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Dynamics of a quantum phase transition in the 1D Bose-Hubbard model: excitations and correlations induced by a quench

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The ground state of the one-dimensional Bose-Hubbard model at unit filling undergoes the Mott-superfluid quantum phase transition. It belongs to the Kosterlitz-Thouless universality class with an exponential divergence of the correlation length in place of the usual power law. We present numerical simulations of a linear quench both from the Mott insulator to superfluid and back. The results satisfy the scaling hypothesis that follows from the Kibble-Zurek mechanism (KZM). In the superfluid-to-Mott quenches there is no significant excitation in the superfluid phase despite its gaplessness. Since all critical superfluid ground states are qualitatively similar, the excitation begins to build up only after crossing the critical point when the ground state begins to change fundamentally. The last process falls into the KZM framework.

I. INTRODUCTION

The study of the dynamics of phase transitions started with the question posed by Kibble^{1,2}. He noted that, in the rapidly cooling post – Big Bang Universe, phase transitions must lead to disparate local choices of the broken symmetry vacuum. The resulting mosaic of domains with independently chosen vacua will in turn precipitate formation of topological defects with observable consequences. For instance, Kibble theorized that the presently observed cosmological structures have been seeded by cosmic strings – an example of such defects.

While the original estimate¹ of defect density was incorrect, as it relied on an equilibrium argument³ (hence, densities did not depend on the phase transition rate), the key question – whether transitions leave relics such as topological defects in their wake – can be also posed in the condensed matter setting and studied in the laboratory. The theory^{4–6} developed to estimate defect density for the second order phase transition relies on their non-equilibrium nature but also on the universality class of the transition: It uses equilibrium critical exponents to predict scaling of the density of defects and other excitations with the quench timescale. The size of domains within which symmetry breaking can be coordinated is given by the size of the “sonic horizon” that estimates how far the local choice of the broken symmetry can be communicated (and plays a similar role as the causal horizon used by Kibble² to set lower bounds on defect density)

The usual estimate of the sonic horizon size relies on a power-law scaling of the relaxation time and the healing length with a distance from the critical point that is characterized by the dynamical and correlation length critical exponents z and ν . A characteristic time-scale

$\hat{t} \sim \tau_Q^{z\nu/(1+z\nu)}$ and a length-scale $\hat{\xi} \sim \tau_Q^{\nu/(1+z\nu)}$ are predicted, where the quench time τ_Q quantifies the rate of the transition. The correlation length $\hat{\xi}$ determines the number of excitations as a function of τ_Q . The Kibble-Zurek mechanism (KZM) has been confirmed by numerical simulations^{7–22} and by experiments^{23–40} in a variety of settings, with most recent results providing evidence of the KZM scaling laws^{33,34,41–47}. Refinements and extensions of KZM include phase transition in inhomogeneous systems (see 48 for a recent overview), generation of winding numbers in the limit when flux trapping is a rare event²⁹ (and KZM scaling must be suitably adjusted⁴⁹), and applications that go beyond topological defect creation (see e.g. 50–53). Recent reviews related to KZM are also available^{54–58}.

In this paper we consider a zero-temperature quantum phase transition. Despite important differences with respect to thermodynamic transitions – where thermal rather than quantum fluctuations act as seeds of symmetry breaking – the KZM can be generalized to quantum phase transitions^{22,59–76}, see also 56–58 for reviews. The quantum regime was also addressed in some of the recent experiments^{46,47,77–81}.

Recently a fully fledged scaling hypothesis was proposed^{22,82,83} as a generalization and extension of the predictive power of KZM, though some of its basic ingredients were known from the beginning^{6,53,84–86}. Since in the adiabatic limit, when $\tau_Q \rightarrow \infty$, both scales \hat{t} and $\hat{\xi}$ diverge, they should become the only relevant time and length scales in the regime of low frequencies and long wavelengths. This in turn suggests a dynamical scaling hypothesis, similar to the (static) one in equilibrium phase transitions, that during the quench all physical ob-

servables depend on a time t through the scaled time t/\hat{t} and on a distance x through the scaled distance $x/\hat{\xi}$. What makes it really powerful is the fact that the two scales are not independent: $\hat{t} \sim \hat{\xi}^z$. This space-time renormalization hypothesis was confirmed in a precise experiment⁴⁷ where it proved useful in organizing the experimental data. In the following we also find it useful in organizing our numerical results.

In this paper we consider the Mott-superfluid quantum phase transition in the 1D Bose-Hubbard model that belongs to the Kosterlitz-Thouless universality class. This problem was touched upon in Ref. 87 where it was argued that, since the correlation length diverges exponentially near the critical point, one cannot ascribe a definite scaling exponent w to $\hat{\xi} \sim \tau_Q^w$ except for $w = 1$ in the limit of exceedingly slow τ_Q that are beyond any realistic experiment. However, when the range of τ_Q is restricted to one or two orders of magnitude, then an effective scaling $\hat{\xi} \sim \tau_Q^w$ with an effective exponent $w < 1$ can be a convenient approximation. In this paper, we readdress the problem, this time with fully fledged numerical DMRG simulations. Like in a real experiment, there are limitations that restrict the range of available quench times and the KZ scaling hypothesis with effective exponents is a convenient approximation. We simulate linear quenches from the Mott insulator to superfluid, where the range of correlations builds up as the tunnelling rate between nearest-neighbor sites is turned on. The spatial profile of the correlators and their time dependence satisfy the scaling hypothesis.

We also simulate reverse linear quenches from the superfluid to Mott insulator. Somewhat surprisingly, we find the excitation during a ramp across the gapless superfluid to be negligible as compared to the excitation that builds up after crossing the critical point to the Mott phase. Apparently, the different critical superfluid ground states that are crossed by the linear ramp are similar enough for the excitation to be negligible despite their vanishing gap. What matters here is that a relatively large change of the tunneling rate during the ramp across the superfluid phase corresponds to a relatively small change of the Luttinger liquid parameter K that determines the ground state of the liquid. When measured by a distance between different ground states, the superfluid phase can be effectively identified as a single critical point. It is only after crossing to the Mott phase that the ground state begins to change fundamentally. This inevitably makes the excitations build up until their growth is halted by the opening Mott gap. The last crossover, that takes place at the time \hat{t} after crossing to the Mott phase, is the essence of the quantum KZM. Our simulations confirm this simple scenario by demonstrating that the excitation energy in the Mott phase satisfies the KZ scaling hypothesis.

The paper is organized as follows. We begin with a general discussion of the textbook version of the quantum Kibble-Zurek mechanism in section II. This generic version assumes gapful phases on both sides of the transi-

tion and a power law divergence of the correlation length at the critical point. These assumptions are relaxed in section III where we introduce the 1D Bose-Hubbard model and consider the phase transition from Mott insulator to superfluid that belongs to the Kosterlitz-Thouless universality class. Here the divergence of the correlation length on the Mott side is exponential and the whole superfluid phase is gapless and critical. We recall relevant predictions of Ref. 87 for this type of transition. In this paper these results are corroborated by numerical simulations described in sections IV and V. In section IV we describe simulations of the Mott-to-superfluid transition and in section V those of the reverse superfluid-to-Mott quench. Finally, we briefly conclude in section VI.

II. QUANTUM KIBBLE-ZUREK MECHANISM

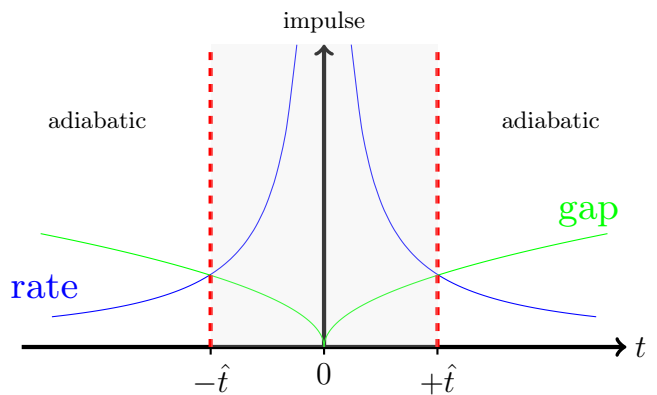


FIG. 1. As a system is driven across a generic quantum critical point with a linear ramp $\epsilon(t) = t/\tau_Q$, the energy gap between the ground state and the first relevant excited state closes like $|\epsilon|^{z\nu}$ and, at the same time, the transition rate diverges like $1/|t|$. The two are equal at $\pm\hat{t}$, where $\hat{t} \sim \tau_Q^{z\nu/(1+z\nu)}$. The evolution must be non-adiabatic (impulse in first approximation) between $-\hat{t}$ and \hat{t} .

A distance from a quantum critical point can be measured with a dimensionless parameter ϵ . The ground state of the Hamiltonian $H(\epsilon)$ undergoes a fundamental change at $\epsilon = 0$ when the correlation length in its ground state diverges like

$$\xi \sim |\epsilon|^{-\nu} \quad (1)$$

and the relevant gap closes like

$$\Delta \sim |\epsilon|^{z\nu}. \quad (2)$$

The system, initially prepared in its ground state, is driven across the critical point by a linear quench,

$$\epsilon(t) = \frac{t}{\tau_Q}, \quad (3)$$

with a quench time τ_Q .

The evolution sufficiently far from the critical point is initially adiabatic. However, the rate of change of epsilon,

$$\left| \frac{\dot{\epsilon}}{\epsilon} \right| = \frac{1}{|t|}, \quad (4)$$

diverges at the gapless critical point. Therefore, the evolution cannot be adiabatic in its neighborhood between $-\hat{t}$ and \hat{t} , see Fig. 1. Here \hat{t} is the time when the gap (2) equals the rate (4), so that:

$$\hat{t} \sim \tau_Q^{z\nu/(1+z\nu)}. \quad (5)$$

Just before the adiabatic-to-non-adiabatic crossover at $-\hat{t}$, the state of the system is still approximately the adiabatic ground state at $\epsilon = -\hat{\epsilon}$, where

$$\hat{\epsilon} = \frac{\hat{t}}{\tau_Q} \simeq \tau_Q^{-1/(1+z\nu)}, \quad (6)$$

with a correlation length

$$\hat{\xi} \sim \hat{\epsilon}^{-\nu} \sim \tau_Q^{\nu/(1+z\nu)}. \quad (7)$$

In a first-order impulse approximation, this state “freezes out” at $-\hat{t}$ and does not change until \hat{t} . At \hat{t} the frozen state is no longer the ground state but an excited state with a correlation length $\hat{\xi}$. It is the initial state for the adiabatic process that follows after \hat{t} .

No matter how quantitatively accurate is the above “freeze-out scenario”, this simple scaling argument establishes $\hat{\xi}$ and \hat{t} , interrelated via

$$\hat{t} \sim \hat{\xi}^z, \quad (8)$$

as the relevant scales of length and time, respectively. What is more, in the adiabatic limit, when $\tau_Q \rightarrow \infty$, both scales diverge becoming the unique scales in the regime of long-wavelength and low-frequency. In analogy to static critical phenomena, this uniqueness suggests a scaling hypothesis^{22,82,83}:

$$\langle \psi(t) | O(x) | \psi(t) \rangle = \hat{\xi}^{-\Delta_O} F_O \left(t/\hat{\xi}^z, x/\hat{\xi} \right). \quad (9)$$

Here $|\psi(t)\rangle$ is the quantum state during the quench, $O(x)$ is an operator depending on a distance x , Δ_O is its scaling dimension, and F_O its scaling function. The same conclusions about scaling are reached when one follows the narrative based on the sonic horizon⁶.

III. KZ MECHANISM IN THE KOSTERLITZ-THOULESS TRANSITION

The power laws (1,2) are not directly applicable in the Kosterlitz-Thouless (KT) transition^{88–90}, where, on the disordered/Mott side of the transition, the correlation length’s divergence is exponential:

$$\xi = \xi_0 \exp(2a/\sqrt{|\epsilon|}). \quad (10)$$

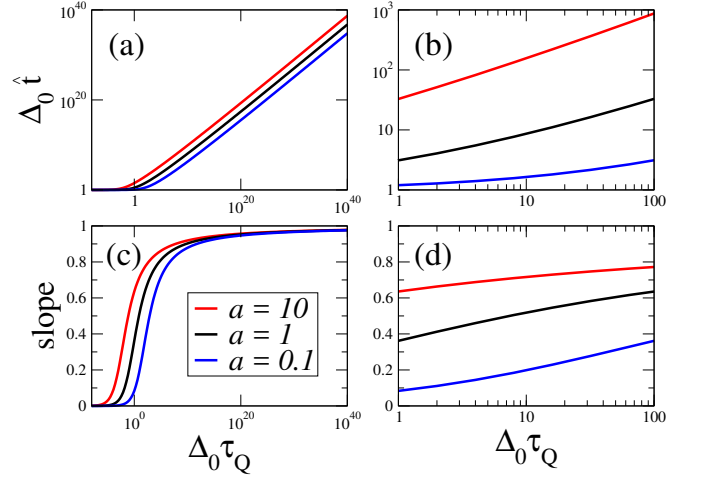


FIG. 2. In (a) log-log scale a generic KZ power-law $\hat{t} \propto \tau_Q^{\nu z/(1+\nu z)}$ in (5) becomes a linear function $\log_{10} \Delta_0 \hat{t} = \frac{\nu z}{1+\nu z} \log_{10} \Delta_0 \tau_Q + \text{const}$, where Δ_0 is a microscopic energy scale. For comparison, in (a), we plot \hat{t} for the Kosterlitz-Thouless transition in the 1D Bose-Hubbard model (14) in function of τ_Q over many decades of the argument. This function becomes linear only asymptotically for $\tau_Q \rightarrow \infty$, but it may appear linear locally, i.e., in a range of one or two decades. Indeed, in (b), we focus on the narrow range of $\Delta_0 \tau_Q = 10^{0 \dots 2}$ that is small enough for a realistic cold-atom experiment. These plots can be reasonably approximated by linear functions, especially when experimental error bars are present. In (c), we plot a local slope $d \log_{10}(\Delta_0 \hat{t}) / d \log_{10}(\Delta_0 \tau_Q)$ of the log-log plot in panel (a) in function of $\Delta_0 \tau_Q$. The slope 1 in Eq. (11) is achieved asymptotically but only for τ_Q that are “cosmological” in magnitude. In (d), a focus on the realistic τ_Q shows that the local slope can be significantly less than 1.

Here $a \simeq 1$ and ξ_0 is a microscopic scale of length. This faster-than-polynomial divergence can be captured by stating that $\nu = \infty$, see *e.g.* 91, but it may tempt one to misuse Eqs. (5, 6, 7) by inserting $\nu = \infty$ together with $z = 1$ to obtain:

$$\hat{t} \sim \tau_Q^1, \quad \hat{\epsilon} \sim \tau_Q^0, \quad \hat{\xi} \sim \tau_Q^1. \quad (11)$$

As shown in Ref. 87, these equations are valid asymptotically for $\tau_Q \rightarrow \infty$ but this asymptote is achieved for unrealistically slow τ_Q (and, hence, astronomically large $\hat{\xi}$ of the order of kilometers⁸⁷). Below we briefly recount the argument.

In the 1D Bose-Hubbard model at commensurate filling, where $z = 1$, the gap $\Delta \sim \xi^{-z}$ on the Mott-insulator side of the transition closes like

$$\Delta = \Delta_0 \exp(-2a/\sqrt{|\epsilon|}), \quad (12)$$

where Δ_0 is a microscopic energy scale. For the linear ramp (3), driving a quench from the Mott insulator to superfluid, this gap equals the rate (4) at $t = -\hat{t}$ when

$$\Delta_0 \exp(-2a/\sqrt{\hat{t}/\tau_Q}) = 1/\hat{t}. \quad (13)$$

A solution of this transcendental equation is

$$\hat{t} = \tau_Q \frac{a^2}{\mathcal{W}^2(a\sqrt{\Delta_0\tau_Q})}, \quad (14)$$

where \mathcal{W} is the Lambert function⁹². The above solution is plotted for different values of $a \simeq 1$ in Figs. 2a,b. A similar relation has been derived for the first time and confirmed by numerical simulations in a dissipative classical model⁹³. It has been also tested in a recent experiment⁹⁴.

Figure 2 shows that the exponent of unity for the dependence of \hat{t} on τ_Q in Eq. (11) is attained only for exceedingly slow quenches that are unlikely to be experimentally or numerically accessible. For any reasonably slow quenches the effective exponent would be significantly less than 1.

The above argument determines the correlation length $\hat{\xi}$ imprinted on the quantum state on the Mott side of the transition. It characterizes the excited state when the quench ramp enters the superfluid phase. In this critical phase the imprinted correlations are spreading with a velocity limited by the speed of quasiparticles.

IV. QUENCH FROM MOTT INSULATOR TO SUPERFLUID

The 1D Bose-Hubbard model at the commensurate filling of 1 particle per site is described by a Hamiltonian

$$H = -J \sum_{s=1}^{L-1} (b_s^\dagger b_{s+1} + b_{s+1}^\dagger b_s) + \frac{U}{2} \sum_{l=1}^L n_l(n_l - 1), \quad (15)$$

where J is the hopping rate and U is the on-site interaction strength. The Kosterlitz-Thouless quantum phase transition from the Mott-insulator to superfluid takes place at $J_c = 0.29$ (see *e.g.* Ref. 97). We consider a quench driven by a linear ramp

$$J(t) = J_c \begin{cases} 0, & \text{for } t \leq -\tau_Q, \\ 1 + t/\tau_Q, & \text{for } t > -\tau_Q. \end{cases} \quad (16)$$

from $J(-\tau_Q) = 0$ to $J(\tau_Q) = 2J_c$. The initial state is the ground state at $J = 0$:

$$|\psi(0)\rangle = |1, 1, 1, \dots, 1\rangle. \quad (17)$$

Figure 3(a) shows that there is a finite density of excitation energy W/L above the adiabatic ground state during a quench, hence the evolution with the ramp is not adiabatic. Furthermore, the system seems to be excited in two stages by two different mechanisms.

The early oscillations visible at small J are excited by the discontinuous time derivative of the ramp at $J = 0$ that is proportional to $1/\tau_Q$. Since this sudden initial jolt excites states well above the Mott gap with a probability proportional to τ_Q^{-2} , the amplitude of the early excitation is roughly $W/L \propto \tau_Q^{-2}$, as can be demonstrated by

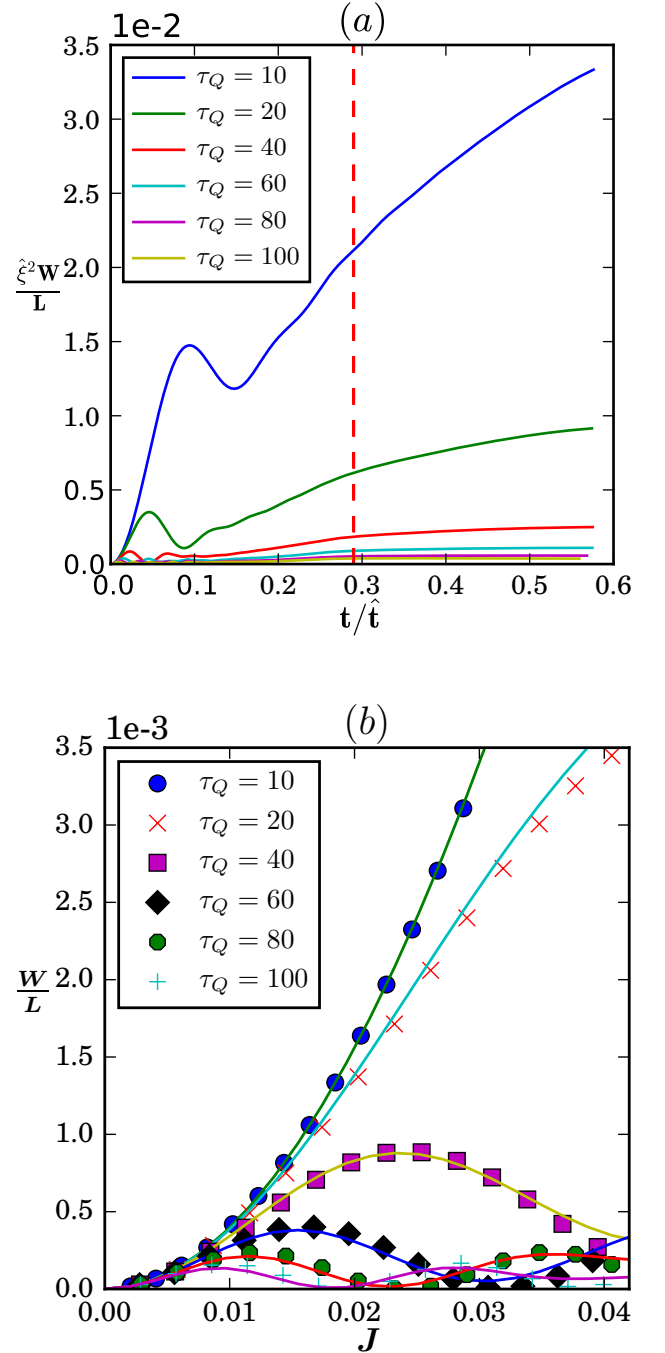


FIG. 3. In (a), density of excitation energy W/L in the center of $L = 100$ sites in function of the hopping frequency J - results from DMRG⁹⁵ simulations for different τ_Q . The early excitation visible at small J originates from the discontinuous slope of the ramp (16) at $J = 0$. This initial excitation is suppressed adiabatically when J is getting closer to $J_c = 0.29$ (the vertical dashed line) and the gap is closing. The following excitation that begins to grow before J_c is attributed to the KZ mechanism. For $\tau_Q = 10$ the initial excitation is too strong and the KZ excitation begins too early for the two mechanisms to be clearly separated in J . In (b), a focus on small J where DMRG (solid lines) can be compared with the Bogoliubov theory for doublons and holons⁹⁶ (data points). The agreement is better for faster τ_Q where the state remains closer to the initial Fock state (17) and has lower density of doublons and holons.

the quantum perturbation theory⁹⁸. For small enough J they compare well with predictions of the Bogoliubov doublon-holon model⁹⁶, see figure 3(b). For larger J , as their energy gap closes with J approaching J_c , the early oscillations are adiabatically suppressed. Closer to J_c the KZ mechanism steps in, see Fig. 4.

At the earlier stages of the project we attempted to mitigate the effect of the initial jolt by initiating the quench more smoothly. For rapid quenches such efforts proved only partly successful. However, similarly as for the linear ramp in Fig. 3(a), long quench times suppress the effect of these early “jolt” excitations compared to these caused by the crossing of the critical region.

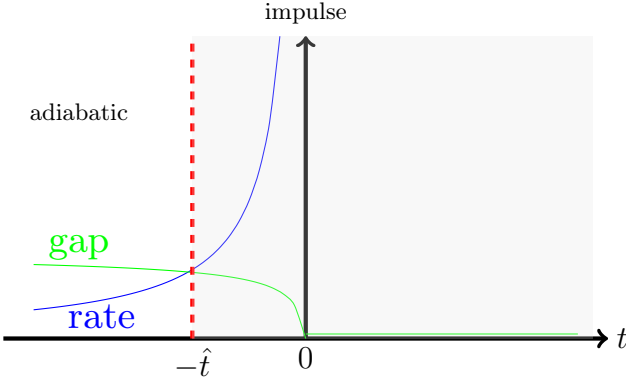


FIG. 4. In (a) quench from the Mott-insulator to superfluid, driven by the linear ramp (16), the initial evolution before $-\hat{t}$ is adiabatic, then it becomes approximately impulse.

The quench times in our numerical simulations are far below the “astronomical” standards that would demand $\hat{\xi}$ of the order of kilometers to approach the scaling $\hat{\xi} \sim \tau_Q$ in Eq. (11). Limited by the system size L they span a narrow range of magnitude. Therefore, we can assume a phenomenological power-law for the ground-state correlation length in the Mott phase:

$$\xi \sim \epsilon^{-\nu_L}. \quad (18)$$

Given the exponential divergence (10), we expect the effective exponent ν_L to be large. To account for finite-size effects, we allow for its dependence on L . Furthermore, we found that the data can be accurately parametrized by letting the gap scale with an effective dynamical exponent:

$$\Delta \sim \xi^{-z_L} \sim \epsilon^{z_L \nu_L}. \quad (19)$$

We expect to recover the exact $z_\infty = 1$ for sufficiently large L .

The correlation length in the ground state is obtained from the best fit to the tail of the correlator:

$$C(x) = \langle b_s^\dagger b_{s+x} \rangle \sim \frac{e^{-x/\xi}}{x^{1/4}}. \quad (20)$$

The lengths and gaps for different ϵ and L are collected in Fig. 5(a,b). They fit well the phenomenological power-laws (18,19). As a self-consistency check, figure 5(c)

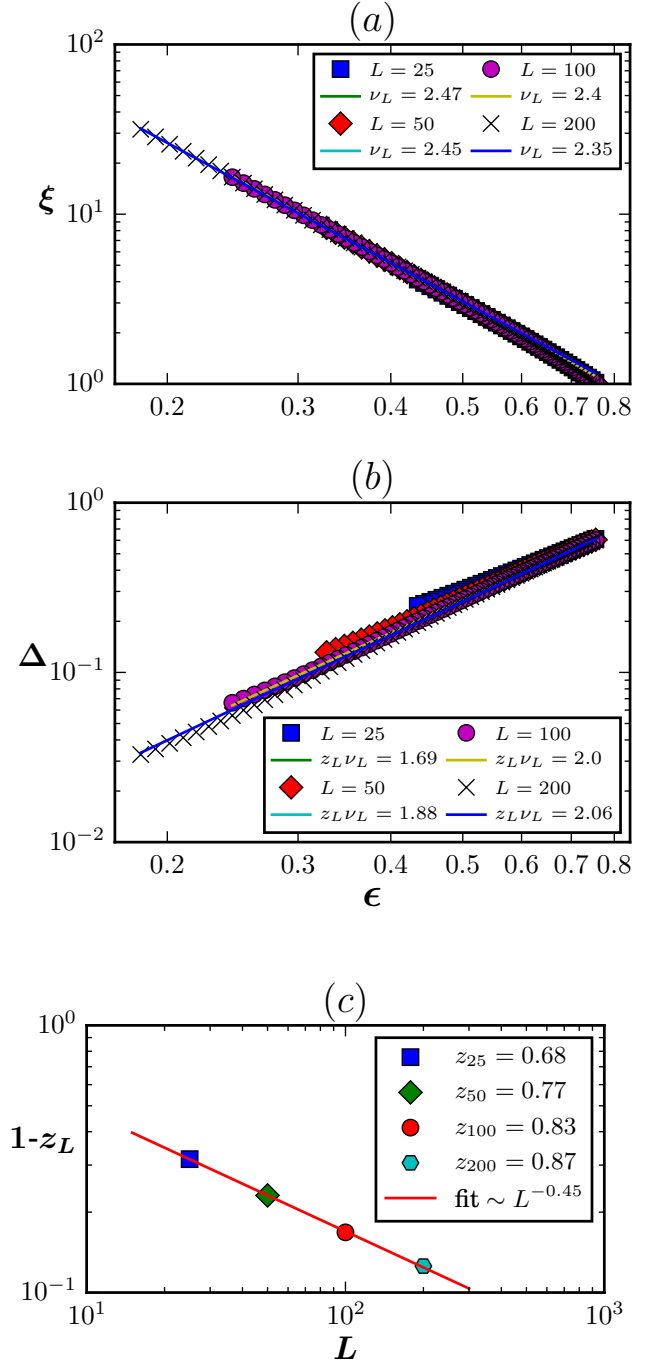


FIG. 5. In (a), a log-log plot of the correlation length ξ in function of ϵ for system sizes $L = 25, 50, 100, 200$. For each L , the data are fitted with a power law $\xi \sim \epsilon^{-\nu_L}$ in the range of ϵ where $1 < \xi < L/6$. The best fits are shown by the solid lines and the best exponents ν_L are listed in the legend. In (b), a log-log plot of the energy gap Δ in function of ϵ for $L = 25, 50, 100, 200$. For each L , the data are fitted with a power law $\Delta \sim \epsilon^{z_L \nu_L}$ in the range of ϵ where $1 < \xi_L < L/6$. Here ν_L are the exponents obtained in panel a. The best fits are shown by the solid lines and the best exponents $z_L \nu_L$ are listed in the legend. In (c), a log-log plot of $1 - z_L$ in function of L . This plot shows the decay of $z_L \rightarrow 1$ with increasing L . It can be fitted with $1 - z_L \sim L^{-0.45}$ suggesting a power law approach of z_L towards $z = 1$ with increasing L .

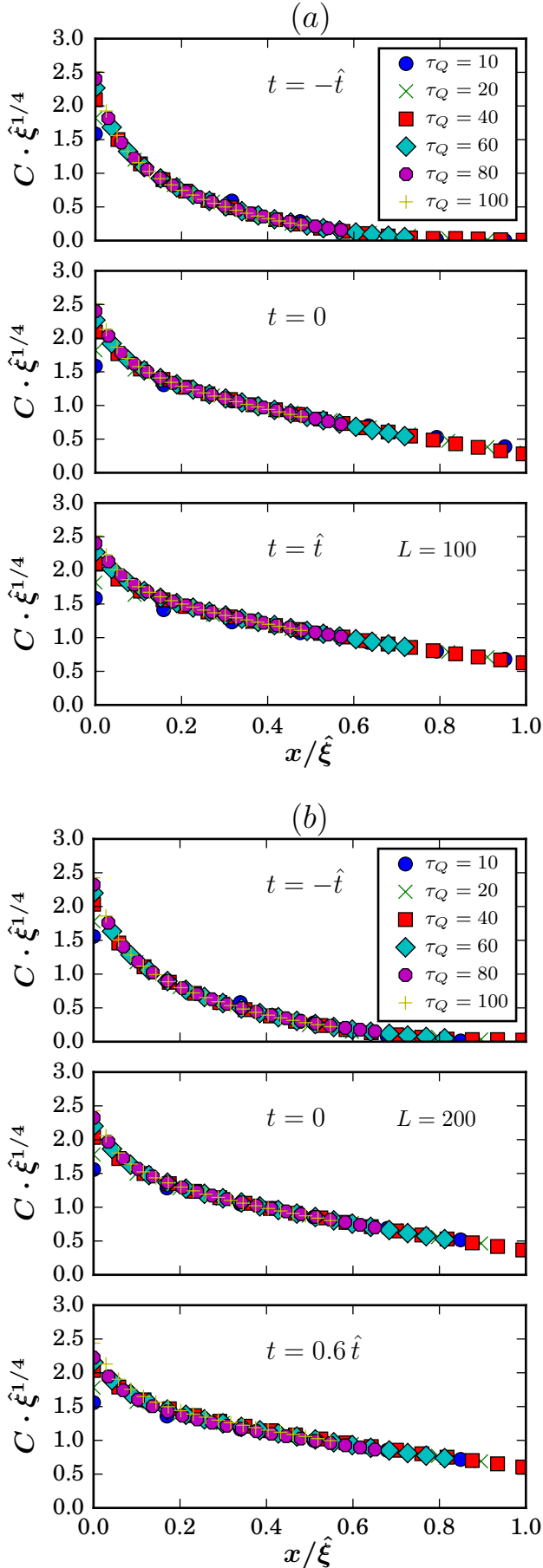


FIG. 6. In (a), linear plots of the scaled correlation functions $\hat{\xi}^{1/4}C(t, x)$ in function of the scaled distance $x/\hat{\xi}$ – measured in the bulk of $L = 100$ lattice sites – at the scaled times $t/\hat{t} = -1, 0, 1$. With increasing τ_Q the plots collapse to the scaling function $F_C(t/\hat{t})$. In (b), the same as in (a) but for $L = 200$ and $t/\hat{t} = -1, 0, 0.6$.

shows how the dynamical exponent z_L decays to 1 with increasing L . For each L we use the effective exponents to find the KZ scales

$$\hat{t} \sim \tau_Q^{\frac{z_L \nu_L}{1+z_L \nu_L}}, \quad \hat{\xi} \sim \tau_Q^{\frac{\nu_L}{1+z_L \nu_L}}. \quad (21)$$

These scales are applied to verify the KZ scaling hypothesis,

$$C(t, x) = \hat{\xi}^{-1/4} F_C \left(t/\hat{\xi}^{z_L}, x/\hat{\xi} \right), \quad (22)$$

in Fig. 6. The plots for different τ_Q collapse demonstrating validity of the hypothesis.

The collapsed plots in Fig. 6 depend on the scaled time t/\hat{t} . This effect reveals that the impulse approximation is not quite correct: since the correlations are spreading, the state cannot be frozen. This behavior suggests a “sonic horizon” paradigm we have also noted earlier. Nevertheless, as discussed and demonstrated by an exact solution of the quantum Ising chain²², the scaling hypothesis still holds because the velocity $2\hat{v}$ of the spreading must also be a combination of the two KZ scales: $\hat{v} \sim \hat{\xi}/\hat{t}$. For a linear dispersion at the critical point, $z = 1$, \hat{v} is expected to be bounded from above by twice the speed c_0 of quasiparticles at the critical point. In figure 7(a), we attempt an estimate of the spreading velocity $2\hat{v}$. Since accurate tails of the correlators cannot be accessed on a finite lattice, our estimate is not robust but at least it is less than the $2c_0$ estimated from the quasiparticle dispersion in figure 7(b).

Neglecting remnants of the early excitation, a distribution of quasiparticle excitations should have a scaling form $f(t/\hat{t}, \hat{\xi}k)$. In the KZ regime close to J_c , their dispersion can be approximated by the gapless linear dispersion at the critical point, $\omega_k \approx c_0|k|$, and the excitation energy (also sometimes referred to as work^{99,100}) density should satisfy a scaling hypothesis:

$$W/L = \int \frac{dk}{2\pi} c_0|k| f(t/\hat{t}, \hat{\xi}k) = \hat{\xi}^{-2} F_W(t/\hat{t}). \quad (23)$$

Here F_W is a scaling function. In particular, at J_c (where $t/\hat{t} = 0$) we expect $W \sim \hat{\xi}^{-2}$. This scaling law and the more general scaling hypothesis (23) are confirmed by the data in Fig. 8.

Finally, we considered the entropy of entanglement between a block of l sites at the end of the chain and the rest of it. Ignoring boundary effects, the entropy should satisfy a scaling hypothesis:

$$\frac{S(t, l)}{\frac{c}{6} \log \hat{\xi} f(t/\hat{t})} = F_S \left(t/\hat{t}, l/\hat{\xi} \right). \quad (24)$$

Here f and F_S are scaling functions and c is the central charge $c = 1$ at the Kosterlitz-Thouless transition. This hypothesis is tested in Fig. 9. Except for $t = -\hat{t}$, where the state is still close to the ground state, the fitted c , though close to 1, are not quite satisfactory, but we have to bear in mind that the range of τ_Q at hand is too narrow

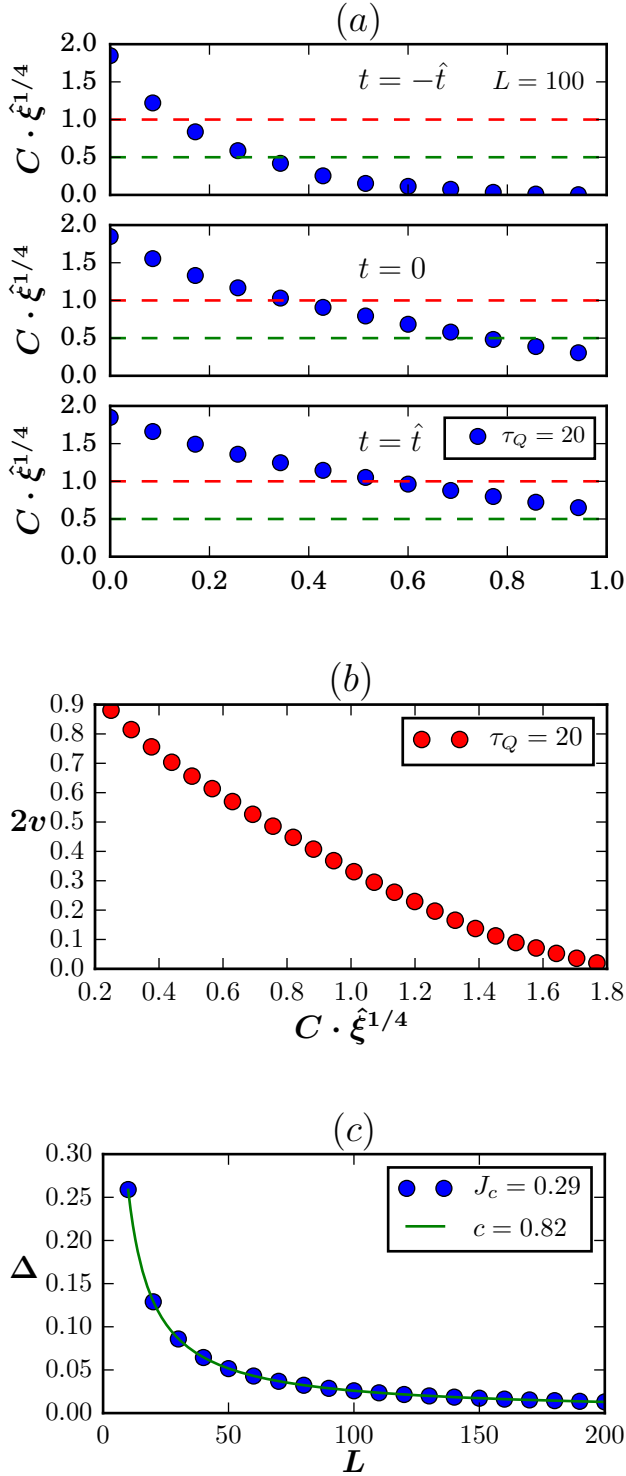


FIG. 7. In (a), scaled correlation functions for $\tau_Q = 20$ in function of the scaled distance $x/\hat{\xi}$ measured in the bulk of $L = 100$ lattice sites. The three plots show the correlator at scaled times: $t/\hat{t} = -1, 0, 1$. A cut by the horizontal dashed line at the level 1 suggests that the correlations are spreading and the sonic horizon grows at a velocity $2\hat{v} = 0.34$, but a similar cut at 0.5 suggests $2\hat{v} = 0.63$. Longer tails of the correlator, that could be probed at lower cut levels, are not available on $L = 100$ sites. In (b), spreading velocity $2\hat{v}$ in function of scaled correlation values (the cut level). In (c), the gap Δ at J_c in function of system size L . The best fit $\Delta = c_0 \frac{\pi}{L}$ yields $2c_0 = 1.64$ as the speed of Luttinger quasiparticles. As expected, $2c_0$ is greater than our crude estimates of $2\hat{v}$.

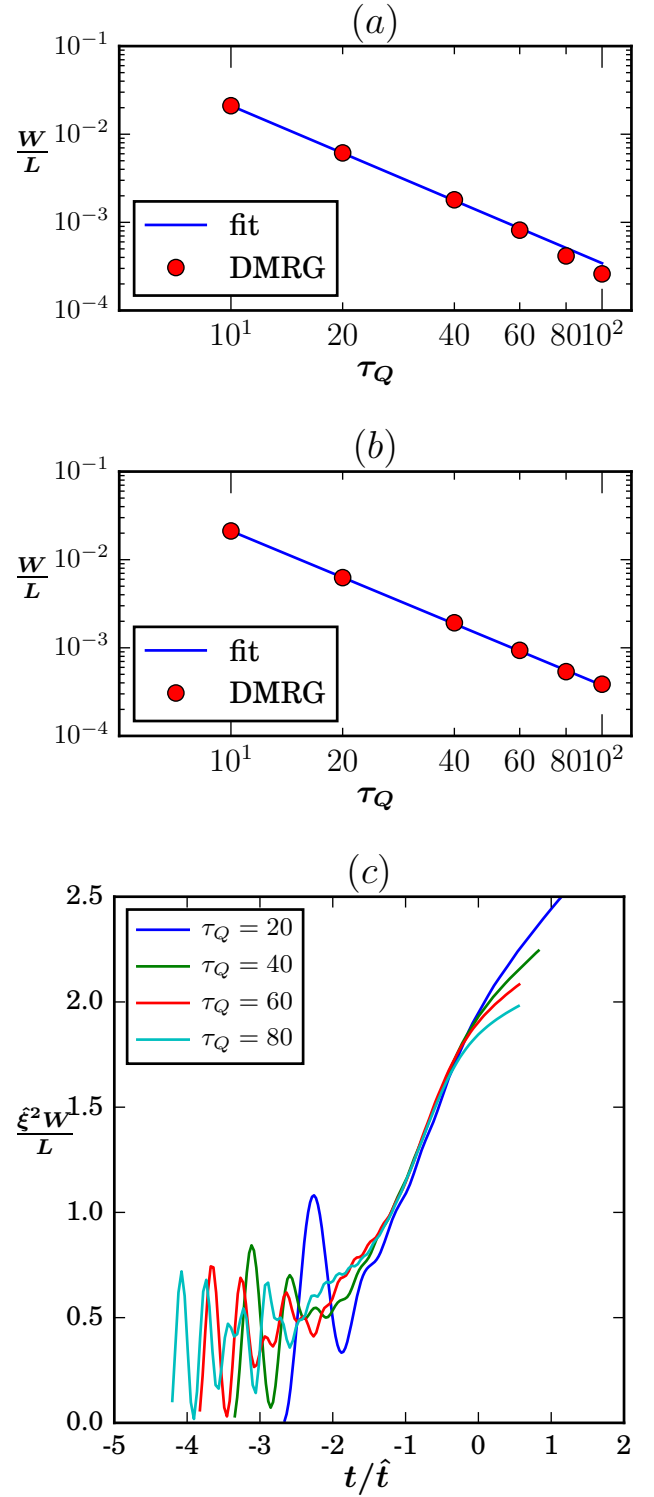


FIG. 8. In (a), log-log plot of the excitation energy density W/L in function of the quench time τ_Q at the center of $L = 100$ lattice at the time $t = 0$ when the quench is crossing the critical point. The solid line is the best fit $W \sim \tau_Q^{-1.79}$ consistent with $W \sim \tau_Q^{-1.80}$ predicted by Eqs. (21,23) with the exponents fitted in Fig. 5In (b), the same as in (a) but for $L = 200$ sites. The solid line is the best fit $W \sim \tau_Q^{-1.75}$ consistent with $W \sim \tau_Q^{-1.77}$ predicted by Eqs. (21,23) with the exponents fitted in Fig. 5b. In (c), Scaled excitation energy density in function of scaled time for $L = 100$ sites. The plots collapse to the scaling function F_W in the KZ regime close to J_c .

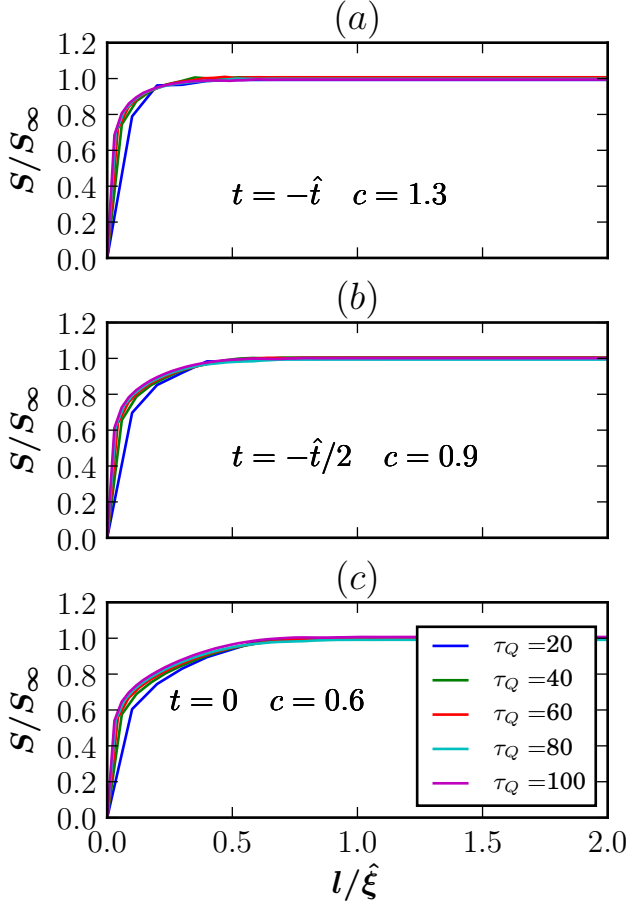


FIG. 9. Entanglement entropy $S(t, l)$ between a block of l sites at the end and the rest of the chain of $L = 200$ sites. Panel (a), (b), and (c) show the entropy at $t = -\hat{t}$, $-\hat{t}/2$, and 0, respectively. All plots are in function of the scaled block size, $l/\hat{\xi}$, and the entropy is divided by $S_\infty(t/\hat{t}) = \frac{c}{6} \log \hat{\xi} f$. This logarithmic function of τ_Q is obtained as the best fit to $S(t/\hat{t}, L/2)$ with fitting parameters f and c . The best central charge c for $t = -\hat{t}$, $-\hat{t}/2$, 0 is $c = 1.3, 0.9, 0.6$, respectively. The plots collapse demonstrating the scaling hypothesis in Eq. (24).

for a better fit with a logarithmic function. However, we find that – in accordance with the scaling hypothesis – the plots collapse to a scaling function F_S after rescaling the block size by $\hat{\xi}$.

V. QUENCH FROM SUPERFLUID TO MOTT INSULATOR

In this section we reverse the quench. Now the linear ramp begins when $t = -\tau_Q$ deep in the superfluid phase at $J = 2J_c$ and ends at $J = 0$ when $t = \tau_Q$:

$$J(t) = J_c (1 - t/\tau_Q). \quad (25)$$

The initial ground state has a quasi-long-range order characterized by a power-law decay of its correlation function $C^{(GS)}(x)$. Small excitations in the superfluid – described by a Luttinger liquid – are gapless, hence it may be tempting to treat the whole evolution in the superfluid phase as impulse. However, different critical ground states in the superfluid phase are sufficiently similar to each other for a significant excitation to be postponed until after the ramp crosses the boundary with the Mott phase where the gap begins to open and the ground state begins to change fundamentally. From this perspective, the similarity between different superfluid critical states means that on the superfluid side of the transition there is little difference between the impulse and adiabatic approximations. The whole superfluid critical phase can be effectively collapsed to the Mott critical point.

Indeed, at $t = 0$ – when the quench is leaving the superfluid at J_c – the correlator $C(t, x)$ quickly converges with increasing τ_Q to the correlator in the adiabatic ground state $C^{(GS)}(x) \sim x^{-1/4}$ at J_c , see figure 10(a). This fast convergence is consistent with the quick decay of the excitation energy density shown in Fig. 10(b). Its decay is steep, almost as steep as predicted for a linear ramp in a Luttinger liquid¹⁰¹: $W \sim \tau_Q^{-2} \log(c_0 \tau_Q)$ with $c_0 \simeq 1$ standing for the speed of Luttinger quasiparticles. It turns out to be steep enough for the excitation in the superfluid phase to give negligible contribution to the final excitation deep in the Mott phase, see Fig. 11. Its panel (a) shows the final excitation energy density after the full ramp from $J = 2J_c$ to $J = 0$ and panel (b) after a shorter ramp from $J = J_c$ to $J = 0$. The two panels are almost indistinguishable and in both the excitation energy density is two orders of magnitude higher than in Fig. 10(b). Therefore, the final excitation after the full ramp originates almost exclusively from the evolution in the Mott phase when the gap opens and the ground state undergoes a fundamental change.

The Mott part of the full ramp, from J_c to 0, falls within the KZ framework, as shown schematically in figure 12. The evolution in the Mott phase crosses over from impulse to adiabatic near $\hat{t} \sim \tau_Q^{\frac{z_L \nu_L}{1+z_L \nu_L}}$ when the correlation length in the ground state is $\hat{\xi} \sim \tau_Q^{\frac{\nu_L}{1+z_L \nu_L}}$. As long as quasiparticle excitations can be considered non-interacting, their distribution $f(t/\hat{t}, \hat{\xi}k)$ satisfies the KZ scaling hypothesis. This scaling form has at least two important consequences.

Before \hat{t} , when the quasiparticle dispersion can be approximated by its critical form $\omega_k = c_0 k$, the excitation energy density should conform to the scaling law (23). Indeed, the scaled plots in Fig. 13 demonstrate a collapse to the scaling function F_W . As predicted, the collapse is perfect up to \hat{t} .

In the adiabatic stage after \hat{t} , the quasiparticle distribution freezes out, $f(t/\hat{t}, \hat{\xi}k) = f(\hat{\xi}k)$, and their dispersion ω_k for the “excited” k between $\pm \hat{\xi}^{-1}$ can be approximated by the finite gap ω_0 . Consequently, the final

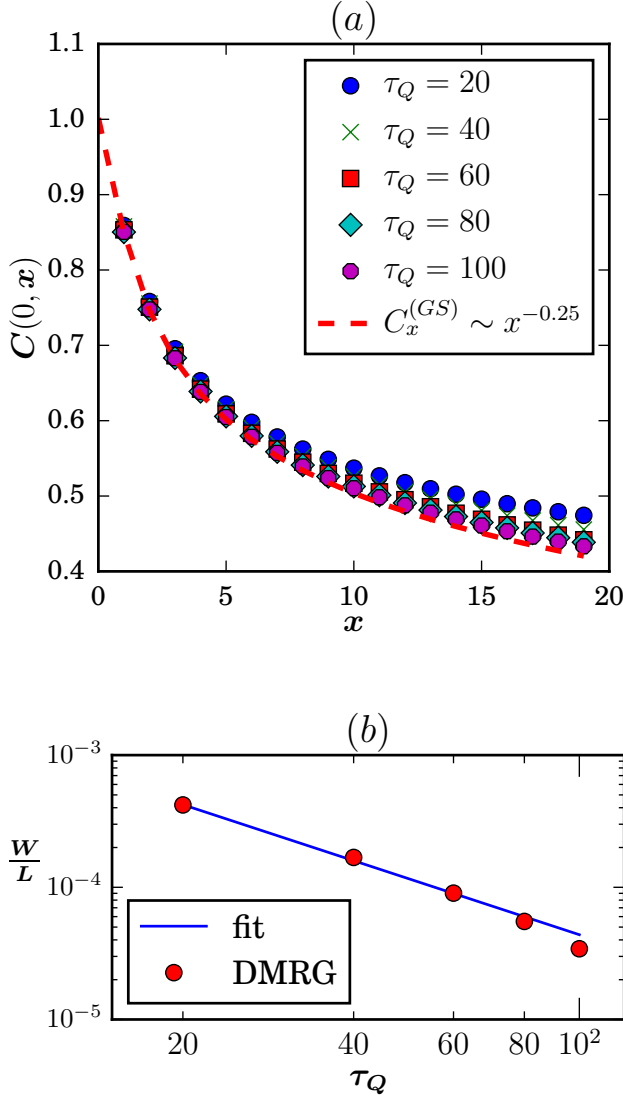


FIG. 10. In (a), correlation functions $C(0, x)$ at the critical point, $J = J_c$, during a linear quench (25) from the superfluid at $J = 2J_c$ to the Mott insulator at $J = 0$. With increasing τ_Q they converge to the adiabatic ground state correlator with a power-law tail: $C_x^{(GS)} \sim x^{-1/4}$. In (b), log-log plot of the corresponding excitation energy density W/L at the center of $L = 100$ lattice sites. These data are fitted with a power law $W \sim \tau_Q^{-1.41}$.

excitation energy density at $J = 0$ should scale as

$$\frac{W}{L} = \int_{-\pi}^{\pi} \frac{dk}{2\pi} \omega_k f(\hat{\xi}k) \sim \hat{\xi}^{-1}. \quad (26)$$

Since in this regime the excited quasiparticles are approximately dispersionless, $\omega_k \approx \omega_0$, the excitation energy density is simply proportional to their density $\hat{\xi}^{-1}$. With our best fits for ν_L and z_L on the $L = 100$ lattice we obtain $W \sim \tau_Q^{-0.80}$. This is roughly consistent with the best fit $W \sim \tau_Q^{-0.96}$ in Fig. 11(b).

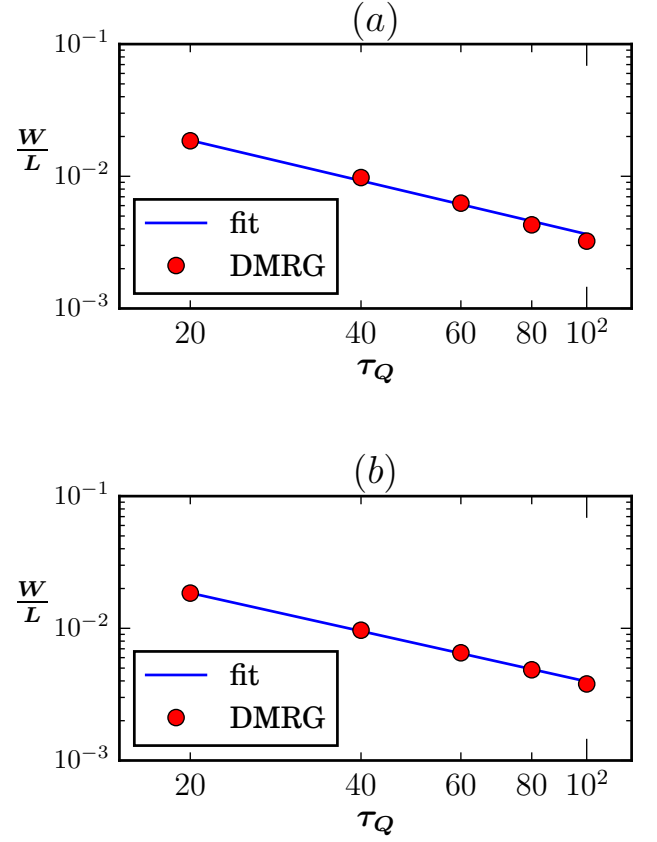


FIG. 11. In (a), the excitation energy density W/L in function of the quench time τ_Q at $J = 0$ after the full linear ramp, $J(t) = J_c (1 - t/\tau_Q)$, starting deep in the superfluid phase at $J = 2J_c$. In (b), the same as in (a) but after a half-ramp starting from the critical point $J = J_c$. The data in (a) and (b) are almost the same as if the initial half of the full ramp in the superfluid phase did not contribute to the final excitation energy deep in the Mott phase.

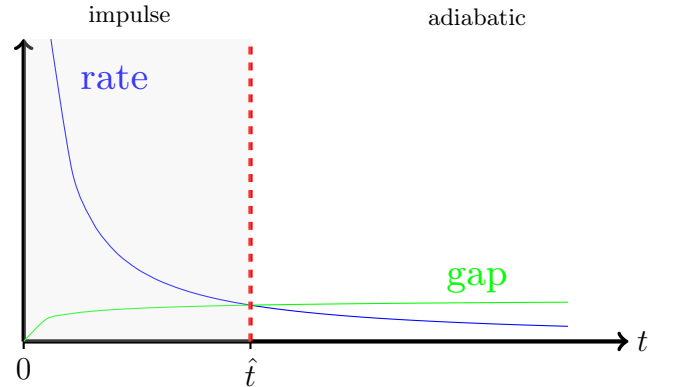


FIG. 12. A ramp from the critical point at $J = J_c$ to $J = 0$ deep in the Mott phase. The evolution crosses over from impulse to adiabatic near \hat{t} .

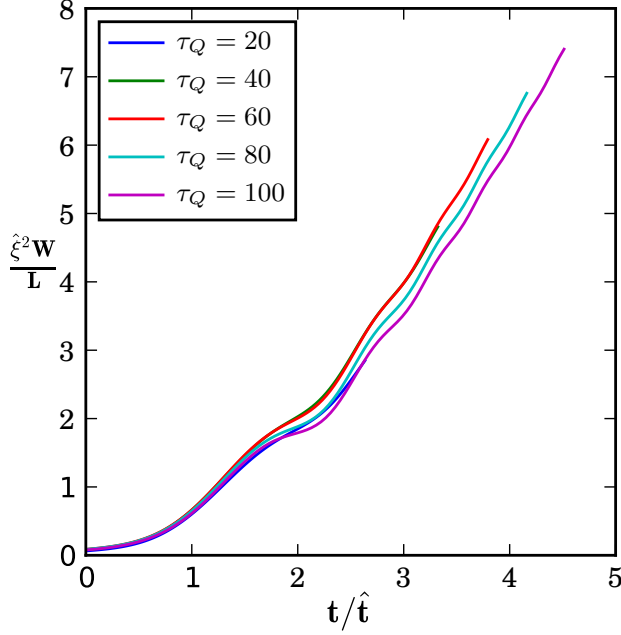


FIG. 13. Scaled excitation energy density $\hat{\xi}^2 W / L$ in function of scaled time t / \hat{t} for different quench times τ_Q . The plots collapse to the scaling function $F_W(t / \hat{t})$ in Eq. (23).

VI. CONCLUSION

In a linear quench from Mott insulator to superfluid the excitation energy density, the entropy of entanglement, and the correlations – that build up as the system is crossing the critical point – satisfy the KZ scaling hypothesis with effective power laws accurate for a limited range of quench times⁸⁷. In particular, the range of correlations scales with an effective power of the transition time.

All superfluid ground states are qualitatively similar. Therefore, in a reverse quench from superfluid to Mott insulator the excitation in the gapless superfluid turns out to be negligible as compared to the excitation that begins to build up just after crossing the critical point when the gap opens and the ground state begins to change fundamentally. The last excitation also falls into the KZ framework. The final excitation energy deep in the Mott phase – proportional to the number of empty and doubly-occupied sites – decays with an effective power of the quench time.

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- ¹ T. W. B. Kibble, *J. Phys. A: Math. Gen* **9**, 1387 (1976).
- ² T. W. B. Kibble, *Phys. Rep.* **67**, 183 (1980).
- ³ M. Hindmarsh, *J. Phys. A: Math. Theor.* **49**, 411001 (2016).
- ⁴ W. H. Zurek, *Nature* **317**, 505 (1985).
- ⁵ W. H. Zurek, *Acta Phys. Pol. B* **24**, 1301 (1993).
- ⁶ W. H. Zurek, *Phys. Rep.* **276**, 177 (1996).
- ⁷ P. Laguna and W. H. Zurek, *Phys. Rev. Lett.* **78**, 2519 (1997).
- ⁸ A. Yates and W. H. Zurek, *Phys. Rev. Lett.* **80**, 5477 (1998).
- ⁹ J. Dziarmaga, P. Laguna, and W. H. Zurek, *Phys. Rev. Lett.* **82**, 4749 (1999).
- ¹⁰ N. D. Antunes, L. M. A. Bettencourt, and W. H. Zurek, *Phys. Rev. Lett.* **82**, 2824 (1999).
- ¹¹ L. M. A. Bettencourt, N. D. Antunes, and W. H. Zurek, *Phys. Rev. D* **62**, 065005 (2000).
- ¹² W. H. Zurek, L. M. A. Bettencourt, J. Dziarmaga, and N. D. Antunes, *Acta Phys. Pol. B* **31**, 2937 (2000).
- ¹³ A. del Campo, G. De Chiara, G. Morigi, M. B. Plenio, and A. Retzker, *Phys. Rev. Lett.* **105**, 075701 (2010).
- ¹⁴ G. D. Chiara, A. del Campo, G. Morigi, M. B. Plenio, and A. Retzker, *New J. Phys.* **12**, 115003 (2010).

- ¹⁵ E. Witkowska, P. Deuar, M. Gajda, and K. Rzażewski, *Phys. Rev. Lett.* **106**, 135301 (2011).
- ¹⁶ H. Saito, Y. Kawaguchi, and M. Ueda, *Phys. Rev. A* **76**, 043613 (2007).
- ¹⁷ J. Dziarmaga, J. Meisner, and W. H. Zurek, *Phys. Rev. Lett.* **101**, 115701 (2008).
- ¹⁸ R. Nigmatullin, A. del Campo, G. De Chiara, G. Morigi, M. B. Plenio, and A. Retzker, *Phys. Rev. B* **93**, 014106 (2016).
- ¹⁹ A. Das, J. Sabbatini, and W. H. Zurek, *Sci. Rep.* **2**, 352 (2012).
- ²⁰ P. M. Chesler, A. M. García-García, and H. Liu, *Phys. Rev. X* **5**, 021015 (2015).
- ²¹ J. Sonner, A. del Campo, and W. H. Zurek, *Nature Communications* **6**, 7406 EP (2015), article.
- ²² A. Francuz, J. Dziarmaga, B. Gardas, and W. H. Zurek, *Phys. Rev. B* **93**, 075134 (2016).
- ²³ I. Chuang, R. Durrer, N. Turok, and B. Yurke, *Science* **251**, 1336 (1991).
- ²⁴ M. J. Bowick, L. Chandar, E. A. Schiff, and A. M. Srivastava, *Science* **263**, 943 (1994).
- ²⁵ V. M. H. Ruutu, V. B. Eltsov, A. J. Gill, T. W. B. Kibble, M. Krusius, Y. G. Makhlin, B. Placais, G. E. Volovik, and W. Xu, *Nature* **382**, 334 (1996).

- ²⁶ C. Bauerle, Y. M. Bunkov, S. N. Fisher, H. Godfrin, and G. R. Pickett, *Nature* **382**, 332 (1996).
- ²⁷ R. Carmi, E. Polturak, and G. Koren, *Phys. Rev. Lett.* **84**, 4966 (2000).
- ²⁸ R. Monaco, J. Mygind, and R. J. Rivers, *Phys. Rev. Lett.* **89**, 080603 (2002).
- ²⁹ R. Monaco, J. Mygind, R. J. Rivers, and V. P. Koshelets, *Phys. Rev. B* **80**, 180501 (2009).
- ³⁰ A. Maniv, E. Polturak, and G. Koren, *Phys. Rev. Lett.* **91**, 197001 (2003).
- ³¹ L. E. Sadler, J. M. Higbie, S. R. Leslie, M. Vengalattore, and D. M. Stamper-Kurn, *Nature* **443**, 312 (2006).
- ³² D. Golubchik, E. Polturak, and G. Koren, *Phys. Rev. Lett.* **104**, 247002 (2010).
- ³³ S. C. Chae, N. Lee, Y. Horibe, M. Tanimura, S. Mori, B. Gao, S. Carr, and S.-W. Cheong, *Phys. Rev. Lett.* **108**, 167603 (2012).
- ³⁴ S. M. Griffin, M. Lilienblum, K. T. Delaney, Y. Kumagai, M. Fiebig, and N. A. Spaldin, *Phys. Rev. X* **2**, 041022 (2012).
- ³⁵ M. Mielenz, J. Brox, S. Kahra, G. Leschhorn, M. Albert, T. Schaetz, H. Landa, and B. Reznik, *Phys. Rev. Lett.* **110**, 133004 (2013).
- ³⁶ S. Ejtemaee and P. C. Haljan, *Phys. Rev. A* **87**, 051401 (2013).
- ³⁷ S. S. Ulm, J. Roßnagel, G. Jacob, C. Degünther, S. T. Dawkins, U. G. Poschinger, R. Nigmatullin, A. Retzker, M. B. Plenio, F. Schmidt-Kaler, and K. Singer, *Nat. Commun.* **4**, 2290 (2013).
- ³⁸ K. Pyka, J. Keller, H. L. Partner, R. Nigmatullin, T. Burgermeister, D. M. Meier, K. Kuhlmann, A. Retzker, M. B. Plenio, W. H. Zurek, A. del Campo, and T. E. Mehlstäubler, *Nat. Commun.* **4**, 2291 (2013).
- ³⁹ C. N. Weiler, T. W. Neely, D. R. Scherer, A. S. Bradley, M. J. Davis, and B. P. Anderson, *Nature* **455**, 948 (2008).
- ⁴⁰ G. Lamporesi, S. Donadello, S. Serafini, F. Dalfovo, and G. Ferrari, *Nat. Phys.* **9**, 656 (2013).
- ⁴¹ L. Corman, L. Chomaz, T. Bienaimé, R. Desbuquois, C. Weitenberg, S. Nascimbène, J. Dalibard, and J. Beugnon, *Phys. Rev. Lett.* **113**, 135302 (2014).
- ⁴² L. Chomaz, L. Corman, T. Bienaimé, R. Desbuquois, C. Weitenberg, S. Nascimbène, J. Beugnon, and J. Dalibard, *Nat. Commun.* **6**, 6162 (2015).
- ⁴³ S.-Z. Lin, X. Wang, Y. Kamiya, G.-W. Chern, F. Fan, D. Fan, B. Casas, Y. Liu, V. Koryukhin, W. H. Zurek, C. D. Batista, and S.-W. Cheong, *Nat. Phys.* **10**, 970 (2014).
- ⁴⁴ N. Navon, A. L. Gaunt, R. P. Smith, and Z. Hadzibabic, *Science* **347**, 167 (2015).
- ⁴⁵ J. Beugnon and N. Navon, “Exploring the Kibble-Zurek mechanism with homogeneous bose gases,” (2016), [arXiv:1611.01145](https://arxiv.org/abs/1611.01145).
- ⁴⁶ M. Anquez, B. A. Robbins, H. M. Bharath, M. Boguslawski, T. M. Hoang, and M. S. Chapman, *Phys. Rev. Lett.* **116**, 155301 (2016).
- ⁴⁷ L. W. Clark, L. Feng, and C. Chin, *Science* **354**, 606 (2016).
- ⁴⁸ A. del Campo, T. W. B. Kibble, and W. H. Zurek, *J. Phys. Condens. Matter* **25**, 404210 (2013).
- ⁴⁹ W. H. Zurek, *J. Phys. Condens. Matter* **25**, 404209 (2013).
- ⁵⁰ B. Damski, H. T. Quan, and W. H. Zurek, *Phys. Rev. A* **83**, 062104 (2011).
- ⁵¹ W. H. Zurek, *Phys. Rev. Lett.* **102**, 105702 (2009).
- ⁵² B. Damski and W. H. Zurek, *Phys. Rev. Lett.* **104**, 160404 (2010).
- ⁵³ L. Cincio, J. Dziarmaga, M. M. Rams, and W. H. Zurek, *Phys. Rev. A* **75**, 052321 (2007).
- ⁵⁴ T. W. B. Kibble, *Patterns of Symmetry breaking* (Kluwer Academic Publishers, London, 2003).
- ⁵⁵ T. W. B. Kibble, *Physics Today* **60**, 47 (2007).
- ⁵⁶ J. Dziarmaga, *Adv. Phys.* **59**, 1063 (2010).
- ⁵⁷ A. Polkovnikov, K. Sengupta, A. Silva, and M. Vengalattore, *Rev. Mod. Phys.* **83**, 863 (2011).
- ⁵⁸ A. del Campo and W. H. Zurek, *Int. J. Mod. Phys. A* **29**, 1430018 (2014).
- ⁵⁹ J. Dziarmaga, A. Smerzi, W. H. Zurek, and A. R. Bishop, *Phys. Rev. Lett.* **88**, 167001 (2002).
- ⁶⁰ B. Damski, *Phys. Rev. Lett.* **95**, 035701 (2005).
- ⁶¹ W. H. Zurek, U. Dorner, and P. Zoller, *Phys. Rev. Lett.* **95**, 105701 (2005).
- ⁶² J. Dziarmaga, *Phys. Rev. Lett.* **95**, 245701 (2005).
- ⁶³ A. Polkovnikov, *Phys. Rev. B* **72**, 161201 (2005).
- ⁶⁴ V. Mukherjee, U. Divakaran, A. Dutta, and D. Sen, *Phys. Rev. B* **76**, 174303 (2007).
- ⁶⁵ U. Divakaran, V. Mukherjee, A. Dutta, and D. Sen, *J. Stat. Mech. Theor. Exp.* **2009**, P02007 (2009).
- ⁶⁶ U. Divakaran, A. Dutta, and D. Sen, *Phys. Rev. B* **78**, 144301 (2008).
- ⁶⁷ D. Chowdhury, U. Divakaran, and A. Dutta, *Phys. Rev. E* **81**, 012101 (2010).
- ⁶⁸ K. Sengupta, D. Sen, and S. Mondal, *Phys. Rev. Lett.* **100**, 077204 (2008).
- ⁶⁹ S. Mondal, D. Sen, and K. Sengupta, *Phys. Rev. B* **78**, 045101 (2008).
- ⁷⁰ U. Divakaran and A. Dutta, *Phys. Rev. B* **79**, 224408 (2009).
- ⁷¹ V. Mukherjee, A. Dutta, and D. Sen, *Phys. Rev. B* **77**, 214427 (2008).
- ⁷² V. Mukherjee and A. Dutta, *J. Stat. Mech. Theor. Exp.* **2009**, P05005 (2009).
- ⁷³ U. Divakaran, A. Dutta, and D. Sen, *Phys. Rev. B* **81**, 054306 (2010).
- ⁷⁴ S. Suzuki, *J. Stat. Mech. Theor. Exp.* **2009**, P03032 (2009).
- ⁷⁵ A. Dutta, R. R. P. Singh, and U. Divakaran, *EPL* **89**, 67001 (2010).
- ⁷⁶ B. Dóra and R. Moessner, *Phys. Rev. B* **81**, 165431 (2010).
- ⁷⁷ K. Baumann, R. Mottl, F. Brennecke, and T. Esslinger, *Phys. Rev. Lett.* **107**, 140402 (2011).
- ⁷⁸ D. Chen, M. White, C. Borries, and B. DeMarco, *Phys. Rev. Lett.* **106**, 235304 (2011).
- ⁷⁹ S. Braun, M. Friesdorf, S. S. Hodgman, M. Schreiber, J. P. Ronzheimer, A. Riera, M. del Rey, I. Bloch, J. Eisert, and U. Schneider, *PNAS* **112**, 3641 (2015).
- ⁸⁰ C. Meldgin, U. Ray, P. Russ, D. Chen, D. M. Ceperley, and B. DeMarco, *Nat Phys* **12**, 646 (2016), letter.
- ⁸¹ J.-M. Cui, Y.-F. Huang, Z. Wang, D.-Y. Cao, J. Wang, W.-M. Lv, L. Luo, A. del Campo, Y.-J. Han, C.-F. Li, and G.-C. Guo, *Sci Rep.* **6**, 33381 (2016).
- ⁸² M. Kolodrubetz, B. K. Clark, and D. A. Huse, *Phys. Rev. Lett.* **109**, 015701 (2012).
- ⁸³ A. Chandran, A. Erez, S. S. Gubser, and S. L. Sondhi, *Phys. Rev. B* **86**, 064304 (2012).
- ⁸⁴ S. Deng, G. Ortiz, and L. Viola, *EPL* **84**, 67008 (2008).
- ⁸⁵ B. Damski and W. H. Zurek, *New J. Phys.* **11**, 063014 (2009).

- ⁸⁶ J. Dziarmaga and M. M. Rams, [New J. Phys. **12**, 055007 \(2010\)](#).
- ⁸⁷ J. Dziarmaga and W. H. Zurek, [Sci. Rep. **4**, 5950 \(2014\)](#).
- ⁸⁸ J. M. Kosterlitz and D. J. Thouless, [J. Phys. C **6**, 1181 \(1973\)](#).
- ⁸⁹ D. J. Resnick, J. C. Garland, J. T. Boyd, S. Shoemaker, and R. S. Newrock, [Phys. Rev. Lett. **47**, 1542 \(1981\)](#).
- ⁹⁰ H. Zoran, K. Peter, C. Marc, B. Baptiste, and J. Dalibard, [Nature **441**, 1118 \(2006\)](#).
- ⁹¹ S. M. Girvin, “The Kosterlitz-Thouless transition,” (unpublished lecture notes Boulder, 2000).
- ⁹² R. Corless, [On the Lambert W Function](#), Research report (University of Waterloo, Computer Science Department, 1993).
- ⁹³ A. Jelić and L. F. Cugliandolo, [J. Stat. Mech. Theor. Exp. **2011**, P02032 \(2011\)](#).
- ⁹⁴ S. Deuschl nder, P. Dillmann, G. Maret, and P. Keim, [PNAS **112**, 6925 \(2015\)](#).
- ⁹⁵ M. L. Wall and L. D. Carr, [New J. Phys. **14**, 125015 \(2012\)](#).
- ⁹⁶ F. M. Cucchietti, B. Damski, J. Dziarmaga, and W. H. Zurek, [Phys. Rev. A **75**, 023603 \(2007\)](#).
- ⁹⁷ N. Elstner and H. Monien, “A numerical exact solution of the Bose-Hubbard model,” (1999), [arXiv:cond-mat/9905367](#).
- ⁹⁸ V. Gritsev and A. Polkovnikov, “Universal dynamics near quantum critical points,” (2010), [arXiv:0910.3692v3](#).
- ⁹⁹ F. Binder, S. Vinjanampathy, K. Modi, and J. Goold, [Phys. Rev. E **91**, 032119 \(2015\)](#).
- ¹⁰⁰ B. Gardas and S. Deffner, [Phys. Rev. E **92**, 042126 \(2015\)](#).
- ¹⁰¹ J. Dziarmaga and M. Tylutki, [Phys. Rev. B **84**, 214522 \(2011\)](#).