Multicomponent meson superfluids in chiral perturbation theory

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We show that multicomponent meson systems can be described by chiral perturbation theory. We chiefly focus on a system of two pion gases at different isospin chemical potential, deriving the general expression of the chiral Lagrangian, the ground state properties and the spectrum of the low-energy excitations. We consider two different kinds of interactions between the two meson gases: one which does not lock the two chiral symmetry groups and one which does lock them. The former is a kind of interaction that has already been discussed in multicomponent superfluids. The latter is perhaps more interesting, because it seems to be related to an instability. Although the pressure of the system does not show any instability, we find that for sufficiently strong locking, the spectrum of one Bogolyubov mode becomes tachyonic. This unstable branch seems to indicate a transition to an inhomogeneous phase.

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I. INTRODUCTION

Cold hadronic matter is an interesting playground for a deep understanding of the properties of the strong interaction. At asymptotic baryonic densities the liberated quarks [1] should pair forming a color superconductor, see [2–4] for reviews. At large isospin densities a different kind of collective phenomenon happens, with mesons forming a Bose-Einstein condensate (BEC) [5-10]. In general, the matter density of the system is controlled by the baryonic chemical potential, μ_B , while the isospin chemical potential, μ_I , is associated to its degree of isospin asymmetry, e.g., indicating that the number of neutrons and protons are different. The properties of matter as a function of μ_I have been the subject of intensive investigation for a number of reasons. Systems with large isospin asymmetry exist in Nature; in particular neutron stars [11] are believed to be compact stellar objects with a large isospin asymmetry. Recently, the possible existence of pion stars has also been proposed [12–14]. Regarding the microscopic properties of matter, the inclusion of μ_I can lead to a better understanding of quantum chromodynamics (QCD) in a regime in which lattice QCD simulations are doable

[15–22]. Remarkably, the lattice QCD simulations of meson gases with vanishing baryonic density are not affected by the sign problem and can be implemented for not too high values of μ_I . These simulations are steadily improving [20–22], reaching increasingly precise results on the thermodynamic properties of the system and thus offering powerful tests for the methods and models developed for the effective description of the strong interaction.

Among the various proposed models, it is worth mentioning the Nambu-Jona Lasinio (NJL) model [23–32] and the quark-meson model [24,33–35], because they can be used in a wide range of values of μ_I . Although these models are useful tools for exploring the properties of hadronic matter, they are based on a number of parameters that have to be phenomenologically fixed. Thus, they lead to results which depend on the choice of these parameters and on the number of degrees of freedom (d.o.f.) used. A relevant limitation of these models is that the obtained results cannot be systematically improved because no expansion parameter can be identified.

A systematic analysis of hadronic matter can be obtained by effective field theories [36–38], which are based on an expansion in a control parameter. Here we focus on chiral perturbation theory (χ PT), which is an effective theory designed to describe the low-energy properties of QCD [39–43]. The χ PT Lagrangian is derived by the global symmetries of QCD, basically integrating out the highenergy part. The effect of the isospin chemical potential is conveniently included in covariant derivatives, see [40,41,44] for a general discussion. This allows us to

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describe the transition to the meson condensed phase in a straightforward way, see [45,46], and to easily compare with lattice QCD results [47]. As the above-mentioned models, the χ PT Lagrangian depends on some constants that have to be fixed in some way, however the obtained results can be systematically improved including higher orders in the γ PT expansion [41,48]. The thermodynamic and low-energy properties of mesons at nonvanishing μ_I have been studied using the χ PT Lagrangian in many different works [12,45–47,49]. In particular, in these works it has been confirmed that the pion condensed phase first discussed in [5–9] sets in at $\mu_I = m_{\pi}$, where m_{π} is the pion mass. The χ PT approach has also been used to study different gauge theories with isospin asymmetry, including 2 color QCD with different flavors [50-56]. In the present paper we extend the χPT Lagrangian to study the condensation of multicomponent meson systems, which is relevant for the core of compact stars where pions and kaons, see e.g., [57], can condense.

The multicomponent condensation can be realized, and it has been studied, in a variety of systems. In some cases the two components are neutral, and in this case one realizes a multicomponent superfluid such as in He³–He⁴ mixtures, see [58,59], or in other ultracold atoms systems [60–64]. In the compact star interior, neutrons and protons are believed to simultaneously condense [11], forming a system which is simultaneously superfluid and superconductor where global and local symmetries are spontaneously broken. If deconfined quark matter is formed, the color-flavor locked (CFL) phase [65] could be realized, which is simultaneously superfluid and superconductor. When supplemented with kaon condensation [66,67], the CFL phase has an additional superfluid component, becoming a phase with two superfluid modes.

In our work we examine a multicomponent meson system in which each component is characterized by a global symmetry. For each component, the spontaneous breaking of a global symmetry leads to the formation of a superfluid. We examine the effect of the possible intraspecies interactions on multicomponent superfluidity. In particular, we focus on the meson condensed phases, which can be realized in the core of compact stars, see e.g., [11], employing the χ PT framework for deriving the relevant low-energy Lagrangian. We discuss a system in which the two meson gases correspond to two fictitious pion systems, paving the way for the discussion of the simultaneous condensation of pions and kaons. We identify two very different types of intraspecies interactions: those that lock the two global symmetries and those that do not lock them. The two considered interaction terms lead to two different symmetry breaking patterns, that could be dynamically generated by the strong dynamics underlying the χPT effective description.

Remarkably, at the leading order (LO) in χ PT, only the interactions that lock the two chiral groups are possible. We

find that the strength of the locking term plays a prominent role: Increasing the locking, the transition to the broken phase is favored; for sufficiently large couplings the system becomes unstable. Analyzing the dispersion laws of the low-energy d.o.f., we find that the instability can be interpreted as a transition toward an inhomogeneous phase.

At the next-to-leading order (NLO) χ PT corrections, it is possible to include those interactions that do not lock the two chiral groups. This type of interaction is akin to the one typically discussed in ultracold atoms systems and indeed in this case we obtain results similar to those of multicomponent Bose gas, see e.g., [68,69].

The present paper is organized as follows. In Sec. II we report known results for meson systems in χ PT. This is useful to fix the notation and for comparison with the multicomponent meson system. In Sec. III we generalize the χ PT Lagrangian to two meson gases, introducing the leading interaction terms. In Sec. IV we analyze the effect of one of the possible interaction terms leading to the chiral locking. In Sec. V we consider the χ PT term that does not lock the two chiral groups. We conclude in Sec. VI. A number of results are collected in the Appendixes. In the Appendix A, we report the low-energy excitations of a single-component pion gas. In the Appendix B, we discuss the low-energy corrections to the mean-field thermodynamic quantities arising from the vacuum energy of the Bogolyubov modes.

II. SINGLE MESON GAS

The χ PT description of the single meson gas is based on the global symmetries

$$G = SU(N_f)_L \times SU(N_f)_R,\tag{1}$$

of massless QCD, with N_f the number of flavors. The meson fields are collected in the Σ field, transforming under *G* as

$$\Sigma \to L\Sigma R^{\dagger},$$
 (2)

where $L \in SU(N_f)_L$ and $R \in SU(N_f)_R$. The leading $\mathcal{O}(p^2) \chi PT$ Lorentz-invariant Lagrangian [41,42,45] is given by

$$\mathcal{L} = \frac{f_{\pi}^2}{4} \operatorname{Tr}(D_{\nu} \Sigma D^{\nu} \Sigma^{\dagger}) + \operatorname{Tr}(M \Sigma^{\dagger} + M^{\dagger} \Sigma), \qquad (3)$$

where the mass matrix, M, and the so-called pion decay constant, f_{π} , are the low energy constants (LECs) that cannot be fixed by the symmetry group G and must be determined in some other way. The χ PT Lagrangian is constructed assuming that the mass term does not break the global symmetries, thus that M transforms as Σ . Then, the locking of the chiral rotations to the vector $SU(N_f)_V$ group is induced by the vev of M, see e.g., the discussion in [39,42]. The covariant derivative in Eq. (3) allows us to take into account the coupling of the meson fields with gauge fields and/or with external currents and/or the effect of different chemical potentials [40,41,44]. In the present work, we will only consider the effect of the isospin chemical potential and we will restrict the analysis to pions, corresponding to the $N_f = 2$ case. Thus, we consider the covariant derivative

$$D_{\nu}\Sigma = \partial_{\nu}\Sigma - \frac{i}{2}\delta_{\nu0}\mu_{I}[\sigma_{3},\Sigma], \qquad (4)$$

where the isospin chemical potential, μ_I , is introduced as the time component of a vector field. Note that the covariant derivative does not include the baryonic chemical potential, μ_B , because mesons do not have a baryonic charge. A useful parametrization is

$$\Sigma = \cos \rho + i \hat{\boldsymbol{\varphi}} \cdot \boldsymbol{\sigma} \sin \rho, \qquad (5)$$

where σ_i are the Pauli matrices. The radial field, ρ , and the unit vector field, $\hat{\varphi}$, encode in a nontrivial way the three pion fields. By this parametrization, the LO χ PT low-energy Lagrangian takes the form obtained in [12]

$$\mathcal{L} = \frac{f_{\pi}^2}{2} (\partial^{\mu} \rho \partial_{\mu} \rho + \sin^2 \rho \partial^{\mu} \hat{\varphi}_i \partial_{\mu} \hat{\varphi}_i - 2m_{\pi} \gamma \sin^2 \rho \epsilon_{3ik} \hat{\varphi}_i \partial_0 \hat{\varphi}_k) - V(\rho), \qquad (6)$$

where

$$V(\rho) = -f_{\pi}^2 m_{\pi}^2 \left(\cos\rho + \frac{\gamma^2}{2}\sin^2\rho\right),\tag{7}$$

is the potential and the control parameter is $\gamma = \mu_I/m_{\pi}$. For $|\gamma| < 1$, the normal phase is favored, the potential minimum is at $\bar{\Sigma} = I$, and

$$V = -f_{\pi}^2 m_{\pi}^2, \tag{8}$$

at the minimum. For $|\gamma| > 1$, the pion condensed phase is favored [5–10,45,49] and in the present parametrization the minimum of the potential is attained for the radial field vev, $\bar{\rho}$, satisfying

$$\cos\bar{\rho} = \frac{1}{\gamma^2}.\tag{9}$$

Therefore, in the broken phase the meson field vev is given by

$$\bar{\Sigma} = \cos\bar{\rho} + i\boldsymbol{n} \cdot \boldsymbol{\sigma} \sin\bar{\rho}, \qquad (10)$$

where n is a unit vector associated to the residual O(2) symmetry of the vacuum. The pressure and the isospin

number density in the broken phase are respectively given by [45,47,49]

$$P = \frac{f_{\pi}^2 m_{\pi}^2}{2} \gamma^2 \left(1 - \frac{1}{\gamma^2} \right)^2, \qquad n_I = f_{\pi}^2 m_{\pi} \gamma \left(1 - \frac{1}{\gamma^4} \right), \qquad (11)$$

leading to the $\mathcal{O}(p^2)$ equation of state [47]

$$\epsilon(P) = -P + 2\sqrt{P(2f_{\pi}^2 m_{\pi}^2 + P)}.$$
 (12)

Close to the phase transition point, $\gamma \gtrsim 1$, the system is dilute and it is possible to expand the pressure *P* and the energy density ϵ , as a function of the isospin number density n_I . If we define the adimensional isospin density $n = n_I/(f_{\pi}^2 m_{\pi})$, we can expand the control parameter as

$$\gamma = 1 + \frac{n}{4} + \frac{3n^2}{32} + \frac{n^3}{32} + \mathcal{O}(n^4), \tag{13}$$

which is a meaningful expansion for $n \ll 1$. The pressure can then be expanded as follows

$$P = \frac{n_I^2}{8f_\pi^2} + \frac{n_I^3}{16f_\pi^4 m_\pi^2} + \mathcal{O}(n_I^4), \qquad (14)$$

where the leading term is the mean-field expression of the pressure of a boson system with coupling $g_0 = 1/4f_{\pi}^2$. This is indeed the correct expression of the coupling close to the phase transition, see Eq. (B2) and the discussion in the Appendix A. The energy density is instead given by

$$\epsilon = m_{\pi} n_{I} + \frac{g_{0} n_{I}^{2}}{2} + \frac{g_{0}^{2} n_{I}^{3}}{2m_{\pi}} + \mathcal{O}(n_{I}^{4}), \qquad (15)$$

which takes into account the energy associated to the mass of the pions. Note that the above expressions are obtained in the mean-field approximation, meaning that the lowenergy fluctuations are not included. Indeed, the order n_I^3 corrections are determined by the χ PT Lagrangian and not by the contribution of the Bogolyubov modes. The vacuum contribution of the Bogolyubov modes is considered in the Appendix B, and is much smaller than the leading meanfield contribution. However, it can play an important role in a multicomponent gas, as we will see below.

III. SYSTEM OF TWO MESON GASES

We now generalize the discussion of the previous section to a system with two mesonic gases, considering the simultaneous effect of the chemical potentials and of the interaction terms. In the second quantization formalism we assume that two meson systems with densities n_1 and n_2 are described by the fields Σ_1 and Σ_2 . As for the single meson gas, we use the global symmetries for constructing the χ PT Lagrangian. As a starting point we consider the noninteracting case with symmetry group

$$G = G_1 \times G_2, \tag{16}$$

where

$$G_a = \{SU(N_f)_L \times SU(N_f)_R\}_a \quad \text{with} \quad a = 1, 2 \qquad (17)$$

is the chiral group of the Σ_a field. For simplicity, we will mainly treat the $N_f = 2$ case, in which the two meson gases correspond to two fictitious pion systems, paving the way for the discussion of the simultaneous condensation of pions and kaons. For a system of kaons and pions one has to consider the $N_f = 3$ case, thus including more complicated interaction terms than the ones discussed below for the $N_f = 2$ case.

In the present framework, we assume that in the noninteracting case the fields Σ_1 and Σ_2 transform independently under two chiral groups as

$$\Sigma_1 \to L_1 \Sigma_1 R_1^{\dagger}$$
 and $\Sigma_2 \to L_2 \Sigma_2 R_2^{\dagger}$, (18)

where $L_a \in SU(2)_{L,a}$ and $R_a \in SU(2)_{R,a}$ with a = 1, 2. The most general $\mathcal{O}(p^2)$ chiral Lagrangian invariant under these symmetries is

$$\mathcal{L} = \frac{f_{1\pi}^2}{4} \operatorname{Tr}(D_{\nu}^1 \Sigma_1 D^{1\nu} \Sigma_1^{\dagger}) + \frac{f_{2\pi}^2}{4} \operatorname{Tr}(D_{\nu}^2 \Sigma_2 D^{2\nu} \Sigma_2^{\dagger}) + \operatorname{Tr}(\Sigma_1 M_1^{\dagger} + M_1 \Sigma_1^{\dagger}) + \operatorname{Tr}(\Sigma_2 M_2^{\dagger} + M_2 \Sigma_2^{\dagger}), \qquad (19)$$

where $f_{1\pi}$ and $f_{2\pi}$, as well as the matrices M_1 and M_2 , are the low energy constants (LECs) of the system. As for a single meson system described by the Lagrangian in Eq. (3), we have constructed this Lagrangian assuming that the mass terms do not break the global symmetries, which means that M_a transforms as Σ_a . The covariant derivative D_{ν}^a takes into account the interaction of the mesons of the *a* system with the external fields. If the two meson systems have different isospin chemical potentials, μ_1 and μ_2 , respectively, this can be encoded in the two covariant derivatives

$$D_a^{\nu} \Sigma_a = \partial^{\nu} \Sigma_a - \frac{i}{2} \delta_{\nu 0} \mu_a [\sigma_3, \Sigma_a], \qquad (20)$$

for a = 1, 2.

We now introduce the interaction between the two gases. Before doing that, let us first recall that under G_a the covariant derivative transforms as the Σ_a fields, that is

$$D_a^{\mu} \Sigma_a \to L_a D_a^{\mu} \Sigma_a R_a^{\dagger}, \qquad (21)$$

and therefore the two covariant derivatives are independently rotated. Let us now consider the possible intraspecies interaction terms. If we add to the noninteracting Lagrangian the term

$$\mathcal{L}_{\text{int},1} = k \frac{f_{1\pi} f_{2\pi}}{2} \operatorname{Tr}(D^{1}_{\nu} \Sigma_{1} D^{2\nu} \Sigma_{2}^{\dagger}), \qquad (22)$$

it locks the two chiral groups, leaving only the group

$$G_D = SU(2)_L \times SU(2)_R,\tag{23}$$

unbroken, corresponding to the simultaneous chiral rotations of the two systems. Under this group, the two fields Σ_1 and Σ_2 transform as in Eq. (18), but with $L_1 = L_2$ and $R_1 = R_2$. In principle, the *k* coefficient is a number that depends on the interaction strength between the two chiral fields and, as any LEC, it is independent of the isospin chemical potentials.

Remarkably, the interaction Lagrangian in Eq. (22) is the only $\mathcal{O}(p^2)$ meaningful coupling leaving the G_D group unbroken. One may think to add a Lagrangian term of the type

$$\operatorname{Tr}(\Sigma_1 \Sigma_2^{\dagger}),$$
 (24)

which indeed locks chiral rotations. However, if one of the two fields vanishes, from Eq. (5) we have that, say, $\Sigma_1 \equiv I$. Then the term in Eq. (24) acts as a mass term for the Σ_2 field, breaking G_2 down to the vector subgroup. Therefore, this kind of term or any term of the type

$$\operatorname{Tr}(\Sigma_1 \Sigma_2^{\dagger})^n,$$
 (25)

with n > 0 is not allowed. For a similar reason the masslike terms

$$\operatorname{Tr}(M_1 \Sigma_1^{\dagger} (\Sigma_2 \Sigma_1^{\dagger})^n), \qquad (26)$$

are not allowed, unless n = 0.

If one wants to preserve the *G* group defined in (16), then one has to consider the $\mathcal{O}(p^4)$ terms. At this order, there are only two derivative intraspecies interaction terms that do not lock the two chiral groups:

$$\mathcal{L}_{\text{int},2} = \tilde{L}_1 \text{Tr}(D^1_{\mu} \Sigma_1 D^{1\mu} \Sigma_1^{\dagger}) \text{Tr}(D^2_{\nu} \Sigma_2 D^{2\nu} \Sigma_2^{\dagger}) + \tilde{L}_2 \text{Tr}(D^1_{\mu} \Sigma_1 D^{1\nu} \Sigma_1^{\dagger}) \text{Tr}(D^2_{\mu} \Sigma_2 D^{2\nu} \Sigma_2^{\dagger}), \qquad (27)$$

where \tilde{L}_1 and \tilde{L}_2 are two LECs analogous to the standard L_1 and L_2 of $\mathcal{O}(p^4) \chi PT$ [42]. When including these contributions, one should consistently include the standard $\mathcal{O}(p^4)$ chiral terms, as well. However, as was shown in [12], the effect of the standard NLO terms on the thermodynamic properties of the system is extremely small and can be accounted for by a renormalization of the LO LECs.

As an aside, we note that in principle one may consider more complicated intraspecies interaction terms, like

$$\mathcal{L}_{\text{int}} \propto k^{\mu\nu} \text{Tr}(D^1_{\mu} \Sigma_1 D^2_{\nu} \Sigma_2^{\dagger}), \qquad (28)$$

with $k^{\mu\nu}$ a Lorentz tensor and a *G* singlet. This kind of interaction term somehow generalizes Eq. (22), however it is not obvious how to fix the values of the $k^{\mu\nu}$ components in general.

In the following, we will discuss the effect of the interaction terms in Eq. (22) and in Eq. (27), separately, focusing on the pion system.

IV. CHIRAL LOCKING

To gain insight on the system described by Eqs. (19) and (22), let us first assume that we are making a partition of an ensemble of undistinguishable pions, dividing it in two (interacting) subsets, in such a way that the Σ_1 field describes the pions of the first subset and Σ_2 field the pions of the second subset. Let us first focus on the kinetic terms at vanishing isospin chemical potentials. Since the pions are indistinguishable, one may naively think that the most general $O(p^2)$ Lagrangian is

$$\mathcal{L} = \frac{f_{\pi}^2}{4} \operatorname{Tr}(\partial_{\nu} \Sigma_1 \partial^{\nu} \Sigma_1^{\dagger}) + \frac{f_{\pi}^2}{4} \operatorname{Tr}(\partial_{\nu} \Sigma_2 \partial^{\nu} \Sigma_2^{\dagger}) + k \frac{f_{\pi}^2}{2} \operatorname{Tr}(\partial_{\nu} \Sigma_1 \partial^{\nu} \Sigma_2^{\dagger}), \qquad (29)$$

where the first term, respectively the second term, describes the propagation and self-interactions of the fields of the subset 1, respectively 2. The third term mixes the two fields and induces the locking between the two subsets. If it were absent, that is for k = 0, there would be no interactions between the two sets.

For subsets made of identical particles there must exist a way of reshuffling them. Since $\Sigma_1 \Sigma_1^{\dagger} + \Sigma_2 \Sigma_2^{\dagger} = 2$, any reshuffling can only correspond to a rotation

$$\begin{split} \Sigma_1 &\to \cos\theta \hat{\Sigma}_1 + \sin\theta \hat{\Sigma}_2, \\ \Sigma_2 &\to -\sin\theta \hat{\Sigma}_1 + \cos\theta \hat{\Sigma}_2, \end{split} \tag{30}$$

transforming the Lagrangian in Eq. (29) in

$$\mathcal{L} = +\frac{f_{\pi}^{2}}{4} (1 - k\sin(2\theta)) \operatorname{Tr}(\partial_{\nu}\Sigma_{1}\partial^{\nu}\Sigma_{1}^{\dagger}) + \frac{f_{\pi}^{2}}{4} (1 + k\sin(2\theta)) \operatorname{Tr}(\partial_{\nu}\Sigma_{2}\partial^{\nu}\Sigma_{2}^{\dagger}) + k\frac{f_{\pi}^{2}}{2} \cos(2\theta) \operatorname{Tr}(\partial_{\nu}\Sigma_{1}\partial^{\nu}\Sigma_{2}^{\dagger}).$$
(31)

To maintain the Lagrangian invariant we have to take k = 0or, more interestingly, k = 1. Indeed, in the latter case

$$\mathcal{L} = \frac{\hat{f}_{1\pi}^2}{4} \operatorname{Tr}(\partial_{\nu}\hat{\Sigma}_1 \partial^{\nu}\hat{\Sigma}_1^{\dagger}) + \frac{\hat{f}_{2\pi}^2}{4} \operatorname{Tr}(\partial_{\nu}\hat{\Sigma}_2 \partial^{\nu}\hat{\Sigma}_2^{\dagger}) + \frac{\hat{f}_{1\pi}\hat{f}_{2\pi}}{2} \operatorname{Tr}(\partial_{\nu}\hat{\Sigma}_1 \partial^{\nu}\hat{\Sigma}_2^{\dagger}),$$
(32)

where $\hat{f}_{1\pi}^2 = f_{\pi}^2(1 - \sin 2\theta)$, $\hat{f}_{2\pi}^2 = f_{\pi}^2(1 + \sin 2\theta)$, and therefore $\hat{f}_{1\pi}\hat{f}_{2\pi} = f_{\pi}^2\cos 2\theta$. Note that one cannot identify $\hat{f}_{a\pi}$ with the pion decay constant of the pions in the subset *a*, because the fields are mixed by the locking interaction term.

For nonvanishing locking, if one takes $k \neq 1$, the O(2) symmetry in Eq. (30) does not hold and the coefficient of the interaction term cannot be expressed as $\hat{f}_{1\pi}\hat{f}_{2\pi}$, meaning that if one makes the rotation, this term would depend on the rotation angle. On the other hand, in the Lagrangian in Eq. (32) it is possible to eliminate the dependence on the unphysical angle θ in the quadratic terms by writing

$$\hat{\Sigma}_a = e^{i\boldsymbol{\sigma}\cdot\hat{\boldsymbol{\varphi}}_a/\hat{f}_{a\pi}},\tag{33}$$

which is a generalization of the standard nonlinear expression of the pion fields. Therefore, the expression in Eq. (32), where k = 1 is set, is the most general χ PT Lagrangian for two gases of undistinguishable pions. We can easily generalize it to N undistinguishable pion gases, writing

$$\mathcal{L} = \sum_{ab} \frac{f_a f_b}{4} \operatorname{Tr}(\partial_{\nu} \Sigma_a \partial^{\nu} \Sigma_b^{\dagger}), \qquad (34)$$

where a = 1, ..., N and f_a are a generalization of the pion decay constant. Note that the propagating d.o.f. are obtained by diagonalizing the quadratic Lagrangian.

We turn now to a system of distinguishable pions. Including the mass terms, formally considering the vevs of the fields M_a in Eq. (19), we can write the total Lagrangian of the system as follows

$$\mathcal{L} = \frac{f_{1\pi}^2}{4} \operatorname{Tr}(\partial_{\nu} \Sigma_1 \partial^{\nu} \Sigma_1^{\dagger}) + \frac{f_{2\pi}^2}{4} \operatorname{Tr}(\partial_{\nu} \Sigma_2 \partial^{\nu} \Sigma_2^{\dagger}) + \frac{f_{1\pi}^2 m_{1\pi}^2}{4} \operatorname{Tr}(\Sigma_1 + \Sigma_1^{\dagger}) + \frac{f_{2\pi}^2 m_{2\pi}^2}{4} \operatorname{Tr}(\Sigma_2 + \Sigma_2^{\dagger}) + k \frac{f_{1\pi} f_{2\pi}}{2} \operatorname{Tr}(\partial_{\nu} \Sigma_1 \partial^{\nu} \Sigma_2^{\dagger}),$$
(35)

where we have assumed that the two fields have different mass parameters, $m_{1\pi}$ and $m_{2\pi}$. These parameters have to be interpreted as the LECs for the coupled system and correspond to the pion masses only in the k = 0 case. The actual masses can be obtained by the dispersion laws

$$E_{\pm}^2 = p^2 + M_{\pm}^2, \tag{36}$$

where the masses, obtained diagonalizing the quadratic part in Eq. (35), are given by

$$M_{\pm}^{2} = \frac{m_{1\pi}^{2} + m_{2\pi}^{2} \pm \sqrt{(m_{1\pi}^{2} - m_{2\pi}^{2})^{2} + 4k^{2}m_{1\pi}^{2}m_{2\pi}^{2}}}{2(1-k^{2})} \quad (37)$$

for $k \neq 1$, and equal to the "reduced mass"

$$M^2 = \frac{m_{1\pi}^2 m_{2\pi}^2}{m_{1\pi}^2 + m_{2\pi}^2},$$
(38)

for k = 1. From the above expressions it is clear that the interaction term in Eq. (22) induces a mass splitting. For clarity we report the behavior of the meson masses as a function of k in Fig. 1.

We remind that k = 0 corresponds to two noninteracting gases, while k = 1 corresponds to a system with only one low-energy mode. For k < 1, the mass splitting induced by the locking term is similar to the one induced by μ_I between the charged pions, see e.g., [46]. However, the system is unstable for k > 1. The instability is signaled by the divergent mass of one mode as $k \to 1^-$, which becomes imaginary for k > 1. In the context of ultracold atoms physics, where boson condensates are mostly considered, the latter feature is generally related to the appearance of spatially inhomogeneous phases, see e.g., [70,71] and references therein. We stress, however, that here we are in the presence of an instability with a completely different origin. Indeed, in ultracold atoms, the instability is triggered by a sufficiently large density-density coupling between the two systems [70,72,73] (a similar phenomenon is known also for fermions, called Stoner instability, see e.g., [74]). Instead, in the present case, the instability is due to a quadratic symmetry locking term in the fields: it is precisely the locking that rules the instability. Since k is the coefficient of a derivative term in Eq. (35), Remarkably, as k varies, the repulsion between the two systems remains



FIG. 1. Mass splitting induced by the locking term in Eq. (22) for the two gases of pions. For simplicity we have assumed that the pions of the two gases have equal mass parameters, i.e., $m_{1\pi} = m_{2\pi}$. The *k* parameter indicates the strength of the intraspecies locking, see Eq. (22); k = 0 corresponds to two non-interacting gases, while k = 1 to a system with one single type low-energy d.o.f. For k > 1 the system is unstable.

fixed and the total potential energy density is given by and reads

$$V = -\sum_{a} f_{a\pi}^2 m_{a\pi}^2, \tag{39}$$

corresponding to the sum of two independent contributions, see Eq. (8), that can be obtained from the nonderivative terms of Eq. (35). In spite of this relevant difference and considering that the locked theory in Eq. (35) is quadratic, it is still quite natural to postulate that the same theory with k > 1 cannot exist with the two species coexisting in the same space domain.

To elucidate the mechanism underlying the locking instability, and its possible resolution, let us consider a simple system consisting of two scalar bosons with a locking term

$$\mathcal{L} = \mathcal{L}_{1} + \mathcal{L}_{2} + \mathcal{L}_{int}$$

$$= \frac{1}{2} \partial_{\mu} \phi_{1} \partial^{\mu} \phi_{1} - \frac{1}{2} m_{1}^{2} \phi_{1}^{2} + \frac{1}{2} \partial_{\mu} \phi_{2} \partial^{\mu} \phi_{2}$$

$$- \frac{1}{2} m_{2}^{2} \phi_{2}^{2} + k \partial_{\mu} \phi_{1} \partial^{\mu} \phi_{2}, \qquad (40)$$

with a manifest discrete $Z_2 \times Z_2$ symmetry for k = 0. This symmetry corresponds to the transformations $\phi_1 \rightarrow -\phi_1$ and $\phi_2 \rightarrow -\phi_2$, separately. For $k \neq 0$ the two discrete symmetries are locked, with the only remaining Z_2 symmetry corresponding to $\phi_1 \rightarrow -\phi_1$ and $\phi_2 \rightarrow -\phi_2$, simultaneously.

This simple system becomes unstable for k > 1, because one of the two eigenmodes has an imaginary mass. One possible solution of the instability corresponds to the realization of an inhomogeneous phase. Let us give an heuristic argument in favor of the inhomogeneous phase. If we assume that one component is realized in the volume V_1 and the other in the volume V_2 , then the action can be written as

$$S = \int d^4 x \mathcal{L} \approx \int_{V_1} d^4 x \mathcal{L}_1 + \int_{V_2} d^4 x \mathcal{L}_2 + \int_{S_{12}} d^4 x \mathcal{L}$$

= $S_1 + S_2 + S_{\text{interface}},$ (41)

where S_a with a = 1, 2 are the actions of the free scalar fields. The effect of the interaction term is only relevant at the interface, S_{12} , of the two volumes. In other words, in the inhomogeneous phase the interaction Lagrangian \mathcal{L}_{int} has only support at the interface and therefore the dispersion laws of the field ϕ_1 , respectively ϕ_2 , in the volumes V_1 , respectively V_2 , are not tachyonic.

A. Two pion gases at different isospin chemical potentials

We now consider the effect of the isospin chemical potentials for the two pion gases. Including them, the Lagrangian reads

$$\mathcal{L} = +\frac{f_{1\pi}^2}{4} \operatorname{Tr}(D_{\nu}^1 \Sigma_1 D^{1\nu} \Sigma_1^{\dagger}) + \frac{f_{2\pi}^2}{4} \operatorname{Tr}(D_{\nu}^2 \Sigma_2 D^{2\nu} \Sigma_2^{\dagger}) + \frac{f_{1\pi}^2 m_{1\pi}^2}{4} \operatorname{Tr}(\Sigma_1 + \Sigma_1^{\dagger}) + \frac{f_{1\pi}^2 m_{1\pi}^2}{4} \operatorname{Tr}(\Sigma_2 + \Sigma_2^{\dagger}) + k \frac{f_{1\pi} f_{2\pi}}{2} \operatorname{Tr}(D_{\nu}^1 \Sigma_1 D^{2\nu} \Sigma_2^{\dagger}),$$
(42)

where the covariant derivatives are given in Eq. (20).

Since the two fields can have different vevs, we generalize Eq. (10) to

$$\Sigma_a = \cos \rho_a + i \boldsymbol{n}_a \cdot \boldsymbol{\sigma} \sin \rho_i \qquad a = 1, 2, \qquad (43)$$

where ρ_a are the two radial fields and n_a are two unit vectors. Upon substituting Eq. (43) in Eq. (42), we obtain the tree-level potential

$$V = -\sum_{a} f_{a\pi}^{2} m_{a\pi}^{2} \left(\cos \rho_{a} + \frac{\gamma_{a}^{2}}{2} \sin^{2} \rho_{a} \right) - k \mathbf{n}_{1} \cdot \mathbf{n}_{2} f_{1\pi} f_{2\pi} \mu_{1} \mu_{2} \sin \rho_{1} \sin \rho_{2}, \qquad (44)$$

where $\gamma_a = \mu_a / m_{a\pi}$ and the last term on the right-hand side originates from the locking term, which explicitly breaks the G symmetry to the diagonal group, G_D . The interesting aspect is that the potential depends on the relative angle between n_1 and n_2 . In the ground state the two unit vectors are locked to be aligned, if the isospin chemical potentials have equal signs, or antialigned, if the isospin chemical potentials have opposite signs. We can clearly restrict the analysis to the case in which both isospin chemical potentials are positive and aligned. Since the vevs of the two fields are not independent but tend to align, it is clear that the condensation of one field favors the condensation of the other; we will discuss this effect in detail below. From the above expression it is also clear that the system has two Nambu Goldstone bosons (NGBs) for k = 0, corresponding to the two independent oscillations of the unit vectors, but only one NGB for $k \neq 0$, corresponding to the locked oscillations of the two fields. The second mode is massive and corresponds to a pseudo-NGB.

B. Phase diagram of the locked pion gases

At the transition to the broken phase, where both gases condense, we can expand

 $\cos \bar{\rho}_1 = 1 - \epsilon_1$ and $\cos \bar{\rho}_1 = 1 - \epsilon_2$, (45)

with $\epsilon_a \ll 1$. Upon replacing this expression in the stationary condition for the potential, we obtain

$$\sqrt{\frac{\epsilon_2}{\epsilon_1}} = \frac{1 - \gamma_1^2}{k\gamma_1\gamma_2}$$
 and $\sqrt{\frac{\epsilon_1}{\epsilon_2}} = \frac{1 - \gamma_2^2}{k\gamma_1\gamma_2}$, (46)

signaling that the condensation of one gas is deeply related to the condensation of the other: as soon as, say, $\epsilon_1 > 0$, it follows that $\epsilon_2 > 0$. The formation of one superfluid necessarily makes the other gas superfluid by a simultaneous condensation (SCO) mechanism.

Upon solving the above system of equations, we easily obtain that the SCO happens for

$$(k^2 - 1)\gamma_1^2\gamma_2^2 + \gamma_1^2 + \gamma_2^2 = 1, \qquad (47)$$

corresponding to the curve, C on the (γ_1, γ_2) plane depicted in Fig. 2 for various values of k. The existence of this curve makes explicit that the onset of one condensate induces the condensation of the other, a manifestation of the interaction between the two. A remarkable aspect is that the SCO happens for any nonvanishing value of k. Clearly, the larger is k, the larger is the effect of one condensate on the other. Moreover, with increasing values of k, the normal phase region shrinks. To better understand this process, let us focus on the $\gamma_1 = \gamma_2 = \gamma$ case. Since the two isospin chemical potentials are equal, it follows has that $\bar{\rho}_1 = \bar{\rho}_2 = \bar{\rho}$,

$$\cos\bar{\rho} = \frac{1}{\gamma^2(1+k)},\tag{48}$$

and the transition happens for $\gamma^2 = 1/(k+1)$. Therefore, with increasing values of *k*, the transition to the SCO phase happens at lower values of γ . One may naively think that increasing *k* would lead to a system that becomes



FIG. 2. Phase diagram for the coupled superfluid system with the intraspecies interaction of Eq. (22). The solid black line corresponds to $k = 10^{-3}$; the dashed red line corresponds to k = 1 and the dotted blue line corresponds to k = 5. For every considered value of k, the broken phase is the region outside the corresponding curve. In that region, indicated with SCO, there is the simultaneous condensation of both fluids. Along the axes, where $\gamma_1 = 0$ and $\gamma_2 > 1$ or $\gamma_2 = 0$ and $\gamma_1 > 1$, only one component is superfluid. The analysis of the low-energy excitations shows that for k > 1 one of the low-energy modes becomes tachyonic, meaning that in this case the mean-field results reported in this figure are not valid.

superfluid for arbitrary values of the isospin chemical potential. As we will see below, this is not the case, because for k > 1 an instability in the low-energy spectrum is triggered.

In general, close to the transition curve, C, one can expand the pressure as

$$P = \frac{1}{2}L_{11}(\gamma_1 - \bar{\gamma}_1)^2 + \frac{1}{2}L_{22}(\gamma_2 - \bar{\gamma}_2)^2 + L_{12}(\gamma_1 - \bar{\gamma}_1)(\gamma_2 - \bar{\gamma}_2),$$
(49)

where $\bar{\gamma}_a \in \mathcal{C}$ and

$$L_{ab} = \frac{\partial^2 P}{\partial \mu_a \partial \mu_b} \bigg|_{\mathcal{C}},\tag{50}$$

are the susceptibilities. Upon expressing the isospin chemical potentials in terms of the number densities, we obtain

$$P = \sum_{ij} \frac{g_{ij}}{2} n_i n_j, \tag{51}$$

where the coupling constants are given by

$$g_{11} = \frac{L_{11}}{D}$$
 $g_{22} = \frac{L_{22}}{D}$ $g_{12} = -\frac{L_{12}}{D}$, (52)

where $D = L_{11}L_{22} - L_{12}^2$, with $L_{ab} > 0$. It turns out that

$$g_{12} + \sqrt{g_{11}g_{22}} \ge 0, \tag{53}$$

the equality corresponding to the case $\gamma_1 = \gamma_2 = 1/\sqrt{2}$. For nonrelativistic distinguishable and dilute superfluid bosons, the equality in Eq. (53) corresponds to the stability threshold against collapse or turn into an inhomogeneous phase (depending on the sign of g_{12}) [70,72,73]. By a similar reasoning, one could argue that, because of the relation in Eq. (53), the locked system at nonvanishing isospin density is stable. Let us discuss this important point more in detail. The condition in Eq. (53) relies on the meanfield approximation. In condensed matter systems it is known that when the relation in Eq. (53) is violated the inclusion of the vacuum energy contribution of the Bogolyubov modes can turn a collapsing system into an inhomogeneous one, made of droplets of coexisting gases [73]. In the present case the condition in Eq. (53) is not violated: the mean-field pressure is well defined, and the system could be expected to be homogeneous and stable. However, for k > 1, we found that in the normal phase there exists a tachyonic mode. It is therefore important to analyze the low-energy spectrum of the system to figure out what is the fate of the tachyonic mode in the SCO phase.

C. Low-energy excitations

The low-energy excitations of the multicomponent system can be determined studying the fluctuations of the radial component and of the Bogolyubov modes of the meson fields. We shall employ the same formalism developed in [12] and briefly discussed in the Appendix A, extending it to the two-component gas of pions.

1. Radial excitations

In the broken phase, the system has two radial excitations, χ_1 and χ_2 , corresponding to the fluctuations around the corresponding vevs:

$$\rho_a = \bar{\rho}_a + \chi_a \quad \text{with} \quad a = 1, 2 \tag{54}$$

where it is assumed that $\chi_a \ll \bar{\rho}_a$. Upon substituting the above expression in Eq. (42) and restricting to the quadratic order in the fields, we obtain the Lagrangian

$$\mathcal{L}_{\chi} = \frac{1}{2} \partial_{\mu} \chi_{1} \partial^{\mu} \chi_{1} + \frac{1}{2} \partial_{\mu} \chi_{2} \partial^{\mu} \chi_{2} + c_{12} \partial_{\mu} \chi_{1} \partial^{\mu} \chi_{2} - \frac{M_{1}^{2}}{2} \chi_{1}^{2} - \frac{M_{2}^{2}}{2} \chi_{2}^{2} + M_{12} \chi_{1} \chi_{2},$$
(55)

where

$$c_{12} = \cos(\bar{\rho}_1 - \bar{\rho}_2) \qquad s_{12} = \sin(\bar{\rho}_1 - \bar{\rho}_2) M_1^2 = m_\pi^2(\cos\bar{\rho}_1 - \gamma_1^2\cos 2\bar{\rho}_1 + k\gamma_1\gamma_2\sin\bar{\rho}_1\sin\bar{\rho}_2) M_2^2 = m_\pi^2(\cos\bar{\rho}_2 - \gamma_2^2\cos 2\bar{\rho}_2 + k\gamma_1\gamma_2\sin\bar{\rho}_1\sin\bar{\rho}_2) M_{12} = km_\pi^2\gamma_1\gamma_2\cos\bar{\rho}_1\cos\bar{\rho}_2.$$
(56)

The corresponding dispersion laws are given by

$$E_{\pm} = \boldsymbol{p}^2 + \frac{c_{12}M_{12} + (M_1^2 + M_2^2)/2 \pm \sqrt{(c_{12}M_{12} + (M_1^2 + M_2^2)/2)^2 + s_{12}^2(M_{12}^2 - M_1^2M_2^2)}}{s_{12}^2},$$
(57)

thus the two modes have non-negative masses and are stable for any value of k. On the transition region to the BEC phase $M_{12} = M_1M_2$, and one of the radial modes becomes massless. The stability in the radial modes for any value of k is clearly a manifestation of the result obtained in the previous section, that the pressure close to the transition region is positive defined.

2. Bogolyubov modes

Neglecting the radial excitations, thus taking $\rho_1 \equiv \bar{\rho}_1$ and $\rho_2 \equiv \bar{\rho}_2$, one has the following low-energy Lagrangian

$$\mathcal{L}_{\hat{\varphi}} = \mathcal{L}_1 + \mathcal{L}_2 + \mathcal{L}_{12},\tag{58}$$

where

$$\mathcal{L}_1 = \frac{f_\pi^2}{2} (\sin^2 \bar{\rho}_1 \partial_\mu \hat{\boldsymbol{\varphi}}^1 \cdot \partial^\mu \hat{\boldsymbol{\varphi}}^1 + 2\mu_1^2 \sin^2 \bar{\rho}_i \epsilon_{3ik} \hat{\varphi}_i^1 \partial_0 \hat{\varphi}_k^1), \quad (59)$$

and with \mathcal{L}_2 given by a similar expression, while

$$\mathcal{L}_{12} = k f_{\pi}^{2} [\sin \bar{\rho}_{1} \sin \bar{\rho}_{2} (\partial_{\mu} \hat{\boldsymbol{\varphi}}^{1} \cdot \partial^{\mu} \hat{\boldsymbol{\varphi}}^{2} + \mu_{1} \mu_{2} \hat{\boldsymbol{\varphi}}^{1} \cdot \hat{\boldsymbol{\varphi}}^{2}) + \epsilon_{3ik} (\sin \bar{\rho}_{1} \mu_{2} \hat{\boldsymbol{\varphi}}_{i}^{2} \partial_{0} \hat{\boldsymbol{\varphi}}_{k}^{1} + \sin \bar{\rho}_{2} \mu_{1} \hat{\boldsymbol{\varphi}}_{i}^{1} \partial_{0} \hat{\boldsymbol{\varphi}}_{k}^{2})], \quad (60)$$

stems from the locking term. The unit vectors fields $\hat{\varphi}_1$ and $\hat{\varphi}_2$ describe the two angular fluctuations of the condensates and can be parametrized as follows

$$\hat{\boldsymbol{\varphi}}^1 = (\cos \alpha, \sin \alpha) \quad \text{and} \quad \hat{\boldsymbol{\varphi}}^2 = (\cos \theta, \sin \theta), \quad (61)$$

which generalize the expression in Eq. (A7). Upon substituting the above expression in the low-energy Lagrangian, we obtain

$$\mathcal{L}_{\hat{\varphi}} = \frac{f_{\pi}^2}{2} [\sin^2 \bar{\rho}_1 \partial_{\mu} \alpha \partial^{\mu} \alpha + \sin^2 \bar{\rho}_2 \partial_{\mu} \theta \partial^{\mu} \theta + 2k \sin \bar{\rho}_1 \sin \bar{\rho}_2 \cos(\alpha - \theta) (\partial_{\mu} \alpha \partial^{\mu} \theta + \mu_1 \mu_2)], \quad (62)$$

where we have not included the terms

$$f_{\pi}^{2}\cos(\alpha-\theta)(\mu_{2}\sin\bar{\rho}_{1}\partial_{0}\alpha+\mu_{1}\sin\bar{\rho}_{2}\partial_{0}\theta) \quad (63)$$

and

$$f_{\pi}^{2}(\mu_{1}\sin^{2}\bar{\rho}_{1}\partial_{0}\alpha + \mu_{2}\sin^{2}\bar{\rho}_{2}\partial_{0}\theta), \qquad (64)$$

leading to interactions and total derivatives. The Lagrangian in Eq. (62) describes two coupled modes. We restrict to the case $\mu_1\mu_2 > 0$; the other case can be treated in a similar way. The potential term is minimized for $\alpha = \theta$, thus expanding in $(\alpha - \theta)$ and keeping only the quadratic terms, we obtain the dispersion laws

$$E_1^2 = p^2$$

$$E_2^2 = p^2 + \mu_1 \mu_2 \frac{k(\sin^2 \bar{\rho}_1 + \sin^2 \bar{\rho}_2) + 2k^2 \sin \bar{\rho}_1 \sin \bar{\rho}_2}{1 - k^2}, \quad (65)$$

corresponding to the massless NGB and the massive pseudo-NGB, respectively. The propagation velocity of the NGB is equal to 1, however integrating out the radial oscillations would lead to a propagation velocity equal to the speed of sound, see [12] and the discussion in the Appendix A. For k = 0 the mass of the pseudo NGB vanishes and thus the system has two NGBs describing the independent fluctuations of the two decoupled superfluids.

We notice that for $k \to 1^-$ the mass of the pseudo-NGB diverges and only one low-energy mode exists, which is consistent with the fact that for k = 1 the system is equivalent to a single superfluid. For k > 1 the mass of the pseudo-NGB becomes imaginary, signaling an instability. This is the same instability we previously discussed in Fig. 1 in the unbroken phase. Thus, the unstable modes is still present in the SCO phase, now appearing as a pseudo-NGB with a tachyonic mass. The presence of this mode indicates that the mean-field approximation breaks down. Therefore, the expression of the pressure in Eq. (51) is incorrect for k > 1. This result is discussed in more detail in the Appendix B, where it is shown that the beyond meanfield contributions are ill defined for k > 1.

V. INDEPENDENT CHIRAL ROTATIONS

We now consider the interaction terms that do not lock the two chiral groups. Upon expanding the Lagrangian given by Eqs. (19) and (27), we obtain the potential

$$V = -\sum_{a} f_{a\pi}^{2} m_{a\pi}^{2} \left(\cos \rho_{a} + \frac{\gamma_{a}^{2}}{2} \sin^{2} \rho_{a} \right) - 4 (\tilde{L}_{1} + \tilde{L}_{2}) \mu_{1}^{2} \mu_{2}^{2} \sin^{2} \rho_{1} \sin^{2} \rho_{2},$$
(66)

where we have assumed the two gases have unequal masses and decay constant parameters. Unlike the locked case in the previous sections, now the tree-level potential is independent of the relative orientation of the two condensates, indeed it does not depend on $n_1 \cdot n_2$. In other words, the potential does not break the degeneracy of the two vacua and the two condensates vectors \mathbf{n}_1 and \mathbf{n}_2 can independently rotate. This is a manifestation of the fact that the interaction term does not lock the two chiral groups and thus the system has two NGBs. Considering $\tilde{L}_1 + \tilde{L}_2 \sim 10^{-3}$, as typical for $\mathcal{O}(p^4)$ corrections (see e.g., [42]) the interaction term has a small impact on the favored ground state. In particular, the onset of the simultaneous condensation is for $\gamma_1 \gtrsim 1$ and $\gamma_2 \gtrsim 1$. In the following we will consider $|\tilde{L}_1 + \tilde{L}_2| = 10^{-2} - 10^{-3}$, also taking into account possible negative values of $(\tilde{L}_1 + \tilde{L}_2).$

In Fig. 3, we report the phase diagrams obtained with positive (left panel) and negative (right panel) values of $\tilde{L}_1 + \tilde{L}_2$. The behavior with the strength of the intraspecies interaction is very similar to the one obtained for a coupled two-fluid system in [68]. The $\tilde{L}_1 + \tilde{L}_2$ parameter has the same effect on the phase diagram of the entrainment parameter of [68]: a positive value of $\tilde{L}_1 + \tilde{L}_2$ favors the SCO, while a negative value disfavors it. In [68] it was also discussed the instability generated by coupled superfluid flows. Although a similar phenomenon might emerge in



FIG. 3. Phase diagram for the coupled superfluid system with the intraspecies interaction of Eq. (27). Left panel: case with $\tilde{L}_1 + \tilde{L}_2 = +10^{-2}$. Right panel: $\tilde{L}_1 + \tilde{L}_2 = -10^{-2}$. In both panels the phase with simultaneous condensation is indicated with SCO. The phases indicated with SF₁ and SF₂, correspond to the phases in which only one component is superfluid. Positive values of $\tilde{L}_1 + \tilde{L}_2$ favor the simultaneous condensation.

our model, we postpone its analysis to future work. In order to generate such an instability in our model, it would presumably require NLO χ PT effects comparable with the LO results, thus in a regime where the χ PT expansion is not under quantitative control. Similar results to the one reported in Fig. 3 have also been reported in [69] by changing the strength of the density-density interaction between two scalar fields. In general, we believe that a phase diagram akin to the one shown in Fig. 3 can be obtained for any system with two condensates by changing the density-density interaction between the two fluids.

In order to infer the effect of one superfluid on the other, we consider the case in which one of the two superfluids is formed, say the superfluid 2, and we seek the critical value $\gamma_{1,c}$ for the onset of the condensation of the superfluid 1. At the leading order in the intra-species interaction, we find that the condensation onset for the first species obeys the equation

$$\gamma_{1,c}^2 = 1 - 8(\tilde{L}_1 + \tilde{L}_2) \frac{\gamma_2^4 - 1}{\gamma_2^2}, \qquad (67)$$

which is depicted in Fig. 4 for $\tilde{L}_1 + \tilde{L}_2 = 10^{-3}$.



FIG. 4. Critical value for the condensation of the fluid 1, once the fluid 2 is in the superfluid phase, obtained by Eq. (67) for $\tilde{L}_1 + \tilde{L}_2 = 10^{-3}$.

In principle, for large values of γ_2 it suffices a small μ_1 isospin chemical potential to drive the system 1 in the condensed phase. However, for reasonable values of the NLO LECs, the influence of one condensate on the other is extremely small. The low-energy spectrum in the broken phase consists of two well-defined NGBs.

VI. CONCLUSIONS

We have discussed multicomponent meson superfluids in the χ PT framework. We have derived the relevant χ PT Lagrangian restricting most of the analysis to the global symmetry group given in Eqs. (16) and (17) with $N_f = 2$, corresponding to two fictitious pion gases with different masses and decay constants. In the noninteracting case, if one of the two isospin chemical potentials exceeds the corresponding pion mass the system becomes superfluid. Turning on the interactions the two condensates influence each other. We have considered two possible interaction terms, one that locks the two chiral groups and one that does not lock them.

The Lagrangian term in Eq. (22) leads to the tree-level potential in Eq. (44), with the peculiar interaction term between the phases of the two condensates. Minimizing the potential we have obtained the phase diagram reported in Fig. 2. With increasing locking parameter k, the region in which the simultaneous condensation is realized becomes larger. It seems that one can arbitrarily shrink the normal phase region by increasing the value of k. However, the locking turns one low-energy mode becomes in a pseudo NGB with dispersion law given in Eq. (65). For k > 1 the mass of the pseudo NGB becomes imaginary and therefore an instability is triggered. The unusual aspect is that even for k > 1 the potential has a well defined minimum, indeed the low-energy radial excitations studied in Sec. A 1 have a well-defined mass. Since no other homogenous phase is energetically favored, this suggests that there exists an energetically favored inhomogeneous phase, where the two gases do not coexist any longer. Though not rigorously proved, this seems an educated guess, also because of the analysis of the simplified model discussed in Sec. IV. It is not obvious to us that this inhomogeneous phase can be treated by a Ginzburg-Landau expansion [75], or any other improved version [76], because in these approaches one expects the appearance of an inhomogeneous phases when the mean-field analysis indicates a first-order phase transition. Instead, in the present case, the tree-level analysis does not show any phase transition or any instability: the only sign of an odd behavior is in the spectrum of the

pseudo-NGB mode. The Lagrangian term in Eq. (27), which does not lock the two global symmetries, is also interesting, because it induces a nontrivial interaction between the two condensates. However, in χ PT this term can only arise at the NLO in the chiral expansion, thus we expect that it is strongly suppressed. The tree-level interaction potential is reported in Eq. (66): since it is independent of n_1 and n_2 , it is clear that in this case the two condensates are free to oscillate and are not locked. The low-energy modes consist of two NGBs which do not show any singular behavior. Upon minimizing the potential in Eq. (66) we obtain the phase diagrams reported in Fig. 3.

The present work can be extended in different ways. As already anticipated, it paves the way for the discussion of a two-component system of pions and kaons. We plan to develop this study shortly. It would also be interesting to realize the locking instability in two-component ultracold atoms system.

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APPENDIX A: LOW-ENERGY EXPANSION

In the following we recap and slightly extend the discussion of the low-energy modes of the pion condensed phase of [46,12], including higher order terms.

1. Radial field

Expanding the radial field around the stationary value as $\rho = \bar{\rho} + \chi$ and neglecting the angular fluctuations we obtain from Eq. (6)

$$\mathcal{L}_{\chi} = \frac{f_{\pi}^{2}}{2} \partial^{\mu} \chi \partial_{\mu} \chi + f_{\pi}^{2} m_{\pi}^{2} \left(\frac{1 - \gamma^{4}}{2\gamma^{2}} \chi^{2} - \frac{1}{2} \sqrt{\frac{\gamma^{4} - 1}{\gamma^{4}}} \chi^{3} + \frac{4\gamma^{4} - 7}{24\gamma^{2}} \chi^{4} \right),$$
(A1)

where the $\mathcal{O}(\chi^5)$ terms and higher have been suppressed. It is convenient to rescale the field with $\chi \to \chi/f_{\pi}$ to put the kinetic term in the canonical form, obtaining

$$\mathcal{L}_{\chi} = \frac{1}{2} \partial^{\mu} \chi \partial_{\mu} \chi - \frac{1}{2} m_{\chi}^2 \chi^2 - g_{3\chi} \chi^3 + g_{4\chi} \chi^4, \quad (A2)$$

where the mass and self-couplings are given by

$$m_{\chi} = m_{\pi}\gamma\sin\bar{\rho},\tag{A3}$$

$$g_{3\chi} = \frac{m_\pi^2 \sin \bar{\rho}}{2f_\pi},\tag{A4}$$

$$g_{4\chi} = \frac{m_{\pi}^2}{f_{\pi}^2} \frac{4\gamma^4 - 7}{24\gamma^2}.$$
 (A5)

We notice that the only nonvanishing term at the phase transition point is the one proportional to χ^4 . Actually, it can be easily proven that any term proportional to χ^{2n+1} vanishes at $\gamma = 1$, because in the unbroken phase the system is symmetric for $\rho \rightarrow -\rho$. Close to the phase transition point, the radial fluctuations can be considered as a self-interacting system of bosons with vanishing mass and cubic interaction but nonvanishing quartic interaction. This Lagrangian for the radial fluctuation is valid in the whole broken phase.

2. Bogolyubov mode

The Lagrangian of the angular field is given by

$$\mathcal{L} = \frac{\hat{f}_{\pi}^2}{2} \partial^{\mu} \hat{\boldsymbol{\varphi}} \cdot \partial_{\mu} \hat{\boldsymbol{\varphi}}, \qquad (A6)$$

with $\hat{f}_{\pi} = f_{\pi} \sin \bar{\rho}$ playing the role of an effective decay constant. Since $\hat{\varphi}$ is a unit vector, we can parametrize it by a Bogolyubov mode α as follows:

$$\hat{\boldsymbol{\varphi}} = (\cos\alpha, \sin\alpha), \tag{A7}$$

leading to

$$\mathcal{L} = \frac{\hat{f}_{\pi}^2}{2} \partial^{\mu} \alpha \partial_{\mu} \alpha, \qquad (A8)$$

which is the Lagrangian of a free scalar field, α . It can be cast in the canonical form by $\alpha \rightarrow \alpha/\hat{f}_{\pi}$. The Bogolyubov field can only feel the medium effect by the interactions with the χ field, as will be discussed below. We note that the NLO chiral terms would be proportional to higher powers of momentum, therefore this is the relevant Lagrangian only for $p^2/\hat{f}_{\pi}^2 \ll 1$. For this reason, this low-energy expansion is not valid close to the phase transition point, corresponding to $\gamma = 1$, where \hat{f}_{π} vanishes and thus all the terms of the effective Lagrangian are equally important. Since the momentum scale is dictated by the temperature of the system, one has to consider the $T/\hat{f}_{\pi} \ll 1$ case.

3. Mixed terms and dispersion laws

The mixed terms can be obtained from the interaction terms in Eq. (6) and considering that upon substituting Eq. (A7) we have the compact expression

$$\epsilon_{3ik}\hat{\varphi}_i\partial_0\hat{\varphi}_k = \partial_0\alpha,\tag{A9}$$

in terms of the Bogolyubov field α . Thus, up to the fourth order in the fields, the mixed interaction terms are

$$\mathcal{L}_{I}^{\chi\alpha} = -g_{2,1}\chi\partial_{0}\alpha + g_{3,1}\chi^{2}\partial_{0}\alpha + g_{3,2}\chi\partial_{\mu}\alpha\partial^{\mu}\alpha \quad (A10)$$

$$+g_{4,1}\chi^3\partial_0\alpha + g_{4,2}\chi^2\partial_\mu\alpha\partial^\mu\alpha, \tag{A11}$$

with the couplings given by:

$$g_{2,1} = \frac{2m_{\pi}}{\gamma} \qquad g_{3,1} = \frac{\gamma^4 - 2}{\gamma^3 \hat{f}_{\pi}} m_{\pi},$$
 (A12)

$$g_{3,2} = \frac{1}{\gamma^2 \hat{f}_{\pi}} \qquad g_{4,1} = \frac{4m_{\pi}}{3f_{\pi}^2 \gamma},$$
 (A13)

$$g_{4,2} = \frac{2 - \gamma^4}{2\gamma^4 \hat{f}_\pi^2},\tag{A14}$$

where the first subscript indicates the total number of fields and the second one the number of α fields.

The quadratic Lagrangian can be written as

$$\mathcal{L} = \frac{1}{2} \partial^{\mu} \chi \partial_{\mu} \chi - \frac{1}{2} m_{\chi}^{2} \chi^{2} + \frac{1}{2} \partial_{\mu} \alpha \partial^{\mu} \alpha - g_{2,1} \chi \partial_{0} \alpha, \quad (A15)$$

where the mixing term allows oscillations between the radial and the angular fields. Integrating out the radial fluctuations one obtains the massless, phononlike, dispersion law

$$E_{\rm ph} = c_s p, \tag{A16}$$

where

$$c_s = \sqrt{\frac{m_{\chi}^2}{m_{\chi}^2 + g_2^2}} = \sqrt{\frac{\gamma^4 - 1}{\gamma^4 + 3}}$$
 (A17)

describes the pressure oscillations propagating at the sound speed.

Alternatively, one can diagonalize the quadratic Lagrangian, obtaining the dispersion laws

$$E_{\pm} = \sqrt{p^2 + \frac{m_{\rm eff}^2}{2} \pm \sqrt{\left(\frac{m_{\rm eff}^2}{2}\right)^2 + g_2^2 p^2}},\qquad(A18)$$

where

$$m_{\rm eff}^2 = m_{\chi}^2 + g_2^2 = m_{\pi}^2 \frac{\gamma^4 + 3}{\gamma^2},$$
 (A19)

which agree with the expressions reported in [46]. In conclusion, the low-energy modes of a single-component pion gas correspond to a NGB with dispersion law in Eq. (A16) (in the limit $p/m_{\chi} \rightarrow 0$) and to a massive mode with mass $m_{\rm eff}$.

APPENDIX B: LHY CORRECTION

Close to the phase transition to the broken phase, the pressure of the single-component pion gas can be approximated with the expression in Eq. (11). Therefore, the χ PT analysis gives a correction to the mean-field value proportional to n^3 , where $n = n_I/(f_\pi^2 m_\pi)$. However, in the context of condensed matter physics, an additional contribution, due to the vacuum energy of the NGBs, is known to play an important role in certain regimes. This contribution is known as the Lee-Huang-Yang (LHY) term, first evaluated for a hard sphere Bose gas in [77]. The LHY term is proportional to $n^{5/2}$ and is therefore the leading correction to the mean-field results close to the phase transition point.

For a single-component pion gas, one can easily obtain the LHY correction using the mapping developed in [47] between the condensed pion gas in χ PT and the Gross-Pitaevskii (GP) Hamiltonian

$$\mathcal{H}_{\rm GP} = \psi^* \frac{\nabla^2}{2M} \psi - \frac{g}{2} |\psi^* \psi|^2, \qquad (B1)$$

where $M = \mu_I$, and

$$g = \frac{4\gamma^2 - 1}{12f_{\pi}^2\gamma^2} = g_0\left(1 + \frac{n}{6}\right) + \mathcal{O}(n^3), \qquad (B2)$$

where $g_0 = 1/(4f_{\pi}^2)$ is the coupling constant at the phase transition point. The LHY correction to the pressure close to the phase transition point is given by

$$\epsilon_{\rm GP,LHY} = \frac{M^{3/2}}{15\pi^2} (gn_I)^{5/2} \propto m_\pi^4 n^{5/2},$$
 (B3)

with the particular dependence on *n* indicating that this is a nontrivial beyond mean-field effect. The LHY contribution is the first one in the series expansion $n_I a^3$, where $a = gM/(4\pi)$ is the s-wave scattering length. Close to the transition point and using the values of the coupling constant and of the mass of the GP expansion, we find that $n_I a^3 \ll 1$ that means the diluition condition for any $\gamma \in [1, 2]$. However, the evaluation of the LHY term by Eq. (B3) assumes that the GP expansion is reliable, implying that $1 \le \gamma \ll 2$.

For a general evaluation of he LHY correction in the χ PT context, we consider the vacuum contribution of the NGB

$$\epsilon_{\rm LHY} \propto \frac{1}{2\pi^2} \int_0^\Lambda dp \; p^2 E_{\rm ph},$$
 (B4)

where $E_{\rm ph}$ is the dispersion law of the NGB obtained integrating out the radial fluctuations, see Eq. (A16). The hard cutoff, Λ , takes into account that the NGB describes the low-energy fluctuations below the mass scale, m_{χ} , of the radial field, see Eq. (A3). Taking for simplicity $\Lambda = m_{\chi}$, considering the expression of the speed of sound in Eq. (A17), and that, close to the phase transition, $\gamma \approx 1 + n/4$, see Eq. (13), we find

$$\epsilon_{\rm LHY} \propto m_{\pi}^4 n^{5/2},$$
 (B5)

in agreement with Eq. (B3). In Fig. 5 we compare the isospin number density evaluated in χ PT (solid red line), with that obtained including the LHY correction (dashed blue line), as well as with the results of the lattice simulations of Refs. [13,21,22] using the same value of their pion mass, $m_{\pi} = 135$ MeV, and of the pion decay constant, $f_{\pi} = 133/\sqrt{2}$.

The χ PT results systematically underestimate the number density. With the inclusion of the LHY term the agreement slightly improves. The χ PT + LHY pressure



FIG. 5. Isospin number density as a function of the isospin chemical potential for a single-component pion gas. The solid red line corresponds to the LO χ PT result. The dashed blue line is obtained adding the LHY contribution. The dots correspond to the lattice results of [13,21,22].

is always larger than the χ PT one. However, the difference between the two is extremely small.

Generalizing the previous discussion to the twocomponent pion gases with the interaction term in Eq. (22), it is clear that there are two relevant low-energy contributions. One from the NGB, and one from the pseudo NGB. Since the latter becomes tachyonic for k > 1, the LHY contribution is ill-defined. Again, this is a signal that the mean-field approximation breaks down for k > 1, and thus the evaluation of the pressure of the system given by the expression in Eq. (51) is incorrect.

- [1] N. Cabibbo and G. Parisi, Phys. Lett. 59B, 67 (1975).
- [2] K. Rajagopal and F. Wilczek, arXiv:hep-ph/0011333.
- [3] M. G. Alford, A. Schmitt, K. Rajagopal, and T. Schafer, Rev. Mod. Phys. 80, 1455 (2008).
- [4] R. Anglani, R. Casalbuoni, M. Ciminale, N. Ippolito, R. Gatto, M. Mannarelli, and M. Ruggieri, Rev. Mod. Phys. 86, 509 (2014).
- [5] A. B. Migdal, Zh. Eksp. Teor. Fiz. 61, 2209 (1971).
- [6] R. F. Sawyer, Phys. Rev. Lett. 29, 382 (1972).
- [7] D. J. Scalapino, Phys. Rev. Lett. 29, 386 (1972).
- [8] A. B. Migdal, Phys. Rev. Lett. 31, 257 (1973).
- [9] J. Kogut and J. T. Manassah, Phys. Lett. 41A, 129 (1972).
- [10] A. B. Migdal, E. Saperstein, M. Troitsky, and D. Voskresensky, Phys. Rep. 192, 179 (1990).
- [11] S. Shapiro and S. Teukolsky, Black Holes, White Dwarfs, and Neutron Stars: The Physics of Compact Objects, A Wiley-Interscience Publication (Wiley, New York, 1983).
- [12] S. Carignano, L. Lepori, A. Mammarella, M. Mannarelli, and G. Pagliaroli, Eur. Phys. J. A 53, 35 (2017).
- [13] B. B. Brandt, G. Endrődi, E. S. Fraga, M. Hippert, J. Schaffner-Bielich, and S. Schmalzbauer, Phys. Rev. D 98, 094510 (2018).

- [14] J.O. Andersen and P. Kneschke, arXiv:1807.08951.
- [15] M. G. Alford, A. Kapustin, and F. Wilczek, Phys. Rev. D 59, 054502 (1999).
- [16] J. B. Kogut and D. K. Sinclair, Phys. Rev. D 66, 034505 (2002).
- [17] W. Detmold, K. Orginos, and Z. Shi, Phys. Rev. D 86, 054507 (2012).
- [18] W. Detmold, K. Orginos, M. J. Savage, and A. Walker-Loud, Phys. Rev. D 78, 054514 (2008).
- [19] P. Cea, L. Cosmai, M. D'Elia, A. Papa, and F. Sanfilippo, Phys. Rev. D 85, 094512 (2012).
- [20] G. Endrödi, Phys. Rev. D 90, 094501 (2014).
- [21] B. B. Brandt, G. Endrodi, and S. Schmalzbauer, Phys. Rev. D 97, 054514 (2018).
- [22] B. B. Brandt, G. Endrodi, and S. Schmalzbauer, arXiv: 1811.06004.
- [23] Y. Nambu and G. Jona-Lasinio, Phys. Rev. 122, 345 (1961).
- [24] S. P. Klevansky, Rev. Mod. Phys. 64, 649 (1992).
- [25] M. Buballa, Phys. Rep. 407, 205 (2005).
- [26] D. Toublan and J. Kogut, Phys. Lett. B 564, 212 (2003).
- [27] A. Barducci, R. Casalbuoni, G. Pettini, and L. Ravagli, Phys. Rev. D 69, 096004 (2004).

- [28] A. Barducci, R. Casalbuoni, G. Pettini, and L. Ravagli, Phys. Rev. D 71, 016011 (2005).
- [29] D. Ebert and K.G. Klimenko, Eur. Phys. J. C 46, 771 (2006).
- [30] D. Ebert and K. G. Klimenko, J. Phys. G 32, 599 (2006).
- [31] L.-Y. He, M. Jin, and P.-F. Zhuang, Phys. Rev. D 71, 116001 (2005).
- [32] L. He, Phys. Rev. D 82, 096003 (2010).
- [33] J. O. Andersen, W. R. Naylor, and A. Tranberg, Rev. Mod. Phys. 88, 025001 (2016).
- [34] P. Adhikari, J. O. Andersen, and P. Kneschke, Phys. Rev. D 95, 036017 (2017).
- [35] P. Adhikari, J. O. Andersen, and P. Kneschke, Phys. Rev. D 98, 074016 (2018).
- [36] S. Weinberg, Physica (Amsterdam) 96A, 327 (1979).
- [37] A. Pich, in Probing the Standard Model of Particle Interactions. Proceedings, Summer School in Theoretical Physics, NATO Advanced Study Institute, 68th session, Les Houches, France, 1997. Pt. 1, 2 (Elsevier, The Netherlands, 1998), pp. 949–1049, http://inspirehep.net/record/471518.
- [38] B. R. Holstein, Nucl. Phys. A689, 135 (2001).
- [39] H. Georgi, Weak Interactions and Modern Particle Theory, Dover Books on Physics Series (Dover Publications, New York, 2009).
- [40] H. Leutwyler, Ann. Phys. (N.Y.) 235, 165 (1994).
- [41] J. Gasser and H. Leutwyler, Ann. Phys. (N.Y.) 158, 142 (1984).
- [42] S. Scherer, Adv. Nucl. Phys. 27, 277 (2003).
- [43] G. Ecker, Prog. Part. Nucl. Phys. 35, 1 (1995).
- [44] H. Leutwyler, Helv. Phys. Acta 70, 275 (1997).
- [45] J. Kogut and D. Toublan, Phys. Rev. D 64, 034007 (2001).
- [46] A. Mammarella and M. Mannarelli, Phys. Rev. D 92, 085025 (2015).
- [47] S. Carignano, A. Mammarella, and M. Mannarelli, Phys. Rev. D 93, 051503 (2016).
- [48] S. Scherer and M. R. Schindler, arXiv:hep-ph/0505265.
- [49] D. Son and M. A. Stephanov, Phys. Rev. Lett. 86, 592 (2001).
- [50] J. B. Kogut, M. A. Stephanov, and D. Toublan, Phys. Lett. B 464, 183 (1999).
- [51] J. B. Kogut, M. A. Stephanov, D. Toublan, J. J. M. Verbaarschot, and A. Zhitnitsky, Nucl. Phys. B582, 477 (2000).
- [52] S. Hands, I. Montvay, S. Morrison, M. Oevers, L. Scorzato, and J. Skullerud, Eur. Phys. J. C 17, 285 (2000).
- [53] J. B. Kogut, D. K. Sinclair, S. J. Hands, and S. E. Morrison, Phys. Rev. D 64, 094505 (2001).
- [54] T. Brauner, Mod. Phys. Lett. A 21, 559 (2006).

- [55] V. V. Braguta, E. M. Ilgenfritz, A. Yu. Kotov, A. V. Molochkov, and A. A. Nikolaev, Phys. Rev. D 94, 114510 (2016).
- [56] P. Adhikari, S. B. Beleznay, and M. Mannarelli, Eur. Phys. J. C 78, 441 (2018).
- [57] A. Ramos, J. Schaffner-Bielich, and J. Wambach, Lect. Notes Phys. 578, 175 (2001).
- [58] J. Tuoriniemi, J. Martikainen, E. Pentti, A. Sebedash, S. Boldarev, and G. Pickett, J. Low Temp. Phys. **129**, 531 (2002).
- [59] J. Rysti, J. Tuoriniemi, and A. Salmela, Phys. Rev. B 85, 134529 (2012).
- [60] E. A. Cornell and C. E. Wieman, Rev. Mod. Phys. **74**, 875 (2002).
- [61] M. Lewenstein, A. Sanpera, and V. Ahufinger, Ultracold Atoms in Optical Lattices: Simulating Quantum Many-Body Systems (Oxford University Press, New York, 2012).
- [62] M. Inguscio and L. Fallani, Atomic Physics: Precise Measurements and Ultracold Matter (Oxford University Press, New York, 2013).
- [63] I. Ferrier-Barbut, M. Delehaye, S. Laurent, A. T. Grier, M. Pierce, B. S. Rem, F. Chevy, and C. Salomon, Science 345, 1035 (2014).
- [64] L. Lepori, A. Trombettoni, and W. Vinci, Europhys. Lett. 109, 50002 (2015).
- [65] M. G. Alford, K. Rajagopal, and F. Wilczek, Nucl. Phys. B537, 443 (1999).
- [66] P.F. Bedaque and T. Schafer, Nucl. Phys. A697, 802 (2002).
- [67] D. Kaplan and S. Reddy, Phys. Rev. D 65, 054042 (2002).
- [68] A. Haber, A. Schmitt, and S. Stetina, Phys. Rev. D 93, 025011 (2016).
- [69] A. Haber and A. Schmitt, Phys. Rev. D 95, 116016 (2017).
- [70] A. F. Andreev and E. P. Bashkin, Zh. Eksp. Teor. Fiz. 69, 164 (1975).
- [71] A. Smerzi, A. Trombettoni, P.G. Kevrekidis, and A.R. Bishop, Phys. Rev. Lett. 89, 170402 (2002).
- [72] B. D. Esry, C. H. Greene, J. P. Burke, Jr., and J. L. Bohn, Phys. Rev. Lett. 78, 3594 (1997).
- [73] D. S. Petrov, Phys. Rev. Lett. 115, 155302 (2015).
- [74] S. Blundell, *Magnetism in Condensed Matter*, Oxford Master Series in Condensed Matter Physics (Oxford University Press, New York, 2001).
- [75] V. L. Ginzburg and L. D. Landau, Zh. Eksp. Teor. Fiz. 20, 1064 (1950).
- [76] S. Carignano, M. Mannarelli, F. Anzuini, and O. Benhar, Phys. Rev. D 97, 036009 (2018).
- [77] T. D. Lee, K. Huang, and C. N. Yang, Phys. Rev. 106, 1135 (1957).