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# Accurate determination of the scattering length of erbium atoms

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An accurate knowledge of the scattering length is fundamental in ultracold quantum gas experiments and essential for the characterisation of the system as well as for a meaningful comparison to theoretical models. Here, we perform a careful characterisation of the s-wave scattering length  $a_s$ for the four highest-abundance isotopes of erbium, in the magnetic field range from 0 G to 5 G. We report on cross-dimensional thermalization measurements and apply the Enskog equations of change to numerically simulate the thermalization process and to analytically extract an expression for the so-called number of collisions per re-thermalization (NCPR) to obtain  $a_s$  from our experimental data. We benchmark the applied cross-dimensional thermalization technique with the experimentally more demanding lattice modulation spectroscopy and find good agreement for our parameter regime. Our experiments are compatible with a dependence of the NCPR with  $a_s$ , as theoretically expected in the case of strongly dipolar gases. Surprisingly, we experimentally observe a dependency of the NCPR on the density, which might arise due to deviations from an ideal harmonic trapping configuration. Finally, we apply a model for the dependency of the background scattering length with the isotope mass, allowing to estimate the number of bound states of erbium.

#### I. INTRODUCTION

The high degree of environmental isolation and the high control over the large parameter-space of ultracold quantum gases are key for their success [1]. One of the most decisive properties in determining the many-body phases of a quantum gas is the interaction force between atoms. Among neutral particles, it can be isotropic and short-range, as in alkali atoms, and/or anisotropic and long-range. Open-shell lanthanides, such as erbium (Er) and dysprosium (Dy), have both interactions in place [2]. Their strong magnetic character is reflected in a large dipole-dipole interaction (DDI), while the contact potential is governed by the well-known scattering length, whose value  $a_s$ , as in alkali atoms, can be largely controlled by so-called Fano-Feshbach resonances [3–5].

Although the concept of the scattering length itself is well known by now, theoretical challenges to calculate  $a_s$ depend on the atomic species of interest. For lanthanides, predicting  $a_s$  remains a major challenge of quantum chemistry and microscopic scattering theories [6]. The complexity of describing such atoms has several reasons: the multiple valence electrons, the strongly anisotropic orbital shells, the strong coupling between core and valence electrons, and the relativistic contributions, also made important by the large atomic mass. To date, there are still no ab-initio models with the capacity for quantitative predictions, although many general properties of the interaction potentials (e. g. Born-Oppenheimer potentials) have been studied and understood [7].

Yet, knowledge of the scattering length remains of prime importance since it is an essential regulator of fewand many-body quantum phenomena. For instance, the fascinating supersolid state, recently discovered in Dy [8–10] and Er [9], lives in a narrow range of only a few  $a_0$  ( $a_0$  is the Bohr radius), or the functional forms of beyond-mean-field corrections, which are still under discussion [11–14], depend on  $a_s$  in a subtle way. In the absence of complete microscopic and ab-initio potential models, the study of  $a_s$  in lanthanides therefore relies on experimental investigations and empirical models.

Several different experimental methods have been applied in previous works to extract  $a_s$  for Er and Dy. These include spectroscopy of the molecular binding energy close to a broad Fano-Feshbach resonance [15, 16], the anisotropic expansion of a thermal gas [17], and the cross-dimensional thermalization technique [18–21]. Furthermore, for the <sup>166</sup>Er isotope,  $a_s$  has been determined with high accuracy based on a measurement of the particle-hole excitation gap in the Mott insulator regime via lattice modulation spectroscopy [22, 23]. These techniques did not always provide consistent values, opening up a number of fundamental questions, e.g. from the validity of the additivity of the interaction pseudo-potentials [24–27] to the appropriateness of the Lee-Huang-Yang form for beyond-mean field effects [12–14, 28].

In this work, we extensively study the scattering length of the four most abundant bosonic isotopes of erbium  $(^{164}\text{Er}, ^{166}\text{Er}, ^{168}\text{Er}, \text{and } ^{170}\text{Er})$  and its magnetic-field de-

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pendence. For each isotope, we perform high-resolution Fano-Feshbach spectroscopy in the low magnetic-field region (0 to 5 G) and identify previously unreported scattering resonances. In this range, we then accurately determine the erbium scattering length,  $a_s$ , by developing a model based on the Enskog equations to extract  $a_s$  from cross-dimensional-thermalization experiments. We benchmark our results with the ones obtained from high-precision lattice-modulation spectroscopy, which has been previously developed for <sup>166</sup>Er [23, 29] and here expanded to <sup>168</sup>Er. Finally, from the magnetic-field mapping of  $a_s$ , we extract for each isotope an effective background scattering length  $a_s^{bg}$  at zero *B* field and we discuss the results in the context of the isotope-mass scaling.

#### II. CROSS-DIMENSIONAL THERMALIZATION

The cross-dimensional-thermalization technique is a very powerful method to experimentally determine the scattering length. First successfully applied to alkali atoms [30–33], this technique has proved to be very general and, more recently, has been used for more complex atomic species, such as chromium [34], specific isotopes of erbium [19] and dysprosium [21], and molecular systems [35, 36].

Starting from a cold thermal cloud, the basic idea of the cross-dimensional thermalization method is to excite the system by increasing the potential energy along one spatial dimension of the atomic cloud and to measure the characteristic time  $\tau$  that the system needs to re-thermalize in the orthogonal directions [30]. In the regime of small excitations, for an atomic cloud at a temperature T and a total atom number N, the characteristic time is related to the total scattering cross section  $\bar{\sigma}$  by

$$\tau = \frac{\alpha}{\bar{n}\bar{\sigma}v_{\rm r}},\tag{1}$$

where  $\bar{n}$  is the mean number density

$$\bar{n} = \frac{N\bar{\omega}^3}{\sqrt{8}} \left(\frac{m}{2\pi k_B T}\right)^{3/2} \tag{2}$$

and  $v_{\rm r}$  the mean relative velocity for two colliding atoms

$$v_{\rm r} = \sqrt{\frac{16k_BT}{\pi m}} \tag{3}$$

Here,  $\bar{\omega}$  is the geometric mean of the harmonic trapping frequencies, m is the atomic mass, and  $k_B$  is the Boltzmann constant. Because multiple collisions, not all contributing equally to re-thermalization, are occurring during the thermalization process, the parameter  $\alpha$  can be interpreted as a re-scaling of  $\bar{\sigma}$  and therefore as a number of collisions per re-thermalization (NCPR). Experimentally, the knowledge of  $\alpha$  is fundamental for the extraction of the total scattering cross section.

Equation (1) has two unknown parameters:  $a_s$  and  $\alpha$ . In contrast to alkali atoms, where the scattering is

isotropic, the situation is more complex for dipolar atoms such as Er and Dy [18, 20]. Here, the total cross section for bosons is not only given by the contact scattering length  $a_s$ , but an additional contribution from the nonisotropic DDI, which for two atoms at a distance r and polarized by an external magnetic field **B**, reads as

$$V_{\rm dd}(r,\theta) = \frac{\mu_0 \mu^2}{4\pi} \frac{1 - 3\cos^2\theta}{|\mathbf{r}|^3}.$$
 (4)

Here,  $\mu_0$  is the magnetic permeability,  $\mu$  is the magnetic dipole moment, and  $\theta$  the angle between **B** and **r**. Taking an angular average of the total cross section leads to

$$\bar{\sigma} = 8\pi a_s^2 + \frac{32\pi}{45} a_d^2,$$
 (5)

where  $a_{\rm d} = \frac{m\mu_0\mu^2}{8\pi\hbar^2}$  is the dipolar length  $(a_{\rm d} = 98.2 a_0 \text{ for } {}^{166}\text{Er})$ , with  $\hbar$  being the reduced Planck constant. Finally, we can rewrite Eq. (1) as

$$\tau = \frac{\alpha}{\bar{n}\bar{\sigma}v_{\rm r}} = \frac{\alpha}{\frac{4Nm\bar{\omega}}{\pi k_{\rm B}T_0}(a_{\rm s}^2 + \frac{4}{45}a_{\rm d}^2)}.$$
 (6)

The interplay between the isotropic scattering length and the anisotropic dipolar cross section leads to a dependence of  $\alpha$  on both, the dipole orientation  $\theta$  and  $a_{\rm s}$  [37]. In the limit of weak excitation, an analytic form of  $\alpha(a_{\rm s}, \theta)$  can be found based on the Enskog equations; see later discussion.

## III. EXPERIMENTAL PROCEDURE

In our experiment, we produce a spin-polarized thermal cloud of Er atoms in the lowest Zeeman sublevel, similarly to Ref. [38]. In brief, after cooling and trapping the Er atomic ensemble in a narrow-line magneto-optical trap [39], we transfer the atoms into a crossed optical dipole trap (cODT). Here, we first further cool the atoms via standard evaporative cooling, and then tighten the trapping confinement to avoid atom loss due to residual evaporation. Simultaneously, we ramp  $\mathbf{B}$  to the desired value. At this stage we typically reach a temperature of  $T = 250 \,\mathrm{nK} - 300 \,\mathrm{nK}$  with  $N \approx 1 \times 10^5$ . The exact numbers depend on the isotope choice and the individual set of measurements. The typical final trap frequencies are  $(\omega_x, \omega_y, \omega_z) = 2\pi \times (65(1), 19(1), 300(2))$  Hz. For all sets of measurements the critical temperature for the onset of Bose-Einstein condensation  $T_{\rm c}$  lies between 150 nK and 200 nK, such that  $T \gtrsim 1.5 \times T_c$ . The orientation of the magnetic dipoles is controlled by the direction of the polarizing **B** and is represented by the angle  $\theta$  between **B** and the vertical direction z, defined by gravity; see inset Fig. 1. We calibrate the value of B with an accuracy of about 1 mG by driving the radio-frequency transition between the two lowest Zeeman sublevels  $m_J = -6$  to  $m_J = -5.$ 

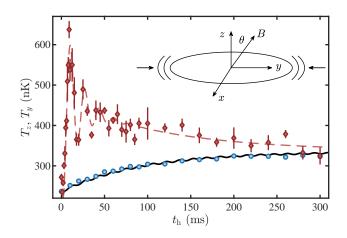


FIG. 1. Effective temperatures  $T_z$  (blue circles) and  $T_y$  (red diamonds) after the increase of the trapping potential along the weakest trapping direction y. The measurement was performed at 1 G and  $\theta = 0^{\circ}$  for the <sup>166</sup>Er isotope. The red dashed line represents a guide to the eye. The black solid line denotes the results of the Enskog simulations for this specific dataset. The errorbars denote the standard error for 3 repetitions. The inset shows a schematic representation of our experimental system.

After preparing the thermal sample, we perform crossdimensional thermalization experiments [19]. In particular, we excite the cloud along the y direction and probe the thermalization dynamics in the z direction. Our excitation scheme relies on a rapid increase in power of one trapping beam, leading to a 60 % increase of the trapping frequency, while leaving the other two directions mostly unaffected. We extract the effective temperature  $T_z$  ( $T_y$ ) for a variable in-trap hold time  $t_h$  from the width of the momentum distribution  $\sigma_z(t_h)$  ( $\sigma_y(t_h)$ ) after a time of flight of  $t_{\text{ToF}} = 25 \text{ ms}$  (20 ms). This scheme, illustrated in the inset of Fig. 1, leads to an out-of-equilibrium cloud with an effective temperature increase along y from about 300 nK to 600 nK.

Figure 1 shows  $T_z$  and  $T_y$  as a function of  $t_h$  at B = 1 G. As we excite the system along y, we observe the expected rapid increase of  $T_y$ . After reaching a maximum effective temperature,  $T_y$  starts to decay, and simultaneously  $T_z$  increases, both reaching the same equilibrium temperature, thus showing thermalization dynamics. We observe oscillations in  $T_y$ , which we attribute to a breathing mode that gets induced by the excitation. For  $T_z$  we observe an exponential-type growth of the form

$$T_z(t) = T_f(1 - \Delta T e^{-t/\tau}). \tag{7}$$

Here,  $T_{\rm f}$  denotes the final temperature and  $\Delta T$  denotes the temperature increase due to the added energy. However, using this simple fit we can not directly extract  $a_{\rm s}$  as additional knowledge on  $\alpha(a_{\rm s}, \theta)$  is needed (see Eq. (6)).

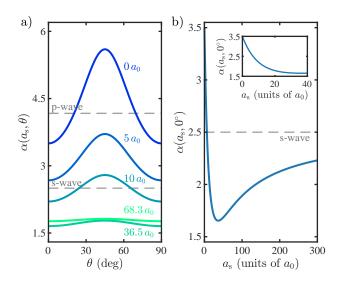


FIG. 2. (a) Dependency of  $\alpha$  on  $\theta$  and  $a_s$  for  $a_s = 0 a_0, 5 a_0, 10 a_0, 36.5 a_0, and 68.3 a_0$ . These values are chosen such that the angle dependence at small  $a_s$  becomes visible. Note that, at 68.3  $a_0$  ( $a_s$  at 1 G, see later measurements) the variation of  $\alpha$  with  $\theta$  is strongly suppressed. (b)  $\alpha$  vs.  $a_s$  for  $\theta = 0^\circ$ . The inset shows an enlargement of the region for  $a_s$  between  $0 a_0$  and  $40 a_0$ . The grey dashed lines show the values of  $\alpha$  for s-wave and p-wave scattering, respectively.

#### IV. THEORETICAL ESTIMATE OF $\alpha(a_s, \theta)$

To compute  $\alpha(a_{\rm s}, \theta)$ , we utilize the Enskog equations of change [40]: a coupled set of differential equations derived in closed-form for dipolar gases, by linearization of the Boltzmann equation, and the assertion of a Gaussian phase-space distribution [41]. These equations permit an analytic derivation of  $\alpha(a_{\rm s}, \theta)$  in the limit of shorttimes and small excitations [37]. For the current experiment with excitation along y and thermalization measured along z, the NCPR is described by a simple analytic formula, which reads

$$\alpha(a_{\rm s},\theta) = \frac{14\left(45a_{\rm s}^2 + 4a_{\rm d}^2\right)}{252a_{\rm s}^2 + 96a_{\rm s}a_{\rm d} + (3\cos(4\theta) + 13)a_{\rm d}^2}.$$
 (8)

The quantity  $\alpha(a_{\rm s}, \theta)$  exhibits an anisotropic character via its angle dependence, as already observed for dipolar fermionic atoms [19] and molecules [36].

Figure 2 shows  $\alpha(a_s, \theta)$  as a function of  $\theta$  (a) and  $a_s$  (b), for our experimental configuration of a pancake shaped trap. Figure 2(a) shows that the anisotropic character of  $\alpha(a_s, \theta)$  competes with the contact one. Indeed, while for small  $a_s$  ( $\leq 10 a_0$ ),  $\alpha(a_s, \theta)$  exhibits a pronounced angledependence with a maximum at 45°, for increasing  $a_s$ such behavior progressively washes out. For  $a_s \approx 70 a_0$ , the thermalization behavior becomes basically independent of  $\theta$ , however  $\alpha(a_s, \theta)$  acquires a number below the one expected for purely contact interacting s-wave collisions. This suggests faster thermalization for dipolar particles, arising from a more efficient diversion of veloc-

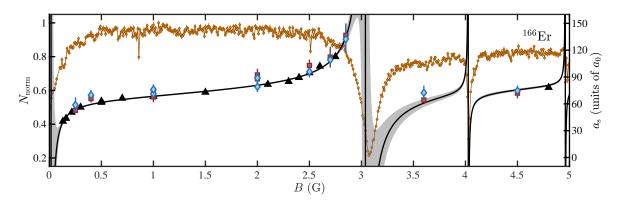


FIG. 3. Atom-loss spectroscopy (orange circles) as a function of B for a fixed holding time of 250 ms. For each B value, the data point is an average of 3 repetitions and it is normalized to the maximum averaged atom number recorded in the explored magnetic field range. Further,  $a_s$  extracted from cross-dimensional thermalization measurements using both, the Enskog equations (red squares) and the analytic formula of Eq. (8) (blue diamonds) is shown for <sup>166</sup>Er. Additionally  $a_s^{\text{LMS}}$  (black triangles) obtained from lattice modulation spectroscopy measurements are given. The solid black lines represent a fit of Eq. (10) to  $a_s^{\text{LMS}}$ . Error bars and the shaded area of the fitting results denote the standard error.

ities of the scattering constituents. In the experiment, we only measure re-thermalization for relatively large values of  $a_{\rm s} \gtrsim 30 a_0$ , and therefore we are not sensitive to the angle dependence of  $\alpha(a_{\rm s}, \theta)$ . In the course of this work, we will thus focus on the case  $\theta = 0^{\circ}$ , simplifying Eq. (8) to

$$\alpha(a_{\rm s}, \theta = 0^{\circ}) = \frac{14\left(45a_{\rm s}^2 + 4a_{\rm d}^2\right)}{252a_{\rm s}^2 + 96a_{\rm s}a_{\rm d} + 16a_{\rm d}^2}.$$
 (9)

As shown in Figure 2(b), after an initial decrease,  $\alpha(a_{\rm s}, 0^{\circ})$  increases for  $a_{\rm s} \gtrsim 36.7 a_0$  – and thus the thermalization loses efficiency – moving to the regime of contact dominated interaction, eventually reaching the  $\alpha(a_{\rm s}, 0^{\circ}) = 2.5$  limit of non-magnetic atoms [18, 42, 43]. We note that by setting  $\theta = 0^{\circ}$  and  $a_{\rm d}/a_{\rm s} \approx 2.7$ , the NCPR is minimized with value  $\alpha \approx 1.65$ , indicating highly efficient collisional thermalization. This is directly attributed to the innate anisotropic differential crosssection in dipolar bosons [18].

# V. MAPPING OF $a_s$ AS A FUNCTION OF B FOR <sup>166</sup>Er

Before taking cross-dimensional thermalization measurements for  $^{166}$ Er, we perform a high resolution scan of the atom number as a function of the magnetic field in order to record the spectrum of Fano-Feshbach resonances, which we know to be exceptionally dense [4, 5]. We record the Fano-Feshbach spectra in a magnetic field region from 0 G to 5 G; see Fig. 3 and Appendix D. In all the measurements the magnetic field is oriented along z.

We then perform thermalization measurements at values of the magnetic field, where the system is not dominated by resonant atom loss. For each thermalization curve, we extract  $a_s$  using two different approaches, one numerical and one semi-analytical. The first, constitutes a direct fit of the full Enskog solutions to the experimental data, leaving  $a_s$  as a float parameter of the theory; see Appendix B for more details. The second method, is based on the exponential growth rate  $\tau$ , from Eq. (1) using the analytic expressin for  $\alpha(a_s, 0^\circ)$  in Eq. (9). For the latter, since  $a_s$  is unknown a priori, we use an iterative approach to determine  $\alpha(a_s, 0^\circ)$  starting from  $\alpha(a_s, 0^\circ) = 1.7$ . We use the calculated  $a_s$  and the analytic formula (see Eq. (9)) to obtain a new value for  $\alpha(a_s, 0^\circ)$ . We stop the iteration once the relative change of  $\alpha(a_s, 0^\circ)$ is  $\leq 1 \times 10^{-7}$ .

Figure 3 summarizes  $a_s$  for <sup>166</sup>Er in the region from 0 G to 5 G. In the studied *B*-field regime, the scattering behavior is essentially dominated by a broad resonance at 3 G and a second one around B = 0 G. The  $a_s$  extracted from the Enskog model and the semi-analytic one are in very good agreement with each other, reflecting the strength of the analytic formula of Eq. (9).

#### VI. BENCHMARKING WITH LATTICE SPECTROSCOPY

To evaluate the robustness of our approach to extract  $a_s$ , we benchmark our cross-dimensional thermalization results with the one obtained using an alternative technique based on lattice modulation spectroscopy (LMS). Such a technique, which we have developed in the past for <sup>166</sup>Er [23, 29] and <sup>167</sup>Er [44], is based on the measurement of the on-site interaction - related to  $a_s$  - of a lattice-confined dipolar gas in a Mott insulator state. The LMS is able to provide accurate values of  $a_s^{\text{LMS}}$ , but at the price of being experimentally more involved due to its requirements of an optical lattice together with a highly degenerate sample. Here we compare the values of  $a_s$  obtained with cross-dimensional thermalization on a low-density thermal sample, with  $a_s^{\text{LMS}}$  obtained from

the lattice modulation spectroscopy obtained in Ref. [29]. In brief we extract  $a_s^{\text{LMS}}$  as follows. We prepare an ultracold sample of <sup>166</sup>Er atoms in a three-dimensional optical lattice, created by two retro-reflected laser beams at 532 nm in the horizontal plane and by one retro-reflected laser beam at 1064 nm along the vertical z direction, defined by gravity. The final lattice depth along the three directions is  $(s_x, s_y, s_z) = (20, 20, 100)$ , in units of  $E_{\text{rec}} = 4.2 \text{ kHz} (1.05 \text{ kHz})$  for 532 nm (1064 nm). The uncertainty on  $s_x$ ,  $s_y$ , and  $s_z$  is about 5%. In such a deep lattice, the atoms are in the Mott insulator phase [23].

We then create particle-hole-excitations by sinusoidally modulating the power of the horizontal lattice beams for 90 ms with a peak-to-peak amplitude of about 30% and measure the recovered BEC fraction after melting of the lattice. At the resonance condition, where the modulation frequency matches the particle-hole excitation gap, we observe a resonant reduction in the BEC fraction [45]. The particle-hole excitation gap is directly given by the on-site interaction  $U = U_c + U_{dd}$ . Here,  $U_c$  is the contact interaction – and thus depends on the unknown  $a_s^{\text{LMS}}$ – while the on-site dipolar interaction,  $U_{dd}$ , can be accurately calculated. We repeat the measurements at various magnetic-field values and, for each, we extract  $a_s^{\text{LMS}}$ .

In Fig. 3, we compare  $a_{\rm s}^{\rm LMS}$  with  $a_{\rm s}$  extracted from the thermalization measurements. We see an overall very good agreement between the value of  $a_{\rm s}$  extracted using the two techniques. This shows that the crossdimensional thermalization approach combined with the Enskog equations is a very reliable method to extract  $a_{\rm s}$ , even in the case of complex atoms for which the knowledge of  $\alpha(a_{\rm s}, \theta)$  is not a priori given.

# VII. DENSITY DEPENDENCE

Our measurements for the <sup>166</sup>Er isotope were performed in a regime of relatively low density ( $\bar{n}$  <  $0.5 \times 10^{13} \,\mathrm{cm}^{-3}$ ). Interestingly, when applying the same method in a regime of high density, we observe a dependence of the thermalization rate on the density which goes beyond the Enskog approach. For instance, we repeat the cross-dimensional thermalization measurements for <sup>166</sup>Er at B = 1 G and variable cloud density,  $\bar{n}$ . We control the density by either increasing N or by applying a tighter trapping configuration of  $(\omega_x^{\rm cyl},\omega_y^{\rm cyl},\omega_z^{\rm cyl})\approx$  $2\pi \times (300, 19, 300)$  Hz before compression, or both. From the lattice modulation spectroscopy, we have extracted a value  $a_s = 68.3(7) a_0$  at B = 1 G. By fixing this value - meaning to impose that the scattering length does not depend on density - and using Eq. (1), we can determine  $\alpha(a_{\rm s},0^{\circ})$  as a function of  $\bar{n}$ . Note that for all measurements we find a Knudsen number, given by the ratio of the mean free path and the size of the atomic cloud, > 10. This implies that we are far away from the hydrodynamic regime, which could otherwise lead to modifications in the thermalization behavior [46, 47].

Figure 4 shows  $\alpha(a_s, 0^\circ)$  for different values of  $\bar{n}$ . We find a pronounced dependency on  $\bar{n}$ , with a rapid increase and an eventual saturation at high densities. Such a behavior is not captured by our theoretical model, which, as reflected in the definition of  $\alpha(a_s, 0^\circ)$  in Eq. (1), predicts no density dependence. To the best of our knowledge, such a dependence has not been reported in previous works on cross-dimensional thermalization. Possible explanations root in various causes, either physical or technical nature. Although being above  $T_{\rm c}$ , precursors of quantum many-body phenomena might influence the scattering behavior. Exemplary, we tried to explicitly include effects coming from Bose-enhancement into our theoretical framework. This did not have significant influence on the thermalization behavior. Note that, in the experiment, we varied the initial temperature of the atomic cloud and the excitation strength which did not show any influence on the observations.

Another possible explanation, based on unavoidable experimental imperfections, roots in deviations from an ideal harmonic trapping condition, leading to a modification of the kinetic energy and the mean density. Such a variation would manifest in an apparent change of  $\alpha(a_{\rm s},0^{\circ})$ ; see Eq. (1). Indeed, Eq. (2) and (3) are only valid for an ideal harmonic trapping confinement. Furthermore, trap anharmonicities lead to a larger kinetic temperature after thermalization compared to the case of purely harmonic traps as seen when formulating the dynamics using the scaling ansatz method [48]. As this effect is stronger for larger densities, this further suggests that the range of low densities is the appropriate one to consider. First Monte-Carlo simulations performed by using a realistic gaussian trapping potential seem to support this assumption; see Appendix C.

We emphasize that, due to the agreement with the lattice modulation spectroscopy results, the above discussion on anharmonicities, and since our measurements to extract  $a_s$  have been performed at low densities, we are confident that our method remains valid.

# VIII. SCATTERING LENGTH FOR ${}^{164}$ Er AND ${}^{170}$ Er

After the detailed study on <sup>166</sup>Er and the benchmarking of the results with high-precision lattice modulation spectroscopy, we confidently apply our cross-dimensional thermalization approach to two other isotopes, <sup>164</sup>Er and <sup>170</sup>Er. Again we start with a Fano-Feshbach spectroscopy between 0 G and 5G to identify the position of the scattering resonances as shown in Fig. 5. We note that this Fano-Feshbach spectra have not been reported previously. For the cross-dimensional thermalization measurements we follow a similar experimental procedure as described above. From the thermalization curve, we again use both, the full fit of the Enskog equations as well as the iterative approach on  $\alpha(a_{\rm s}, 0^{\circ})$  to determine  $a_{\rm s}$  from the exponential growth rate  $\tau$ .

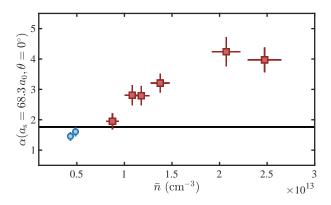


FIG. 4. Measurements of  $\alpha(a_s, \theta)$  as a function of  $\bar{n}$ . The blue circles correspond to the datasets at 1 G, shown in Fig. 3. The black solid line marks the value given by the analytic formula in Eq. (9). All measurements are performed with  $\theta = 0^{\circ}$ . Errorbars denote the standard error.

Figure 5 shows  $a_{\rm s}$  for the isotopes (a)  $^{164}{\rm Er}$  and (b)  $^{170}{\rm Er}$ . While the scattering behavior for  $^{164}{\rm Er}$  is, similarly to  $^{166}{\rm Er}$ , dominated by two broad resonances at 1.5 G and 3.3 G,  $^{170}{\rm Er}$  features several narrow overlapping resonances, providing different test scenarios for our cross-dimensional thermalization. Although minor deviations can be observed in the vicinity of Fano-Feshbach resonances, for both isotopes, the extracted  $a_{\rm s}$  using the two approaches are once more in good agreement.

### IX. SCALING OF BACKGROUND SCATTERING LENGTH WITH MASS

The knowledge on  $a_{\rm s}$  as a function of the magneticfield allows us to extract an effective background scattering length  $a_{\rm s}^{\rm bg}$  for each isotope. The general behavior of  $a_{\rm s}$  with *B* can be described by generalizing the wellknown formula [49]

$$a_{\rm s}(B) = (a_{\rm s}^{\rm bg} + sB) \times \prod_{i=1}^{N_{\rm res}} \left(1 - \frac{\Delta B_i}{B - B_i}\right), \qquad (10)$$

to the case of  $N_{\rm res}$  overlapping resonances of position  $B_i$ and width  $\Delta B_i$ , and allowing for a smooth off-resonant variation of  $a_{\rm s}$  with B. We observe that a linear variation of slope s already well reproduces the data with  $a_{\rm s}(0)$  defined as the effective  $a_{\rm s}^{\rm bg}$ . We note that different mechanisms could lead to an off-resonant variation of  $a_{\rm s}$ . For instance, the influence of broad Fano-Feshbach resonances, which are not within our measurement range, could lead to a smooth variation of the background behavior, similar to that observed for cesium [50]. Alternatively, the effect could be due to the coupling induced by DDI between the incident scattering channel and Zeeman states that lie higher in energy. As a consequence this results in a perturbation of the molecular potential, whose strength depends on the magnetic field, leading to an increasing value of the van der Waals  $C_6$  coefficient [51].

To parametrize  $a_s$  as a function of B, we fit Eq. (10) to the measured  $a_s$  for <sup>164</sup>Er, <sup>166</sup>Er, <sup>168</sup>Er, and <sup>170</sup>Er. For <sup>166</sup>Er and <sup>168</sup>Er, we use the scattering lengths obtained from the lattice modulation spectroscopy, corresponding to our most accurate determination; see solid lines in Fig. 3 and Fig. 3. For <sup>164</sup>Er and <sup>170</sup>Er, we fit Eq. (10) to the  $a_s$  data obtained by applying the Enskog equations to the cross-dimensional thermalization measurements; see solid lines in Fig. 5. More details on the fitting procedure as well as the complete list of the fit parameters is given in Appendix F. In general, we observe that the fitting function reproduces very well the behavior of  $a_s$  for every isotope.

Figure 6 shows the value of  $a_{\rm s}^{\rm bg}$  from the fit as a function of the isotope mass. We observe a monotonic rising of  $a_{\rm s}^{\rm bg}$  with increasing m, which might be compatible with different functional forms, including a simple linear increase. Under the assumption that erbium has a similar behavior to ytterbium and cesium, we can use the model for the mass scaling as developed in Ref. [52–54]. Such a model assumes that  $a_{\rm s}$  is only given by the Van der Waals potential  $U(r) = -C_6/r^6$ , with  $C_6$  being the Van der Waals coefficient. This might be a rather severe approximation for magnetic atoms but, in absence of alternative models, it is interesting to compare the simple mass-scaling approach to erbium.

As introduced in Ref. [52],  $a_{\rm s}$  can be written as

$$a_{\rm s} = \bar{a} \left[ 1 - \tan\left(\phi - \frac{\pi}{8}\right) \right],\tag{11}$$

with  $\bar{a} = 2^{-3/2} \frac{\Gamma(3/4)}{\Gamma(5/4)} \left(\frac{mC_6}{\hbar^2}\right)^{1/4}$  being the characteristic length and

$$\phi = \frac{\sqrt{m}}{\hbar} \int_{R_0}^{\infty} \sqrt{-U(r)} dr.$$
 (12)

Here,  $\Gamma(x)$  is the gamma-function and  $R_0$  is the classical turning point of U(r). Although the exact shape of U(r) is unknown, Eq. (11) can be employed to extract a mass-scaling due to the dependence of  $\phi \propto \sqrt{m}$  [53]. Such a scaling is valid, as long as the mass dependent modification of U(r) is negligible. Furthermore,  $\phi$  allows for the calculation of the number of bound states  $N_{\rm B}$  via relation  $N_{\rm B} = \lfloor \phi/\pi - 5/8 \rfloor$ , where  $\lfloor \ \rfloor$  denotes the floor integer function.

We now apply this model to our Er case. Figure 6 shows the fit of Eq. (11) to the experimental data; see Appendix G for details. We obtain the best agreement for  $\phi/\pi = 144(1)$ , leading to  $N_{\rm B} = 143(1)$  for <sup>168</sup>Er. Despite the similar  $C_6$  coefficient,  $N_B$  is approximately a factor of 2 larger than for ytterbium [53]. Note that,  $N_{\rm B}$  is in agreement with the result obtained when using the same approach but assuming a hard core potential; see Appendix H. We would like to emphasize once more that this model does not consider any contribution arising from the DDI. An improved description calls for the development of advanced theoretical models.

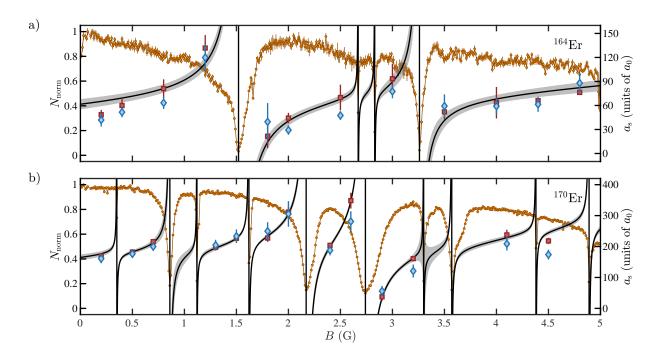


FIG. 5. Atom-loss spectroscopy (orange circles) as a function of B for a fixed holding time of (a) 250 ms for <sup>164</sup>Er and (b) 500 ms for <sup>170</sup>Er. For each B value, the data point is an average of 4-5 repetitions and it is normalized to the maximum averaged atom number recorded in the explored magnetic field range.

Further,  $a_s$  extracted from cross-dimensional thermalization measurements using both, the Enskog equations (red squares) and the analytic formula of Eq. (8) (blue diamonds) is shown. The solid black lines represent a fit of Eq. (10) to  $a_s$  obtained using the Enskog equations. Error bars and the shaded area of the fitting results denote the standard error.

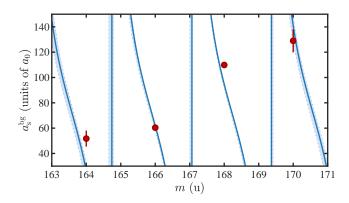


FIG. 6. Background scattering length  $a_s^{\text{bg}}$  for four bosonic isotopes (red circles). The solid line represents the best fit with  $\phi/\pi = 144(1)$ ; see text. The shaded area, enclosed by the dotted lines, represents the fitting function for  $\phi = 143$ and  $\phi = 145$ . The errorbars denote the standard error of the fit of Eq. (10) to the experimental data.

#### X. CONCLUSION

In conclusion, we report on an accurate study of the scattering length of four different isotopes of erbium. Our work focuses on the low magnetic field region, which is the range of most interest in current experiments. Our experimental survey combines two different techniques: a high-precision, yet demanding, approach based on the measurement of the onsite interaction in a Mott insulator phase, and another one based on measuring the re-equilibration time in cross-dimensional thermalization experiments. From the latter, we extract the the value of  $a_s$  by both numerically applying the full Enskog equations and using the analytic formulation for  $\alpha(a_s, \theta)$ . All these different approaches, benchmarked one with respect to the others, provide a very consistent measure of the scattering length in the region of interest. These results will be relevant for current experiments and moreover point to a practical manner to extract  $a_s$  with reduced experimental effort, which can be readily generalized to other magnetic lanthanides.

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#### Appendix A: Analytic number of collisions per re-thermalization

Analytic expressions for  $\alpha(a_s, \theta)$  can be derived under a short-time approximation, with the Enskog equations

$$\frac{d\langle q_j^2 \rangle}{dt} - \frac{2}{m} \langle q_j p_j \rangle = 0, \qquad (A1a)$$

$$\frac{d\langle p_j^2 \rangle}{dt} + 2m\omega_j^2 \langle q_j p_j \rangle = \mathcal{C}[p_j^2], \qquad (A1b)$$

$$\frac{d\langle q_j p_j \rangle}{dt} - \frac{1}{m} \left\langle p_j^2 \right\rangle + m \omega_j^2 \langle q_j^2 \rangle = 0, \qquad (A1c)$$

where  $r_j$  and  $p_j$  are positions and momenta respectively (j = x, y, z), and C is the collision integral. The derivation follows from Ref. [37], but we present a brief outline here for completeness. The gas is assumed close-to-equilibrium, allowing us to treat  $r_j$  and  $p_j$  as Gaussian distributed. Thermalization trajectories are then tracked using the Gaussian widths along each axis, to compute the energy differential

$$\langle \chi_j \rangle \equiv E_j - k_B T_f,$$
 (A2)

where  $T_f = (T_x + T_y + T_z)/3$  is the final equilibration temperature (obtained from the equipartition theorem),  $\langle \ldots \rangle$  denotes an ensemble average assuming a Gaussian phase space distribution whose widths are allowed to vary, and  $E_j = \langle p_j^2 \rangle/(2m) + m\omega_j^2 \langle r_j^2 \rangle/2$  is the sum of kinetic and potential energies in the *j*-th direction. The Enskog equations dictate that the relaxation of  $\langle \chi_j \rangle$  follows the differential equation

$$\frac{d\langle\chi_j\rangle}{dt} = \mathcal{C}[\chi_j]. \tag{A3}$$

For small deviations from equilibrium and at short times, re-thermalization can be approximated with a single decay rate  $\gamma$ , such that  $\mathcal{C}[\chi] \approx -\gamma \langle \chi \rangle$ . This results in the relation

$$\frac{dE_j}{dt} = -\gamma_{yj} \left( E_j - k_B T_f \right) = \mathcal{C}[E_j], \qquad (A4)$$

where the subscript on  $\gamma_{yj}$  indicates that the gas was excited along y, and re-thermalization measured along j. This then permits us to compute

$$\alpha_{yj} = \frac{\bar{n}\bar{\sigma}v_r}{\gamma_{yj}} = \left(\frac{E_j - k_B T_f}{\mathcal{C}[E_j]}\right)\bar{n}\bar{\sigma}v_r,\qquad(A5)$$

which for j = z, has the form in Eq. (8).

## Appendix B: Fitting Enskog equations to experimental data

The extraction of the scattering lengths  $a_s$ , from crossdimensional thermalization data was done here by means of full numerical solutions to the Enskog equations. To

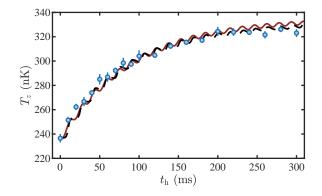


FIG. 7. Benchmarking of the Enskog simulation results for  $T_z$  (red solid line) with Monte-Carlo simulations (black dashed line). The dataset is the same as in Fig. 1.

do so,  $a_s$  was left as a float parameter in the theory, then varied until a best fit between the theory and experimental data was obtained. A feature we noticed during fitting was the high sensitivity of thermalization rates to variations in the trapping frequencies  $\boldsymbol{\omega}$ , over the finite-time quench. Measurement uncertainties therefore motivate us to also leave  $\boldsymbol{\omega}$  a float parameter, with allowed values within its 1-sigma errorbars. This is applied both to the trapping frequencies before and after the quench.

We performed fits using a  $\chi^2$  optimization criterion

$$\min_{\boldsymbol{\omega}, a_s} \sum_{t=t_0}^{t_{\text{end}}} \left( \frac{T(t) - \mathcal{T}_{\text{E}}\left[T(0); \boldsymbol{\omega}, a_s\right](t)}{\delta T(t)} \right)^2, \quad (B1)$$

where the sum runs over measurement time instances t, T(t) is the temperature data from the experiment,  $\delta T(t)$  is the temperature measurement uncertainty, and  $\mathcal{T}_{\rm E}[T(0); \boldsymbol{\omega}, a_s]$  is the solution to the Enskog equations with initial condition T(0), and fit parameters  $a_s$  and  $\boldsymbol{\omega}$ .

To reduce biasing of the fits, we run an iterative algorithm that recursively fits  $\boldsymbol{\omega}$  and  $a_s$  in succession until they converge to stable values. Such a procedure would take exceedingly long times (~ weeks) with full Monte Carlo (MC) simulations, but can be done in minutes with the Enskog equations on a current-day computing device.

Solutions to the Enskog equations have shown themselves accurate when compared to MC simulations [37, 41]. We show their accuracy here yet again, using the parameters from the current experimental set-up. An illustrative example is provided in the plot of Fig. 7, comparing an instance of the Enskog solutions (red solid line), MC simulations (black dashed line) and the experimental data (blue circles).

## Appendix C: Monte-Carlo simulations including trap anharmonicities

Optical dipole traps are, in many studies, assumed to be well modeled by purely harmonic potentials. This may however be inadequate in regimes with significant trap anharmonicity effects, which we currently attribute the density dependence of  $\alpha$  to. In such cases, the potential is better modeled as two cross-propagating Gaussianprofile beams along the y and z axes (with gravity). This produces the confinement potential

$$V_{\text{ODT}}(\boldsymbol{r}) = -\frac{2\tilde{U}_1 P_1}{\pi w_{1,x}(z) w_{1,y}(z)} e^{-2\left(\frac{x^2}{w_{1,x}^2(z)} + \frac{y^2}{w_{1,y}^2(z)}\right)} - \frac{2\tilde{U}_2 P_2}{\pi w_{2,x}(y) w_{2,y}(y)} e^{-2\left(\frac{x^2}{w_{2,x}^2(y)} + \frac{z^2}{w_{2,z}^2(y)}\right)} + mgz, \tag{C1}$$

where P is the laser power U is an atomic polarizability parameter and

$$w(z) = w_0 \sqrt{1 + \frac{z^2}{z_R^2}},$$
 (C2)

with  $z_R$  and  $w_0$  denoting Rayleigh lengths and beam widths respectively.

Such a potential limits the applicability of the aforementioned Enskog equations as formulated in Ref. [41]. Instead, more robust MD methods are required to accurately predict thermalization trajectories. We implement a MD simulation similar to that in Ref. [20], which evolves simulation particles under the action of  $V_{\text{ODT}}$  via the Verlet symplectic integrator

$$\boldsymbol{q}_{k} = \boldsymbol{r}_{k}(t) + \frac{\Delta t}{2m} \boldsymbol{p}_{k}(t), \qquad (C3a)$$

$$\boldsymbol{p}_k(t + \Delta t) = \boldsymbol{p}_k(t) + \boldsymbol{F}_k \Delta t,$$
 (C3b)

$$\boldsymbol{r}_{k}(t+\Delta t) = \boldsymbol{q}_{k} + \frac{\Delta t}{2m}\boldsymbol{p}_{k}(t+\Delta t),$$
 (C3c)

where subscripts k denote the k-th simulation particle,  $\Delta t$  is the simulation time-step, t is the time and

$$\boldsymbol{F}_k = -\nabla V_{\text{ODT}}(\boldsymbol{r}_k). \tag{C4}$$

Dipolar collisions are then computed with the direct simulation Monte Carlo method [56], that determines postcollision momenta via stochastic sampling of the differential cross section.

In a preliminary study of the density dependence, ideal Gaussian beam profiles are assumed, along with perfectly accurate beam widths and Rayleigh lengths. Following a trap quench, thermalization of the out-of-equilibrium gas in  $V_{\rm ODT}$  indeed shows an apparent increase of  $\alpha$  with density, qualitatively similar to that observed in the experiment. This effect is absent in simulations with an ideal harmonic trap. Furthermore, in higher density regimes, the simulations with  $V_{\rm ODT}$  predict the experimentally observed equilibration temperatures more accurately compared to the harmonic trap case. These early findings on density dependence from trap anharmonicities are intriguing, and a cautionary tale for future experiments. However, we do not develop this idea further here and leave such analysis for future works.

#### Appendix D: Fano-Feshbach spectroscopy

To identify the positions of the Fano-Feshbach resonances we perform high-resolution loss spectroscopy in a cylindrically symmetric trap. We evaporatively cool the atoms until they reach a temperature between T =300 nK and 400 nK. At this stage, the atom number is between  $6 \times 10^4$  and  $1.2 \times 10^4$  with typical trap frequencies of  $(\omega_x, \omega_y, \omega_z) = 2\pi \times (300, 30, 300)$  Hz. The exact values depend on the isotope choice. After reaching thermal equilibrium, we change B, oriented along the z axis, in 1 ms to the desired value and wait for a holding time between 250 ms and 500 ms. We use different holding times for different datasets to avoid saturation effects of the resonances for higher densities. After the holding time, we measure the atom number using absorption imaging after a time of flight expansion of 25 ms. The results of the loss-spectroscopy measurements are shown in Fig. 3, Fig. 5 and Fig. 8.

#### Appendix E: Scattering length for <sup>168</sup>Er

To obtain  $a_s$  for the <sup>168</sup>Er isotope, we follow a similar approach as for <sup>166</sup>Er. First, we perform loss spectroscopy to identify the position of Fano-Feshbach resonances. We then transfer the atoms into an optical lattice with a depth of  $(s_x, s_y, s_z) = (20, 20, 40)E_{\rm rec}$  and apply the lattice modulation spectroscopy technique to extract  $a_s$ . The lattice modulation spectroscopy follows the same lines as for the <sup>166</sup>Er isotope; see main text. Fig. 8 summarizes the results for <sup>168</sup>Er and shows the Fano-Feshbach spectroscopy result as well as  $a_s$  as a function of B in the magnetic field range from 0 G to 5 G.

# Appendix F: Extracting background scattering length

To obtain a value for  $a_{\rm s}^{\rm bg}$ , we fit Eq. (10) either to  $a_{\rm s}$  obtained from the full Enksog equations (<sup>164</sup>Er and <sup>170</sup>Er) or  $a_{\rm s}^{\rm LMS}$  (<sup>166</sup>Er and <sup>168</sup>Er). Due to the different numbers of Fano-Feshbach resonances compared to the number of available data points for  $a_{\rm s}$ , we slightly vary the fitting approach for the individual isotopes. Depending on the position and the width of the resonance, for some resonances, we fix the position  $B_i$  to the minimum of the loss feature and keep only the width  $\Delta B_i$  as a floating parameter. For the very narrow resonances, which have a negligible influence on the overall scattering behavior, we fix both  $B_i$  and  $\Delta B_i$ .

Table I gives the results for the background scattering lengths  $a_{\rm s}^{\rm bg}$  and the slopes *s* for all four isotopes. Moreover, Tables II–V contain a detailed listing of all Fano-Feshbach resonances and how they are included in the fitting procedure. Note that for <sup>170</sup>Er, we are aware of the existence of a particularly broad resonance at 6.91 G [57], which we include with variable width. When looking

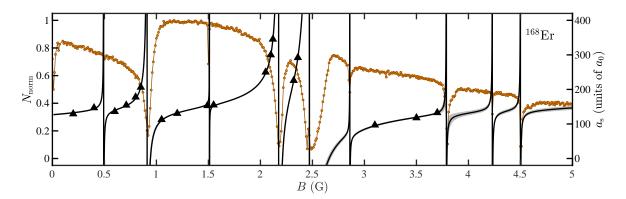


FIG. 8. Atom-loss spectroscopy (orange circles) as a function of B for a fixed holding time of 500 ms. For each B value, the data point is an average of 6-7 repetitions and it is normalized to the maximum averaged atom number recorded in the explored magnetic field range. Further, the measured scattering lengths  $a_s$  obtained for <sup>168</sup>Er from lattice modulation spectroscopy measurements are shown. The solid black line represents a fit to  $a_s^{LMS}$ . The shaded area as well as the error bars denote the standard error.

TABLE I. Values for  $a_{\rm s}^{\rm bg}$  and s obtained from the fit of Eq. (10) to  $a_{\rm s}$  for the four bosonic isotopes. The error denotes the fit error of one standard deviation.

isotope	$a_{ m s}^{ m bg} \; (a_0)$	$s~(a_0/{ m G})$
164	52(6)	9(3)
166	61(3)	5.4(9)
168	110(2)	11(2)
170	129(9)	20(10)

TABLE II. Parameters for the Fano-Feshbach resonances included into the fit of Eq. (10) to  $a_{\rm s}$  for <sup>164</sup>Er. The error denotes the fit error of one standard deviation. Values without error are fixed in the fitting procedure.

Position $B_i$ (G)	Width $\Delta B_i$ (G)	
1.52	0.22(3)	
2.67	0.005	
2.83	0.005	
3.26	0.10(3)	

closely, the onset of this resonance can actually be seen as a reduction of N towards higher magnetic field values in the loss spectroscopy (see Fig. 5(b)).

## Appendix G: $\chi^2$ analysis for mass scaling

In this section, we describe our analysis of the background  $a_s$  of the 4 Er isotopes (Fig. 6) with Eq. (10). To find the best fitting parameter  $\phi$ , we analyze the agreement of the theoretical model in Eq. (11) with our experimental data. For each value of  $\phi$ , we calculate the  $\chi^2$ 

TABLE III. Parameters for the Fano-Feshbach resonances included into the fit of Eq. (10) to  $a_{\rm s}$  for <sup>166</sup>Er. The error denotes the fit error of one standard deviation. Values without error are fixed in the fitting procedure.

Position $B_i$ (G)	Width $\Delta B_i$ (G)	
0.02(5)	0.05(2)	
3.04(5)	0.15(2)	
4.208	0.01	
4.96	0.005	

TABLE IV. Parameters for the Fano-Feshbach resonances included into the fit of Eq. (10) to  $a_{\rm s}$  for <sup>168</sup>Er. The error denotes the fit error of one standard deviation. Values without error are fixed in the fitting procedure.

Position $B_i$ (G)	Width $\Delta B_i$ (G)
0.49	0.005
0.911(6)	0.032(2)
1.51	0.01
2.174(4)	0.038(2)
2.471(9)	0.19(1)
2.86	0.005
3.79	0.006(5)
4.23	0.005
4.5	0.005

via

$$\chi^{2} = \sum_{i=1}^{4} \left( \frac{a_{s}^{\text{mod}} - a_{s}^{i}}{\sigma_{s}^{i}} \right)^{2}.$$
 (G1)

Here,  $a_s^{\text{mod}}$  is the scattering length given by the model for the corresponding  $\phi$  and  $a_s^i$  and  $\sigma_s^i$  are the measured  $a_s$  with the corresponding standard error.

TABLE V. Parameters for the Fano-Feshbach resonances included into the fit of Eq. (10) to  $a_{\rm s}$  for <sup>170</sup>Er. The error denotes the fit error of one standard deviation. Values without error are fixed in the fitting procedure.

Position $B_i$ (G)	Width $\Delta B_i$ (G)
0.35	0.005
0.86	0.028(12)
1.12	0.005
1.62	0.01
2.17	0.067(7)
2.74	0.134(9)
3.3	0.01(1)
3.57	0.01
4.38	0.005
4.49	0.01
6.91	0.8(7)

The behavior of  $\chi^2$  is non-monotonic with the appearance of several minima. We identify the absolute minimum of  $\chi^2$  for  $\phi = 144.03$ . To further obtain an estimate for the error of  $\phi$  we fit a quadratic function to the local minima. We extract the limits of the confidence interval by considering the region where  $\chi^2 \leq \chi^2 + 1$ .

### Appendix H: Hard-core potential for mass scaling

The model contains the assumption, that the s-wave scattering length is given at large distances by the vander-Waals potential scaling with  $U_{\rm vdW}(r) \propto -C_6/r^6$ , with  $C_6$  being the Van der Waals coefficient, and at short distances  $r < r_c$  by a hard core potential [52]. In this specific case, the scaling of  $a_{\rm s}^{\rm bg}$  can be described by

$$a_{\rm s}^{\rm bg} = \bar{a}\tan(\Phi),\tag{H1}$$

where  $\bar{a} = \frac{\Gamma(3/4)}{2\sqrt{(2)}\Gamma(5/4)}a_c$  with  $a_c = \left(\frac{2m_rC_6}{\hbar^2}\right)^{1/4}$  being the characteristic scattering length scale of the potential, and  $\Phi = \frac{a_c^2}{2r_c^2} - \frac{3\pi}{8}$  is the semi-classical phase [52].

From theoretical calculations in Ref. [6] we use  $C_6 = 1723$  a.u. and we estimate from the theoretical interaction potential given in Ref. [6] that  $r_c \approx 4 - 8 a_0$ . We fit Eq. (H1) to  $a_s^{bg}$  of the four bosonic isotopes. Due to a large number of possible local minima, we combine the fitting with a minimization of the  $\chi^2$ -value while varying the start parameter for  $r_c$ . We obtain the best agreement for  $r_c = 5.05(5) a_0$ .

In addition, the Levinson theorem [58] allows us to estimate the number of bound states  $N_B$ , which can be calculated from the semi-classical phase  $\Phi$  using

$$N_B = \left[\frac{\Phi}{\pi} - \frac{3}{8}\right] + 1,\tag{H2}$$

where the square brackets mean the integer part. For the current fitting we obtain  $N_B$  ranging from 141 to 144, in agreement with the approach in the main text. We want to emphasize, that this modelling of  $a_{\rm s}^{\rm bg}$  is a simple approach and a more thorough analysis could add deeper valuable insights.