



This is the accepted manuscript made available via CHORUS. The article has been published as:

Band-gap engineering in TiO_{2}-based ternary oxides

J. A. McLeod, R. J. Green, E. Z. Kurmaev, N. Kumada, A. A. Belik, and A. Moewes

Phys. Rev. B **85**, 195201 — Published 4 May 2012

DOI: 10.1103/PhysRevB.85.195201

Band gap Engineering in TiO₂-based Ternary Oxides

J. A. McLeod

Department of Physics and Engineering Physics, University of Saskatchewan, 116 Science Place, Saskatoon, Saskatchewan S7N 5E2, Canada*

R. J. Green

Department of Physics and Engineering Physics, University of Saskatchewan, 116 Science Place, Saskatoon, Saskatchewan S7N 5E2, Canada

E. Z. Kurmaev

Institute of Metal Physics, Russian Academy of Sciences-Ural Division, 620990 Yekaterinburg, Russia

N. Kumada

Department of Research Interdisciplinary Graduate School of Medicine and Engineering, University of Yamanashi, Miyamae-Cho 7, Kofu 400-8511, Japan

A. A. Belik

International Center for Materials Nanoarchitectonics (MANA), National Institute for Materials Science (NIMS), 1-1 Namiki, Tsukuba, Ibaraki 305-0044, Japan

A. Moewes

Department of Physics and Engineering Physics, University of Saskatchewan, 116 Science Place, Saskatoon, Saskatchewan S7N 5E2, Canada (Dated: February 17, 2012)

The electronic structure of several ternary oxides (Sn_2TiO_4 , $PbTiO_3$, $Bi_2Ti_4O_{11}$ and $Bi_4Ti_3O_{12}$) based on binary lone pair oxides (SnO, PbO, and Bi_2O_3) and a d^0 oxide (TiO_2) is investigated using soft X-ray spectroscopy and electronic structure calculations. We find that the valence band of these ternary oxides is bounded by bonding (at the bottom of the valence band) and antibonding (at the top of the valence band) O 2p - lone pair ns (Sn 5s, Pb 6s, Pb 6s) hybridized states, while the conduction band is dominated by unoccupied Ti 3d states. The existence of these two features is found to be independent of crystal structure or stoichiometry. The calculated hybridization in the bonding Palpenter Conduction and the valence bands in the ternary oxides is due to different aspects of the electronic structure in the parent binary oxides, and we consequently find that the band gap of the ternary oxide is found to be a stoichiometric-weighed addition of the band gaps of the parent oxides.

I. INTRODUCTION

The search for more efficient host materials for solar-driven hydrogen production has brought metal oxides (such as TiO₂, ZnO, and Fe₂O₃) into the light of renewed scientific attention in the last few years¹. Of these binary metal oxides, TiO₂ is one of the most promising photoanode materials for water splitting using solar radiation because of its suitable band structure — including a band gap straddling the reduction and oxidation potential of water — and low cost and high corrosion resistance in a wide variety of aqueous electrolytes². However the band gap of TiO₂ is too large (3.03 eV for rutile³ and 3.2 eV for anatase⁴) to absorb a significant part of the visible spectrum. Only UV radiation, which is about 4-5% of the solar spectrum, can drive photoconversion in TiO₂⁵. To achieve an acceptable hydrogen conversion efficiency (i.e. 15% or more) the band gap must be reduced below 2.2 eV⁶. Therefore, the band gap reduction of TiO₂ is a key requirement for effective utilization of solar radiation. One way of reducing the bandgap for the next generation of TiO₂-based photocatalysts is to combine multiple anions or cations to form ternary (or even more complex) oxides. Of all the possible ways of forming ternary oxides, four areas of research are of immediate promise for band gap reduction:

- 1. Doping TiO_2 by replacing oxygen atoms with anions having higher 2p orbital energy (such as B, C and N), creating bands^{7,8} or defect states within the nominal band gap^{9-12} .
- 2. Replacing Ti (d^0) with $3d^n$ and $4d^n$ cations (such as V, Cr, Nb, or Mo) introducing d-character to the valence band 13,14 .
- 3. Replacing Ti atoms with with $3d^{10}$ cations (such as Zn, Cd, or Ga) inducing $3d^{10}$ O 2p repulsion¹⁵.
- 4. Replacing Ti atoms with heavy metal ns^2 cations (such as Sn, Sb, Pb, or Bi) because of participation of 5s or 6s lone pairs in chemical bonding with O 2p states^{16,17}.

Of course a combination of these approaches is also possible ^{18,19}.

The focus of the present paper is to explore the last possibility. We study the electronic structure of ternary oxides formed by the precursors TiO₂ and SnO, PbO or Bi₂O₃. Soft X-ray spectroscopy measurements and electronic structure calculations of TiO₂, SnO, PbO, Bi₂O₃ and ternary Sn₂TiO₄, PbTiO₃, Bi₄Ti₃O₁₂ and Bi₂Ti₄O₁₁ are presented and the driving factors behind the reduction in band gap of the ternary oxides from that of pure TiO₂ are examined.

II. SYNTHESIS, MEASUREMENT, AND CALCULATION DETAILS

Synthesis: Single crystals of Sn_2TiO_4 were prepared by heating a mixture of K_2CO_3 and TiO_2 with a molar ratio of 1:1 in air at 600 - 700 °C for 6 h. The product was then mixed with $SnCl_2$ with a molar ratio of Sn/Ti = 1 - 2. This mixture was then put into a quartz tube and dehydrated at 200 °C for 1 h in vacuum. The tube was then sealed and heated at 700 °C for 24 h. The product was washed with distilled water to remove chlorides and dried at 50 °C. The full details of the preparation and structural characterization of this material are published elsewhere.²⁰

The $Bi_4Ti_3O_{12}$ powder sample was prepared from a stoichiometric mixture of Bi_2O_3 (Alfa Aesar, 99.9999 %) and TiO_2 (Alfa Aesar, 99.999 %) by annealing in air at 850 °C for 50 h with two intermediate grindings. The phase purity was checked with XRD, our results were the same as those available in the literature. PbTiO₃ and $Bi_2Ti_4O_{11}$ were both commercially available powders (Alfa Aesar, 99.9%). Rutile TiO_2 , SnO, litharge PbO, and Bi_2O_3 powders were also commercially available (Alfa Aesar, 99.9%).

Spectroscopy Measurements: The X-ray emission spectroscopy (XES) measurements were performed at Beamline 8.0.1 of the Advanced light Source (ALS) at Lawrence Berkeley National Laboratory. The endstation uses a Rowland circle geometry X-ray spectrometer with spherical gratings and an area sensitive multichannel detector²². The O K XES were excited near the O 1s ionization threshold (~ 540 eV) to suppress the high-energy satellite structure²³. The spectrometer resolving power (E/ Δ E) for emission measurements was about 10^3 .

The X-ray absorption spectroscopy (XAS) measurements were performed at the Spherical Grating Monochromator (SGM) beamline of the Canadian Light Source (CLS) at the University of Saskatchewan. The absorption measurements were acquired in total fluorescence mode (TFY) using a channel plate fluorescence detector²⁴. The monochromator resolving power (E/ Δ E) for absorption measurements was 2×10^3 . All absorption spectra were normalized to the incident photon current using a highly transparent gold mesh in front of the sample to correct for intensity fluctuations in the incident photon beam. For all X-ray measurements the above samples were pressed into clean indium foil and measured in ultrahigh vacuum ($\sim10^{-7}$ - 10^{-8} torr).

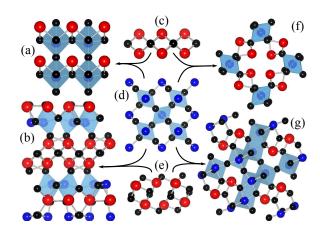


FIG. 1. (Color online) Crystal structures of the oxides studied herein. (a) PbTiO₃, (b) Bi₄Ti₃O₁₂, (c) SnO or PbO, (d) TiO₂, (e) Bi₂O₃, (f) Sn₂TiO₄, (g) Bi₂Ti₄O₁₁. In the case of Bi₄Ti₃O₁₂ only half of the unit cell is shown, the full unit cell is formed by mirroring the shown atoms upwards in the plane of the page. In the other cases slightly more than a single unit cell is shown to emphasize the site symmetry of the cations. These crystal structure diagrams were produced using VESTA 2.1.0³⁶.

Electronic Structure Calculations: Density functional theory (DFT) calculations for the experimentally determined crystal structures of ternary oxides Sn_2TiO_4 (space group $P4_2/mbc)^{20}$, PbTiO₃ (space group $P4mm)^{25}$, $Bi_4Ti_3O_{12}$ (space group $B2cb)^{26}$, $Bi_2Ti_4O_{11}$ (α phase, space group $C2/c)^{27}$, TiO_2 (rutile, space group $P4_2/mnm)^{28}$, SnO (space group $P4/nmm)^{29}$, PbO (litharge, space group $P/nmm)^{30}$, and Bi_2O_3 (α phase, space group $P2_1/c)^{31}$ based on the full-potential augmented plane-wave method with scalar-relativistic corrections were performed using the WIEN2k code³². For the exchange-correlational functional the Perdew-Burke-Ernzerhof (PBE) variant of the generalized gradient approximation³³ was used. To improve the calculated band gaps, we have also added a calculation cycle to the PBE calculation using the modified Becke Johnson (mBJ) exchange potential³⁴. The atomic sphere radii (used to determine whether plane waves or sperical harmonics will be used in the expansion of the wavefunction) for all cases were chosen such that they were nearly touching, and $R_{MT}^{min}K_{max} = 7$ (the product of the smallest of the atomic sphere radii R_{MT} and the plane wave cutoff parameter K_{max}) was used for the expansion of the basis set. The Brillouin zone integrations were performed on a k-point grid of up to 1000 points (the exact number of k-points depends on the symmetry of the unit cell). In all cases the energy convergence was better than 10^{-4} Ry/unit cell and the charge convergence was better than 10^{-4} e/unit cell.

The oxygen K XES and 1s XAS was calculated from the ground state bandstructure using the "XSPEC" package included with WIEN2k, which simulates the measured spectra by multiplying the partial density of states with a dipole transition matrix and a radial transition probability³⁵. The calculated XES and XAS were broadened with a Lorentzian function (with an energy dependent width varying quadratically from 0.1 eV at the upper band edge to 0.5 eV at the lower band edge) to mimic the core hole lifetime broadening, and a Gaussian function (with an energy dependent width based on the aforementioned experimental resolving power) to mimic the instrumental resolution.

The energy scale of the calculated XES spectra were shifted so the main spectral peak was aligned with the main spectral peak in the measured XES spectra. The calculated XAS spectra were shifted by the same amount; in this manner the energy of the most intense measured and calculated XES spectral feature coincides for each material by design, any agreement between other calculated and measured XES or XAS features may be regarded as an indication of the accuracy of the calculation.

III. RESULTS AND DISCUSSION

The crystal structure of the ternary oxides Sn_2TiO_4 , $PbTiO_3$, $Bi_4Ti_3O_{12}$ and $Bi_2Ti_4O_{11}$, in particular the local symmetry of the constituent atoms, bears some similarity to their precursor binary oxides. As shown in Figure 1, we may separate the ternary oxides into two groups: those consisting of alternating layers of TiO_6 octahedra and cation - oxygen planes (namely, $PbTiO_3$ and $Bi_4Ti_3O_{12}$), and those consisting of groups of TiO_6 octahedra separated by cation - oxygen channels (namely Sn_2TiO_4 and $Si_2Ti_4O_{11}$).

To what degree does the similarity in local structure between the ternary oxides and their precursor binary oxides indicate a similarity in electronic structure? For an arbitrary material we may not expect any general similarity, but for these ternary oxides there are two key factors to keep in mind:

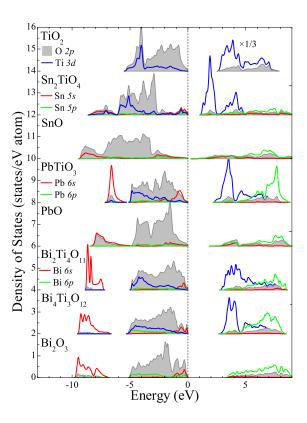


FIG. 2. (Color online) Calculated DOS of the materials studied herein. The top of the valence band is at 0 eV. ^a The calculations predict all materials have a band gap, although the gap for SnO is quite small. The conduction band Ti 3d states have been reduced in amplitude by a factor of 3. Note how the lone pair cation states (Sn 5s, Pb and Bi 6s) are split into bonding and antibonding states at the bottom and top of the valence band, respectively.

- ^a For brevity, we will call the top of the valence band the "Fermi level", even though technically the Fermi level is undefined for band gap materials.
 - 1. We expect the conduction band to be dominated by Ti 3d states (since in all cases the Ti is nominally in a $3d^0$ state), and
 - 2. We expect the lone pair Sn 5s, Pb/Bi 6s states to strongly affect the valence band.

In regard to the former point, the distribution of transition metal 3d states is strongly driven by the local symmetry^{37,38}, and since in all cases the Ti atoms are in distorted octahedra we may expect the distribution of these states in all ternary oxides considered herein to be similar to those of the precursor TiO_2 . With regard to the latter point, because the lone pair Sn 5s, Pb/Bi 6s states are nominally directed away from the ligands, and all stereochemical activity is confined to the same species of ligand, we may expect some similarity in the distribution of these states in the ternary oxides compared to the precursor binary oxides.

To examine these qualitative predictions in more detail, we turn to our X-ray spectroscopy measurements and DFT calculations of the materials in question. XES measurements involve the transitions from occupied electron states to core levels, while XAS measurements involve the transitions from core levels to unoccupied states. This process is dominated by dipole selection rules, and XES measurements involving a core level with angular momentum ℓ therefore probe the local $\ell \pm 1$ projected character of the valence band, and XAS measurements probe the local $\ell \pm 1$ projected character of the conduction band. Note that the conduction band probed by XAS is distorted from the true ground state conduction band by the presence of a local core hole in the final state of the absorption process.

Oxygen K XES measurements (probing the $2p \to 1s$ transition) provide a good experimental probe of the ground state oxygen 2p valence band. The oxygen states strongly hybridize with bonding cation states, $^{39-41}$ and consequently these measurements can provide insight into the cation electronic structure. Further, although the presence of an O 1s core hole in oxygen 1s XAS measurements (probing the $1s \to 2p$ transition) distorts the shape of the local conduction band, when oxygen is bonded to heavy cations the core hole does not usually shift the measured XAS in energy relative to the ground state conduction band⁴¹. Therefore the combination of O K XES and 1s XAS can provide an estimation of the band gap (be it direct or indirect).

Electronic Structure: The calculated density of states (DOS) for the ternary oxides and their precursors is shown in Figure 2. It is immediately clear that the electronic structures of these materials have many aspects in common. As we expect, the valence band in all of these materials is dominated by O 2p states. Where Ti 3d states are present, they are hybridized with the O 2p states, primarily at 4-5 eV below the Fermi level. These Ti 3d states dominate the onset of the conduction band as well, and typically are separated into two different features (although the separation is weak for Bi₂Ti₄O₁₁); a rather sharp feature directly at the onset of the conduction band, and a somewhat less intense feature at 2 - 3 eV higher in the conduction band. Since the Ti sites are almost octahedral, the low and high energy Ti 3d - O 2p conduction band hybridization regions are often referred to as e_g and t_{2g} bands, respectively. ^{42,43} The heavy cation p states (Sn 5p, Pb and Bi 6p) have an almost negligible contribution to the valence band, and only make a significant contribution to the conduction band at about 5 eV above the Fermi level.

The lone pair cation states (Sn 5s, Pb and Bi 6s) provide a significant contribution to the valence band. These states are also hybridized with O 2p states, and bonding between O and the heavy cation create a O 2p - cation nsband at the bottom of the valence band. This band is actually separate from the main valence band for materials with Bi and to a lesser extent, Pb. There is also a weaker, but still significant, O 2p - cation ns antibonding region at the Fermi level. This lone-pair hybridization is consistent with previous studies. 44-46 It is worth noting that one can qualitatively see the effects of combining the lone-pair oxides with TiO_6 octahedra in the DOS of the ternary oxides. In all cases the bonding O 2p - cation ns states of the ternary oxides are shifted to slightly higher energies than the corresponding states in the binary oxides. Similarly, the Ti 3d states at the onset of the conduction band are shifted towards the onset of the conduction band in the lone pair oxide. This is easiest to see in the case of Sn₂TiO₄: SnO is calculated to have a very narrow band gap, and the conduction Ti 3d states in Sn_2TiO_4 are about 2 eV lower in energy than the corresponding states in TiO₂. This trend also occurs in PbTiO₃, Bi₂Ti₄O₁₁, and Bi₄Ti₃O₁₂, although it is harder to see (the Ti 3d conduction states in PbTiO₃ start ~ 0.4 eV lower in energy than in TiO₂, while the corresponding states in $Bi_2Ti_4O_{11}$ and $Bi_4Ti_3O_{12}$ are ~ 0.2 eV higher in energy). Finally, in the case of PbTiO₃ and Sn_2TiO_4 , the antibonding cation ns states at the Fermi level are more localized in energy, this is probably due to the greater concentration of states at the Fermi level in the TiO₆ octahedra (relative to pure SnO and PbO) limiting the curvature of the band structure. These observations suggest there is some merit to our initial hypothesis that the electronic structure of a d^0 -lone pair ternary oxide may be described in terms of the electronic structure of both binary precursors.

X-ray Spectroscopy: The measured and calculated O K XES and 1s XAS spectra are shown in Figure 3. The calculated O K XES spectra generally agree quite well with the measured ones. There is also reasonable agreement between the calculated and measured O 1s XAS spectra, although since our calculations did not include a core hole we do not expect the agreement to be perfect. In fact, since the XES measurements agree with the calculations, we can attribute the differences between the measured and calculated XAS spectra to the influence of the O 1s core hole. We have previously found that the main influence of a O 1s core hole is to increase the intensity of the XAS at the bottom of the conduction band, 41 this seems to be true for the ternary and binary oxides studied herein as well, although the O 1s core hole seems to shift the XAS to lower energies in $\mathrm{Bi}_2\mathrm{O}_3$ (by about 1 eV), and to a lesser extent, in $\mathrm{Bi}_4\mathrm{Ti}_3\mathrm{O}_{12}$ (by about 0.5 eV). In contrast, the O 1s core hole seems to have a minor effect on the XAS of $\mathrm{Sn}_2\mathrm{TiO}_4$ and SnO .

The measured spectra have several features in common. The TiO₂ XES spectrum shows two clear features (labelled A and B in Figure 3), which we can attribute to O 2p - Ti3d hybridization for A, and O 2p states for B with the help of the DOS calculations shown in Figure 2. The TiO₂ XAS spectrum also has two clear near edge features, labeld C and D, these can be attributed to O 2p hybridization with the Ti 3d e_g and t_{2g} bands. The ternary and binary oxides containing lone pairs also show low energy structure in their XES spectrum that is related to bonding O 2p - cation ns states, labelled S_1 in Figure 3. There is also high energy structure in the O XES spectrum of SnO, Sn₂TiO₄, PbO, and Bi₂O₃ that is related to antibonding O 2p - cation ns states (labelled S_2). This identification of the S_1 and S_2 features follows from analysis of the calculated DOS discussed previously. Identifying S_2 in the XES spectrum of PbTiO₃, Bi₂Ti₄O₁₁ and Bi₄Ti₃O₁₂ is difficult, and not unambiguous. However since the calculated XES spectra for these materials do not suggest any discontinuous behaviour from the main O 2p states at B to the Fermi level, this is not unexpected. Small vertical bars in Figure 3 denote the features S_1 , A, B, S_2 , C, and D in the XES and XAS for each compound. Note that in the binary oxides SnO, PbO, and Bi₂O₃, we attribute feature B to O 2p - cation np hybridization, as suggested by the DOS in Figure 2.

There are two general observations one may make about the measured compared to the calculated XES spectra: The main peak B is typically sharper and more intense relative to the other spectral features in the measured spectrum than predicted by the calculated spectrum. This is likely due to lattice defects in the materials destroying the perfect crystalline periodicity assumed by the calculations, and consequently introducing a greater localization of states. The other general observation is that, save for PbO and perhaps SnO, the S_1 spectral feature is at lower energies than predicted by the calculated spectrum. This may be due to the calculation underestimating O 2p - cation np bonding affinity. We also note that S_1 in the ternary oxides S_{12} TiO₄ and PbTiO₃ is closer in energy to S_1 in their parent

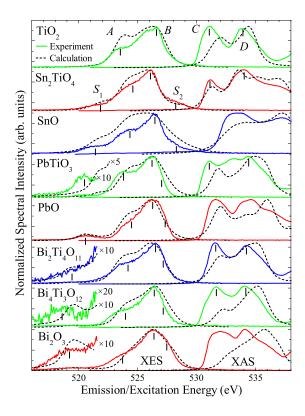


FIG. 3. (Color online) Measured and calculated O K XES and 1s XAS spectra of the materials studied herein. The calculated spectra were shifted so the main O K XES peak was aligned with the experimental peak. The key spectral features (labelled S_1 , A, B, S_2 , C, and D) are denoted by small black lines on each spectrum. Note TiO₂ lacks the $S_{1,2}$ XES features, while SnO, PbO, and Bi₂O₃ lack the C, D XAS features.

binary oxides (SnO and PbO, respectively), than suggested by the calculation, this indicates that the calculation may overestimate the influence of TiO_6 octahedra on the bonding structure. Dispite these differences, however, the basic agreement in features, relative intensities, and energy spacings of the calculated and measured XES and XAS spectra suggest that our calculated DOS is accurate for these systems.

O 2p Hybridization with Cations: We can use our experimental XES spectra to quantitatively estimate the strength of the O 2p - cation ns bonding by taking the ratio of the integrated intensity of the S_1 XES feature I_{S_1} to the integrated intensity of the total XES spectrum I_{tot} , and dividing by the ratio of the number of cation atoms n and the number of electron pairs p per formula unit of the lone pair binary oxide. This analysis for Bi_2O_3 and PbO is available in the literature, 45 and is commonly called the $f_{O\,2p}$ value:

$$f_{O2p} = \frac{I_{S_1}}{I_{tot}} \left(\frac{p}{n}\right)$$

Our estimates for f_{O2p} are shown in Figure 4(a). These empirical values can be compared directly to the ratio of the total O 2p states to the total O 2p and cation ns states in this region (using the calculated DOS).⁴⁵

The technique of using $f_{O\,2p}$ to estimate the degree of O 2p - cation ns hybridization is justified, given the agreement shown in Figure 4(a). Importantly, we can also estimate whether $f_{O\,2p}$ will over- or underestimate the hybridization by a qualitative inspection of the XES. Referring back to Figure 3, we see that in the case of SnO and Sn₂TiO₄ the S_1 feature is relatively intense and is more of a shoulder to the main XES emission than a separate feature. We can therefore predict that $f_{O\,2p}$ will be an underestimate of the true hybridization, since the the S_1 feature probably extends somewhat into the main XES. On the other hand, for Bi₂O₃ and Bi₂Ti₄O₁₁ the S_1 feature is relatively quite weak. We can therefore predict that $f_{O\,2p}$ will be an overestimate of the true hybridization, since the broadened emission from the main DOS region will extend into S_1 , providing a significant contribution to the weak hybridization emission. Bi₄Ti₃O₁₂ has lowest intensity S_1 feature, but this feature is also quite low in energy. The separation of this feature from the emission band minimizes the amount of "leakage" that can occur through broadening. We therefore expect that $f_{O\,2p}$ will be quite close to the true DOS hybridization, as indeed it is.

Indeed, the impact of spectral broadening on f_{O2p} is evident in Bi₄Ti₃O₁₂: the calculated XES has roughly double the S_1 intensity as the measured XES (refer back to Figure 3), although f_{O2p} is in excellent agreement with the

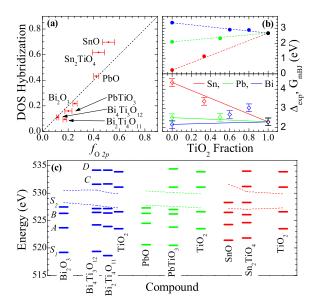


FIG. 4. (Color online) (a) $f_{O\,2p}$ derived from the XES compare to the calculated DOS hybridization (O 2p / (O 2p + cation ns) using the calculated DOS). The error bars denote the approximate spread of $f_{O\,2p}$ if the estimated end of region S_1 is shifted by ± 0.2 eV, the straight line indicates perfect agreement between $f_{O\,2p}$ and the DOS hybridization. (b) The calculated band gaps (G_{mBJ} , top) and estimated spectral gaps (Δ_{exp} , bottom) for these materials, expressed as a fraction of the TiO₂ content. The lines indicate a linear change in gap with TiO₂ content. (c) The energy of the measured XES spectral features as a function of TiO₂ content, the lines the Fermi level and the bottom of the conduction band estimated using the second derivative maxima.

DOS hybridization (refer to Figure 4(a)). Note that S_1 in the calculated XES is also about 1.5 higher in energy; it is therefore likely that the broadening of the main A and B features into S_1 is responsible for the larger intensity (this also probably applies to PbTiO₃, where a similar situation occurs). Finally, we should mention that the broadening applied to the calculated spectra was merely a "best guess", it could easily have been somewhat too large.

Band Gap Estimates: In our previous research on binary and ternary oxides, we have shown that the separation between the local maxima in the second derivatives of the O K XES and 1s XAS spectra provides a reasonably consistent method for estimating the band gap. 41,47,48 This is not a particularly rigourous method of determining the band edges, indeed it is only correct for DOSes which exhibit a limited range of behaviours near the band edge (here "near" means within the effective range of the spectral broadening), such as being linear or parabolic in energy. Looking back at Figure 2, we may anticipate that the second derivative method will provide reasonable estimates of the band gap for TiO₂ and the Bi-based compounds, since they all have relatively sharp and smooth behaviour near the Fermi level and the bottom of the conduction band. In the Sn-based materials, however, the valence band and (and conduction band for SnO) have very gradual onsets, it is unlikely that a single second derivative maximum can be unambiguously determined at the band edges. Note also that even if the DOS is shaped appropriately for the second derivative method, the aforementioned core hole shift will cause the X-ray "spectral gap" to be less than the true ground state band gap.

With these caveats in mind, it is worthwhile mentioning the estimated gap ($\Delta_{\rm exp}$) for these materials to compare with the calculated band gap ($G_{\rm mBJ}$) and values obtained from previous studies. These values are shown in Table I. As we expected, the $\Delta_{\rm exp}$ spectral gap for SnO is considerably larger than both our calculated gap and those obtained from optical measurements, because of the extremely gradual onset of both the valence and the conduction bands, we expect this value to be an over estimate. Indeed, the second derivative maximum of the SnO O K XES is at 527.2 eV, which is above the S_2 antibonding feature, and therefore clearly not at the Fermi level. In the same sense the Fermi level of Sn₂TiO₄ is also not accurately estimated by the XES second derivative maximum (this is at 526.9 eV, again above S_2), although the XAS second derivative maximum is within 0.1 eV of the calculated conduction band onset. For TiO₂, PbO, PbTiO₃, and Bi₂Ti₄O₁₁ the spectral gaps $\Delta_{\rm exp}$ are quite similar to the calculated gaps. If we add the estimated core hole shifts for Bi₂O₃ and Bi₄Ti₃O₁₂ mentioned above to $\Delta_{\rm exp}$ then we again get good agreement, but since the core hole shifts here were estimated by comparing the calculated and measured XAS the agreement is admittedly a bit contrived.

The band gaps reported from the literature in Table I are all derived from optical measurements, most of which were only probing the direct band gap. For SnO, the indirect gap was estimated as ~ 0.7 eV from the optical data,

Compound	$\Delta_{\rm exp}~[\pm~0.2~{\rm eV}]$	$G_{\rm mBJ}\ [{\rm eV}]$	Literature $[eV]$
${ m TiO_2}$	2.6	2.679	3.03^{3}
$\mathrm{Sn_2TiO_4}$	3.4	1.170	2.09^{49}
SnO	4.2	0.254	$2.7, \sim 0.7^{50}$
$PbTiO_3$	2.6	2.336	3.4^{51}
PbO	2.5	2.116	2.75^{52}
$\mathrm{Bi}_{2}\mathrm{Ti}_{4}\mathrm{O}_{11}$	2.8	2.887	3.75^{53}
$\mathrm{Bi_{4}Ti_{3}O_{12}}$	2.4	2.919	3.08^{54}
$\mathrm{Bi}_{2}\mathrm{O}_{3}$	2.2	3.388	2.91^{55}

TABLE I. Measured, calculated, and previously reported band gaps for all materials studied herein. The measured gaps are denoted by $\Delta_{\rm exp}$ and the calculated band gaps are denoted by $G_{\rm mBJ}$. The literature band gaps are all based on optical measurements save for Sn_2TiO_4 , which is exclusively based on a calculation. The second value for SnO is the estimated indirect gap, from the measurement of 2.7 eV for the direct gap.

but this analysis is not available for the other lone pair oxides Bi_2O_3 and PbO. Calculated band structures suggest these are all indirect band gap materials, so we may expect the literature values for PbO and Bi_2O_3 in Table I to overestimate the gap. For Bi_2O_3 , this perhaps suggests that our larger calculated gap may be a rare overestimate of a band gap by DFT (it is well known that DFT underestimates band gaps⁵⁶).

Trends in Electronic Structure: Due to the previously discussed features in the electronic structure particular to these binary and ternary oxides, namely the lone pair ns states bounding the valence band and the Ti 3d states defining the bottom of the conduction band, if we treat the ternary oxides as a mixture of the two precursor binary oxides (i.e. Sn_2TiO_4 is treated as $(SnO)_{2/3}(TiO_2)_{1/3}$ the calculated band gaps show an almost linear dependence on the mixing fraction, as shown in Figure 4(b). Interestingly, the estimated spectral gaps $\Delta_{\rm exp}$ for the Sn- and Pb-based compounds also show an almost linear dependence — despite the aforementioned inaccuracy of $\Delta_{\rm exp}$ as an estimate of the band gap for SnO and Sn_2TiO_4 . There is no linear trend in $\Delta_{\rm exp}$ for the Bi-based compounds, but this is perhaps due to the anomalous core-hole effect on Bi_2O_3 , as discussed above ("anomalous" in the fact that it appears to disproportionally affect Bi_2O_3 over any of the other Bi-based compounds).

Further, while Δ_{exp} does not necessarily indicate the band gap of a compound with lone pair s states, it can be used to determine whether a ternary oxide abides by this rule. If the Δ_{exp} of the ternary compound obtained from O K XES and 1s XAS is close to the "ideal" value (again by stoichiometry-weighted addition of Δ_{exp} of the parent oxides), then the band gap of the ternary compound may be inferred to be the stoichiometry-weighted addition of the band gaps of the parent oxides. This is useful because ternary oxides with a dilute stoichiometric addition of a particular parent oxide may have complicated crystal structures and make calculating their properties directly a computationally intensive task.

This is demonstrated by the energies of the XES spectral features, shown in Figure 3 and summarized in Figure 4(c). In Figure 4(c) we see how the TiO_2 content gradually lowers the energy of the antibonding S_2 feature at the top of the valence band (note that the dotted lines in Figure 4(c) show the band edges estimated using the second derivative maxima, here the problem with using the method for SnO and Sn_2TiO_4 is clear since S_2 is at higher energies than the second derivative maxima). Conversely, the presence of TiO_2 seems to stabilize the conduction band (features C and D).

This simple agreement in both calculated and measured data suggests that the true band gap of a ternary oxide formed from lone pair and a d^0 binary oxide is directly related to the stoichiometry-weighted addition of the band gaps of the precursors: in the examples studied herein the conduction band is qualitatively defined by the Ti $3d^0$ states, but increasing the stoichiometric content of lone pair oxides is responsible for progressively "dragging" of the conduction band edge to energies closer to that of the pure lone pair oxide. The same argument applies to the valence band edge; while it is qualitatively defined by antibonding s^2 states, increasing stoichiometric content of TiO₂ can "drag" the valence band edge closer to that of pure TiO₂.

Conclusions: To summarize, we have probed the electronic structure of ternary oxides Sn_2TiO_4 , $PbTiO_3$, $Bi_4Ti_3O_{12}$ and $Bi_2Ti_4O_{11}$ using O K XES, 1s XAS, and DFT calculations. These ternary oxides are formed from precursor binary lone pair oxides (SnO, PbO, and Bi_2O_3) and the d^0 oxide TiO_2 . We find that key features in the electronic structure of the precursor oxides, such as the valence band being bounded by bonding and antibonding Sn 5s, Pb 6s, or Bi 6s states hybridized with O 2p states in SnO, PbO, and Bi_2O_3 , and the conduction band being primarily defined by Ti 3d states in TiO_2 , are preserved in the ternary oxides. Importantly we find that the band gap of a ternary oxide with a TiO_2 fraction of x basically falls on the straight line between that of TiO_2 (x = 1) and the lone pair parent oxide (x = 0). A similar trend in the gap between the peaks in the second derivative of the O K XES and 1s XAS of these oxides suggests that this trend is a property of the general electronic structure, and

that O K XES and 1s XAS spectroscopy can be an experimental way of testing the idea that the band gap of a lone pair, d^0 ternary oxide is the stoichiometry-weighted addition of the band gaps of the parent lone pair and d^0 binary oxides. These findings can be used to predict the band gap of new ternary oxides if the band gaps of the parent lone pair and d^0 binary oxides are known.

ACKNOWLEDGMENTS

We acknowledge support of the Natural Sciences and Engineering Research Council of Canada (NSERC), the Canada Research Chair program, and the Russian Foundation for Basic Research (Projects 11-02-00022). The Canadian Light Source is supported by the NSERC, the National Research Council (NSC) Canada, the Canadian Institutes of Health Research (CIHR), the Province of Saskatchewan, Western Economic Diversification Canada, and the University of Saskatchewan. The Advanced Light Source is supported by the Director, Office of Science, Office of Basic Energy Sciences, of the U.S. Department of Energy under Contract No. DE-AC02-05CH11231.

- * john.mcleod@usask.ca
- ¹ R. Asahi, T. Morikawa, T. Ohwaki, K. Aoki, and Y. Taga, Science **293**, 269 (2001).
- ² A. Fujishima and K. Honda, Nature **238**, 37 (1972).
- ³ J. Pascual, J. Camassel, and H. Mathieu, Phys. Rev. Lett. **39**, 1490 (1977).
- ⁴ H. Tang, H. Berger, P. E. Schmid, and F. Lévy, Solid State Commun. **92**, 267 (1994).
- ⁵ R. Sharma, P. P. Das, M. Misra, V. Mahajan, J. P. Bock, S. Trigwell, A. S. Biris, and M. K. Mazumder, Nanotechnology 20, 075704 (2009).
- ⁶ D. Y. C. Leung, X. Fu, C. Wang, M. Ni, M. K. H. Leung, X. Wang, and X. Fu, ChemSusChem 3, 681 (2010).
- ⁷ H. Irie, Y. Watanabe, and K. Hashimoto, J. Phys. Chem B **107**, 5483 (2003).
- ⁸ S. Khan, M. Al-Shahry, and W. B. Ingler Jr., Science **297**, 2243 (2002).
- ⁹ S.-C. Moon, H. Mametsuka, S. Tabata, and E. Suzuki, Catalysis Today **58**, 125 (2000).
- ¹⁰ T.-H. Xu, C.-L. Song, Y. Liu, and G.-R. Han, J. Zhejiang Univ. Sci. B 7, 299 (2006).
- ¹¹ N. Serpone, J. Phys. Chem. B **110**, 24287 (2006).
- ¹² L. Tsetseris, Phys. Rev. B **81**, 165205 (2010).
- ¹³ T. Umebayashi, T. Yamaki, H. Itoh, and K. Asai, J. Phys. Chem. Sol. **63**, 1909 (2002).
- ¹⁴ Y. Yang, X.-J. Li, J.-T. Chen, and L.-Y. Wang, J. Photochemistry and Photobiology A: Chemistry 163, 517 (2004).
- ¹⁵ A. Schleife, F. Fuchs, J. Furthmüller, and F. Bechstedt, Phys. Rev. B 73, 245212 (2006).
- 16 A. Walsh, SPIE Newsroom: Solar & Alternative Energy , 1 (2009).
- ¹⁷ J. Moon, H. Takagi, Y. Fujishiro, and M. Awano, J. Mater. Sci. **36**, 949 (2001).
- ¹⁸ W.-J. Yin, H. Tang, S.-H. Wei, M. M. Al-Jassim, J. Turner, and Y. Yan, Phys. Rev. B 82, 045106 (2010).
- ¹⁹ P. Wang, Z. Liu, F. Lin, G. Zhou, J. Wu, W. Duan, B.-L. Gu, and S. B. Zhang, Phys. Rev. B 82, 193103 (2010).
- ²⁰ N. Kumada, Y. Yonesaki, T. Takei, N. Kinomura, and S. Wada, Mater. Res. Bull. 44, 1298 (2009).
- ²¹ Y. Noguchi, T. Goto, M. Miyamama, A. Hoshikawa, and T. Kamiyama, J. Electroceram. **21**, 49 (2008).
- ²² J. J. Jia, T. A. Callcott, J. Yurkas, A. W. Ellis, and F. J. Himpsel, Rev. Sci. Instr. **66**, 1394 (1995).
- ²³ J. Valjakka, J. Utriainen, T. Åberg, and J. Tulkki, Phys. Rev. B **32**, 6892 (1985).
- ²⁴ T. Regier, J. Krochak, T. K. Sham, Y. F. Hu, J. Thompson, and R. I. R. Blyth, Nulc. Instrum. Meth. A 582, 93 (2007).
- ²⁵ A. M. Glazer and S. A. Mabud, Acta Cryst. B **34**, 1065 (1978).
- While Bi₄Ti₃O₁₂ was recently reported to have B2cb symmetry⁵⁷, there are a few earlier reports that Bi₄Ti₃O₁₂ is best represented by the B1a1 spacegroup at ambient conditions^{58,59}. Since other studies have noted that there is very little difference between the two structures^{60,61}, and ab initio calculations reveal little difference in the energy stability⁶² between the two phases, we feel justified in using the B2cb phase herein. Indeed, there is little difference in the bulk electronic structure of the B1a1 phase and the higher symmetry I4/mmm phase (realized above 700° C)²¹.
- ²⁷ V. Kahlenberg and H. Böhm, Acta Cryst. B **51**, 11 (1995).
- ²⁸ R. J. Swope, J. R. Smyth, and A. C. Larson, Am. Mineral. **80**, 448 (1995).
- ²⁹ X. Wang, F. X. Zhang, I. Loa, K. Syassen, M. Hanfland, and Y.-L. Mathis, phys. stat. sol. (b) **241**, 3168 (2004).
- ³⁰ P. Boher, P. Garnier, J. R. Gavarri, and A. W. Hewat, J. Solid State Chem. **57**, 343 (1985).
- ³¹ G. Malmros, Acta Chem. Scand. **24**, 384 (1970).
- ³² P. Blaha, K. Schwarz, G. K. H. Madsen, D. Kvasnicka, and J. Luitz, WIEN2k, An Augmented Plane Wave + Local Orbitals Program for Calculating Crystal Properties (Karlheinz Schwarz, Techn. Universität Wien, Austria, 2001) ISBN 3-9501031-1-2
- ³³ J. P. Perdew, K. Burke, and M. Ernzerhof, Phys. Rev. Lett. 77, 3865 (1996).
- ³⁴ F. Tran and P. Blaha, Phys. Rev. Lett. **102**, 226401 (2009).
- ³⁵ K. Schwarz, A. Neckel, and J. Nordgren, J. Phys. F: Metal Phys. 9, 2509 (1979).
- ³⁶ K. Momma and F. Izumi, J. Appl. Crystallogr. **41**, 653 (2008).
- ⁷ J. D. Dunitz and L. E. Orgel, J. Phys. Chem. Solids **3**, 20 (1957).
- ³⁸ G. W. Ludwig and H. H. Woodbury, Phys. Rev. Lett. **5**, 98 (1960).
- ³⁹ A. Walsh and G. W. Watson, Phys. Rev. B **70**, 235114 (2004).
- ⁴⁰ M. W. Stoltzfus, P. W. Woodward, R. Seshadri, J.-H. Klepeis, and B. Bursten, Inorg. Chem. 46, 3839 (2007).
- ⁴¹ J. A. McLeod, R. G. Wilks, N. A. Skorikov, L. D. Finkelstein, M. Abu-Samak, E. Z. Kurmaev, and A. Moewes, Phys. Rev. B 81, 245123 (2010).
- ⁴² F. M. F. de Groot, M. Grioni, J. C. Fuggle, J. Ghijsen, G. A. Sawatzky, and H. Petersen, Phys. Rev. B 40, 5715 (1989).
- ⁴³ F. M. F. de Groot, J. Faber, J. J. M. Michiels, M. T. Czyžyk, M. Abbate, and J. C. Fuggle, Phys. Rev. B 48, 2074 (1993).
- ⁴⁴ A. Walsh, D. J. Payne, R. G. Egdell, and G. W. Watson, Chem. Soc. Rev. 40, 4455 (2011).
- ⁴⁵ A. Walsh, G. W. Watson, D. J. Payne, R. G. Edgell, J. Guo, P.-A. Glans, T. Learmonth, and K. E. Smith, Phys. Rev. B 73, 235104 (2006).
- ⁴⁶ U. V. Waghmare, N. A. Spaldin, H. C. Kandpal, and R. Seshadri, Phys. Rev. B 67, 125111 (2003).
- J. A. McLeod, Z. V. Pchelkina, L. D. Finkelstein, E. Z. Kurmaev, R. G. Wilks, A. Moewes, I. V. Solovyev, A. A. Belik, and E. Takayama-Muromachi, Phys. Rev. B 81, 144103 (2010).
- ⁸ E. Z. Kurmaev, R. G. Wilks, A. Moewes, L. D. Finkelstein, S. N. Shamin, and J. Kuneš, Phys. Rev. B 77, 165127 (2008).
- ⁴⁹ C. R. A. Catlow, Z. X. Guo, M. Miskufova, S. A. Shevlin, A. G. H. Smith, A. A. Sokol, A. Walsh, D. J. Wilson, and S. M. Woodley, Phil. Trans. R. Soc. A 368, 3379 (2010).

- ⁵⁰ Y. Ogo, H. Hiramatsu, K. Nomura, H. Yanagi, T. Kamiya, M. Hirano, and H. Hosono, Appl. Phys. Lett. **93**, 032113 (2008).
- ⁵¹ W. L. Warren, J. Robertson, D. Dimos, B. A. Tuttle, G. E. Pike, and D. A. Payne, Phys. Rev. B **53**, 3080 (1996).
- ⁵² R. C. Keezer, D. L. Bowmand, and J. H. Becker, J. Appl. Phys. **39**, 2062 (1968).
- ⁵³ T. Kidchob, L. Malfatti, D. Marongiu, S. Enzo, and P. Innocenzi, J. Am. Ceram. Soc. 93, 2897 (2010).
- ⁵⁴ H. Zhang, G. Chen, and X. Li, Solid State Ionics 180, 1599 (2009).
- ⁵⁵ H. Gobrecht, S. Seeck, H.-E. Bergt, A. Märtens, and K. Kossmann, phys. stat. sol. **33**, 599 (1969).
- P. Dufek, P. Blaha, and K. Schwarz, Phys. Rev. B 50, 7279 (1994).
 K. R. Chakraborty, S. N. Achary, S. J. Patwe, P. S. R. Krishna, A. B. Shinde, and A. K. Tyagi, Ceram. International 33,
- A. D. Rae, J. G. Thompson, R. L. Withers, and A. C. Willis, Acta Cryst. B 46, 474 (1990).
- ⁵⁹ Y.-I. Ki and M. K. Jeon, Mater. Lett. **58**, 1889 (2004).
- ⁶⁰ C. H. Hervoches and P. Lightfoot, Chem. Mater. **11**, 3359 (1999).
- ⁶¹ T. Hashimoto and H. Moriwake, Phys. Rev. B **78**, 092106 (2008).
- ⁶² A. Shrinagar, A. Garg, R. Prasad, and S. Auluck, Acta Cryst. A **64**, 368 (2008).