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Hall Viscosity and the Acoustic Faraday Effect

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For more than 20 years, observation of the non-dissipative Hall viscosity in the quantum Hall effect has been impeded by the difficulty to probe directly the momentum of the two-dimensional electron gas. However, in three-dimensional systems such as superfluid ³He–B, the momentum density is readily probed through transverse acoustic waves. We show that in a three-dimensional elastic medium supporting transverse waves, a non-vanishing Hall viscosity induces circular bire-fringence. Such an effect has been observed in ³He–B in the presence of a weak magnetic field, and is known as the acoustic Faraday effect. The acoustic Faraday effect has been understood in terms of the Zeeman splitting of the excited order parameter modes which support the transverse wave propagation in the superfluid. We show that the Zeeman effect can generically lead to a non-zero Hall viscosity coefficient, and confirm this prediction using a simple phenomenological model for the ³He–B collective modes. Therefore, we claim that the observation of the acoustic Faraday effect can be leveraged to make a direct observation of the Hall viscosity in superfluid ³He–B in a magnetic field and other systems such as the crystalline Tb3Ga5O12 material.

I. INTRODUCTION

Liquids with broken time-reversal symmetry can exhibit a dissipationless response to a strain rate known as the Hall viscosity. In direct analogy to the Hall conductivity, the Hall viscosity in quantum systems may be understood as an effect of a Berry curvature of the ground state^{1,2}. The early predictions of Hall viscosity pertained to the integer quantum Hall effect, but the Hall viscosity response is far more general, and is present in any time-reversal breaking fluid³ including the fractional quantum Hall effect⁴⁻⁶, the quantum anomalous Hall effect^{7,8}, and chiral superconductors/superfluids^{4,9,10}. It has also been predicted in gapless systems where strain fields couple to the low-energy electronic degrees of freedom as gauge fields (emergent elastic gauge fields) such as graphene¹¹, borophene¹², and Weyl semimetals¹³⁻¹⁵.

Despite the initial predictions more than two decades ago, there are no experiments that have observed the Hall viscosity in a two-dimensional electron gas (2DEG). The viscosity drives momentum transport, and the difficulty in directly probing the 2DEG momentum is a primary obstacle to this observation. However, the charge current and density responses at finite wavevector^{16,17}, and the density response to spatial curvature^{6,18} have been predicted to have corrections due to the Hall viscosity, and may lead to more realistic experimental proposals in the integer and fractional quantum Hall contexts. Indeed, proposals exist to measure the Hall viscosity response¹⁹ and the electron viscosity more generally²⁰, but none have yet been realized.

Interestingly, the effects of the Hall viscosity have been observed in a non-electronic system, i.e., a photonic analog of the quantum Hall effect²¹, and other recent articles propose to measure the Hall viscosity in superfluids²² and (classical) chiral active fluids²³. Such systems may be a more expedient route to observing this response because the momentum can be probed directly. In this article we predict another consequence of the Hall viscosity in superfluids. We claim that the Hall viscosity will generate a Faraday effect in the transverse acoustic waves in superfluid 3 He-B when subjected to a magnetic field. The connection between the Hall viscosity and the acoustic Faraday effect is motivated by an experiment that has already observed this Faraday effect²⁴. We will focus on the connection to 3 He-B here, but we note that the general concepts can be applied to other contexts where the acoustic Faraday effect has been observed or predicted, e.g., in the crystal Tb3Ga5O12^{25,26} or in superconductor vortex lattices²⁷⁻²⁹.

Transverse acoustic waves (TAWs) provide a direct probe of momentum transport in three-dimensional paired superfluids such as ³He–B, and are an unusual collective excitation for a fluid (since simple fluids usually only support longitudinal acoustic waves). TAWs were predicted early in the history of ${}^{3}\text{He}-\text{B}$ research 30,31 , but were not described in detail theoretically 32 , or observed²⁴, until much later. This is unlike conventional longitudinal acoustic waves that were observed much $earlier^{33,34}$. In a typical experiment the superfluid resides in a cavity with one wall that acts as an TAW transducer. The collective mode excitations of the superfluid pairs are largely responsible for supporting the propagation of TAWs in the superfluid. If we subject the (intrinsically time-reversal invariant) superfluid to a weak magnetic field, the degeneracy of the collective mode excitations is broken due to the Zeeman effect. Since the TAWs are coupled to the collective modes, the Zeeman effect, as we will see below, generates changes in the relative phase velocity of the circular polarized components of the TAWs (i.e., circular birefringence). For example, if a linearly-polarized TAW is injected into the fluid, then the linear-polarization rotates as it propagates through ³He–B due to the relative phase velocity of each component, a phenomenon known as the acoustic Faraday

effect. We will show that this effect can be interpreted as arising directly from the Hall viscosity generated by the application of the magnetic field.

Our article is organized as follows. We begin with an overview of TAW propagation in generic fluids, and subsequently show that a Hall viscosity manifests in an acoustic Faraday effect. With this understanding, which is valid for a generic visco-elastic medium, we use the known properties of TAWs and collective modes in ³He–B to estimate the Hall viscosity coefficient at frequencies near the collective mode resonance. Finally, we present a simple phenomenological model of the relevant ³He–B collective modes and carry out a Kubo-formula momentum response calculation to show that it exhibits the same Hall viscosity coefficient, thus corroborating our discussion.

II. TRANSVERSE ACOUSTIC WAVES

We begin by reviewing TAW propagation in viscoelastic media. The conservation law for the momentum density field g^a is the constitutive equation

$$\partial_t g^a + \partial_b T^{ba} = 0 \tag{1}$$

where T^{ab} is the stress tensor, and a, b = 1, 2, 3. Designating the mass density by ρ , and the displacement field as u^a , the momentum density is simply

$$g^a = \rho \partial_t u^a. \tag{2}$$

Throughout we will assume that the density is uniform as we are concerned with transverse, not longitudinal waves. The stress tensor is related to the strain,

$$w_{ab} = \frac{1}{2}(\partial_a u_b + \partial_b u_a),\tag{3}$$

and the strain rate $\partial_t w_{ab}$, through the elasticity λ^{abcd} and viscosity η^{abcd} tensors:

$$T^{ab} = -\lambda^{abcd} w_{cd} - \eta^{abcd} \partial_t w_{cd}.$$
 (4)

Now let us choose a plane-wave ansatz for a TAW, i.e., let the displacement field be a plane wave with wavevector \mathbf{q} and frequency ω :

$$\mathbf{u}(\mathbf{x},t) = \mathbf{u} \, e^{i(\mathbf{q} \cdot \mathbf{x} - \omega t)}.\tag{5}$$

In three dimensions, we can decompose the polarization vector \mathbf{u} using an oriented triad of real, orthonormal vectors $\{\mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3\}$ (the linear polarization basis) such that \mathbf{e}_3 is chosen to be the direction of propagation: $\mathbf{q} = q \mathbf{e}_3$. We can proceed directly to the constitutive equation evaluated in this ansatz:

$$-\rho\omega^2 u^a + q^2 \left(\lambda^{ac} - i\omega\eta^{ac}\right) g_{cf} u^f = 0, \qquad (6)$$

where, for brevity, we have designated nine components each of the elasticity and viscosity tensors as

$$\lambda^{3a3c} = \lambda^{ac}$$
 and $\eta^{3a3c} = \eta^{ac}, a, c \in \{1, 2, 3\}.$

Note that the components with a = 3 or c = 3 do not contribute because the transverse waves have $u^3 = 0$.

In terms of classical forces, transverse waves propagate in an elastic medium because the elasticity provides a restoring force against a strain gradient. The ordinary viscosity appears as an imaginary term in Eq. (6), and generates dissipation by producing a drag force. The anti-symmetric part of the viscosity is often neglected (e.g., in time-reversal invariant systems it vanishes), but it leads to a force perpendicular to the strain rate gradient. To highlight the relationship between Hall viscosity and angular momentum, we may prefer to think of this as a non-dissipative *torque*. To see this clearly, we will work in the circular polarization basis, $\{\mathbf{e}_+, \mathbf{e}_-, \mathbf{e}_3\}$, where

$$\mathbf{e}_{\pm} = \frac{1}{\sqrt{2}} (\mathbf{e}_1 \pm i \mathbf{e}_2).$$

This basis has a metric given by

$$g_{++} = g_{--} = \mathbf{e}_+ \cdot \mathbf{e}_+ = \mathbf{e}_- \cdot \mathbf{e}_- = 0$$
, and
 $g_{+-} = g_{-+} = \mathbf{e}_+ \cdot \mathbf{e}_- = \mathbf{e}_- \cdot \mathbf{e}_+ = 1$,

so we must take care when raising and lowering indices. The components of a 2-index tensor A in the circular basis are related to the linear basis by

$$A^{++} = (A^{--})^* = \frac{1}{2}(A^{11} - A^{22}) - \frac{i}{2}(A^{12} + A^{21})$$
(7)

$$A^{+-} = (A^{-+})^* = \frac{1}{2}(A^{11} + A^{22}) - \frac{i}{2}(A^{12} - A^{21}) \quad (8)$$

where A may be the (reduced) elasticity λ or viscosity η . The components A^{++} and A^{--} violate rotation symmetry and generate coupling between right-circularly polarized waves; therefore, we demand that they vanish. Additionally, transverse waves in ³He–B only couple to non-dissipative order parameter fluctuations³², so we are justified to neglect the dissipative terms (Im λ^{+-} , Re η^{+-}). Including the ordinary shear viscosity through the Re η^{+-} coefficient would introduce a contribution to the constitutive equation out of phase with the other terms, leading to a damped solution instead of an undamped plane wave. In combination with rotation symmetry, this yields the constraints

$$\begin{array}{ll} \lambda^{12} &= \lambda^{21} = 0 \\ \eta^{11} &= \eta^{22} = 0 \end{array}$$

The only surviving terms are the *shear modulus*,

$$G = \frac{1}{2} \left(\lambda^{11} + \lambda^{22} \right)$$

and the Hall viscosity,

$$\eta_H = \frac{1}{2} \left(\eta^{12} - \eta^{21} \right).$$

Under these conditions, the dispersion relations for the circular polarization components \mathbf{e}_{\pm} are

$$\rho\omega^2 = (G \mp \omega\eta_H) q^2. \tag{9}$$

Immediately we see that each component has a different phase velocity: the fluid exhibits circular birefringence when $\eta_H \neq 0$, i.e., when time-reversal symmetry is broken. Hence, an observation of circular birefringence in transverse acoustic waves would enable the direct determination of the Hall viscosity coefficient. We will consider this possibility in the context of ³He–B below.

III. SUPERFLUID HELIUM 3-B

The superfluid *B*-phase of ³He is described by the spintriplet, *p*-wave pairing Balian-Werthamer state^{32,35,36}. The order parameter is a 3×3 complex matrix,

$$d_{ij} = \frac{\Delta}{\sqrt{3}} R_{ij}(\hat{\mathbf{n}}, \theta) e^{i\phi} \tag{10}$$

parameterized by the self-energy amplitude Δ , the phase ϕ , and the rotation R_{ij} of the spin by an angle θ around an axis $\hat{\mathbf{n}}$ specified relative to the orbital angular momentum³². The ground state and collective mode excitations are eigenstates of the *twisted* total angular momentum operator,

$$\mathbf{J} = \mathbf{L} + R^{-1} \mathbf{S}. \tag{11}$$

The states are additionally labeled by their signature under particle-hole symmetry; the real part of d_{ij} transforms with signature +1, and the imaginary with signature -1.

The six families of states $J = \{0, 1, 2\}^{\{+,-\}}$ comprise 18 states in all. The $J = 0^-$ and $J = 1^+$ states are the Goldstone modes, coupling to longitudinal zero sound and spin waves respectively^{37,38}. The counterparts of these modes, $J = 0^+$ and $J = 1^-$, are not relevant to our considerations here: these branches of modes are at, and beyond, the pair-breaking edge and hence they are strongly damped³⁹. The real $J = 2^+$ excitations do not couple to transverse waves in the quasiclassical linear response theory, so we do not consider them^{36,39,40}. Thus, it is only the $J = 2^-$ imaginary squashing modes that couple to TAWs below the pair-breaking frequency, and are primarily responsible for collision-less transverse sound.

Circularly-polarized TAWs transform under the J = 2angular momentum representation and carry $m = \pm 1$; therefore, to conserve angular momentum, they couple only to the subset of the $J = 2^{-}$ multiplet with $m = \pm 1^{32}$. The dispersion relation for TAWs with frequency ω and wavevector q is given explicitly in Refs. 32 and 41 as

$$\frac{\omega^2}{v_F^2 q^2} = \Lambda_0 + \Lambda_{(2^-)} \frac{\omega^2}{\omega^2 - \omega_{(2^-)}^2 - \frac{2}{5} v_F^2 q^2}, \quad (12)$$

where, v_F is the Fermi velocity, Λ_0 is the effective quasiparticle restoring force which is insensitive to the magnetic field, and $\Lambda_{(2^-)}$ is the coupling to the $J = 2^-$ collective modes. The denominator of the second term depends on the dispersion relation of the imaginary-squashing collective modes,

$$\omega^2 = \omega_{(2^-)}^2 + \frac{2}{5}v_F^2 q^2,$$

where $\omega_{(2^-)}$ is the frequency edge for the $J = 2^-$ modes, which will be modified by a magnetic field. In fact, the sensitivity of the TAWs to the magnetic field is primarily due to the Zeeman splitting of the $J = 2^-$ collective modes. In the following, we will show that the Zeeman effect may also be thought of as generating a contribution to the Hall viscosity, which we have already seen to be responsible for creating acoustic circular birefringence.

In zero magnetic field, the lowest energy, fully degenerate $J = 2^-$ modes have frequency $\omega_{(2^-)} = \sqrt{\frac{12}{5}}\Delta$. We can determine the shear modulus due to the $J = 2^-$ collective modes using Eq. (9),

$$\rho^{-1}G_{(2^-)} = \frac{\Lambda_{(2^-)}v_F^2\omega^2}{\omega^2 - \frac{12}{5}\Delta^2 - \frac{2}{5}v_F^2q^2}$$

Note that we are neglecting the quasiparticle contributions to the shear modulus because the quasiparticle contribution is insensitive to the magnetic field (i.e., only contains time-reversal invariant contributions) and so it does not contribute to the Hall viscosity. In the longwavelength limit, which is valid near resonance,

$$\lim_{q \to 0} \rho^{-1} G_{(2^-)} = \frac{\Lambda_{(2^-)} v_F^2 \omega^2}{\omega^2 - \frac{12}{5} \Delta^2}.$$

Applying a weak magnetic field along the propagation direction breaks the $J = 2^-$ degeneracy by Zeeman splitting:

$$\omega_{(2^-)} = \sqrt{\frac{12}{5}}\Delta + m_{(2^-)}g_{(2^-)}\omega_L$$

where $g_{(2^-)}$ is the Landé g-factor⁴², $m_{(2-)}$ is the angular momentum quantum number along the applied field direction, and ω_L is the Larmor frequency of the $J = 2^$ modes given by

$$g_{(2^-)}\hbar\omega_L = \gamma B_z,$$

where γ is the effective coupling constant of the collective modes to the magnetic field⁴³. The dispersion relation (12) now differs for each of the $m_{(2^-)} = \pm 1$ components,

$$\frac{\omega^2}{v_F^2 q^2} = \Lambda_0 + \Lambda_{(2^-)} \frac{\omega^2}{\omega^2 - \left[\sqrt{\frac{12}{5}}\Delta \pm g_{(2^-)}\omega_L\right]^2 - \frac{2}{5}v_F^2 q^2}$$

Let us consider the limit near resonance in which the magnetic field B is weak enough that we can consider expanding to linear order in B, i.e.,

$$g_{(2^-)}\omega_L \ll \sqrt{\omega^2 - \frac{12}{5}\Delta^2} \ll \Delta.$$

Comparing with Eq. (9), we find the Hall viscosity for weak magnetic fields in the long-wavelength limit to be

$$\lim_{q \to 0} \eta_H = -2G_{(2^-)} \frac{g_{(2^-)}\omega_L}{\omega^2 - \frac{12}{5}\Delta^2},$$
(13)

where we have neglected terms of order $(\omega^2 - \frac{12}{5}\Delta^2)/\omega^2$. This shows explicitly that the Zeeman splitting of the $J = 2^-$ collective modes may be interpreted as a direct contribution to the Hall viscosity coefficient and hence affects the TAW propagation. We will further justify this claim by providing corroborating evidence from a Kubo formula calculation of the Hall viscosity in a phenomenological model of the collective modes. This will show the origin of the Hall viscosity from an alternate perspective.

IV. PHENOMENOLOGICAL MODEL

Let us consider a simple phenomenological model of the low-energy superfluid collective modes as an ensemble of non-interacting bosons. We will see that this suffices to derive a Hall viscosity coefficient that agrees with the circular birefringence result above. The success of this model further supports our interpretation of the acoustic Faraday effect as a Zeeman-induced Hall viscosity.

We will describe an effective theory of the collective modes using a model of non-interacting bosons with orbital angular momentum $\hat{\mathbf{L}}$ and spin $\hat{\mathbf{S}}$. The orbital angular momentum represents, roughly, the orbital angular momentum of the quasiparticles making up the superfluid pairs. At the mean field level, the effective interaction-experienced by the quasiparticles as they orbit-is perturbed by the strain on the system; therefore, the orbital angular momentum will be coupled to strain. However, the spin-an internal degree of freedom-is not coupled to the strain in this context. Indeed, the Cooper pair has a spatial extent, so its orbital angular momentum is sensitive to the (effective) spatial metric induced by the strain. Spin, on the other hand, is the intrinsic angular momentum of a point-like particle and insensitive to the effective strain metric.

With this assumption the dynamics of the collective modes are given by the model Hamiltonian

$$H = H_0(\mathbf{\hat{J}}) + \gamma \mathbf{B} \cdot \mathbf{\hat{J}} + \mu^{-1} w^{mn} \hat{L}_m \hat{L}_n + 2\gamma w^{mn} B_m \hat{L}_n,$$

where μ is the effective pair moment of inertia, and γ is the effective pair coupling to the magnetic field. The symmetrized strain tensor w_{mn} is defined in Eq. (3). The Hamiltonian $H_0(\hat{\mathbf{J}})$ is the zero-field, zero-strain Hamiltonian for the bosons. With $\hat{\mathbf{J}}$ given by Eq. (11), and the knowledge that the low-energy collective modes of ³He–B arise from the L = S = 1 pairs, it suffices for our purposes to take

$$H_0(\hat{\mathbf{J}}) = \sum_{J=0}^{2} \sum_{m=-J}^{J} \hbar \omega_{(J^-)} |J, m\rangle \langle J, m|$$

with the well-established spectrum of particle-hole antisymmetric collective modes,

$$\omega_{(0^-)} = 0, \ \omega_{(1^-)} = 2\Delta, \ \text{and} \ \omega_{(2^-)} = \sqrt{\frac{12}{5}}\Delta.$$

However, we could easily have chosen H_0 to model another system, or to also include the particle-hole symmetric modes, and it should apply in more general contexts.

The strain susceptibility is given by the Kubo formula 9,44

$$\begin{split} \chi^{abmn} &= -\frac{i}{\omega} \frac{1}{\sqrt{g}} \left\langle \frac{\delta T^{ab}}{\delta w_{mn}} \right\rangle \\ &+ \lim_{\epsilon \to 0^+} \frac{1}{\hbar \omega^+} \int_0^\infty \mathrm{d}t \, e^{i\omega^+ t} \left\langle \left[T^{ab}(t), T^{mn} \right] \right\rangle \end{split}$$

where the symmetric stress tensor is

$$\begin{split} T^{ab} &= -\frac{1}{\sqrt{g}} \frac{\delta H}{\delta w_{ab}} \\ &= \frac{1}{\sqrt{g}} \left(g^{am} g^{bn} + g^{an} g^{bm} \right) \left[\frac{1}{2\mu} \hat{L}_m \hat{L}_n + \gamma B_m \hat{L}_n \right], \end{split}$$

and the effective metric is $g_{mn} = \delta_{mn} + 2w_{mn}$. The stress tensor is independent of H_0 since it is insensitive to strain. The variational term of the susceptibility gives the infinite-frequency (contact) contribution to the transport coefficients; if the zero-field ground state is isotropic in space, then it gives no contribution to η_H . The commutator term gives the finite-frequency contribution; it vanishes unless the orbital angular momentum L > 0 and time-reversal symmetry is broken. This calculation is described in detail in the Appendix, but the results are given below.

For the specific case of ³He–B, the shear modulus is

$$\begin{split} G_{(2^{-})} &= -\frac{1}{2}i\omega\left(\chi^{1313} + \chi^{2323}\right) \\ &= \frac{1}{3}\frac{n}{\hbar}\left[\frac{\hbar^2}{2\mu}\right]^2\sqrt{\frac{12}{5}}\frac{\Delta}{\left(\omega^2 - \frac{12}{5}\Delta^2\right)} + \frac{4}{3}\frac{\hbar^2}{2\mu}n, \end{split}$$

where n is the boson number density. The Hall viscosity coefficient is

$$\eta_{H} = \frac{1}{2} \left(\chi^{1323} - \chi^{2313} \right) \\ = -\frac{2}{3} \frac{n}{\hbar} \left[\frac{\hbar^{2}}{2\mu} \right]^{2} \sqrt{\frac{12}{5}} \frac{\Delta}{\left(\omega^{2} - \frac{12}{5}\Delta^{2}\right)^{2}} \frac{\gamma B_{z}}{\hbar}$$

where $\gamma B_z/\hbar = g_{(2^-)}\omega_L$. Near resonance, i.e., when $\omega^2 - \frac{12}{5}\Delta^2 \ll \Delta^2$, the shear modulus is entirely dominated by its divergent term and the Hall viscosity coefficient generated by Zeeman splitting agrees with the prediction from the previous section,

$$\eta_H = -2G_{(2^-)}\frac{g_{(2^-)}\omega_L}{\omega^2 - \frac{12}{5}\Delta^2}.$$

We conclude that the Zeeman-split collective modes induce an effective Hall viscosity, which is responsible for the acoustic Faraday effect in ³He–B.

V. DISCUSSION AND CONCLUSION

From our results, it may be possible to directly observe the Hall viscosity in future experiments. Several studies have been undertaken that demonstrate the acoustic Faraday effect in ³He–B using acoustic cavity interferometry. In this type of experiment, one wall of the cavity serves as a transducer that generates and detects TAWs with a particular linear polarization. In the lowattenuation regime, waves reflected from the opposite wall of the cavity interfere with the waves emitted by the transducer to modify the acoustic impedance. The relative phase of the emitted and reflected waves depends on the dimensions of the cavity and on the phase velocity of the waves, which is controlled through the temperature. As the temperature varies, the acoustic impedance oscillates in response to the changing TAW wavelength. We have seen that, in a magnetic field, the Hall viscosity induces a relative phase velocity between polarization components, and the relative phase velocity causes the polarization of the wave to rotate as it traverses the cavity. Rotation of the reflected TAW away from the polarization axis of the transducer reduces the detected interference; if the polarization of the reflected wave has rotated 90° when it returns to the transducer, no change in the acoustic impedance is detected. The vanishing impedance os5

cillations have been used to confirm the presence of the acoustic Faraday effect²⁴ and acoustic cavity interferometry has been used further to determine precise values of the Landé g-factor^{41,45}. If one can modify such an experiment to collect data on the relative phase velocities then it would be possible to directly extract the Hall viscosity.

We would also like to remark briefly on the observation of the acoustic Faraday effect in other systems. Several magneto-acoustic phenomena have been observed in Tb₃Ga₅O₁₂^{25,26}. In this crystal, a magnetic field breaks the degeneracy of phonon modes through interaction with the Tb³⁺ ions²⁶. Such degeneracy breaking is reminiscent of the collective mode Zeeman effect in ³He-B, which suggests this system for future studies of the Hall viscosity. The acoustic Faraday effect is also predicted in vortex lattices in type-II superconductors²⁷⁻²⁹, where the Magnus force between vortices enters the acoustic wave dispersion relation as a Hall viscosity term. We are not aware of any experiment observing this effect, but the similarity to the phenomenon described here indicates it as a promising avenue of investigation.

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Appendix: Calculation of the Kubo Formula for Hall Viscosity

To calculate the Hall viscosity near resonance, we treat the collective mode excitations as a non-interacting gas of bosons with dynamics given by the Hamiltonian

$$H = H_0(\hat{\mathbf{J}}) + \gamma \mathbf{B} \cdot \hat{\mathbf{J}} + \mu^{-1} w^{mn} \hat{L}_m \hat{L}_n + 2\gamma w^{mn} B_m \hat{L}_n.$$

Our interpretation of this Hamiltonian in terms of the superfluid collective modes is given in the main text. We remind the reader of the following facts:

• The Hamiltonian acts on the Hilbert space of eigenstates of J^2 and J_3 where

$$\mathbf{\hat{J}} = \mathbf{\hat{L}} + R^{-1}\mathbf{\hat{S}}$$

and
$$L = S = 1$$
 so that $J \in \{0, 1, 2\}$. These
represent the particle-hole antisymmetric collective
modes of ³He–B.

• At zero strain in zero field, we model the collective modes with the Hamiltonian

$$H_0(\mathbf{\hat{J}}) = \sum_{J=0}^2 \sum_{m=-J}^J \hbar \omega_{(J^-)} |J, m\rangle \langle J, m|.$$

with the collective mode spectrum $\omega_{(0^-)} = 0$, $\omega_{(1^-)} = 2\Delta$, and $\omega_{(2^-)} = \sqrt{\frac{12}{5}}\Delta$.

The stress susceptibility is given by the Kubo formula

$$\begin{split} \chi^{abmn} &= -\frac{i}{\omega} \frac{1}{\sqrt{g}} \left\langle \frac{\delta T^{ab}}{\delta w_{mn}} \right\rangle \\ &+ \lim_{\epsilon \to 0^+} \frac{1}{\hbar \omega^+} \int_0^\infty \mathrm{d}t \, e^{i\omega^+ t} \left\langle \left[T^{ab}(t), T^{mn} \right] \right\rangle \end{split}$$
 The symmetric stress tensor is

$$T^{ab} = -\frac{1}{\sqrt{g}} \frac{\delta H}{\delta w_{ab}}$$
$$= \frac{1}{\sqrt{g}} \left(g^{am} g^{bn} + g^{an} g^{bm} \right) \left[\frac{1}{2\mu} \hat{L}_m \hat{L}_n + \gamma B_m \hat{L}_n \right]$$

in terms of the effective metric $g_{mn}=\delta_{mn}+2w_{mn}.$ At zero strain, this yields

$$T^{ab}\big|_{w=0} = -\frac{1}{2\mu} \left(\hat{L}^{a} \hat{L}^{b} + \hat{L}^{b} \hat{L}^{a} \right) - \gamma \left(B^{a} \hat{L}^{b} + B^{b} \hat{L}^{a} \right).$$

The instantaneous term in the susceptibility is the expectation value at zero strain of

$$\frac{1}{\sqrt{g}} \frac{\delta T^{ab}}{\delta w_{mn}} \bigg|_{w=0} = -T^{ab} \delta^{mn} - \left(\delta^{ma} \left[\frac{1}{2\mu} \left(\hat{L}^n \hat{L}^b + \hat{L}^b \hat{L}^n \right) + \gamma (B^n \hat{L}^b + B^b \hat{L}^n) \right] + [m \leftrightarrow n] + [a \leftrightarrow b] + [ma \leftrightarrow nb] \right).$$

The Kubo formula enables us to calculate the following response coefficients defined in the text:

$$G_{(2^{-})} = \frac{1}{3} \frac{n}{\hbar} \left[\frac{\hbar^2}{2\mu} \right]^2 \sqrt{\frac{12}{5}} \frac{\Delta}{\left(\omega^2 - \frac{12}{5}\Delta^2\right)} + \frac{4}{3} \frac{\hbar^2}{2\mu} n,$$
$$\eta_H = -\frac{2}{3} \frac{n}{\hbar} \left[\frac{\hbar^2}{2\mu} \right]^2 \sqrt{\frac{12}{5}} \frac{\Delta}{\left(\omega^2 - \frac{12}{5}\Delta^2\right)^2} \frac{\gamma B_z}{\hbar}$$

where n is the boson number density and $\gamma B_z/\hbar = g_{(2^-)}\omega_L$. Near resonance, the shear modulus is entirely

dominated by its first term,

$$G_{(2^-)} = \frac{1}{3} \frac{n}{\hbar} \left[\frac{\hbar^2}{2\mu} \right]^2 \sqrt{\frac{12}{5}} \frac{\Delta}{\left(\omega^2 - \frac{12}{5}\Delta^2\right)}$$

when $\omega^2 - \frac{12}{5}\Delta^2 \ll \Delta^2$,

and the Hall viscosity is, in terms of the shear modulus,

$$\eta_H = -2G_{(2^-)} \frac{g_{(2^-)}\omega_L}{\omega^2 - \frac{12}{5}\Delta^2}.$$