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Shubnikov-de Haas oscillations in 2D electron gas under sub-terahertz radiation

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We report on magnetotransport measurements in a 2D electron gas subject to sub-terahertz radiation in the regime where Shubnikov-de Haas oscillations (SdHO) and microwave-induced resistance oscillations (MIRO) coexist over a wide magnetic field range, spanning several harmonics of the cyclotron resonance. Surprisingly, we find that the SdHO amplitude is modified by the radiation in a non-trivial way owing to the oscillatory correction which has the same period and phase as MIRO. This remarkable finding challenges our current understanding of microwave photoresistance in 2D electron gas, calling for future investigations.

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When a 2D electron gas (2DEG) is subject to a perpendicular magnetic field B and low temperature T , the longitudinal resistivity ρ exhibits Shubnikov-de Haas oscillations (SdHO), owing to a quantum correction

$$\delta\rho_{\text{SdH}} = -S \cos \pi\nu, \quad S = 4\rho_0\lambda\mathcal{D}_T. \quad (1)$$

Here, ν is the filling factor, ρ_0 is the resistivity at $B = 0$, $\lambda = \exp(-\pi/\omega_c\tau_q)$ is the Dingle factor, τ_q is the quantum lifetime, $\mathcal{D}_T = \mathcal{X}_T/\sinh \mathcal{X}_T$, $\mathcal{X}_T = 2\pi^2k_B T/\hbar\omega_c$, $\omega_c = eB/m^*$ is the cyclotron frequency, and m^* is the effective mass. When a 2DEG is subject to radiation of frequency $\omega = 2\pi f$, ρ also reveals microwave-induced resistance oscillations (MIRO) [1–11] which, according to Refs. [12, 13], are given by

$$\delta\rho_{\text{MIRO}} = -2\pi\epsilon\rho_0\mathcal{P}\eta\lambda^2 \sin 2\pi\epsilon, \quad (2)$$

where $\epsilon = \omega/\omega_c$, \mathcal{P} is the dimensionless radiation intensity [13–17], $\eta = \tau/2\tau_* + 2\tau_{\text{in}}/\tau$, τ is the transport lifetime, τ_{in} is the inelastic lifetime, and $\tau_*^{-1} = 3\tau_0^{-1} - 4\tau_1^{-1} + \tau_2^{-1}$ [18]. When the photoresistance $\delta\rho_{\text{MIRO}}$ approaches the dark resistivity ρ by absolute value, the MIRO minima evolve into zero-resistance states [19–27], which are understood in terms of current domains [28–31].

The majority of MIRO studies have been performed at relatively high T and low f , at which SdHO are strongly suppressed. Extending experiments to higher f [23, 32–39] and lower T yields a regime where SdHO and MIRO coexist, allowing to explore possible mixing between these two types of quantum oscillations and to investigate the effect of radiation on SdHO in general.

It has been known for some time that microwaves suppress SdHO in the vicinity of the cyclotron resonance,

$\epsilon \approx 1$ [32, 33, 35]. As SdHO are sensitive to the thermal smearing of the Fermi surface, this suppression can be directly linked to absorption, which is indeed the strongest near the cyclotron resonance [40–42]. Away from the cyclotron resonance, our understanding of how microwaves affect SdHO is definitely lacking. Some experiments have shown that the effect of microwaves on the SdHO is the weakest near half-integer ϵ , which was attributed to the suppression of both inter- and intra-Landau level absorption [34, 35]. Another experiment [37] found that as the MIRO minima approach zero, the SdHO amplitude vanishes in proportion with the background resistance. Reference 37 then argued that in an irradiated 2DEG, ρ_0 in Eq. (1) should be replaced by $\rho_{\text{MIRO}} \approx \rho_0 + \delta\rho_{\text{MIRO}}$.

There exist several mechanisms that could lead to modification of the SdHO by radiation. First, the absorption coefficient \mathcal{A} is expected to acquire an oscillatory quantum correction [6, 40, 43–46] which, according to Ref. 40 and 46, is given by

$$\delta\mathcal{A}_q \simeq 2\mathcal{A}_D\lambda^2 \cos 2\pi\epsilon, \quad (3)$$

where \mathcal{A}_D is a classical absorption described by a Drude formula [16, 46, 47]. Since oscillations in \mathcal{A} translate to oscillations in T [41, 42], Eq. (3) suggests that the microwave-induced suppression of SdHO is maximized near the cyclotron resonance and, to a much lesser extent, near its harmonics,

$$\epsilon = n = 1, 2, 3, \dots \quad (4)$$

In addition, theory also predicts a radiation-induced oscillatory correction, of the order $\mathcal{O}(\lambda)$ [48], to the dc resistivity. While the inelastic mechanism produces no

such contribution [49], the displacement mechanism dictates that S in Eq. (1) acquires an oscillatory correction and should be replaced by [49, 50]

$$S_\omega = S \left[1 - \mathcal{P} \frac{\tau}{\tau_*} \sin^2(\pi\epsilon) \right], \quad (5)$$

suggesting that the SdHO amplitude is minimized at

$$\epsilon = n + 1/2 = 3/2, 5/2, 7/2, \dots, \quad (6)$$

a condition orthogonal to Eq. (4). Finally, the same condition, Eq. (6), can be expected from classical oscillations in magnetoabsorption [51], $\delta\mathcal{A}_c/\mathcal{A}_D \sim -\cos 2\pi\epsilon$, which can be stronger than quantum oscillations, given by Eq. (3), in a typical 2DEG.

In this Rapid Communication we experimentally investigate the photoresistance in high-quality 2DEGs. Using high f , low \mathcal{P} , and low T allows us to overlap MIRO and SdHO over a wide range of ϵ and to investigate the SdHO waveform near multiple harmonics of the cyclotron resonance. Our data reveal pronounced modulation of the SdHO amplitude which persists to $\epsilon \approx 6$. Surprisingly, even though the modulation is periodic in ϵ , it cannot be described by either Eqs. (3),(4) or Eqs. (5),(6). Instead, the radiation-modified SdHO amplitude closely replicates the MIRO waveform, see Eq. (2), suggesting a non-trivial mixing of MIRO and SdHO. While it is well established that quantum oscillations of the order $\mathcal{O}(\lambda^2)$ interfere with each other [14, 15, 52–58], the observed correlation between $\text{MIRO} \sim \mathcal{O}(\lambda^2)$ and $\text{SdHO} \sim \mathcal{O}(\lambda)$ is totally unexpected [59].

While we have obtained similar findings from a variety of samples grown at Princeton and Purdue, in what follows we present the results from two Purdue-grown Hall bars, I and II, of width $w_A = 300 \mu\text{m}$ and $w_B = 200 \mu\text{m}$, respectively. Sample I is fabricated from a 30 nm-wide $\text{Al}_{0.0015}\text{Ga}_{0.9985}\text{As}/\text{Al}_{0.24}\text{Ga}_{0.76}\text{As}$ quantum well, with density $n_e \approx 3.1 \times 10^{11} \text{ cm}^{-2}$ and mobility $\mu \approx 3.6 \times 10^6 \text{ cm}^2/\text{Vs}$. Sample II contains a 30 nm-wide $\text{GaAs}/\text{Al}_{0.24}\text{Ga}_{0.76}\text{As}$ quantum well, with $n_e \approx 2.6 \times 10^{11} \text{ cm}^{-2}$ and $\mu \approx 2.1 \times 10^7 \text{ cm}^2/\text{Vs}$. The resistivity was measured using a standard low-frequency lock-in technique, in sweeping B , with f from 0.2 to 0.4 THz, generated by backward wave oscillators.

To facilitate the discussion of our results, we first define the relevant quantities and introduce convenient notations. In the absence of microwave radiation, the resistivity can be represented as

$$\rho = \rho_{\text{sm}} + \delta\rho_{\text{SdH}}, \quad (7)$$

where ρ_{sm} is the smooth part of the resistivity [60] and $\delta\rho_{\text{SdH}}$ is given by Eq. (1). When the radiation is present, we write the resistivity as

$$\rho_\omega = \rho_{\text{MIRO}} + \delta\rho_{\text{SdH}\omega}, \quad (8)$$

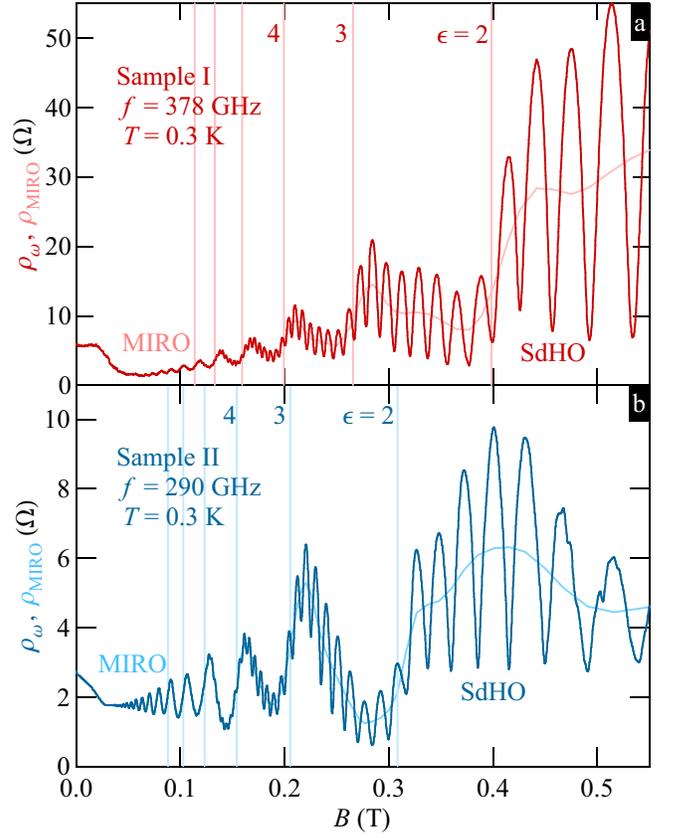


FIG. 1. (Color online) (a) [(b)] Magnetoresistivity $\rho_\omega(B)$ (dark curves) measured in sample I [B] irradiated by microwaves of $f = 378$ GHz [$f = 290$ GHz] at $T = 0.3$ K. Vertical lines are drawn at the cyclotron resonance harmonics, $\epsilon = \omega/\omega_c = 2, 3, 4, \dots$. Both panels also show $\rho_{\text{MIRO}}(B)$ (light curves) obtained by averaging out SdHO, see Eq. (8).

where we have introduced $\rho_{\text{MIRO}} = \rho_{\text{sm}\omega} + \delta\rho_{\text{MIRO}}$ containing slowly varying background $\rho_{\text{sm}\omega}$ [61] and MIRO photoresistance, see Eq. (2). The main goal of our study is to examine if and how $\delta\rho_{\text{SdH}\omega}$ is different from $\delta\rho_{\text{SdH}}$.

In Fig. 1(a) [(b)] we present the magnetoresistivity $\rho_\omega(B)$ measured in sample I [B] irradiated by microwaves of $f = 378$ [290] GHz at $T = 0.3$ K. The vertical lines are drawn at the cyclotron resonance harmonics, $\epsilon = \omega/\omega_c = 2, 3, 4, \dots$. In both samples, the data reveal pronounced MIRO, persisting down to $B \approx 0.05$ T, and SdHO, terminating around $B \approx 0.15$ T. Owing to high f and low T , there exists a wide range of B where SdHO and MIRO coexist. Most importantly, this range extends over several harmonics of the cyclotron resonance, spanning up to $\epsilon \approx 6$ and $\epsilon \approx 5$ for sample I and B, respectively. We immediately notice that, under irradiation, the SdHO amplitude S_ω is *not* a monotonic function of B , in contrast to the “dark” amplitude S described by Eq. (1). We thus conclude that the effect of radiation on SdHO is not limited to non-resonant heating.

To get more insight into the radiation-induced changes of SdHO, one needs to separate ρ_{MIRO} and $\delta\rho_{\text{SdH}\omega}$, en-

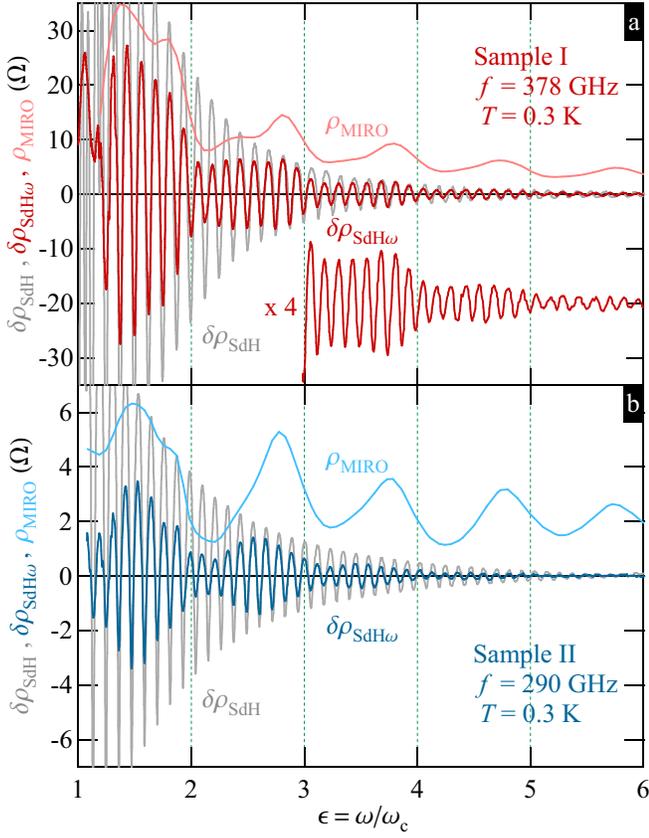


FIG. 2. (Color online) (a) [(b)] $\delta\rho_{\text{SdH}}$, $\delta\rho_{\text{SdH}\omega}$, and ρ_{MIRO} , obtained from the data in Fig. 1(a) [Fig. 1(b)], versus ϵ . Panel (a) also shows $\delta\rho_{\text{SdH}\omega}$, multiplied by 4 and offset by 20 Ω .

tering Eq.(8). Since ρ_{MIRO} oscillate much slower than SdHO [cf. Eq. (7)], it can be easily obtained by averaging out faster SdHO. Obtained in this way, $\rho_{\text{MIRO}}(B)$ is shown in both panels of Fig.1 by light curves running midway between the SdHO maxima and minima.

Having found ρ_{MIRO} , we now use Eq. (8) to obtain $\delta\rho_{\text{SdH}\omega}$ by subtracting ρ_{MIRO} from ρ_{ω} , both shown in Fig. 1. The results for sample I and II are presented as a function of ϵ in Fig. 2(a) and (b), respectively. For comparison, we also include ρ_{MIRO} and $\delta\rho_{\text{SdH}}$, as marked. The latter was found using Eq. (7) by subtracting the smooth part of the resistivity ρ_{sm} from $\rho(B)$ measured without irradiation. Direct examination of the SdHO reveals that $S_{\omega} \leq S$ in the entire range of ϵ , which is expected because the radiation elevates the temperature. In addition, one can now clearly see that, in contrast to $\delta\rho_{\text{SdH}}$, which monotonically decays with ϵ in accordance with Eq. (1), $\delta\rho_{\text{SdH}\omega}$ exhibits clear signs of modulation with the period close to unity. Furthermore, a comparison of $\delta\rho_{\text{SdH}\omega}$ and ρ_{MIRO} hints on strong correlation between the two quantities.

We next extract the amplitude of $\delta\rho_{\text{SdH}\omega}$, shown in Fig. 2, and examine it in more detail. In Fig. 3 we present the extracted amplitude S_{ω} (open circles) and ρ_{MIRO} (solid circles) as a function of ϵ on a log-linear scale.

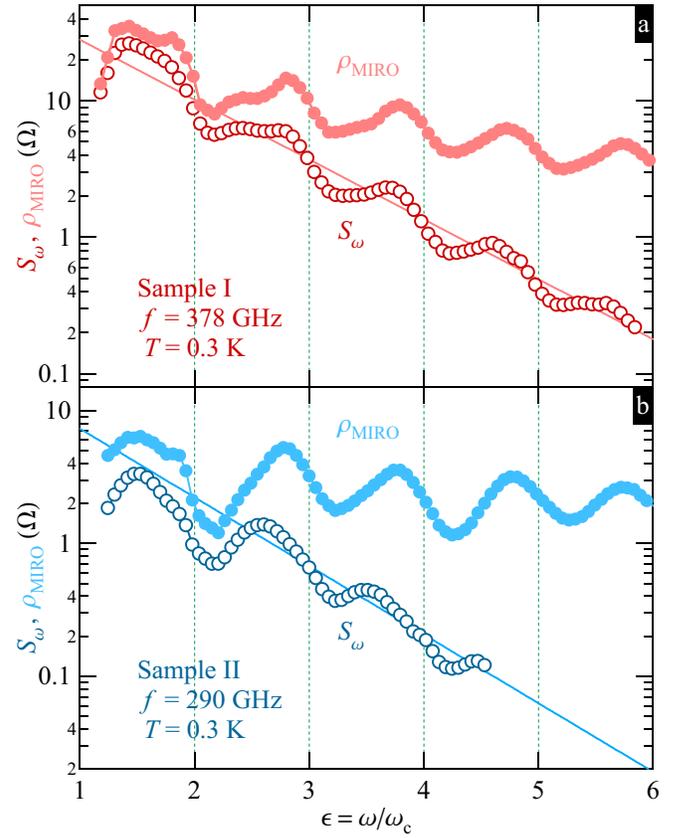


FIG. 3. (Color online) (a) [(b)] ρ_{MIRO} (solid circles) and S_{ω} (open circles), extracted from the data in Fig. 1(a) [Fig. 1(b)], as a function of ϵ .

Once plotted together, the correlation between S_{ω} and ρ_{MIRO} becomes very clear – both quantities oscillate in phase with each other. In other words, radiation induces minima in SdHO amplitude at

$$\epsilon \approx n + 1/4 = 5/4, 9/4, 13/4, \dots, \quad (9)$$

in contrast to both the scenario considering oscillations in magnetoabsorption [Eq. (3)] and the one predicting direct modification of the SdHO [Eq. (5)].

In the remaining part of this Rapid Communication we search for an empirical relation describing the SdHO amplitude in the presence of radiation. To this end, we extract and compare the oscillatory parts in S_{ω} and in ρ_{MIRO} . More specifically, we introduce the dimensionless quantity $\delta S_{\omega}/S = S_{\omega}/S - 1$, where S is the smooth, non-oscillating part of the SdHO amplitude shown in Fig. 3 by straight lines, and present the results in Fig. 4. We note that S is somewhat lower than S_0 as a result of a somewhat higher electron temperature under irradiation. For comparison, we also plot the oscillatory part of MIRO, $\delta\rho_{\text{MIRO}}/(2\pi\epsilon\lambda^2\rho_0)$.

We immediately see that both quantities oscillate around zero without noticeable decay, confirming that exponential factors have been properly eliminated. As already anticipated, a very good agreement in both the

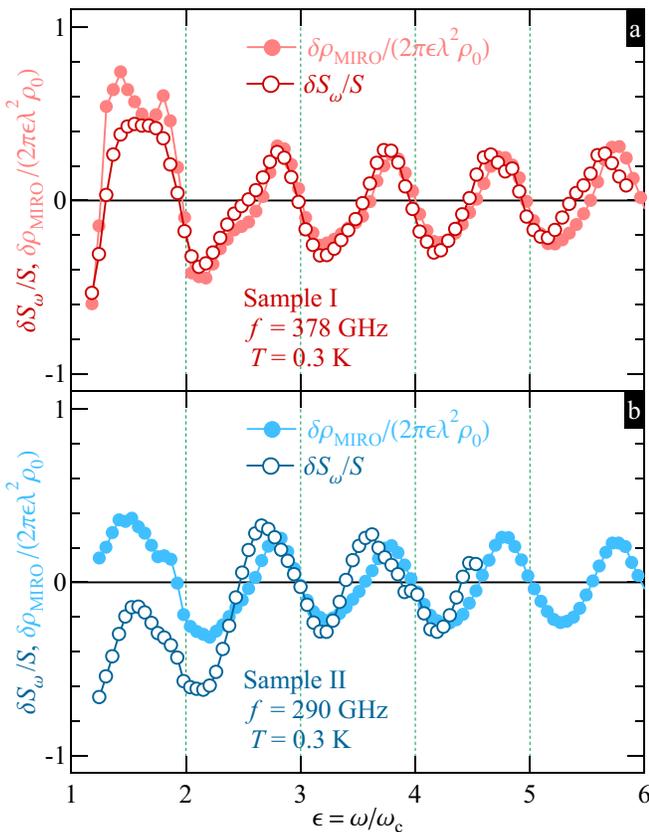


FIG. 4. (Color online) (a) [(b)] $\delta S_\omega/S$ (open circles) and $\delta\rho_{\text{MIRO}}/(2\pi\epsilon\lambda^2\rho_0)$ (solid circles), extracted from the data in Fig. 1(a) [Fig. 1(b)], as a function of ϵ .

period and the phase is found for the whole range of ϵ studied. This finding indicates that observed oscillations in SdHO amplitude are closely related to MIRO and thus are unlikely to originate from resonant heating caused by oscillations in magnetoabsorption [Eq. (3)]. We thus conclude that under presence of radiation the SdHO amplitude is given by

$$S_\omega \approx S(1 - \alpha \sin 2\pi\epsilon), \quad (10)$$

where α is a dimensionless ϵ -independent constant, showing that δS_ω is a correction of order $\mathcal{O}(\lambda)$, just like SdHO themselves.

The most remarkable finding, however, is an excellent *quantitative* correlation between δS_ω and $\delta\rho_{\text{MIRO}}$, namely

$$\frac{\delta S_\omega}{S} \approx \frac{\delta\rho_{\text{MIRO}}}{2\pi\epsilon\lambda^2\rho_0} \approx -\alpha \sin 2\pi\epsilon, \quad (11)$$

where all parameters have been obtained experimentally. The observed correlation is completely unexpected and is found almost everywhere, except at $\epsilon \lesssim 2$ in sample II. The latter can be linked to increased absorption close to the cyclotron resonance, where SdHO are suppressed due to resonant heating [32, 33, 35, 62]. The absence of

such deviation in sample I can be attributed to considerably higher f which reduces the influence of the cyclotron absorption peak. Interestingly, combining Eq. (11) with Eq. (2) one finds that $\alpha \approx \mathcal{P}\eta$.

Finally, we examine the proposal of Ref. 37 that the SdHO under irradiation can be described by Eq. (1) with ρ_0 replaced by $\rho_0 + \delta\rho_{\text{MIRO}}$ [63]. Taking this approach, one obtains $\delta S_\omega/S = \delta\rho_{\text{MIRO}}/\rho_0$, a result similar to Eq. (11), but with an extra factor $2\pi\epsilon\lambda^2$, which has significant dependence on ϵ . Indeed, as ϵ increases from 2 to 6, $2\pi\epsilon\lambda^2$ decreases by nearly a factor of 5 for sample I. In contrast, our data shown in Fig. 4(a) show virtually no decay at $\epsilon \gtrsim 2$. In addition, if this factor were actually present, the correction to SdHO would have been up to > 3 (> 5) times larger than observed in sample I (sample II). We thus conclude that the proposal of Ref. 37 is irrelevant to our findings.

In summary, we have studied the photoresistance in high-quality 2DEG subject to low temperatures and high microwave frequencies, which allowed us to overlap MIRO and SdHO over multiple harmonics of the cyclotron resonance. Our data revealed pronounced modulation of the SdHO which is periodic in ϵ , with the period equal to unity, and the phase matching that of MIRO. This result does not fit existing theories considering either magnetoabsorption or photoresistance. Most remarkably, we have found that once the exponential factors are eliminated, the oscillatory part of the SdHO amplitude matches that of MIRO *quantitatively*, without any adjustable parameters. This finding allowed us to deduce an empirical relation for the SdHO amplitude in irradiated 2DEG, given by Eqs. (10),(11). Taken together, our study reveals that current understanding of SdHO in irradiated 2DEG is lacking, calling for further investigations.

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- [1] M. A. Zudov, R. R. Du, J. A. Simmons, and J. L. Reno, Phys. Rev. B **64**, 201311(R) (2001).
 - [2] P. D. Ye, L. W. Engel, D. C. Tsui, J. A. Simmons, J. R.

- Wendt, G. A. Vawter, and J. L. Reno, Appl. Phys. Lett. **79**, 2193 (2001).
- [3] V. I. Ryzhii, Sov. Phys. Solid State **11**, 2078 (1970).
- [4] A. C. Durst, S. Sachdev, N. Read, and S. M. Girvin, Phys. Rev. Lett. **91**, 086803 (2003).
- [5] X. L. Lei and S. Y. Liu, Phys. Rev. Lett. **91**, 226805 (2003).
- [6] M. G. Vavilov and I. L. Aleiner, Phys. Rev. B **69**, 035303 (2004).
- [7] I. A. Dmitriev, M. G. Vavilov, I. L. Aleiner, A. D. Mirlin, and D. G. Polyakov, Phys. Rev. B **71**, 115316 (2005).
- [8] A. T. Hatke, M. A. Zudov, L. N. Pfeiffer, and K. W. West, Phys. Rev. Lett. **102**, 066804 (2009).
- [9] D. Konstantinov and K. Kono, Phys. Rev. Lett. **103**, 266808 (2009).
- [10] D. Konstantinov, Y. Monarkha, and K. Kono, Phys. Rev. Lett. **111**, 266802 (2013).
- [11] M. A. Zudov, O. A. Mironov, Q. A. Ebner, P. D. Martin, Q. Shi, and D. R. Leadley, Phys. Rev. B **89**, 125401 (2014).
- [12] I. A. Dmitriev, M. Khodas, A. D. Mirlin, D. G. Polyakov, and M. G. Vavilov, Phys. Rev. B **80**, 165327 (2009).
- [13] I. A. Dmitriev, A. D. Mirlin, D. G. Polyakov, and M. A. Zudov, Rev. Mod. Phys. **84**, 1709 (2012).
- [14] M. Khodas and M. G. Vavilov, Phys. Rev. B **78**, 245319 (2008).
- [15] M. Khodas, H. S. Chiang, A. T. Hatke, M. A. Zudov, M. G. Vavilov, L. N. Pfeiffer, and K. W. West, Phys. Rev. Lett. **104**, 206801 (2010).
- [16] Q. Zhang, T. Arikawa, E. Kato, J. L. Reno, W. Pan, J. D. Watson, M. J. Manfra, M. A. Zudov, M. Tokman, M. Erukhimova, et al., Phys. Rev. Lett. **113**, 047601 (2014).
- [17] $\mathcal{P} = \mathcal{P}_+ + \mathcal{P}_-$, $\mathcal{P}_\pm = e^2 E^2 v_F^2 / 2\epsilon_{\text{eff}} \hbar^2 \omega^4 [(1 \pm \epsilon^{-1})^2 + (\omega\tau_{\text{em}})^{-2}]$, where $\tau_{\text{em}}^{-1} = n_e e^2 / 2\sqrt{\epsilon_{\text{eff}} \epsilon_0 m^* c}$, $2\sqrt{\epsilon_{\text{eff}}} = \sqrt{\epsilon} + 1$, $\epsilon = 12.8$ is the dielectric constant of GaAs, v_F is the Fermi velocity, and E is the microwave electric field.
- [18] The rate of scattering on angle θ can be expressed in terms of angular harmonics, $\tau_n = \tau_{-n}$, as $\tau_\theta^{-1} = \sum \tau_n^{-1} e^{in\theta}$.
- [19] R. G. Mani, J. H. Smet, K. von Klitzing, V. Narayana-murti, W. B. Johnson, and V. Umansky, Nature **420**, 646 (2002).
- [20] M. A. Zudov, R. R. Du, L. N. Pfeiffer, and K. W. West, Phys. Rev. Lett. **90**, 046807 (2003).
- [21] C. L. Yang, M. A. Zudov, T. A. Knuttila, R. R. Du, L. N. Pfeiffer, and K. W. West, Phys. Rev. Lett. **91**, 096803 (2003).
- [22] R. L. Willett, L. N. Pfeiffer, and K. W. West, Phys. Rev. Lett. **93**, 026804 (2004).
- [23] J. H. Smet, B. Gorshunov, C. Jiang, L. Pfeiffer, K. West, V. Umansky, M. Dressel, R. Meisels, F. Kuchar, and K. von Klitzing, Phys. Rev. Lett. **95**, 116804 (2005).
- [24] M. A. Zudov, R. R. Du, L. N. Pfeiffer, and K. W. West, Phys. Rev. B **73**, 041303(R) (2006).
- [25] M. A. Zudov, R. R. Du, L. N. Pfeiffer, and K. W. West, Phys. Rev. Lett. **96**, 236804 (2006).
- [26] D. Konstantinov and K. Kono, Phys. Rev. Lett. **105**, 226801 (2010).
- [27] S. I. Dorozhkin, L. Pfeiffer, K. West, K. von Klitzing, and J. H. Smet, Nat. Phys. **7**, 336 (2011).
- [28] A. V. Andreev, I. L. Aleiner, and A. J. Millis, Phys. Rev. Lett. **91**, 056803 (2003).
- [29] A. Auerbach, I. Finkler, B. I. Halperin, and A. Yacoby, Phys. Rev. Lett. **94**, 196801 (2005).
- [30] I. G. Finkler and B. I. Halperin, Phys. Rev. B **79**, 085315 (2009).
- [31] I. A. Dmitriev, M. Khodas, A. D. Mirlin, and D. G. Polyakov, Phys. Rev. Lett. **111**, 206801 (2013).
- [32] A. E. Kovalev, S. A. Zvyagin, C. R. Bowers, J. L. Reno, and J. A. Simmons, Solid State Commun. **130**, 379 (2004).
- [33] R. R. Du, M. A. Zudov, C. L. Yang, Z. Q. Yuan, L. N. Pfeiffer, and K. W. West, Int. J. Mod. Phys. B **18**, 3465 (2004).
- [34] P. S. Dorozhkin, S. V. Tovstonog, S. A. Mikhailov, I. V. Kukushkin, J. H. Smet, and K. von Klitzing, Appl. Phys. Lett. **87**, 092107 (2005).
- [35] S. I. Dorozhkin, A. A. Bykov, I. V. Pechenezhskii, and A. K. Bakarov, JETP Lett. **85**, 543 (2007).
- [36] A. Wirthmann, B. D. McCombe, D. Heitmann, S. Holland, K.-J. Friedland, and C.-M. Hu, Phys. Rev. B **76**, 195315 (2007).
- [37] R. G. Mani, Appl. Phys. Lett. **91**, 132103 (2007).
- [38] L.-C. Tung, C. L. Yang, D. Smirnov, L. N. Pfeiffer, K. W. West, R. R. Du, and Y.-J. Wang, Solid State Commun. **149**, 1531 (2009).
- [39] Z. Kvon, D. Kozlov, S. Danilov, C. Zoth, P. Vierling, S. Stachel, V. Belov, A. Bakarov, D. Dmitriev, A. Toropov, et al., JETP Lett. **97**, 41 (2013).
- [40] I. A. Dmitriev, A. D. Mirlin, and D. G. Polyakov, Phys. Rev. Lett. **91**, 226802 (2003).
- [41] X. L. Lei and S. Y. Liu, Phys. Rev. B **72**, 075345 (2005).
- [42] X. L. Lei and S. Y. Liu, Appl. Phys. Lett. **86**, 262101 (2005).
- [43] T. Ando, J. Phys. Soc. Jpn. **38**, 989 (1975).
- [44] G. Abstreiter, J. P. Kotthaus, and J. F. Koch, Phys. Rev. B **14**, 2480 (1976).
- [45] O. E. Raichev, Phys. Rev. B **78**, 125304 (2008).
- [46] O. M. Fedorych, M. Potemski, S. A. Studenikin, J. A. Gupta, Z. R. Wasilewski, and I. A. Dmitriev, Phys. Rev. B **81**, 201302 (2010).
- [47] For detailed description of \mathcal{A}_D see Eq. (6) in Ref. 46.
- [48] This contribution, which directly affects SdHO, was not considered in relation to MIRO, originating from corrections of order $\mathcal{O}(\lambda^2)$.
- [49] I. A. Dmitriev, J. of Phys.: Conf. Ser. **334**, 012015 (2011).
- [50] X. L. Lei and S. Y. Liu, Appl. Phys. Lett. **94**, 232107 (2009).
- [51] I. A. Dmitriev, A. D. Mirlin, and D. G. Polyakov, Phys. Rev. B **70**, 165305 (2004).
- [52] W. Zhang, M. A. Zudov, L. N. Pfeiffer, and K. W. West, Phys. Rev. Lett. **98**, 106804 (2007).
- [53] A. T. Hatke, H.-S. Chiang, M. A. Zudov, L. N. Pfeiffer, and K. W. West, Phys. Rev. B **77**, 201304(R) (2008).
- [54] A. T. Hatke, H.-S. Chiang, M. A. Zudov, L. N. Pfeiffer, and K. W. West, Phys. Rev. Lett. **101**, 246811 (2008).
- [55] W. Zhang, M. A. Zudov, L. N. Pfeiffer, and K. W. West, Phys. Rev. Lett. **100**, 036805 (2008).
- [56] S. Wiedmann, G. M. Gusev, O. E. Raichev, A. K. Bakarov, and J. C. Portal, Phys. Rev. Lett. **105**, 026804 (2010).
- [57] O. E. Raichev, Phys. Rev. B **81**, 165319 (2010).
- [58] I. A. Dmitriev, R. Gellmann, and M. G. Vavilov, Phys. Rev. B **82**, 201311(R) (2010).
- [59] Under conditions of our experiment, in sample I (II) we

estimate τ_q to be 2 (3) ps from SdHO and 4.4 (8.2) ps from MIRO. However, accurate determination of SdHO τ_q requires extremely slow sweeps at very low B which were not attempted because the exact knowledge of τ_q is not crucial for our study.

- [60] ρ_{sm} is not significantly different from ρ_0 , except in 2DEG exhibiting strong negative magnetoresistance [64–67].
- [61] Strictly speaking, $\rho_{sm\omega} \neq \rho_{sm}$ because of radiation-induced heating and T -dependence of ρ_{sm} .
- [62] While all samples surveyed have shown correlation between MIRO and SdHO, in many of them MIRO rapidly decayed with increasing f . As a result, modulation of the

SdHO was observed only up to $\epsilon \approx 2$, where the excess heating due to the proximity to the cyclotron resonance precluded quantitative analysis of the waveforms.

- [63] We are not suggesting that this scenario is justified.
- [64] A. T. Hatke, M. A. Zudov, J. L. Reno, L. N. Pfeiffer, and K. W. West, Phys. Rev. B **85**, 081304(R) (2012).
- [65] L. Bockhorn, P. Barthold, D. Schuh, W. Wegscheider, and R. J. Haug, Phys. Rev. B **83**, 113301 (2011).
- [66] Q. Shi, P. D. Martin, Q. A. Ebner, M. A. Zudov, L. N. Pfeiffer, and K. W. West, Phys. Rev. B **89**, 201301(R) (2014).
- [67] Q. Shi, M. A. Zudov, L. N. Pfeiffer, and K. W. West, Phys. Rev. B **90**, 201301(R) (2014).