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Phys. Rev. B 91, 020101 — Published 15 January 2015

DOI: 10.1103/PhysRevB.91.020101

## Prediction of a Stable Post-Post-Perovskite Structure from First Principles

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## Abstract

A novel stable crystallographic structure is discovered in a variety of  $ABO_3$ ,  $ABF_3$  and  $A_2O_3$ compounds, via the use of first principles. This novel structure appears under hydrostatic pressure, and can be considered to be a post-post-perovskite phase. It provides a successful solution to experimental puzzles in important systems, and is characterized by one-dimensional channels linked by group of two via edge-sharing oxygen/fluorine octahedra. Such organization automatically results in anisotropic elastic properties and new magnetic arrangements. Depending on the system of choice, this post-post-perovskite structure also possesses electronic band gaps ranging from zero to  $\simeq 10$  eV being direct or indirect in nature.

PACS numbers: 61.50.Ks,71.20.-b,75.25.-j

 $ABX_3$  perovskites (Pv) form an important class of crystal structures for which A and B are cations and X is typically the oxygen or fluorine anion. Perovskites display a wealth of phenomena, such as ferroelectricity, magnetism, multiferroicity, piezoelectricity, magnetoelectricity, charge and orbital orderings, superconductivity, etc.... As a result, they constitute a rich playground for research and are important for various technologies, which explains the flurry of activities that have been devoted to them<sup>1</sup>. Interestingly, recent works have shown that applying a hydrostatic pressure in some  $ABX_3$  materials can result in the transformation from the Pv structure to the so-called "post-perovskite" (pPv) structure which can have important physical consequences<sup>2–9</sup>. For instance, the pPv structure discovered in  $MgSiO_3$  explains the existence of anisotropic features in the D" layer of Earth<sup>8-10</sup>. Moreover, CaRhO<sub>3</sub> was recently found to adopt a polymorph that was described as being an intermediate phase between perovskite and post-perovskite<sup>11</sup>. Based on these discoveries as well as recent findings of new high-pressure phases in  $ABO_3$  and  $ABF_3$  systems<sup>12–14</sup>, one may wonder if there is another crystal structure (to be termed as "post-post-perovkite" (ppPv)) for which Pv or pPv materials can evolve to under hydrostatic pressure. Positively answering such question will deepen the current knowledge of crystallography and high pressure. Moreover, if such structure does exist, one may also wonder about its structural characteristics and if they can lead to novel physical properties – which is obviously interesting for fundamental reason but also for the design of original devices. It is also of high importance to determine what precise compounds may possess such hypothetical structure.

The goal of this Communication is to address all these aforementioned unknown questions, via the use of first-principles calculations. As we will see, surprises are in store since we, e.g., (1) predict that many and various  $ABX_3$  and  $A_2O_3$  materials can transform to a common, novel and stable ppPv structure (from a pPv structure or even directly from a Pv structure) under hydrostatic pressure; and (2) reveal its unusual structural, magnetic and electronic properties. Moreover, this ppPv structure is likely the "N-phase" that has been observed in Refs.<sup>2,3</sup>.

As detailed in the Supplementary Materials  $(SM)^{15}$  (see, also, references<sup>2,3,8–11,16–39</sup> therein), first-principles calculations are performed on many  $ABX_3$  and  $A_2O_3$  materials, with different A and B atoms and with X = O or F anion, under hydrostatic pressure. A list of these materials is indicated in Fig. 1.

Crystal structures. Let us first concentrate on a specific material that has been experi-

mentally explored under pressure, namely  $NaMgF_3$ . Figure 2(a) shows that the orthorhombic Pv Pnma phase (Pv-Pnma) is predicted to be its ground state up to  $\simeq 20$  GPa, as consistent with measurements $^{2,3,40}$ . Such phase is common to many perovskites<sup>20</sup> and is schematized in Fig. 2(b). In this phase, any fluorine (or oxygen) octahedra share corners with their neighboring octahedra along the pseudo-cubic [100], [010] and [001] directions. Figure 2(a) further reveals that  $NaMgF_3$  is predicted to experience a phase transition to the (orthorhombic) post-perovskite Cmcm phase (pPv-Cmcm) at  $\simeq 20$  GPa, for which not only the space group but also the crystallographic structure change, as schematized in Fig. 2(c). Interestingly, the Pv-Pnma-to-pPv-Cmcm transition has been observed to occur for pressure around 27-30 GPa and under laser heating (likely, to overcome the kinetic barrier inherent to first-order transitions)<sup>2,3</sup> in NaMgF<sub>3</sub>, which is rather consistent with our prediction of a corresponding critical pressure of  $\simeq 20$  GPa at 0 Kelvin. As indicated by Fig. 3 and Table I of the  $SM^{15}$ , the pPv-*Cmcm* phase differs from the Pv-*Pnma* structure by the existence of two-dimensional sheets formed by octahedra that share edges along the a-axis and corners along the *c*-axis. These two-dimensional sheets are stacked along the *b*-axis with an interlayer made of A atoms separating any two neighboring sheets. As a result, the elastic (stiffness) constant of pPv-Cmcm is much lower along the b-axis than along the a or c axis for any material, including NaMgF<sub>3</sub> (see Table II of the  $SM^{15}$  and MgSiO<sub>3</sub> – which, for this latter compound, is consistent with the seismic anisotropy observed in the so-called D" layer of  $Earth^{8,10}$ .

As also revealed by Fig. 2(a), we further found that NaMgF<sub>3</sub> undergoes another transition at  $\simeq 51$  GPa, for which the space group and crystallographic structure both change again: the resulting phase re-adopts the *Pnma* space group but within a different crystallographic structure that is termed "post-post-perovskite"<sup>3,41,42</sup> and that is denoted as ppPv-*Pnma* in the following. Its structural characteristics are shown in Figs. 2(d) and 2(e). Interestingly, while ppPv-*Pnma* has never been previously reported in any material, its present discovery solves a puzzle: it likely is the so-called "N-phase" that has been observed in Ref.<sup>2,3</sup>, based on the facts that (i) it experimentally appears as a result of a phase transformation from the pPv-*Cmcm* phase at 56 GPa under laser-heating of about 2000 K (as consistent with our predicted pPv-*Cmcm*-to-ppPv-*Pnma* transition for a critical pressure  $\simeq 51$  GPa at T = 0 K); (ii) the "N-phase" has been assigned an orthorhombic symmetry<sup>2</sup>, in line with the *Pnma* space group we presently found for our ppPv structure<sup>43</sup>; and (iii) our simulated X-Ray Diffraction pattern of ppPv-Pnma is consistent with the one experimentally found in Ref.<sup>3</sup> for this N-phase (see Fig. 4 of the SM<sup>15</sup>).

Remarkably, comparing Figs. 2(c) with 2(d) and 2(e) reveals that the transformation from pPv-Cmcm to ppPv-Pnma results in the breaking of the two-dimensional octahedra sheet at the shared corners in favor of one-dimensional channels that are elongated along the *b*-axis of the ppPv-*Pnma* structure. These channels organize themselves by group of two (with the two channels forming the double channel being parallel to each other along the *b*-axis), as a result of edge-sharing octahedra. As shown in Table I of the  $SM^{15}$ , for a given pressure of 60 GPa (which is rather close to the predicted pPv-Cmcm-to-ppPv-Pnma transition), the formation of these double channels leads, in  $NaMgF_3$ , to the b and c lattice constants of ppPv-Pnma being larger by 4.3% and 25.3%, respectively, than the a and b lattice constants of pPv-Cmcm. On the other hand, the *a* lattice parameter of ppPv-Pnmais smaller by 24.9% than the c lattice constant of pPv-Cmcm (note that the b-axis is parallel to the direction of the one-dimensional channel in ppPv-Pnma while it is perpendicular to the octahedra sheets in pPv-Cmcm, implying that comparisons have to be made between the (a, b, c) triad axis of ppPv-Pnma and the (c, a, b) triad axis of pPv-Cmcm). As shown in the inset of Fig. 2(a), such changes in lattice constants result in a decrease of 1.84% of the volume at the pPv-Cmcm-to-ppPv-Pnma transition in NaMgF<sub>3</sub>, which is a prediction that can be easily checked by measurements. Note that the octahedra are more distorted in ppPv-Pnma than in pPv-Cmcm, as evidenced by the facts that the six Mg-F bonds of the octahedra in ppPv-Pnma adopt four different values equal to 1.813 Å, 1.871 Å (doubly degenerate), 1.888 Å (doubly degenerate) and 1.942 Å, respectively, while those of pPv-*Cmcm* only split between two values of 1.785 Å (doubly degenerate) and 1.846 Å (four times degenerate), respectively, for a pressure of 60 GPa. The fluorine octahedra therefore become 0.84% larger in ppPv-Pnma than in pPv-Cmcm (even if the volume decreases), as edge-sharing allows for more compact packing. Moreover, in the ppPv-Pnma phase, any Mg ion belonging to one channel gets rather close to a specific F ion belonging to the adjacent channel (indicated by the dashed line in Fig. 2(d)) forming the double channels and therefore leads to an increase in coordination number from 6 to "6+1". For instance, at 60 GPa, the bond between these Mg and specific F ions is about 2.103 Å, which is comparable to the distances of 1.813 Å- 1.942 Å between Mg and F ions belonging to the same octahedra<sup>44</sup>.

Figure 1(a) also shows that many materials are also predicted to exhibit the aforemen-

tioned Pv-Pnma-to-pPv-Cmcm and pPv-Cmcm-to-ppPv-Pnma transitions, but at different critical pressures. On the other hand, Fig. 1(a) further indicates that some materials are predicted to *directly* transform from Pv-Pnma to ppPv-Pnma without adopting the intermediate pPv-Cmcm phase, as the pressure increases. Examples include (i) BiFeO<sub>3</sub> and  $BiCrO_3$  that are often studied for their multiferroic properties<sup>45,46</sup>; (ii) CaMnO<sub>3</sub> that has been predicted to exhibit both magnetic and electric orderings when grown as a strained  $film^{47,48}$ ; and (iii) the rare-earth ferrites  $RFeO_3^{28-30}$  with small or intermediate ionic radius. For instance and as shown in Fig. 1(a) of the SM<sup>15</sup>, GdFeO<sub>3</sub> directly undergoes a transition from Pv-Pnma to ppPv-Pnma at the pressure of  $\simeq 56.5$  GPa. Conversely, there are some materials, such as  $CaBO_3$  with B = Ru, Ir, Rh, Pt (that have been investigated because of their analogy with  $MgSiO_3^{4-7,49,50}$  or because of the strong effect of spin orbit interactions on some of their physical properties<sup>51</sup>) that do not exhibit the Pv-Pnma phase but rather evolve from pPv-Cmcm to ppPv-Pnma, as a hydrostatic pressure is applied and increased (see Fig. 1(b) of the  $SM^{15}$  for  $CaPtO_3$ ). In particular, we predict that the ppPv-Pnma phase of  $CaRuO_3$  will appear at a pressure of 33.8 GPa, which should make its observation rather easily feasible. Note that Fig. 1 also indicates that the volume is typically reduced by an amount varying between 0.9% and 3.8% (depending on the chemistry) at the critical pressure at which the post-post-perovskite structure becomes the most stable phase in our studied compounds. Moreover and as shown in Fig. 1(b), no ppPv structure was found upto 120 GPa in some other systems, such as  $RFeO_3$  compounds with large ionic radius (i.e., R = Nd, Pr, Ce and La, see Fig. 1(c) of the SM<sup>15</sup>), MgSiO<sub>3</sub>, Mn<sub>2</sub>O<sub>3</sub> or Al<sub>2</sub>O<sub>3</sub> – as consistent with measurements and previous computations<sup>8,9,32-34</sup> (note that the SM<sup>15</sup> provides a more detailed comparison between our predictions and these previous works).

Dynamical stability. The ppPv-Pnma structure is dynamically stable in its pressure range of stability for all the materials shown in Fig. 1(a). Two examples are shown in Figs. 2(a) and 2(c) of the SM<sup>15</sup> for NaMgF<sub>3</sub> and GdFeO<sub>3</sub>, respectively, both under a pressure of 60 GPa. In fact, we also numerically found that, in several studied compounds, ppPv-Pnma does not have any unstable phonon even in pressure regions for which this phase is not the lowest one in enthalpy. For instance and as shown in Fig. 1(b) of the SM<sup>15</sup>, ppPv-Pnma is dynamically stable even at zero pressure in, e.g., CaPtO<sub>3</sub>. As consistent with previous works<sup>52,53</sup>, this stability likely implies that this phase can be quenched to ambient pressure in this material (especially because the difference in enthalpy between pPv-Cmcm and ppPvPnma is found to be as small as 181 meV/5-atom at zero pressure in CaPtO<sub>3</sub>). Conversely, other phases, such as pPv-Cmcm, can also have no unstable phonon in the pressure range for which ppPv-Pnma has the lowest enthalpy, which implies that (i) pPv-Cmcm may still be experimentally found in some materials at pressure higher than the predicted pPv-Cmcm-to-ppPv-Pnma transition pressure and (ii) observing ppPv-Pnma phase in these materials may require the use of laser heating (to overcome kinetic barrier). Note that the aforementioned Pv-Pnma-to-pPv-Cmcm and pPv-Cmcm-to-ppPv-Pnma transitions are of reconstructive type, and that the Pv-Pnma, pPv-Cmcm and ppPv-Pnma phases form local minima that are linked by bond-breaking, which explains their dynamical stability.

Electronic structure. We also numerically found that, within ppPv-Pnma, the electronic band gap can be rather quantitatively different between investigated materials (see Table III and Fig. 2 of the SM<sup>15</sup>). For instance, the calculated band gap of NaMgF<sub>3</sub> is as large as 9.04 eV for a pressure of 60 GPa while it is dramatically reduced to 0.83 eV for GdFeO<sub>3</sub> under the same pressure. In fact, a few systems are even metallic above the pressure at which the ppPv-Pnma phase begins to appear. Examples include CaRhO<sub>3</sub> at 70 GPa and CaIrO<sub>3</sub> at 90 GPa. Equally striking and as shown in Fig. 2 of the SM<sup>15</sup> too, even the character of the band gap (that is direct versus indirect) can be altered when going from one material to another within ppPv-Pnma. Such electronic flexibility may result, in the future, to the discovery of anomalous properties (such as metal-insulator transitions<sup>54</sup>) or highly-desired features (such as a direct-band gap in the frequency spectrum needed for photovoltaic devices<sup>55</sup> or light-emitting devices<sup>56</sup>) in materials possessing the ppPv-Pnma

Magnetic ordering. Interestingly, some  $ABO_3$  materials, that are predicted to exhibit ppPv-Pnma structure, possess A and/or B atoms that are magnetic. As a result, novel or striking magnetic arrangements should emerge, especially when recalling that ppPv-Pnma adopts unusual "double" one-dimensional channels inside which A and B bond with O atoms (see Figs. 2(d) and 2(e)). Let us, for instance, consider the case of the ppPv-Pnma phase of GdFeO<sub>3</sub> at 60 GPa and include the 4f electrons of Gd in the valence in the calculations, thus allowing both Gd and Fe ions to adopt localized magnetic moments (that are found to be 6.90  $\mu_B$  and 4.12  $\mu_B$ , respectively). Practically, enthalpies of different collinear magnetic configurations are computed and used to extract the coupling coefficients  $(J_{BB,chain}, J_{BB,across}, J_{BA,single}, J_{BA,four})$  of the model described by  $H = \frac{1}{2}J_{BB,chain}\sum_{i,j} \mathbf{S}_i \cdot \mathbf{S}_j$ 

 $+\frac{1}{2}J_{BB,across}\sum_{i,j}\boldsymbol{S}_{i}\cdot\boldsymbol{S}_{j}+\frac{1}{2}J_{BA,single}\sum_{i,j}\boldsymbol{S}_{i}\cdot\boldsymbol{S}_{j}+\frac{1}{2}J_{BA,four}\sum_{i,j}\boldsymbol{S}_{i}\cdot\boldsymbol{S}_{j},$  where the sums over irun over all Fe atoms while the first (respectively, last) two sums over j run over specific Fe (respectively Gd) atoms that will be indicated below. As depicted in Fig. 3, the strongest coupling coefficient (denoted by  $J_{BB,chain}$ ) is found to be 2.86 meV (that is antiferromagnetic in nature) and is between Fe ions that are distant (by  $\simeq 3.02$  Å) along the *b*-axis. Interestingly, the coupling between Fe ions that belong to two adjacent and parallel one-dimensional channels (and are distant by  $\simeq 2.67$  Å) is also antiferromagnetic in nature but is of smaller magnitude since it is equal to 1.52 meV (this parameter is denoted here as  $J_{BB,across}$ ). As a result and as shown in Fig. 3(a), the magnetic ground state of GdFeO<sub>3</sub> possesses onedimensional antiferromagnetic chains elongating along the *b*-axis and formed by Fe ions with each of these Fe ions having two neighboring Fe ions of opposite spins and that belong to the adjacent parallel chain. Note that the particular triangular-like geometry seen by any magnetic B ion (see Fig. 3(a)) because of the formation of the double one-dimensional channels inherent to ppPv-Pnma in  $ABO_3$  materials is a perfect "recipe" to generate the so-called geometric frustration<sup>57,58</sup> in the specific (and presently hypothetical) case that  $J_{BB,chain}$  and  $J_{BB,across}$  would still be antiferromagnetic in nature but would now be close to each other in magnitude (unlike in  $GdFeO_3$ ). Searching for such compounds or the hypothetical pressure giving rise to such condition in some materials therefore constitutes a promising avenue to pursue in the future. Note also that we numerically found that, in the ppPv-Pnma phase of GdFeO<sub>3</sub> at 60 GPa, magnetic interactions between Gd ions are negligible (as consistent with the deep f-shell of Gadolinium) but Fe ions are antiferromagnetically coupled with their closest Gd ions. As indicated in Fig. 3(b), the resulting coupling is  $J_{BA,single} = 1.22$ meV between Fe and Gd ions that form single bond (and are distant by 2.783 Å) while it is  $J_{BA,four} = 0.73$  meV between Fe and Gd ions that are tetrahedrally bonded (and distant by 3.121 Å). As a result, the magnetic ordering of Gd ions is governed by their interaction with Fe ions and is the one depicted in Fig. 3(b).

In summary, we used first-principles techniques to discover a common and stable ppPv crystal structure in a variety of  $ABX_3$  and  $A_2O_3$  materials under pressure. Such phase exhibits one-dimensional structural characteristics which naturally lead to strong anisotropy and emergence of novel magnetic orderings, and provides a plausible explanation for the "N-phase" that has been reported in Refs.<sup>2,3</sup>. Moreover, the electronic band gap of this phase is highly dependent on the system and can be of rather different nature and magnitude.

We hope that this Communication will encourage researchers to confirm the predictions presently reported and to determine properties associated with such novel crystal structure.

This work is financially supported by the Department of Energy, Office of Basic Energy Sciences, under contract ER-46612 (B.X.) and ONR Grant N00014-12-1-1034 (Y.Y.). It is also supported by the Ministry of Science and Technology of China (Grant No. 2011CB606405) and National Natural Science Foundation of China (Grant No. 11174173). We also acknowledge ARO Grant W911NF-12-1-0085 and NSF grant DMR-1066158 for discussions with scientists sponsored by these grants. The calculations were performed on the "Razor" (Univ. of Arkansas) and "Explorer 100" (Tsinghua Univ.) cluster systems.

- <sup>1</sup> A. S. Bhalla, R. Guo, and R. Roy, *Material Research Innovations* 4, 3 (2000).
- <sup>2</sup> C. D. Martin, W. A. Crichton, H. Liu, V. Praparkenka, J. Chen, and J. B. Parise, American Mineralogist 91, 1703 (2006).
- <sup>3</sup> C. D. Martin, W. A. Crichton, H. Liu, V. Praparkenka, J. Chen, and J. B. Parise, *Geophys. Res. Lett.* 33, L11305 (2006).
- <sup>4</sup> C. L. McDaniel, and S. J. Schneider, J. Solid State Chem 4, 275 (1972).
- <sup>5</sup> H. Kojitani, Y. Shirako, and M. Akaogi, *Phys. Earth Planet. Inter.* **165**, 127 (2007).
- <sup>6</sup> Y. Shirako, H. Kojitani, M. Akaogi, K. Yamaura, and E. Takayama-Muromachi, *Phys. Chem. Miner.* **36**, 455 (2009).
- <sup>7</sup> K. Ohgushi, Y. Matsushita, N. Miyajima, Y. Katsuya, M. Tanaka, F. Izumi, H. Gotou, Y. Ueda, and T. Yagi, *Phys. Chem. Miner.* **35**, 189 (2008).
- <sup>8</sup> A. R. Oganov, and S. Ono, *Nature* **430**, 445 (2004).
- <sup>9</sup> M. Murakami, K. Hirose, K. Kawamura, N. Sata, and Y. Ohishi, *Science* **304**, 855 (2004).
- <sup>10</sup> A. R. Oganov, R. Martonak, A. Laio, P. Raiteri, and M. Parrinello, *Nature* **438**, 1144 (2005).
- <sup>11</sup> Y. Shikaro et al., American Mineralogist **97**, 159 (2012).
- <sup>12</sup> L. Zhang *et al.*, *Science* **344**, 877 (2014)
- <sup>13</sup> R. E. Cohen and Y. Lin, *Phys. Rev. B* **90**, 140102 (2014)
- <sup>14</sup> W. A. Crichton, F. L. M. Bernal, J. Guignard, M. Hanfland, S. Margadonna *arXiv*:1410.2783
- <sup>15</sup> See Supplemental Material at \*\*\* for details about the methods used, as well as additional information.

- <sup>16</sup> A. R. Oganov and C. W. Glass, *J. Chem. Phys.* **124**, 244704 (2006).
- <sup>17</sup> A. O. Lyakhov, A. R. Oganov, H. T. Stokes, and Q. Zhu, *Comput. Phys. Commun.* 184, 1172-1182 (2013).
- <sup>18</sup> A. R. Oganov, A. O. Lyakhov, and M. Valle, Acc. Chem. Res. 44, 227 (2011).
- <sup>19</sup> A. R. Oganov, Y. Ma, A. O. Lyakhov, M. Valle, and C. Gatti, *Rev. Mineral. Geochem.* **71**, 271 (2010).
- <sup>20</sup> J. Zhao, N. L. Ross, and R. J. Angel, Acta Cryst. B60, 263 (2004) and references therein.
- <sup>21</sup> G. Kresse and D. Joubert, *Phys. Rev. B* **59**, 1758 (1999).
- <sup>22</sup> J. P. Perdew, A. Ruzsinszky, G. I. Csonka, O. A. Vydrov, G. E. Scuseria, L. A. Constantin, X. Zhou, and K. Burke, *Phys. Rev. Lett.* **100**, 136406 (2008).
- <sup>23</sup> O. Diéguez, O. E. González-Vázquez, J. C. Wojdel, and J. Íñiguez, *Phys. Rev. B* 83, 094105 (2011).
- <sup>24</sup> P. E. Blochl, *Phys. Rev. B* **50**, 17953 (1994).
- <sup>25</sup> A. Togo, F. Oba, and I. Tanaka, *Phys. Rev. B* **78**, 134106 (2008).
- <sup>26</sup> W. S. Choi, D. G. Kim, S. S. A. Seo, S. J. Moon, D. Lee, J. H. Lee, H. S. Lee, D.-Y. Cho, Y. S. Lee, P. Murugavel, J. Yu, and T. W. Noh, *Phys. Rev. B* 77, 045137 (2008).
- <sup>27</sup> P. Xiao, J.-G. Cheng, J.-S. Zhou, J. B. Goodenough, and G. Henkelman, *Phys. Rev. B* 88, 144102 (2013).
- <sup>28</sup> R. L. White, *J. Appl. Phys.* **40**, 1061 (1969).
- <sup>29</sup> H. J. Zhao, W. Ren, Y. Yang, X. Chen, and L. Bellaiche, *J. Phys.: Condens. Matter* 25, 466002 (2013).
- <sup>30</sup> C. Xu, Y. Yang, S. Wang, W. Duan, B. Gu, and L. Bellaiche, *Phys. Rev. B* **89**, 205122 (2014).
- <sup>31</sup> S. Yakovlev, M. Avdeev, and M. Mezouar, Journal of Solid State Chemistry 182, 1545-1549 (2009).
- <sup>32</sup> J. Santillan and S. H. Shim, AGU Fall Meeting No. MR23B-0050 (2005).
- <sup>33</sup> S. Ono, A. R. Oganov, T. Koyama, and H. Shimizu, *Earth and Planetary Science Letters* 246, 326 (2006).
- <sup>34</sup> R. Caracas and R. E. Cohen, *Geophys. Res. Lett.* **32**, L16310 (2005).
- <sup>35</sup> K. Umemoto, R. M. Wentzcovitch, and P. B. Allen, *Science*, **311**, 983 (2006).
- <sup>36</sup> K. Umemoto, R. M. Wentzcovitch, D. J. Weidner, and J. B. Parise, *Geophys. Res. Lett.* 33, L15304 (2006).

- <sup>37</sup> B., Grocholski, S.-H. Shim, and Prakapenka, *Geophys. Res. Lett.* **37**, L14204 (2010).
- <sup>38</sup> F. Gao, Y. Yuan, K. F. Wang, X. Y. Chen, J.-M. Liu, and Z. F. Ren, *Appl. Phys. Lett.* 89, 102506 (2006).
- <sup>39</sup> A. V Krukau, O. A. Vydrov, A. F. Izmaylov, G. E. Scuseria, J. Chem. Phys. **125**, 224106 (2006).
- <sup>40</sup> M. OKeeffe, B. G. Hyde, and Bovin, *Phys. Chem. Minerals* 4, 299 (1979).
- <sup>41</sup> R. Caracas and R. E. Cohen, *Phys. Rev. B* **76**, 184101 (2007).
- <sup>42</sup> H. Yusa, Y. Shirako, M. Akaogi, H. Kojitani, N. Hirao, Y. Ohishi, and T. Kikegawa, *Inorg. Chem.* **51** 6559 (2010).
- <sup>43</sup> Note, however, that Ref.<sup>2</sup> tentatively assigns a *Pnnm* space group rather than *Pnma* for their N-phase. It will thus be interesting if the authors of Ref.<sup>2</sup> can check if *Pnma* better describes their experimental data, especially once realizing that *Pnnm* and *Pnma* share the same *mmm* point group.
- <sup>44</sup> Note that, for comparison, the closest distance between Mg and F ions that do not belong to the same octahedra is about 2.929 Å in the pPv-*Cmcm* phase at around 60 GPa, while the distances between Mg and F ions belonging to the same octahedra range between 1.785 Å and 1.846 Å.
- <sup>45</sup> J. Wang, J. B. Neaton, H. Zheng, V. Nagarajan, S. B. Ogale, B. Liu, D. Viehland, V. Vaithyanathan, D. G. Schlom, U. V. Waghmare, N. A. Spaldin, K. M. Rabe, M. Wuttig, and R. Ramesh, *Science* **299**, 1719 (2003).
- <sup>46</sup> M. Murakami, S. Fujino, S.-H. Lim, C. J. Long, L. G. Salamanca-Riba, M. Wuttig, I. Takeuchi, V. Nagarajan, and A. Varatharajan, *Appl. Phys. Lett.* 88, 152902 (2006).
- <sup>47</sup> S. Bhattacharjee, E. Bousquet, and P. Ghosez, *Phys. Rev. Lett.* **102**, 117602 (2009).
- <sup>48</sup> E. Bousquet, and N. Spaldin, *Phys. Rev. Lett.* **107**, 197603 (2011).
- <sup>49</sup> K. Hirose and Y. Fujita, *Geophys. Res. Lett.* **32**, L13313 (2005).
- <sup>50</sup> Y. Inaguma, K. Hasumi, M. Yoshida, T. Ohba, and T. Katsumata, *Inorg. Chem.* 47, 1868 (2008).
- <sup>51</sup> K. Ohgushi, J. I. Yamaura, H. Ohsumi, K. Sugimoto, S. Takeshita, A. Tokuda, H. Takagi, M. Takata, and T. H. Arima, *Phys. Rev. Lett.* **110**, 217212 (2013).
- Y. Shirakoa, b, Y.G. Shib, c, A. Aimia, D. Moria, H. Kojitania, K. Yamaurab, Y. Inagumaa,
  M. Akaogia, *Journal of Solid State Chemistry* 191 167-174 (2012)

- <sup>53</sup> A. C. Garcia-Castro, A. H. Romero, and E. Bousquet, *Phys. Rev. B* **90**, 064113 (2014).
- <sup>54</sup> M. Imada, A. Fujimori, and Y. Tokura, *Rev. Mod. Phys.* **70**, 1039 (1998).
- <sup>55</sup> J. M. Pearce, *Futures* **34**, 663 2012.
- <sup>56</sup> J. H. Burroughes, D. D. C. Bradley, A. R. Brown, R. N. Marks, K. Mackay, R. H. Friend, P. L. Burns, and A. B. Holmes, *Nature* **347**, 539 (1990).
- <sup>57</sup> R. Moessner and R. P. Ramirez, *Phys. Today* **59**, 24 (2006).
- <sup>58</sup> N. Choudhury, L.Walizer, S. Lisenkov, and L. Bellaiche, *Nature* **470**, 513 (2011).



FIG. 1: (Color online). Pressure range of stability of the Pv-Pnma, pPv-Cmcm and ppPv-Pnma phases in the  $ABX_3$  and  $A_2O_3$  materials under study. Panels (a) and (b) report materials possessing or missing, respectively, the presently discovered ppPv-Pnma structure for pressure up to 120 GPa. Numbers in these panels indicate the decrease in volume (in percent) at the transitions.



FIG. 2: (Color online). Pressure dependence of the enthalpy of the Pv-Pnma, pPv-Cmcm and ppPv-Pnma phases of NaMgF<sub>3</sub> (Panel (a)), along with the schematization of (b) the Pv-Pnma, (c) pPv-Cmcm and (d) and (e) ppPv-Pnma crystallographic structures. Note that the enthalpy of the Pv-Pnma phase has been set to be zero for any pressure in Panel (a), and that the inset of Fig. 2a displays the behavior of the volume versus pressure in the Pv-Pnma, pPv-Cmcm and ppPv-Pnma phases (in the pressure ranges they have the lowest enthalpy).



FIG. 3: (Color online). Magnetic ground state of the ppPv-Pnma phase of GdFeO<sub>3</sub> under 60 GPa. Panel (a) reports the strength of the magnetic interactions and the resulting magnetic ordering between Fe ions, while Panel (b) depicts the coupling coefficients associated with Fe and Gd magnetic interactions as well as the spin pattern adopted by these two types of ions.