 shift is proportional to the apparent velocity $U_{S,app}$ of a ficterface reflectivity is $R = |n(HDC) - n(D_2)|^2 / |n(HDC) +$ 118 tional mirror moving into vacuum. It can be shown²⁷ that



FIG. 2. Experimental VISAR data and inferred shock velocity. (Top) Example VISAR record for a two-axis keyhole experiment with an HDC ablator. (Bottom) Corresponding shock velocity along the equator in the HDC (t < 3.1ns) and in the D_2 (t > 3.1ns) showing the sudden acceleration of the shock as it is transmitted into the deuterium (near 3.1 ns, vertical dotted line), including a weighted average of the velocity histories from VISAR channels Leg A and Leg B. (Inset) Raw data for the equator region before and after ghost fringe subtraction³.

¹¹⁹ $U_{S,app}$ is the product of the true shock velocity U_S and the index of refraction n of the unshocked medium at the VISAR 120 wavelength (here: 659.5 nm) so that $U_S = U_{S,app}/n$. 121

Fig. 2 shows an example VISAR record for one of these 122 experiments showing the shock velocity discontinuity due to 123 shock impedance mismatch between HDC and the liquid 124 deuterium. Optimized high-Z doping in the capsules success-125 fully mitigated blanking issues and allowed the collection of 126 VISAR patterns clearly exhibiting fringe motion due to reflec-127 tion at the shock front in HDC. Before 2.8ns, VISAR only 128 records the static *ghost fringes* due to the Fresnel reflection at ¹⁵⁰ ing procedure^{34,35} (see Fig. 3). As the HDC ablator is first 129 130 denly changes. This indicates the superposition of an addi-131 tional fringe pattern due to the reflection of the VISAR probe 132 at the shock front in the HDC ablator. A simple algorithm³ 133 can be used to eliminate the ghost fringes due to the Fresnel 134 eflection at the ablator/fuel interface and isolate the signal 135 carrying the information on the HDC shock velocity (Inset of 136 Fig. 2). Retrieving the fringe phase and using the two chan-137 nels with different sensitivities, we obtain the shock velocity 138 history $U_S(t)$ (Bottom of Fig. 2) showing the sudden accelera-139 tion of the shock as it is transmitted into the deuterium. Linear 140 fits of $U_S(t)$ over $\sim 200 ps$ before and after breakout provide 141 142 143 in leg A – yields $U_S(HDC)$ and $U_S(D_2)$. 144

145 146 147 148 ¹⁴⁹ mismatch between these two media by an impedance match-¹⁷⁰ range.



FIG. 3. Graphical construction of the impedance matching procedure. Calculated Hugoniot curves are shown in $U_S - u_p$ and $P - u_p$ spaces (solid lines), together with calculated isentropic release curves (dotted lines). For a given incident velocity $U_S(HDC)$, one can determine - for a given HDC EOS - the incident particle velocity in HDC $u_p(HDC)$ using the top panel, then follow a vertical line at constant u_p to the lower panel to determine the incident shock pressure P(HDC). Then following a release curve until it crosses the D_2 Hugoniot, one determines the approximate transmitted shock pressure $P(D_2)$ and particle velocity $u_p(D_2)$. Following again a vertical, constant u_p line back up into the upper panel until the intersection with the deuterium $U_S - u_p$ Hugoniot line, one obtains an approximate transmitted shock velocity $U_S(D_2)$. Dashdotted and dashed lines show that for a given incident shock velocity Us(HDC) = 25 km/s, LLNL 65 and 9061 models predict two very different transmitted shock velocities $Us(D_2)$ resp. near 28 and 31 km/s.

the HDC/D_2 interface. Near 2.8ns the fringe contrast sud- 151 shock-compressed and then releases to ensure the continuity ¹⁵² of the pressure and particle velocity at the interface^{34,35}, the ¹⁵³ predicted $U_S(HDC) - U_S(D_2)$ for a given model therefore 154 depends on both the shock and release behaviors as illustrated 155 in Fig. 3.

Fig. 4 shows most high-pressure experimental shock data 156 ¹⁵⁷ previously collected at large laser^{10–12} and pulse-power¹³ fa-158 cilities along with calculated theoretical Hugoniot curves. Be-159 low 500 GPa (5 Mbar) most current tabular EOS models pre-160 dict very similar shock compressibility, likely because they ¹⁶¹ were fit to the same data in this range, well below melting. 162 Above 500 GPa, as shown on Fig. 4, theoretical Hugoniot the velocities at breakout for both leg A and B. Weighted av- 163 curves rapidly span a much broader region of the pressureerages of these velocities – accounting for higher sensitivity 164 density space. Near 12 Mbar, LLNL 9069 (similar to LANL 165 7834) and LLNL 65/64 predict a much stiffer response, with Using the Rankine-Hugoniot equations and tabular EOS $_{166}$ densities near 6.8 q/cm^3 compared to 7.4 q/cm^3 according to models for HDC and D₂, one can compute the expected 167 LLNL 9061 and LANL Sesame 7830. The high accuracy data change in shock velocity at the interface between the HDC 168 from Hicks et al. 12 and Knudson et al. 13 are in agreement and ablator and the liquid D_2 resulting from the shock impedance 169 suggest that the stiffer models are not valid in the 6-20 Mbar



FIG. 4. Diamond Hugoniot: experimental data and tabular EOS models. Grey shaded area indicates the pressure range of the shock states explored in the present experiments. Data from Brygoo et al.¹¹ and Hicks et al.¹² measured relative to a quartz standard, and the data from Nagao et al.¹⁰ and Bradley et al.¹² using an Al standard were reanalyzed using updated quartz and Al Hugoniot models³¹ Knudson et al.¹³ used a Cu standard. Data from Ref.¹⁵ collected with a quartz standard are also included.



FIG. 5. Shock velocity jump at the interface between the HDCablator and the liquid D_2 fuel surrogate. Experimental data (blue squares) assuming 2.40 as the refractive index of unshocked HDC. Calculated velocity jumps for various HDC EOS models assuming LLNL 1014 (based on Kerley's 2003 model³², solid lines) or the included.

171 ¹⁷² the NIF are shown in Fig. 5 (blue squares). These experiments ²³⁰ at appeasing agreement with both the melt temperature and 173 provide 28 sets of velocities tracked either along the equato- 231 the specific heat inferred in Ref.¹⁹ do not lift the discrepancy. 174 rial or the polar direction, together with the predicted velocity 232 It is therefore possible that the *ab-initio* calculations, as per-175 jump between HDC and D_2 for several different EOS mod- 233 formed, used to constrain 9061³⁶, lack accuracy in the low- $_{176}$ els. Strong variations between the model predictions are no- $_{234}$ ρ , high-T regimes of importance for the impedance matching

177 ticeable. Fig. 5 also indicates that no EOS model matches the experimental data over the explored velocity range between ~ 23 to ~ 27 km/s corresponding to ~ 8 Mbar to 14 Mbar. 179 In particular, a strong discrepancy is found between the data 180 and the most detailed and recent EOS model to date³⁶ (LLNL 9061). In contrast, the LLNL 9069 model seems to match the 182 data up to 25 km/s due to a cancellation of errors, as it predicts 183 a stiffer principal Hugoniot than that of experiment. 184

An important parameter of the calculations is the descrip-185 tion of the shock compressibility of deuterium. As with most 186 simulations of ICF research experiments at NIF, we used pri-187 marily LLNL 1014 (based on G. Kerley's 2003 model³²). Since several experiments suggest that the compressibility of 189 shocked deuterium in the regime of interest for ICF implo-190 sions might be slightly higher than predicted^{35,37–39} by LLNL 1014, we also computed the predicted shock velocity jumps 192 using a recent model constructed using density functional the-193 ory based molecular dynamics simulations³³ (dashed lines on 194 ¹⁹⁵ Fig. 5) that is found to correctly capture the $P - \rho$ behavior of shocked D_2 . However, this does not bring the data in agree-196 ment with LLNL 9061.

We verified that the discrepancy in shock velocity jump was 198 ¹⁹⁹ not due to uncertainty in the refractive index n of the HDC as the primary observable in our experiment in the apparent 200 velocity $U_{S,app}$ from which we extract the shock velocity U_S 201 using $U_S = U_{S,app}/n$ (see Supplementary Material). 202

Unexpected significant preheat can also be ruled out since 203 calculations considering either preheated deuterium, or pre-204 heated HDC as well as the case of preheat in the two materi-205 als further enhances the discrepancy in impedance mismatch 206 shown on Fig. 5. Note that no early fringe motion is detected 207 on the VISAR records, suggesting the absence of significant preheat near the ablator/fuel interface. Fig. 7 also shows no 209 clear correlation between the level of W-doping and the dis-210 agreement between the data and the 9061 model. As one could expect that increasing the W-doping level should contribute to 212 213 mitigate photoionization and/or preheat effects, the absence of correlation suggests that either preheat/photoionization ef-214 fects are negligible or than the doping variations are too small 215 to matter. 216

The large discrepancy between the experimental $U_S(D_2)$ – 21 $U_S(HDC)$ values and the predicted curves with LLNL 9061 218 ²¹⁹ suggests that this model, which accurately described the shock 220 compressibility fails to correctly capture the release behavior of HDC. To investigate whether this discrepancy was 221 222 caused by the parametrization of the EOS, we computed the ²²³ impedance matching $U_S(D_2) - U_S(HDC)$ curves directly us-DFT-MD based CEA table³³ (dashed lines) as the deuterium EOS 224 ing the pressure and internal energy values obtained from the with an initial density $\rho_0 = 168mg/cm^3$. Data from Ref.¹⁵ are also ²²⁵ DFT-MD simulations that were used³⁶ to fit the broad range 226 analytical free energy model for the LLNL 9061 table. This 227 does not provide a better agreement between the experimen-228 tal shock velocity jump and the predicted ones. Similarly, The results from 15 different shock-timing experiments at 229 slight modifications of the liquid free energy of 9061 aimed 235 with deuterium.

236 237 238 239 240 the melting line causing non trivial release behavior includ- 296 ments are performed. 241 ing kinetic effects. Indeed, an increase in the scatter of the 297 242 243 244 245 246 247 space and seed hydrodynamic instabilities at the ablator/fuel 248 interface. Further investigations also include testing the influ-249 ence of the exchange correlation functional on the DFT-MD 303 250 calculations, in particular at the lowest densities for which no 304 251 experimental data are available. 252

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Appendix A: Experimental details

256 257 258 259 260 261 and 34.065 km/s/fringe, while others used 8.006 mm and 262 263 264 accounting for the discrepancies between the two legs the ve-265 locity resolution for most shots was $\sim 5\%$ of a fringe for 266 the most resolving etalon that is $\sim 0.28 km/s$ in HDC and 267 $0.60 km/s \text{ in } D_2.$ 268

At the conditions of the experiments, T = 21.5K and 269 $P \sim 800$ Torr, the refractive index of deuterium at 660 nm $_{_{322}}$ 270 271 $168mg/cm^{3}$. 272

273 274 275 277 278 279 280 281 density41-44. 282

283 284 285 286 288 Fig. 5. A quick estimate brings the *necessary* index to be 340 models, for states in the liquid (P ; 1000 GPa; see Fig. 3), 289 around 2.50.

5

Note that the cryogenic temperature required for the liquid 290 Previous shock temperature measurements¹⁹ showed that ²⁹¹ deuterium does not affect the HDC refractive index. The temmelting starts near 7-8 Mbar and is complete near 10 Mbar 292 perature dependence of the far-infrared refractive index of dialong the diamond Hugoniot. Since theoretical simulations⁴⁰ 293 amond has been measured both optically⁴⁵ and electrically⁴⁶: suggest a negative Clapeyron slope near 7-10 Mbar, release 294 the index increases very slowly with temperature, by less than isentropes from shock states right above 10 Mbar might cross 295 0.001 between 21.5 K and 300 K where most optical measure-

Modification of the refractive index of the unshocked mate- $U_S(D_2) - U_S(HDC)$ data, perhaps even suggesting a break ²⁹⁸ rial ahead of the shock front by ablation generated x-rays have in the slope appears on Fig. 5 near $U_S(HDC) = 25km/s$ (10 ²⁹⁹ been reported but found to cause a small (0.001) decrease of Mbar). Such a scenario could have dramatic implications in 300 diamond real part of the refractive index⁴⁷ so that it is unlikely the ICF implosion context as it could significantly affect the 301 that this could be the cause for the discrepancy in shock veoughness of the shock front both in physical and in velocity 302 locity jump between the experiment and the models (Fig. 5).

Appendix B: Re-analysis or previous Diamond Hugoniot equation of state data

Most high pressure Hugoniot equation of state measure-306 ments for diamond were performed relative to a standard using an impedance-matching procedure³⁴. As the knowledge 307 about standards is improved, re-analyzing existing data is necessary and provides more accurate and precise data than the 309 ³¹⁰ original analysis. All the data from Nagao *et al.*¹⁰, and four The velocity resolution of the VISAR for the HDC/ D_2 311 shots reported in Hicks *et al.*¹² (from an unpublished study impedance match event is somehow limited because the fused 312 by Bradley et al.) used an aluminum standard. Reanalyzed silica etalons were selected to allow tracking the subsequent 313 data using the latest fit to high precision aluminum Hugoniot shocks in the deuterium fluid traveling much faster, up to 150 314 data^{31,34} are reported in Fig. 4 and Table IV. The bulk of km/s. Most shots used 4.5799 mm on Leg A and 1.8332 mm ³¹⁵ the data from Hicks *et al.*¹² and the measurements from on Leg B giving a vacuum velocity per fringe (vvpf) of 13.635 316 Brygoo et al. 11 were performed with a quartz standard and 317 therefore reanalyzed using the constant Gruneisen re-shock 3.2196 mm yielding vvpf of 7.8001 and 19.296 km/s/fringe. ³¹⁸ model in Hicks *et al.*¹² and a fit of all available data for Taking into account the ability to resolve the fringe phase and ³¹⁹ quartz shocked into the liquid phase for the quartz Hugoniot³⁵. 320

Appendix C: Current tabular equations of state for diamond

The available tabular EOSs for modeling ICF im $is^{24} n(D_2) = 1.136$ and the initial density is $\rho_0(D_2) = 323$ plosions with HDC ablators are designed to describe 324 full density diamond, including the Lawrence Livermore The shock velocity measurements reported here are con- 325 National Laboratory (LLNL) multiphase model # 9061 ducted when the shock wave is in the innermost HDC layer 326 (Benedict/Correa/Sterne- 2014) (see Table I). This EOS is with density near $3.44g/cm^3$. No published data exist on 327 based entirely on DFT and Quantum Monte Carlo *ab-initio* the refractive index on this fine-grained HDC capsule mate- 328 electronic structure calculations. The exchange-correlation rial but preliminary characterization measurements on $8\mu m_{329}$ functional used in the DFT (PBE) renders the equilibrium denthick layers of undoped HDC grown in the same conditions $_{330}$ sity at 300 K for the model, 3.456 g/cm^3 , to be slightly less than the capsules used in the experiments reported here indi-³³¹ than the experimental value for full-density diamond (3.51 cated $n \simeq 2.35 - 2.40$ at 660 nm, in agreement with previ- $_{332} q/cm^3$). It is therefore ~ 2.5 % higher than the measured ously published data for nanocrystalline diamond with similar 333 value for HDC (3.37 g/cm^3). However, when we use the 334 9061 model in this work, we force the initial density in the The analysis reported on all figures assumed n = 2.40. 335 principal Hugoniot calculation to be $\rho_0 = 3.37 q/cm^3$, pro-This appears as a conservative choice because as the actual 336 ducing the red curves in Figs. 3 and 4 (and likewise for the shock velocity is given by: $U_S = U_{app}/n$, one would need 337 further analysis pertaining to the impedance matching with to assume a higher index for HDC in order to reconcile the 338 deuterium, shown in Fig. 5). The noticeable difference bepredictions using LLNL 9061 and the experimental data on 339 tween the principal Hugoniot of 9061 and that of the other $_{341}$ is not due to these small differences in the tables' ρ_0 values, ³⁴² but is rather due to the specific construction of the liquid free ³⁴³ energy, described in Ref.³⁶.

Previous diamond EOS tables include Los Alamos Na-344 tional Laboratory (LANL) Sesame 7830 (B. Bennett - 1982) 345 and 7834 (S. Crockett -2007), and LLNL 65 (LLNL-1999), 346 64 (LLNL-2007). LLNL 9069 is a QEOS table con-347 structed to match the multiphase table 9067 (Ref.⁴⁸) (Cor-348 rea/Benedict/Young - 2007). LANL Sesame 7834 was fit to 349 DFT-MD results and experimental Hugoniot data¹³ but no at-350 tempt to accurately fit the melt region was made, in contrast to 351 the explicitly multiphase LLNL 9067 and 9061 tables. San-352 dia National Laboratory (SNL) Sesame 7830 is a multicompo-353 nent and multiphase table for carbon (graphite is the reference 354 state) build with the code Panda (Kerley/Chhabildas)⁴⁹. 356

Appendix D: Shock and release experiments between HDC and quartz

We also report on Fig. 6 and Table III the result of a pla-359 360 nar, impedance-matching event between a transparent slab of chemical vapor deposition (CVD) grown HDC and a z-361 cut α -quartz plate. This datum was obtained using data col-362 ected at the NIF to characterize drive conditions for proton-363 stopping power experiments⁵⁰. We used $\rho_0 = 3.493 g/cm^3$, 364 n[HDC](660nm) = 2.41 (Ref.⁵¹), n[Quartz](660nm) =365 1.517 (Ref.⁵². Using the LLNL 9061 model, the incident 366 shock pressure in the CVD diamond slab was about 9 Mbar. 367

In addition to one new experimental datum for the shock 368 impedance mismatch between CVD diamond and quartz, we 369 report on Fig. 6 the prediction using various EOS models for 370 HDC and a fit to high precision experimental data³¹. Simi-371 lar to the observation for the single shock states above, only 372 373 LLNL 9061 and LANL Sesame 7830 models match well the experimental datum. Since the pressure drop at the dia-374 mond/quartz is only $\sim 30\%$, the release wave into the dia-375 mond ablator does not lower the density too much and it is 376 not too surprising that the model which describes well single 377 shock states manages to capture this behavior as well. 378



FIG. 6. Shock velocity jump at the interface between a CVD diamond slab and a quartz plate. Experimental data and predictions using various models for HDC and a fit to experimental data for the quartz Hugoniot³¹. Data from Ref.¹⁵ are also included.



FIG. 7. **Influence of W-doping on the shock impedance mismatch**. Same data than reported in Fig.5 with the level of W-doping as a color code. One could expect that increasing the W-doping level should contribute to mitigate photoionization and/or preheat effects. No clear correlation appears between the level of doping and the disagreement between the data and the 9061 model.

TABLE I. Diamond and HDC tabular equation of state models commonly used for ICF research

Table	ρ_0	Date	Multiphase	Comments
LANL Sesame 7830	$3.51g/cm^{3}$	1982	No	Very coarse temperature grid.
LLNL 65	$3.51g/cm^{3}$	1999	No	Poor fit to cold curve data.
SNL Sesame 7830	$2.20g/cm^{3}$	2001	Yes	Multicomponent and multiphase. Graphite initial state.
LANL Sesame 7834	$3.51g/cm^{3}$	2007	No	Fit to DFT-MD and Hugoniot ¹³ data. No attempt to accurately fit the melt region.
LLNL 9067	$3.51g/cm^{3}$	2007	Yes	Total functions P_t , E_t based on DFT-MD.
				Cold and thermal functions P_c , P_e and P_i based on QEOS.
				P_t does not equal $P_c + P_e + P_i$.
LLNL 64	$3.51g/cm^{3}$	2007	No	Post-dates L9067, but uses cruder EOS models
				for solid/melt region. Same as LLNL 65 along the Hugoniot.
LLNL 9061	$3.37g/cm^{3}$	2014	Yes	Best available EOS based on DFT-MD ³⁶ and a range of experimental data.
LLNL 9069	$3.42g/cm^{3}$	2015	Yes	Same as the LLNL 9067 cold, ion-thermal and electron-thermal terms.

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- D. E. Fratanduono, T. R. Boehly, M. a. Barrios, D. D. Meyerhofer, 441 380
- J. H. Eggert, R. F. Smith, D. G. Hicks, P. M. Celliers, D. G. Braun, 442 381 and G. W. Collins, J. Appl. Phys. 109, 123521 (2011). 382 443
- 2 F. Coppari, R. F. Smith, J. H. Eggert, J. Wang, J. R. Rygg, a. Laz-383
- icki, J. a. Hawreliak, G. W. Collins, and T. S. Duffy, Nat. Geosci. 384 445 385 6, 926 (2013). 446
- 3 M. Millot, N. Dubrovinskaia, A. Cernok, S. Blaha, L. Dubrovin-386 447 sky, D. G. Braun, P. M. Celliers, G. W. Collins, J. H. Eggert, and 387 448 R. Jeanloz, Science (80-.). 347, 418 (2015). 449 388
- 4 A. J. MacKinnon, N. B. Meezan, J. S. Ross, S. Le Pape, 450 389 L. Berzak Hopkins, L. Divol, D. Ho, J. Milovich, A. Pak, J. Ralph, 451 390
- T. Döppner, P. K. Patel, C. Thomas, R. Tommasini, S. Haan, 452 391
- A. G. MacPhee, J. McNaney, J. Caggiano, R. Hatarik, R. Bionta, 453 392
- T. Ma, B. Spears, J. R. Rygg, L. R. Benedetti, R. P. J. Town, 454 393
- D. K. Bradley, E. L. Dewald, D. Fittinghoff, O. S. Jones, H. R. 455 394
- Robey, J. D. Moody, S. Khan, D. A. Callahan, A. Hamza, J. Bi- 456 395
- ener, P. M. Celliers, D. G. Braun, D. J. Erskine, S. T. Prisbrey, R. J. 457 396
- Wallace, B. Kozioziemski, R. Dvlla-Spears, J. Sater, G. Collins, 458 397
- E. Storm, W. Hsing, O. Landen, J. L. Atherton, J. D. Lindl, M. J. 459 398
- Edwards, J. A. Frenje, M. Gatu-Johnson, C. K. Li, R. Petrasso, 460 399
- H. Rinderknecht, M. Rosenberg, F. H. Séguin, A. Zylstra, J. P. 461 400
- Knauer, G. Grim, N. Guler, F. Merrill, R. Olson, G. A. Kyrala, 462 401
- J. D. Kilkenny, A. Nikroo, K. Moreno, D. E. Hoover, C. Wild, 463 402
- and E. Werner, Phys. Plasmas 21, 056318 (2014). 403
- J. S. Ross, D. Ho, J. Milovich, T. Döppner, J. McNaney, A. G. 465 404 MacPhee, A. Hamza, J. Biener, H. F. Robey, E. L. Dewald, 466 405 R. Tommasini, L. Divol, S. Le Pape, L. B. Hopkins, P. M. Cel- 467 406
- liers, O. Landen, N. B. Meezan, and A. J. Mackinnon, Phys. Rev. 468 407 E **91**, 021101 (2015). 408
- L. F. Berzak Hopkins, N. B. Meezan, S. Le Pape, L. Divol, A. J. 470
- 409 Mackinnon, D. D. Ho, M. Hohenberger, O. S. Jones, G. Kyrala, 471 410
- J. L. Milovich, A. Pak, J. E. Ralph, J. S. Ross, L. R. Benedetti, 472 411
- J. Biener, R. Bionta, E. Bond, D. Bradley, J. Caggiano, D. Calla- 473 412
- han, C. Cerjan, J. Church, D. Clark, T. Döppner, R. Dylla-Spears, 474 413
- M. Eckart, D. Edgell, J. Field, D. N. Fittinghoff, M. Gatu Johnson, 475 414
- G. Grim, N. Guler, S. Haan, A. Hamza, E. P. Hartouni, R. Hatarik, 476 415
- H. W. Herrmann, D. Hinkel, D. Hoover, H. Huang, N. Izumi, 477 416
- S. Khan, B. Kozioziemski, J. Kroll, T. Ma, A. MacPhee, J. Mc- 478 417
- Naney, F. Merrill, J. Moody, A. Nikroo, P. Patel, H. F. Robey, J. R. 479 418
- Rygg, J. Sater, D. Sayre, M. Schneider, S. Sepke, M. Stadermann, 480 419
- W. Stoeffl, C. Thomas, R. P. J. Town, P. L. Volegov, C. Wild, 481 420
- C. Wilde, E. Woerner, C. Yeamans, B. Yoxall, J. Kilkenny, O. L. 482 421
- Landen, W. Hsing, and M. J. Edwards, Phys. Rev. Lett. 114, 483 422 175001 (2015). 484 423
- N. B. Meezan, L. F. Berzak Hopkins, S. Le Pape, L. Divol, a. J. 485 424
- MacKinnon, T. Döppner, D. D. Ho, O. S. Jones, S. F. Khan, T. Ma, 486 425 J. L. Milovich, a. E. Pak, J. S. Ross, C. a. Thomas, L. R. Benedetti, 487 426
- 427 D. K. Bradley, P. M. Celliers, D. S. Clark, J. E. Field, S. W. Haan, 488
- N. Izumi, G. a. Kyrala, J. D. Moody, P. K. Patel, J. E. Ralph, 489 428
- J. R. Rygg, S. M. Sepke, B. K. Spears, R. Tommasini, R. P. J. 490 429
- Town, J. Biener, R. M. Bionta, E. J. Bond, J. a. Caggiano, M. J. 491 430
- Eckart, M. Gatu Johnson, G. P. Grim, a. V. Hamza, E. P. Hartouni, 492 431
- R. Hatarik, D. E. Hoover, J. D. Kilkenny, B. J. Kozioziemski, J. J. 493 432
- Kroll, J. M. McNaney, a. Nikroo, D. B. Sayre, M. Stadermann, 494 433
- C. Wild, B. E. Yoxall, O. L. Landen, W. W. Hsing, and M. J. 495 434
- Edwards, Phys. Plasmas 22, 062703 (2015). 435
- L. F. Berzak Hopkins, S. Le Pape, L. Divol, N. B. Meezan, A. J. 497 436
- 437 Mackinnon, D. D. Ho, O. S. Jones, S. Khan, J. L. Milovich, J. S. 498
- Ross, P. Amendt, D. Casey, P. M. Celliers, A. Pak, J. L. Peterson, 438 499 500
- J. Ralph, and J. R. Rygg, Phys. Plasmas 22, 056318 (2015). 439

S. L. Pape, L. F. B. Hopkins, L. Divol, A. Pak, E. Dewald, N. B. Meezan, D. D.-m. Ho, S. F. Khan, A. J. Mackinnon, C. Weber, C. Goyon, J. S. Ross, M. Millot, L. R. Bennedetti, N. Izumi, G. A. Kyrala, T. Ma, S. R. Nagel, D. Edgell, A. G. Macphee, B. J. Macgowan, D. Strozzi, D. N. Fittinghoff, R. Hatarik, D. Sayre, P. Volegov, C. Yeamans, J. Biener, A. Nikroo, N. G. Rice, M. Stadermann, S. Bhandarkar, S. Haan, D. Callahan, W. W. Hsing, O. L. Landen, P. Patel, A. O. Hurricane, and M. J. Edwards, Submitted (2017).

444

469

496

- 10 H. Nagao, K. G. Nakamura, K. Kondo, N. Ozaki, K. Takamatsu, T. Ono, T. Shiota, D. Ichinose, K. a. Tanaka, K. Wakabayashi, K. Okada, M. Yoshida, M. Nakai, K. Nagai, K. Shigemori, T. Sakaiya, and K. Otani, Phys. Plasmas 13, 052705 (2006).
- 11 S. Brygoo, E. Henry, P. Loubeyre, J. Eggert, M. Koenig, B. Loupias, A. Benuzzi-Mounaix, and M. Rabec Le Gloahec, Nat. Mater. 6, 274 (2007).
- 12 D. G. Hicks, T. R. Boehly, P. M. Celliers, D. K. Bradley, J. H. Eggert, R. S. McWilliams, R. Jeanloz, and G. W. Collins, Phys. Rev. B 78, 174102 (2008).
- 13 M. D. Knudson, M. P. Desjarlais, and D. H. Dolan, Science (80-.). 322, 1822 (2008).
- 14 R. S. McWilliams, J. H. Eggert, D. G. Hicks, D. K. Bradley, P. M. Celliers, D. K. Spaulding, T. R. Boehly, G. W. Collins, and R. Jeanloz, Phys. Rev. B 81, 014111 (2010).
- 15 464 M. C. Gregor, D. E. Fratanduono, C. A. McCoy, D. N. Polsin, A. Sorce, J. R. Rygg, G. W. Collins, T. Braun, P. M. Celliers, J. H. Eggert, D. D. Meyerhofer, and T. R. Boehly, Phys. Rev. B 95, 144114 (2017).
 - D. Bradley, J. Eggert, R. Smith, S. Prisbrey, D. Hicks, D. Braun, J. Biener, a. Hamza, R. Rudd, and G. Collins, Phys. Rev. Lett. 102, 075503 (2009).
 - 17 R. F. Smith, J. H. Eggert, R. Jeanloz, T. S. Duffy, D. G. Braun, J. R. Patterson, R. E. Rudd, J. Biener, a. E. Lazicki, a. V. Hamza, J. Wang, T. Braun, L. X. Benedict, P. M. Celliers, and G. W. Collins, Nature 511, 330 (2014).
 - D. Bradley, J. Eggert, D. Hicks, P. Celliers, S. Moon, R. Cauble, and G. Collins, Phys. Rev. Lett. 93, 195506 (2004).
 - J. H. Eggert, D. G. Hicks, P. M. Celliers, D. K. Bradley, R. S. McWilliams, R. Jeanloz, J. E. Miller, T. R. Boehly, and G. W. Collins, Nat. Phys. 6, 40 (2009).
 - J. Biener, D. Ho, A. Hamza, and E. Al, Nucl. Fusion 49, 112001 (2009).
 - C. Dawedeit, S. O. Kucheyev, S. J. Shin, T. M. Willey, M. Bagge-Hansen, T. Braun, Y. M. Wang, B. S. El-Dasher, N. E. Teslich, M. M. Biener, J. Ye, L. Kirste, C.-C. C. Roehlig, M. Wolfer, E. Woerner, A. W. van Buuren, A. V. Hamza, C. Wild, and J. Biener, Diam. Relat. Mater. 40, 75 (2013).
 - H. F. Robey, P. M. Celliers, J. L. Kline, a. J. Mackinnon, T. R. Boehly, O. L. Landen, J. H. Eggert, D. Hicks, S. Le Pape, D. R. Farley, M. W. Bowers, K. G. Krauter, D. H. Munro, O. S. Jones, J. L. Milovich, D. Clark, B. K. Spears, R. P. J. Town, S. W. Haan, S. Dixit, M. B. Schneider, E. L. Dewald, K. Widmann, J. D. Moody, T. D. Döppner, H. B. Radousky, A. Nikroo, J. J. Kroll, a. V. Hamza, J. B. Horner, S. D. Bhandarkar, E. Dzenitis, E. Alger, E. Giraldez, C. Castro, K. Moreno, C. Haynam, K. N. LaFortune, C. Widmayer, M. Shaw, K. Jancaitis, T. Parham, D. M. Holunga, C. F. Walters, B. Haid, T. Malsbury, D. Trummer, K. R. Coffee, B. Burr, L. V. Berzins, C. Choate, S. J. Brereton, S. Azevedo, H. Chandrasekaran, S. Glenzer, J. a. Caggiano, J. P. Knauer, J. a. Frenje, D. T. Casey, M. Gatu Johnson, F. H. Séguin, B. K. Young, M. J. Edwards, B. M. Van Wonterghem, J. Kilkenny, B. J. Mac-

- 501
 Gowan, J. Atherton, J. D. Lindl, D. D. Meyerhofer, and E. Moses,
 556

 502
 Phys. Rev. Lett. 108, 215004 (2012).
 557
- ⁵⁰³ ²³ H. F. Robey, T. R. Boehly, P. M. Celliers, J. H. Eggert, D. Hicks, ⁵⁵⁸
 ⁶⁰⁴ R. F. Smith, R. Collins, M. W. Bowers, K. G. Krauter, P. S. Datte, ⁵⁵⁹
- ⁵⁰⁵ D. H. Munro, J. L. Milovich, O. S. Jones, P. A. Michel, C. A. 560
- Thomas, R. E. Olson, S. Pollaine, R. P. J. Town, S. Haan, D. Calla- 561
- han, D. Clark, J. Edwards, J. L. Kline, S. Dixit, M. B. Schnei- 562
- der, E. L. Dewald, K. Widmann, J. D. Moody, T. Doppner, H. B. 563
- Radousky, A. Throop, D. Kalantar, P. DiNicola, A. Nikroo, J. J. 564
- 510 Kroll, A. V. Hamza, J. B. Horner, S. D. Bhandarkar, E. Dzenitis, 565
- 511 E. Alger, E. Giraldez, C. Castro, K. Moreno, C. Haynam, K. N. 566
- 512 LaFortune, C. Widmayer, M. Shaw, K. Jancaitis, T. Parham, D. M. 567
- Holunga, C. F. Walters, B. Haid, E. R. Mapoles, J. Sater, C. R. 568
- 514 Gibson, T. Malsbury, J. Fair, D. Trummer, K. R. Coffee, B. Burr, 569
- L. V. Berzins, C. Choate, S. J. Brereton, S. Azevedo, H. Chan- 570
- drasekaran, D. C. Eder, N. D. Masters, A. C. Fisher, P. A. Sterne, 571
 B. K. Young, O. L. Landen, B. M. Van Wonterghem, B. J. Mac- 572
- B. K. Young, O. L. Landen, B. M. Van Wonterghem, B. J. Mac-518 Gowan, J. Atherton, J. D. Lindl, D. D. Meyerhofer, and E. Moses, 573
 Phys. Plasmas 19, 042706 (2012). 574
- 520²⁴ S. Hamel, L. Benedict, P. Celliers, M. Barrios, T. Boehly, 575
- 521 G. Collins, T. Döppner, J. Eggert, D. Farley, D. Hicks, J. Kline, 576
- ⁵²² a. Lazicki, S. LePape, a. Mackinnon, J. Moody, H. Robey, ⁵⁷⁷ E. Schwegler, and P. Sterne, Phys. Rev. B **86**, 1 (2012). ⁵⁷⁸
- ⁵²³ E. Schwegler, and I. Sterne, Phys. Rev. D 60, 1 (2012).
- ⁵²⁴ ²⁵ J. D. Moody, H. F. Robey, P. M. Celliers, D. H. Munro, D. a. ⁵⁷⁹
 ⁵²⁵ Barker, K. L. Baker, T. Döppner, N. L. Hash, L. Berzak Hop⁵²⁶ kins, K. LaFortune, O. L. Landen, S. LePape, B. J. MacGowan, ⁵⁸¹
- J. E. Ralph, J. S. Ross, C. Widmayer, a. Nikroo, E. Giraldez, and 582
- ⁵²⁸ T. Boehly, Phys. Plasmas **21**, 092702 (2014).
- ⁵²⁹ ²⁶ H. F. Robey, P. M. Celliers, J. D. Moody, J. Sater, T. Parham, 584
 ⁵³⁰ B. Kozioziemski, R. Dylla-Spears, J. S. Ross, S. LePape, J. E. 585
- Ralph, M. Hohenberger, E. L. Dewald, L. Berzak Hopkins, J. J. 586
- Kroll, B. E. Yoxall, a. V. Hamza, T. R. Boehly, a. Nikroo, O. L. 587
- Landen, and M. J. Edwards, Phys. Plasmas **21**, 022703 (2014).
- ²⁷ P. M. Celliers, D. K. Bradley, G. W. Collins, D. G. Hicks, T. R. 589
 Boehly, and W. J. Armstrong, Rev. Sci. Instrum. **75**, 4916 (2004). 590
- 536 ²⁸ A. Pak, L. Divol, A. L. Kritcher, T. Ma, J. E. Ralph, B. Bach-591
- mann, L. R. Benedetti, D. T. Casey, P. M. Celliers, E. L. Dewald, 592
 T. Döppner, J. E. Field, D. E. Fratanduono, L. F. Berzak Hopkins, 593
- N. Izumi, S. F. Khan, O. L. Landen, G. A. Kyrala, S. LePape, 594
- M. Millot, J. L. Milovich, A. S. Moore, S. R. Nagel, H.-S. Park, 595
- J. R. Rygg, D. K. Bradley, D. A. Callahan, D. E. Hinkel, W. W. 596
- Hsing, O. A. Hurricane, N. B. Meezan, J. D. Moody, P. Patel, H. F. 597
- Robey, M. B. Schneider, R. P. J. Town, and M. J. Edwards, Phys. 598
 Plasmas 24, 056306 (2017). 599
- 545²⁹ L. Divol, A. Pak, L. F. Berzak Hopkins, S. Le Pape, N. B. Meezan, 600
- 546 E. L. Dewald, D. D.-M. Ho, S. F. Khan, A. J. Mackinnon, J. S. 601
- 547 Ross, D. P. Turnbull, C. Weber, P. M. Celliers, M. Millot, L. R. 602
- 548 Benedetti, J. E. Field, N. Izumi, G. A. Kyrala, T. Ma, S. R. 603
- 549 Nagel, J. R. Rygg, D. Edgell, A. G. Macphee, C. Goyon, M. Ho- 604
- henberger, B. J. MacGowan, P. Michel, D. Strozzi, W. Cassata, 605
- D. Casey, D. N. Fittinghoff, N. Gharibyan, R. Hatarik, D. Sayre, 606
- P. Volegov, C. Yeamans, B. Bachmann, T. Döppner, J. Biener, 607
- J. Crippen, C. Choate, H. Huang, C. Kong, A. Nikroo, N. G. 608
- Rice, M. Stadermann, S. D. Bhandarkar, S. Haan, B. Kozioziem- 609
- ski, W. W. Hsing, O. L. Landen, J. D. Moody, R. P. J. Town, D. A. 610

Callahan, O. A. Hurricane, and M. J. Edwards, Phys. Plasmas 24, 056309 (2017).

- ³⁰ E. L. Dewald, F. Hartemann, P. Michel, J. Milovich, M. Hohenberger, A. Pak, O. L. Landen, L. Divol, H. F. Robey, O. A. Hurricane, T. Döppner, F. Albert, B. Bachmann, N. B. Meezan, A. J. Mackinnon, D. Callahan, and M. J. Edwards, Phys. Rev. Lett. **116**, 1 (2016).
- ³¹ M. D. Knudson and M. P. Desjarlais, Phys. Rev. B 88, 184107 (2013).
- ³² G. I. Kerley, *Equations of state for hydrogen and deuterium.*, Tech. Rep. (Sandia National Laboratories (SNL), Albuquerque, NM, and Livermore, CA, 2003).
- ³³ L. Caillabet, S. Mazevet, and P. Loubeyre, Phys. Rev. B **83**, 094101 (2011).
- ³⁴ P. M. Celliers, G. W. Collins, D. G. Hicks, and J. H. Eggert, J. Appl. Phys. **98**, 113529 (2005).
- S. Brygoo, M. Millot, P. Loubeyre, A. E. Lazicki, S. Hamel, T. Qi,
 P. M. Celliers, F. Coppari, J. H. Eggert, D. E. Fratanduono, D. G.
 Hicks, J. R. Rygg, R. F. Smith, D. C. Swift, G. W. Collins, and
 R. Jeanloz, J. Appl. Phys. **118**, 195901 (2015).
 - ³⁶ L. X. Benedict, K. P. Driver, S. Hamel, B. Militzer, T. Qi, A. a. Correa, a. Saul, and E. Schwegler, Phys. Rev. B **89**, 224109 (2014).
 - ³⁷ D. G. Hicks, T. R. Boehly, P. M. Celliers, J. H. Eggert, S. J. Moon,
 D. D. Meyerhofer, and G. W. Collins, Phys. Rev. B 79, 014112 (2009).
- ³⁸ P. Loubeyre, S. Brygoo, J. Eggert, P. M. Celliers, D. K. Spaulding,
 J. R. Rygg, T. R. Boehly, G. W. Collins, and R. Jeanloz, Phys.
 Rev. B 86, 144115 (2012).
 - ³⁹ M. D. Knudson and M. P. Desjarlais, Phys. Rev. Lett. **118** (2017), 10.1103/PhysRevLett.118.035501.
- ⁴⁰ A. a. Correa, S. a. Bonev, and G. Galli, Proc. Natl. Acad. Sci. U.
 S. A. 103, 1204 (2006).
 - ⁴¹ Z. G. Hu and P. Hess, Appl. Phys. Lett. **89**, 081906 (2006).
 - ⁴² F. Klauser, D. Steinmuller-Nethl, R. Kaindl, E. Bertel, and N. Memmel, Chem. Vap. Depos. 16, 127 (2010).
 - ⁴³ S. Potocky, A. Kromka, J. Potmesil, Z. Remes, Z. Polackova, and M. Vanecek, Phys. Status Solidi Appl. Mater. Sci. 203, 3011 (2006).
 - ⁴⁴ Z. G. Hu, P. Prunici, P. Hess, and K. H. Chen, J. Mater. Sci. Mater. Electron. **18**, 37 (2007).
 - ⁴⁵ T. Ruf, M. Cardona, C. S. J. Pickles, and R. Sussmann, Phys. Rev. B 62, 16578 (2000).
- ⁴⁶ J. Fontanella, R. L. Johnston, J. H. Colwell, and C. Andeen, Appl.
 Opt. 16, 2949 (1977).
 - ⁴⁷ W. Theobald, J. E. Miller, T. R. Boehly, E. Vianello, D. D. Meyerhofer, T. C. Sangster, J. Eggert, and P. M. Celliers, Phys. Plasmas 13, 122702 (2006).
 - ⁴⁸ A. Correa, L. Benedict, D. Young, E. Schwegler, and S. Bonev, Phys. Rev. B **78**, 024101 (2008).
 - ⁴⁹ G. I. KERLEY and L. C. CHHABILDAS, *Sandia Rep.*, Tech. Rep.
 September (Sandia National Laboratories (SNL), Albuquerque,
 NM, and Livermore, CA, 2001).
- ⁵⁰ J. R. Rygg, , in preparation (2018).
- ⁵¹ F. Peter, Zeitschrift fur Phys. **15**, 358 (1923).
- ⁵² G. Ghosh, Opt. Commun. **163**, 95 (1999).

Dataset ID	$U_S(HDC)$	$U_S(D_2)$	Batch	W-Doping	Density
	km/s	km/s		at.%	g/cm^3
N151004-005-E	25.40 ± 0.51	29.74 ± 0.90	KC038	0.22	3.59
N151004-005-E	24.34 ± 0.44	27.27 ± 0.91	KC038	0.22	3.59
N151005-003-P	24.05 ± 0.53	27.01 ± 0.93	KC038	0.22	3.59
N151101-001-E	24.67 ± 0.59	28.58 ± 0.91	KC038	0.22	3.59
N151101-001-E	23.28 ± 0.57	25.56 ± 1.15	KC038	0.22	3.59
N160110-002-E	26.08 ± 0.40	31.8 ± 0.81	KC038	0.22	3.59
N160114-003-E	24.85 ± 0.40	30.01 ± 0.81	KC038	0.22	3.59
N160114-003-P	24.95 ± 0.58	29.83 ± 1.12	KC038	0.22	3.59
N160131-001-E	26.86 ± 0.40	33.5 ± 0.81	KC047	0.24	3.60
N160131-001-P	26.20 ± 0.41	32.40 ± 0.80	KC047	0.24	3.60
N160419-001-E	25.51 ± 0.28	31.47 ± 0.47	K252	0.31	3.63
N160419-001-P	25.23 ± 0.24	30.83 ± 0.47	K252	0.31	3.63
N160523-001-E	24.66 ± 0.51	28.73 ± 0.84	KC047	0.24	3.60
N160523-001-P	24.12 ± 0.60	27.52 ± 1.03	KC047	0.24	3.60
N160724-005-E	25.12 ± 0.41	30.78 ± 0.81	KC047	0.24	3.60
N160724-005-P	25.32 ± 0.41	30.68 ± 0.81	KC047	0.24	3.60
N161115-002-E	24.58 ± 0.39	28.38 ± 0.82	KC047	0.24	3.60
N161115-002-P	24.93 ± 0.43	29.06 ± 0.82	KC047	0.24	3.60
N170119-001-E	25.02 ± 0.24	30.32 ± 0.46	KC206	0.16	3.56
N170119-001-P	25.84 ± 0.26	31.25 ± 0.47	KC206	0.16	3.56
N170228-002-E	25.47 ± 0.41	30.60 ± 0.82	KC210	0.13	3.55
N170228-002-P	25.64 ± 0.41	30.58 ± 0.82	KC210	0.13	3.55
N170409-002-E	26.42 ± 0.40	32.40 ± 0.82	KC210	0.13	3.55
N170409-002-P	26.85 ± 0.41	32.92 ± 0.82	KC210	0.13	3.55
N170809-001-E	24.78 ± 0.43	28.15 ± 0.82	K252	0.31	3.63
N170809-001-P	25.45 ± 0.43	30.15 ± 0.81	K252	0.31	3.63
N171016-002-Е	24.48 ± 0.40	29.02 ± 0.84	KC245	0.25	3.60
N171016-002-P	24.51 ± 0.43	29.24 ± 0.83	KC245	0.25	3.60

 $TABLE III. \ \ \ Experimental \ \ details \ and \ result \ summary \ for \ a \ HDC/Quartz \ shock \ impedance \ matching \ experiment. \ The \ target \ package \ details \ with \ thicknesses \ in \ microns \ are \ Al[0.2]HDC[28]Au[2]HDC[492]Quartz[525]. \).$

ShotID	$U_S(HDC)$	$U_S(Quartz)$
N151214-003	23.61 ± 0.24	21.70 ± 0.36

TABLE IV. Re-analyzed Diamond Hugoniot equation of state data

Shot ID	Standard	ρ_0	U_S	u_p	Р	ρ	Source
		g/cm^3	km/s	km/s	GPa	g/cm^3	
26332	Al	3.51	29.00 ± 1.16	15.04 ± 1.03	1530.78 ± 104.88	7.29 ± 0.54	Nagao
26396	Al	3.51	19.90 ± 0.86	7.43 ± 0.96	518.91 ± 67.32	5.60 ± 0.43	Nagao
27300	Al	3.51	32.70 ± 1.70	18.79 ± 2.31	2156.99 ± 264.62	8.25 ± 1.37	Nagao
28693	Al	3.51	27.20 ± 1.44	14.72 ± 1.81	1405.35 ± 173.13	7.65 ± 1.11	Nagao
28695	Al	3.51	21.40 ± 1.33	8.38 ± 0.96	629.38 ± 72.15	5.77 ± 0.43	Nagao
28697	Al	3.51	24.60 ± 2.29	11.94 ± 1.14	1031.01 ± 98.41	6.82 ± 0.61	Nagao
20547	Al	3.51	23.12 ± 0.34	11.83 ± 0.75	959.69 ± 61.17	7.19 ± 0.48	Bradley
24278	Al	3.51	24.58 ± 0.26	11.73 ± 0.60	1012.34 ± 52.16	6.72 ± 0.32	Bradley
24365	Al	3.51	23.82 ± 0.26	13.16 ± 0.43	1100.03 ± 36.36	7.84 ± 0.32	Bradley
24294	Al	3.51	26.18 ± 0.26	13.43 ± 0.69	1233.69 ± 63.43	7.20 ± 0.39	Bradley
24288	Al	3.51	30.70 ± 0.27	17.95 ± 0.94	1933.79 ± 100.89	8.45 ± 0.62	Bradley
20541	Al	3.51	34.95 ± 0.35	23.86 ± 1.82	2926.57 ± 222.95	11.06 ± 1.81	Bradley
26397	Mo	3.51	39.08 ± 0.33	26.10 ± 1.42	3580.00 ± 192.10	10.57 ± 1.17	Bradley
10076	0	2.51	20.20 1.0.24	0.00 + 0.10	600 5 0 1 0 0 4		TT' 1
49976	Quartz	3.51	20.39 ± 0.24	8.39 ± 0.12	600.72 ± 8.84	5.97 ± 0.09	Hicks
50364	Quartz	3.51	21.66 ± 0.13	9.42 ± 0.16	716.30 ± 12.45	6.21 ± 0.09	Hicks
49974	Quartz	3.51	22.05 ± 0.11	9.80 ± 0.11	758.30 ± 8.58	6.32 ± 0.07	Hicks
49614	Quartz	3.51	22.46 ± 0.08	10.22 ± 0.11	805.62 ± 8.66	6.44 ± 0.06	Hicks
51565	Quartz	3.51	23.36 ± 0.07	10.84 ± 0.13	889.11 ± 10.50	6.55 ± 0.07	Hicks
49616	Quartz	3.51	23.46 ± 0.07	10.88 ± 0.13	895.67 ± 10.55	6.54 ± 0.07	Hicks
49615	Quartz	3.51	23.72 ± 0.07	10.98 ± 0.12	914.11 ± 10.29	6.53 ± 0.07	Hicks
48882	Quartz	3.51	24.22 ± 0.07	11.57 ± 0.13	983.96 ± 10.83	6.72 ± 0.07	Hicks
48448	Quartz	3.51	25.02 ± 0.09	12.06 ± 0.14	1058.95 ± 12.42	6.78 ± 0.08	Hicks
48880	Quartz	3.51	26.33 ± 0.07	13.94 ± 0.11	1288.34 ± 10.16	7.46 ± 0.07	Hicks
49450	Quartz	3.51	28.42 ± 0.08	16.12 ± 0.15	1608.03 ± 14.92	8.11 ± 0.11	Hicks
49447	Quartz	3.51	30.19 ± 0.09	17.76 ± 0.19	1881.81 ± 19.72	8.53 ± 0.14	Hicks
D1	Quartz	2 5 1	26.20 ± 0.10	14.89 ± 0.91	$1267,7600\pm10,62$	9.04 ± 0.16	Dragoo
	Qualtz	2.51	20.30 ± 0.10	14.02 ± 0.21	1307.7000 ± 19.03	6.04 ± 0.10	Drygoo
D2 D2	Quartz	2.51	20.00 ± 0.20	0.04 ± 0.10	399.7220 ± 7.00	0.13 ± 0.08	Divgoo
B3 D4	Quartz	3.51	23.10 ± 0.10	12.53 ± 0.12	1010.1300 ± 10.29 1206 0400 ± 10.46	7.07 ± 0.10	Brygoo
B4 D5	Quartz	3.51	26.10 ± 0.10	14.48 ± 0.21	1326.2400 ± 19.46	7.88 ± 0.15	Brygoo
BO	Quartz	3.51	22.00 ± 1.00	8.77 ± 0.73	095.0390 ± 39.41	5.75 ± 0.39	Brygoo
B0 D7	Quartz	3.31	23.10 ± 1.00	12.83 ± 1.33	1039.5500 ± 109.86	8.01 ± 1.39	Brygoo David
B/	Quartz	3.51	20.00 ± 1.30	10.33 ± 1.05	1431.3200 ± 157.89	8.45 ± 1.76	ыrygoo
B8 D0	Quartz	5.51	21.60 ± 0.90	11.38 ± 1.10	802.5200 ± 84.81	1.49 ± 1.03	ыrygoo
89	Quartz	3.51	22.50 ± 0.90	8.55 ± 0.73	$0/4.8500 \pm 58.58$	5.68 ± 0.36	Brygoo