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Probing Stark-induced nonlinear phase variation with opto-optical modulation

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We extend the recently developed technique of opto-optical modulation (OOM) to probe state-resolved AC-Stark-induced phase variations of a coherently excited ensemble of helium atoms. In a joint experimental and theoretical study, we find that the spatial redirection of the resonant emission from the OOM process is different for the low-lying 1s2p state as compared with the higher-lying Rydberg states, and that this redirection can be controlled through the spatial characteristics of the infrared (IR) probe beam. In particular, we observe that the intensity dependence of the IR-induced Stark phase on the 1s2p emission is nonlinear, and that the phase accumulation changes sign for moderate intensities. Our results suggest that OOM, combined with precise experimental shaping of the probe beam, could allow future measurements of Stark-induced phase shifts of excited states.

I. INTRODUCTION

Light-matter interactions can be addressed from two complementary points of view [1]. Just as light can be used as a tool to probe and control matter [2–7], atoms can be exploited to probe and control light [8–14]. The recently demonstrated technique of opto-optical modulation (OOM) [8, 9] is an example of this duality in the realm of ultrafast extreme ultraviolet (XUV) sources. OOM relies on the combination of two coherent femtosecond pulses with different properties. First, an XUV pump pulse resonantly excites an atomic target producing a coherent superposition of ground and excited states. This triggers a long-lived emission of coherent XUV light at the resonant transition frequencies. Subsequently, a strong, infrared (IR) probe pulse arrives and modifies the XUV emission, altering its spatio-temporal profile.

The effect of the IR probe pulse on the coherent XUV emission is mediated by the AC-Stark shift [15]. This IR-induced shift of the excited state energies yields an additional state-dependent *phase* that is imprinted on the dipole and thus on the emitted XUV light [2, 16, 17]. The OOM technique translates the spatial intensity variation of the IR beam into a state-specific spatial phase gradient that results in the redirection of the XUV emission.

Previously OOM has been used to redirect ultrafast XUV light pulses in an argon gas, from both Rydberg and autoionizing states [8, 9]. Further details of the technique using also helium and neon gases can be found in reference [9]. The direction of emission in these experiments was explained via the known, approximately linear AC-Stark shift of high lying Rydberg states. For these states the AC-Stark shift approaches the average kinetic energy of a free electron oscillating in an electric field, namely the ponderomotive energy $U_p = e^2 F^2 / 4m_e \omega^2$, where e and m_e are the electron charge and mass, and F and ω are the electric field amplitude and angular frequency.

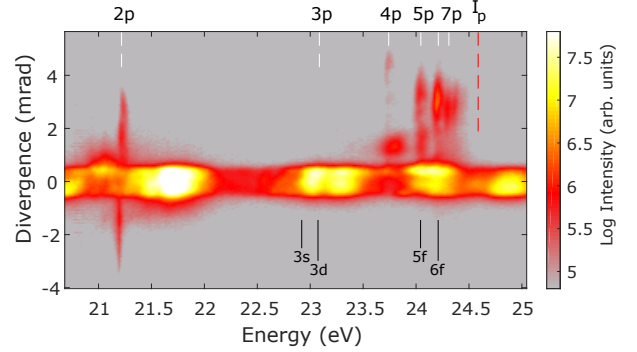


FIG. 1. Example of XUV spatial control using OOM from a manifold of excited np states in helium. The unperturbed 2p–7p energies are indicated in white, and the ionization energy in red. The 2p emission is redirected both up and down by the 800 nm IR pulse, whereas the high-lying np emission is only redirected up. States pertinent to later discussions are shown in black. He state energy levels are taken from [18].

The ponderomotive Stark phase depends linearly on the IR intensity, which acts as a control parameter on the XUV spatio-temporal properties.

In this article we demonstrate that the OOM technique can be used to probe unknown, non-linear Stark phases. In particular, we reveal the intensity dependence of the Stark phase for the low-lying 1s2p state in helium (hereafter we omit the passive 1s occupation label). We coherently excite the manifold of higher energy np Rydberg states as a reference and observe that the spatial redirection of the XUV light from the 2p transition is different relative to the higher-lying np states. Significantly, we find that the 2p energy shift changes sign as a function of intensity, so that if the 2p emission is redirected down at low intensity, it will be redirected up at high intensity. In practice we observe 2p emission in both directions at higher peak intensities, because both high and low intensity regions of the IR beam contribute to

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the redirection, as illustrated in Fig. 1. Solutions of the coupled Maxwell wave equation and the time-dependent Schrödinger equation (MWE-TDSE) reproduce the main features of the experimental results. They allow us to understand the observed 2p state behavior in terms of a transition from a regime of strong near-resonant coupling with nearby states at low intensity, to a regime of non-resonant free-electron-like behavior at high intensity.

II. PRINCIPLE

The principle of OOM [8, 9] and how it may be used to probe the intensity dependence of the Stark phase is illustrated in Fig. 2. A broadband, coherent XUV pump pulse excites a time-dependent dipole moment, which leads to coherent emission in the forward direction at a number of resonant frequencies [19, 20]. The long lifetime of the resonances is reflected as sharp absorption features in the spectral domain. An IR probe pulse following the XUV excitation interacts with the target and produces a spatial phase gradient through the intensity-dependent Stark phase, thereby modifying the XUV wavefront and redirecting the emission. This happens because the phase gradient yields a transverse contribution to the wave vector, $k_{\perp} = d\phi_s/dr$, where ϕ_s is the accumulated Stark phase, which alters the direction of wave-vector phase matching. Since the AC Stark shift is state specific, the emission associated with different excited states can be redirected in different ways by the IR interaction.

To understand the expected behavior of the OOM redirection we consider the spatial dependence of the accumulated Stark phase in the limit where ionization can be ignored [8]:

$$\phi_s(\mathbf{r}) = \frac{1}{\hbar} \int_{\tau_{\text{IR}}} \Delta E(\mathbf{r}, t) dt, \quad (1)$$

where $\Delta E(\mathbf{r}, t)$ is the intensity-dependent Stark shift of a specific resonance, \hbar is the reduced Planck constant, and τ_{IR} is the total duration of the IR probe pulse. For Rydberg states the shift in energy with increasing field intensity is positive and close to linear. Spatially offsetting a smaller pump beam and a larger probe beam, imprints an approximately linear phase gradient across the pump beam so that all the np emission is redirected in the same direction, as observed in [8, 9] [upward in Fig. 2(a)]. If, however, the intensity-dependent phase shift for a state as a function of intensity is nonlinear, as in Fig. 2(b), the phase front of the emission can be altered in a more complex way. In particular, if the Stark phase decreases for low intensity and increases at high intensity, the XUV emission can be redirected through both negative and positive divergence angles, resulting in an effective beamsplitter for XUV light.

With the pump and probe beams offset as in Fig. 2, the redirected light can be adjusted via the focal overlap between the pump and probe beams, and the spatial intensity profile of the probe pulse at the target. For resonances long-lived with respect to the duration of the

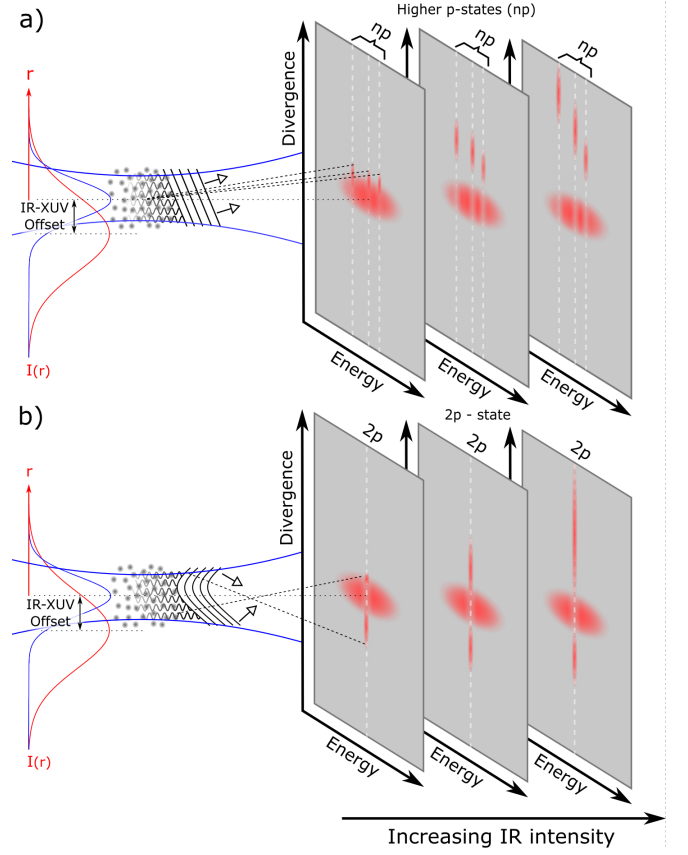


FIG. 2. Illustration of OOM redirection for (a) linear and (b) nonlinear Stark phase behavior. A small (blue in color version online) pump XUV beam excites the atoms. (a) Following interaction with a spatially offset, larger (red in color version online) probe IR beam, the XUV emission phase front can become tilted if the Stark phase response is approximately linear, as for the np states. (b) A nonlinear Stark response can result in the phase front being tilted in one direction at low intensity, and the other direction at high intensity. The amount of phase accumulation, and consequently the phase gradient spatial profile, is determined by the IR intensity and spatial distribution across the XUV pump focus.

pulses, redirection can occur many tens or hundreds of fs after the excitation pulse has passed, allowing this measurement to be performed outside of temporal overlap of the pump and probe pulses. For the OOM technique, the lifetime of the excited state must be sufficient for an appreciable Stark-shifting to occur, enabling redirection. Redirection from short-lived states could require shorter pulse durations to satisfy the condition for IR-free XUV excitation of the excited ensemble.

III. EXPERIMENTAL SETUP

The experimental setup is a pump-probe scheme where both pulses are derived from the same 1 kHz repetition rate, 800 nm titanium-sapphire laser system producing pulses of ~ 20 fs duration. Annular mirrors are used to spatially separate and recombine the pump and probe

beam-paths. The outer, annular part of the IR beam is focused into a pulsed gas jet of argon atoms to produce the pump XUV light through high-order harmonic generation (HHG) [21–25]. To shift the 13th harmonic into resonance with the 1s-2p transition in the helium target gas, the HHG process is driven at sufficient intensity to induce blue-shifting of the generated harmonics [26]. This blue-shifting, along with overlaid second order diffraction components from the diffraction grating, produces the observed near continuous harmonic spectrum detected on axis (Fig. 1). An iris is positioned downstream in the HHG beam-path to limit the divergence of the XUV beam and thereby suppress any off-axis emission in the far field that is not due to the IR interaction. This iris also acts to reduce the residual fundamental light from the HHG process. The inner part of the IR beam bypasses the HHG gas and serves as the probe. Both pump and probe beams are focused into the target helium gas using a toroidal mirror. Through imaging we measure the probe focus to be $\sim 160\mu\text{m}$ full width at half maximum (FWHM). From the ability to redirect the XUV np emission either up (as in Fig. 1) or down by adjusting the XUV-IR spatial offset, we deduce that the XUV focus is smaller than this. The beams are recombined at a small angle, and the probe is offset spatially from the pump in the interaction region to capture the steepest slope of the IR spatial intensity distribution. The delay between the pump and probe pulses is controlled using a precision translation stage, and the delay of the IR probe used in the following measurements is several tens of fs after temporal overlap. The helium pressure has been adjusted to optimize the 2p emission and avoid effects of resonant pulse propagation (RPP) [27, 28]. The spectrally resolved spatial profile of the XUV light is recorded in the far field using a flat-field spectrometer, with a micro-channel plate detector, imaged by a CCD camera. The probe intensity in the interaction region is controlled using a motorized, zero-aperture iris after the focus in the IR beam path.

IV. RESULTS

A. Experiment

Fig. 3(a) shows the evolution of the XUV emission from a narrow energy region around the 2p-state excitation energy with an iris-opening parameter that varies from 0 (fully closed) to 1 (fully open). Note that the exact mapping between this opening parameter and the actual iris diameter is not perfectly known. The estimated IR peak intensity for the fully open iris is $9 \times 10^{12} \text{ W/cm}^2$. The effect of the iris is two-fold since it changes both the total energy in the probe beam and its confocal parameter. The figure shows that at low intensity (up to iris opening ≈ 0.35), the 2p emission is redirected only downward (opposite to the np emission), whereas at higher intensities it splits and is redirected both up and down. This indicates that the intensity dependence of the accumulated Stark phase changes sign, or, equivalently, that the shift in energy changes from being negative to being positive.

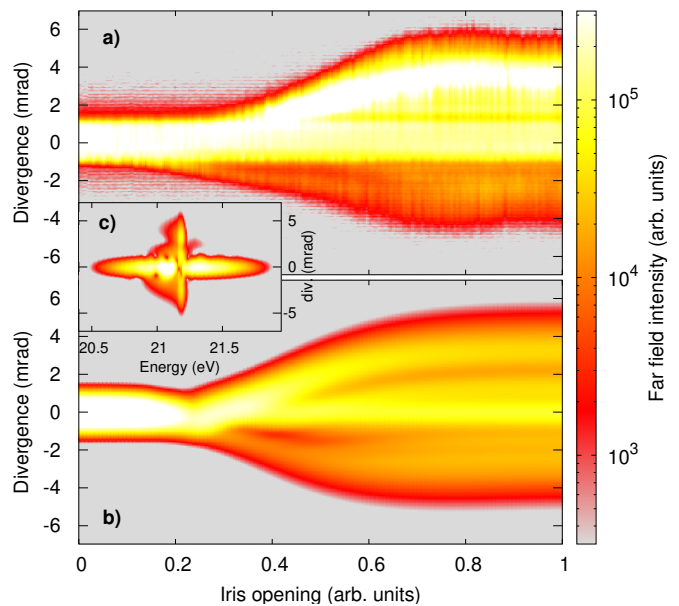


FIG. 3. Far field divergence of the 2p emission in (a) the experiment, and (b) the calculation. (c) shows the calculated spatio-spectral profile of the 2p emission for a fully open iris.

B. Theory

For comparison with experiment, Fig. 3(b) shows the 2p emission calculated by solving the coupled MWE-TDSE equations for a He gas interacting with two spatially offset XUV and IR fields [29]. The 2p-resonant XUV pump pulse duration is 4 fs, with a focus of $28\mu\text{m}$ FWHM and a peak intensity of 10^{11} W/cm^2 , and the 800 nm probe pulse duration is 27 fs, with a focus of $56\mu\text{m}$ FWHM and a peak intensity of 10^{13} W/cm^2 when the iris is fully opened. The two pulses are delayed with respect to each other by 40 fs and spatially offset by $35\mu\text{m}$. We use a thin $10\mu\text{m}$ He gas medium with a density of $5 \times 10^{18} \text{ cm}^{-3}$ to avoid effects of RPP. To account for the non-cylindrical symmetry, the MWE calculations were performed in one transverse direction (1D). This means that the iris in the calculations, which is applied before focusing the IR beam, does not exactly replicate the effect of the experimental iris on the two-dimensional (2D) beam. In particular, the intensity of the 1D beam increases too slowly as the 1D iris diameter is increased as compared to the experiment. To compensate for this, we multiply the intensity after the aperture, I_a , by the square of the intensity loss, I_a/I_0 , where I_0 is the intensity before the aperture. The two factors of I_a/I_0 mimic the extra drop in intensity due to the energy loss and the increased confocal parameter.

The calculations can also provide further insight into the 2p emission. Fig. 3(c) shows the calculated far-field, spatio-spectral profile of the XUV light near the 2p state for a fully open iris, clearly exhibiting both up- and down-directed emission. In the calculation, we can block out selected parts of the near field interaction region, which alters the far-field signal. From this we confirm the pic-

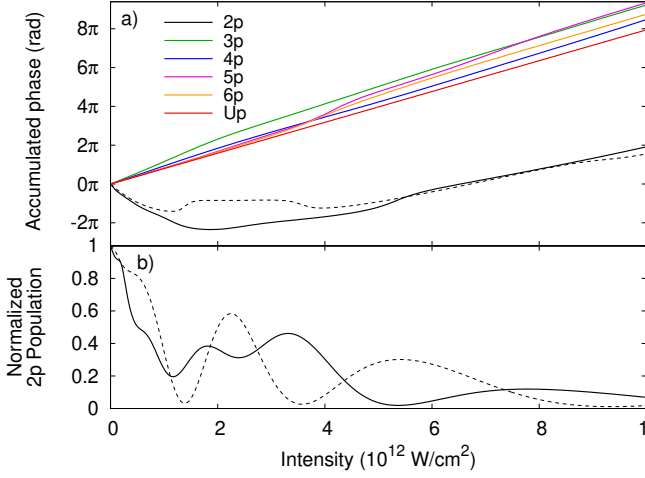


FIG. 4. (a) TDSE calculation of the total IR-induced phase accumulation for the different np states in helium after interaction with a resonant 4 fs, 10^{11} W/cm^2 pump and a subsequent 40 fs delayed, 800 nm, 27 fs IR probe pulse for a range of different probe peak intensities. The lower solid line corresponds to the 2p state, while the upper solid lines correspond to the higher p states, and U_p . (b) 2p population at the end of the IR pulse normalized by the 2p population at the end of the XUV pulse. The dashed lines correspond to the 2p (a) phase and (b) population in the same conditions but with an 829 nm IR pulse which drives near-resonant two-photon Rabi oscillations between the 2p and the 5f states. Rapid phase variations are observed at intensities matching near zeros in the 2p state populations.

ture illustrated in Fig. 2(b): the downward 2p emission comes from the upper part of the probe beam where the intensity is low, and the upward 2p emission comes from the lower part of the probe beam where the intensity is high. Fig. 3(b) shows the calculated behavior as a function of iris opening. Allowing for the differences between the experiment and theory discussed above, the general features of the calculated behavior agree very well with those of the experiment, both in terms of the down-only redirection at low intensity, and the asymmetry between up and down-directed emission at higher intensities. From the calculations we find that the detailed behavior as a function of iris opening, especially in terms of the up/down emission asymmetry, is sensitive to the peak IR intensity, the relative sizes of the pump and probe beams, and in particular to the spatial offset of the pump-probe foci. This suggests that more precise experimental control over the probe spatial profile, for example through the use of spatial light modulators [30, 31], could allow for future reconstruction of the intensity-dependence of state-resolved Stark shifts from the experimental results, and to finely control and tailor the XUV emission in space and time.

Finally, to understand the observed intensity dependence of the 2p-emission redirection, Fig. 4(a) shows the TDSE-calculated accumulated Stark phase for each of

the excited states discussed in this paper. The intensity axis denotes the peak intensity of the same 800 nm, 27 fs IR pulse used in Fig. 3(b), and the phase is extracted at the end of the IR pulse by projecting onto the field free states. The accumulated phase due to a Stark shift equal to the ponderomotive energy $\Delta E = U_p$ is shown for comparison, and marks the simplest possible linear Stark phase. This figure shows that the accumulated phase increases approximately as $U_p \tau_{\text{IR}}$ for the 3p and higher-lying np states (upper solid lines). The phase of the 2p state (lower solid line), however, exhibits a completely different behavior. It drops rapidly at low intensity, below approximately $1.9 \times 10^{12} \text{ W/cm}^2$, then reverses and increases almost linearly at higher intensity, although slower than the higher np states. These general trends are in good agreement with the results discussed above, and can be understood with the following considerations. At low intensity, the 2p state couples strongly to the 3s and 3d states, which are in close to one-photon resonance with it. Indeed, we find that the low intensity behavior of the 2p phase can be accurately reproduced with a three-level model using only the 2p, 3s and 3d states (not shown). We also find, as expected for near-resonant interactions, that the sign of the 2p-3s and 2p-3d detuning controls the sign of the low intensity phase shift. The 2s state, which is below the 2p state by about half an IR photon, is too far detuned to play a significant role. Conversely, at high intensities, the electric field strongly distorts the potential felt by the electron so that it behaves increasingly like a free electron in an oscillating field, and the 2p state presents a near-linear phase more similar to the higher-lying np states.

At low and moderate intensities, the IR field also enables near-resonant two-photon coupling between the 2p and higher-lying nf states [32] that drives Rabi oscillations between these states, as can be seen in the 2p population shown in Fig. 4(b). These oscillations are highly sensitive to the IR wavelength and are best observed at a slightly longer wavelength (829 nm) than the one used in the experiment. The longer wavelength 2p population and phase are shown as dashed lines in Fig. 4(a) and (b). Note that the minima in the 2p population are associated with rapid variations of the phase (near 1.5 and $3.5 \times 10^{12} \text{ W/cm}^2$), as expected for Rabi flopping [33]. This provides another interesting perspective on XUV spatial control through OOM: in the resonant case, both the phase and the amplitude of the XUV field can be modulated through IR-control of the Stark shift and the population of the resonant state.

V. SUMMARY

In summary, we have used the all-optical OOM technique to probe the Stark-induced phase change of excited states in matter. We have experimentally observed the change of sign of the 2p-state phase accumulation as the intensity of the non-resonant IR field is increased, in good agreement with MWE-TDSE based calculations. This result opens the possibility for the future study of Stark

phases in more complicated atoms or molecules, where the states and/or their dipole couplings may be less well known, and even allow for reconstruction of the phase accumulation from the experimental result given tighter control over the experimental parameters. We also emphasize the potential for the OOM technique to be used to probe unknown Stark phases of states embedded in the continuum, which although beyond the scope of the work presented here, would be interesting to study in future experiments. This work also highlights the potential for the OOM technique to control XUV frequency light in different ways, such as by creating variable beam-splitters in the XUV by exploiting the nonlinear response of states to IR intensity changes.

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- [1] F. Calegari, G. Sansone, S. Stagira, C. Vozzi, and M. Nisoli, *J. Phys. B: At. Mol. Opt. Phys.* **49**, 062001 (2016).
 - [2] M. Wu, S. Chen, S. Camp, K. J. Schafer, and M. B. Gaarde, *Journal of Physics B: Atomic, Molecular and Optical Physics* **49**, 062003 (2016).
 - [3] W. Cao, E. R. Warrick, A. Fidler, D. M. Neumark, and S. R. Leone, *Physical Review A* **94**, 053846 (2016).
 - [4] W. Cao, E. R. Warrick, A. Fidler, S. R. Leone, and D. M. Neumark, *Physical Review A* **94**, 021802(R) (2016).
 - [5] S. Baker, J. S. Robinson, C. A. Haworth, H. Teng, R. A. Smith, C. C. Chiril, M. Lein, J. W. G. Tisch, and J. P. Marangos, *Science* **312**, 424 (2006).
 - [6] M. Chini, X. Wang, Y. Cheng, Y. Wu, D. Zhao, D. A. Telnov, S.-I. Chu, and Z. Chang, *Scientific Reports* **3**, 10.1038/srep01105 (2013).
 - [7] N. Shivaram, H. Timmers, X.-M. Tong, and A. Sandhu, *Physical Review Letters* **108**, 193002 (2012).
 - [8] S. Bengtsson, E. W. Larsen, D. Kroon, S. Camp, M. Miranda, C. L. Arnold, A. L'Huillier, K. J. Schafer, M. B. Gaarde, L. Rippe, and J. Mauritsson, *Nature Photonics* **11**, 252 (2017).
 - [9] S. Bengtsson and J. Mauritsson, *J. Phys. B: At. Mol. Opt. Phys.* **52**, 063002 (2019).
 - [10] L. Drescher, O. Kornilov, T. Witting, G. Reitsma, N. Monserud, A. Rouze, J. Mikosch, M. J. J. Vrakking, and B. Schütte, *Nature* **564**, 91 (2018).
 - [11] A. Kaldun, C. Ott, A. Blättermann, M. Laux, K. Meyer, T. Ding, A. Fischer, and T. Pfeifer, *Phys. Rev. Lett.* **112**, 103001 (2014).
 - [12] A. Blättermann, C. Ott, A. Kaldun, T. Ding, V. Stoss, M. Laux, M. Rebholz, and T. Pfeifer, *Opt. Lett.* **40**, 3464 (2015).
 - [13] A. Fleischer, O. Kfir, T. Diskin, P. Sidorenko, and O. Cohen, *Nature Photonics* **8**, 543 (2014).
 - [14] O. Kfir, P. Grychtol, E. Turgut, R. Knut, D. Zusin, D. Popmintchev, T. Popmintchev, H. Nembach, J. M. Shaw, A. Fleischer, H. Kapteyn, M. Murnane, and O. Cohen, *Nature Photonics* **9**, 99 (2015).
 - [15] N. B. Delone and V. P. Krainov, *Physics-Uspekhi* **42**, 669 (1999).
 - [16] C. Ott, A. Kaldun, P. Raith, K. Meyer, M. Laux, J. Evers, C. H. Keitel, C. H. Greene, and T. Pfeifer, *Science* **340**, 716 (2013).
 - [17] H. Mashiko, T. Yamaguchi, K. Oguri, A. Suda, and H. Gotoh, *Nature Communications* **5** (2014), 10.1038/ncomms6599.
 - [18] W. C. Martin, *Physical Review A* **36**, 3575 (1987).
 - [19] R. G. Brewer and R. L. Shoemaker, *Physical Review A* **6**, 2001 (1972).
 - [20] F. A. Hopf, R. F. Shea, and M. O. Scully, *Physical Review A* **7**, 2105 (1973).
 - [21] A. McPherson, G. Gibson, H. Jara, U. Johann, T. S. Luk, I. A. McIntyre, K. Boyer, and C. K. Rhodes, *JOSA B* **4**, 595 (1987).
 - [22] M. Ferray, A. L'Huillier, X. F. Li, L. A. Lompre, G. Mainfray, and C. Manus, *Journal of Physics B: Atomic, Molecular and Optical Physics* **21**, L31 (1988).
 - [23] K. J. Schafer, B. Yang, L. F. DiMauro, and K. C. Kulander, *Physical Review Letters* **70**, 1599 (1993).
 - [24] P. B. Corkum, *Physical Review Letters* **71**, 1994 (1993).
 - [25] M. Lewenstein, P. Balcou, M. Y. Ivanov, A. L'Huillier, and P. B. Corkum, *Physical Review A* **49**, 2117 (1994).
 - [26] W. M. Wood, C. W. Siders, and M. C. Downer, *Physical Review Letters* **67**, 3523 (1991).
 - [27] U. van Bürck, *Hyperfine Interactions* **123/124**, 483 (1999).
 - [28] C.-T. Liao, A. Sandhu, S. Camp, K. J. Schafer, and M. B. Gaarde, *Phys. Rev. Lett.* **114**, 143002 (2015).
 - [29] M. B. Gaarde, C. Buth, J. L. Tate, and K. J. Schafer, *Physical Review A* **83**, 013419 (2011).
 - [30] S. Hasegawa, H. Ito, H. Toyoda, and Y. Hayasaki, *Optics Express* **24**, 18513 (2016).
 - [31] B. sun, P. S. Salter, C. Roider, A. Jesacher, J. Strauss, J. Heberle, M. Schmidt, and M. J. Booth, *Light: Science and Applications* **7**, 17117 (2018).
 - [32] M. Fushitani, C.-N. Liu, A. Matsuda, T. Endo, Y. Toida, M. Nagasono, T. Togashi, M. Yabashi, T. Ishikawa, Y. Hikosaka, T. Morishita, and A. Hishikawa, *Nature Photonics* **10**, 102 (2016).
 - [33] M. Wu, S. Chen, K. J. Schafer, and M. B. Gaarde, *Physical Review A* **87**, 013828 (2013).