Hydrodynamic Gradient Expansion Diverges beyond Bjorken Flow

Michal P. Heller, 1,2,3,* Alexandre Serantes, 2,4,† Michał Spaliński, 2,5,‡ Viktor Svensson, 2,1,§ and Benjamin Withers ¹Max Planck Institute for Gravitational Physics (Albert Einstein Institute), 14476 Potsdam-Golm, Germany ²National Centre for Nuclear Research, 02-093 Warsaw, Poland ⁵Department of Physics and Astronomy, Ghent University, 9000 Ghent, Belgium ⁴Departament de Física Ouàntica i Astrofísica, Institut de Ciències del Cosmos (ICCUB), Facultat de Física, Universitat de Barcelona, Martí i Franquès 1, ES-08028 Barcelona, Spain

⁵Physics Department, University of Białystok, 15-245 Białystok, Poland ⁶Mathematical Sciences and STAG Research Centre, University of Southampton, Highfield, Southampton SO17 1BJ, United Kingdom

(Received 28 October 2021; revised 21 January 2022; accepted 28 February 2022; published 25 March 2022)

The gradient expansion is the fundamental organizing principle underlying relativistic hydrodynamics, yet understanding its convergence properties for general nonlinear flows has posed a major challenge. We introduce a simple method to address this question in a class of fluids modeled by Israel-Stewart-type relaxation equations. We apply it to (1+1)-dimensional flows and provide numerical evidence for factorially divergent gradient expansions. This generalizes results previously only obtained for (0+1)dimensional comoving flows, notably Bjorken flow. We also demonstrate that the only known nontrivial case of a convergent hydrodynamic gradient expansion at the nonlinear level relies on Bjorken flow symmetries and becomes factorially divergent as soon as these are relaxed. Finally, we show that factorial divergence can be removed using a momentum space cutoff, which generalizes a result obtained earlier in the context of linear response.

DOI: 10.1103/PhysRevLett.128.122302

Introduction.—Hydrodynamics plays a pivotal role in the description of nonequilibrium phenomena, with applications ranging from condensed matter systems [1] to scenarios in astrophysics [2–4] or nuclear physics [5,6]. The reason is that hydrodynamics captures the infrared behavior of any medium endowed with conserved quantities. For a given set of conserved currents, the expression of hydrodynamic behavior rests on the derivative expansion in the spirit of an effective field theory [7–11]. For a neutral relativistic fluid, the natural choice of dynamical variables are the energy density $\mathcal{E}(x)$ and the unit-normalized fluid velocity $U(x) = U^{\mu}(x)\partial_{\mu}$, with the conserved currents, $T^{\mu\nu}$, given by the constitutive relation

$$T^{\mu\nu} = \mathcal{E}U^{\mu}U^{\nu} + P(\mathcal{E})(g^{\mu\nu} + U^{\mu}U^{\nu}) + \Pi^{\mu\nu}. \tag{1}$$

Here, the first two terms describe ideal flow with g being the Minkowski metric and $\Pi^{\mu\nu}$ captures dissipative effects organized as

$$\Pi_{\mu\nu} = \sum_{n=1}^{\infty} \epsilon^n \Pi_{\mu\nu}^{(n)}[\mathcal{E}, U], \tag{2}$$

Published by the American Physical Society under the terms of the Creative Commons Attribution 4.0 International license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP³.

where $\Pi_{\mu\nu}^{(n)}$ contains *n* spacetime derivatives of \mathcal{E} , *U* and we have introduced ϵ as a formal derivative-counting parameter. The gradient expansion in Eq. (2) is defined up to redundancies associated with frame choice and current conservation $\nabla_{\mu}T^{\mu\nu}=0$.

Understanding the character of the expansion (2) constitutes a fundamental open problem. Is it convergent, in such a way that subsequent truncations are progressively more accurate? If not, how does its divergent nature relate to the empirical success of low-order truncations?

Studies of comoving flows in Refs. [12–31], in which all fluid flow lines can be mapped to each other under symmetry transformations, rendering the problem effectively (0+1) dimensional, have been instrumental in advancing our understanding of the hydrodynamic expansion (2). Among these, Bjorken flow [32] in conformally invariant theories is the most thoroughly explored example due to its role in studies of quark-gluon plasma. In these cases a particular strategy to solve the dynamical equations is an expansion in the Knudsen number, 1/w [33]. It is possible to compute a sufficient number of terms to assess that these expansions are factorially divergent. The expansion in ϵ defined in Eq. (2) encapsulates the expansion in 1/w, as we review in the Supplemental Material [34]. Another well-studied example of a comoving flow is the Gubser flow which is reached by Weyl transformation from a (0 + 1) flow on $dS_3 \times \mathbb{R}$ [35,36].

Outside the realm of comoving flows, the only generic result on (2) was restricted to the linear response regime [37]. It showed that depending on how the momentum space support $k_{\rm max}$ of ${\cal E}$ and U^μ compares to an intrinsic scale of the underlying microscopic theory k_* [38], the gradient-expanded constitutive relations could either be convergent $(k_{\rm max} < k_*)$, geometrically divergent $(k_{\rm max} < \infty)$, or factorially divergent $(k_{\rm max} \to \infty)$.

In this Letter, we break free both from the symmetry constraints of comoving flows and the conveniences of linearization to address, for the first time, the large-order behavior of the hydrodynamic gradient expansion beyond comoving flows at the fully nonlinear level. Specifically, we ask the following question: Given a generic nonequilibrium, nonlinear configuration of \mathcal{E} , U^{μ} arising on-shell, what is nature of the expansion in ϵ (2) when evaluated on this solution? We answer this question by introducing a simple method that allows one to calculate Eq. (2) up to high order on a desktop computer. In this Letter we illustrate it in two examples. The first is the model put forward by Baier, Romatchske, Son, Starinets, and Stephanov (BRSSS) in Ref. [39], while the second is the model originally introduced by Denicol and Noronha (DN) in Ref. [27]. Both theories are representative examples within a broad class of models employing the Israel-Stewart approach to embed hydrodynamics in a framework compatible with relativistic causality [40,41]. Our method applies to any member of this class, such as [42], and covers also equations with more than one derivative of $\Pi^{\mu\nu}$, such as [43–45]. Each member of this class generally gives rise to infinitely many transport coefficients in the gradient expansion (2) that are specific to it.

Results in the BRSSS model.—In this Letter, we consider the restriction of the BRSSS model to conformal fluids in d spacetime dimensions (d=4 in our numerics). We work in the Landau frame and take $\Pi_{\mu\nu}U^{\nu}=0$. The model is defined by promoting $\Pi_{\mu\nu}$ to a set of independent degrees of freedom subject to a relaxation equation,

$$\Pi^{\mu\nu} = -2\eta \mathcal{D}^{\langle\mu} u^{\nu\rangle} - \tau_{\Pi} U^{\alpha} \mathcal{D}_{\alpha} \Pi^{\mu\nu} + \frac{\lambda_{1}}{\eta^{2}} \Pi_{\lambda}^{\langle\mu} \Pi^{\nu\rangle\lambda}, \quad (3)$$

where we neglected terms not relevant to the flow we consider. \mathcal{D}_{μ} is the Weyl-covariant derivative [46], and the angle brackets instruct one to take the symmetric, transverse, and traceless part of the tensor they act upon. The relaxation time τ_{Π} , the shear viscosity η , and λ_1 are transport coefficients. Conformal invariance demands that these quantities depend on the local temperature T, defined by the relation $\mathcal{E} = \mathcal{E}_0 T^d$, as

$$\eta = C_{\eta} \frac{4\mathcal{E}}{3T}, \qquad \tau_{\Pi} = \frac{C_{\tau_{\Pi}}}{T}, \qquad \lambda_{1} = C_{\lambda_{1}} \frac{\eta}{T}, \qquad (4)$$

where C_{η} and $C_{\tau_{\Pi}} > 0$. The equations of motion of BRSSS theory are given by (3) and the conservation equation

 $\nabla_{\mu}T^{\mu\nu}=0$, where the energy-momentum tensor is specified in terms of \mathcal{E} , U, and $\Pi_{\mu\nu}$ as in (1).

The fluid flows we focus on are characterized as follows. We separate the spatial coordinates into one longitudinal direction, x, and d-2 transverse directions, x_1, \ldots, x_{d-2} , demanding isotropy and translational invariance in the transverse hyperplane spanned by x_i . Hence, the nontrivial dynamics is confined to the longitudinal plane spanned by t and t, and our fluid flows are t dimensional. We refer to these fluid flows as t longitudinal. At the linearized level such flows would correspond to sound wave propagation. See, e.g., Ref. [47] for a study of longitudinal flows in a quark-gluon plasma context [48].

The most general fluid velocity for a longitudinal flow is parameterized by a single degree of freedom, u, as

$$U = U^{\mu} \partial_{\mu} = \cosh u \partial_{t} + \sinh u \partial_{x}. \tag{5}$$

Furthermore, any tensor which is symmetric, transverse to U^{μ} , and traceless is described in terms of a single additional degree of freedom that we pick as

$$\Pi^{\mu\nu} = (2 - d)\Pi_{\star}\Sigma^{\mu\nu} \tag{6}$$

where $\Sigma^{\mu\nu}\equiv g^{\mu\nu}+U^{\mu}U^{\nu}-[(d-1)/(d-2)]P_T^{\mu\nu}$ and $\Pi_{\star}=[1/(d-2)]P_T^{\mu\nu}\Pi_{\mu\nu}$, with $P_T^{\mu\nu}$ being the projector in the transverse directions.

We now consider the expansion (2) applied to the conformal BRSSS model for longitudinal flows. This is facilitated by a numerical algorithm which makes a computation of (2) to large orders tractable. Since ϵ counts derivatives it can be introduced by taking (3) and replacing $\nabla_{\mu} \rightarrow \epsilon \nabla_{\mu}$ together with positing a perturbative ansatz for Π_{\star} , as follows,

$$\mathcal{D}_{\alpha} \to \epsilon \mathcal{D}_{\alpha}, \qquad \Pi_{\star} \to \sum_{n=1}^{\infty} \Pi_{\star}^{(n)} \epsilon^{n}.$$
 (7)

This leads to the following recursion relation:

$$\Pi_{\star}^{(1)} = -\frac{2}{d-2} \eta P_T^{\mu\nu} \mathcal{D}_{\langle \mu} U_{\nu \rangle},\tag{8a}$$

$$\Pi_{\star}^{(n+1)} = -\tau_{\Pi}(U \cdot \partial)\Pi_{\star}^{(n)} - \frac{d(\partial \cdot U)}{d-1}\tau_{\Pi}\Pi_{\star}^{(n)}
- (d-3)\frac{\lambda_{1}}{\eta^{2}} \sum_{m=1}^{n} \Pi_{\star}^{(m)}\Pi_{\star}^{(n+1-m)}, \qquad n > 1. \quad (8b)$$

Here, \mathcal{E} and U are not expanded in ϵ . Therefore to proceed to evaluate (8) we must first find \mathcal{E} and U for a given choice of flow. These (as well as the exact Π_{\star}) can be obtained by numerically solving the BRSSS equations of motion as an initial value problem without invoking an ϵ expansion. Once \mathcal{E} and U are known, the recursion relation (8) can be

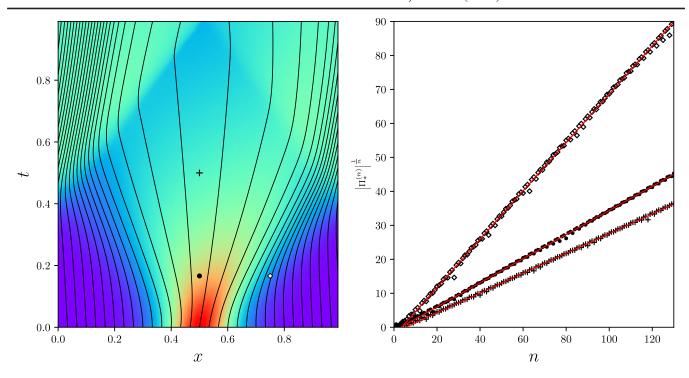


FIG. 1. Left panel: solution to an initial value problem in BRSSS. Color shows the temperature profile T(t,x), and the black solid lines are flow lines of velocity U. Here t, x are obtained from standard Minkowski coordinates by periodically identifying $x \sim x + 1$. Initial conditions were provided at t = 0 corresponding to a strong Gaussian overdensity. The color scale ranges from min T = 0.76 (violet) to max T = 1.76 (red). Also included are three marked points (disk, plus, open diamond). Right panel: root plot for the hydrodynamic expansion of the constitutive relation (2) for BRSSS, evaluated using the recursion relation (8) for the solution shown in the left panel, at each of the marked points with matching marker shapes. To guide the eye, in red we show straight lines fit to the range $n \in [20, 130]$, which correspond to $\Pi_*^{(n)} \sim \Gamma(n)$ at large n. Numerical convergence of our results is demonstrated in the Supplemental Material.

efficiently evaluated numerically to high orders. Careful consideration of resolution and precision is required, as this procedure involves high numbers of successive derivatives of the background solution $\mathcal E$ and U. This is further discussed in the Supplemental Material, where we show that our numerical results are convergent. Our approach applies to the whole class of theories which build on the Israel-Stewart approach.

Note that solving the recursion relation (8) analytically is prohibitively expensive due to the fast growth in the number of terms contributing at each order. In particular, for the case $\lambda_1 = 0$ we observed an exponential growth of the number of individual contributions at each order. Our method circumvents this difficulty.

We have applied our approach in the BRSSS model across a wide variety of initial conditions, transport coefficients, and spacetime points. We find factorial growth in all cases considered. We illustrate this in Fig. 1 with one representative example in which we consider a strong Gaussian overdensity for our initial conditions and adopt a periodic compactification of the spatial direction. The left panel shows T and flow lines of U and highlights three spacetime point samples. The right panel root plot demonstrates factorial growth at these sampled points. Further details are provided in the Supplemental Material.

Momentum cutoff.—In previous work [37] we showed that a momentum-space cutoff gives at most a geometrically divergent hydrodynamic expansion for linear deviations from equilibrium. This result naturally extends to strongly nonlinear scenarios, as we now demonstrate. So far in this Letter we have used a numerical grid simply as a tool to approximate the continuum, but we now push beyond the continuum picture and reevaluate the grid in a new role as a physical lattice which naturally enforces a momentum-space cutoff. In the BRSSS model, when $\lambda_1 = 0$ the recursion relation (8) can be written as

$$\Pi_{\star}^{(n+1)} = \mathcal{M}\Pi_{\star}^{(n)} \qquad n > 1, \tag{9}$$

where $\mathcal{M} = -\tau_{\Pi}(U \cdot \partial) - [d(\partial \cdot U)/(d-1)]$ is a differential operator independent of n, depending only on the background solution \mathcal{E} , U. For a grid of dimensions $N_x \times N_t$, each $\Pi_{\star}^{(j)}$ can be written as a $N_x N_t$ -sized vector, and \mathcal{M} accordingly as a $N_x N_t \times N_x N_t$ square matrix. Thus, on a lattice the expansion is ultimately only geometrically growing at a rate set by the largest eigenvalue of \mathcal{M} , which scales with the inverse lattice spacing. In Fig. 2 this is demonstrated by utilizing a deliberately low resolution lattice to allow for evaluating the hydrodynamic expansion

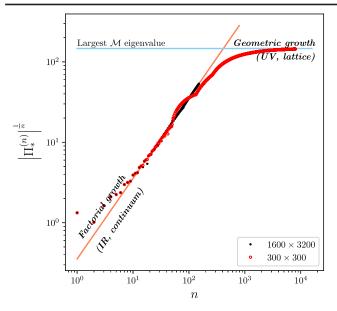


FIG. 2. Working on a lattice gives a window of factorial growth—where it successfully approximates the continuum—before yielding to a geometrically growing hydrodynamic expansion regulated by the lattice. The asymptotic value reached is governed by the lattice spacing as discussed in the text. Black disks are those from Fig. 1 corresponding to the disk spacetime marker point. Red circles are the same simulation and spacetime point, but on a coarse numerical grid and evaluated to hydrodynamic order n = 8000. This plot serves also as an indication of a convergence of our approach.

to order n=8000. It shows the transition from factorial growth where the continuum approximation holds, to the geometrically divergent asymptotic behavior governed by the aforementioned eigenvalue. We have also verified numerically that this result holds at $\lambda_1 \neq 0$, with the definition of \mathcal{M} as given.

Resolving the DN model tension.—Bjorken flow is a boost-invariant longitudinal flow such that the dynamics depends on the proper time $\tau = \sqrt{t^2 - x^2}$ only. In Ref. [27] the authors analyzed a Knudsen number expansion for an ultrarelativistic gas of hard spheres undergoing Bjorken flow. While in all the other models such expansions have been observed to be factorially divergent, in Ref. [27] the terms grow only geometrically, with convergence ensuing for a Knudsen number smaller than a critical value. Our objective is to reanalyze this physical scenario using the expansion in ϵ . For Bjorken flow we find analogous results, namely, geometric growth; however, our method allows us to explore what happens when these symmetries are relaxed.

We work in d=4. As in the conformal BRSSS model, the energy-momentum tensor in the DN model is traceless and decomposed as in Eq. (1), with $\Pi_{\mu\nu}$ still obeying Eq. (3) with $\lambda_1=0$. Hence, the recursion relations giving $\Pi_{\star}^{(n)}$ still take the form (8), again with zero λ_1 . The differences start with the inclusion of a conserved current $J^{\mu}=\rho U^{\mu}$, where

 ρ is the particle density. Furthermore, τ_{Π} and η are not fixed purely in terms of the local temperature T, but rather obey

$$\mathcal{E} = 3\rho T, \qquad \eta = \frac{a}{\sigma_T} T, \qquad \tau_{\Pi} = \frac{ab}{4\sigma_T} \frac{1}{\rho}, \quad (10)$$

where σ_T is the total cross section and a, b are positive dimensionless constants.

For Bjorken flow, the conservation of the particle current J^{μ} entails that the particle density ρ decouples from the energy-momentum tensor. One has that

$$\rho(\tau) = \frac{\rho_0 \tau_0}{\tau},\tag{11}$$

where $\rho_0 = \rho(\tau_0)$ is the initial particle density. Hence,

$$\tau_{\Pi} = \frac{1}{4} ab \text{Kn}\tau, \tag{12}$$

where $\text{Kn} = 1/(\rho_0 \tau_0 \sigma_T)$ is the Knudsen number. In the DN model for Bjorken flow it is time independent.

To assess the large-order behavior of the expansion in ϵ , Eq. (2), we first note that one can find a closed-form expression for $\Pi_{\star}^{(n)}$,

$$\Pi_{\star}^{(n)}(\tau) = \frac{2}{3} a \operatorname{Kn} \rho_0 \left(\frac{\tau_0}{\tau}\right)^{\frac{4}{3}} \left(-\frac{ab \operatorname{Kn}}{4}\right)^{n-1} \times (\tau \partial_{\tau})^{n-1} \left[\left(\frac{\tau}{\tau_0}\right)^{\frac{1}{3}} T(\tau)\right], \tag{13}$$

a fact that relies crucially on Eq. (12). Second, we recall that T can also be determined exactly [27],

$$T(\tau) = T_{0,+} \left(\frac{\tau_0}{\tau}\right)^{\alpha_+} + T_{0,-} \left(\frac{\tau_0}{\tau}\right)^{\alpha_-},$$
 (14)

where $T_{0,\pm}$ depend on initial conditions and α_{\pm} on a,b, and Kn. Together, Eqs. (13) and (14) entail that $\Pi_{\star}^{(n)}$ cannot grow factorially with n at fixed τ , since the repeated action of the differential operator $\tau \partial_{\tau}$ on terms of the form $(\tau/\tau_0)^{(1/3)-\alpha_{\pm}}$ only grows geometrically. Furthermore, Eqs. (13) and (14) also imply that $\Pi_{\star}^{(1)}$ (and therefore all $\Pi_{\star}^{(n>1)}$) is a linear combination of eigenfunctions of the differential operator $\mathcal M$ defined as in Eq. (9), providing another perspective on why the gradient expansion grows geometrically in this case. We refer the reader to the Supplemental Material for further details.

The analysis above relies crucially on the symmetry restrictions of Bjorken flow. Empirically, when relaxing these symmetry restrictions in all cases studied we find that the large-order geometric growth is destroyed and the factorial divergence is restored. We illustrate this in Fig. 3

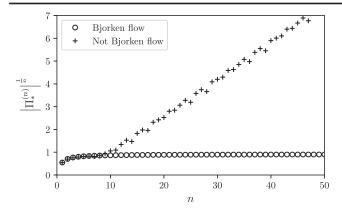


FIG. 3. Deviations from Bjorken symmetries restore factorial growth in the DN model. Open circles represent $|\Pi_{\star}^{(n)}|^{(1/n)}$ for Bjorken flow in the DN model evaluated at $\tau=1$, where we have chosen initial conditions given by $\rho(\tau_0)=4$, $T(\tau_0)=2$ and $\Pi_{\star}(\tau_0)/[\rho(\tau_0)T(\tau_0)]=0.436\,461$ at $\tau_0=0.1$. Crosses represent $|\Pi_{\star}^{(n)}|^{(1/n)}$ at $\tau=1$ and zero rapidity for a longitudinal flow in the DN model which obeys the same initial conditions as Bjorken flow, but with a rapidity-even Gaussian overdensity in ρ of amplitude 0.1 and unit width. We have set a=0.5, b=10 and $\sigma_T=1$. The Bjorken flow coefficients have been obtained from the analytic solution, while the non-Bjorken flow ones from a simulation discussed in the Supplemental Material.

for a longitudinal flow corresponding to a small perturbation of Bjorken flow.

Summary and outlook.—Understanding the behavior of the hydrodynamic gradient expansion at large orders is a challenging question in the foundations of relativistic hydrodynamics. We have proposed a method to compute such series in a large class of models. Applying this to nonlinear longitudinal flows reveals factorially divergent series which we have illustrated with a number of examples. This shows that previously observed instances of factorial growth were not reliant on Bjorken symmetries.

It is natural to ask what are the generic conditions that lead to factorial growth. We have here established that the ability of a system to support arbitrarily large momentum is important; ways around this include working on a lattice, and appropriate restriction of initial data in the linearized case [37], both of which naturally lead to geometric growth. Second, as our analysis of the DN model shows, imposing special symmetries such as boost invariance can also lead to geometric rather than factorial growth. It would be interesting to explore this question further including other natural momentum cutoffs such as microscopic physics and turbulent cascades. For models with recursion relations of the form (9), it may be possible to engineer further examples where the underlying equations of motion give rise to $\Pi_{\star}^{(n)}$ which are eigenfunctions of \mathcal{M} such that the hydrodynamic gradient expansion grows geometrically. This could form the basis of a rigorous mathematical formulation for investigating the genericity of factorial growth.

The picture that is emerging from this work and results in linear response [37] is that the origin of factorial growth at large n in the hydrodynamic gradient expansion is the successive action of n derivatives on the hydrodynamic variables \mathcal{E} , U. This is intimately connected with support of a solution in momentum space. It suggests that having a factorial growth in the *number* of transport coefficients at each order is not necessary.

The factorial divergence of asymptotic series is not an impediment to their practical utility: such series typically provide excellent approximations as long as one does not exceed the so-called order of optimal truncation. Our work makes such investigations possible for a much wider set of flows than previously tractable.

The Gravity, Quantum Fields and Information group at AEI. is supported by the Alexander von Humboldt Foundation and the Federal Ministry for Education and Research through the Sofja Kovalevskaja Award. A. S. was supported by the Polish National Science Centre Grant No. 2018/29/B/ST2/02457 and by Grant No. CEX2019-000918-M funded by MCIN/AEI/10.13039/501100011033. M. S. is supported by the National Science Centre, Poland, under Grants No. 2018/29/B/ST2/02457 and No. 2021/41/B/ST2/02909. B. W. is supported by a Royal Society University Research Fellowship.

- [1] D. A. Bandurin, A. V. Shytov, L. S. Levitov, R. K. Kumar, A. I. Berdyugin, M. Ben Shalom, I. V. Grigorieva, A. K. Geim, and G. Falkovich, Fluidity onset in graphene, Nat. Commun. 9, 4533 (2018).
- [2] M. Shibata, K. Kiuchi, and Y.-i. Sekiguchi, General relativistic viscous hydrodynamics of differentially rotating neutron stars, Phys. Rev. D 95, 083005 (2017).
- [3] M. G. Alford, L. Bovard, M. Hanauske, L. Rezzolla, and K. Schwenzer, Viscous Dissipation and Heat Conduction in Binary Neutron-Star Mergers, Phys. Rev. Lett. 120, 041101 (2018).
- [4] F. S. Bemfica, M. M. Disconzi, and J. Noronha, Causality of the Einstein-Israel-Stewart Theory with Bulk Viscosity, Phys. Rev. Lett. 122, 221602 (2019).
- [5] U. Heinz and R. Snellings, Collective flow and viscosity in relativistic heavy-ion collisions, Annu. Rev. Nucl. Part. Sci. 63, 123–151 (2013).
- [6] W. Busza, K. Rajagopal, and W. van der Schee, Heavy ion collisions: The big picture, and the big questions, Annu. Rev. Nucl. Part. Sci. 68, 339 (2018).
- [7] P. Kovtun, Lectures on hydrodynamic fluctuations in relativistic theories, J. Phys. A **45**, 473001 (2012).
- [8] S. A. Hartnoll, A. Lucas, and S. Sachdev, Holographic quantum matter, arXiv:1612.07324.

^{*}michal.p.heller@aei.mpg.de †alexandre.serantes@ub.edu †michal.spalinski@ncbj.gov.pl \$viktor.svensson@aei.mpg.de b.s.withers@soton.ac.uk

- [9] W. Florkowski, M. P. Heller, and M. Spaliński, New theories of relativistic hydrodynamics in the LHC era, Rep. Prog. Phys. 81, 046001 (2018).
- [10] P. Romatschke and U. Romatschke, *Relativistic Fluid Dynamics in and Out of Equilibrium*, Cambridge Monographs on Mathematical Physics (Cambridge University Press, Cambridge, England, 2019).
- [11] A. Soloviev, Hydrodynamic attractors in heavy ion collisions: A review, arXiv:2109.15081.
- [12] M. P. Heller, R. A. Janik, and P. Witaszczyk, Hydrodynamic Gradient Expansion in Gauge Theory Plasmas, Phys. Rev. Lett. **110**, 211602 (2013).
- [13] A. Buchel, M. P. Heller, and J. Noronha, Entropy production, hydrodynamics, and resurgence in the primordial quark-gluon plasma from holography, Phys. Rev. D 94, 106011 (2016).
- [14] J. Casalderrey-Solana, N. I. Gushterov, and B. Meiring, Resurgence and hydrodynamic attractors in Gauss-Bonnet holography, J. High Energy Phys. 04 (2018) 042.
- [15] M. Baggioli and A. Buchel, Holographic viscoelastic hydrodynamics, J. High Energy Phys. 03 (2019) 146.
- [16] A. Buchel, Non-conformal holographic Gauss-Bonnet hydrodynamics, J. High Energy Phys. 03 (2018) 037.
- [17] I. Aniceto, B. Meiring, J. Jankowski, and M. Spaliński, The large proper-time expansion of Yang-Mills plasma as a resurgent transseries, J. High Energy Phys. 02 (2019) 073.
- [18] M. P. Heller and M. Spaliński, Hydrodynamics Beyond the Gradient Expansion: Resurgence and Resummation, Phys. Rev. Lett. **115**, 072501 (2015).
- [19] G. Başar and G. V. Dunne, Hydrodynamics, resurgence, and transasymptotics, Phys. Rev. D 92, 125011 (2015).
- [20] I. Aniceto and M. Spaliński, Resurgence in extended hydrodynamics, Phys. Rev. D **93**, 085008 (2016).
- [21] G. S. Denicol and J. Noronha, Divergence of the Chapman-Enskog expansion in relativistic kinetic theory, arXiv:1608 .07869.
- [22] W. Florkowski, R. Ryblewski, and M. Spaliński, Gradient expansion for anisotropic hydrodynamics, Phys. Rev. D 94, 114025 (2016).
- [23] M. P. Heller, A. Kurkela, M. Spaliński, and V. Svensson, Hydrodynamization in kinetic theory: Transient modes and the gradient expansion, Phys. Rev. D 97, 091503 (2018).
- [24] G. S. Denicol and J. Noronha, Analytical attractor and the divergence of the slow-roll expansion in relativistic hydrodynamics, Phys. Rev. D **97**, 056021 (2018).
- [25] M. P. Heller and V. Svensson, How does relativistic kinetic theory remember about initial conditions?, Phys. Rev. D **98**, 054016 (2018).
- [26] J.-P. Blaizot and L. Yan, Emergence of hydrodynamical behavior in expanding ultra-relativistic plasmas, Ann. Phys. (Amsterdam) **412**, 167993 (2020).
- [27] G. S. Denicol and J. Noronha, Exact Hydrodynamic Attractor of an Ultrarelativistic Gas of Hard Spheres, Phys. Rev. Lett. **124**, 152301 (2020).
- [28] A. Behtash, C. N. Cruz-Camacho, and M. Martinez, Far-from-equilibrium attractors and nonlinear dynamical systems approach to the Gubser flow, Phys. Rev. D 97, 044041 (2018).

- [29] G. S. Denicol and J. Noronha, Hydrodynamic attractor and the fate of perturbative expansions in Gubser flow, Phys. Rev. D 99, 116004 (2019).
- [30] A. Behtash, S. Kamata, M. Martinez, and H. Shi, Global flow structure and exact formal transseries of the Gubser flow in kinetic theory, J. High Energy Phys. 07 (2020) 226.
- [31] Z. Du, X.-G. Huang, and H. Taya, Hydrodynamic attractor in a Hubble expansion, Phys. Rev. D 104, 056022 (2021).
- [32] J. Bjorken, Highly Relativistic nucleus-nucleus collisions: The central rapidity region, Phys. Rev. D 27, 140 (1983).
- [33] M. P. Heller, R. A. Janik, and P. Witaszczyk, The Characteristics of Thermalization of Boost-Invariant Plasma From Holography, Phys. Rev. Lett. 108, 201602 (2012).
- [34] See Supplemental Material at http://link.aps.org/supplemental/ 10.1103/PhysRevLett.128.122302 for a comment on the relation between the expansions in ε and 1/w in Bjorken flow, a discussion of our numerical computations and associated convergence tests, and further details on the DN model.
- [35] S. S. Gubser, Symmetry constraints on generalizations of Bjorken flow, Phys. Rev. D 82, 085027 (2010).
- [36] S. S. Gubser and A. Yarom, Conformal hydrodynamics in Minkowski and de Sitter spacetimes, Nucl. Phys. B846, 469 (2011).
- [37] M. P. Heller, A. Serantes, M. Spaliński, V. Svensson, and B. Withers, Hydrodynamic gradient expansion in linear response theory, Phys. Rev. D 104, 066002 (2021).
- [38] B. Withers, Short-lived modes from hydrodynamic dispersion relations, J. High Energy Phys. 06 (2018) 059.
- [39] R. Baier, P. Romatschke, D. T. Son, A. O. Starinets, and M. A. Stephanov, Relativistic viscous hydrodynamics, conformal invariance, and holography, J. High Energy Phys. 04 (2008) 100.
- [40] W. Israel, Nonstationary irreversible thermodynamics: a causal relativistic theory, Ann. Phys. (N.Y.) **100**, 310 (1976).
- [41] W. Israel and J. Stewart, Transient relativistic thermodynamics and kinetic theory, Ann. Phys. (N.Y.) 118, 341 (1979).
- [42] G. S. Denicol, H. Niemi, E. Molnar, and D. H. Rischke, Derivation of transient relativistic fluid dynamics from the Boltzmann equation, Phys. Rev. D **85**, 114047 (2012); Erratum, **91**, 039902 (2015).
- [43] A. Jaiswal, Relativistic third-order dissipative fluid dynamics from kinetic theory, Phys. Rev. C 88, 021903 (2013).
- [44] J. Noronha and G. S. Denicol, Transient fluid dynamics of the quark-gluon plasma according to AdS/CFT, arXiv:1104 .2415.
- [45] M. P. Heller, R. A. Janik, M. Spaliński, and P. Witaszczyk, Coupling Hydrodynamics to Nonequilibrium Degrees of Freedom in Strongly Interacting Quark-Gluon Plasma, Phys. Rev. Lett. 113, 261601 (2014).
- [46] R. Loganayagam, Entropy current in conformal hydrodynamics, J. High Energy Phys. 05 (2008) 087.
- [47] W. Florkowski, R. Ryblewski, M. Strickland, and L. Tinti, Non-boost-invariant dissipative hydrodynamics, Phys. Rev. C 94, 064903 (2016).
- [48] Note that longitudinal flows are in general distinct from comoving flows such as Bjorken and Gubser flow. While Bjorken flow can be found as a special case of a longitudinal flow, Gubser flow cannot since it features transverse dynamics as dictated by symmetries.