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Collapse of Skyrmions in 2d Ferro- and Antiferromagnets

Liufei Cai, Eugene M. Chudnovsky, and D. A. Garanin Physics Department, Lehman College, City University of New York 250 Bedford Park Boulevard West, Bronx, New York 10468-1589, USA (Dated: July 11, 2012)

Collapse of a skyrmion due to the discreteness of a crystal lattice in isotropic two-dimensional ferro- and antiferromagnets has been studied analytically and by numerical solution of equations of motion for up to 2000×2000 classical spins on a square lattice coupled via Heisenberg exchange interaction. Excellent agreement between analytical and numerical results has been achieved. The lifetime of the skyrmion scales with its initial size, λ_0 , as $(\lambda_0/a)^5$ in ferromagnets and as $(\lambda_0/a)^{2.15}$ in antiferromagnets, with a being the lattice parameter. This makes antiferromagnetic skyrmions significantly shorter lived than ferromagnetic skyrmions.

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Skyrmions^{1,2} are topologically stable configurations of a fixed-length three-component vector field $\mathbf{n}(\mathbf{r})$ in the coordinate space of two dimensions. Due to the constraint $\mathbf{n}^2 = 1$ the \mathbf{n} -field has two independent components. This permits unique mappings of $\mathbf{n} = (n_x, n_y, n_z)$ onto $\mathbf{r} = (x, y)$, described by classes of homotopy³. Each homotopy class corresponds to a non-trivial field configuration characterized by a conserved topological charge. The emergence of a conserved charge from a continuous field theory prompted numerous studies of skyrmions in problems of high-energy and condensed matter physics⁴. They include cosmology⁵, Bose-Enstein condensates⁶, quantum Hall effect^{7,9} and anomalous Hall effect⁸, liquid crystals¹⁰.

The interest to skyrmions in ordered spin systems had received much attention soon after the discovery of high-temperature superconductivity in copper oxides $^{11-13,15-18}$, and further explored recently $^{19-22}$. It is related to the fact that superconductivity in copper oxides occurs in doped ${\rm CuO_2}$ layers that, when undoped, are square lattices of antiferromagnetically ordered spins. Initially there was some hope that interaction of electrons and holes with skyrmions could play some role in Cooper pairing but this was never successfully demonstrated. Some indirect evidence of skyrmions in the magnetoresistence of lanthanum copper oxide has been recently reported 23 but direct observation of skyrmions in 2d antiferromagnetic lattices is still lacking.

In a continuous field model like, e.g., the non-linear σ -model, the ground-state energy of the skyrmion does not depend on its size, λ . This follows from the invariance of the model with respect to the scale transformation $\mathbf{r} \to k\mathbf{r}$, where k is an arbitrary constant. If the skyrmion lives on a lattice, however, the scale invariance becomes broken due to the presence of a lattice parameter a. Thus the energy of the skyrmion depends on its size. This, in general, must lead to the collapse or expansion of the skyrmion, making it unstable. The nature of the exchange interaction on a lattice makes the skyrmion energy decreasing with its size, that leads to skyrmion collapse. A number of authors looked for interactions that could stabilize skyrmions in 2d ferromagnets^{24–26}.

It was argued that anisotropic crystal field added to the isotropic exchange model may, in principle, dynamically stabilize the skyrmion. In reality, however, anisotropic interactions are of relativistic origin, while the lattice effect that leads to the collapse of the skyrmion is of the exchange origin and thus much greater. Same is true about magnetic dipole-dipole interactions. In ferromagnets macroscopic skyrmions could, in principle, be a part of a stable domain structure. This, however, would not apply to microscopic skyrmions. In antiferromagnets long-range dipole-dipole interactions are negligible and they cannot stabilize skyrmions of any size. Therefore, it is important first to understand what is the mechanism of skyrmion collapse in a generic exchange model.

In this paper we study the dynamics of skyrmions and the dependence of their collapse time t_c on their initial size in a 2d square lattice of classical spins coupled via Heisenberg ferromagnetic (FM) or antiferromagnetic (AFM) exchange interaction. The accuracy of the continuous approximation increases with the size of the skyrmion, λ . One should, therefore, expect that the lattice skyrmion becomes stable in the limit of $\lambda \to \infty$. We find that t_c of the AFM skyrmion scales as $t_c \propto (\lambda_0/a)^{2.15}$ with its initial size λ_0 . We compute the dynamics of the collapse using both the analytical field model for the Neel vector and a direct numerical calculation on lattices of up to 2000×2000 exchange-coupled spins. The two approaches show excellent agreement with each other. For a 2d ferromagnet we obtain (up to logarithmic corrections) the $(\lambda_0/a)^5$ scaling of the lifetime. This makes skyrmions significantly shorter lived in a 2d AFM than in a 2d FM.

We begin with an antiferromagnet described by the Hamiltonian for the Néel vector L:

$$\mathcal{H}_0 = \frac{1}{2}JS^2 \int dxdy \left[\frac{1}{c^2} \dot{\mathbf{L}}^2 + (\nabla \mathbf{L})^2 \right]. \tag{1}$$

Here **L** is normalized as $\mathbf{L}^2 = 1$, $(\nabla \mathbf{L})^2 \equiv (\partial_x \mathbf{L})^2 + (\partial_y \mathbf{L})^2$, $JS^2 > 0$ is the exchange energy associated with the interaction of spins of length S, and c is the speed of AFM spin waves that equals $2\sqrt{2}Ja/\hbar$ in a square lattice. The term with $\dot{\mathbf{L}}^2$ can be understood as a kinetic energy responsible for the inertia of antiferromagnets.

The absolute minimum of the energy corresponds to the uniform AFM background, $\mathbf{L}=\mathrm{const.}$ Non-uniform configurations of \mathbf{L} are characterized by the topological charge

$$Q = \frac{1}{4\pi a^2} \int dx dy \, \mathbf{L} \cdot (\partial_x \mathbf{L} \times \partial_y \mathbf{L}) \tag{2}$$

that takes values $Q=0,\pm 1,\pm 2,\ldots$ Within, e.g., the homotopy class Q=-1 the minimum energy, static configuration is a skyrmion given by

$$\mathbf{L} = \left(\frac{2\lambda x}{r^2 + \lambda^2}, \frac{2\lambda y}{r^2 + \lambda^2}, \frac{r^2 - \lambda^2}{r^2 + \lambda^2}\right), \tag{3}$$

where $r^2 = x^2 + y^2$. Its energy, $E = 4\pi J S^2$, is independent of λ .

Equation (1) can be derived from the Heisenberg exchange interaction between nearest-neighbor classical spins $|\mathbf{s}^A| = |\mathbf{s}^B| = 1$,

$$\mathcal{H} = S^2 \sum_{ij} J_{ij} \mathbf{s}_i^A \cdot \mathbf{s}_j^B = -\frac{S}{2} \sum_{i \in A} \mathbf{s}_i^A \cdot \mathbf{H}_i^A - \frac{S}{2} \sum_{j \in B} \mathbf{s}_j^B \cdot \mathbf{H}_j^B,$$

where A and B denote AFM sublattices and $\mathbf{H}_{i}^{A,B} = -\delta \mathcal{H}/\delta(S\mathbf{s}_{i}^{B,A}) = -S\sum_{j}J_{ij}\mathbf{s}_{j}^{B,A}$ are the effective fields acting on the spins. As spins in each sublattice rotate smoothly through space, one can expand the effective fields as

$$\mathbf{H}_{i}^{A} = -JS \left[4\mathbf{s}_{i}^{B} + a^{2}\nabla^{2}\mathbf{s}_{i}^{B} + \frac{a^{4}}{12} \left(\partial_{x}^{4} + \partial_{y}^{4} \right) \mathbf{s}_{i}^{B} + \ldots \right]$$
(5)

and similar for \mathbf{H}_{i}^{B} . This allows one to go over to the continuum description in which there are two spin fields \mathbf{s}^{A} and \mathbf{s}^{B} . Switching to the magnetization $\mathbf{M} = (\mathbf{s}^{A} + \mathbf{s}^{B})/2$ and the Néel vector $\mathbf{L} = (\mathbf{s}^{A} - \mathbf{s}^{B})/2$, satisfying $\mathbf{M}^{2} + \mathbf{L}^{2} = 1$ and $\mathbf{M} \cdot \mathbf{L} = 0$, with the help of equations of motion $\hbar \dot{\mathbf{s}}^{A,B} = [\mathbf{s}^{A,B} \times \mathbf{H}^{A,B}]$ one obtains

$$\mathcal{H} = \mathcal{H}_0 - \frac{1}{24} J S^2 a^2 \int dx dy \left[\left(\partial_x^2 \mathbf{L} \right)^2 + \left(\partial_y^2 \mathbf{L} \right)^2 \right] , \quad (6)$$

which differs from Eq. (1) by the second term due to the discreteness of the lattice. If the size of the skyrmion λ is large compared to a, this term can be treated as a perturbation. Using the "rigid" skyrmion profile of Eq. (3), one obtains the energy due to this term

$$\mathcal{E}_{\text{discr}} = -(2\pi J S^2/3)(a/\lambda)^2 \tag{7}$$

that violates the scale invariance of the skyrmion. Eq. (7) can be interpreted as a potential energy responsible for the skyrmion collapse. During the collapse it is transformed into the kinetic energy defined by the integral of $\dot{\mathbf{L}}^2 = 4r^2(r^2 + \lambda^2)^{-2}\dot{\lambda}^2$. With account of energy conservation, Eq. (6) gives

$$\frac{3}{c^2} \left(\ln \frac{r_{\text{max}}^2 + \lambda^2}{\lambda^2} - \frac{r_{\text{max}}^2}{r_{\text{max}}^2 + \lambda^2} \right) \dot{\lambda}^2 = \left(\frac{a}{\lambda} \right)^2 - \left(\frac{a}{\lambda_0} \right)^2, \tag{8}$$

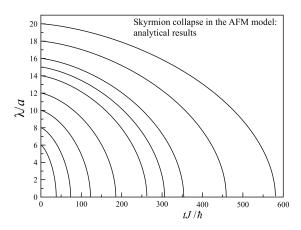


Figure 1: Collapse of antiferromagnetic skyrmions as described by the numerical solution of Eq. (8).

where λ_0 is the initial size of the skyrmion and $r_{\rm max}$ has been introduced because of the logarithmic divergence of the integral in the kinetic energy. The natural choice is $r_{\rm max} = \lambda_0 + ct$, which describes a front of AFM spin waves propagating away from the collapsing skyrmion. This is confirmed by direct numerical calculations, see Fig. 4 below. The logarithmic terms with time-dependent $r_{\rm max}$ require numerical integration of Eq. (8). The resulting collapse curves are shown in Fig. 1.

We now turn to the direct numerical solution of the dynamics of the skyrmion provided by the microscopic Hamiltonian (4). The dynamics is determined by the coupled equations of motion for spins, $\hbar \dot{\mathbf{s}}_i = -\left[\mathbf{s}_i \times \delta \mathcal{H}/\delta(S\mathbf{s}_i)\right]$. We chose initial state as a staggered skyrmion texture, \mathbf{s}^{st} , given by Eq. (3) for the A sublattice and by the same formula but with a minus sign for the B sublattice. The size of the skyrmion numerically can be defined as $\lambda_m^2 = (m-1)(2^m\pi)^{-1}\sum_i (1-s_{zi}^{\text{st}})^m$, where m>1 is an integer. If one replaces summation by integration over $dxdy/a^2$ and uses the skyrmion texture (3) for s_z^{st} , this formula becomes an identity, $\lambda_m = \lambda$. The results presented below have been obtained with m=4. Other options make little difference.

As the dynamics of the skyrmion is entirely due to small terms arising from the lattice discreteness, the time dependence is slow and sufficient accuracy can be achieved even for a large time step of integration. Increasing the step is limited by stability rather then by required accuracy. The challenge of the numerical solution is the 1/r decay of the skyrmion profile that requires rather big lattice sizes even for moderate values of λ/a . Free or periodic boundary conditions introduce spurious λ -dependent energies that compete with the small energy due to the lattice discreteness, leading to the expansion of the skyrmion instead of collapse. To make boundary conditions more resembling an infinite lattice, we have included the missing outside neighbors of the boundary spins with the values approximated by the second-

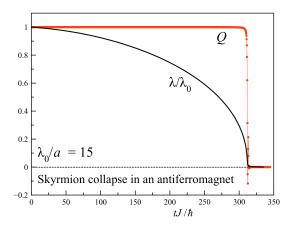


Figure 2: Skyrmion collapse in an antiferromagnet. Whereas the skyrmion size λ is decreasing continuously, the topological charge Q decays only during a short final stage of the collapse.

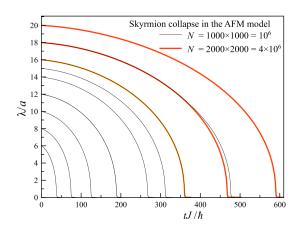


Figure 3: Skyrmion collapse in an antiferromagnet for different initial skyrmion sizes.

order extrapolation from the inside of the working region. Still, the lattice size has to be large: 1000×1000 for λ_0/a up to 16 and 2000×2000 for $\lambda_0/a=18$ and 20. The program was implemented in Wolfram Mathematica with a compiled vectorized fixed-step fourth-order Runge-Kutta routine. One AFM skyrmion-collapse event required about one hour computer time.

The collapse of an AFM skyrmion with $\lambda_0/a=15$ is shown in Fig. 2. Whereas the skyrmion size λ is decreasing continuously, the topological invariant Q changes only during a short final stage of the collapse when the continuous approximation fails. Fig. 3 shows skyrmion collapse curves for different values of λ_0/a . For $\lambda_0/a=18$ the lattice size of one million spins is too small and computation with four millions of spins is needed. For $\lambda_0/a=16$ these both lattice sizes yield the same collapse curve. These results compare very well with the semi-analytical results shown in Fig. 1. The collapse time can be fitted as $t_c \propto \lambda_0^{2.15}$ in this range of λ_0 . The considerable deviation

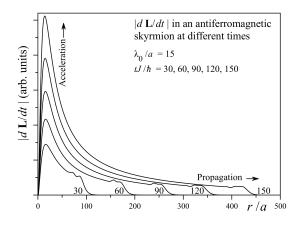


Figure 4: The front propagating from the center of antiferromagnetic skyrmion at the beginning of its collapse.

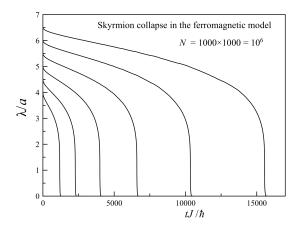


Figure 5: Skyrmion collapse in a ferromagnet. The collapse time scales as $t_c \propto \lambda_0^5$.

from the square law can be traced back to the logarithmic term in Eq. (8). Fig. 4 shows $|d\mathbf{L}/dt|$ in an antiferromagnetic skyrmion at different times. The region of skyrmion motion where $|d\mathbf{L}/dt| > 0$ is expanding with the speed of antiferromagnetic spin waves c. The reason for this is that the lattice-discreteness terms that drive the skyrmion collapse are very short ranged while the skyrmion itself is long-ranged. The action of the former is transferred to the whole skyrmion with a speed c in accordance with the causality. The front position can be estimated as $r_{\rm max} = \lambda_0 + ct$, as was argued after Eq. (8).

Along the same lines we have numerically studied the dynamics of ferromagnetic skyrmions. It turns out to be much slower than the collapse of AFM skyrmions, so that up to one day of computations is needed for one collapse event. Fig. 5 shows time dependences of the size of FM skyrmions during the collapse. The collapse time scales as $t_c \propto \lambda_0^5$.

The λ_0^5 scaling of the collapse time of the FM skyrmion can be qualitatively understood as follows. The exchange

interaction conserves the total spin of the system. The infinitesimal increase of the (negative) skyrmion spin in the course of its collapse is

$$dS = S \int \frac{d^2r}{a^2} \frac{ds_z}{d\lambda} d\lambda = -8\pi S \frac{\lambda d\lambda}{a^2} \ln \frac{R}{\lambda}, \qquad d\lambda > 0.$$
(9)

Here we used s_z in the skyrmion form given by Eq. (3) and introduced the long range cut-off R. Because of the conservation of the total spin, the increase of the skyrmion spin by $d\mathcal{S}$ generates $d\mathcal{S}$ magnons. Since in this process the spin is being carried by large distances, the skyrmion collapse is very slow. The average energy of emitted magnons can be estimated as $\hbar\omega \sim -\hbar\lambda/a$. This yields the emitted magnon power

$$P = \hbar \omega \frac{dS}{dt} = 8\pi \hbar S \frac{\lambda \dot{\lambda}^2}{a^3} \ln \frac{R}{\lambda} \,. \tag{10}$$

On the other hand, the rate of change of the energy (7) due to discreteness of the lattice is $\dot{\mathcal{E}}_{\mathrm{discr}} \propto \dot{\lambda}$. From the energy conservation, $\dot{\mathcal{E}}_{\mathrm{discr}} + P = 0$, one obtains

$$\frac{d\lambda}{dt} = -\frac{JSa^5}{6\hbar} \cdot \frac{1}{\lambda^4 \ln(R/\lambda)}, \qquad (11)$$

yielding the collapse time

$$t_c = \frac{6\hbar}{5JS} \left(\frac{\lambda_0}{a}\right)^5 \ln\left(\frac{R}{\lambda_0}\right). \tag{12}$$

The condition $\hbar\omega \ll SJ$ for the energy of the magnons translates to $5(\lambda/a)^4 \ln(R/\lambda) \gg 1$, which is well satisfied during the collapse.

In conclusion, we have studied the collapse of skyrmions due to the discreteness of the lattice in generic models of isotropic 2d ferro- and antiferromagnets with Heisenberg exchange interaction. The results obtained within continuous field model are in excellent agreement with the direct numerical calculation on lattices of up to 2000×2000 coupled spins. The collapse time of antiferromagnetic skyrmions obtained by both methods scales as $(\lambda_0/a)^{2.15}$. For ferromagnetic skyrmions, the numerical calculation gives the $(\lambda_0/a)^5$ scaling of the collapse time. It is explained by the emission of magnons. Thus, AFM skyrmions are much shorter lived than FM skyrmions. This can be understood in the following terms. The collapse of an AFM skyrmion occurs via transformation of its potential energy due to the discreteness of the lattice into the kinetic energy defined by $\dot{\mathbf{L}}^2$. The FM skyrmion does not possess such a kinetic energy, so that its potential energy has to be dissipated into magnons, which is a much slower process. In the expression for t_c the time constant in front of the power of the ratio λ_0/a is of order $\hbar/(JS)$. For, e.g., $JS \sim 100K$ and $\lambda_0 \sim 10a$, this gives $t_c \sim 10$ ns for the lifetime of the skyrmion in a ferromagnet and $t_c \sim 10$ ps in an antiferromagnet. Skyrmions of size exceeding one thousand lattice spacings would be practically stable in a ferromagnet. In antiferromagnets,

however, even macroscopic skyrmions would decay rather fast.

An interesting question is whether a skyrmion can be stabilized by the exchange interaction with an itinerant electron or a hole. If such interaction is stronger than the exchange interaction between magnetic atoms in a 2d lattice, the electron polarizes the background and the problem becomes one of the magnetic polaron. Only when the exchange interaction of the electron with the background is weak it can be considered as a perturbation of the skyrmion problem. In the case of a ferromagnet the energy of the electron in the uniform ferromagnetic background at infinity would always be lower than its energy in the vicinity of the skyrmion, thus ruling out the stability of any bound state. For an antiferromagnet, the discreteness of the lattice generates small uncompensated spin of the skyrmion. The electron can, in principle, couple to that spin by the exchange interaction. Our preliminary study shows, however, that this cannot prevent skyrmion from collapsing. Full analysis of this problem will be presented elsewhere.

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- ¹ T. H. R. Skyrme, Proc. Roy. Soc. London, Ser. A **247**, 260 (1958).
- A. A. Belavin and A. M. Polyakov, Pis'ma Zh. Eksp. Teor.
 Fiz 22, 503 (1975) [JETP Lett. 22, 245 (1975).
- ³ A. M. Polyakov, Gauge Fields and Strings, Harwood Academic Publishers 1987.
- ⁴ The Multifaceted Skyrmion, edited by G. E. Brown and M. Rho (World Scientific, 2010).
- ⁵ R. Durrer, M. Kunz, and A. Melchiorri, Phys. Rep. **364**, 1 (2002).
- ⁶ U. Al'Khawaja, and H. T. C. Stoof, Nature **411**, 918(2001).
- ⁷ S. L. Sondhi, A. Karlhede, S. A. Kivelson, and E. H. Rezayi, Phys. Rev. B 47, 16419 (1993).
- ⁸ Jinwu Ye, Y. B. Kim, A. J. Millis, B. I. Shraiman, P. Majumdar, and Z. Tesanovic, Phys. Rev. Lett. 83, 3737 (1999).
- ⁹ M. Stone, Phys. Rev. B **53**, 16573 (1996).
- ¹⁰ D. C. Wright, and N. D. Mermin, Rev. Mod. Phys. **61**, 385 (1989).
- ¹¹ P. B. Wiegmann, Phys. Rev. Lett. **60**, 821 (1988).
- ¹² B. I. Shraiman and E.D. Siggia, Phys. Rev. Lett. **61**, 467 (1988).
- ¹³ X. G. Wen and A. Zee, Phys. Rev. Lett. **61**, 1025 (1988).

- ¹⁴ F. D. M. Haldane, Phys. Rev. Lett. **61**, 1029 (1988).
- ¹⁵ S. Chakravarty, B. I. Halperin, and D. R. Nelson, Phys. Rev. B **39**, 2344 (1989).
- ¹⁶ P. Voruganti and S. Doniach, Phys. Rev. B **41**, 9358 (1990).
- ¹⁷ R. J. Gooding, Phys. Rev. Lett. **66**, 2266 (1991).
- ¹⁸ S. Haas, F.-C. Zhang, F. Mila, and T. M. Rice, Phys. Rev. Lett. **77**, 3021 (1996).
- E. C. Marino and M. B. Silva Neto, Phys. Rev. B 64, 092511 (2001).
- ²⁰ T. Morinari, Phys. Rev. B **65**, 064513 (2002).
- ²¹ T. Morinari, Phys. Rev. B **72**, 104502 (2005).
- ²² Z. Nazario and D. I. Santiago, Phys. Rev. Lett. **97**, 197201 (2006).
- ²³ I. Raicevic, D. Popovic, C. Panagopoulos, L. Benfatto, M. B. Silva Neto, E. S. Choi, and T. Sasagawa, Phys. Rev. Lett. **106**, 227206 (2011).
- ²⁴ A. Abanov and V. L. Pokrovsky, Phys. Rev. B **58**, R8889 (1998).
- ²⁵ B. A. Ivanov, A. Y. Merkulov, V. A. Stepanovich, C. E. Zaspel, Phys. Rev. B **74**, 224422 (2006).
- ²⁶ E. G. Galkina, E. V. Kirichenko, B. A. Ivanov, V. A. Stephanovich, Phys. Rev. B **79**, 134439 (2009).