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Spin texture induced by oxygen vacancies in strontium perovskite (001) surfaces: A theoretical comparison between SrTiO_{3} and SrHfO_{3} A. C. Garcia-Castro, M. G. Vergniory, E. Bousquet, and A. H. Romero Phys. Rev. B 93, 045405 — Published 7 January 2016 DOI: 10.1103/PhysRevB.93.045405

Spin-texture induced by oxygen vacancies in Strontium perovskites (001) surfaces: A theoretical comparison between SrTiO₃ and SrHfO₃

A. C. Garcia-Castro¹,^{*} M. G. Vergniory², E. Bousquet¹, and A. H. Romero^{3†}

¹Physique Théorique des Matériaux, Université de Liège, B-4000 Sart-Tilman, Belgium

²Donostia International Physics Center, 20018 Donostia/San Sebastian, Spain and

³Physics Department, West Virginia University, WV-26506-6315, Morgantown, USA

The electronic structure of $SrTiO_3$ and $SrHfO_3$ (001) surfaces with oxygen vacancies is studied by means of first-principles calculations. We reveal how oxygen vacancies within the first atomic layer of the $SrTiO_3$ surface (i) induce a large antiferrodistortive motion of the oxygen octahedra at the surface, (ii) drive localized magnetic moments on the Ti-3d orbitals close to the vacancies and (iii) form a two-dimensional electron gas localized within the first layers. The analysis of the spintexture of this system exhibits a splitting of the energy bands according to the Zeeman interaction, lowering of the $Ti-3d_{xy}$ level in comparison with d_{xz} and d_{yz} and also an in-plane precession of the spins. No Rashba-like splitting for the ground state neither for ab-initio molecular dynamics trajectory at 400K is recognized as suggested recently by A. F. Santander-Syro *et al.*¹. Instead, a sizeable Rashba-like splitting is observed when the Ti atom is replaced by a heavier Hf atom with a much larger spin-orbit interaction. However, we observe the disappearance of the magnetism and the surface two-dimensional electron gas when full structural optimization of the SrHfO₃ surface is performed. Our results uncover the sensitive interplay of spin-orbit coupling, atomic relaxations and magnetism when tuning these Sr-based perovskites.

PACS numbers: 71.15.Mb, 73.20.-r, 68.47.Gh, 71.70.Ej

I. INTRODUCTION

Oxide perovskites exhibit a very large variety of physical properties, which allow their use in a wide range of applications such as spintronics or thermoelectricity²⁻⁴. In particular, during the last decade it has been observed new and exotic functional properties at the surface of oxide crystals with no parallel to their bulk counterpart^{5,6}. In this field of multifunctional oxide materials, $SrTiO_3$ (STO) has attracted a lot of attention due to its wide range of physical and chemical properties that goes from superconductivity⁷ to engineered ferroelectricity⁸. In various studies a two-dimensional electron gas (2DEG) has been reported at the (001) surface of STO crystal, both theoretically⁹ and experimentally by using angle-resolved photoemission spectroscopy $(ARPES)^{10,11}$, which appeared to be induced by the presence of oxygen vacancies and explained by an electronic charge redistribution at the surface. The dynamics of the Ti-d electron quasiparticle on this surface have been experimentally characterized and the appearance of cooperative effects of the electrons has been reported, giving rise to an orbital ordering itself leading to a Rashba-like splitting at the d_{xy} - $d_{xz/yz}$ bands crossing close to the Fermi level¹¹. Recently, a mixed character of the confined electrons of the (001) surface of STO displaying a complex spin texture at the Fermi surface has been reported. This unexpected result was ascribed to a Zeeman interaction plus Rashba-like splitting of the lowest d_{xy} owed as well to oxygen vacancies¹.

Similar Rashba-effect¹² in oxides has been observed in the $LaAlO_3/SrTiO_3$ interface¹³⁻¹⁶ or in $SrTiO_3$ and $KTaO_3$ -based transistors^{17,18}. For these single perovskite-based devices, a k-cubic dependence of the splitting was found in contrast to the linear+cubic dependence of the Rashba splitting at the LaAlO₃/SrTiO₃ interface due to the $d_{xy} - d_{xz/yz}$ multiorbital nature of the lowest bands¹⁵. As in the case of Au(111) surface, where it has been shown that even when the surface states exhibit a mainly *p*-orbital character, the *d*-orbitals drives the Rashba splitting and dictates the direction of the spin by means of the orbital-angular momentum $(OAM)^{19,20}$. Therefore, the *d*-orbitals of the transition metal atom of perovskite oxides can be the key parameter for a large and controllable electron band splitting in future spintronic devices 19,21 . This later property gives an additional degree of freedom to the rich field of multifunctional perovskites. However, the recent findings of a giant spin-splitting at the STO surface, opens some questions about the nature of the spin-splitting in this type of materials, such as: What is the origin of this giant-splitting? Could it be manipulated or observed in similar compounds? What is the role of the spin-orbit coupling strength in the Rashba splitting of d-orbitals in oxide perovskites? Since magnetism at the surface was reported, how does it couple with the splitting?. From the theoretical point of view, it is well known that density-functional theory (DFT) formalism provides invaluable information related to the structural, magnetic and electronic structure of oxide interfaces, superlattices and $surfaces^{5,15,22,23}$. This has been particularly true for the $SrTiO_3/LaAlO_3$ interface^{23,24} or the STO surface²⁴ where DFT helped in understanding the origin of the 2DEG through the polar catastrophe scenario. In the

^{*} a.c.garcia.castro@gmail.com

[†] alromero@mail.wvu.edu

present paper we address the previous questions by performing such DFT calculations on the (001) STO surface with oxygen vacancies in order to (i) shed light onto this novel experimental finding and (ii) to scrutinize the microscopic origin of the giant Rashba spin-splitting and the Ti *d*-orbitals spin-texture.



FIG. 1. (Color online) Schematic view of the SrTiO₃: TiO_{2- δ} (001) surface used in the calculations with an oxygen vacancy (OV) at the yz- (left) and xz-plane (right) respectively. The OV is located at the $(\frac{1}{4}, \frac{1}{4}, z)$ position with respect to the computational x'- and, y'- coordinates). The amplification of the AFD distortions due to the oxygen vacancy can be clearly appreciated. Additionally, a surface induced off-centering of the Ti along the z direction is observed with respect to the oxygens (rumpling), inducing an electric dipole. In the right bottom part we present a top view of the unit cell in the xy-plane. This figure was elaborated with the support of the VESTA code²⁵.

II. COMPUTATIONAL DETAILS

The first-principles calculations have been performed within the DFT as implemented in the Vienna Ab-initio Simulation Package $(VASP)^{26,27}$. The interaction between ion cores and valence electrons was treated by the projector augmented wave method²⁸. The orbitals considered as valence electrons in the atomic pseudopotentials were: Sr $(4s^24p^65s^2)$, Ti $(3s^23p^64s^23d^2)$, Hf $(5s^25p^65d^26s^2)$, and O $(2s^22p^4)$. To describe the exchange-correlation energy, we used the General Gradient Approximation (GGA) with the Perdew-Burke-Ernkzerhof for Solids (PBEsol) parameterization²⁹. In order to account for the magnetic character of the system we performed spin polarized calculations. The Hamiltonian contained the scalar relativistic corrections and the spin-orbit coupling was taken into account by the second variation method³⁰.

A Monkhorst-Pack k-point grid of $(8 \times 8 \times 1)$ for reciprocal space integration and 600 eV energy cut-off of the plane wave expansion have been used to get a residual error on the forces below 1 meV/Å and a fully converged electronic structure including SOC. To correctly describe the (001) surface and avoid interactions between periodic slabs, we used a supercell with a vacuum space of 12 Å. Additionally, in order to take into account the octahedra rotations and tilting, we considered a $(\sqrt{2} \times \sqrt{2} \times 7)$ cell as shown in Fig. 1, enforcing the cubic symmetry at the cell center. Aiming to prevent a fake polarization along the z-axis, the central plane is used as m_z mirror symmetry such as the slab is repeated along the z direction.

III. RESULTS

In what follows, the electronic characterization of the Rashba-like splitting for $SrBO_3$ with B = Ti and Hf, including surface oxygen vacancies at the BO_2 termination according to experiments¹ is discussed. Great emphasis is done on the size of the splitting and the spin structure.

A. STO:TiO₂ terminated surface:

Fully relaxed STO surface reveals a change in the interatomic distances close to the surface due to the missing Ti–O surface bonds (rumpling). The Ti–Ti distance calculated between the Ti surface layer and the next just below is 3.821 Å, which is much smaller than the STO bulk distance of 3.905 Å. Additionally, the out-of-plane Ti-O bond length at the surface changes its distance from 1.948 to 1.950 Å, leading to a local polar like displacement. We also found a substantial lowering of the calculated electronic gap, which goes from 2.0 eV at the bulk level to 0.8 eV in the slab geometry. We computed the electronic band structure with and without SOC and we did not observe substantial differences, which implies that SOC has a small contribution to the changes in the electron band structure compared with the geometrical changes. Furthermore, no octahedral rotations or tilting were observed in the fully relaxed surface when the cubic symmetry is imposed at the central slab plane.

B. STO:TiO_{2- δ} surface with oxygen vacancies:

Experimentally, STO surface samples become conducting in the presence of oxygen vacancies^{32,33}. In order to analyze their effect, we performed full structural relaxation of the STO:TiO₂ surface by removing one oxygen in the last TiO₂ termination layer ($\delta = 0.5$), imposing a bulk cubic symmetry to allow us a direct comparison with the pristine case. As we illustrate it in Fig. 1, octahedral tilting and rotations are now observed in the



FIG. 2. (Color online) Band structures around Γ -point for SrTiO₃ and SrHfO₃ with an oxygen vacancy on the BO_2 terminated surface. Red and blue denote positive and negative projection of spin vector \vec{s} on Cartesian axes, units of the k-vector in units of π/a . The non spin-polarized bands are depicted in dark grey. The effect of the SOC and a Rashba-type splitting is observed in the d_{xy} and $d_{xz/yz}$ bands crossing (noted with the dashed circles) as well as the vertical energy shift $\Delta \equiv \Delta E_{Zeeman} + \Delta E_l$. For SHO a spin-splitting of the lower bands is observed for the x and y components around E = -0.1 eV. Figures were created by using the PyProcar script³¹.

presence of oxygen vacancies (the rotation angle is 10.9° with respect to the z-axis for the tetrahedral closest to the vacancy), in contrast to the surface without oxygen vacancies, where no rotation was observed.

In what follows we study the electronic and spin structure of the (001) STO:TiO_{2- δ} surface. In Fig. 2, we report the calculated electronic band structure around the Γ -point including non-collinear magnetism and SOC. We found a conducting behavior very similar to what has been observed at the LaAlO₃/SrTiO₃ interface²⁴, where the lowest energy bands are the Ti-3d orbitals. At the bulk level, the Ti–3d orbitals present in the TiO₆ octahedra are expected to split into $t_{2\mathbf{g}}$ (d_{xy}, d_{xz}, d_{yz}) and $e_{\mathbf{g}}(d_{z^2}, d_{x^2-y^2})$ due to the crystal-field splitting. Here, we remark that small distortions in the octahedral geometry, as the one observed at the surface, generate a small shifting in the energy-values of the electron levels (ΔE_l) and a possible mixing of the $t_{2\mathbf{g}}$ and $e_{\mathbf{g}}$ states, mainly due to the symmetry breaking and the uncompensated TiO_6 octahedra at the surface³⁴. In the oxygen vacancy STO surface we found that the lower d-levels bands corresponds to the d_{xy} orbitals in the paths that runs from $-0.5k_x - \Gamma - +0.5k_x$ and $-0.5k_y - \Gamma - +0.5k_y$. In addition, we also found a magnetic moment on the Ti atoms close to the surface. The amplitude of these magnetic moments is around 0.212 μ_B /Ti-atom in the

surface and 0.026 $\mu_B/\text{Ti-atom}$ in the sub-surface layer. These results are in agreement with the experimental observation of magnetism in STO with oxygen vacancies³⁵ and other theoretical works of the same system^{36–38}. The spin polarization is analyzed by plotting the s_x , s_y and s_z components of the magnetization for each path, where up and down spin orientations are depicted by red and blue colors respectively. We can appreciate that lower light bands close to the Fermi level are separated vertically due to an induced Zeeman splitting (ΔE_{Zeeman}) as expected from the magnetism found at the surface and the energy splitting ΔE_l . This magnetic moment is induced by the effect of the electron surface density rearrangement as a response of the uncompensated orbitals, thus oxygen vacancy acts as a magnetic impurity^{36,38}.

From Fig. 2 we can also see the parabolic-like shape of the *d*-bands, which is a characteristic of a 2DEG, as reported in previous studies^{10,33,39}. Here, it is important to note that this 2DEG is induced by the electronic reconstruction related to the confinement of the electrons at the surface^{32,38}, rather than a polar catastrophe origin as mentioned for the SrTiO₃/LaAlO₃ interface²³. Thus, we can expect that the behavior of the conducting electrons in STO differs from the one found in the oxide/oxide interfaces.

The effect of the SOC can be clearly appreciated in the

Rashba-like spin-splitting at the d_{xy} and $d_{xz/yz}$ crossing (pointed out by the dashed circles in Fig. 2), similarly to what has been previously reported in this surface¹¹. For the d_{xy} state, we obtain an effective mass of 1.6 m_e for the orbital located at the second layer of TiO_2 . This calculated effective mass is two times larger than the one reported in Ref.11 (0.65 m_e), which is too small for d like states⁴⁰. However, a spin-splitting such as in a classical Rashba picture is not observed (*i.e.* crossing of the spin up and down bands) between the light branches or a spin inversion within the d_{xy} bands for any of the spinorientations along the k_x and k_y paths. This result is in contrast with the interpretation of the experimental measurements of Santander-Syro $et \ al.^1$ who reported a giant spin-splitting in the lowest d-bands of the (001) STO surface. One possible explanation of this splitting is that it can be a consequence of the surface ionic vibrations. In that respect, we have performed ab-initio molecular dynamics calculations at constant temperature of 400 K for the slab geometry. Selected uncorrelated frames were taken from the simulation and electronic band structures were calculated for each case. As for the 0 K ground state, we did not observe any Rashba splitting like the one reported in Ref 1, which indicates that at least for low temperatures there is no coupling between the ionic vibrations and the electronic response.

Another possibility could be that it comes from anharmonic vibrations at much higher temperatures, a possibility we do not explore in this work. The absence of a large spin-splitting can be understood from the small SOC present in the 3d-Ti orbitals. To test this hypothesis, we performed a computer experiment by replacing all the Ti atoms by a heavier cation with much larger SOC such as Hafnium. Taking the advantage that the Hf atom has the same electronic configuration as the Ti but with 5d orbitals⁴¹. In these calculations cell lattice parameters and internal coordinates of the optimized geometry of the STO surface were fixed in order to keep the same structure. The 2DEG generated at the HfO_2 surface with oxygen vacancies presents the same main features as in the STO case. However, a large change in the spin-splitting state is observed in the electron dispersion as shown in Fig. 2, and somehow closer to the experimental suggestion of A. F. Santander-Syro *et al.*¹, who claimed for a spin-inversion direction within the lowest *d*-band. In both, k_x and k_y paths in Fig. 2 and at the Fermi level the bands are mainly composed by s_z spin direction. The spin switching is observed for the light d_{xy} Hf band at E = -0.1 eV and reciprocal points close to -0.1 and 0.1 for s_x and s_y spin components respectively. This suggests that a strong surface polarization is not enough to induce a Rashba-like spin-splitting and the presence of high-Z cations in the structure is also necessary to achieve this spin state as previously reported for Ag, Cu, and Au surfaces⁴². The Rashba-like splitting in the SrHfO₃ (SHO) is only observed for the s_x and s_y spin components of the xy-plane perpendicular to z-axis as expected from the Rashba Hamiltonian $\mathcal{H}_{\mathcal{R}}$ =

 $\alpha_R(\vec{e}_z \times \vec{k}) \cdot \vec{s}^{12}.$

In addition to the band analysis, we have also extracted the in-plane spin-texture as shown in the Fig. 3. In Fig. 3 we can see two concentric rings for up and down spin polarization respectively for the STO (at E = 0 eV) and one single ring for SHO (E = -0.1 eV) surface. The rings in the STO appears as a result of the vertical splitting due to $\Delta E_{Zeeman} + \Delta E_l$, where ΔE_l is induced by the lowering of the d_{xy} energy bands in comparison to d_{yz} and d_{zx} orbitals as discussed before. In the case of SHO (Fig. 3), spin precession in the lower ring is observed in addition to a spin switching of the d_{xy} band in agreement with the bands in Fig. 2) as expected for the Rashba-type splitting.



FIG. 3. (Color online) Band structure at the k_x - k_y plane (in π/a) where the spin-texture is calculated for each spin component along the cartesian axis (*i.e.* s_x , s_y and s_z). For SrTiO₃ compound (above, E = 0 eV) simple concentric rings for up (red) and down (blue) spin polarizations are observed. In contrast a Rashba-like splitting is observed by replacing the Ti cation by the heavier Hf in the SrHfO₃ (bottom, E = -0.1 eV). Figures were created by using the script PyProcar³¹.

The absence of the spin-splitting in the STO surface could be also linked to magnetism, as observed at the the Gd(0001) surface⁴³ where the presence of magnetism and a Rashba-type splitting has been reported. For this system the splitting is lower than 0.1 \AA^{-1} even with the large SOC present in Gd (around 9 times larger than in Ti)⁴³. Moreover, due to the presence of ΔE_{Zeeman} , there is no crossing of the bands and the direction of the magnetization vector (\vec{M}) determine the presence of the spinsplitting. This is due to the fact that the Rashba-splitting vanishes for $\vec{k} \parallel \vec{M}$ and only ΔE_{Zeeman} remains⁴³. The later statement needs to be carefully taken into account in the STO case for which we find a magnetic moment components at the surface $m_x = 0.213$, $m_y = 0.212$ and $m_z = 0.212 \ \mu_B / \text{Ti-atom}$ respectively and thus after considering several possible magnetic orientations. Different starting magnetic configurations, along x- y- and z-axis respectively were taken into account and full structural and energy minimizations were applied. Nonetheless, the system always relaxed toward the same magnetic ground

state in which a diagonal direction of the magnetic moment is observed with respect to the surface. No traces of magnetization were observed at the oxygen sites confirming the Ti–localized magnetic moments. Thus, this rather large in- and out-of-plane magnetic moments have the tendency to prevent the possibility to have a giant Rashba splitting, in addition to the small SOC discussed previously.

In the case of SHO, we obtain that the Hf magnetic moments are strongly reduced to $m_x = 0.034$, $m_y = 0.034$ and $m_z = 0.032 \,\mu_B/\text{Hf}$ -atom respectively close to the surface, which strongly reduced the ΔE_{Zeeman} . Fig. 3 shows that the lowest *d*-band in SHO only exhibit spin components along the *xy*-plane. Nonetheless for higher energy values, the bands are mainly spin-polarized along the *z*axis, demonstrating a very complex spin structure with a unique spin-texture. Therefore, only a deep understanding of this subtle competition between the polar distortions, the symmetry conditions, SOC, and magnetism would allow us to tune the Rashba-splitting at the (001) BO_2 surface.

In order to test the validity of our results against the exchange-correlation functional, we also carried out the same calculations with both the Local Density Approximation⁴⁴ (LDA) and the GGA PBEsol+U functionals. Using the LDA approximation, we did not find substantial differences with the GGA PBEsol calculations. Due to the strong correlation of the *d*-orbitals that drives the magnetism, we might expect the DFT+Uformalism⁴⁵ to enhance the electron correlation and change the physical picture discussed previously. Indeed, above U = 2 eV, we observe an insulating state of the STO surface with the oxygen vacancy, which do not agree with the experimental measurements 1,10,11 . Then, for whatever U value below 2 eV, we do not observe a large Rashba+Zeeman (RZ) type spin-splitting. In order to test the oxygen vacancy position effects, we performed a second set of calculations considering the oxygen vacancy in the second inner layer from the surface instead of the first layer and, as before, we did not observe a Rashba spin-splitting in the STO surface within this configuration.

C. *Pbnm* SrHfO₃:HfO₂ surface with oxygen vacancies:

We completed the analysis of the Rashba splitting mentioned above by analysing the spin textures in the SrHfO₃ ground state, *i.e.* in the *Pbnm* phase with a =5.7516 Å, b = 5.7646 Å, and c = 8.1344 Å⁴⁶. To analyze the new SHO surface in its *Pnma* phase, we fully relaxed the surface including an oxygen vacancy in the HfO_{2- δ} termination slab. Our results reveal an increase of octahedral tilting and rotations as a result of the free surface plus the oxygen vacancy, which induce a structural reconstruction with respect to the bulk (see Fig 4).

From the electronic structure analysis we observe a



FIG. 4. (Color online) Band structures around the Γ-point along k_x and k_y for SrHfO₃ surface with an oxygen vacancy on the top of the HfO₂ layer for each direction of the spin. Red and blue colors represent the up and down spin directions respectively. A larger Rashba-type splitting of the bands is observed for the z component as appreciated for the $k_y - s_z$ extracted image. The SHO slab structure is also shown in which the large octahedral tilting can be observed. Figures were created by using the script PyProcar³¹.

strong reduction of the band gap, from 3.22 eV in the bulk to 0.06 eV in the surface with oxygen vacancy and thus an absence of the 2DEG. Additionally, from the band structure plotted for each spin orientation (see Fig. 4, bands around the Γ -point along k_x and k_y paths) we can appreciate that the lowest Hf bands corresponds to 5dlevels, and more precisely to the $5d_{xy}$ states. According to Fig. 4, we note that the major spin components for the 5d bands are along the z-axis and a Rashba-type splitting is also observed for this orientation in most of the branches, which happen to be out-of-plane and perpendicular to the surface according to our results. This behavior can be explained in terms of the polarization close to the surface and the Rashba Hamiltonian expressed before. Here, the bands present a spin-splitting for a dispersion along a \vec{k} -vector perpendicular to the electric field \vec{e}_{xy} . To confirm this hypothesis, we computed the spontaneous polarization by means of the Berry phase approach⁴⁷. We obtained that the P_z component is zero by construction and the in-plane components are equal to $0.44 \ \mu \text{C} \cdot \text{cm}^{-2}$ and $1.15 \ \mu \text{C} \cdot \text{cm}^{-2}$ found along the x-axis (P_x) and y-axis (P_y) respectively. This in-plane polarization in orthorhombic perovskites is quite well known and is related to the presence of anti-polar motion of the A-site in the xy-plane, which, in the case of the surface symmetry breaking, do not cancel out from layer to

layer and thus resulting in an effective spontaneous polarization along the xy-plane^{48,49}. Interestingly, and in contrast to the STO case, we do not find any presence of magnetism in the *Pbnm* SHO surface with oxygen vacancies and at the same time the Rashba-splitting is considerably enhanced, confining the rule discussed above. Interestingly, when replacing the Hf by Ti in the relaxed *Pbnm* SrHfO₃ surface, the magnetism appears again and the Rashba splitting dissapear. All of that confirms the influence of the magnetic ordering and the SOC strength of the *B*-site at the spin-texture of SrBO₃ surfaces with oxygen vacancies.

IV. SUMMARY AND DISCUSSIONS

We have analyzed from first principles calculations the (001) $SrBO_3$, (B = Ti and Hf) surface with oxygen vacancies. We found an insulator to metal transition behavior when an oxygen vacancy is included in STO at the TiO_2 surface, which produces an electronic charge rearrangement that results into a 2DEG located at the surface and magnetization of the Ti-d orbitals close to the vacancy. In the limits of our DFT calculations, it is unlikely that a large Rashba-type splitting in the light lowest d_{xy} bands exists in the presence of an oxygen vacancy at the $STO:TiO_2$ surfaces. This is in contrast with the giant spin-splitting of about 0.1 Å⁻¹ reported by Santander-Syro $et \ al.^1$ in their experiments. This discrepancy can come from different origins. On the experimental side, due to the methodology of the spin-ARPES technique used by Santander-Syro et al., excited states could have been induced at the surface, with a different spin-polarization of the photoelectrons. Then, these states can become accessible and be measured by this technique as already reported before 50-52. There is also a possibility of high temperature gradients at the surface due to the high energy from the laser field. This could create large ionic vibrations, which can create large ionic vibrations, even if the adiabatic approximation still holds 53 . On the calculation side, it is possible that the configuration chosen for the oxygen vacancy does not correspond to what is present experimentally. In that respect imaging techniques such as STM would be of use in order to check the surface structure.

It has been shown from photoemission spectroscopy

that the Rashba spin-splitting in these 2DEG systems is extremely sensitive to the charge and orbital ordering, the SOC or the lattice degrees of freedom¹¹. We observed such a sensitivity when replacing Ti by Hf atom with a much larger SOC, which induces a larger change in the spin-splitting landscape at fixed lattice degrees of freedom. Indeed, when relaxing toward the SHO ground state, we observe a totally different magnetic and Rashba spin-splitting states, all of that confirming the strong spin structure sensitivity of these surfaces. However, in all the configurations we considered, we never obtained a Rashba spin-splitting as large as the one reported by Santander-Syro et al. and the fact that it is much larger than the one observed in the $LaAlO_3/SrTiO_3$ interfaces (< 0.01 Å^{-113,16} and supported by theoretical analysis^{14,15}) call for additional studies (theoretical and experimental) in order to understand this unexpected large splitting. Going beyond this discrepancy, we proved that sizeable Rashba-type splitting can be achieved in multifunctional perovskite oxides with large SOC, such as SrHfO₃ not observed before. These systems thus opens the possibility to tune exotic properties for new technological applications such as spin-orbitronics 54,55 .

ACKNOWLEDGMENTS

We acknowledge the support from the Extreme Science and Engineering Discovery Environment (XSEDE), which is supported by National Science Foundation Grant No. ACI-1053575. AHR acknowledge the support of DMREF-NSF 1434897 and the Donors of the American Chemical Society Petroleum Research Fund for partial support of this research under contract 54075-ND10. EB thanks the FRS-FNRS and the Consortium des Equipements de Calcul Intensif (CECI) funded by the FRS-FNRS (grant 2.5020.11) and the PRACE project TheDeNoMo. MGV acknowledge partial support from the Basque Country Government, Departamento de Educación, Universidades e Investigación (Grant No. IT-756-13), the Spanish Ministerio de Economía e Innovación (Grant No. FIS2010-19609-C02-01 and FIS2013-48286-C2-1-P) and the FEDER funding. Additionally, the authors gratefully acknowledge the referees comments and suggestions that lead to the substantial improvement of the present paper.

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